



STRUCTURES
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Demography of planetary systems

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Content

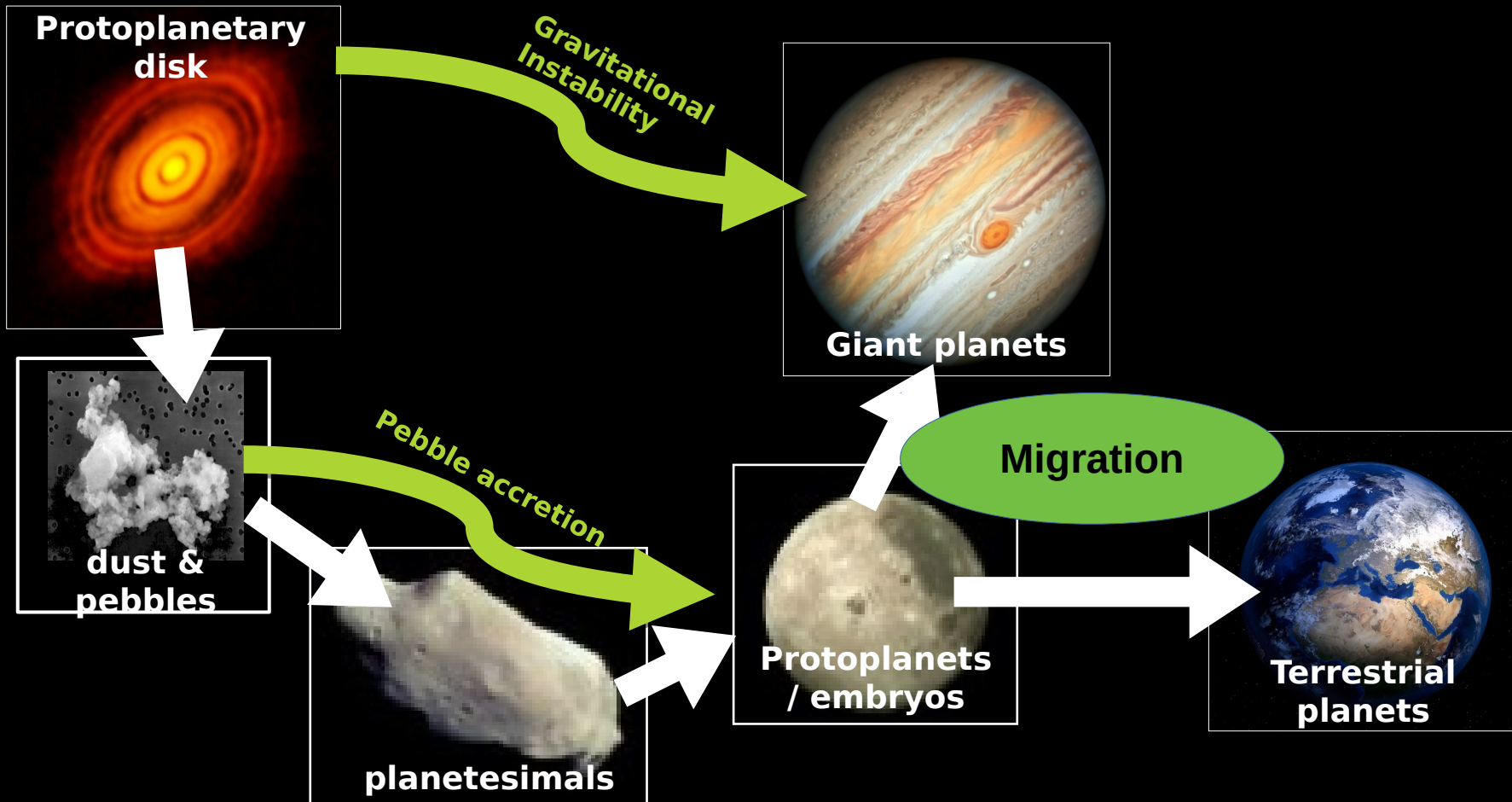


STRUCTURES
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- **Planet Formation**
- **Exoplanet Demographics**
- **Planetary population synthesis**
- **Exoplanetary radius valley**



Planet Formation Overview

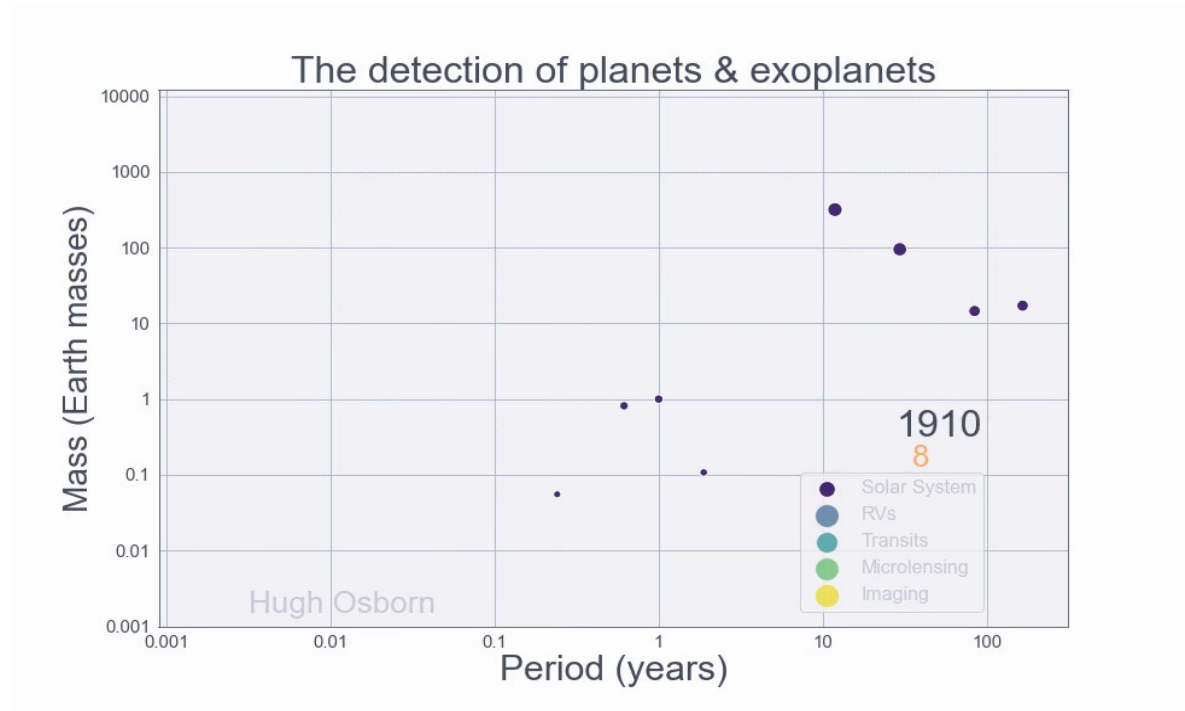


Exoplanet Demographics



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- **Thousands of discoveries**
- **Enables Statistical studies**



