

Status of LUX-ZEPLIN

WIMP Search 2024

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WIMP Direct Detection - Briefly

□ Well motivated candidate is a "WIMPy" thermal relic

- □ MeV 100 TeV scale particle (cosmological bounds)
- Weak scale interactions leads to correct relic density
- Theoretical models independently predict particles with the correct properties

Direct Detection

- Solar System travels through Milky Way's Dark Matter Halo
- □ Very high flux of WIMPs pass through the earth
- Challenge Interaction is an exceptionally rare process
- Challenge Background radiation is everywhere



Fill a detector with your favorite material and wait for WIMPs to scatter off it



LZ Collaboration - 38 Institutions 250 scientists, engineers, and technical staff

- Black Hills State University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- King's College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Los Angeles
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Sydney
- University of Texas at Austin
- University of Wisconsin, Madison
- University of Zürich
- US Europe Asia Oceania







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Thanks to our sponsors and participating institutions!

LZ at SURF

LUX-ZEPLIN at Sanford Underground Research Facility

- SURF Homestake Mine, Lead, South Dakota
- □ ~1 mile underground, Davis Cavern
- **\Box** Rock overburden reduces cosmic ray muon flux by O(10⁶)





LZ - Experiment for Direct Detection of WIMP Dark

Calibration Source Deployment Tubes (3 Total)

17T Gd-loaded liquid scintillator

> 120 Outer Detector PMTs

2T LXe Skin Veto

> 131 Skin PMTs

60,000 gallons of ultrapure water

494 LXe PMTs

7T Active LXe Target

Neutron Calibration Conduit (2 total)

LXe TPC Principles



Why LXe?:

- Highest charge & light yields of all noble elements
- Commercially available & easily purified
- $\label{eq:scalable} \square \quad \text{Scalable} \to \text{potential for large target mass}$



Time Projection Chamber:

- □ S2 hit pattern, Drift Time (xy, z)
- Energy deposited reconstructed from S1, S2
- Particle discrimination S1:log(S2)

LXe TPC Principles







LXe Skin



- **2 tonne** of LXe surrounding the TPC
- 131 1" or 2" PMTs
- Anti-coincidence veto for **γ-rays** with **78±5% efficiency**
- Reduction of important ER background rates
 E.g. ¹²⁷Xe decay via electron capture





Courtesy of Jack Bargemann APS April 2023 11

Outer Detector



- 17 tonne of Gd-loaded liquid scintillator (120 8" PMTs)
- Infinite volume of Gd-LS has a 30us capture time for a thermal neutron
- LZ has ~90% Single Scatter neutron veto efficiency
- ~50Hz of background above 100-200 keV threshold
 Gd alphas, ¹⁴C betas

Waveform example of a tagged neutron multiple scatter:





WIMP Search 2024

WIMP Search 2024



- 220 live-days, one calendar year March '23 to March '24
- Milestones:
 - Bias mitigation ("salting") began July 3rd
 - □ Circulation change July 12th

Calibrations

- ER calibration: 156k evts radiolabeled methane
 ³H (18.6 keV β⁻) & ¹⁴C (156 kev β⁻)
- □ NR calibration: 11k neutron evts!
 - Collimated DD 2.45 MeV & CSD AmLi (α,n)
- **Q** 99.8% discrimination of β^{-} under flat NR band median
- Injected ^{83m}Kr, ^{131m}Xe sources, spatial & temporal corrections
 (xy, z) resolution at 1σ (<1cm, <1mm)
- Electron lifetime >8ms (depends strongly on LXe purity)
 max e⁻ drift time ~1ms.

