

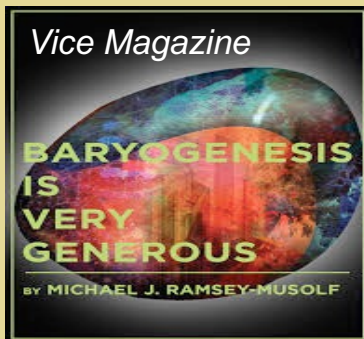
Fundamental Interactions: Inter-frontier Connections

M.J. Ramsey-Musolf

- *T.D. Lee Institute/Shanghai Jiao Tong Univ.*
- *UMass Amherst*
- *Caltech*

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- 微信 : mjrm-china
- <https://michaelramseymusolf.com/>

About MJRM:



Science



Family

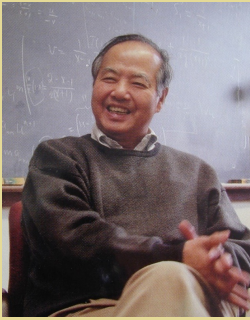


Friends

My pronouns: he/him/his
MeToo

PASCOS 2024 Quy Nhon
July 13, 2024

T. D. Lee Institute / Shanghai Jiao Tong U.



Director



Prof Jie Zhang

A point of convergence of the world's top scientists

A launch pad for the early-career scientists

A world famous source of original innovation



Founded 2016

100+

faculty members from 17 countries and regions, with over 40% of them foreign (non-Chinese) citizens

Theory & Experiment

Particle & Nuclear Physics

Astronomy & Astrophysics

Quantum Science

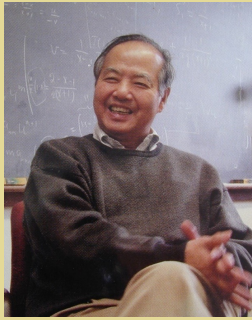
Dark Matter & Neutrino

Laboratory Astrophysics

Topological Quantum Computation

<https://tdli.sjtu.edu.cn/EN/>
<https://www.youtube.com/watch?v=z0awD6q8FTI>

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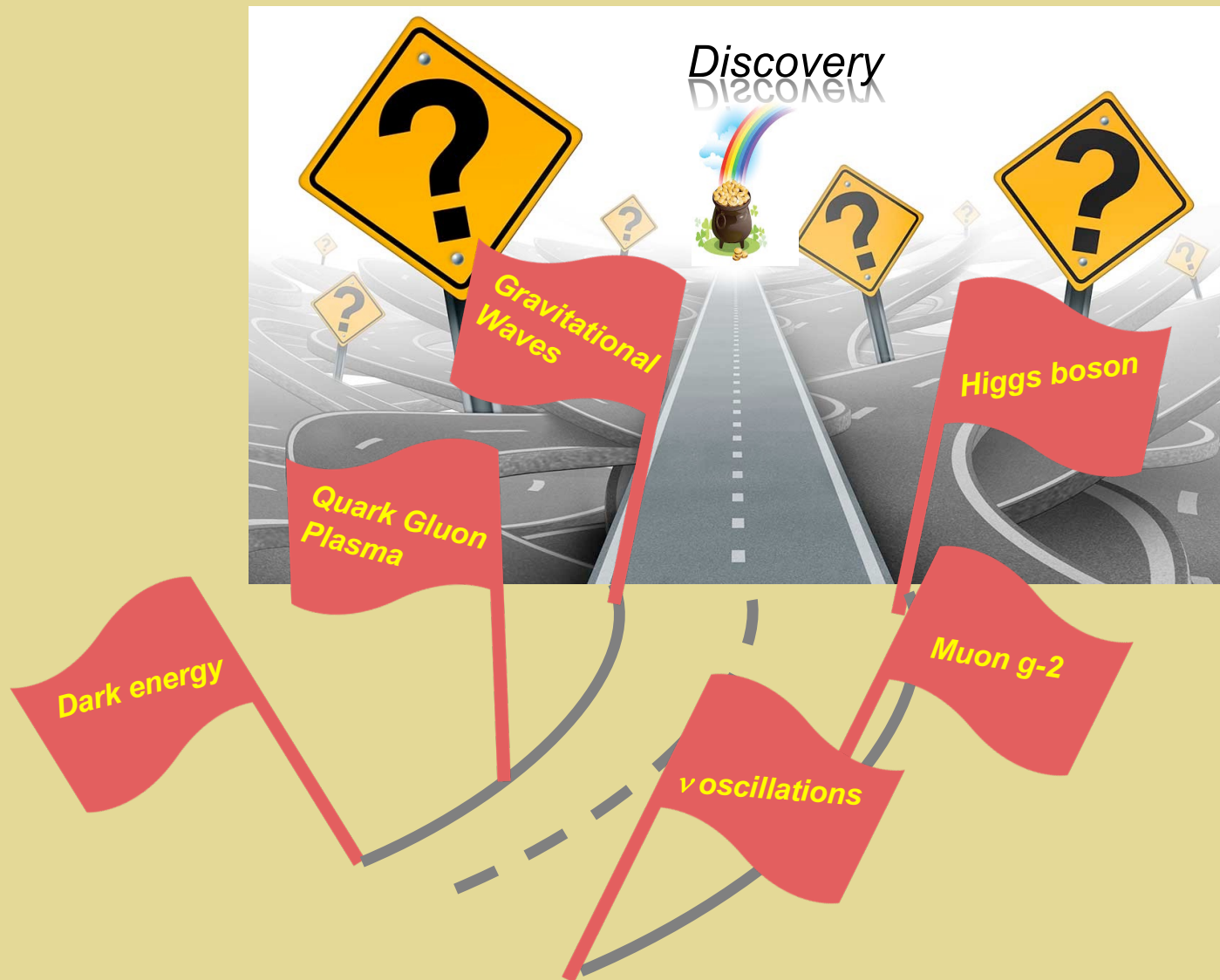
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Goals for this Talk

- *What I won't do: Give a theoretical summary overview*
- *What I will do: Share my perspective on two themes*

Fundamental Physics: Past & Future



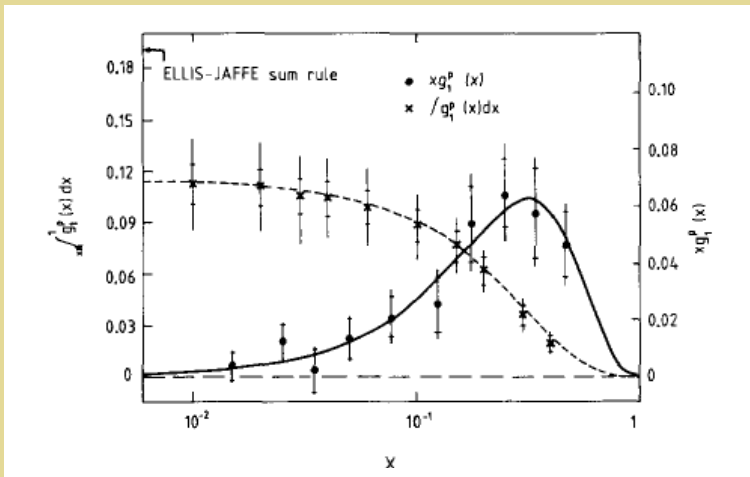
Two Themes

- *Progress will come from focusing on well-posed scientific questions → potential for insights and/or discoveries*

“Historical” Examples

“Spin Crisis”

EMC ‘88



$$\langle S_z \rangle_u = 0.373 \pm 0.019 \pm 0.039 ,$$

$$\langle S_z \rangle_d = -0.254 \pm 0.019 \pm 0.039 ,$$

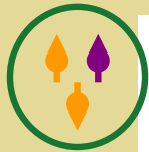
$$\langle S_z \rangle_s = -0.113 \pm 0.019 \pm 0.039 ,$$

$$\langle S_z \rangle_{u+d+s} = 0.006 \pm 0.058 \pm 0.117$$

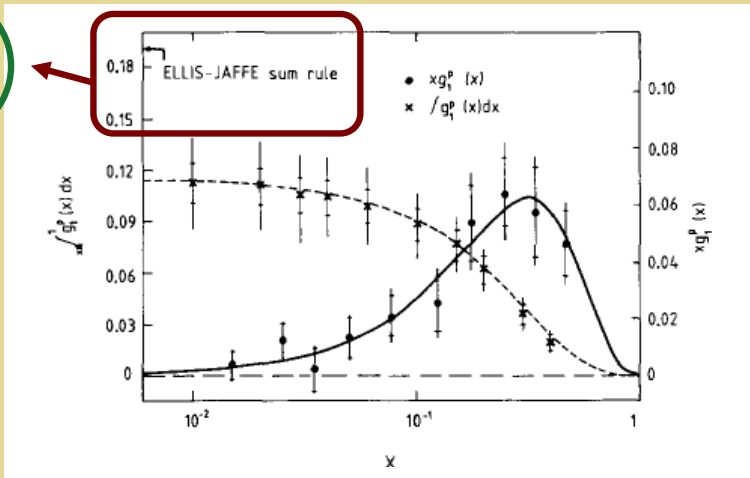
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$$\Delta s = 0$$



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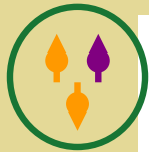
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Large negative strange sea polarization & small $\Delta\Sigma$

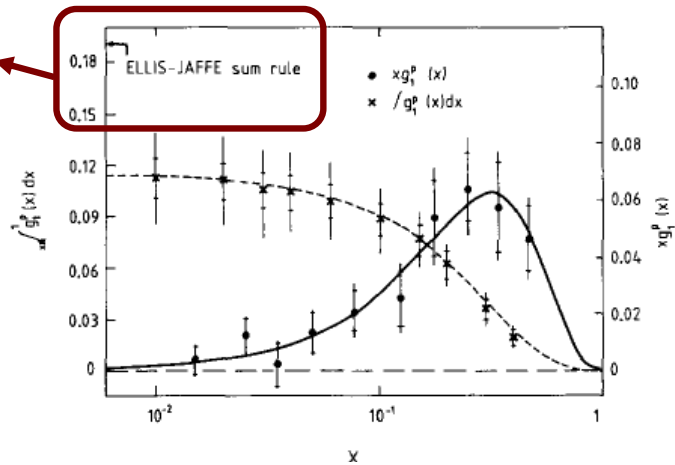
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Large negative strange sea polarization & small $\Delta\Sigma$

What is the s-quark component of the nucleon magnetic moment and charge distribution ?

- Kaplan & Manohar ‘88
- Jaffe ‘89
- McKeown & Beck ‘89
- Holstein & MJRM ‘90

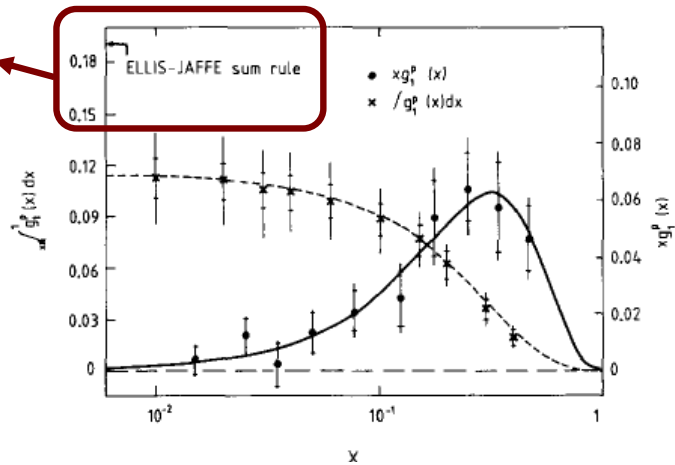
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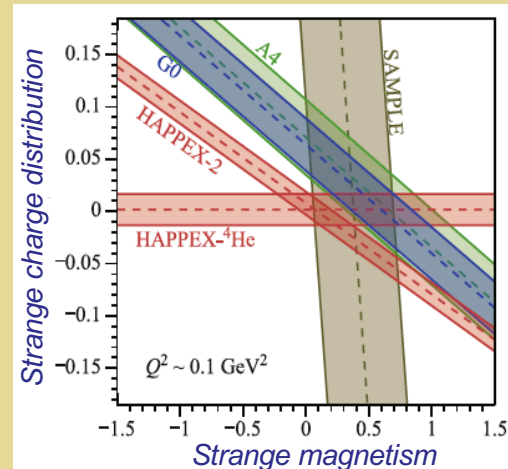
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25 years of PV
ep, eA scattering



Two Themes

- *Progress will come from focusing on well-posed scientific questions → potential for insights and/or discoveries*
- *It's important to think beyond boundaries of funding agency priorities and conventional sub-field categories → rich opportunities from inter-frontier connections*

“Historical” Examples

*Nuclear Science Long Range Plan
Pre-Town Meeting: Los Alamos 2006*

Hamish Robertson

$$V_{\text{PMNS}} =$$

$$\begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \\ \times \text{diag}(1, e^{i\frac{\alpha_{21}}{2}}, e^{i\frac{\alpha_{31}}{2}}) .$$

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

“Atmospheric”

“Reactor”

“Solar”

“Historical” Examples

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$$V_{\text{PMNS}} =$$

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“Atmospheric”

“Reactor”

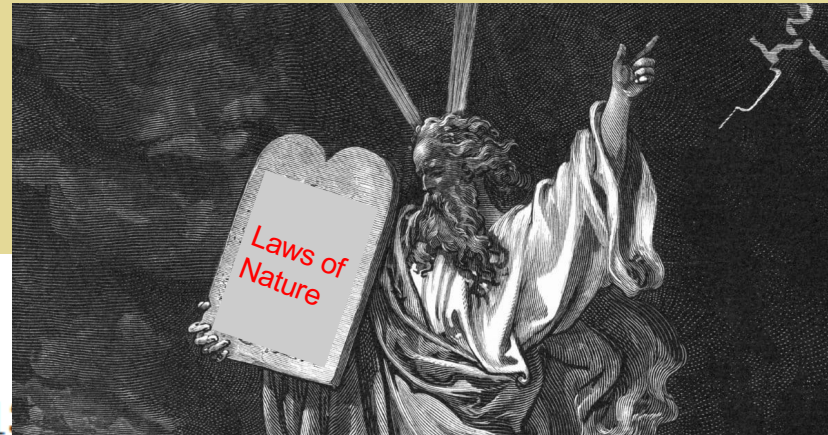
“Solar”

“Historical” Examples

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$$\begin{bmatrix} c_{12}c_{13} & & & \\ -s_{12}c_{23} - c_{12}s_{23}s_{13} & & & \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & & \\ & & & \end{bmatrix} \times \text{diag}(1, e^{i\frac{\alpha_{21}}{2}}, e^{i\frac{\alpha_{31}}{2}}).$$



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“High Energy”
“Atmospheric”

“Reactor”

“Solar”

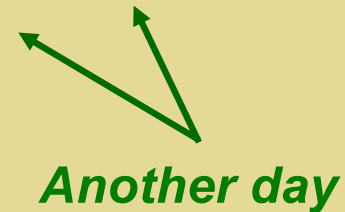
The BSM Dog Race



This talk: I'll focus on a problem I'm particularly immersed in to illustrate → other areas of inquiry equally important & interest

Outline

- I. Questions & Frontiers*
- II. Origin of matter: EW scale*
- III. Outlook*
- IV. Origin of matter & neutrino physics*
- V. Electroweak precision tests*



I. Questions & Frontiers

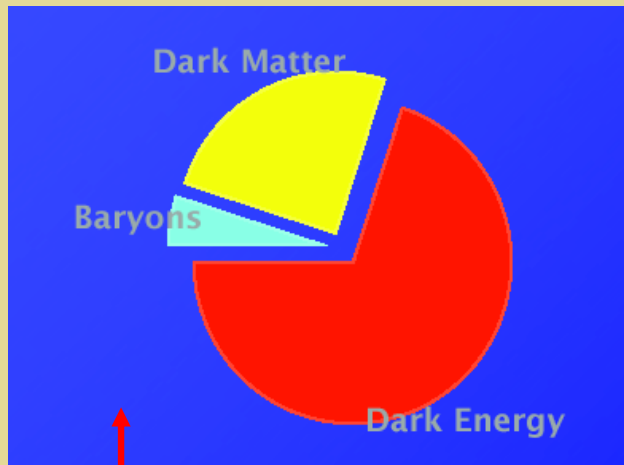
Fundamental Questions

***MUST** answer*

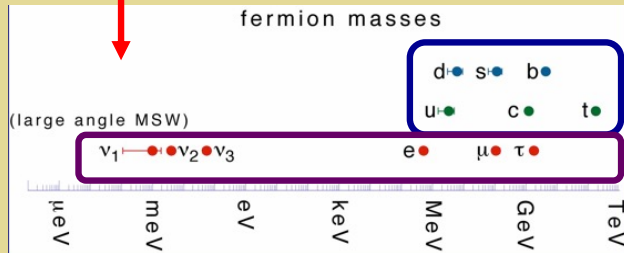
***SHOULD** answer*

Fundamental Questions

MUST answer



?

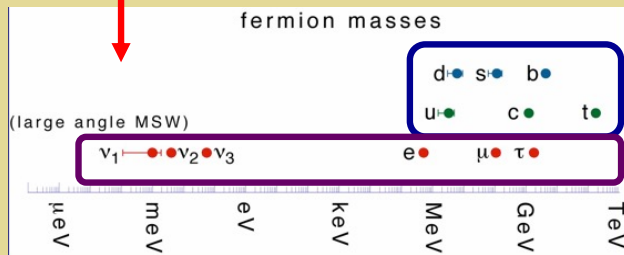
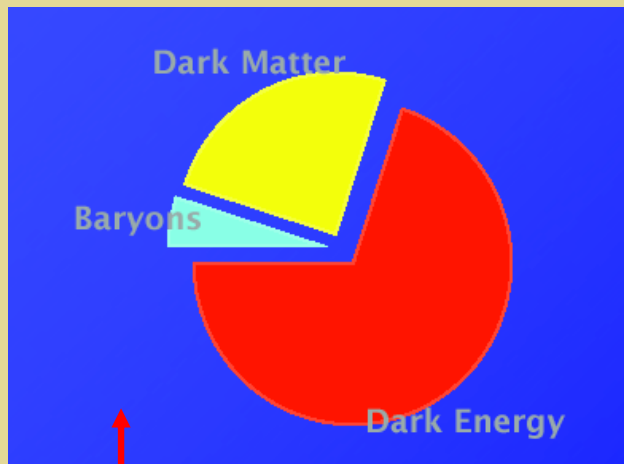


Origin of m_ν

SHOULD answer

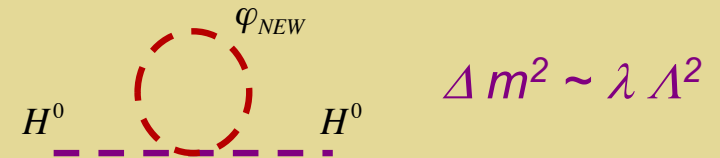
Fundamental Questions

MUST answer



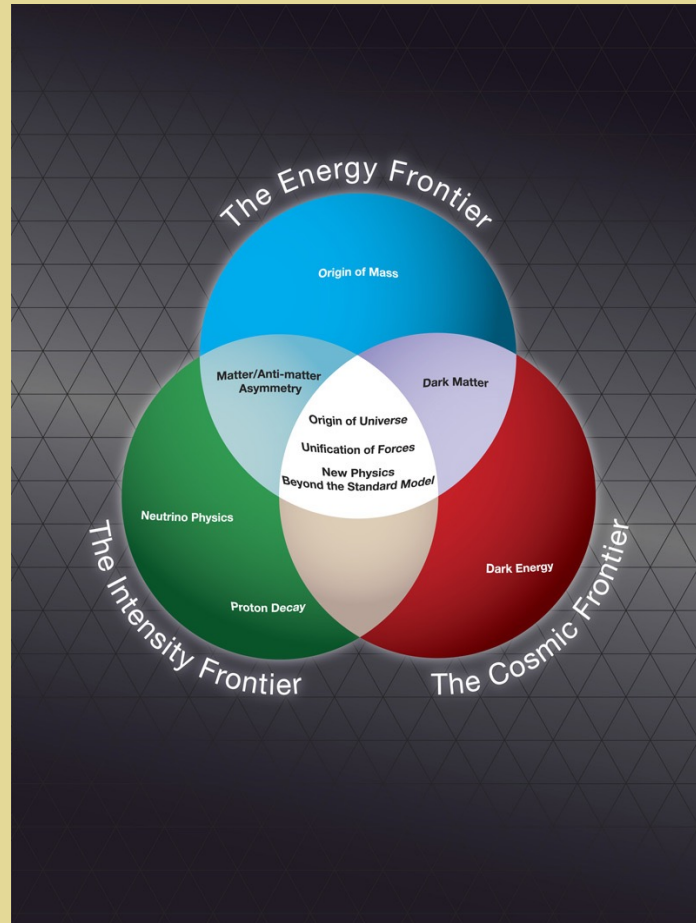
Origin of m_ν

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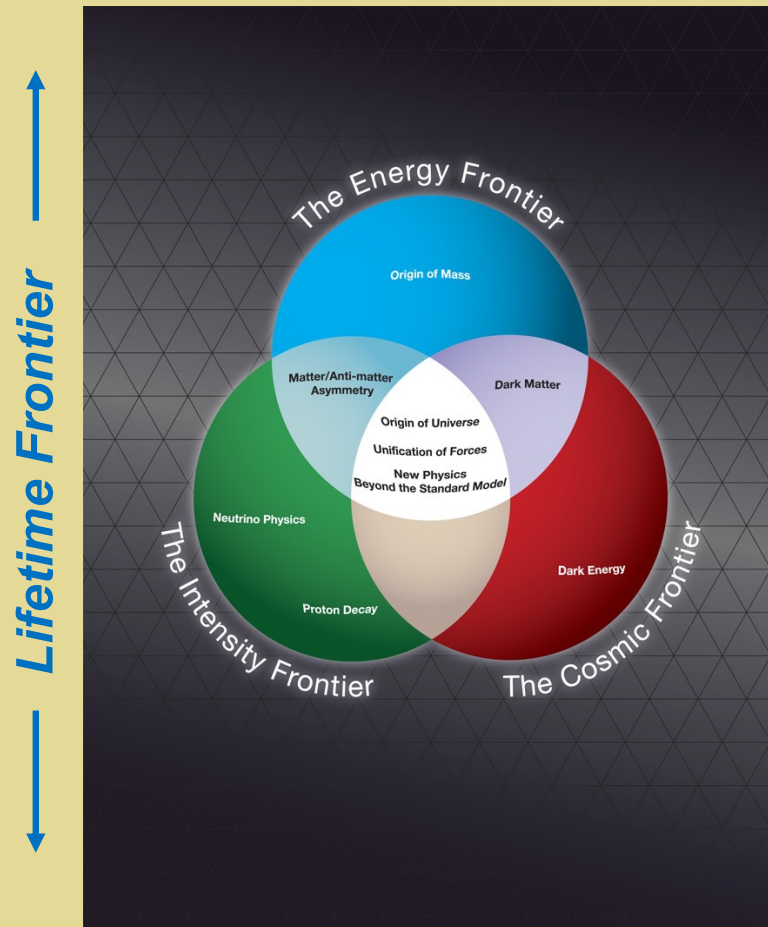
Λ *Cosmological*

Frontiers



*Historical artifact: US HEP
vision → still useful mnemonic*

Frontiers



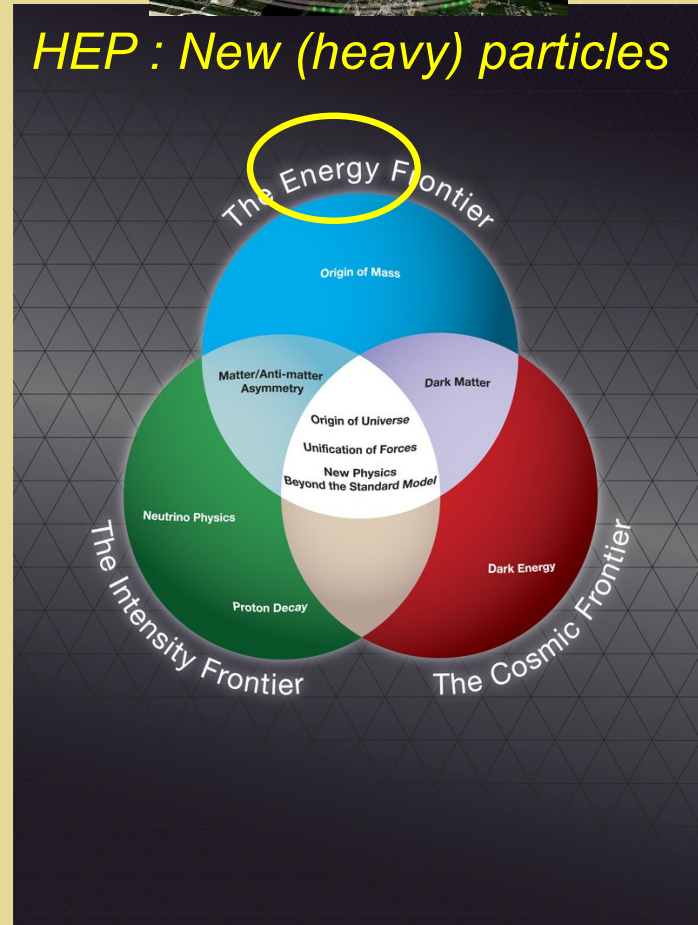
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Frontiers



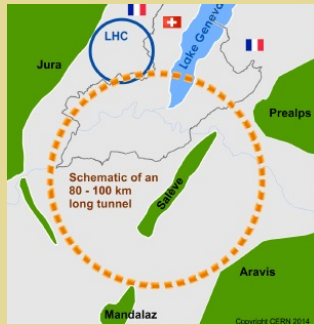
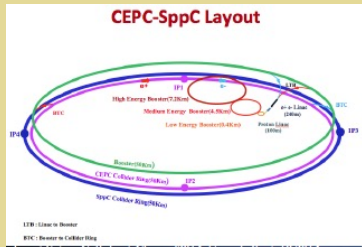
HEP : New (heavy) particles

Lifetime Frontier



Historical artifact: US HEP vision → still useful mnemonic

Frontiers



Linear Colliders
 ILC & CLIC specs

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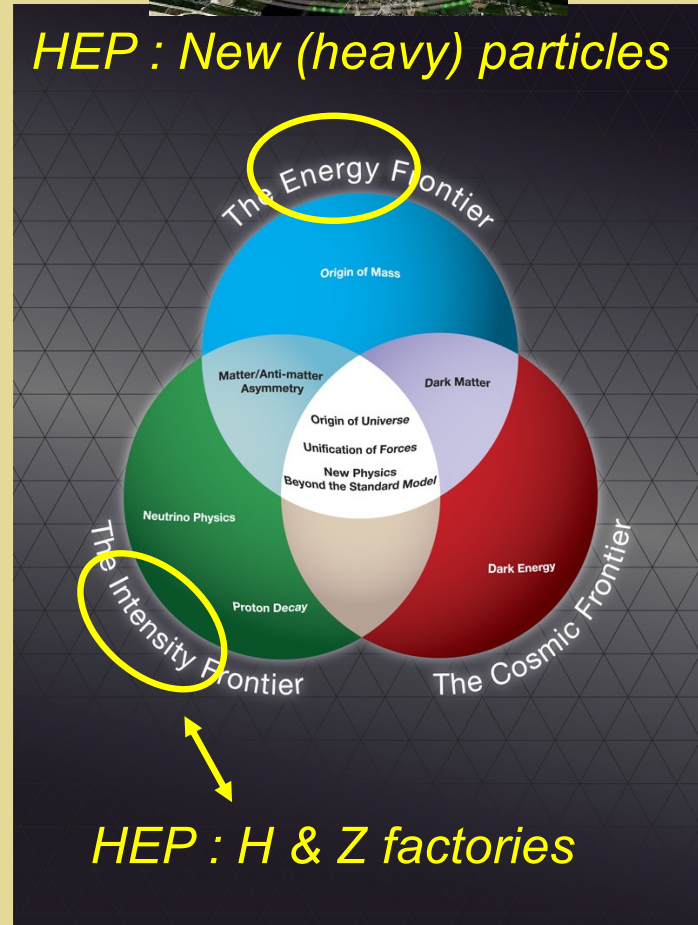
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ILC250 ~ 20km

Kitakami Mountains in Tohoku of Japan

Lifetime Frontier

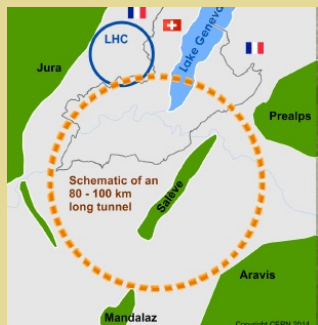
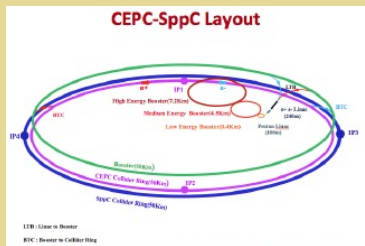
HEP : New (heavy) particles



HEP : H & Z factories

Historical artifact: US HEP vision → still useful mnemonic

Frontiers



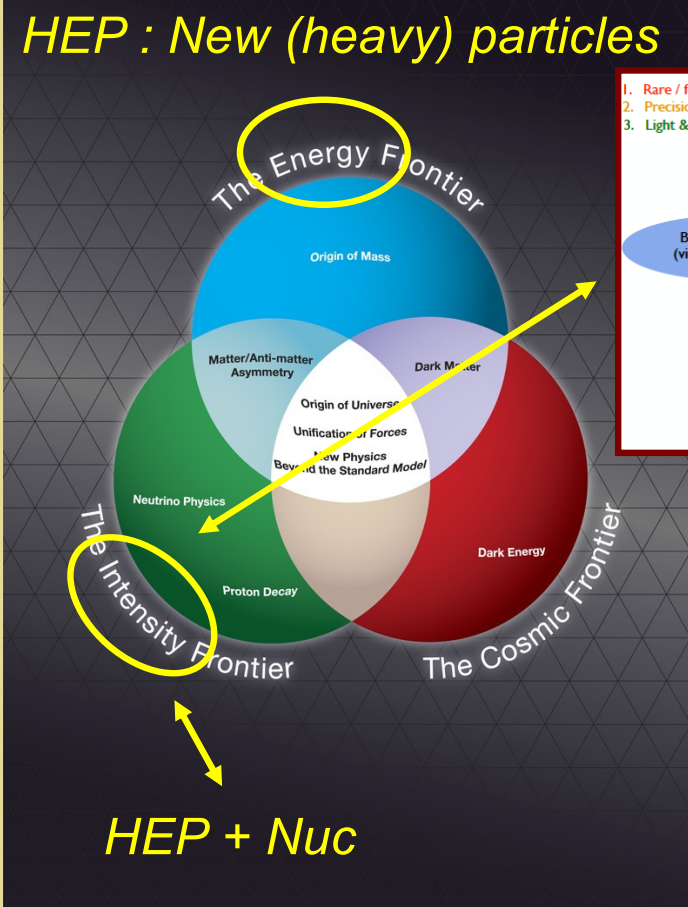
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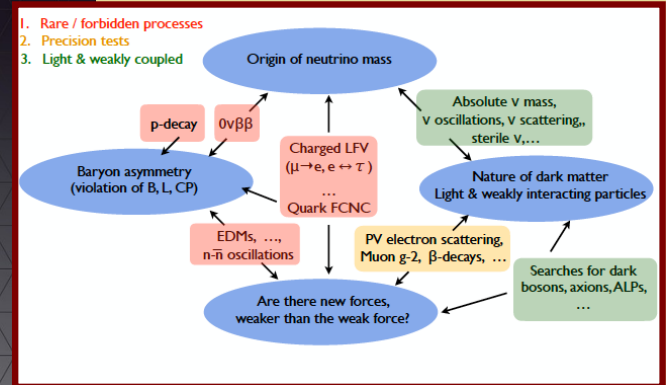
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Lifetime Frontier