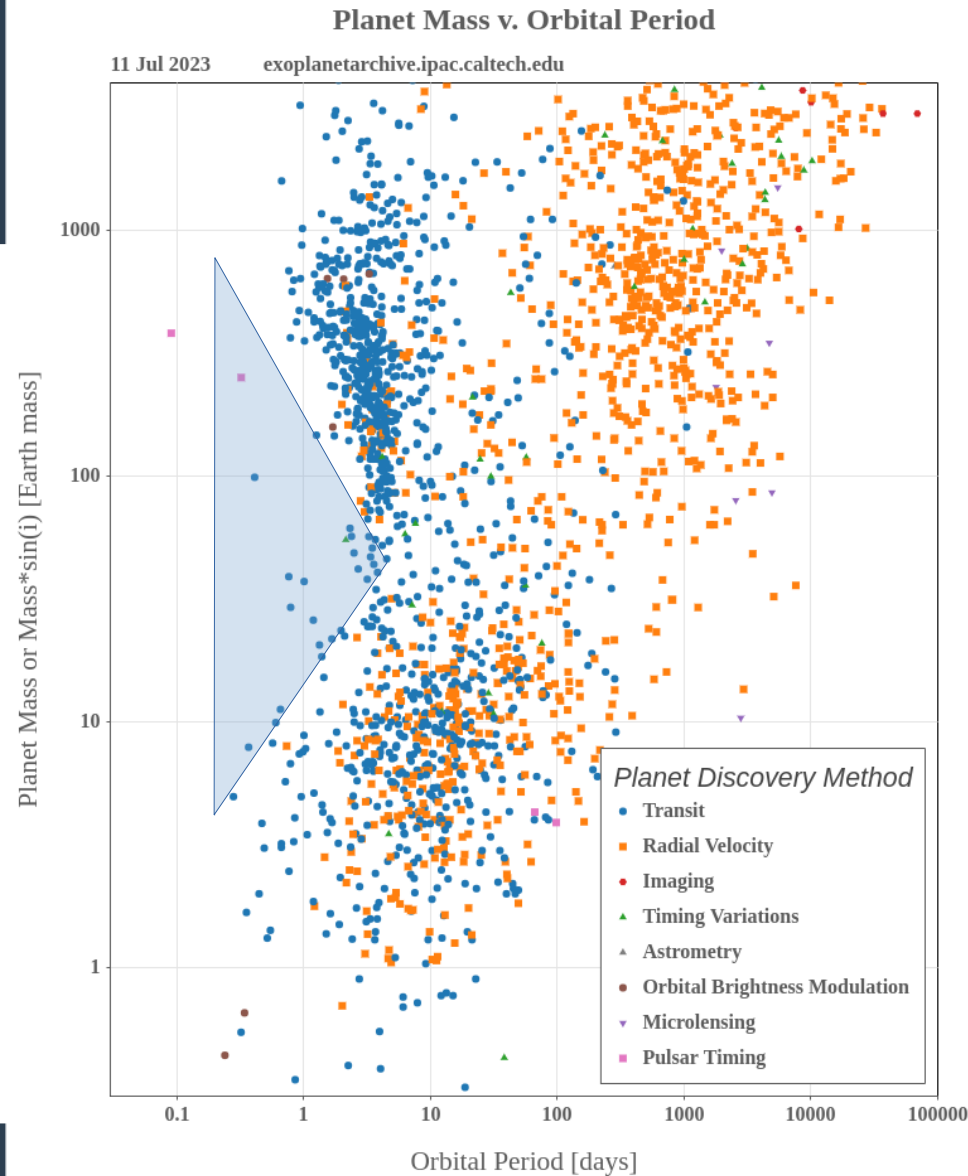
Exoplanet Demographics

- **Known and disputed demographic features**



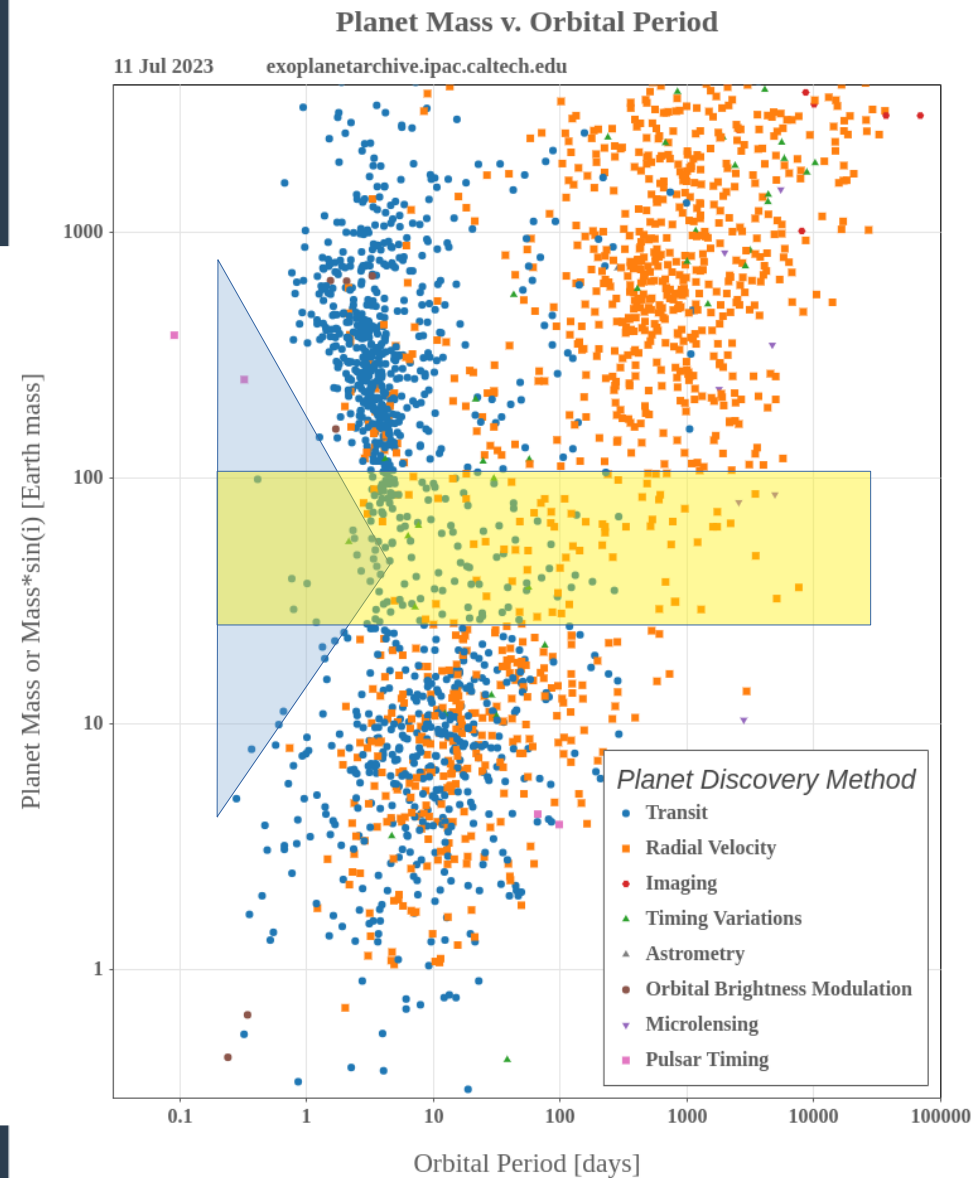
Exoplanet Demographics

- **Known and disputed demographic features**
 - Sub-Jovian deserts
 - Neptunian desert / hot-Neptune desert (Mazeh+ 2005/2016)
 - Most likely shaped by photoevaporation (Owen&Lai 2018)