Science Run	C/G/A Voltage [kV]	Drift Field [V/cm]	Analysis live time [days]	g1 (phd/photon)	g2 (phd/e⁻)
WS2022	-32/-4/+4	193	60	0.114 ± 0.002	47 ± 1.1
WS2024	-18/-4/+3.5	97	220	0.112 ± 0.002	34 ± 0.9





Bias Mitigation

- General Salting" fake signal injected randomly
- □ Manufactured from sequestered ER & NR calibration data
- □ Unknown number of salt events
 - Rate capped by WS2022 cross-section limit
- Exponential WIMP recoil spectrum
 - + flat energy spectrum to cover high energy NR



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Backgrounds Overview

- **Dissolved** β emitters:
 - ²¹⁴Pb (²²²Rn), ²¹²Pb (²²⁰Rn), ⁸⁵Kr, ¹³⁶Xe (ββ)
- Dissolved EC decays (x-rays/Auger electrons)
 ¹²⁷Xe & ¹²⁵Xe produced by activation from neutron calibration
 - □ ¹²⁴Xe (double EC), 0.095% natural abundance
- **Given Solar** ν 's: ⁸B (NR), pp+⁷Be (ER)
- Detector ER, γ emitters from detector materials
 ²³⁸U, ²³²Th, ⁴⁰K, ⁶⁰Co decay chains
- **Neutrons** from USF and (α, n) in detector materials
- Accidentals random coincidence of isolated S1 and S2s



Liquid Xenon Flow

- Fine control of LXe circulation and temperature allows control of LXe flow state
- □ **High-mixing** state: Turbulent flow, better circulation across TPC, uniform injection of calibration sources
- □ **Low-mixing** state: Slower, laminar flow. Lower activity in FV, predictable "streamlines".





160

140

120 100

80

60

40 20

Z [cm]

Radon Tag



- "Naked" ²¹⁴Pb β make up majority of ER background
- Combine LXe flow maps & electric field maps to predict volumes likely to contain charged or neutral ²¹⁴Pb.
- Reduces ²¹⁴Pb to **1.8 ± 0.3 \muBq/kg** in untagged sample **3.9 ± 0.4 \muBq/kg in total exposure**
- Tagged and untagged samples both fitted for WS2024

Sample	% ²¹⁴ Pb	% Exposure
Tagged	60 ± 4	15
Untagged	40 ± 4	85

Electron Capture Backgrounds

- □ ¹²⁵Xe & ¹²⁷Xe decay by electron capture (EC)
- Produced by activation from neutrons calibration and cosmogenics
 Rate in WS2024 << WS2022
- L-shell electron capture (5.2 keV) is a WS background
- Electron capture has a field-dependent suppressed charge yield
- \Box Higher ionization density \rightarrow increased recombination
- LZ WS2022 (193 V/cm): Q₁/Q₈=0.88±0.01*
- LZ WS2024 (97 V/cm): $Q_L/Q_\beta = 0.87 \pm 0.03^*$
- Charge suppression first measured in XELDA
 - PRD 104, 112001 ('21)

*Preliminary, dedicated publication in progress



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¹²⁴Xe Double Electron Capture

- ¹²⁴Xe 2vECEC T_{1/2}=(1.09±0.14^{stat}±0.05^{sys})×10²² years
 Worlds longest directly measured half-life
- In-situ T_{1/2} measurement of KK (64.3 keV) captures $(65 \pm 5)\%$
- LL (10 keV) and LM (6 keV) are WS backgrounds
- □ LM modelled with same as single L-shell charge suppression □ $Q_{LM}/Q_{\beta} = Q_{L}/Q_{\beta}$
- For LL expect further charge yield suppression due to higher ionization density.
- □ Vary Q_{LL}/Q_{β} in fitting of our data 0.65 < Q_{LL}/Q_{β} < 0.87

2x ionization density

Best fit to WS2024 data: $Q_{LL}/Q_{\beta} = 0.70 \pm 0.04$



 Q_1/Q_R

Accidentals Background

- Accidental coincidence of uncorrelated isolated S1 and S2 pulses
- LZ's accidentals model:
 - **Rate:** Measure rate of unphysical drift time (UDT) single scatters. Analysis cut efficiencies analysed using manufactured accidental events.
 - Shape: Manufactured by combining isolated S1 and S2 waveforms and applying analysis cut efficiencies.
- Analysis cuts have an 99.5% rejection power





Data Analysis

- Region of interest (ROI)
 - \Box 3 < S1c < 80 photons detected (phd)
 - □ S2 > 14.5 electrons (645 phd)
 - $\Box \log_{10}(S2c) < 4.5$
- Fiducial volume: azimuthally & drift time-dependent
 - < 0.01 charge loss ("wall") events</p>
 - □ 5.5 ± 0.2 tonne mass
- □ Single scatter
- □ Anti-coincidence Skin & OD vetoes
- Cuts on S1 & S2 parameters to remove accidentals