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HEP + Nuc



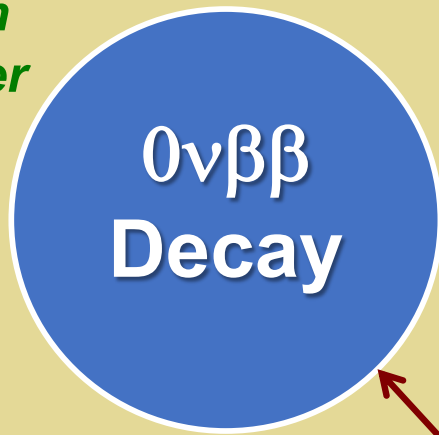
V. Cirigliano, INT EIC '24

- Precision tests: muon $g-2$, PV ee ...
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- Neutrino properties
- Flavor physics

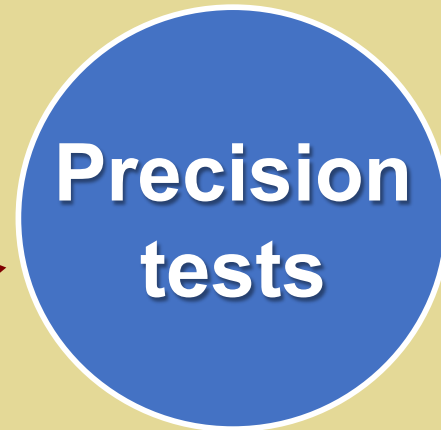
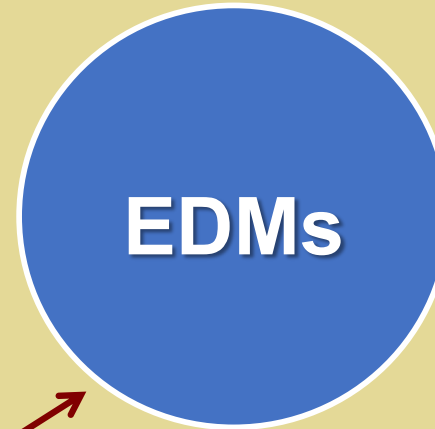
Historical artifact: US HEP vision \rightarrow still useful mnemonic

Nuclear Physics Connections

Lepton number



CP & T

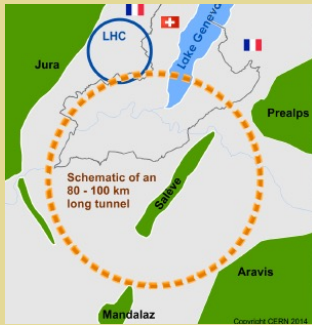
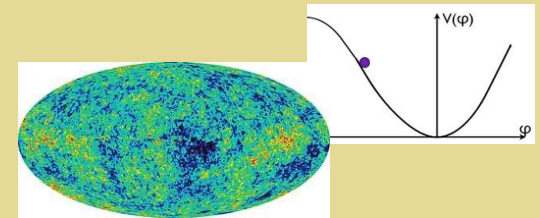
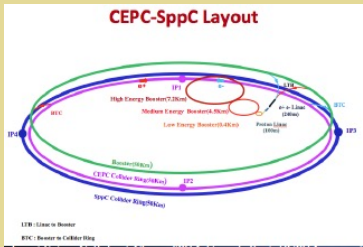


Precision tests

Muon $g-2$, PV ee , β decay...

Fundamental symmetries & neutrinos: "Intensity Frontier"

Frontiers



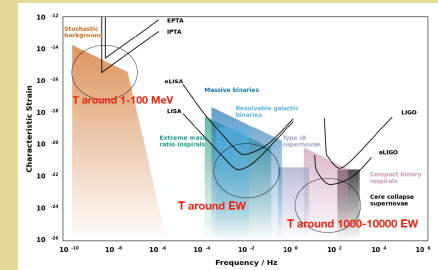
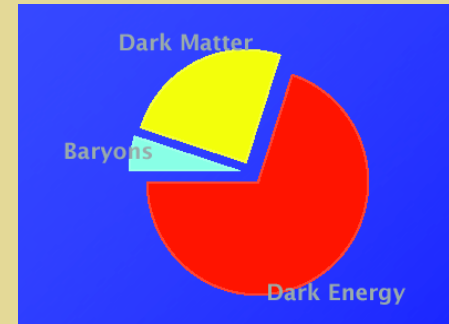
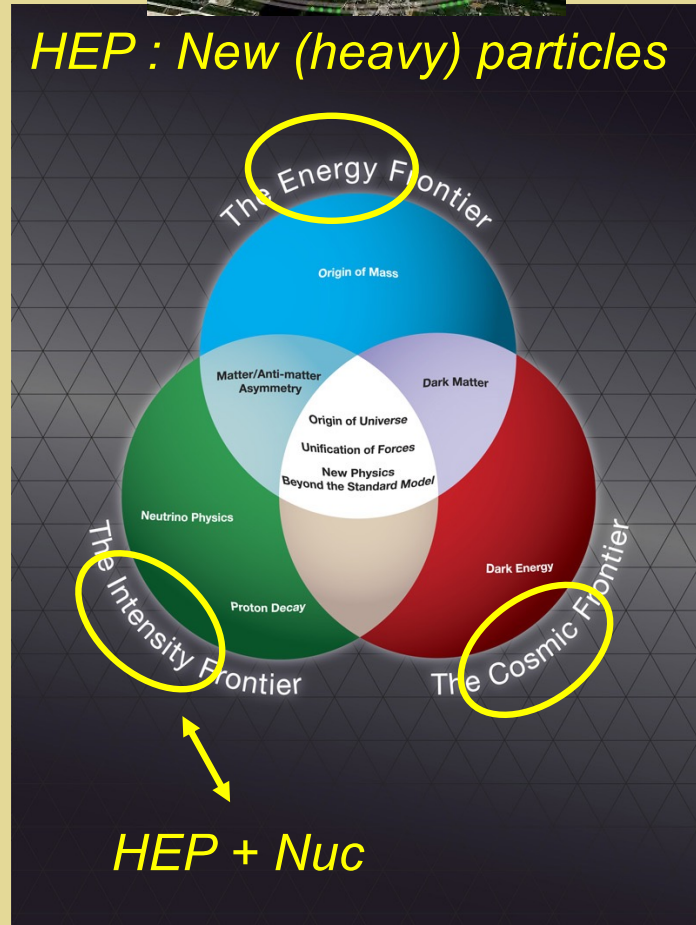
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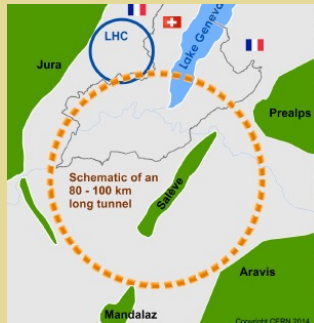
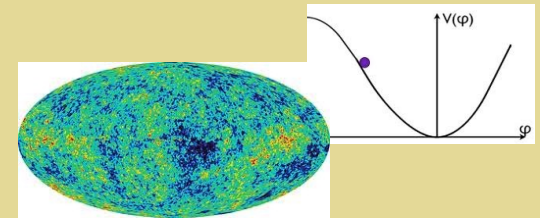
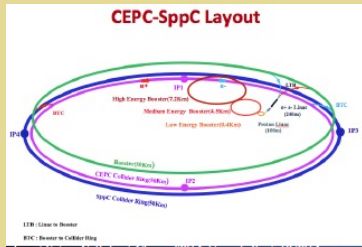


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HEP + Nuc

Historical artifact: US HEP vision \rightarrow still useful mnemonic

Frontiers



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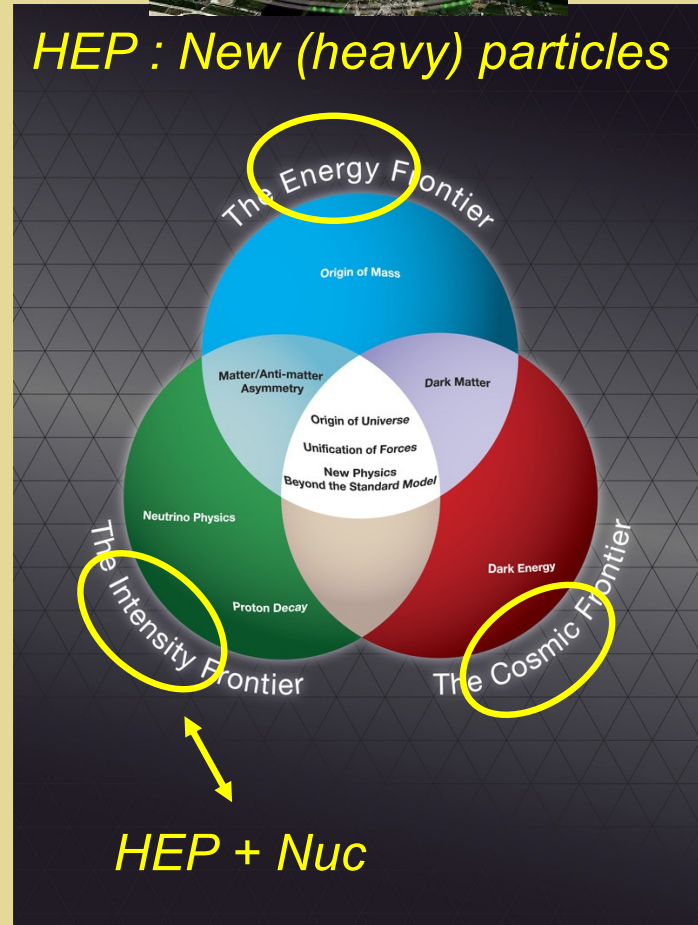
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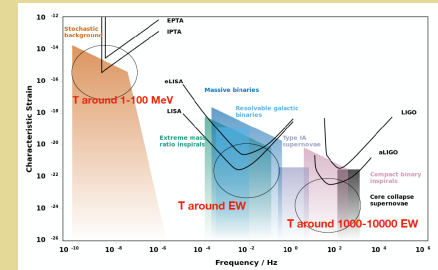
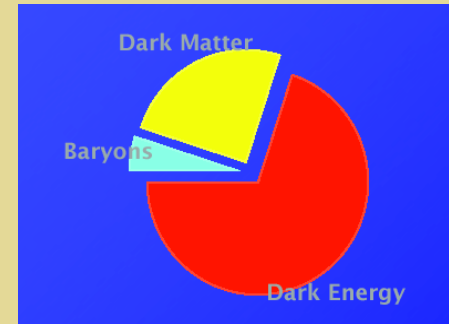
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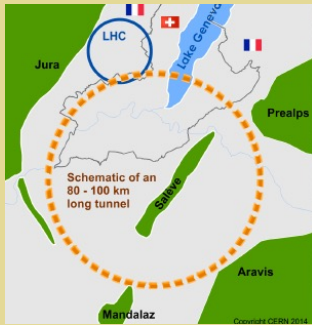
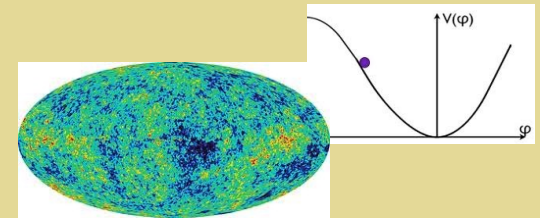
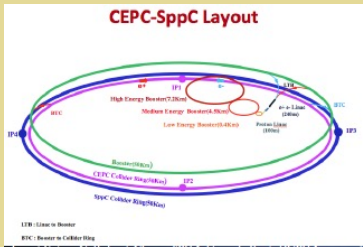


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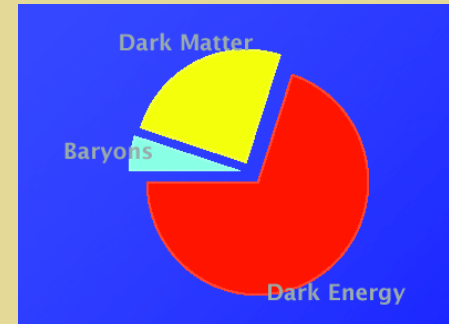
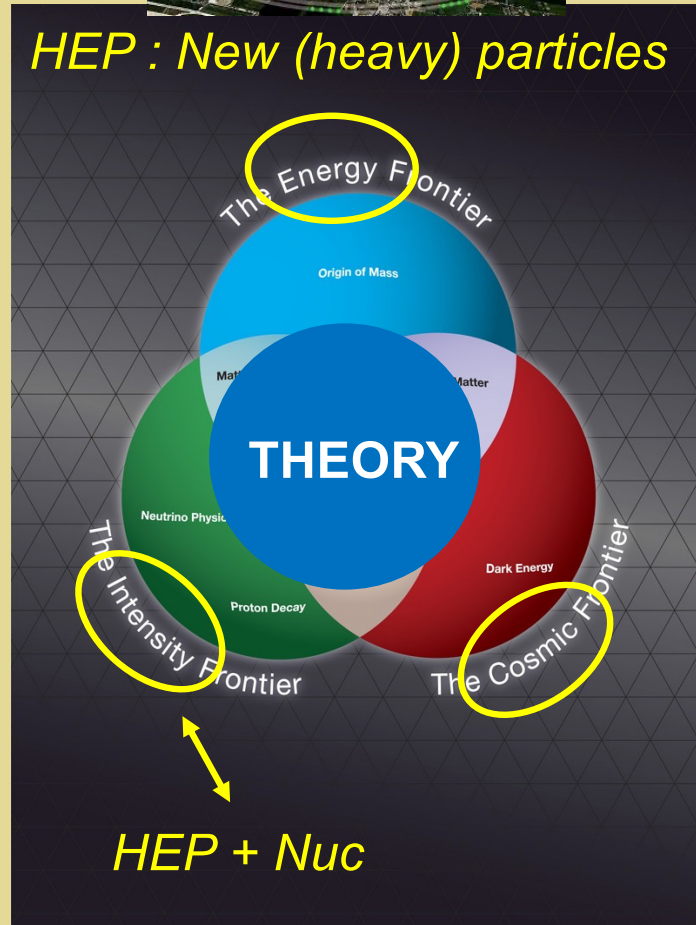
- Atomic, Molecular, Optical
- Condensed Matter

Historical artifact: US HEP vision → still useful mnemonic

Frontiers



HEP : New (heavy) particles

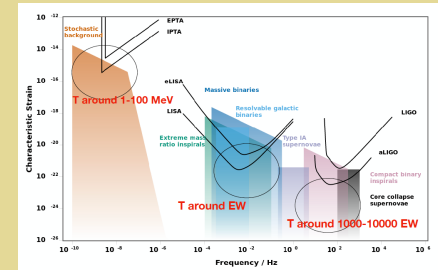


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Lifetime Frontier

HEP + Nuc

- Atomic, Molecular, Optical
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More Matter than Antimatter ?

When & how was the baryon asymmetry generated ?

Paradigmatic inter-frontier challenge

Ingredients for Baryogenesis



Scenarios: leptogenesis, EW baryogenesis, Affleck-Dine, asymmetric DM, cold baryogenesis, post-sphaleron baryogenesis...

- *B violation (sphalerons)*
- *C & CP violation*
- *Out-of-equilibrium or CPT violation*

Standard Model

BSM



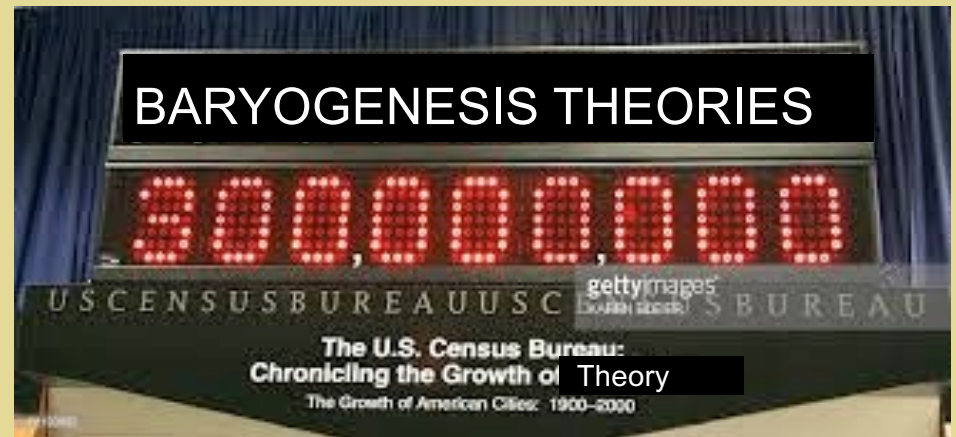
Cosmic Baryon Asymmetry

$$Y_B = \frac{n_B}{s} = (8.66 \pm 0.04) \times 10^{-11}$$

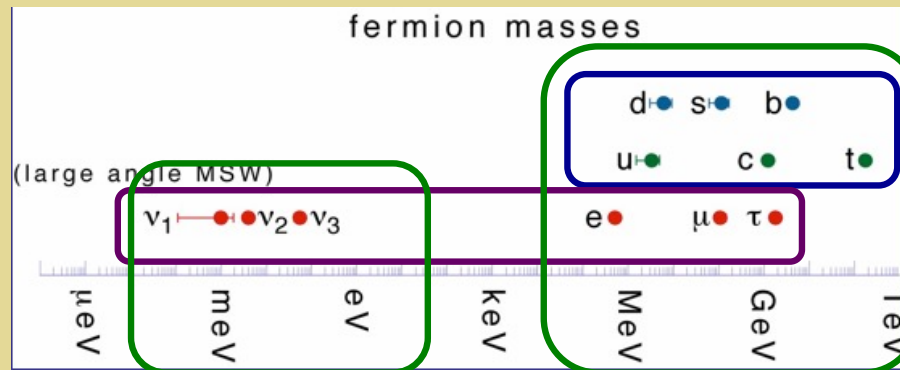
One number → ~~!!!~~ ~~!!!~~ ~~!!!~~ ... *Explanations*

Experiment can help:

- *Discover ingredients*
- *Falsify candidates*



Fermion Masses & Baryon Asymmetry



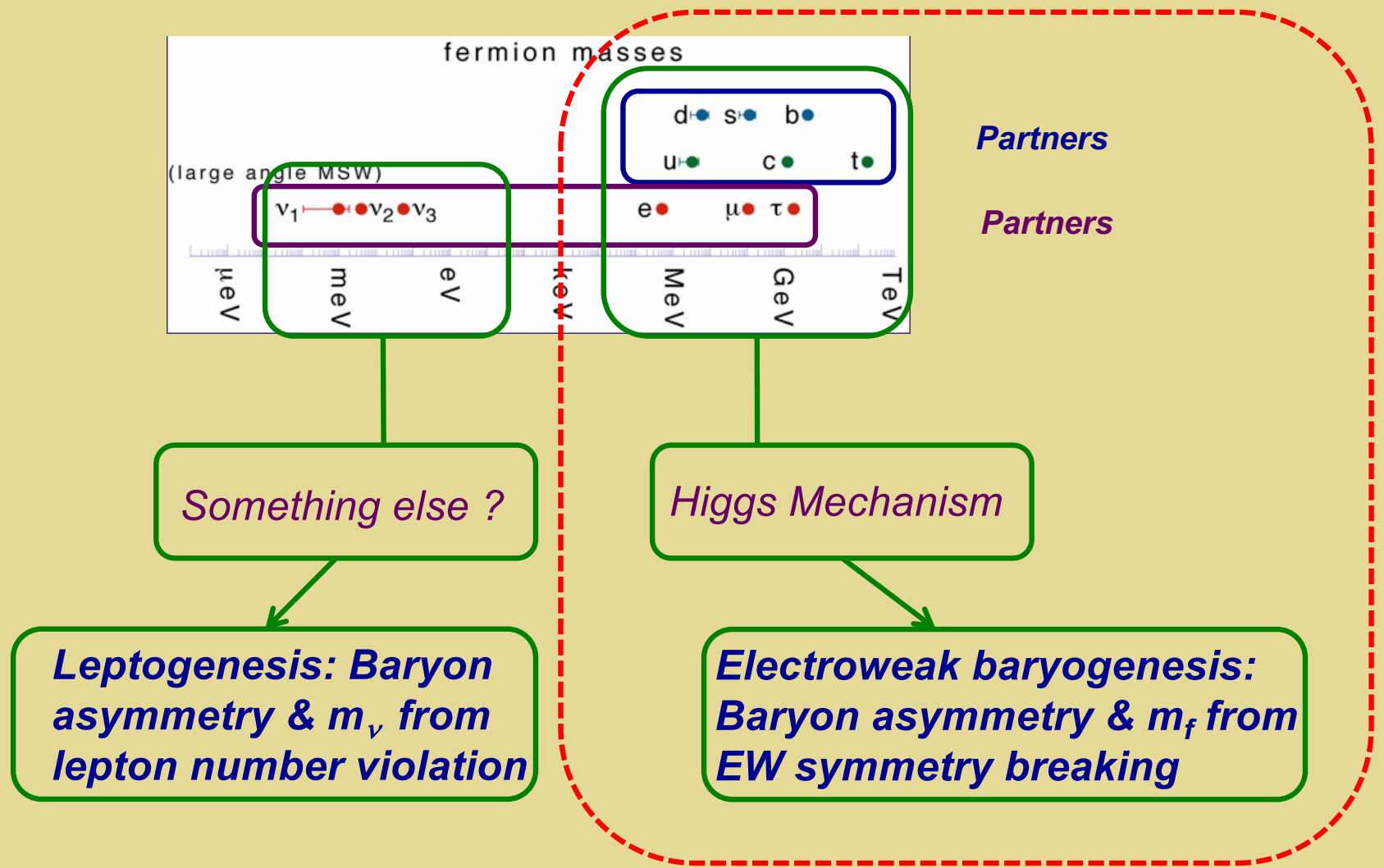
Something else ?

Higgs Mechanism

Leptogenesis: Baryon asymmetry & m_ν from lepton number violation

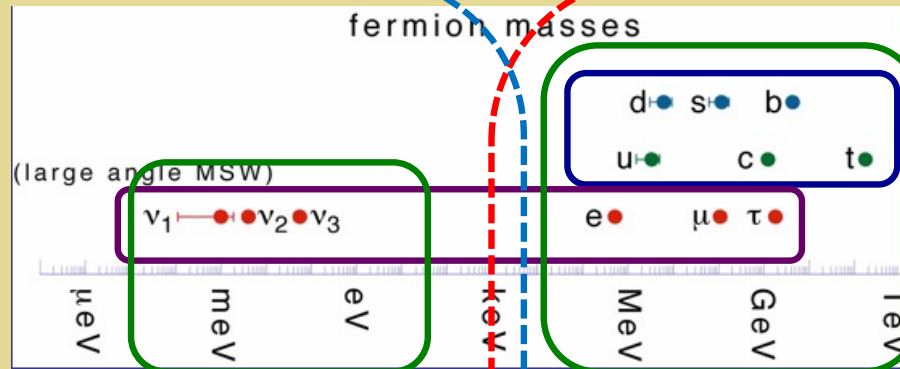
Electroweak baryogenesis: Baryon asymmetry & m_f from EW symmetry breaking

Fermion Masses & Baryon Asymmetry



This talk

Fermion Masses & Baryon Asymmetry



Partners

Partners

Something else ?

Higgs Mechanism

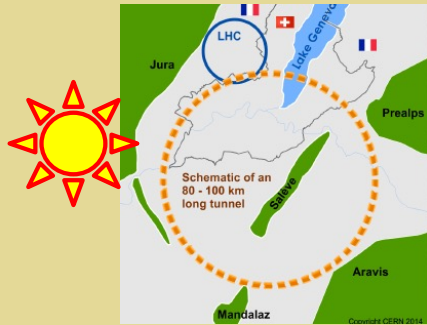
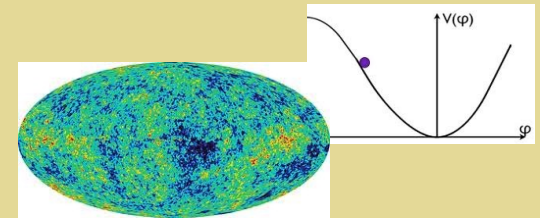
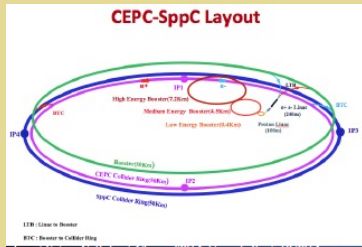
Leptogenesis: Baryon asymmetry & m_ν from lepton number violation

Electroweak baryogenesis: Baryon asymmetry & m_f from EW symmetry breaking

Another day

This talk

Frontiers



Linear Colliders
ILC & CLIC specs

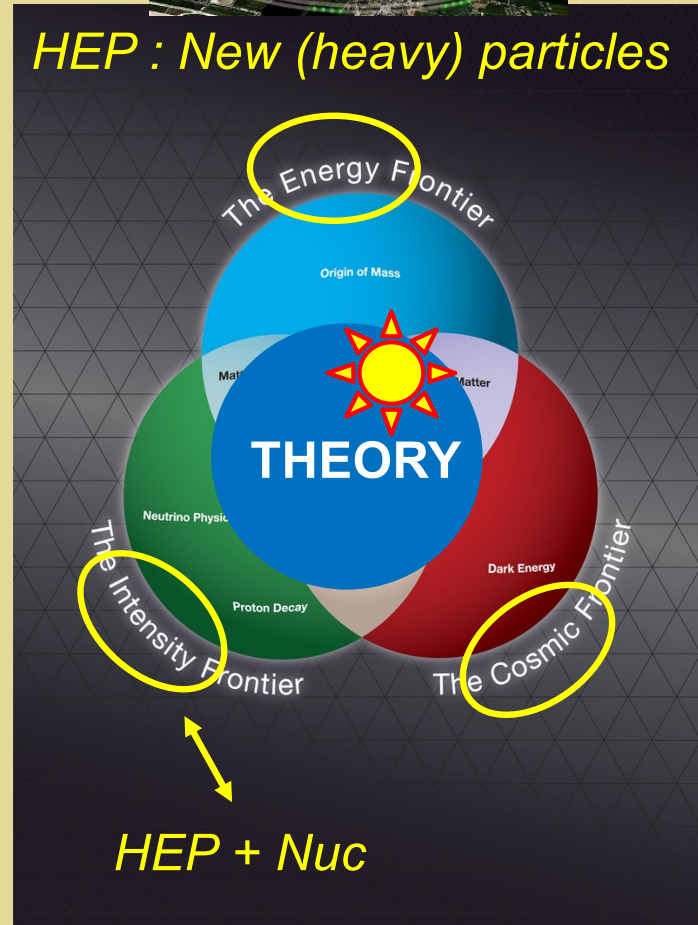
- Energy extendability to TeV scale lies in the heart of linear colliders: ILC focuses on \sqrt{s} from 250 GeV to 1 TeV; CLIC 380 GeV to 3 TeV, keeping options to run at Z-pole ("GigaZ")
- Complementary approaches: "Warm" & "Cold" accelerating technologies; 72MeV/m @ CLIC380; 51.5MeV/m @ ILC250
- Polarized beams: both offering 80% for electron; 30% for positron in ILC default design

東京大学 THE UNIVERSITY OF TOKYO

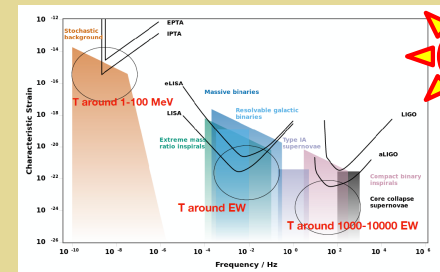
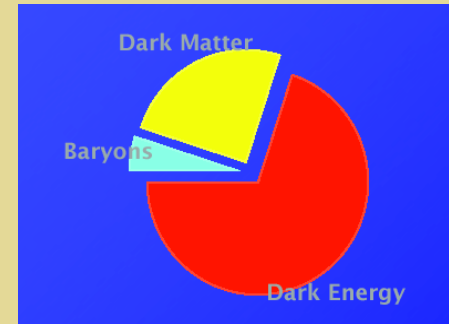
ILC250 - 20km

Lifetime Frontier

HEP : New (heavy) particles



HEP + Nuc



- Precision tests: muon $g-2$, PV $ee...$
- Fundamental symmetry tests (CP, Lepton number...)
- Neutrino properties
- Flavor physics

- Atomic, Molecular, Optical
- Condensed Matter

Historical artifact: US HEP vision \rightarrow still useful mnemonic




II. Origin of Matter: EW Scale

Was the baryon asymmetry generated in conjunction with spontaneous EW symmetry breaking ?