Exoplanet Demographics

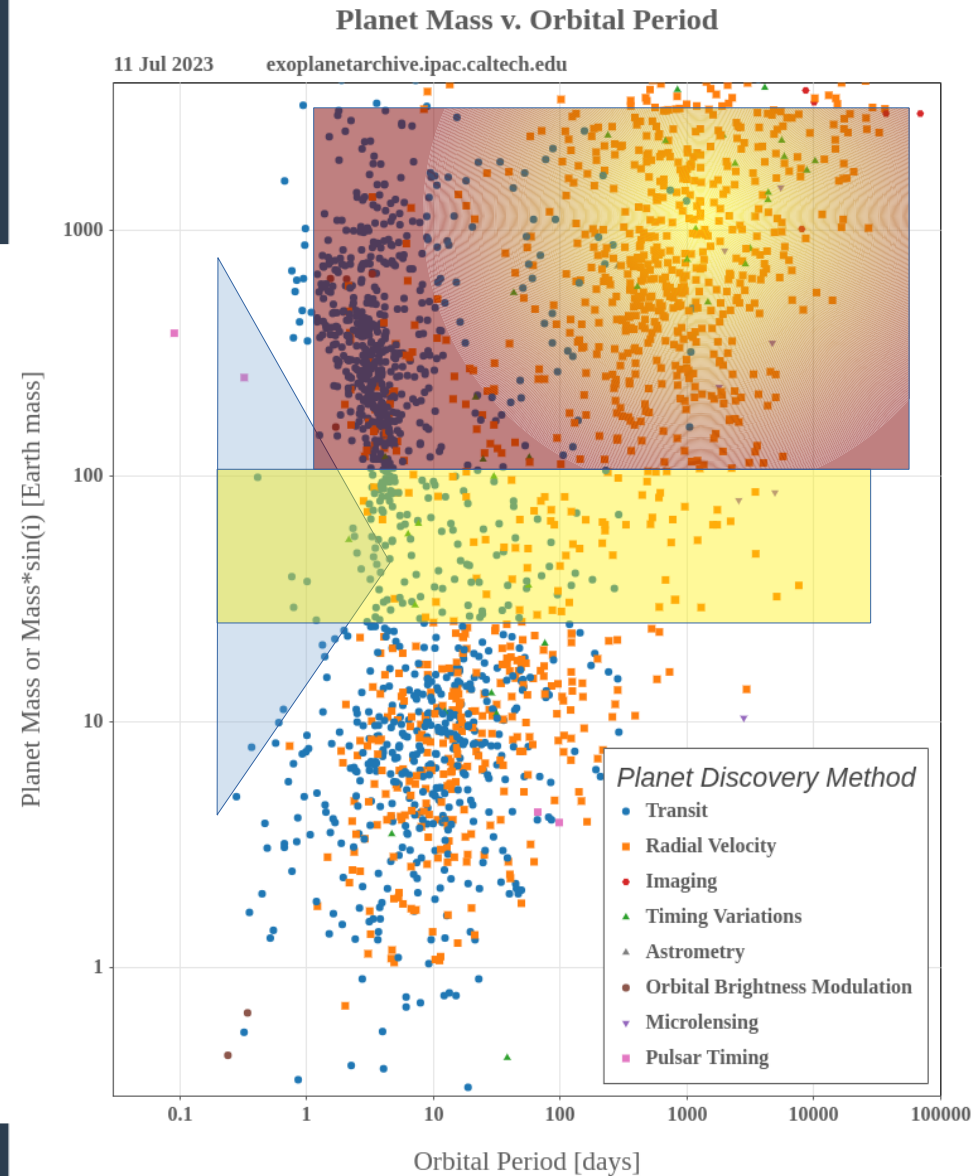
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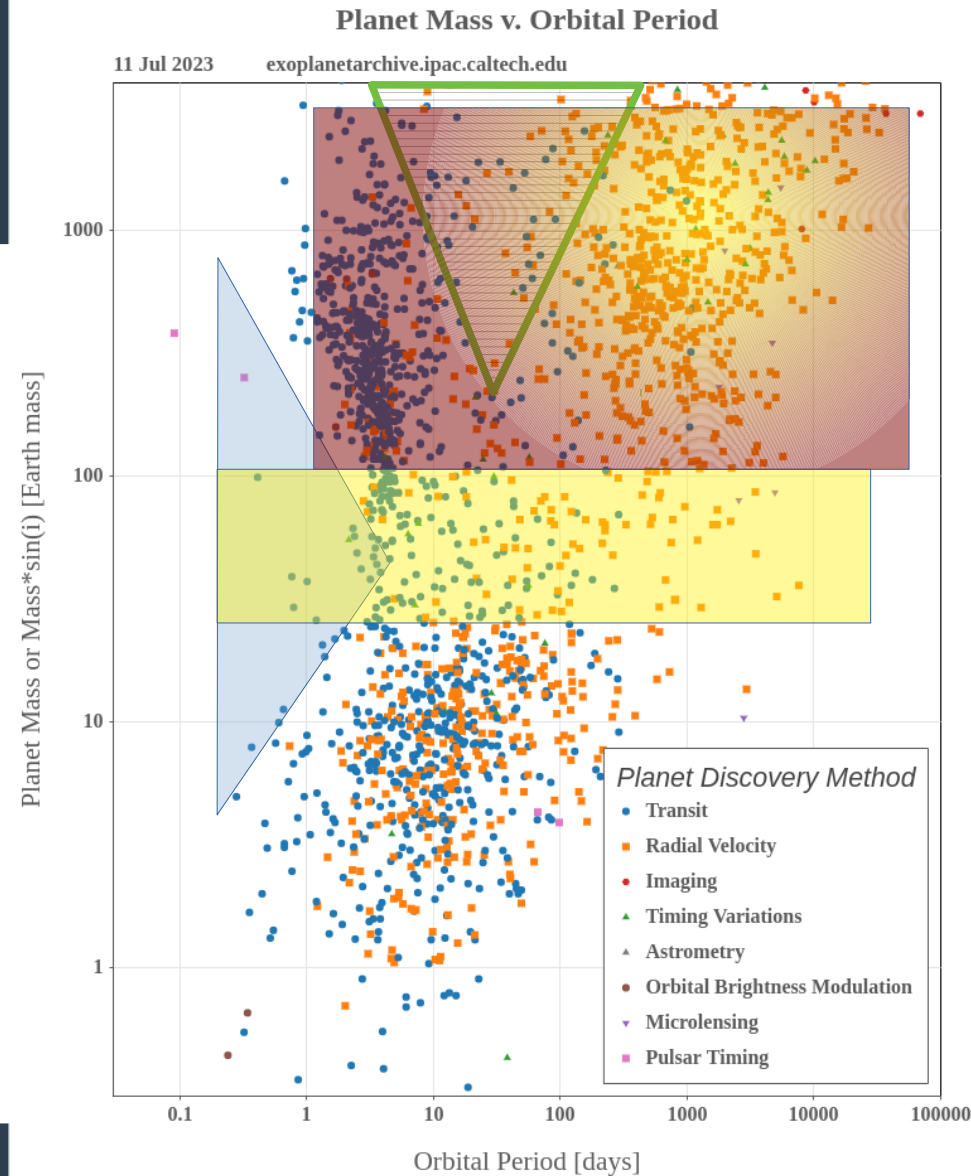
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- Hot Jupiter/Cold Jupiters (Mayor&Queloz 1995)
 - Turnover at iceline (Suzuki+ 2016, Pascucci+ 2018, Fernandes+ 2019)