WS2024 Data - Salted

□ Final exposure:

□ 220 livedays * 5.5 tonnes = 3.3 tonne years



WS2024 Data - Unsalting

- □ Final exposure:
 - □ 220 livedays * 5.5 tonnes = 3.3 tonne years
- 7 salt events present after all cuts
 - □ 8 injected in WS2024
 - Consistent with evaluated signal efficiency



WS2024 Data - Unsalted

- □ Final exposure:
 - 220 livedays * 5.5 tonnes = 3.3 tonne years
- 7 salt events present after all cuts
 - □ 8 injected in WS2024
 - Consistent with evaluated signal efficiency
- 1220 events remain after unsalting
- No changes to model required post-unsalting
- Next step: Statistical inference of this data



Radon Tagged vs Untagged



Exposure in each sample [tonne years]

High Mixing	Radon Tag	Radon	Radon	OD/Skin	WS2022
State	Inactive	Tagged	Untagged	Vetoed	
0.6	0.6	0.3	1.8	n/a	0.9

- Likelihood fit contains six samples
- □ WS2024 represented by first four, totalling 3.3 tonne years
- OD/Skin vetoed full WS2024 3.3 tonne years constraint on neutron background
- \Box WS2022 unmodified since original publication \rightarrow Maximise exposure

WS2024 Fit Results

		20 88
Source	Pre-fit Constraint	Fit Result
²¹⁴ Pb β s	743 ± 88	733 ± 34
85 Kr + 39 Ar β s + det. γ s	162 ± 22	161 ± 21
Solar ν ER	102 ± 6	102 ± 6
212 Pb + 218 Po β s	62.7 ± 7.5	63.7 ± 7.4
Tritium+ ¹⁴ C β s	58.3 ± 3.3	59.7 ± 3.3
136 Xe $2\nu\beta\beta$	55.6 ± 8.3	55.8 ± 8.2
124 Xe DEC	19.4 ± 3.9	21.4 ± 3.6
127 Xe + 125 Xe EC	3.2 ± 0.6	2.7 ± 0.6
Accidental coincidences	2.8 ± 0.6	2.6 ± 0.6
Atm. ν NR	0.12 ± 0.02	0.12 ± 0.02
$^{8}\mathrm{B}+hep \ \nu \ \mathrm{NR}$	0.06 ± 0.01	0.06 ± 0.01
Detector neutrons	_	$0.0^{+0.2}$
$40 \text{ GeV}/c^2 \text{ WIMP}$	_	$0.0^{+0.6}$
Total	1210 ± 91	1203 ± 42



□ Strong agreement with background only hypothesis



WS2024 Only Sensitivity



- Two-sided profile likelihood ratio test statistic
- Power constrained at -1σ as per recommended conventions EPJC 81, 907 ('21)
- Under-fluctuation in sensitivity results from arrangement of accidentals background events

WS2024 + WS2022 Sensitivity



- Two-sided profile likelihood ratio test statistic
- Power constrained at -1σ as per recommended conventions
 EPJC 81, 907 ('21)
- Further under-fluctuation from
 WS2022 ER background
- □ Best limit from combined analysis of σ_{sl} =2.1×10⁻⁴⁸cm² for 36GeV/c²
- World leading sensitivity!

Spin Dependent Sensitivity



- □ Spin-dependent limit using odd Xe isotopes
 - ¹²⁹Xe, spin 1/2, 26.4% natural abundance
 - □ ¹³¹Xe, spin 3/2, 21.2% natural abundance

Conclusions

- LZ is the world's most sensitive WIMP direct detection experiment
 - □ Total exposure 4.2 tonne-years
 - □ New constraint exceeds previous best constraint by factor >4
- □ Radon tag developed and deployed for the first time
 - □ 60% reduction in main ER background ²¹⁴Pb
- □ First observation of suppressed charge yield from LL-shell captures of ¹²⁴Xe
- LZ will continue to take data until 2028, towards 1000 live days
- □ Many physics searches on the horizon!:
 - \square ⁸B CE*v*NS, low mass WIMPs, ER based searches, $0\nu\beta\beta$