EWBG Ingredients

- ***EW Sphalerons***
- ***Strong 1st Order EW Phase Transition***
- ***Left-handed number density***

EWBG Ingredients

- ***EW Sphalerons*** 
- ***Strong 1st Order EW Phase Transition***  ***BSM Higgs***
- ***Left-handed number density***  ***BSM CPV***

EWBG Ingredients

- ***EW Sphalerons***



- ***Strong 1st Order EW Phase Transition***



BSM Higgs

- ***Left-handed number density***

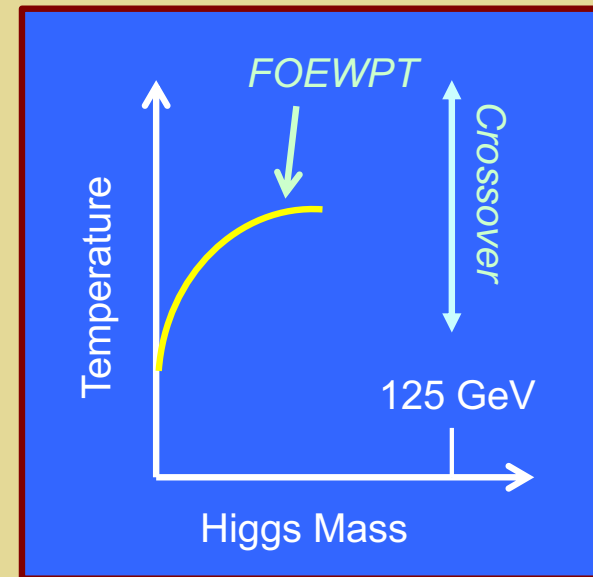
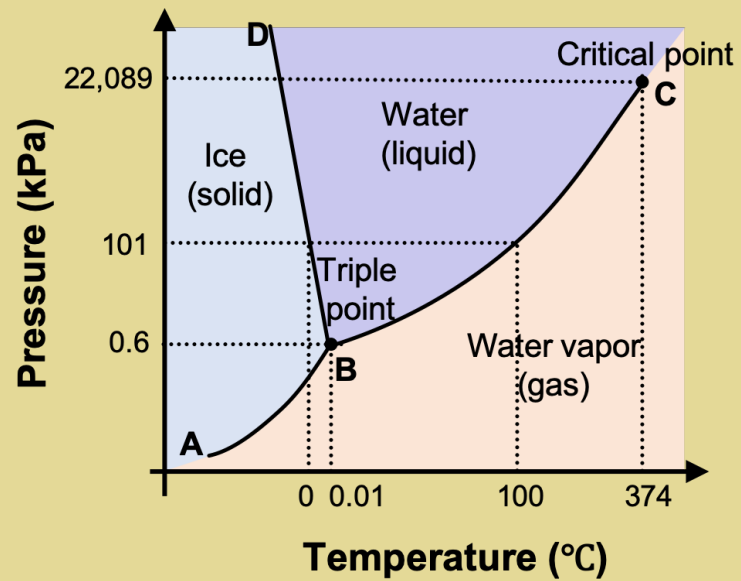


BSM CPV

Was There an Electroweak Phase Transition ?

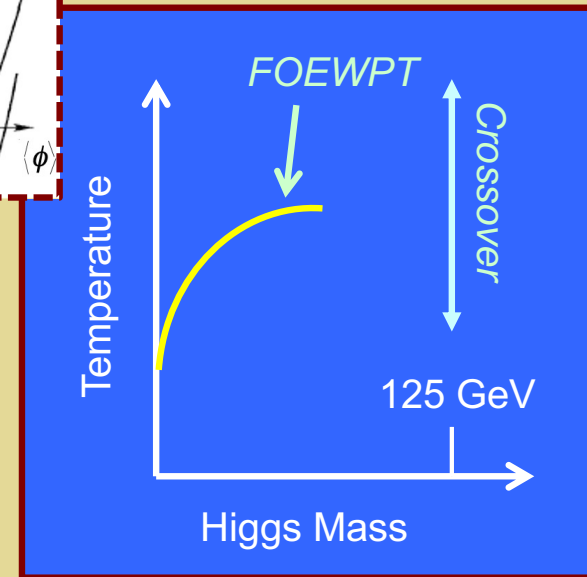
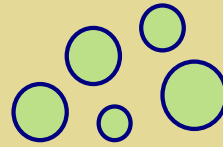
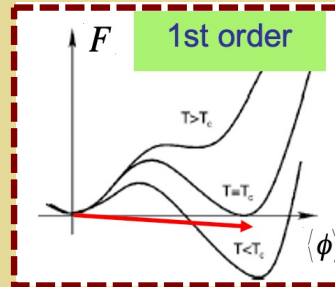
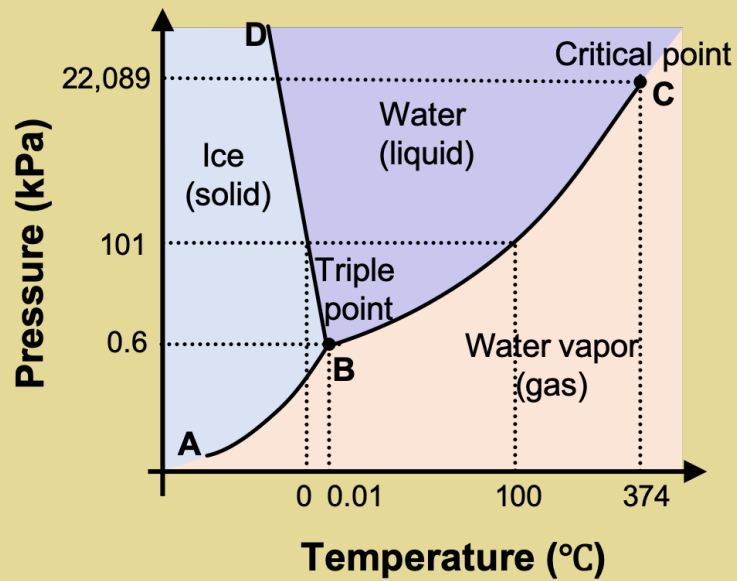
- ***Interesting in its own right***
- ***Key ingredient for EW baryogenesis***
- ***Source of gravitational radiation***

Was There an Electroweak Phase Transition ?



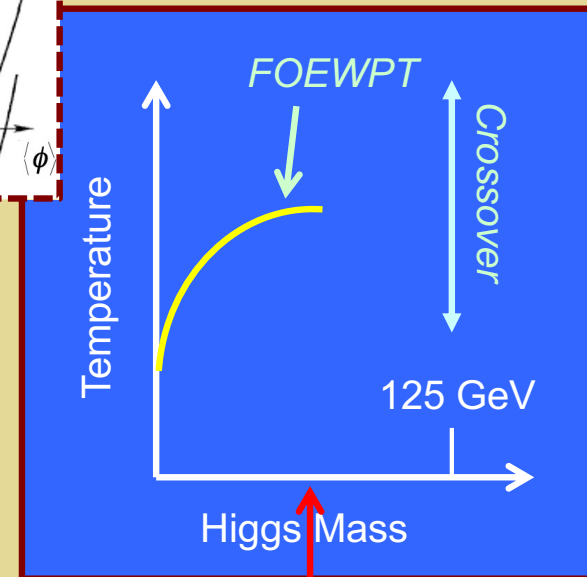
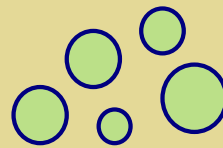
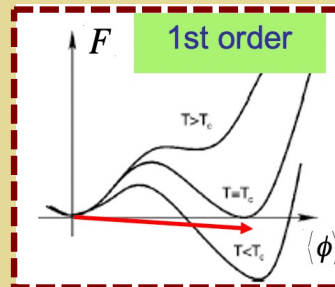
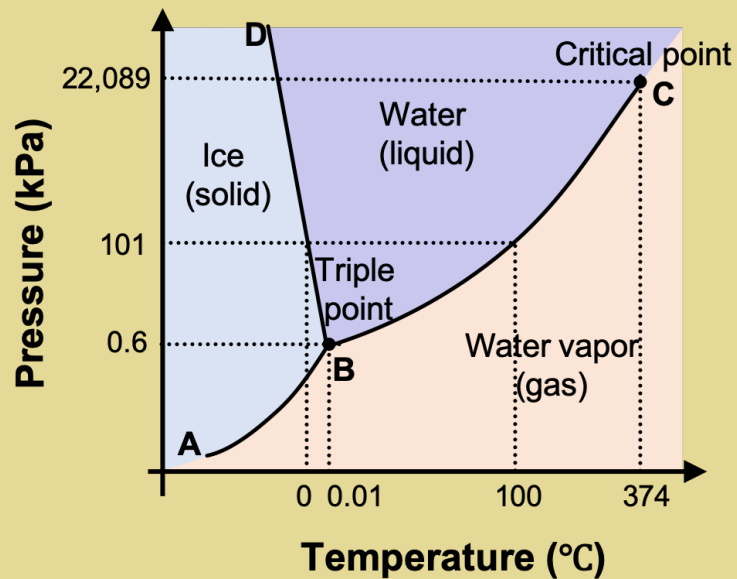
EW Phase Diagram

Was There an Electroweak Phase Transition ?



EW Phase Diagram

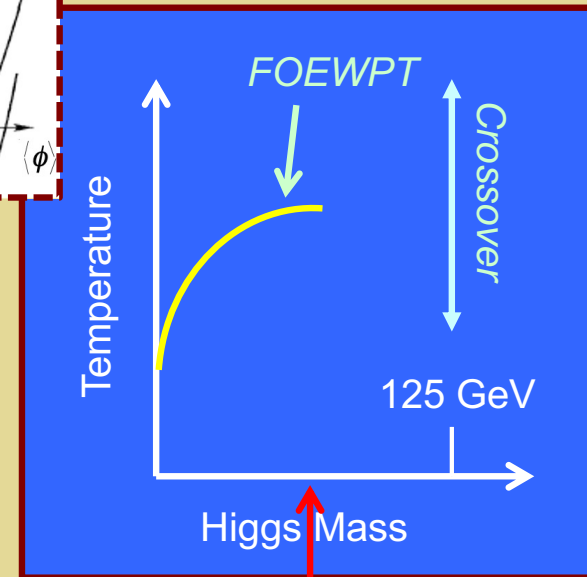
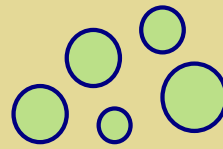
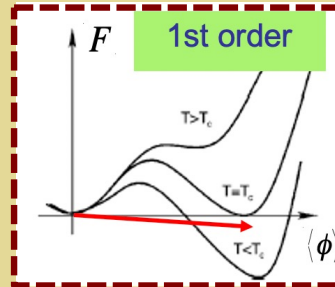
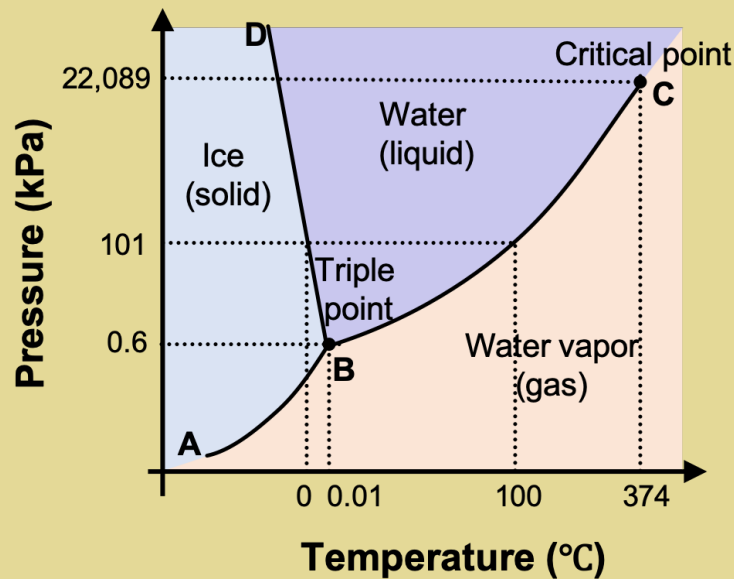
Was There an Electroweak Phase Transition ?



EW Phase Diagram

| Lattice | Authors | M_h^C (GeV) |
|----------------|---------|----------------|
| 4D Isotropic | [76] | 80 ± 7 |
| 4D Anisotropic | [74] | 72.4 ± 1.7 |
| 3D Isotropic | [72] | 72.3 ± 0.7 |
| 3D Isotropic | [70] | 72.4 ± 0.9 |

Was There an Electroweak Phase Transition ?

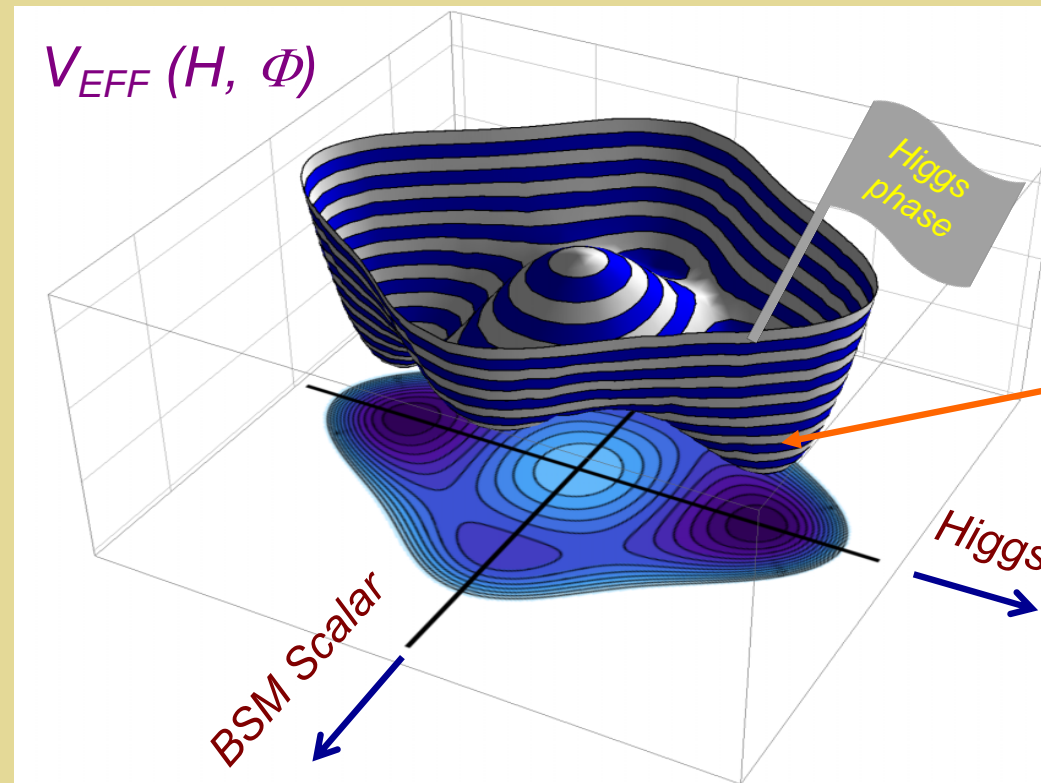


EW Phase Diagram

How does this picture change in presence of new TeV scale physics ? What is the phase diagram ? SFOEWPT ?

| Lattice | Authors | M_h^C (GeV) |
|----------------|---------|----------------|
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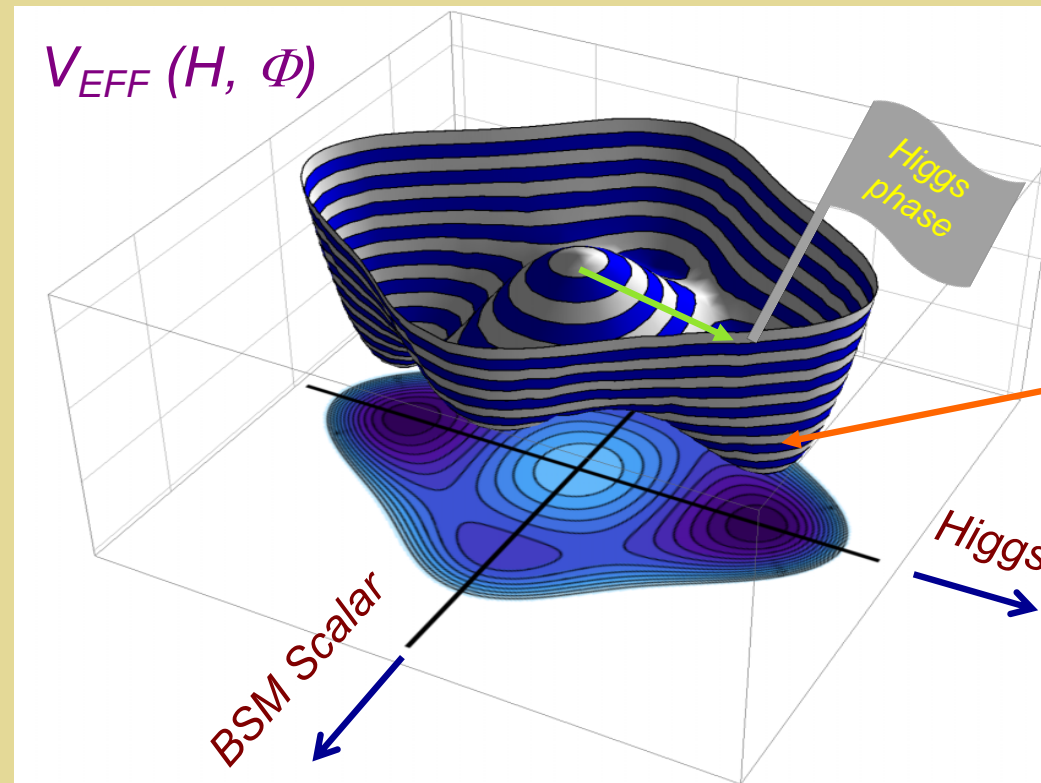
What Was the EWSB Thermal History ?



How did we end up here ?

Extrema can evolve differently as T evolves \rightarrow rich possibilities for symmetry breaking

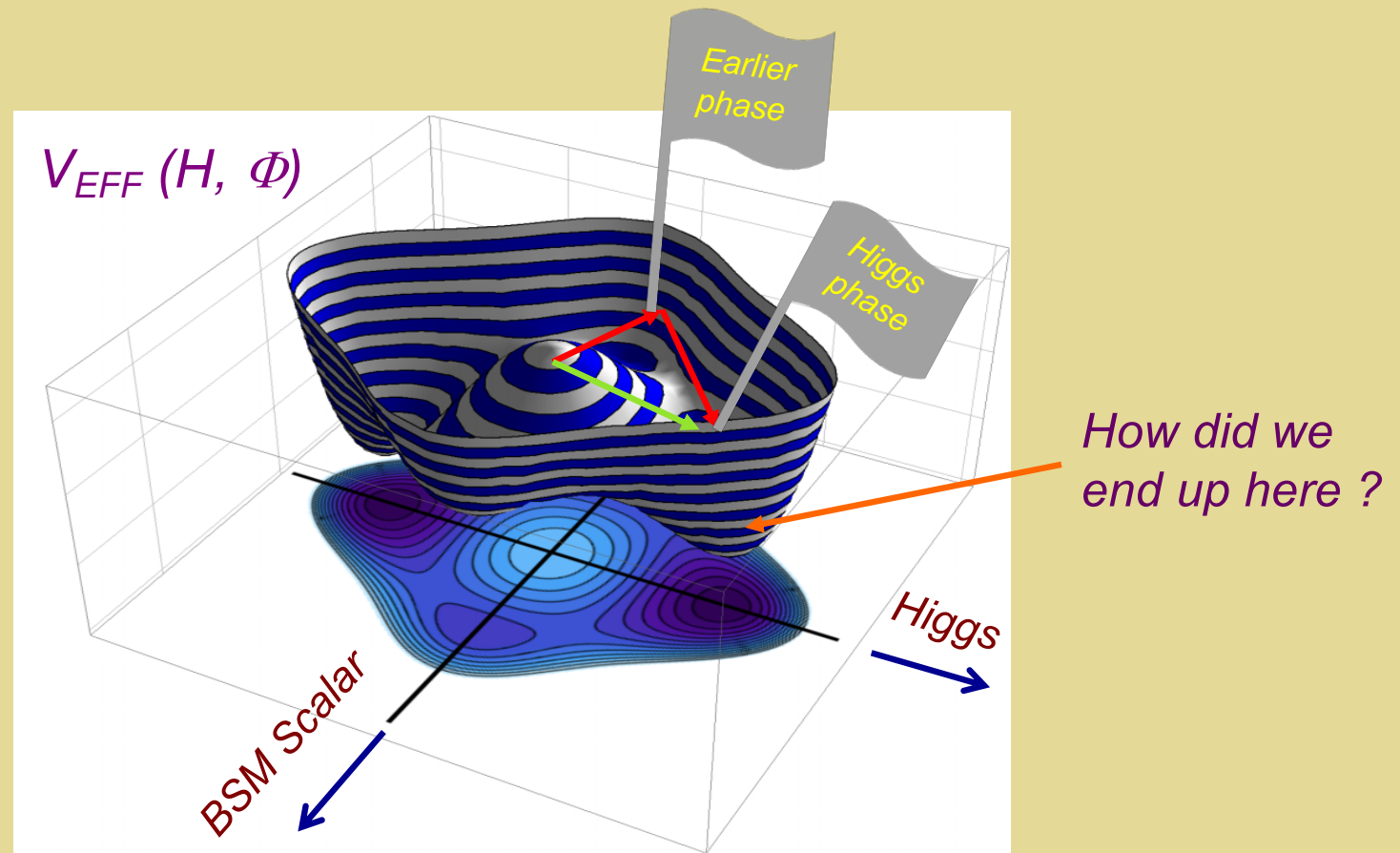
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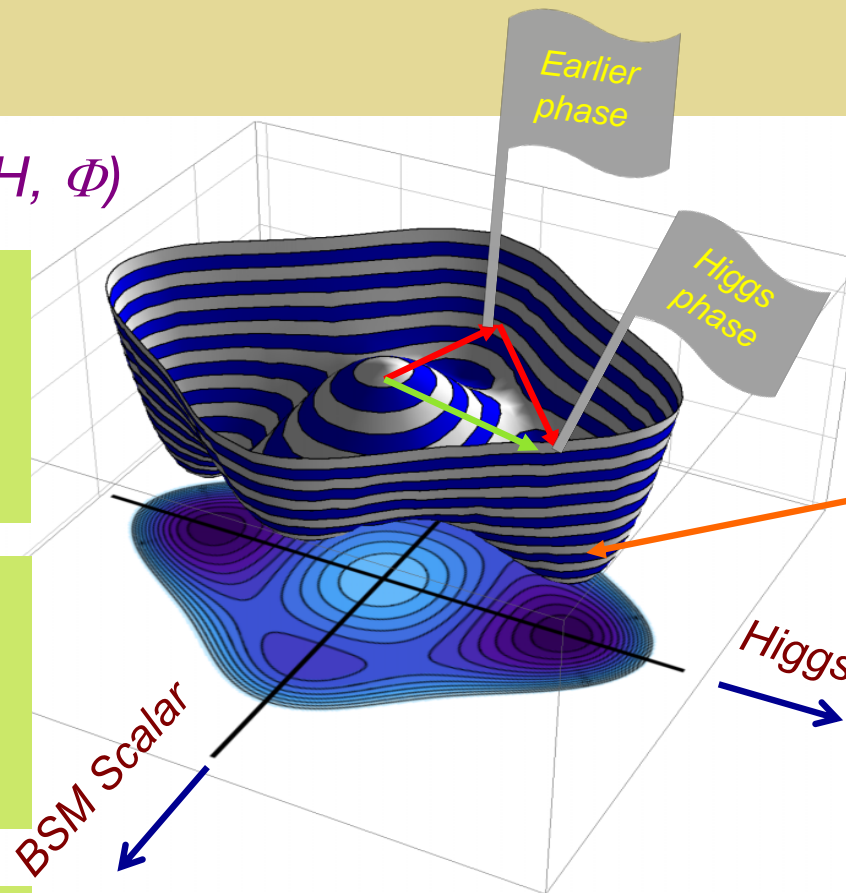
What Was the EWSB Thermal History ?



**Extrema can evolve differently as T evolves \rightarrow
rich possibilities for symmetry breaking**

What Was the EWSB Thermal History ?

$$V_{\text{EFF}}(H, \Phi)$$



- What is the landscape of potentials and their thermal histories?

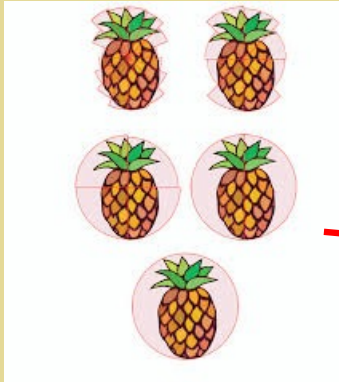
- How can we probe this $T > 0$ landscape experimentally?

- How reliably can we compute the thermodynamics?

**n evolve differently as T evolves \rightarrow
abilities for symmetry breaking**

Was There an EW Phase Transition?

Bubble Collisions

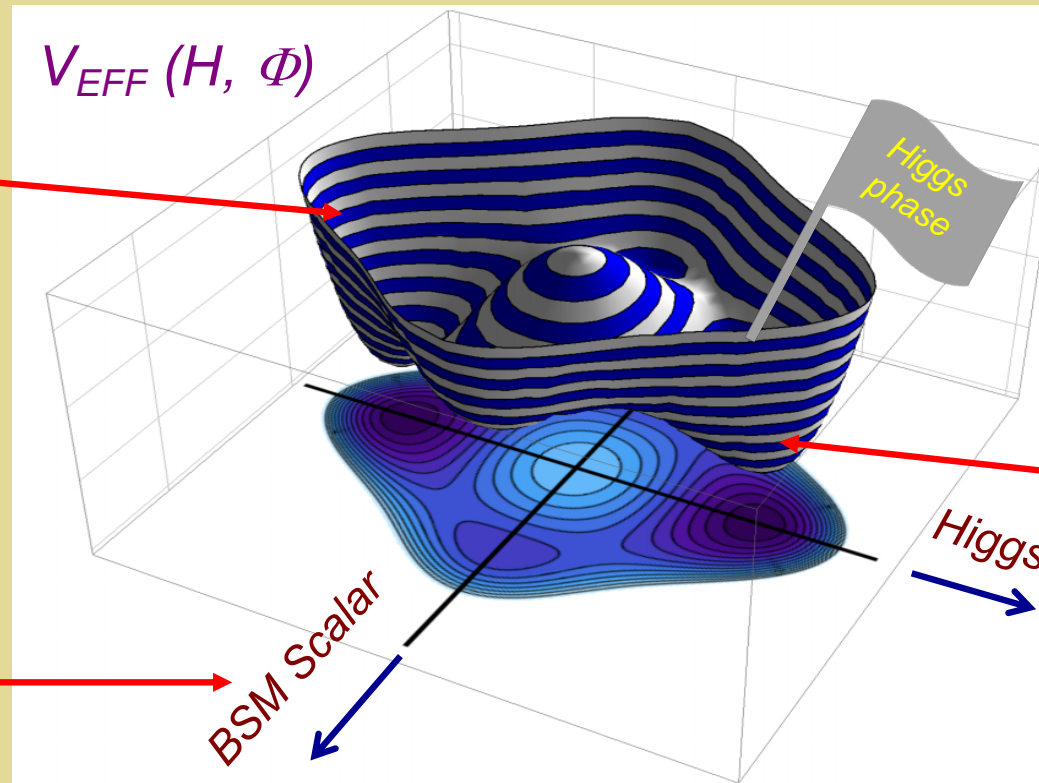


Grav Radiation

Direct Production



BSM Higgs



Higgs precision tests

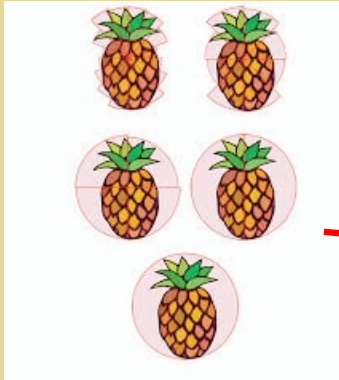


SM Higgs BSM Higgs

Extrema can evolve differently as T evolves \rightarrow rich possibilities for symmetry breaking

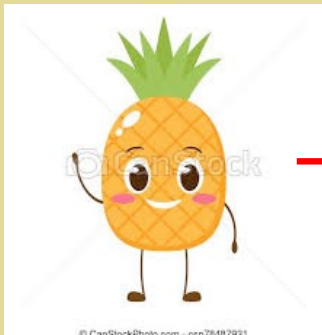
Was There an EW Phase Transition?

Bubble Collisions

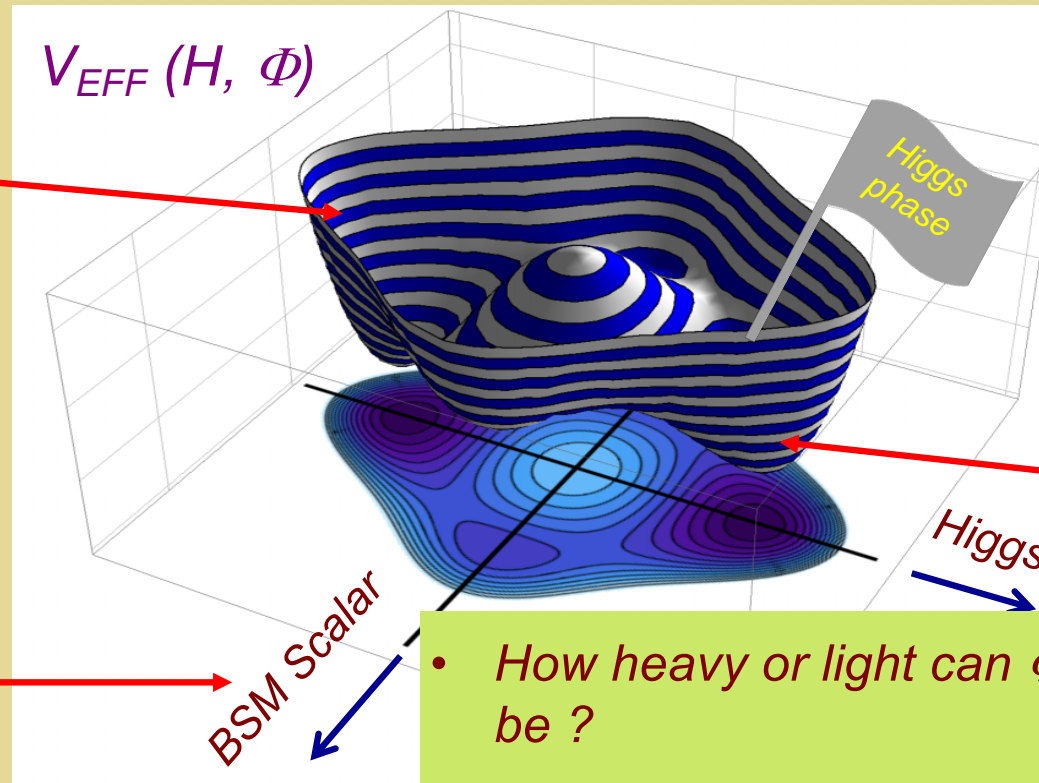


Grav Radiation

Direct Production



BSM Higgs



Higgs precision tests



SM Higgs BSM Higgs

- How heavy or light can Φ be ?
- How coupled to H ?
- Can it be discovered with colliders & GW probes ?