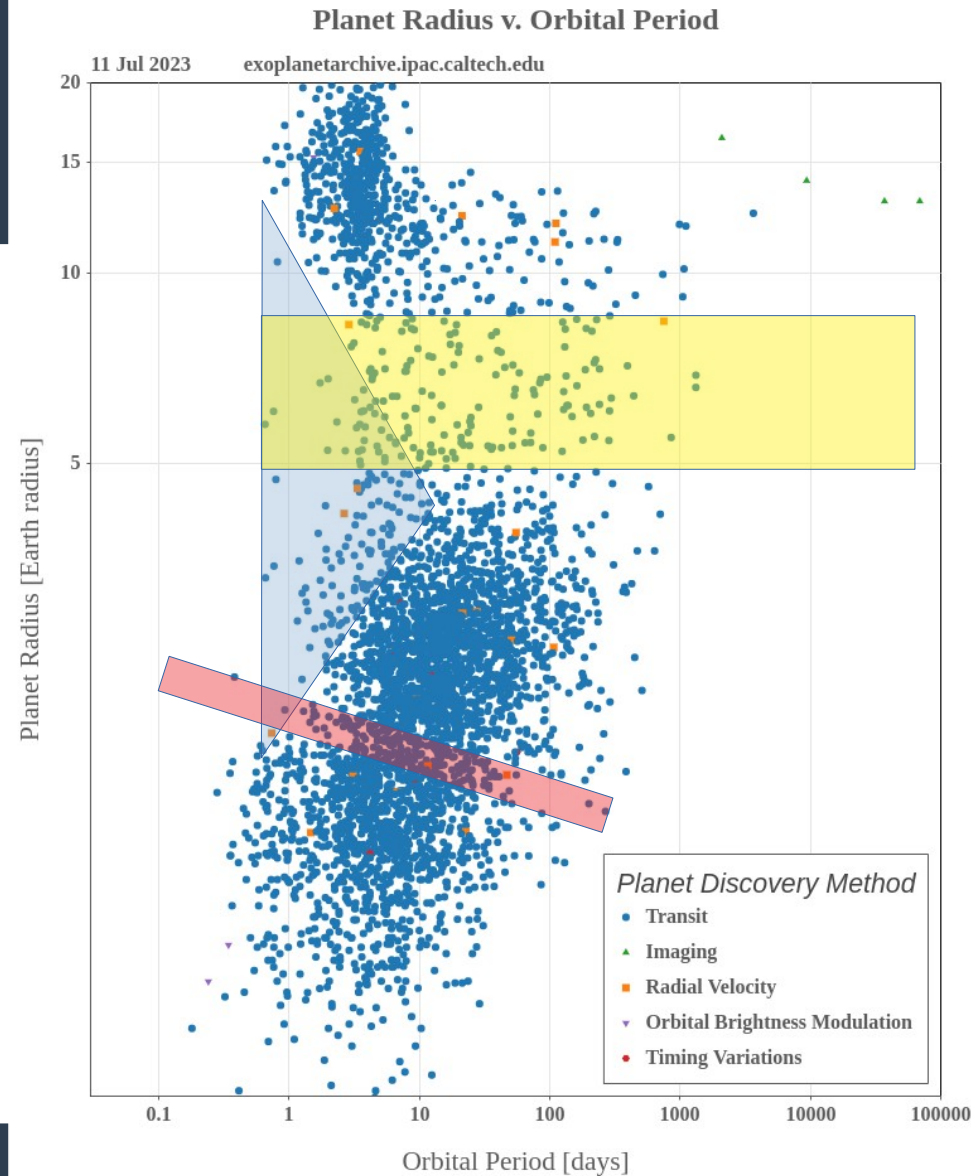
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 - 'Apparent separation of hot and cold Jupiters'
 - Radius Valley (Fulton+2017)
 - Slope with stellar mass & orbital period/irradiation?



Exoplanet Demographics



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- **Trends**

- with stellar $[Fe/H]$
 - Increasing with giant planet incidence (e.g. Santos+2001)
 - Relatively flat for Earth-type planets (Buchhave+2013)
 - Eccentricity of giant planets increases (Dawson+2013)



Exoplanet Demographics



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- **Trends**

- with stellar $[Fe/H]$
 - Increasing with giant planet incidence (e.g. Santos+2001)
 - Relatively flat for Earth-type planets (Buchhave+2013)
 - Eccentricity of giant planets increases (Dawson+2013)
- with stellar mass
 - Increasing giant planet occurrence (Endl+2006, Butler+2006, Gaidos+2013, Ghezzi+2018, Jordán et al. 2022)
 - But exceptional giants around late M dwarfs exist (Sabotta+2021, Bryant+2023)



Exoplanet Demographics



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- **Characterization stage:
Compositional statistics**
 - Over-density of water-worlds
(Luque&Palle 2022)

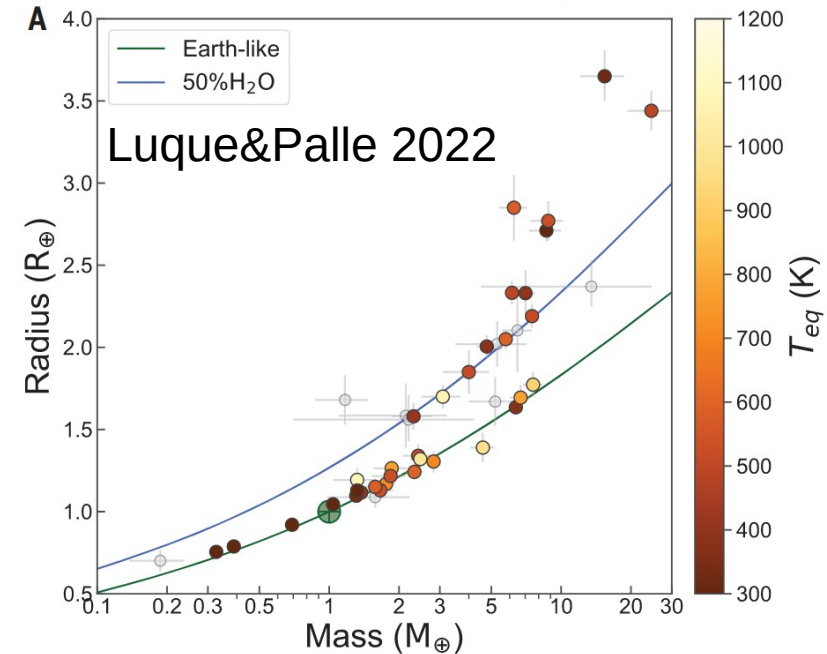


Exoplanet Demographics

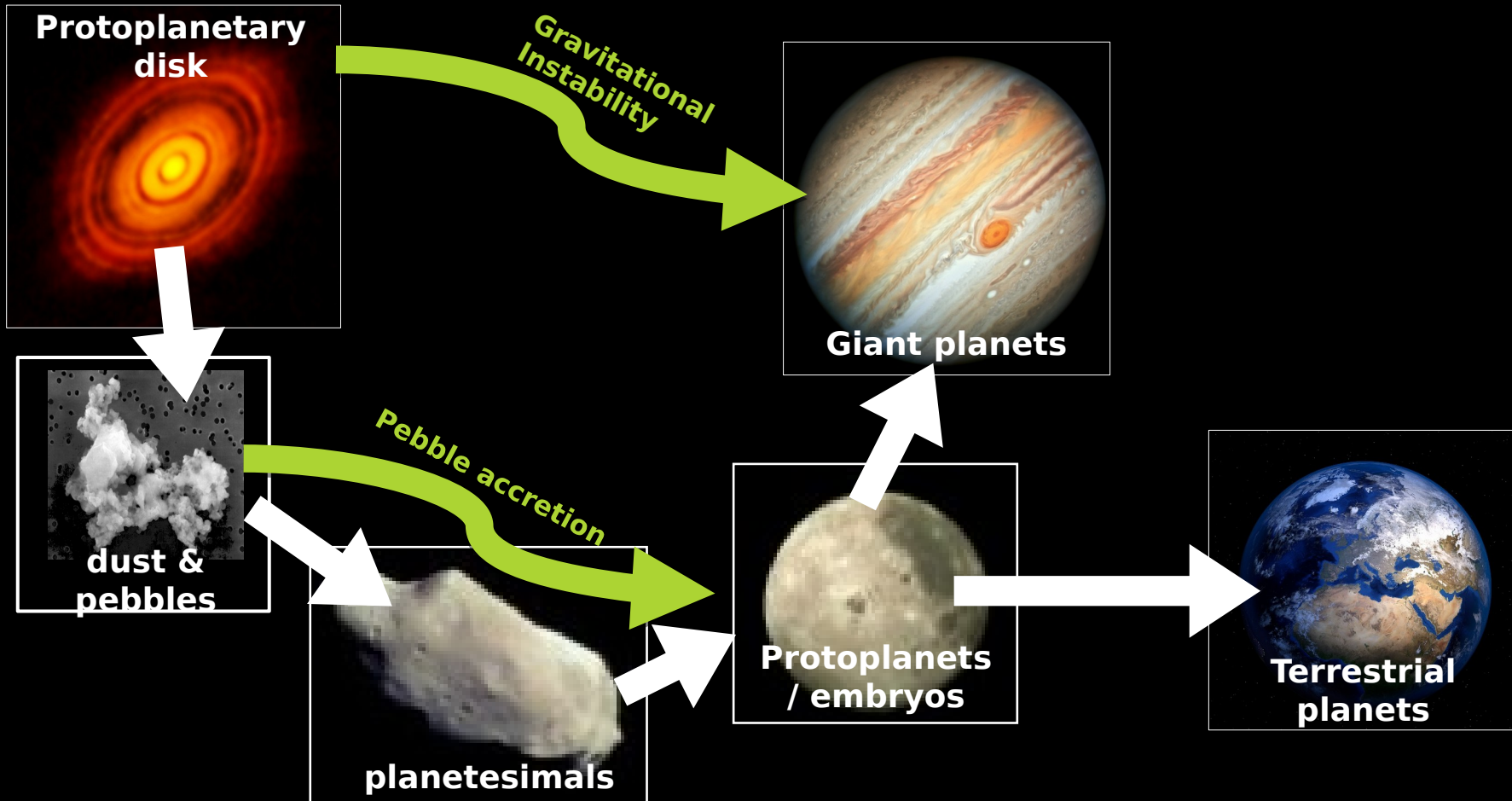


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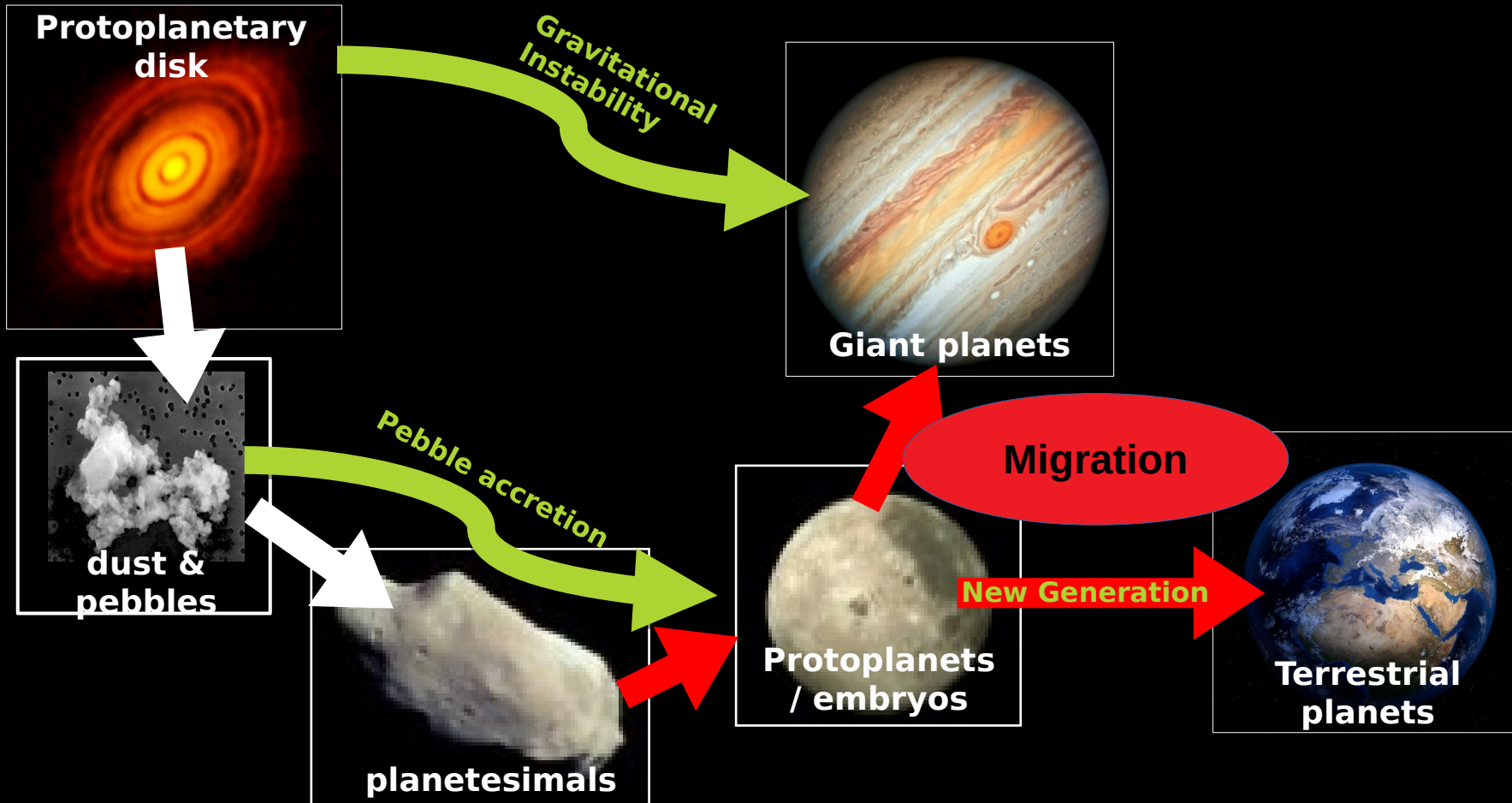
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Planet Formation Overview



Planet Formation Overview



Planetary Population Synthesis



STRUCTURES
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- **Connects observed disks with observed planets via formation models**
- **New generation of simulations**