Extrema can evolve
rich possibilities for



$T_{EW} \rightarrow$ Scale for Colliders & GW probes

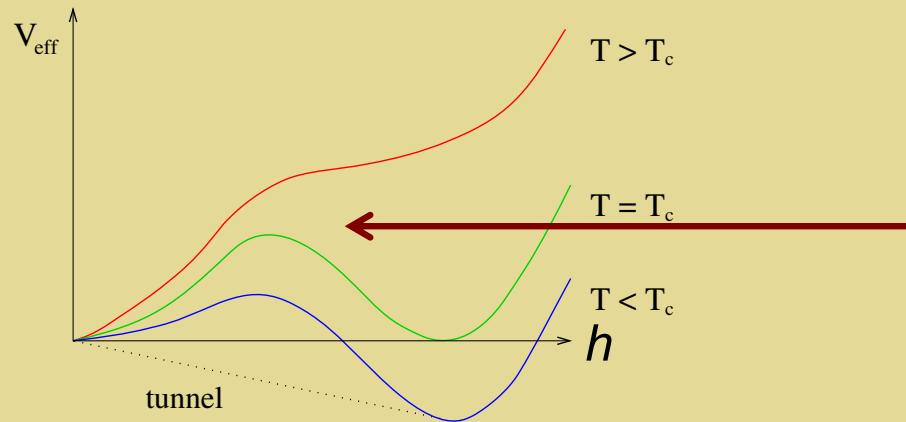
High- T SM Effective Potential

$$V(h, T)_{\text{SM}} = D(T^2 - T_0^2) h^2 + \lambda h^4 + \dots$$

$$T_0 \sim 140 \text{ GeV}$$

$$\equiv T_{EW}$$

First Order EWPT from BSM Physics



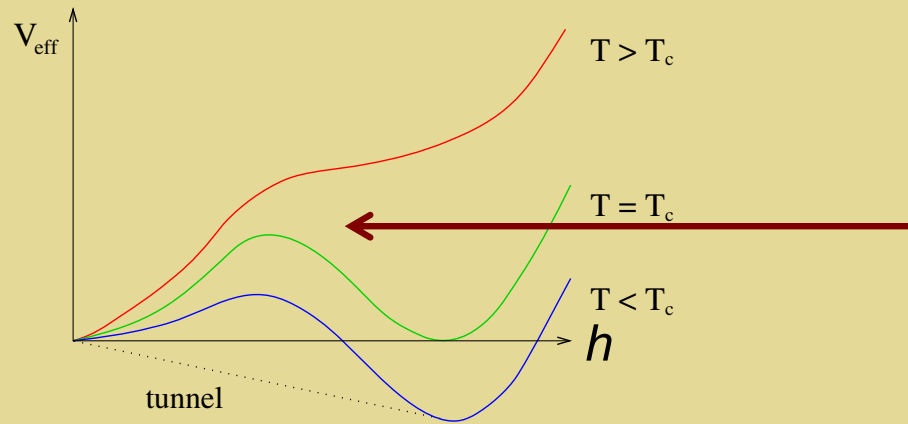
Generate finite- T barrier



Introduce new scalar ϕ interaction with h via the Higgs Portal

- $M_\phi \lesssim 700 \text{ GeV}$
- h - ϕ mixing: $|\sin\theta| \gtrsim 0.01$

First Order EWPT from BSM Physics



Generate finite-T barrier

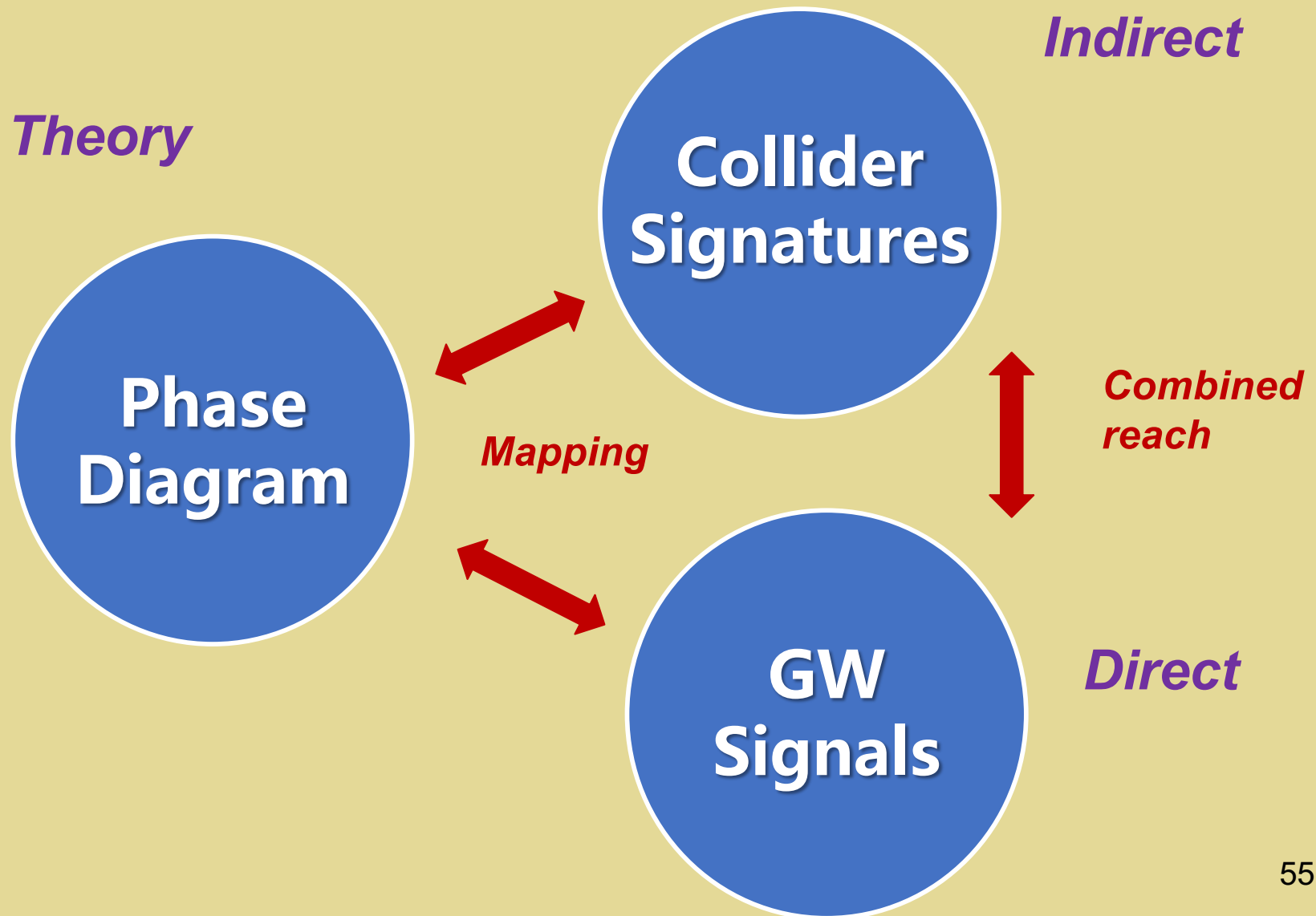


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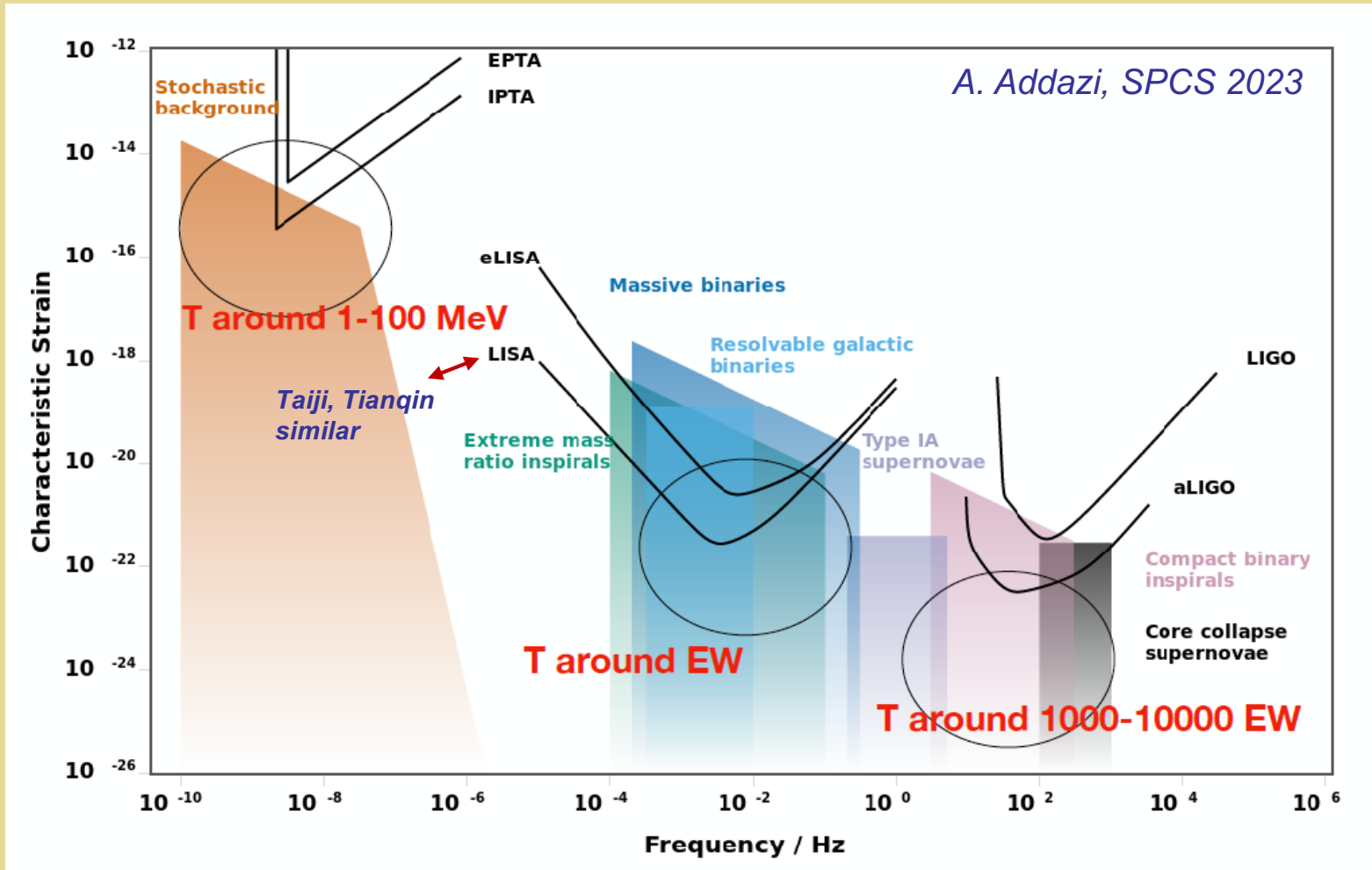
- $M_\phi \lesssim 700 \text{ GeV}$
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Collider target

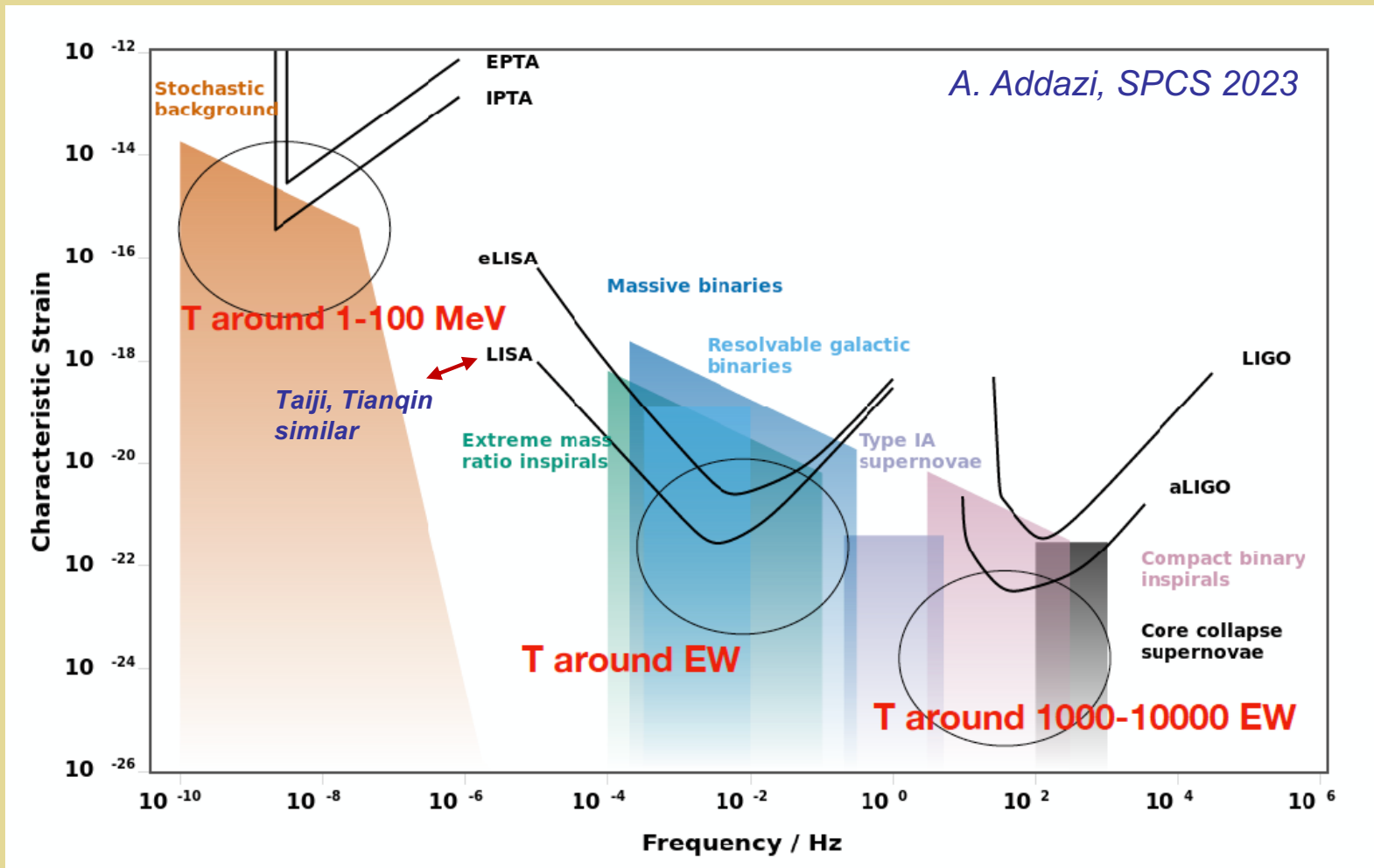
BSM EWPT: Inter-frontier Connections



Gravitational Waves

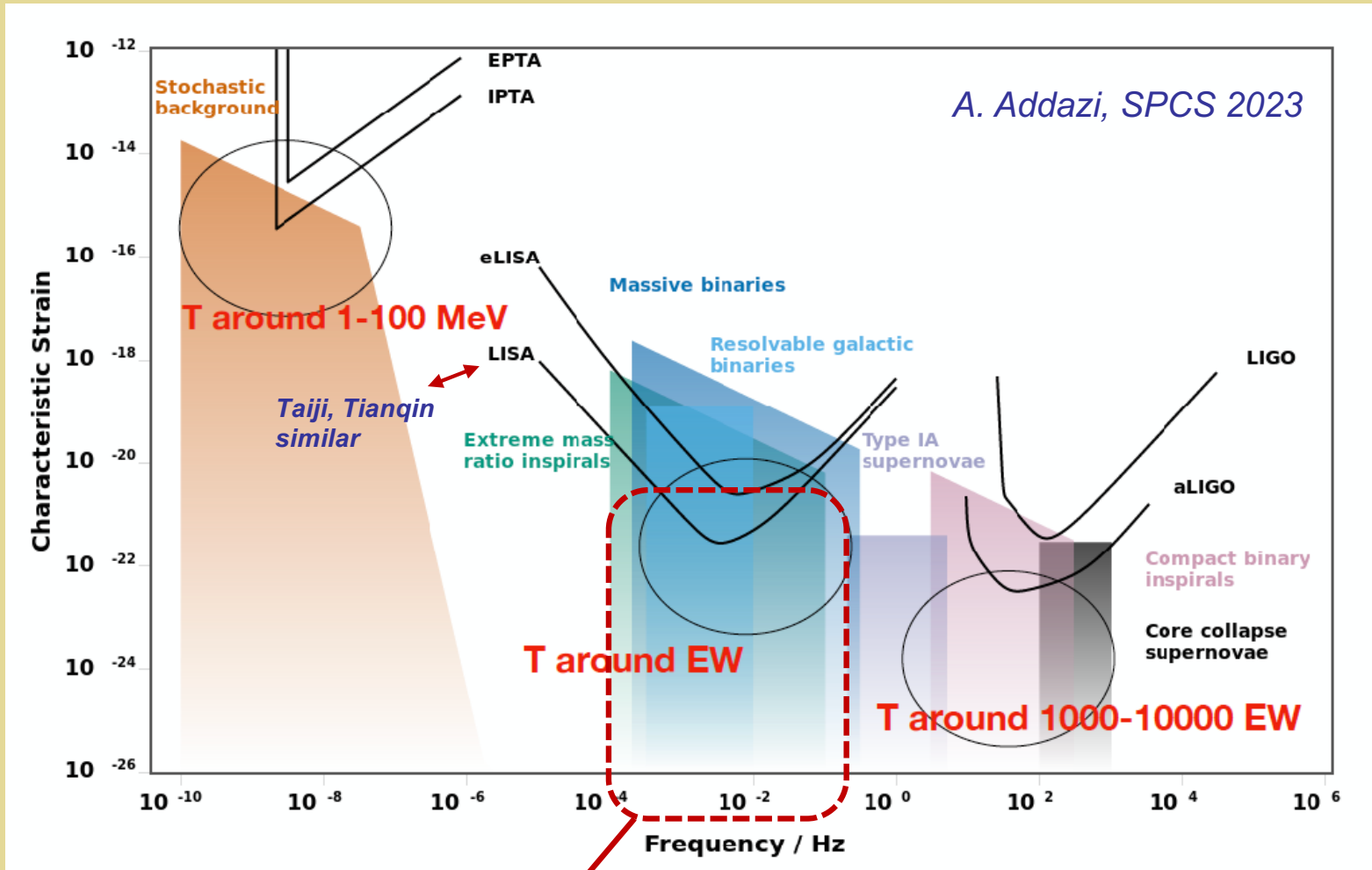


Gravitational Waves



Particle physics (phase transitions, domain walls...)
→ possible GW sources at range of frequencies

Gravitational Waves



EWPT laboratory for GW micro-physics: colliders can probe particle physics responsible for non-astro GW sources → test our framework for GW microphysics at other scales

EWPT: Theory-Pheno Interface

***Theoretical developments →
phenomenological implications***

Models & Phenomenology

What BSM Scenarios?

SM + Scalar Singlet

Espinosa, Quiros 93, Benson 93, Choi, Volkas 93, Vergara 96, Branco, Delepine, Emmanuel-Costa, Gonzalez 98, Ham, Jeong, Oh 04, Ahriche 07, Espinosa, Quiros 07, Profumo, Ramsey-Musolf, Shaughnessy 07, Noble, Perelstein 07, Espinosa, Konstandin, No, Quiros 08, Barger, Langacker, McCaskey, Ramsey-Musolf, Shaughnessy 09, Ashoorioon, Konstandin 09, Das, Fox, Kumar, Weiner 09, Espinosa, Konstandin, Riva 11, Chung, Long 11, Barger, Chung, Long, Wang 12, Huang, Shu, Zhang 12, Fairbairn, Hogan 13, Katz, Perelstein 14, Profumo, Ramsey-Musolf, Wainwright, Winslow 14, Jiang, Bian, Huang, Shu 15, Kozaczuk 15, Cline, Kainulainen, Tucker-Smith 17, Kurup, Perelstein 17, Chen, Kozaczuk, Lewis 17, Gould, Kozaczuk, Niemi, Ramsey-Musolf, Tenkanen, Weir 19...

SM + Scalar Doublet
(2HDM)

Turok, Zadrozny 92, Davies, Froggatt, Jenkins, Moorhouse 94, Cline, Lemieux 97, Huber 06, Froome, Huber, Seniuch 06, Cline, Kainulainen, Trott 11, Dorsch, Huber, No 13, Dorsch, Huber, Mimasu, No 14, Basler, Krause, Muhlleitner, Wittbrodt, Wlotzka 16, Dorsch, Huber, Mimasu, No 17, Bernon, Bian, Jiang 17, Andersen, Gorda, Helset, Niemi, Tenkanen, Tranberg, Vuorinen, Weir 18...

SM + Scalar Triplet

Patel, Ramsey-Musolf 12, Niemi, Patel, Ramsey-Musolf, Tenkanen, Weir 18 ...

MSSM

Carena, Quiros, Wagner 96, Delepine, Gerard, Gonzalez Felipe, Weyers 96, Cline, Kainulainen 96, Laine, Rummukainen 98, Carena, Nardini, Quiros, Wagner 09, Cohen, Morrissey, Pierce 12, Curtin, Jaiswal, Meade 12, Carena, Nardini, Quiros, Wagner 13, Katz, Perelstein, Ramsey-Musolf, Winslow 14...

NMSSM...

Pietroni 93, Davies, Froggatt, Moorhouse 95, Huber, Schmidt 01, Ham, Oh, Kim, Yoo, Son 04, Menon, Morrissey, Wagner 04, Funakubo, Tao, Yokoda 05, Huber, Konstandin, Prokopec, Schmidt 07, Chung, Long 10, Kozaczuk, Profumo, Stephenson Haskins, Wainwright 15...

Models & Phenomenology

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MSSM

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NMSSM...

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Models & pheno: how reliable?

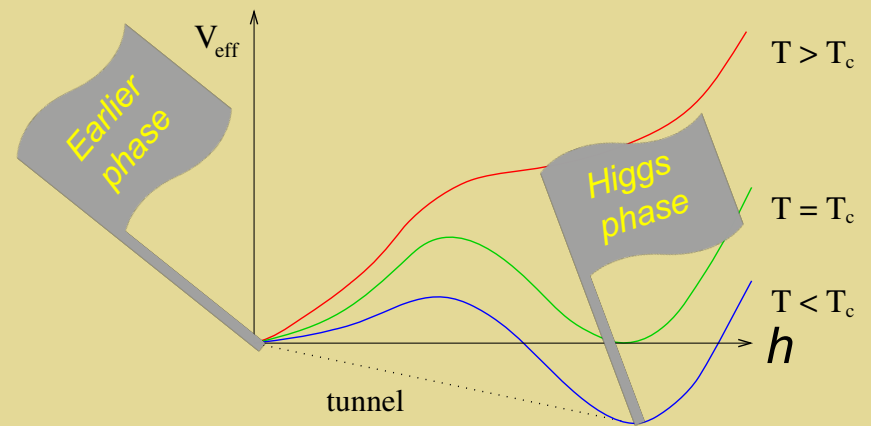
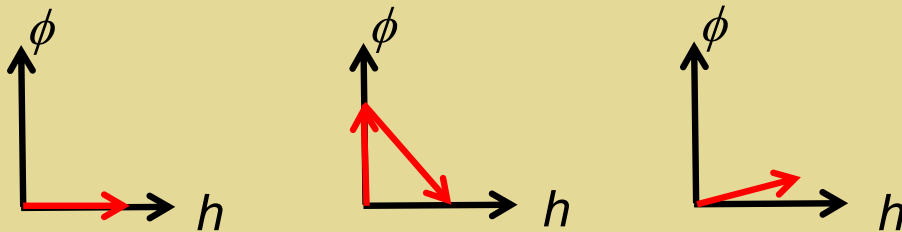
Theory-Pheno Interface



Simple Higgs portal models:

- *Real gauge singlet (SM + 1)*
- *Real EW triplet (SM + 3)*

$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



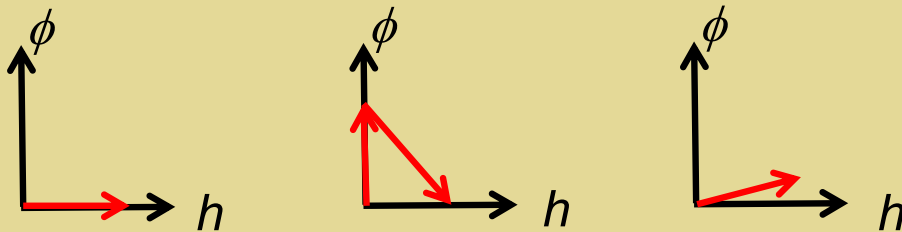
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Phenomenology

$$h_1 = \sin \theta s + \cos \theta h$$

$$h_2 = \cos \theta s - \sin \theta h$$

$m_{1,2}; \theta; h_i h_j h_k$ couplings

Collider Probes

- *Resonant di-Higgs ($h_1 h_1$) production **
- *Heavy h_2 production **
- *Associated production ($Z h_1$) and non-resonant di-Higgs production **
- *Exotic Higgs decays ***

** Heavy h_2*

*** Light h_2*

Collider Probes

- *Resonant di-Higgs ($h_1 h_1$) production **

- *Heavy h_2 production **

- *Associated production ($Z h_1$) and non-resonant di-Higgs production **

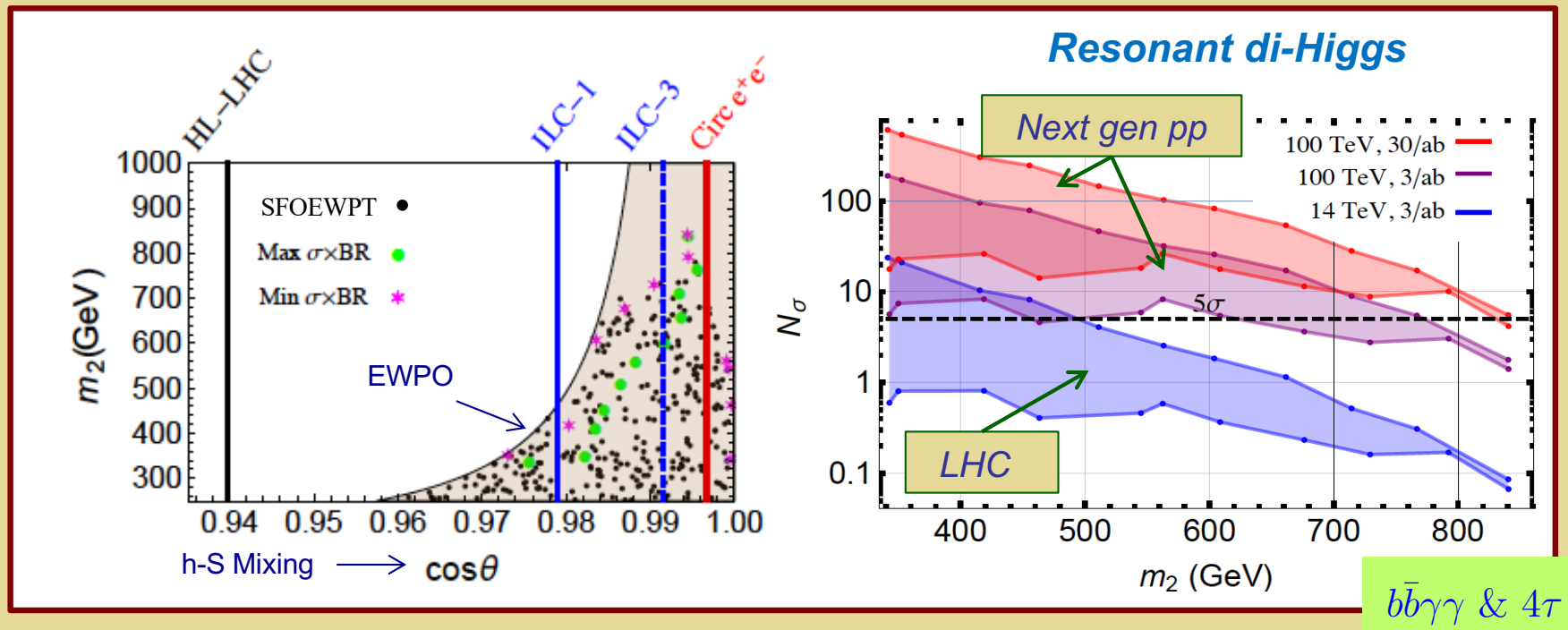
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** Heavy h_2*