 - Emsenhuber+ 2021 (a,b), Schlecker+ 2021 (a,b), Burn+ 2021
 - Start with already formed, small planetesimals and moon-sized embryos in smooth disks
 - Single stars
 - Global models including evolution of planets



Planetary Population Synthesis



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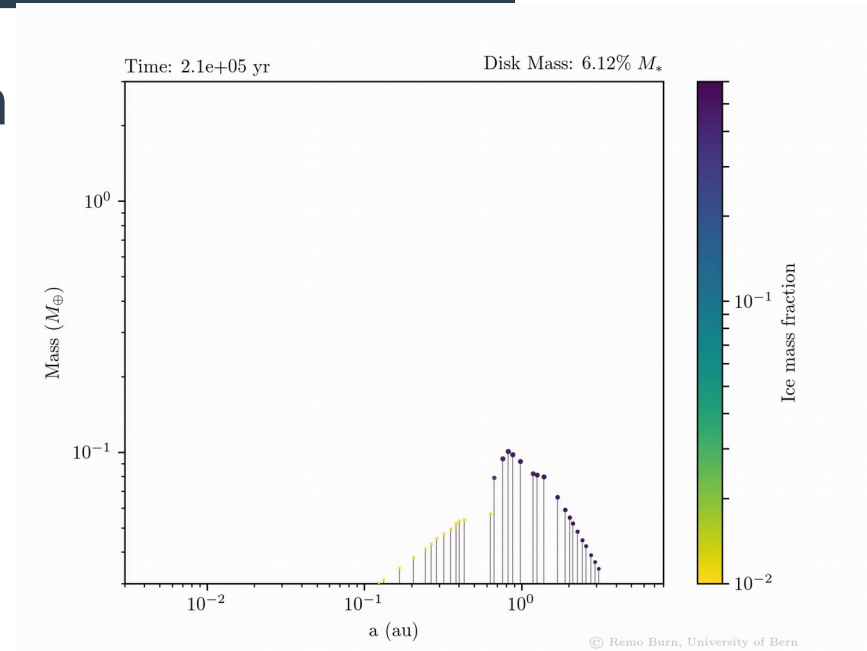


Planetary Population Synthesis



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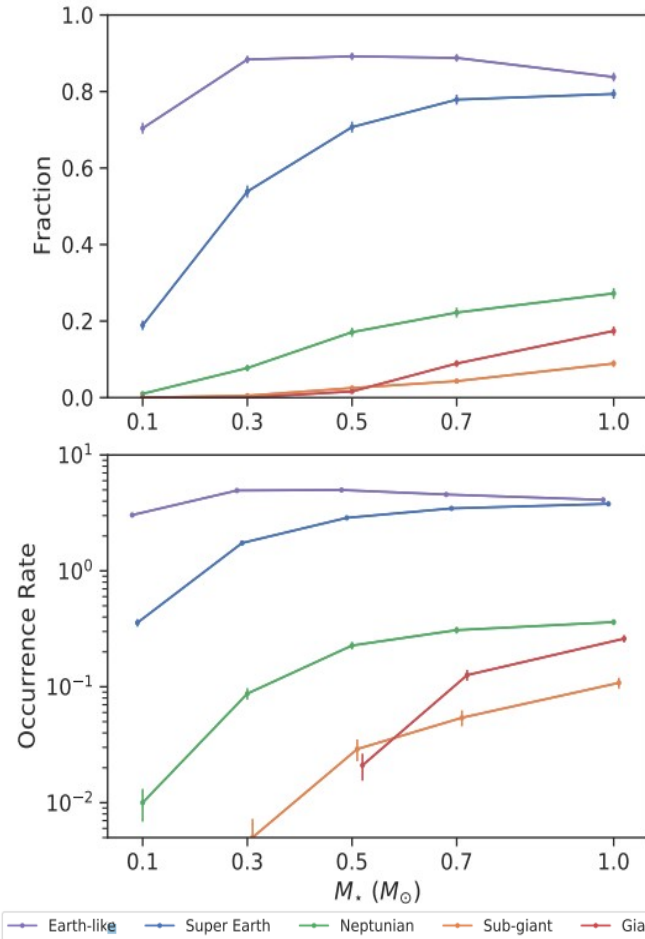
Planetary Population Synthesis



STRUCTURES
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- **Trends with stellar mass**
 - Qualitative agreement with observations
 - Tens of percent chance for each star to host a planet in the habitable zone

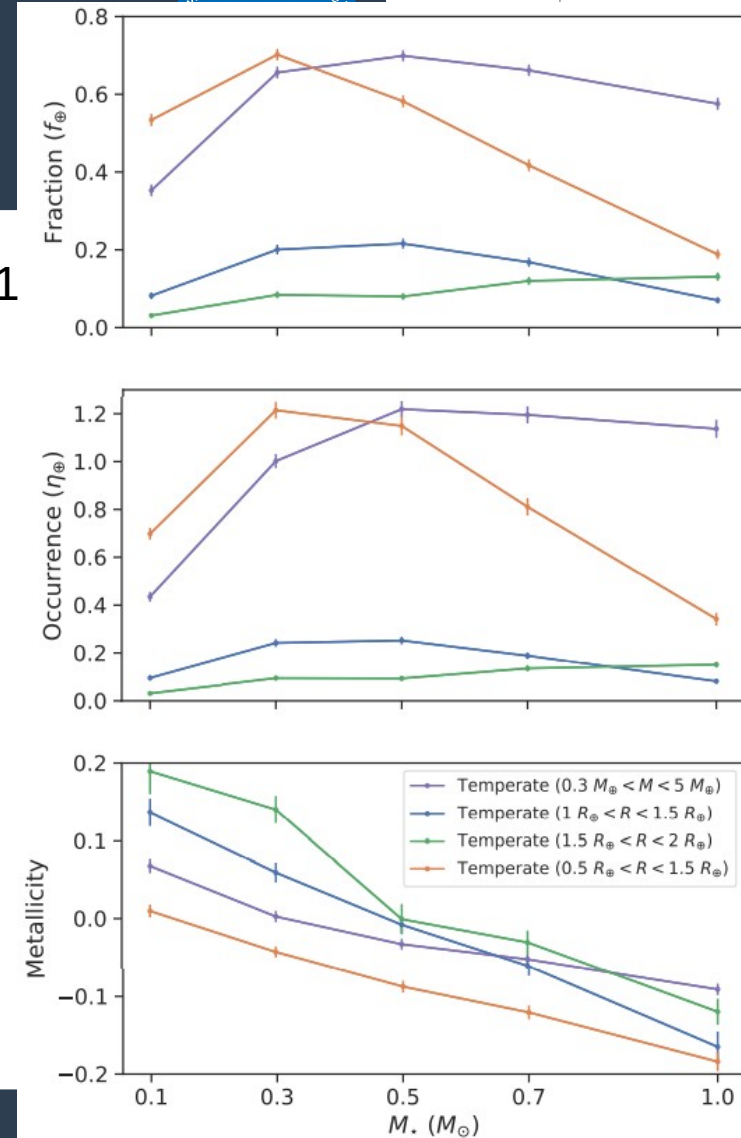
Burn+ 2021



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Burn+ 2021

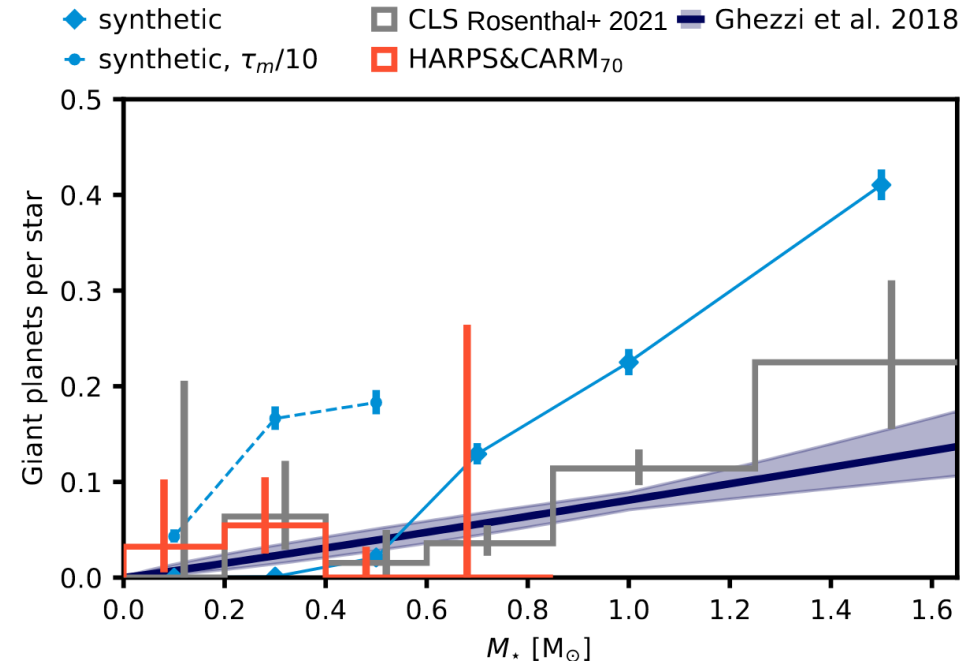


Planetary Population Synthesis



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- **Giant Planets exist around low-mass stars**
 - Not predicted by standard models
 - zero giants for stellar Masses < 0.3 solar masses
 - Not necessarily a mass problem \rightarrow Reduced/inhibited migration allows for some giant planet formation