*** Light h_2*

Singlets: Precision & Res Di-Higgs Prod

SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies

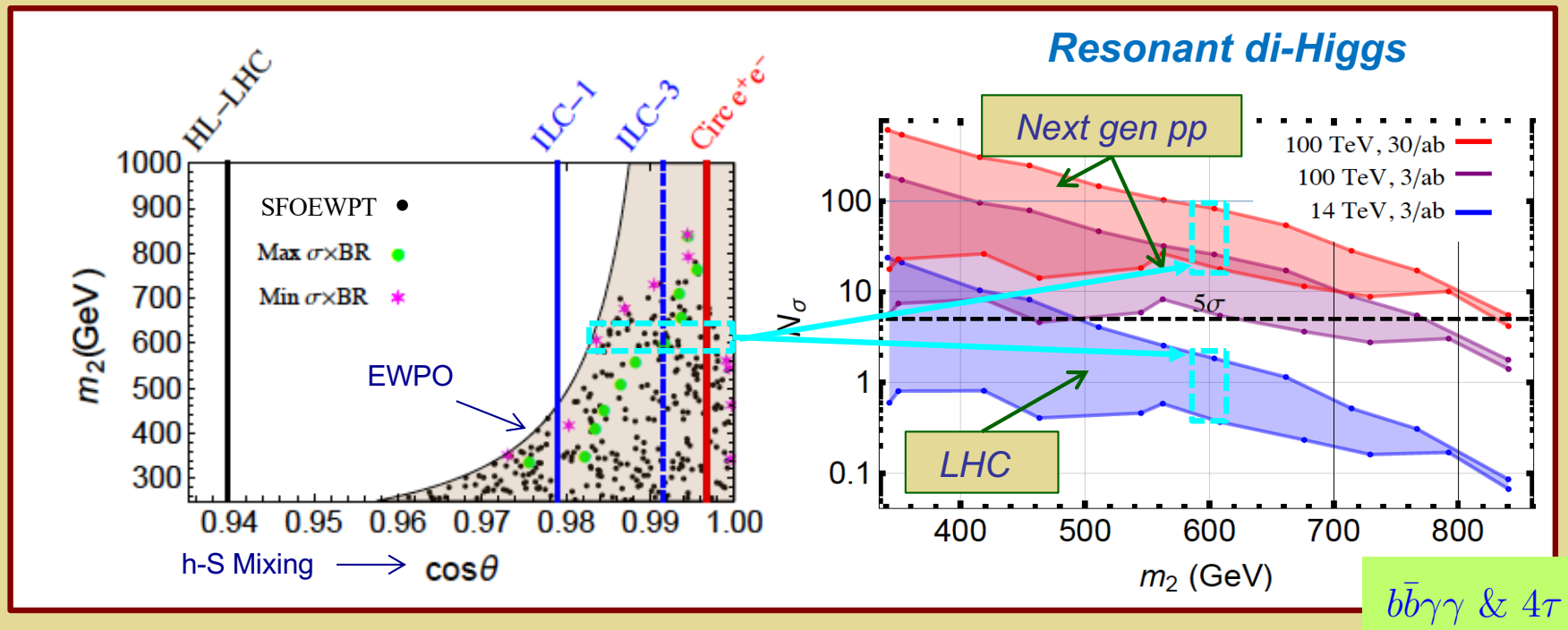


Kotwal, No, R-M, Winslow 1605.06123

See also: Huang et al, 1701.04442;
Li et al, 1906.05289

Singlets: Precision & Res Di-Higgs Prod

SFOEWPT Benchmarks: Resonant di-Higgs & precision Higgs studies

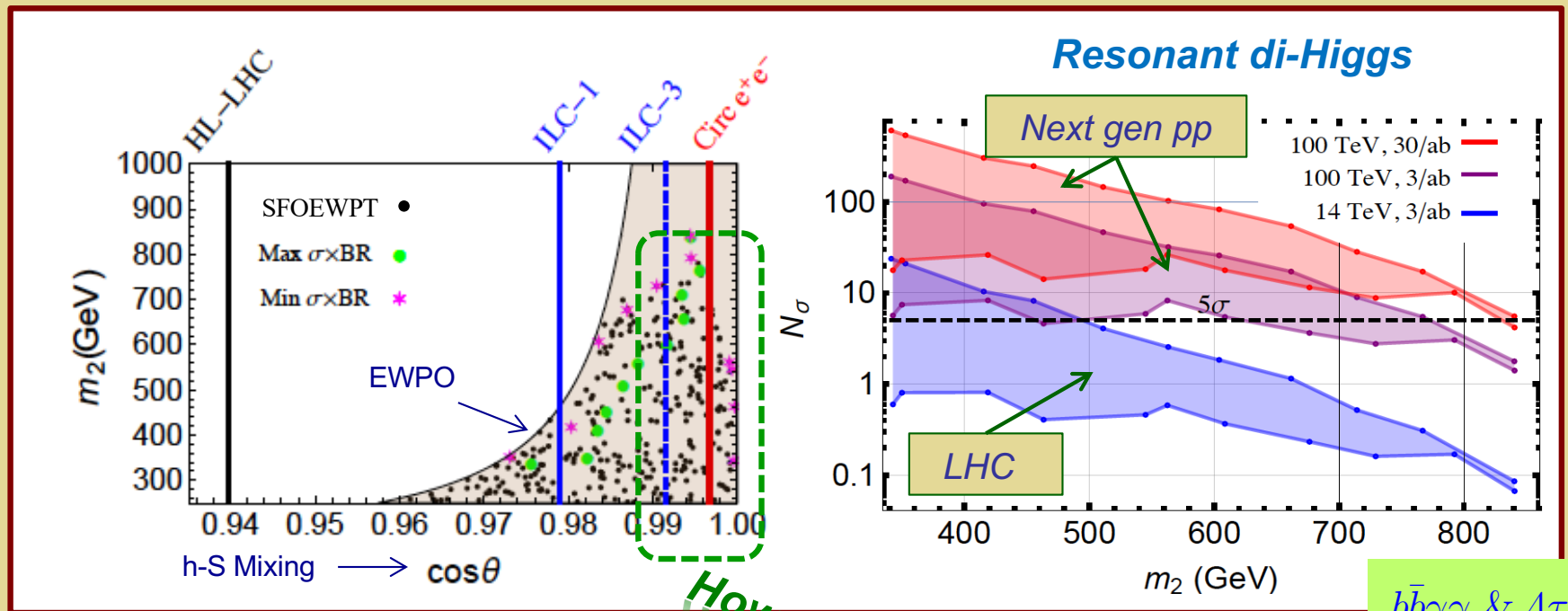


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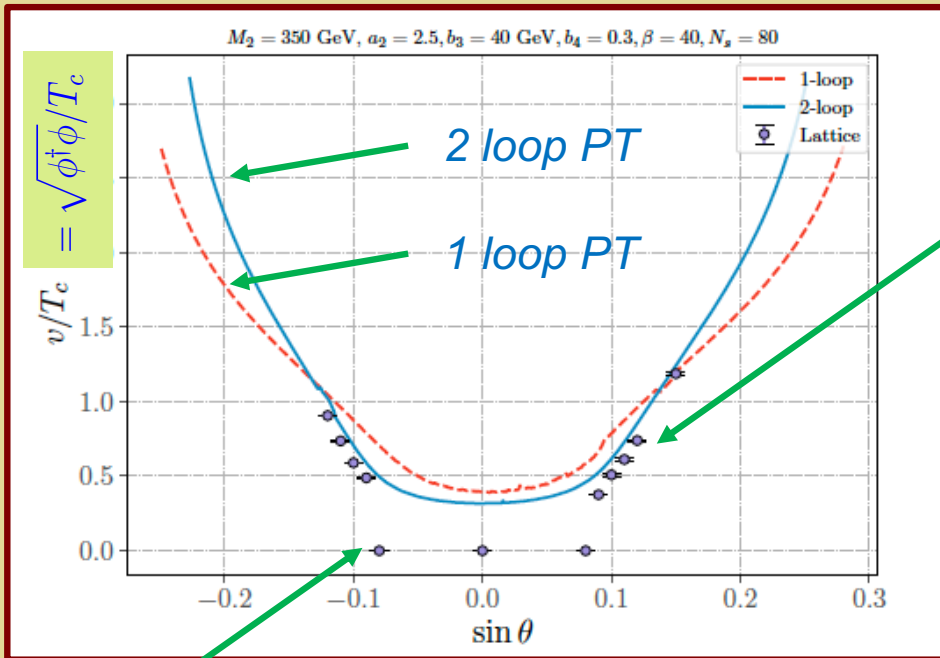
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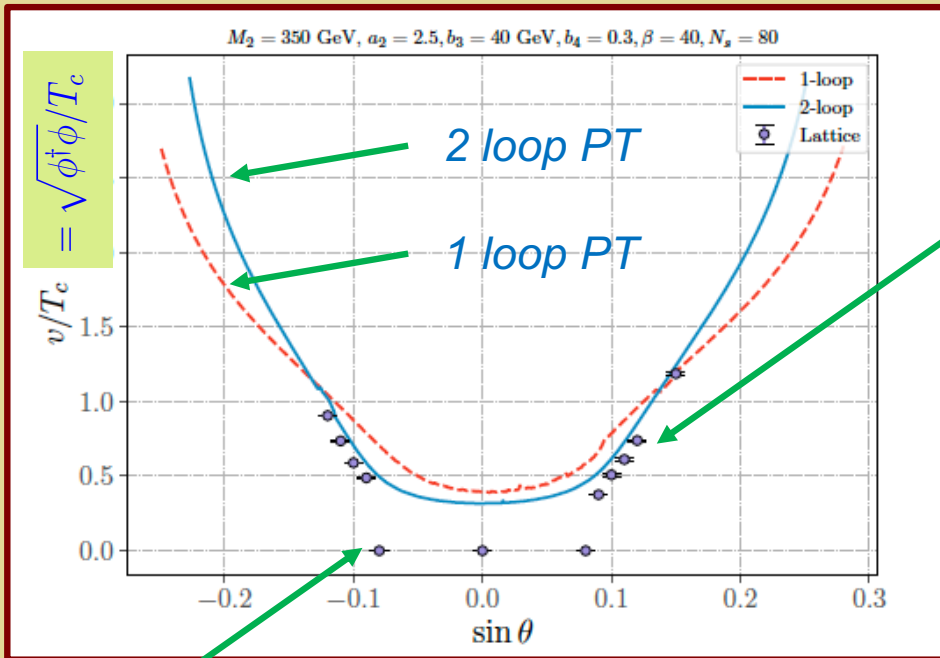
Singlets: Lattice vs. Pert Theory



Lattice:
FOEWPT

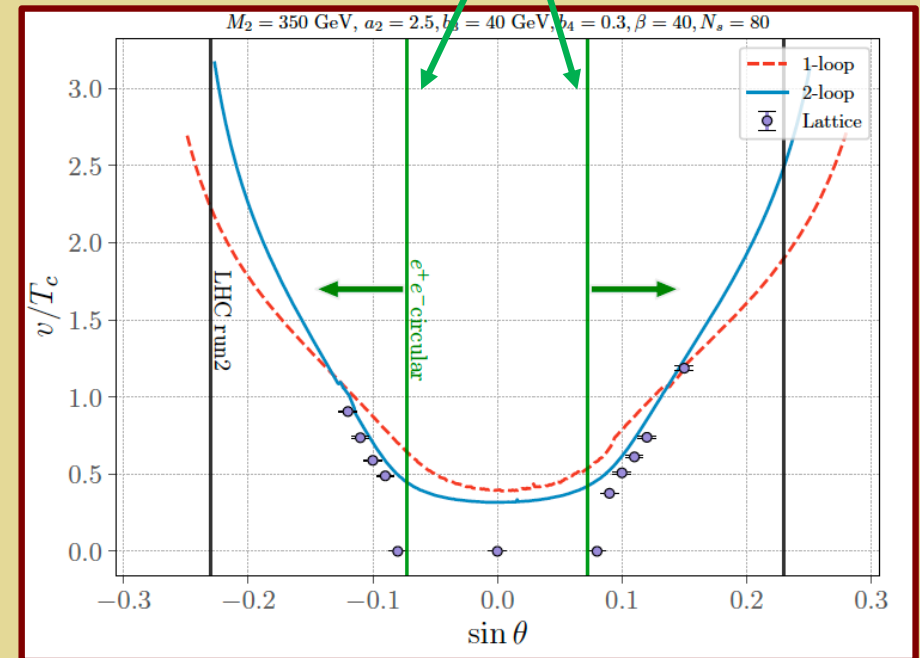
Lattice:
Crossover

Singlets: Lattice vs. Pert Theory



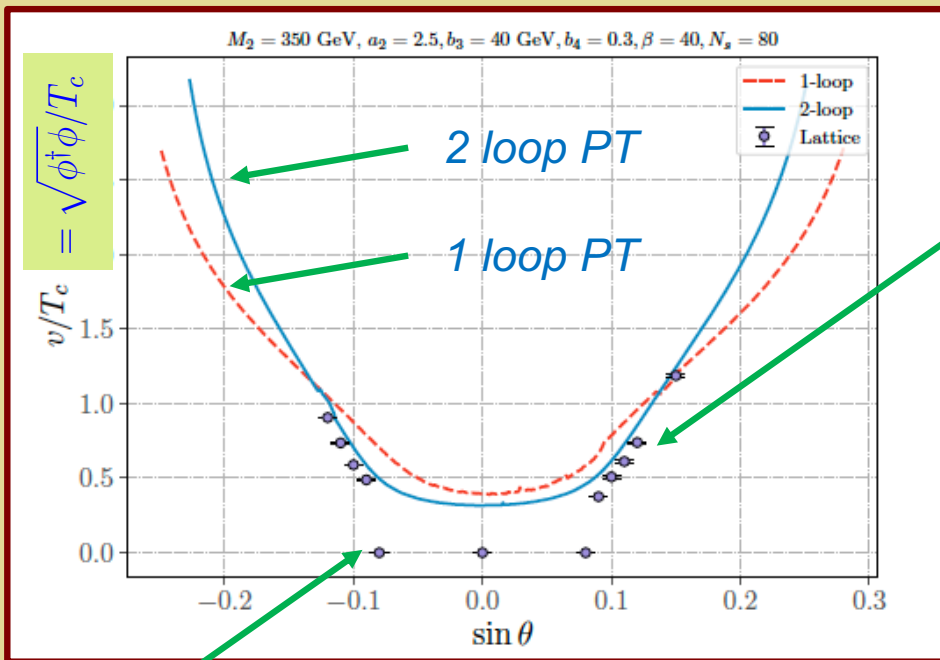
Lattice:
FOEWPT

Future e^+e^-



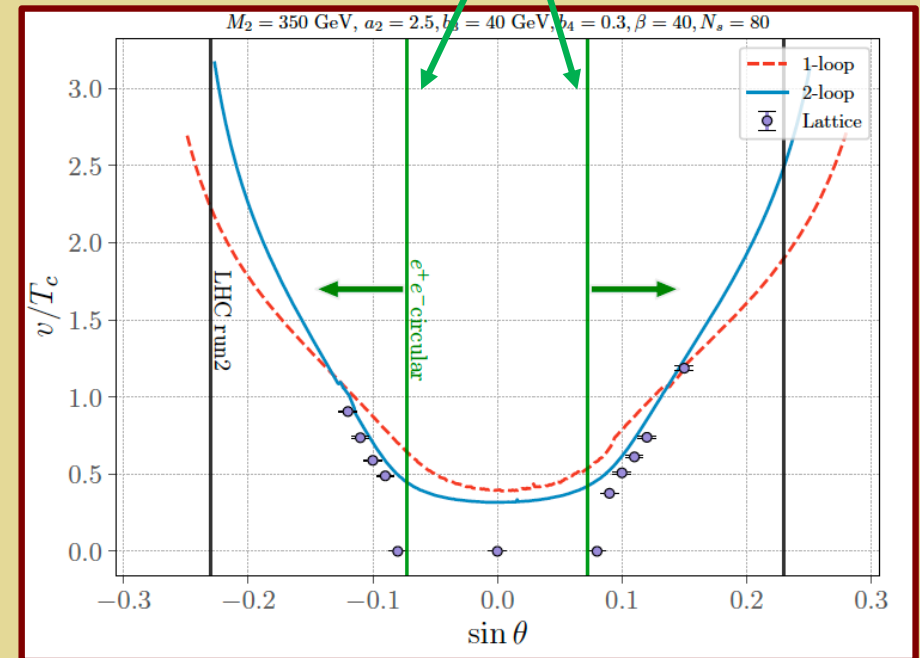
Lattice:
Crossover

Singlets: Lattice vs. Pert Theory



Lattice:
FOEWPT

Future e^+e^-



Lattice:
Crossover

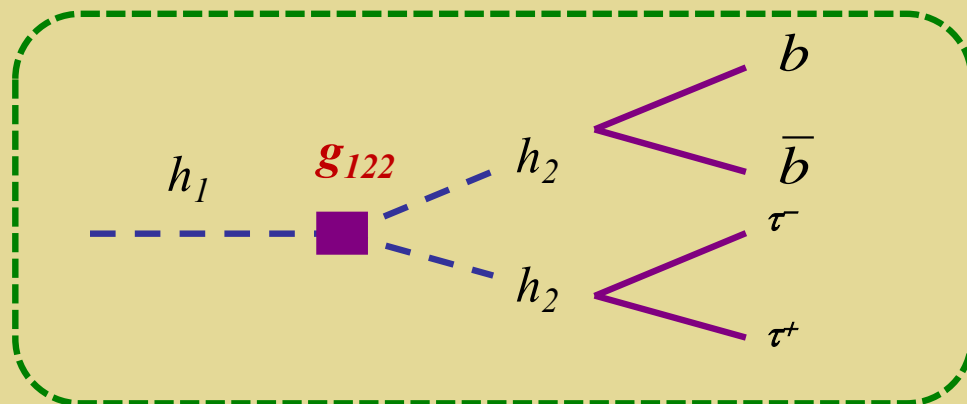
- Lattice: crossover-FOEWPT boundary
- FOEWPT region: PT-lattice agreement
- Pheno: precision Higgs studies may be sensitive to a greater portion of FOEWPT-viable param space than earlier realized

Collider Probes

- Resonant di-Higgs ($h_1 h_1$) production *
- Heavy h_2 production *
- Associated production ($Z h_1$) and non-resonant di-Higgs production *
- Exotic Higgs decays **

* Heavy h_2

** Light h_2



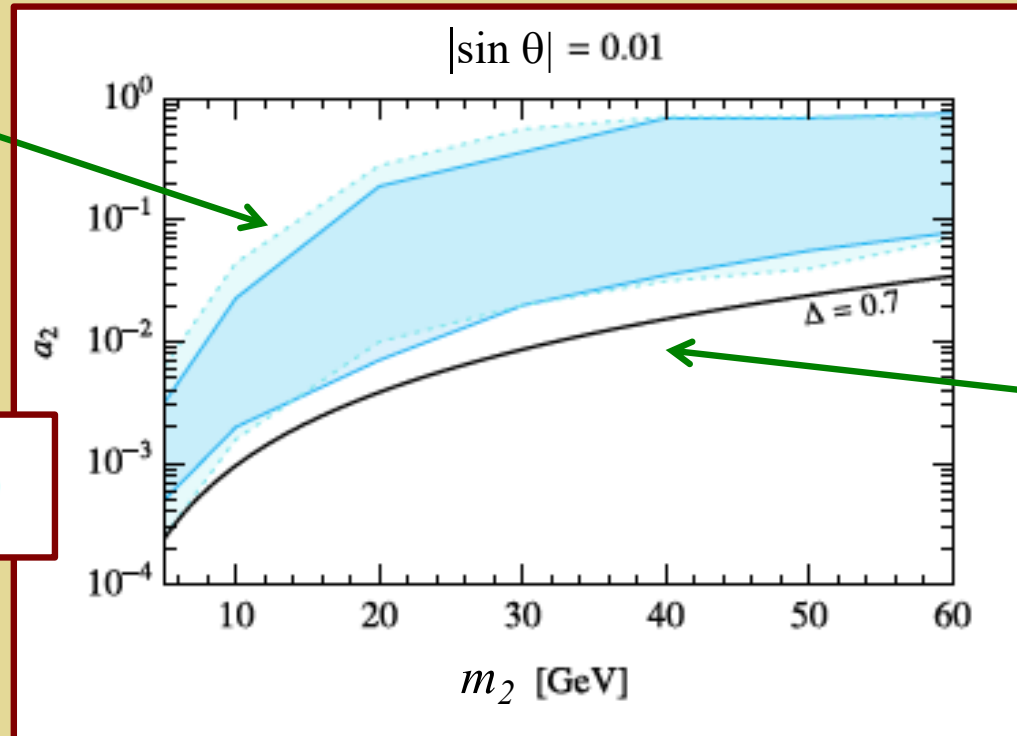
Light Singlets: Exotic Higgs Decays

One loop perturbation theory

EWPT viable:
numerical

$$g_{122} = \frac{1}{2}va_2 + \mathcal{O}(\theta^2)$$

$$h_1 \rightarrow h_2 h_2$$



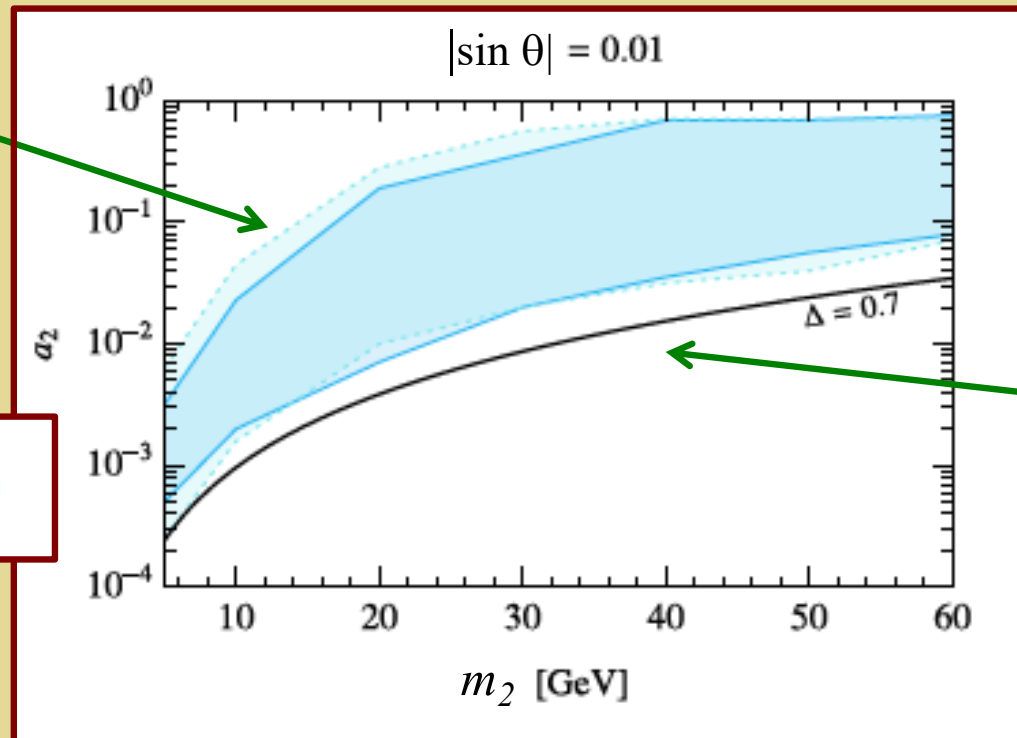
EWPT viable:
Semi analytic
 \rightarrow nucleation
decisive

J. Kozaczuk, MR-M, J. Shelton 1911.10210
See also: Carena et al 1911.10206, Carena et al
2203.08206, Wang et al 2203.10184

Light Singlets: Exotic Higgs Decays

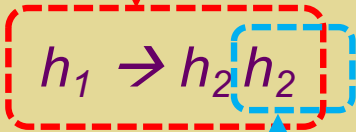
One loop perturbation theory

EWPT viable:
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EWPT viable:
Semi analytic
→ nucleation
decisive

$$g_{122} = \frac{1}{2} v a_2 + \mathcal{O}(\theta^2)$$



$$\Gamma(h_2, m_2) = \sin^2 \theta \Gamma(h_{SM}, m_2)$$

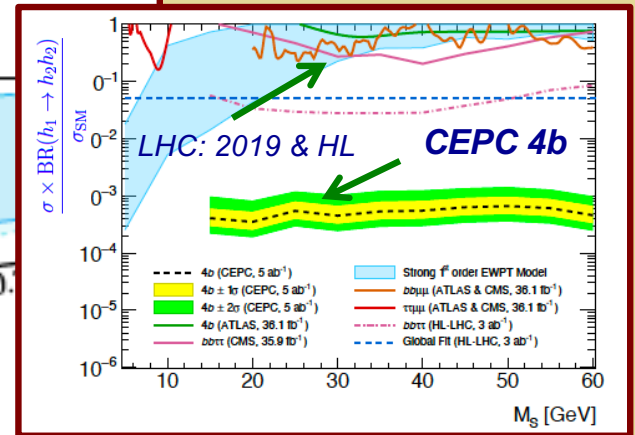
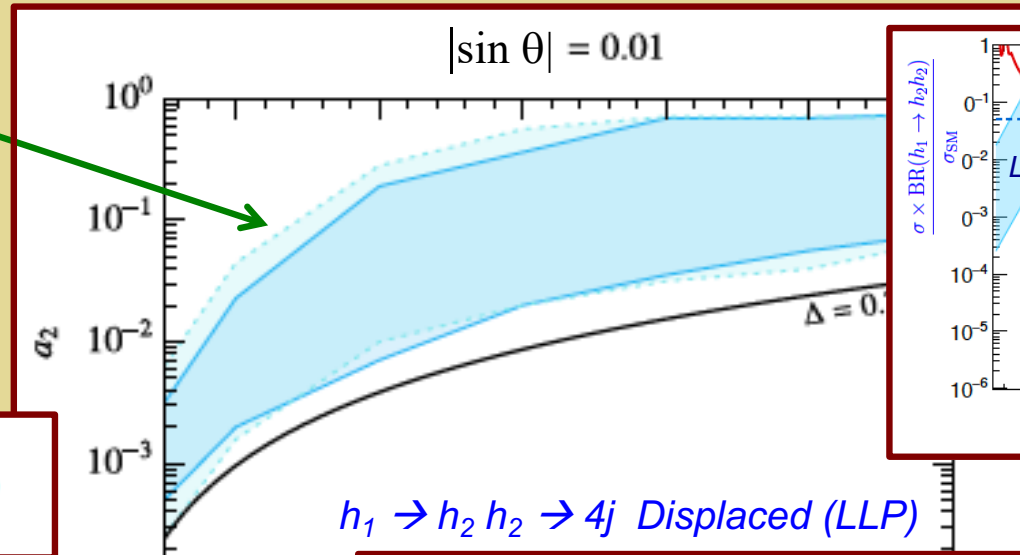
J. Kozaczuk, MR-M, J. Shelton 1911.10210
See also: Carena et al 1911.10206, Carena et al 2203.08206, Wang et al 2203.10184

Light Singlets: Exotic Higgs Decays

One loop perturbation theory

$h_1 \rightarrow h_2 h_2 \rightarrow 4b$ (prompt)

EWPT viable:
numerical



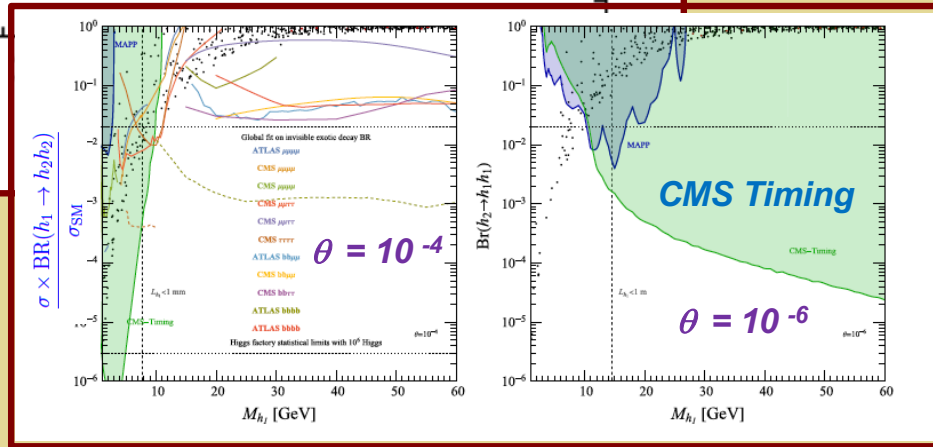
$$g_{122} = \frac{1}{2} v a_2 + \mathcal{O}(\theta^2)$$

$$h_1 \rightarrow h_2 h_2$$

$$\Gamma(h_2, m_2) = \sin^2 \theta \Gamma(h_{SM}, m_2)$$

$h_1 \rightarrow h_2 h_2 \rightarrow 4j$ Displaced (LLP)

J. Wang et al (Snowmass) 2203.10184



Lifetime frontier

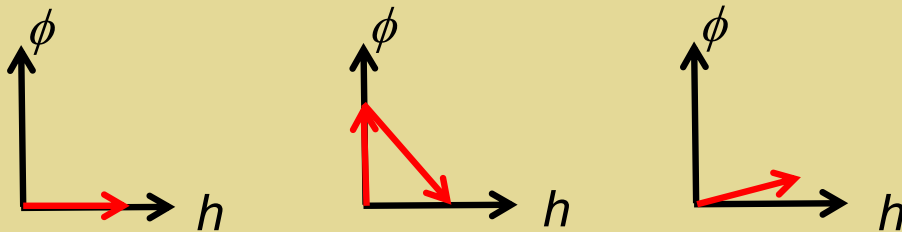
Theory-Pheno Interface



Simple Higgs portal models:

- *Real gauge singlet (SM + 1)*
- *Real EW triplet (SM + 3)*

$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



Theory-Pheno Interface

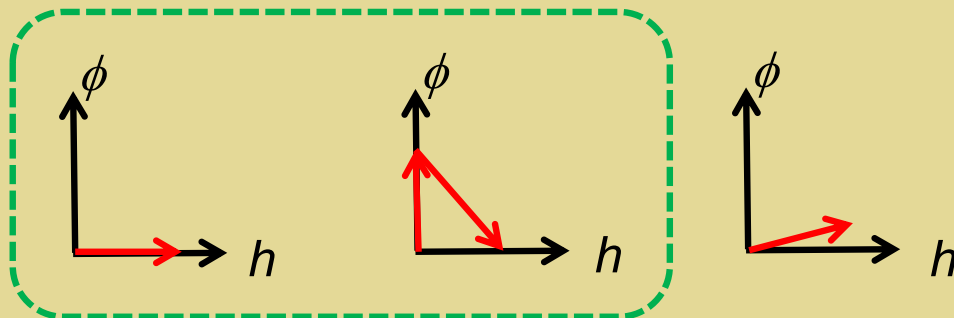


Simple Higgs portal models:

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small

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Theory-Pheno Interface

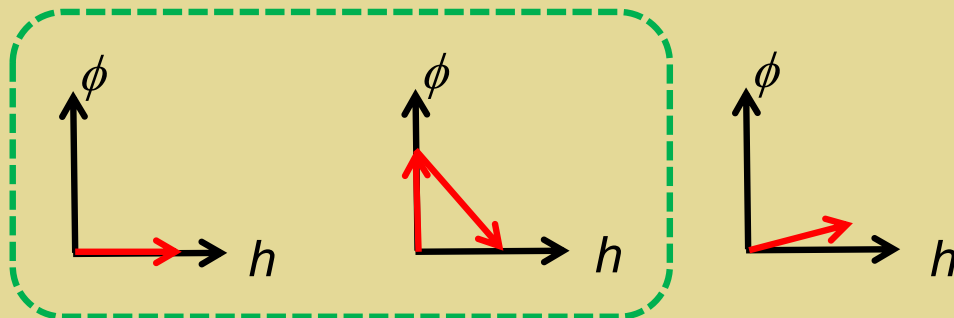


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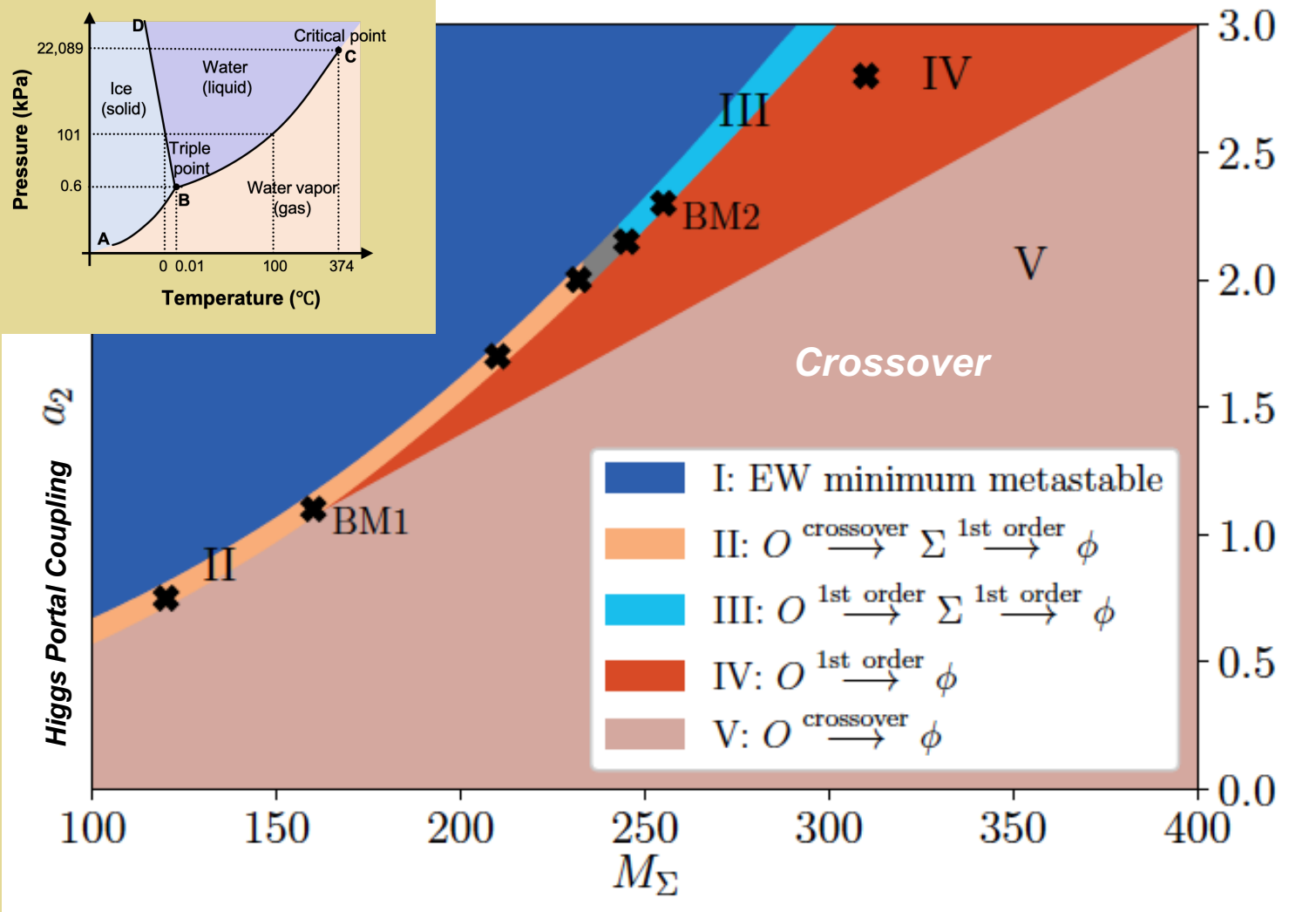
$$V \subset a_1 H^2 \phi + a_2 H^2 \phi^2$$



Phenomenology

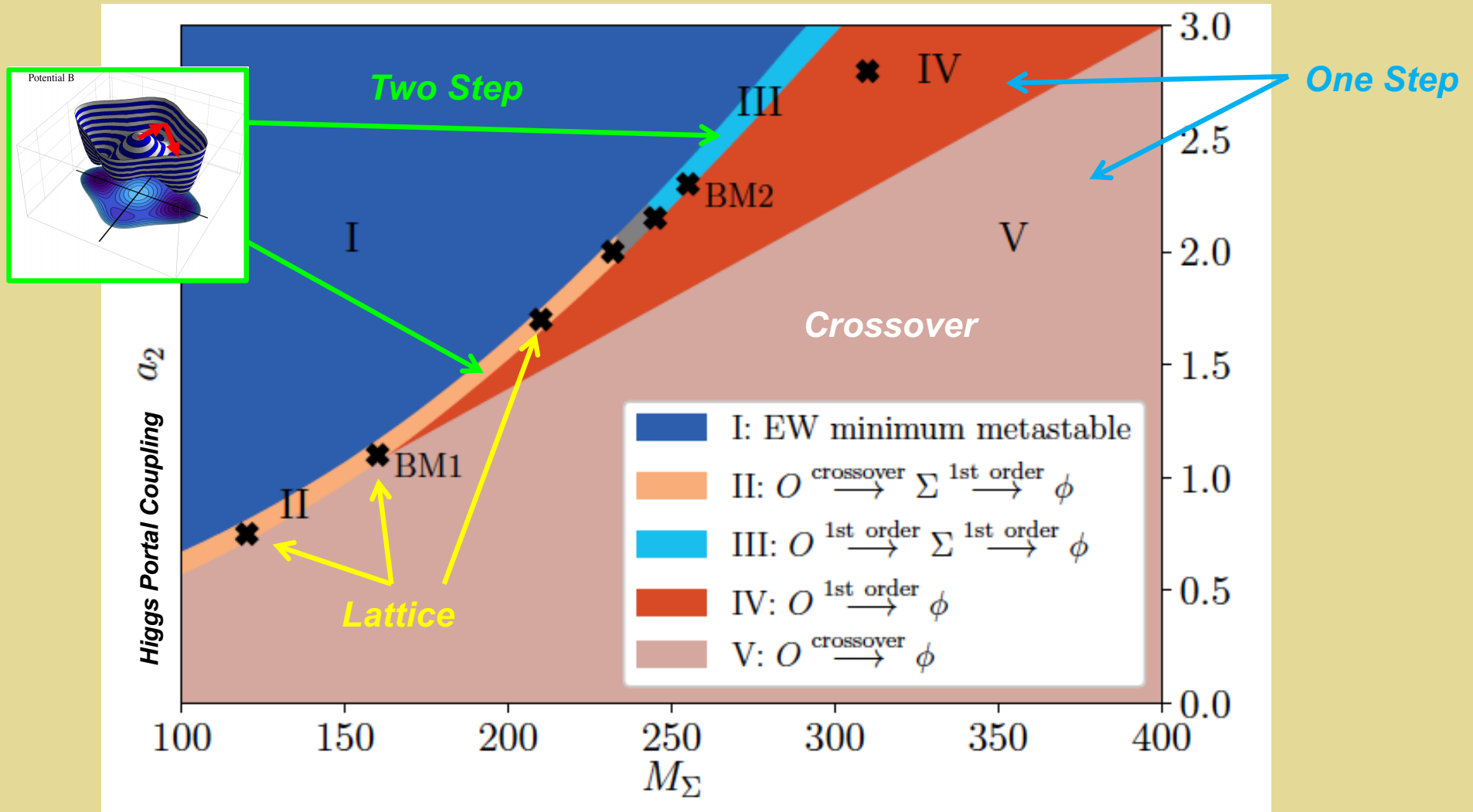
- Gravitational waves
- Collider: $h \rightarrow \gamma\gamma$, dis charged track, NLO $e^+e^- \rightarrow Zh...$

Real Triplet & EWPT: Novel EWSB



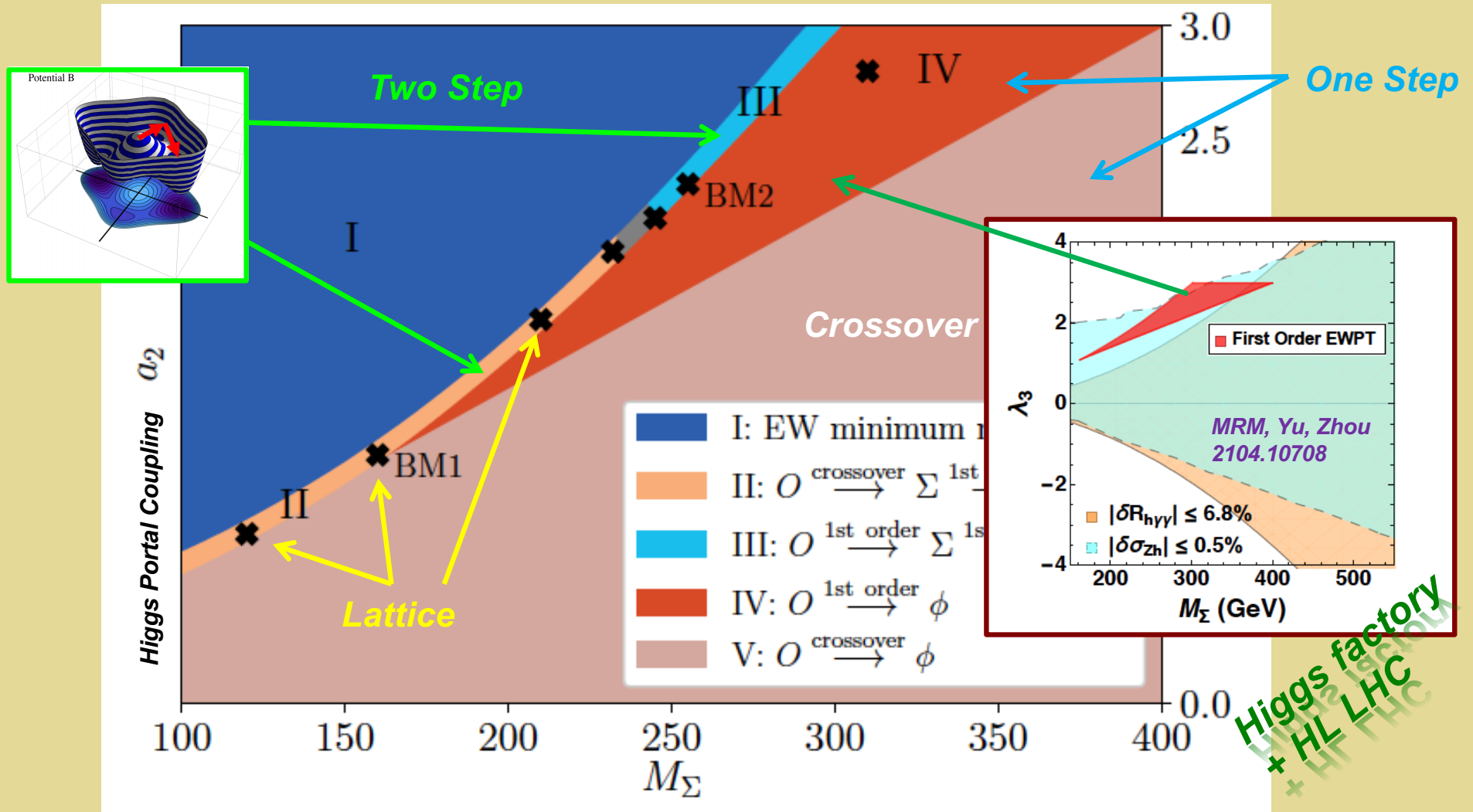
- 1 or 2 step
- Non-perturbative

Real Triplet & EWPT: Novel EWSB



- 1 or 2 step
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Real Triplet & EWPT: Novel EWSB



- 1 or 2 step
- Non-perturbative

BSM EWPT: Inter-frontier Connections

*Robust theory:
EFT + lattice*



*Observables:
model specific*



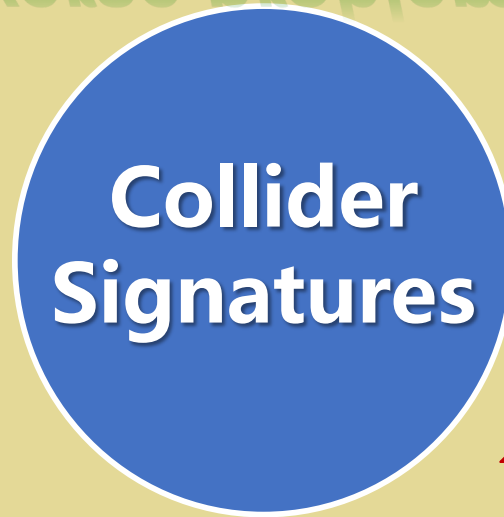
*Hydro:
 $\alpha, \beta / H_*$*

BSM EWPT: Inter-frontier Connections

GW – Collider “inverse problem” **

**Robust theory:
EFT + lattice**

**Observables:
model specific**

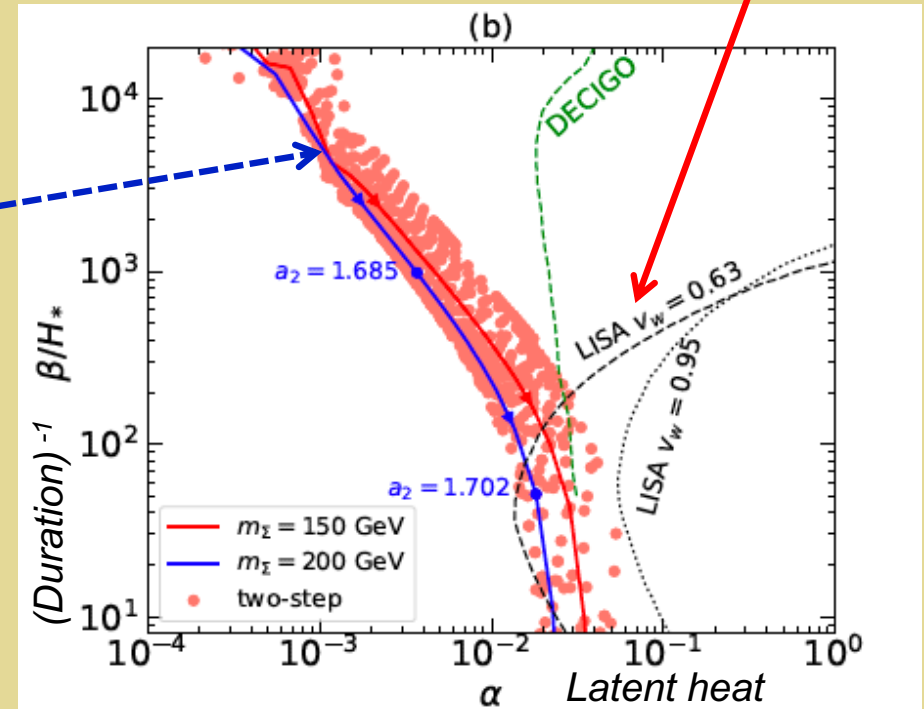
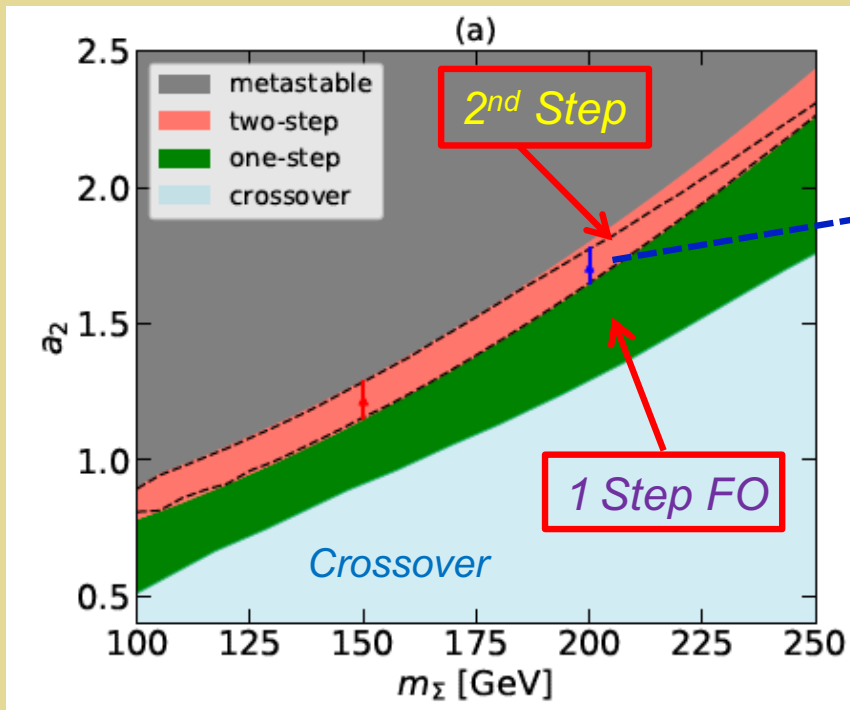


**Combined
reach**

**Hydro:
 $\alpha, \beta / H^*$**

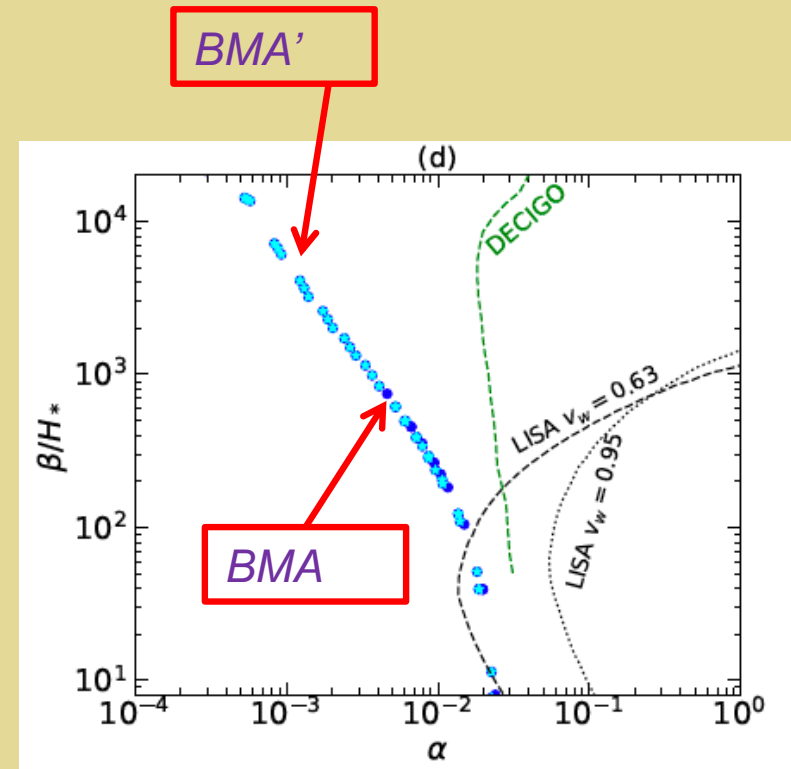
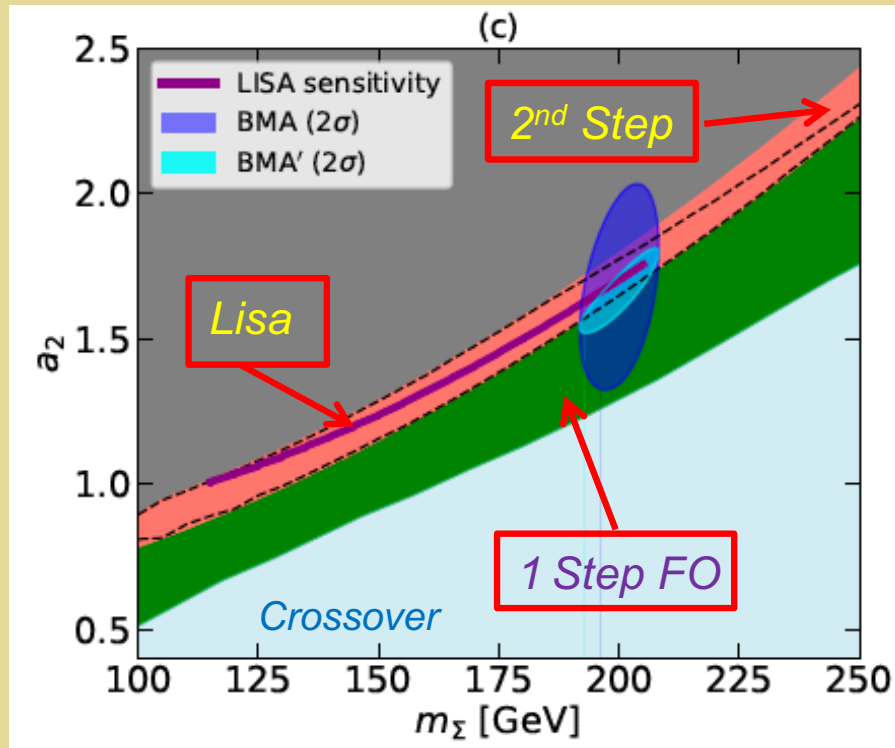
**** How can we exploit experiment to identify EWPT-viable models & parameters ?**

GW & EWPT Phase Diagram



- *Single step transition: GW well outside LISA sensitivity*
- *Second step of 2-step transition can be observable*
- *Significant GW sensitivity to portal coupling*

GW & EWPT Phase Diagram



$$BMA: m_\Sigma + h \rightarrow \gamma\gamma$$

$$BMA': BMA + \Sigma^0 \rightarrow ZZ$$

- *Two-step*
- *EFT+ Non-perturbative*

EWBG Ingredients

- ***EW Sphalerons***



- ***Strong 1st Order EW Phase Transition***



BSM Higgs

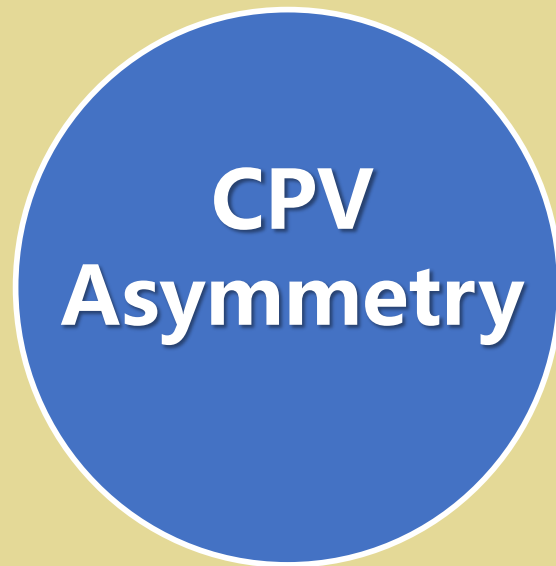
- ***Left-handed number density***



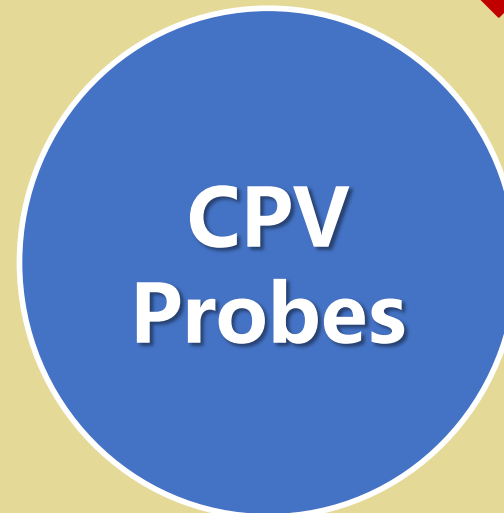
BSM CPV

BSM CPV: Inter-frontier Connections

*Robust theory:
Quantum transport,
bubble dynamics*



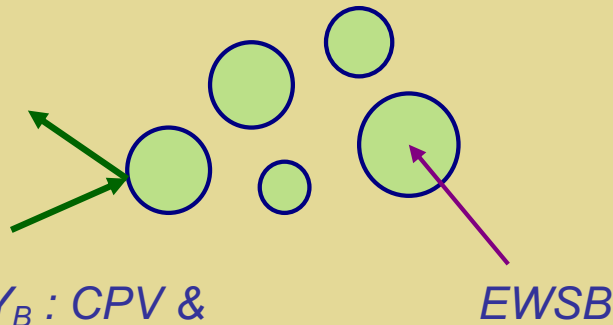
*Models, other
pheno...*



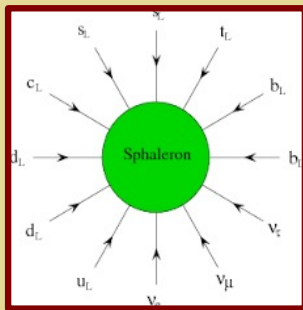
*EDM, heavy
flavor...*

CPV for EWBG

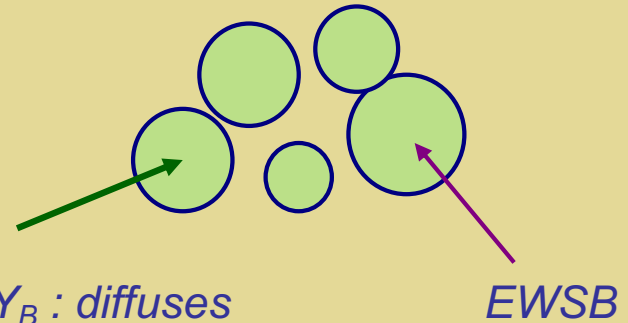
1st order EWPT



Y_B : CPV & EW sphalerons

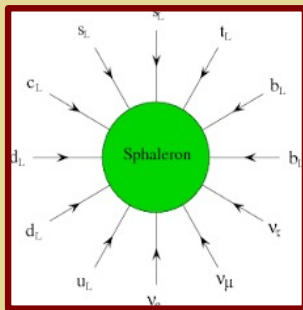
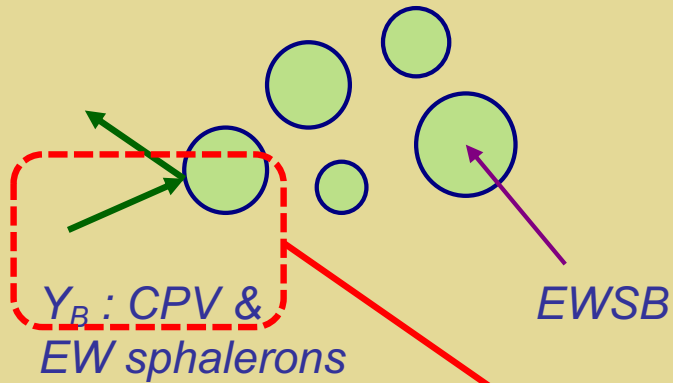


1st order EWPT →
 “strong” to preserve Y_B



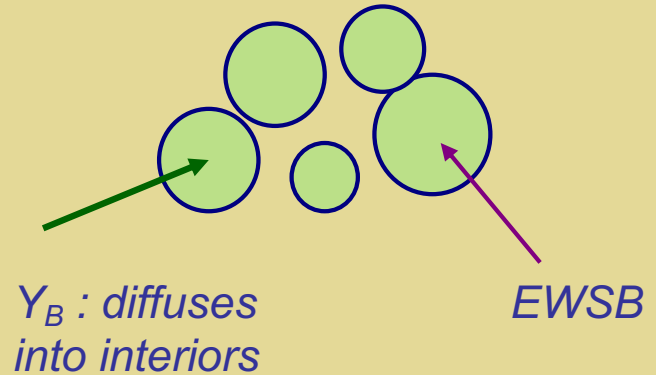
CPV for EWBG

1st order EWPT



- Electric dipole moments
- Heavy flavor CPV
- Higgs CPV

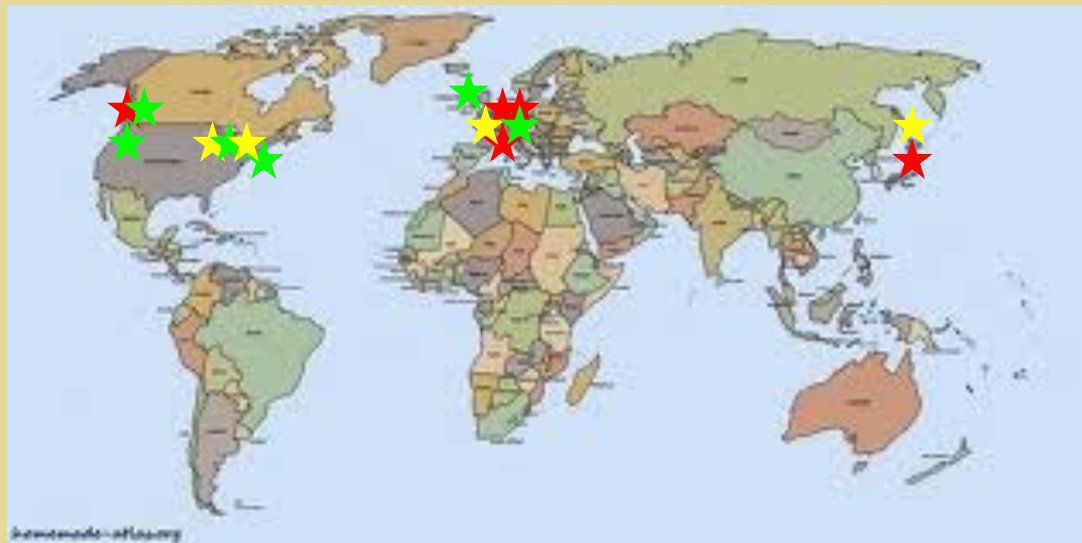
1st order EWPT →
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EDMs: New CPV?

| System | Limit (e cm)* | SM CKM CPV | BSM CPV |
|-------------------|--------------------------|------------|------------|
| ^{199}Hg | 7.4×10^{-30} | 10^{-35} | 10^{-30} |
| HfF ⁺ | 4.1×10^{-30} ** | 10^{-38} | 10^{-30} |
| n | 1.8×10^{-26} | 10^{-31} | 10^{-26} |

* 95% CL ** e⁻ equivalent



- ★ neutron
- ★ proton & nuclei
- ★ atoms

~ 100 x better sensitivity

Not shown:
muon

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Challenge for nuc & AMO



- ★ neutron
- ★ proton & nuclei
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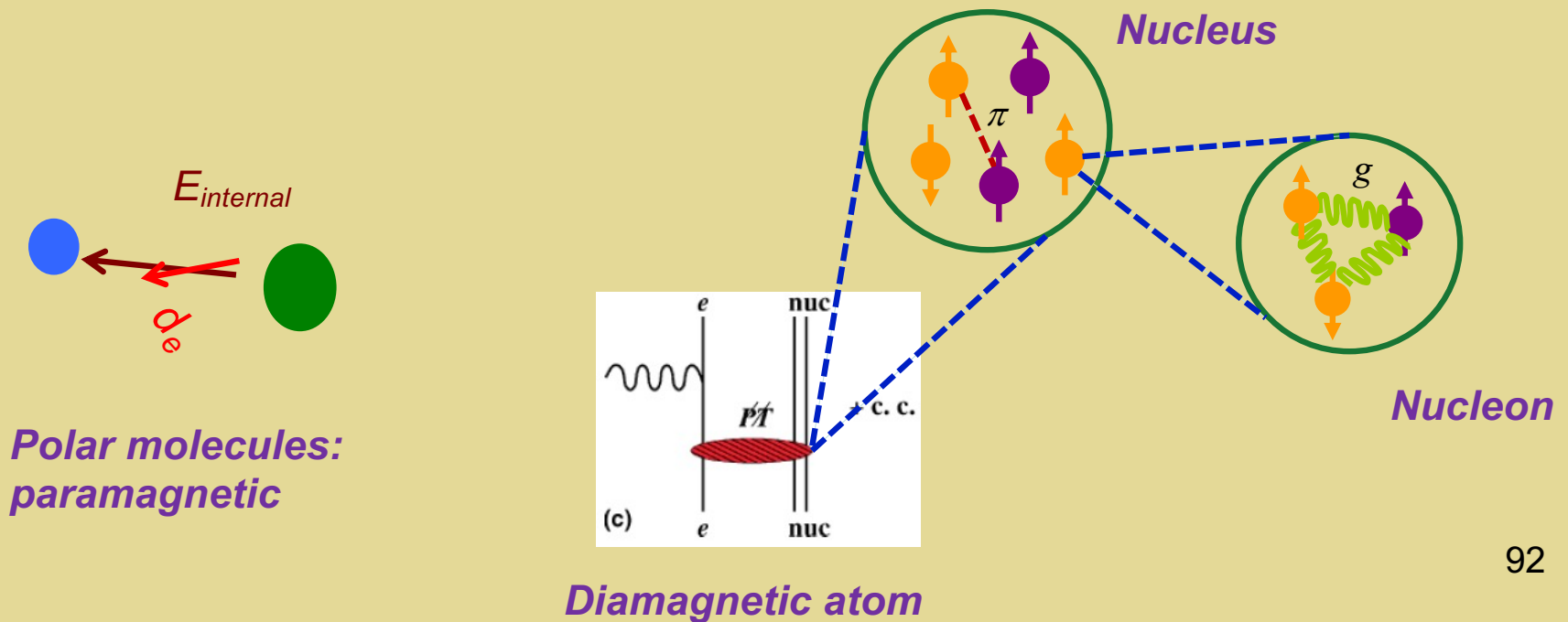
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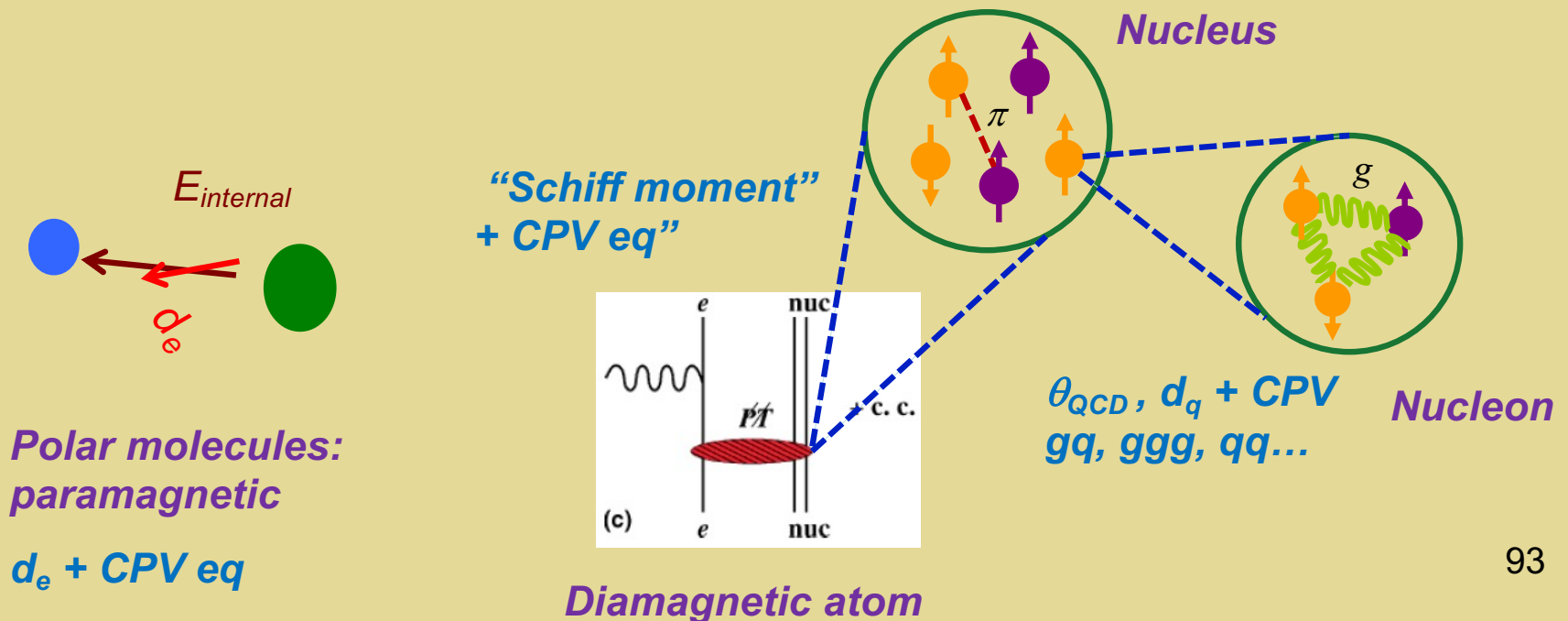
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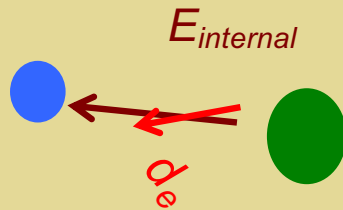
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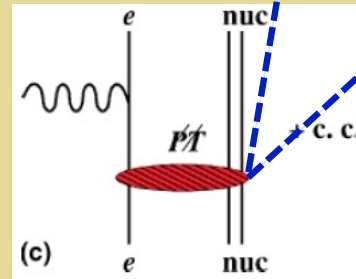
THEORY



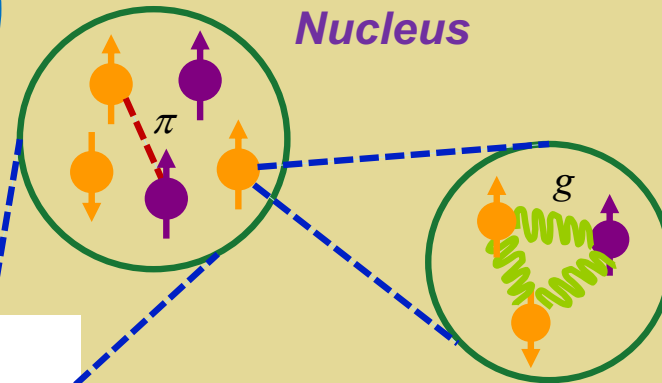
Polar molecules:
paramagnetic

$d_e + CPV eq$

“Schiff moment”
+ CPV eq”



Diamagnetic atom



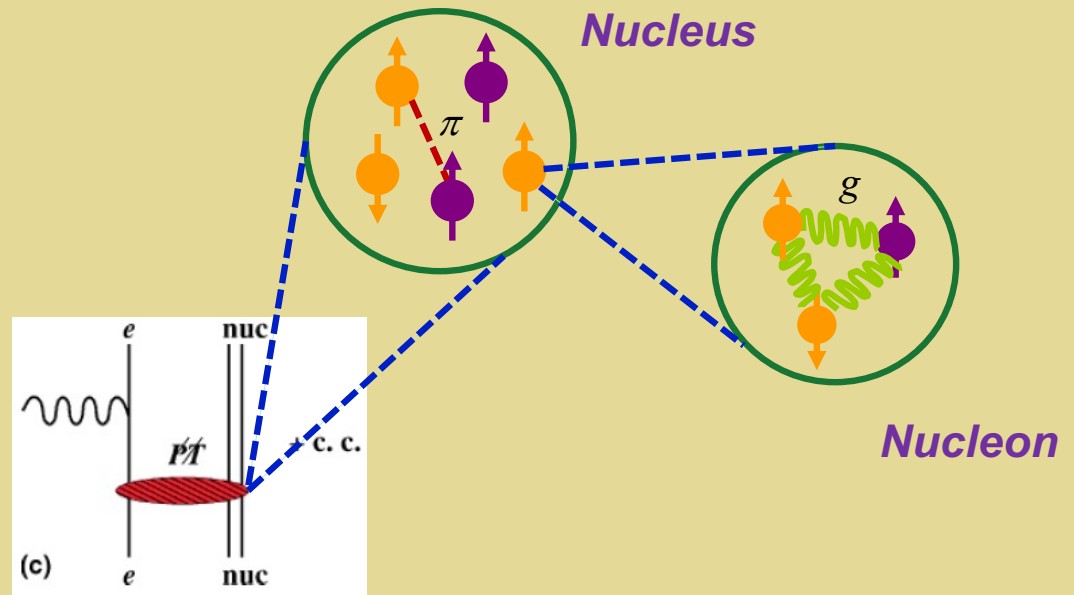
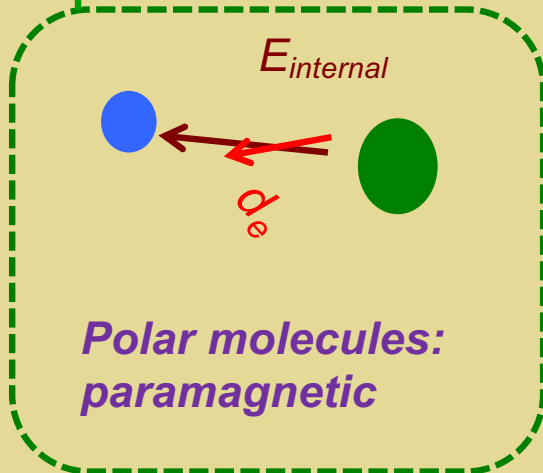
$\theta_{QCD}, d_q + CPV$
 $gg, ggg, qq...$ Nucleon

EDMs: New CPV?

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AMO



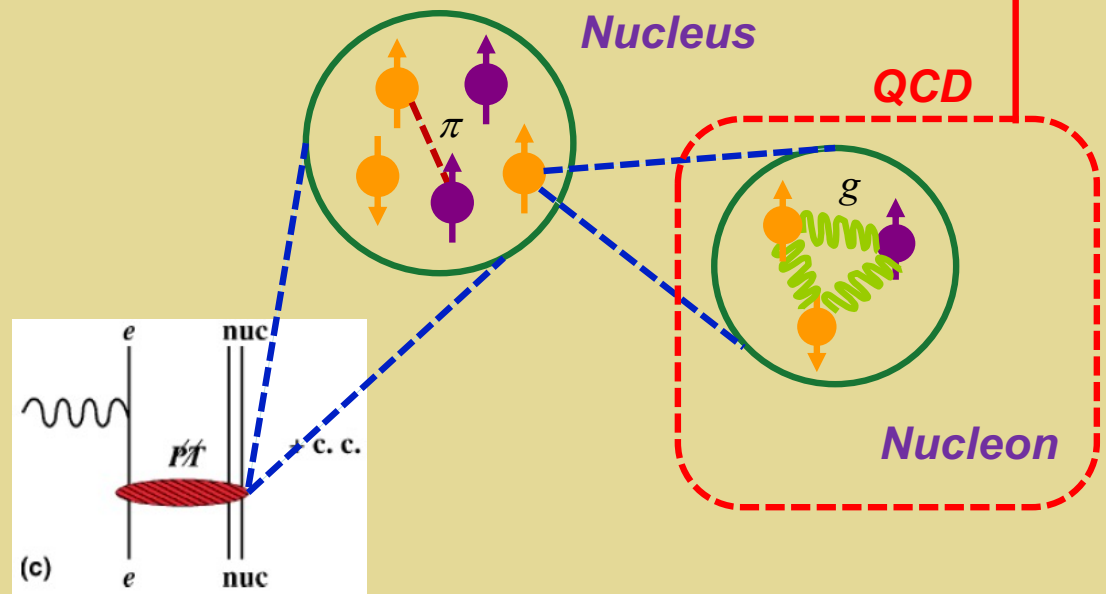
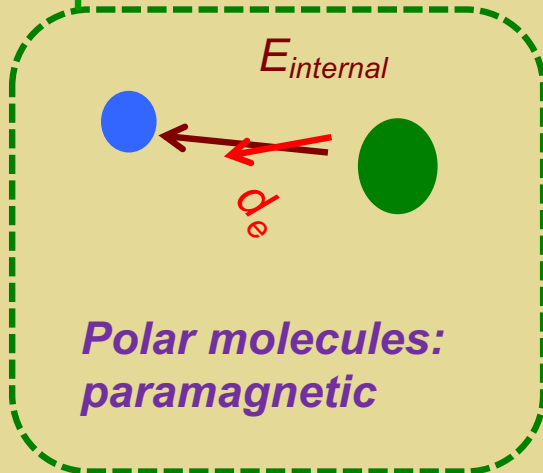
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AMO



Diamagnetic atom

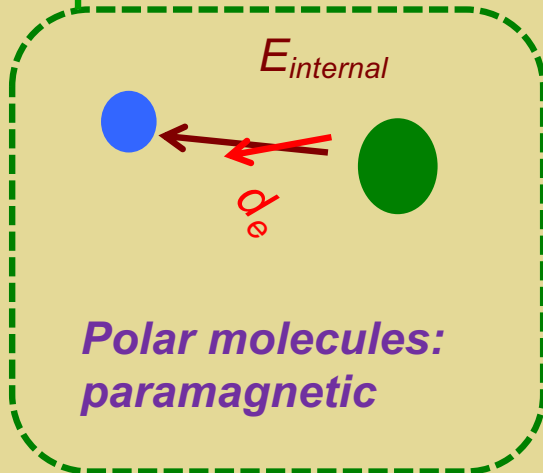
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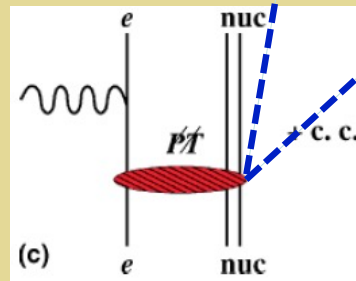
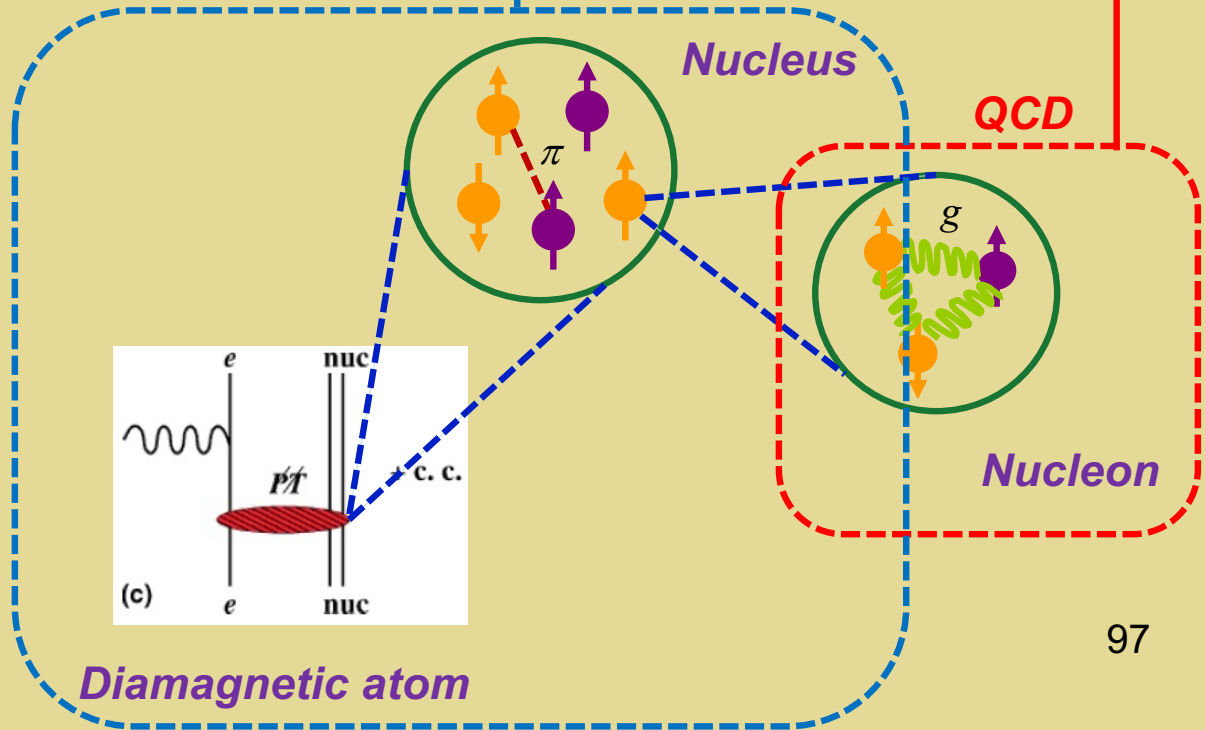
NUC

AMO



Nucleus

QCD



EDMs: New CPV?

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** e⁻ equivalent

Challenge for EWBG
Challenge for EMBC



- ★ neutron
- ★ proton & nuclei
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~ 100 x better sensitivity

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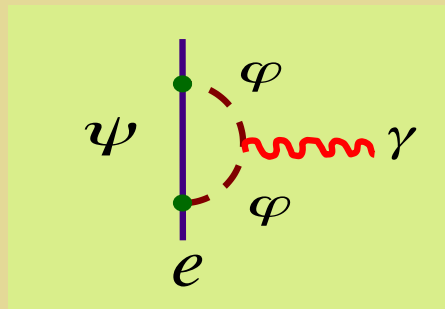
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Challenge for EWBG
Challenge for EMBC

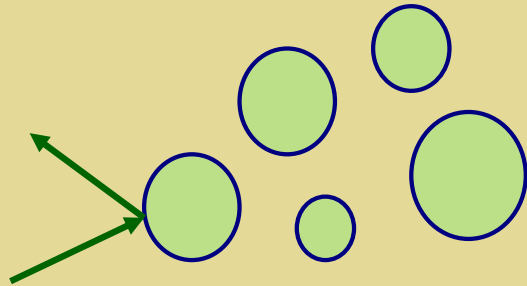
Mass Scale Sensitivity



$$\sin\phi_{\text{CP}} \sim 1 \rightarrow M > 5000 \text{ GeV}$$

$$M < 500 \text{ GeV} \rightarrow \sin\phi_{\text{CP}} < 10^{-2}$$

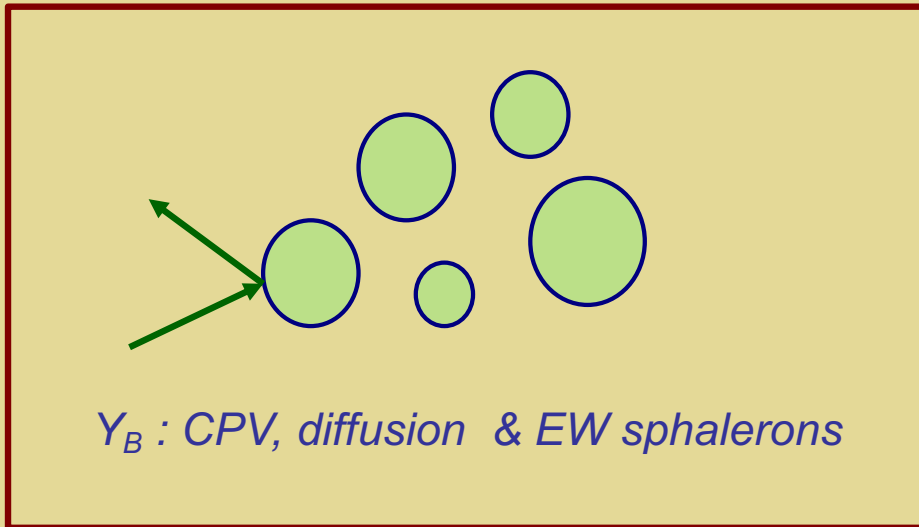
CPV for EWBG



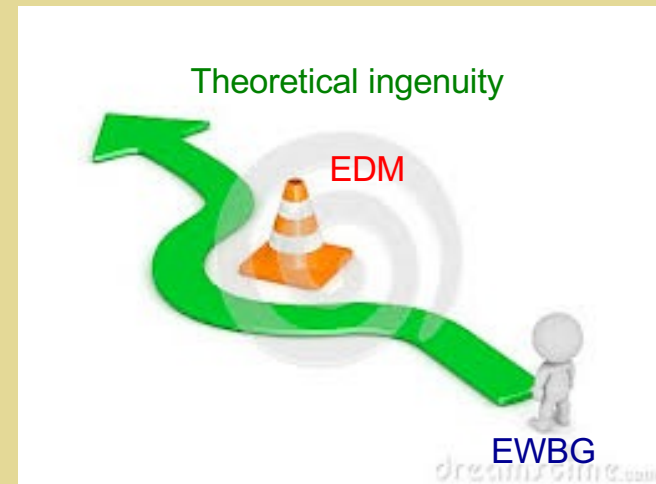
Y_B : *CPV, diffusion & EW sphalerons*

- *Can the CPV interactions be sufficiently “large” ?*
- *Reliable quantum transport computations ?*

CPV for EWBG



- **Can the CPV interactions be sufficiently “large” ?**
- **Reliable quantum transport computations ?**

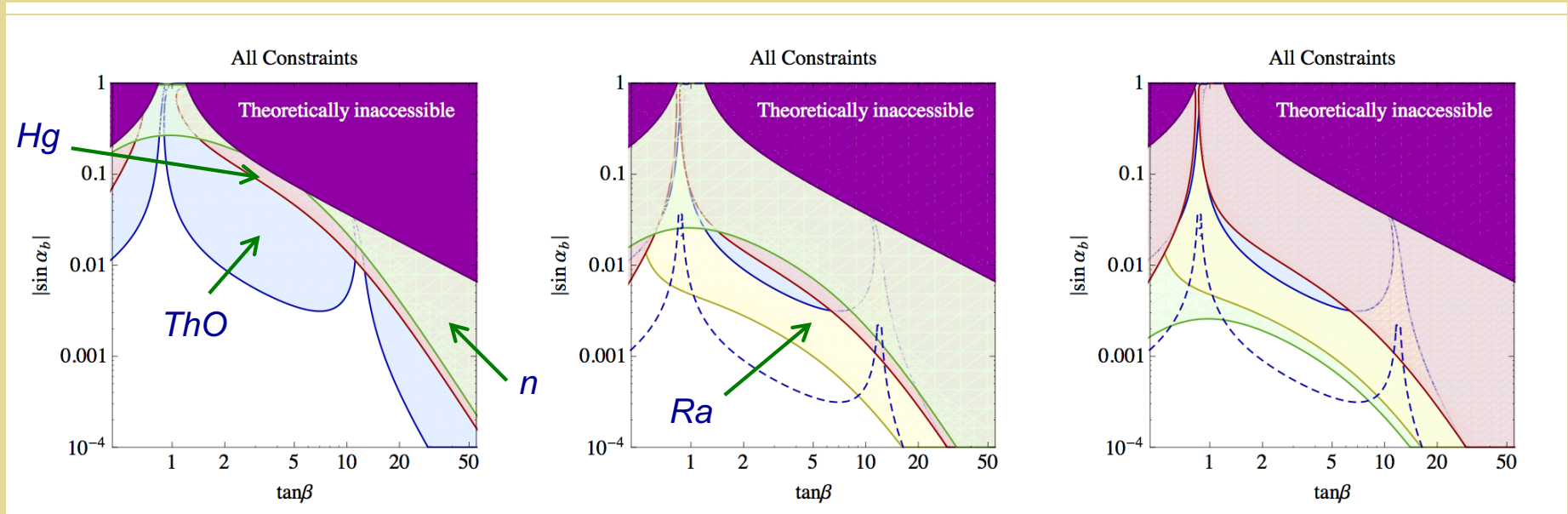


Illustrations

2HDM CPV : EDMs 2014

CPV & 2HDM: Type II illustration

$\lambda_{6,7} = 0$ for simplicity



2014

New HfF⁺

$\sin \alpha_b$: CPV
scalar mixing

Present & Future:

- $d_n \times 0.1$
- $d_A(\text{Hg}) \times 0.1$
- $d_{\text{ThO}} \times 0.1$
- $d_A(\text{Ra}) [10^{-27} \text{ e cm}]$

Future:

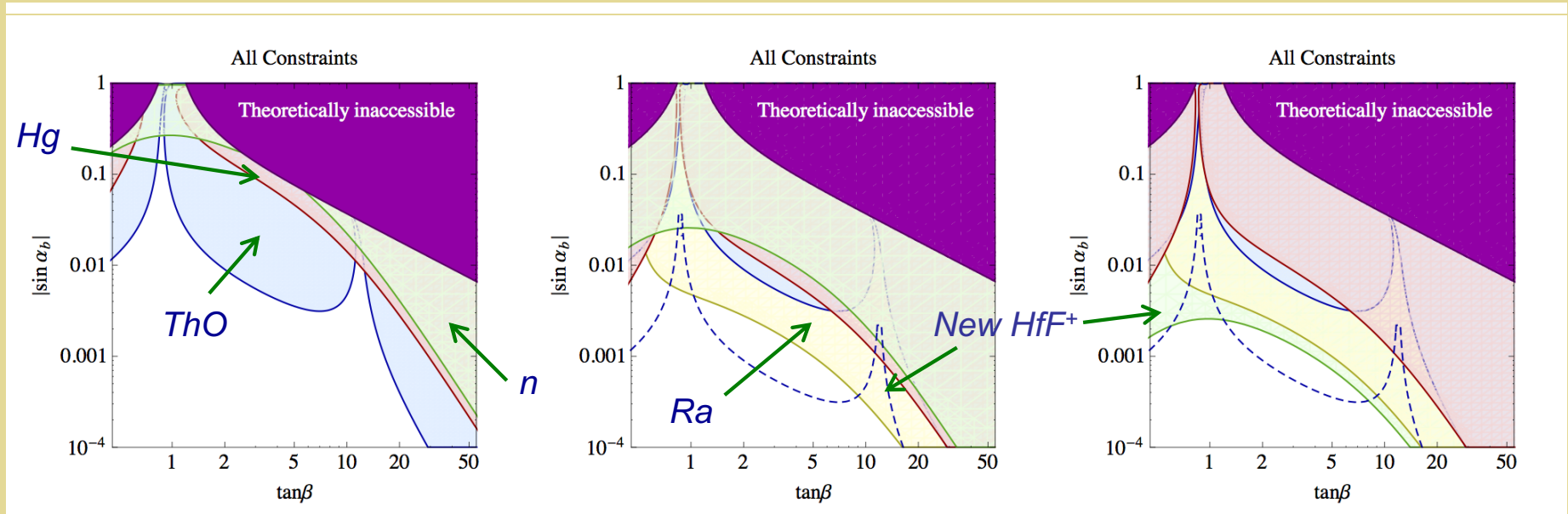
- $d_n \times 0.01$
- $d_A(\text{Hg}) \times 0.1$
- $d_{\text{ThO}} \times 0.1$
- $d_A(\text{Ra})$

Inoue, R-M, Zhang: 1403.4257

2HDM CPV : EDMs & LHC 2024

CPV & 2HDM: Type II illustration

$\lambda_{6,7} = 0$ for simplicity



Present

New HfF+

$\sin \alpha_b$: CPV
scalar mixing

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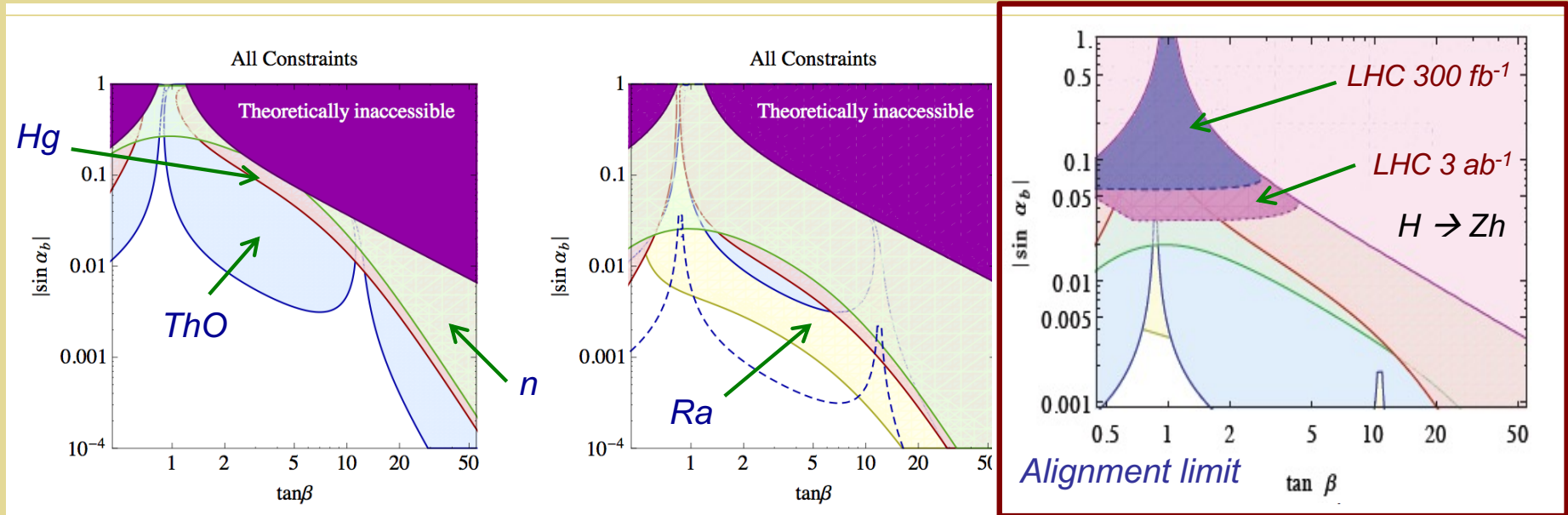
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2HDM CPV : EDMs & LHC 2024

CPV & 2HDM: Type II illustration

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Chen, Li, R-M: 1708.00435

Present

New HfF^+

$\sin\alpha_b$: CPV
scalar mixing

Present & Future:

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Future:

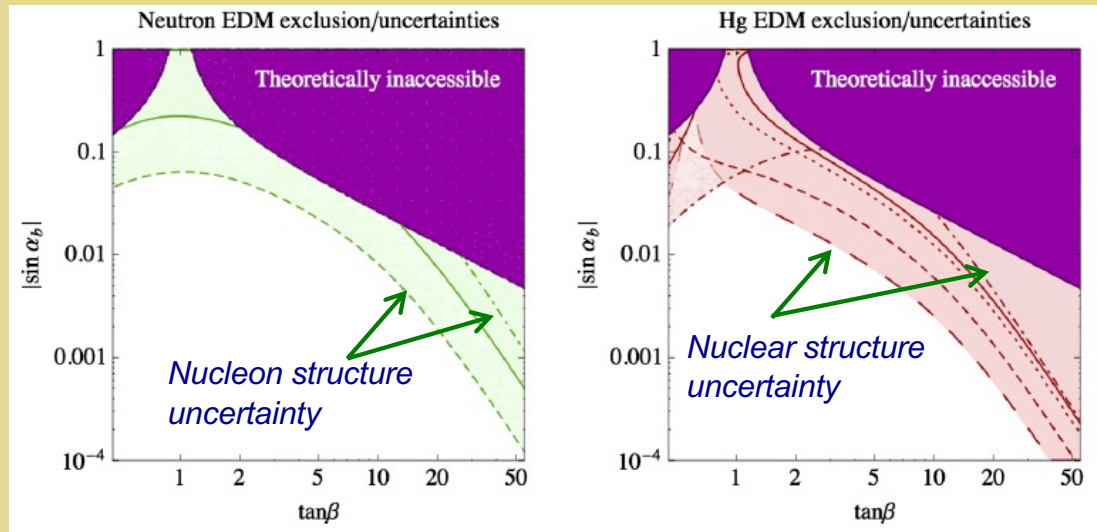
- $d_n \times 0.01$
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- $d_A(Ra)$

Inoue, R-M, Zhang: 1403.4257

2HDM CPV : Had & Nuc Structure

CPV & 2HDM: Type II illustration

$\lambda_{6,7} = 0$ for simplicity



2014

New HfF^+

$\sin \alpha_b$: CPV
scalar mixing

Present & Future:

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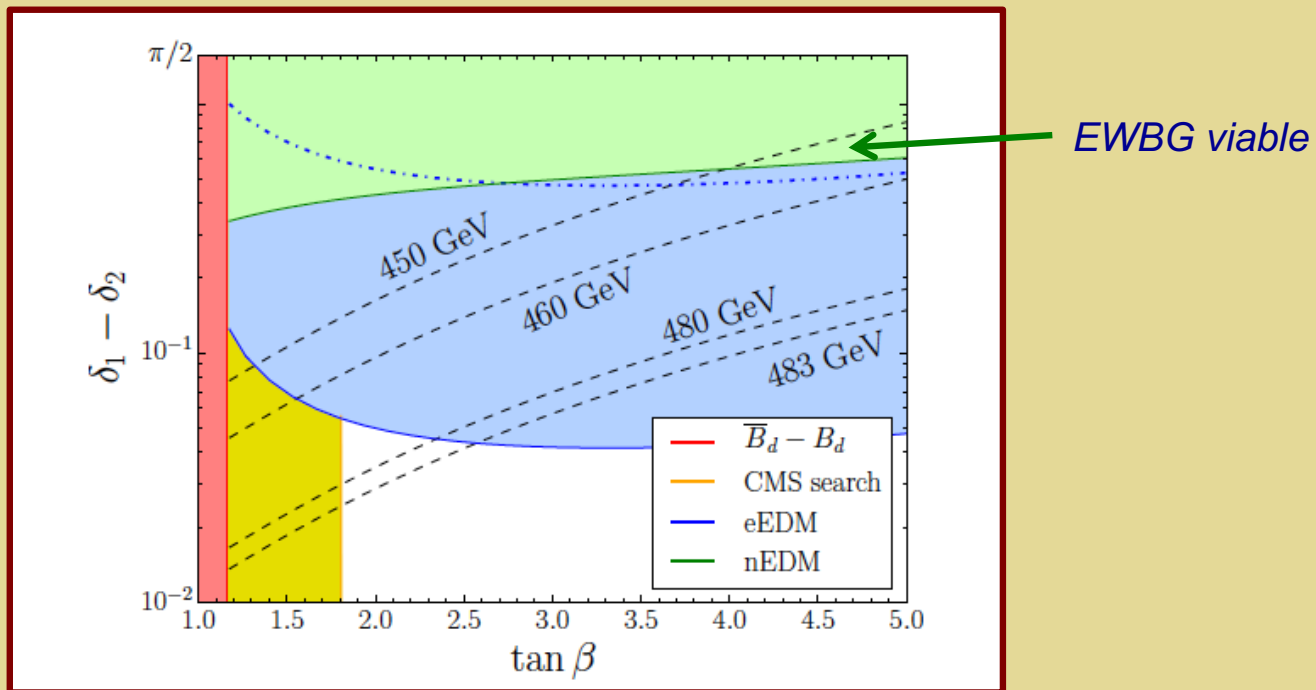
Future:

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- $d_A(Hg) \times 0.1$
- $d_{ThO} \times 0.1$
- $d_A(Ra)$

2HDM CPV & EWBG

2HDM CPV: Source for EWBG?

Dorsch et al, 1611.05874

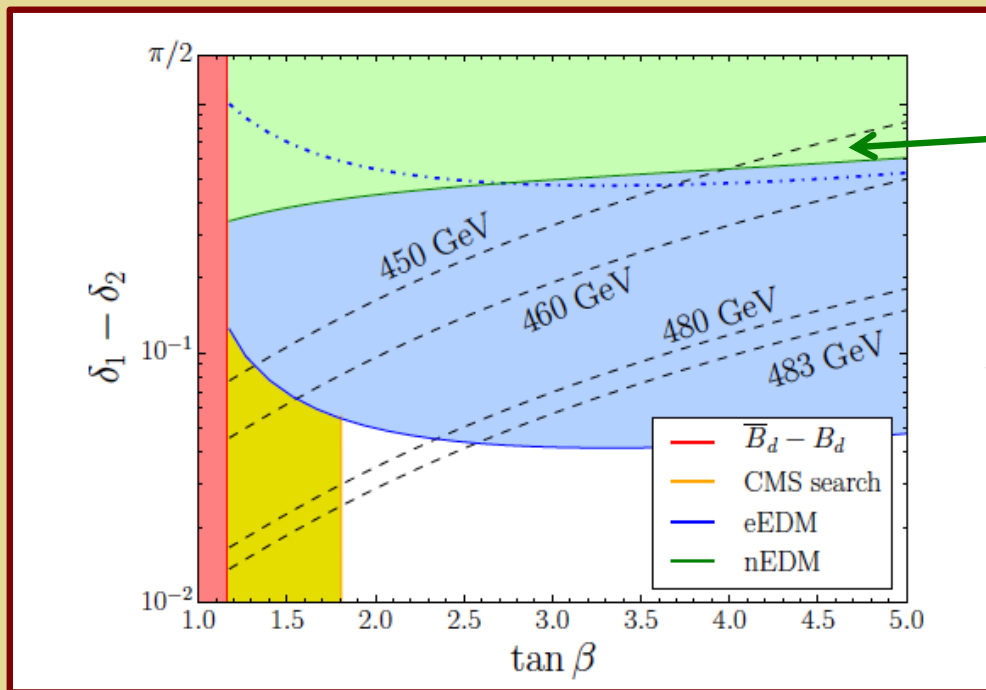


$$\alpha_b \propto \delta_1 - \delta_2$$

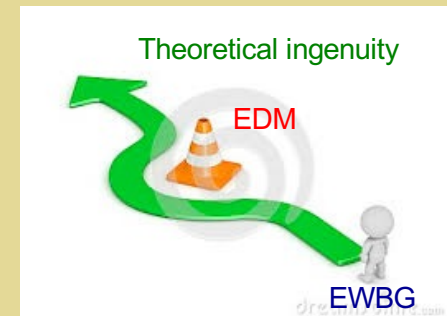
2HDM CPV & EWBG

2HDM CPV: Source for EWBG?

Dorsch et al, 1611.05874

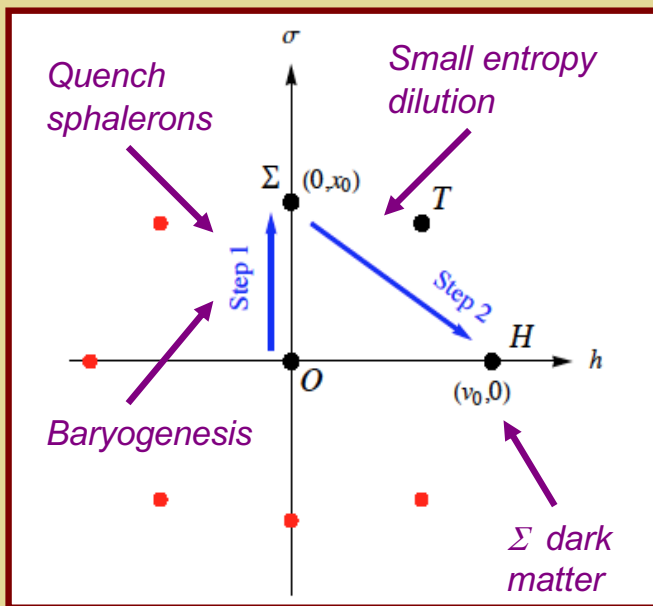
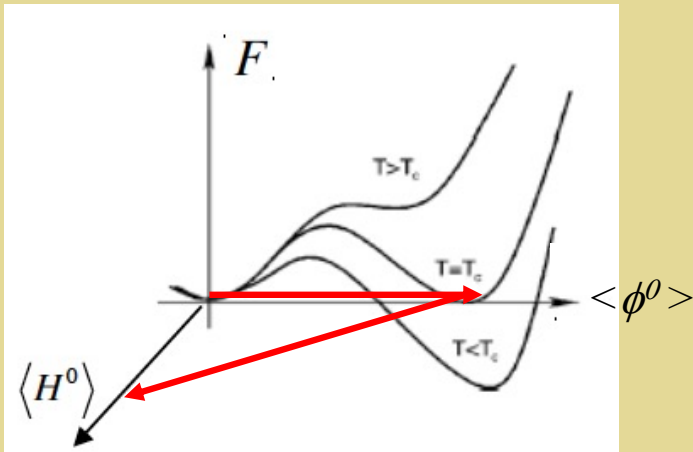


But new transport theory → see ahead

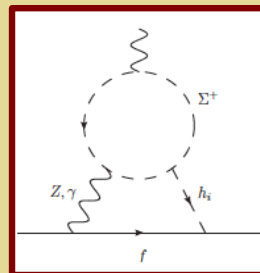
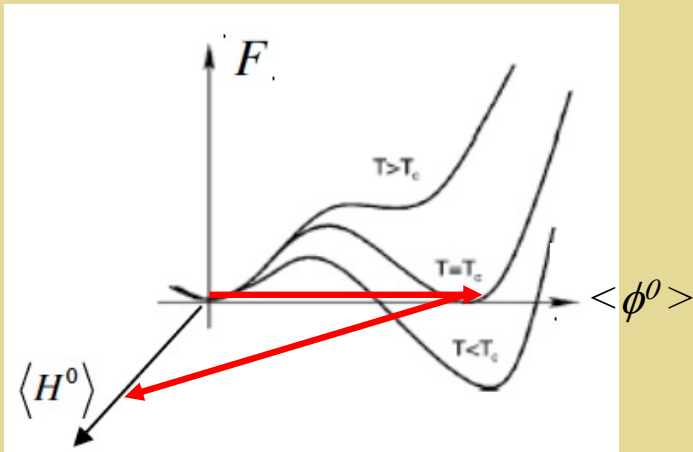


$$\alpha_b \propto \delta_1 - \delta_2$$

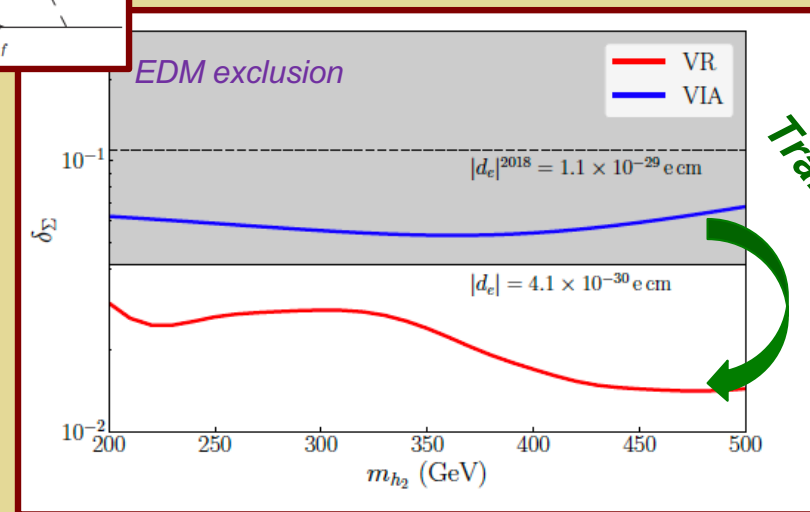
CPV for EWBG: Transport Theory & EDMs



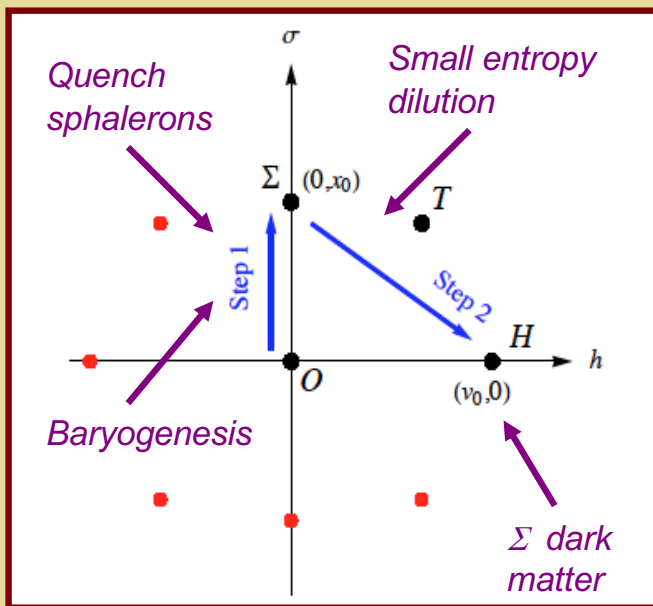
CPV for EWBG: Transport Theory & EDMs



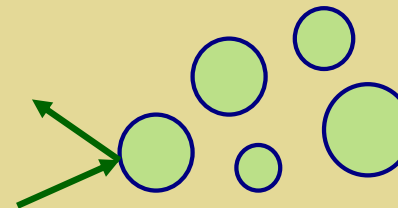
$$a_2 H_1^* H_2 \Sigma^2 + \text{c.c.}$$



Yuan-Zhen Li, MJRM, Jiang-Hao Yu 2404.19197

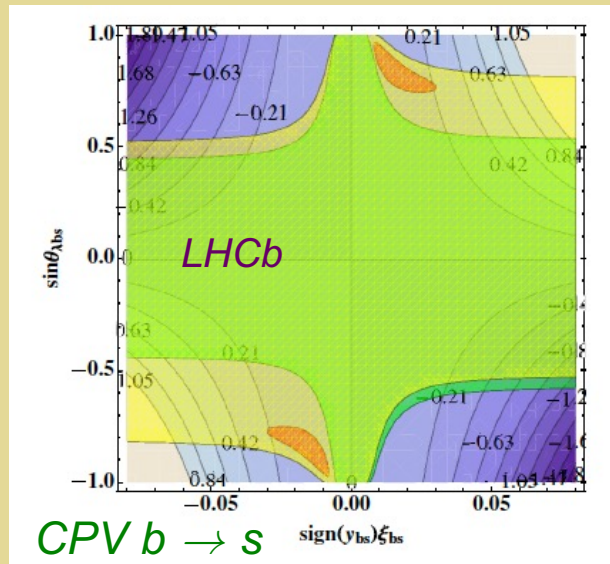
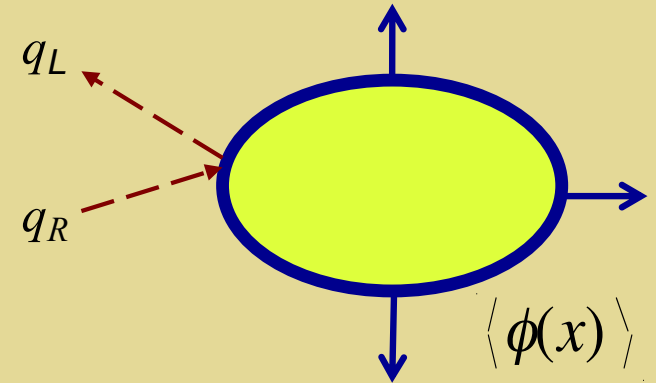


Inoue, Ovanesyan, R-M: 1508.05404



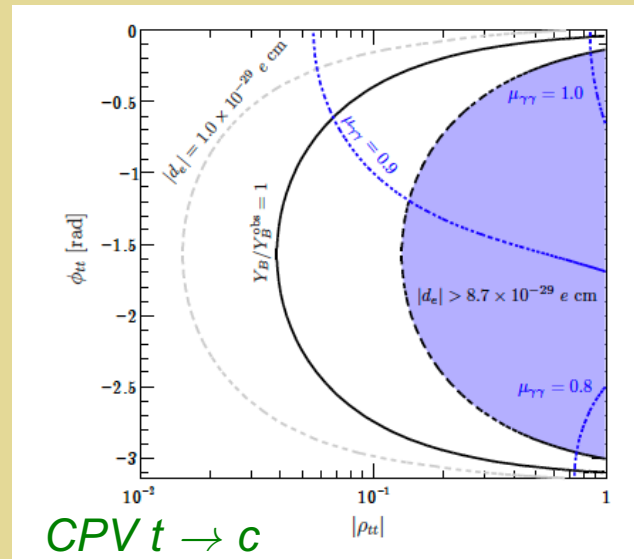
Y_B : CPV, diffusion & EW sphalerons

Flavored EW Baryogenesis



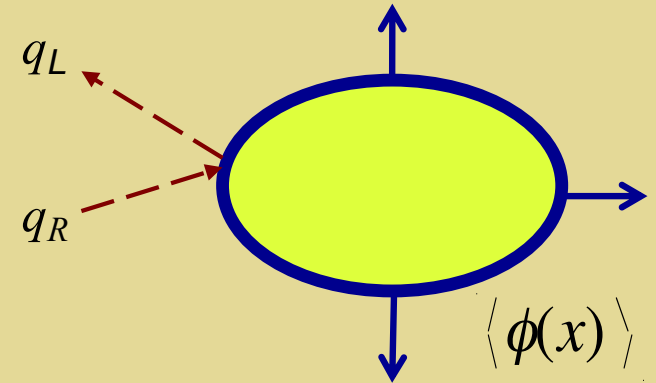
Liu, RM, Shu '12

$$\mathcal{L} = \lambda_{ij}^u \bar{Q}^i (\epsilon H_d^\dagger) u_R^j - \lambda_{ij}^d \bar{Q}^i H_d d_R^j - y_{ij}^u \bar{Q}^i H_u u_R^j + y_{ij}^d \bar{Q}^i (\epsilon H_u^\dagger) d_R^j + \text{H.c.}$$

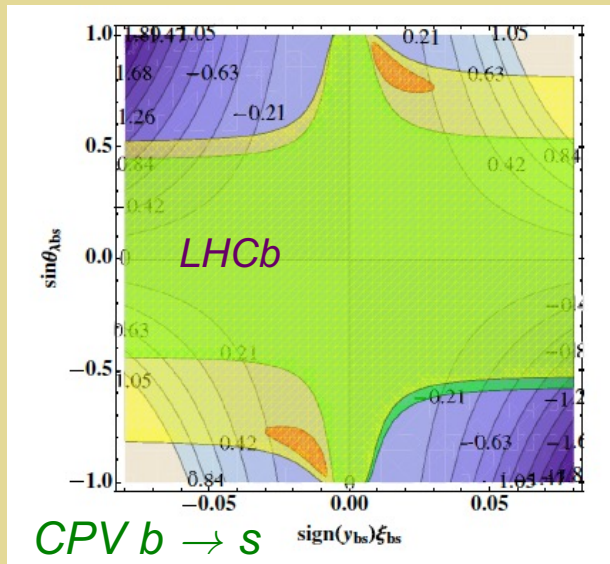


Fuyuto, Hou, Senaha '17

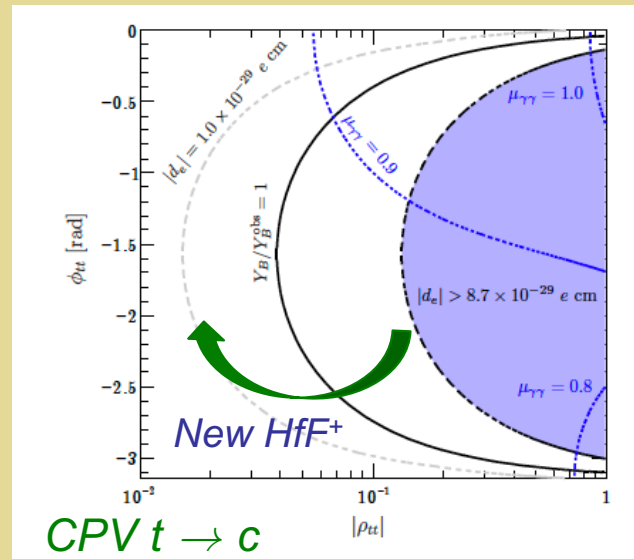
Flavored EW Baryogenesis



$$\mathcal{L} = \lambda_{ij}^u \bar{Q}^i (\epsilon H_d^\dagger) u_R^j - \lambda_{ij}^d \bar{Q}^i H_d d_R^j - y_{ij}^u \bar{Q}^i H_u u_R^j + y_{ij}^d \bar{Q}^i (\epsilon H_u^\dagger) d_R^j + \text{H.c.}$$

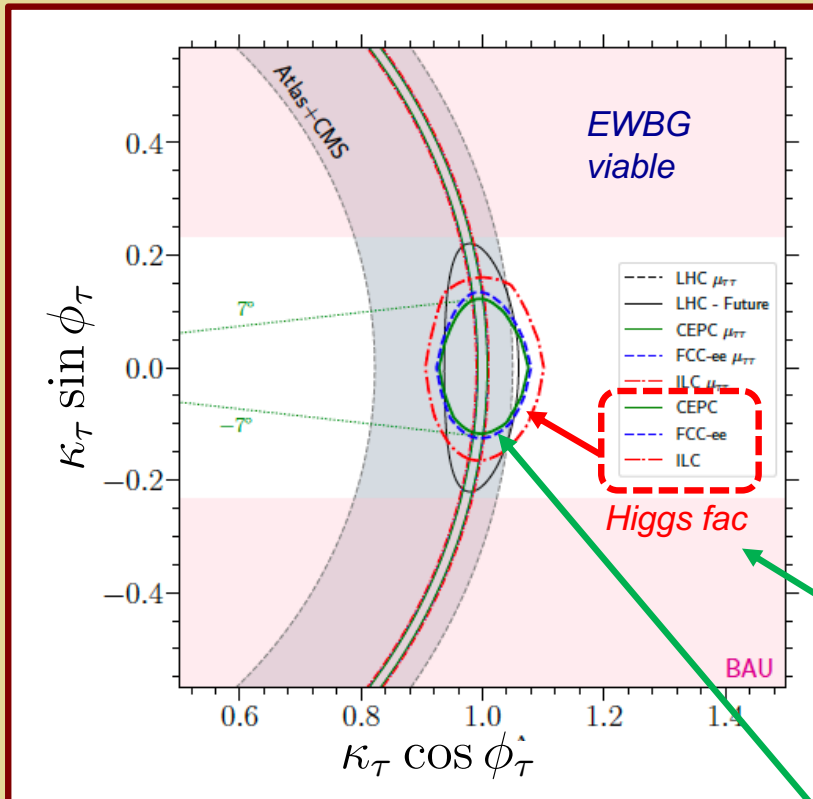
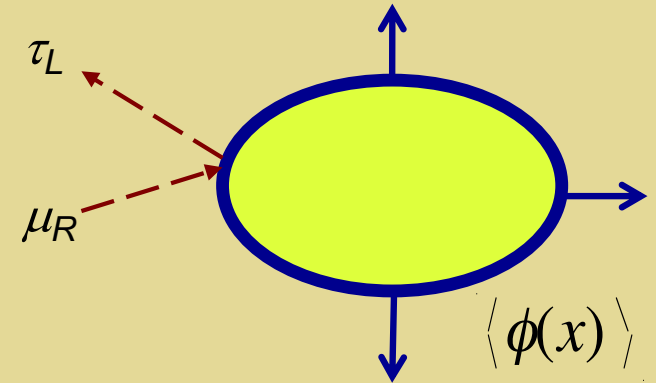


Liu, RM, Shu '12



Fuyuto, Hou, Senaha '17

Flavored EW Baryogenesis



Flavor basis (high T)

$$\mathcal{L}_{\text{Yukawa}}^{\text{Lepton}} = -\overline{E}_L^i [(Y_1^E)_{ij} \Phi_1 + (Y_2^E)_{ij} \Phi_2] e_R^j + h.c.$$

Mass basis ($T=0$)

$$\frac{m_f}{v} \kappa_\tau (\cos \phi_\tau \bar{\tau} \tau + \boxed{\sin \phi_\tau \bar{\tau} i \gamma_5 \tau}) h$$

$CPV h \rightarrow \tau\tau$

Guo, Li, Liu, R-M, Shu 1609.09849

Ge, Li, Pasquini, R-M, Shu 2021.13922

Two Themes

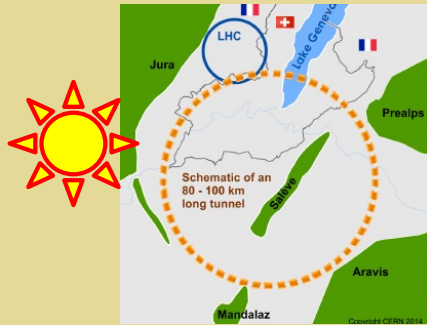
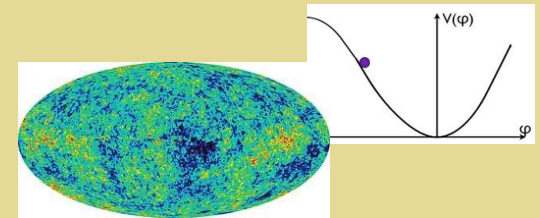
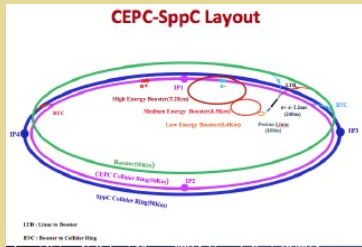


- *Progress will come from focusing on well-posed scientific questions → potential for insights and/or discoveries*

Was the matter-antimatter asymmetry generated in conjunction with electroweak symmetry-breaking ?

- *It's important to think beyond boundaries of funding agency priorities and conventional sub-field categories → rich opportunities from inter-frontier connections*

Frontiers



Linear Colliders
ILC & CLIC species

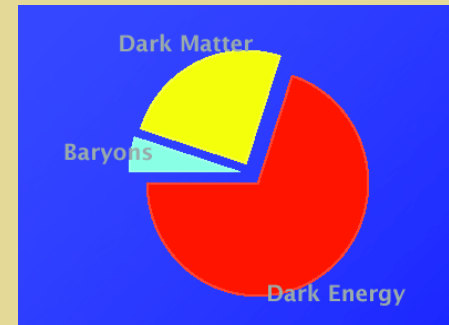
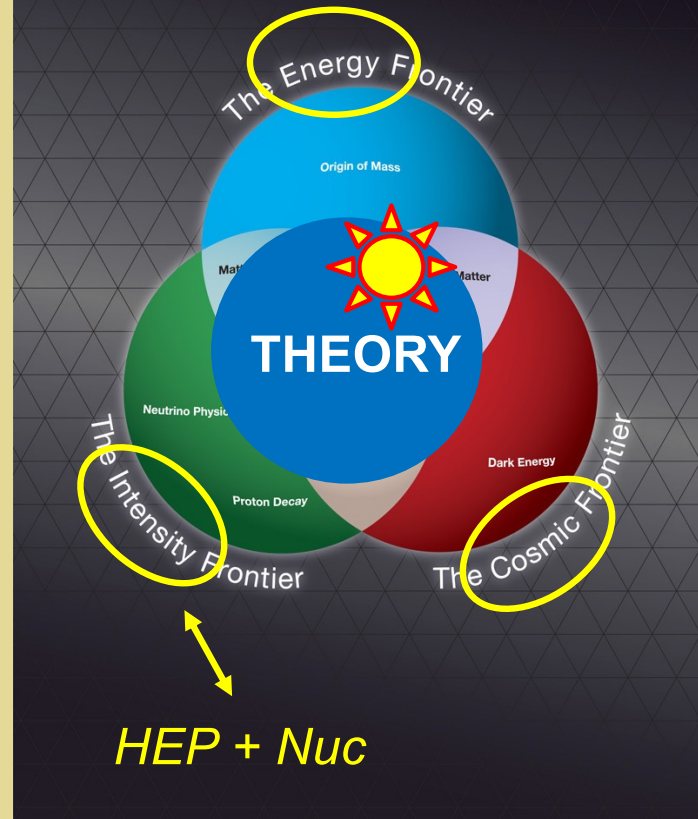
- Energy extendability to TeV scale lies in the heart of linear colliders: ILC focuses on \sqrt{s} from 250 GeV to 1 TeV; CLIC 380 GeV to 3 TeV, keeping options to run at Z-pole ("GigaZ")
- Complementary approaches: "Warm" & "Cold" accelerating technologies; 72MeV/m @ CLIC380; 51.5MeV/m @ ILC250
- Polarized beams: both offering 80% for electron; 30% for positron in ILC default design

ILC250 - 20km

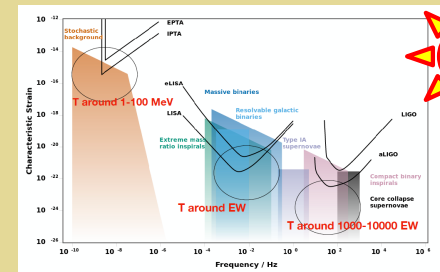
東京大学 THE UNIVERSITY OF TOKYO

Lifetime Frontier

HEP : New (heavy) particles



- Precision tests: muon $g-2$, PV $ee...$
- Fundamental symmetry tests (CP, Lepton number...)
- Neutrino properties
- Flavor physics



- Atomic, Molecular, Optical
- Condensed Matter

Historical artifact: US HEP vision → still useful mnemonic

III. Outlook

Fundamental Physics: Past & Future