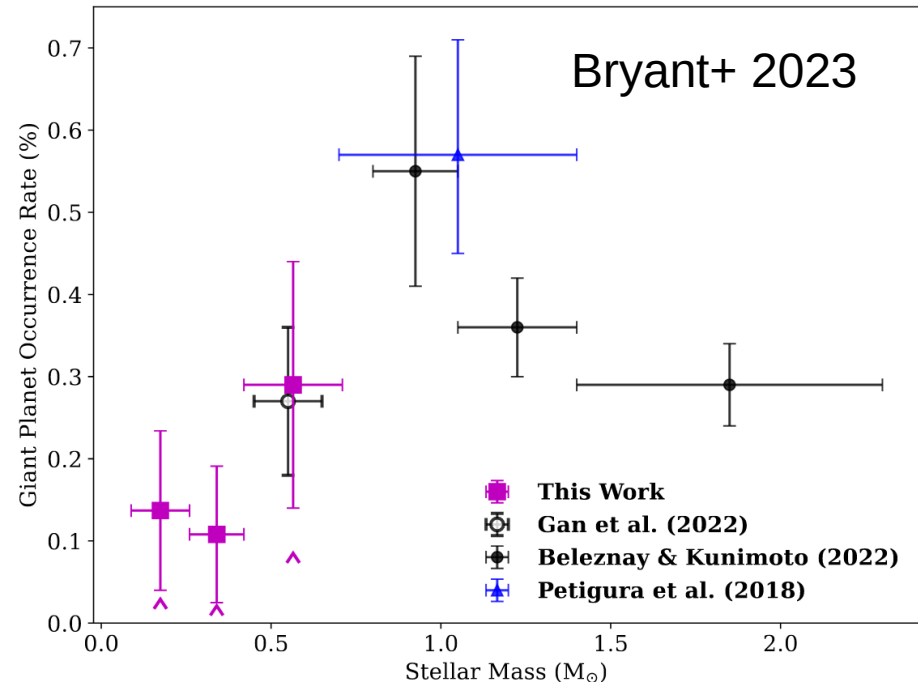


Planetary Population Synthesis



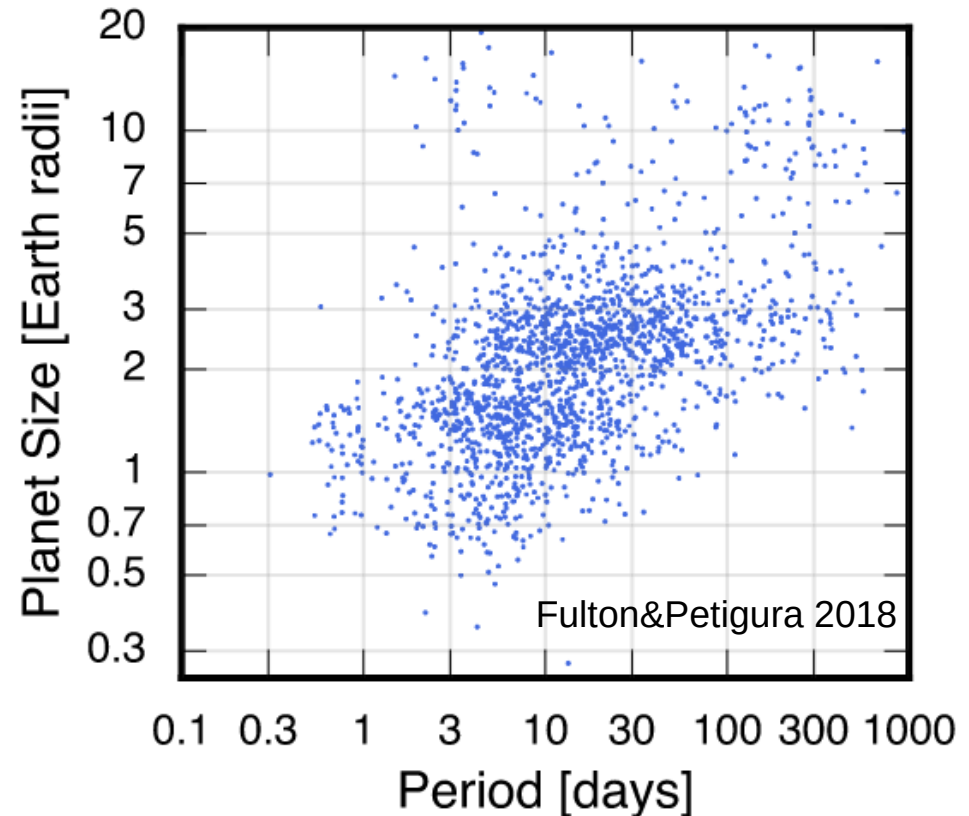
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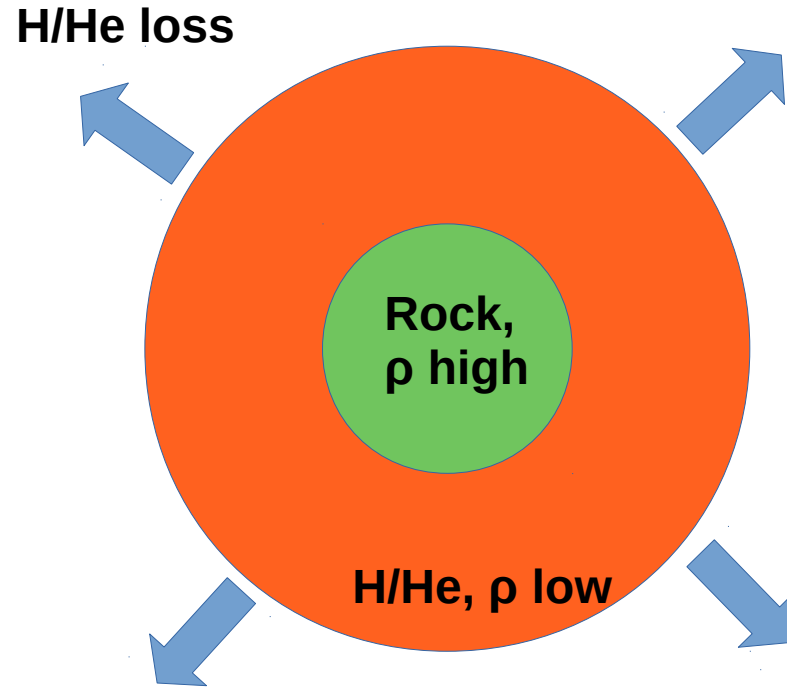
Radius Valley

- **Observed radius valley (Fulton+2017,2018)**
- **Observational constraints**
 - Inconclusive direct evidence He absorption (Kasper+2020, Zhang+2022), H₂O detection (Benneke+2019)
 - Indirect evidence for water worlds (Diamond-Lowe+2022, Luque&Palle 2022, for M stars)
- **Mass loss process can shape the valley (core- or XEUV-powered, e.g. Owen 2019 for a review)**



Radius Valley

Possible sub-Neptune



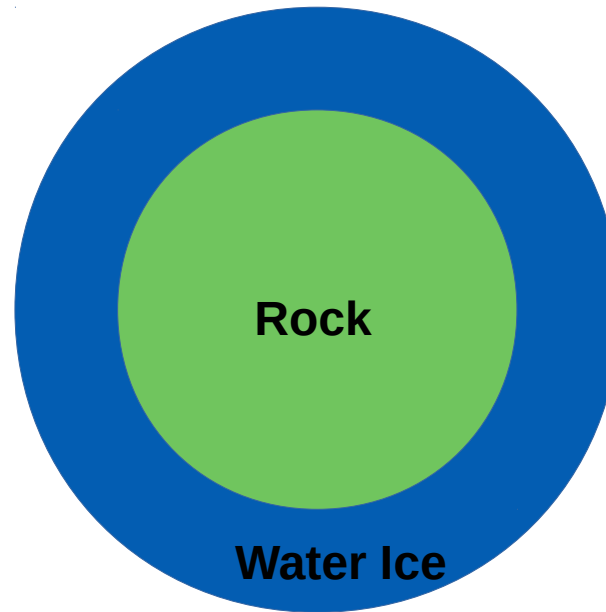
Orbital Migration



- **Indications for orbital migration of planets**
 - (near-)resonant systems
 - super-Earth formation
 - theoretical necessity
- **Leads to water world formation**



Possible sub-Neptune?

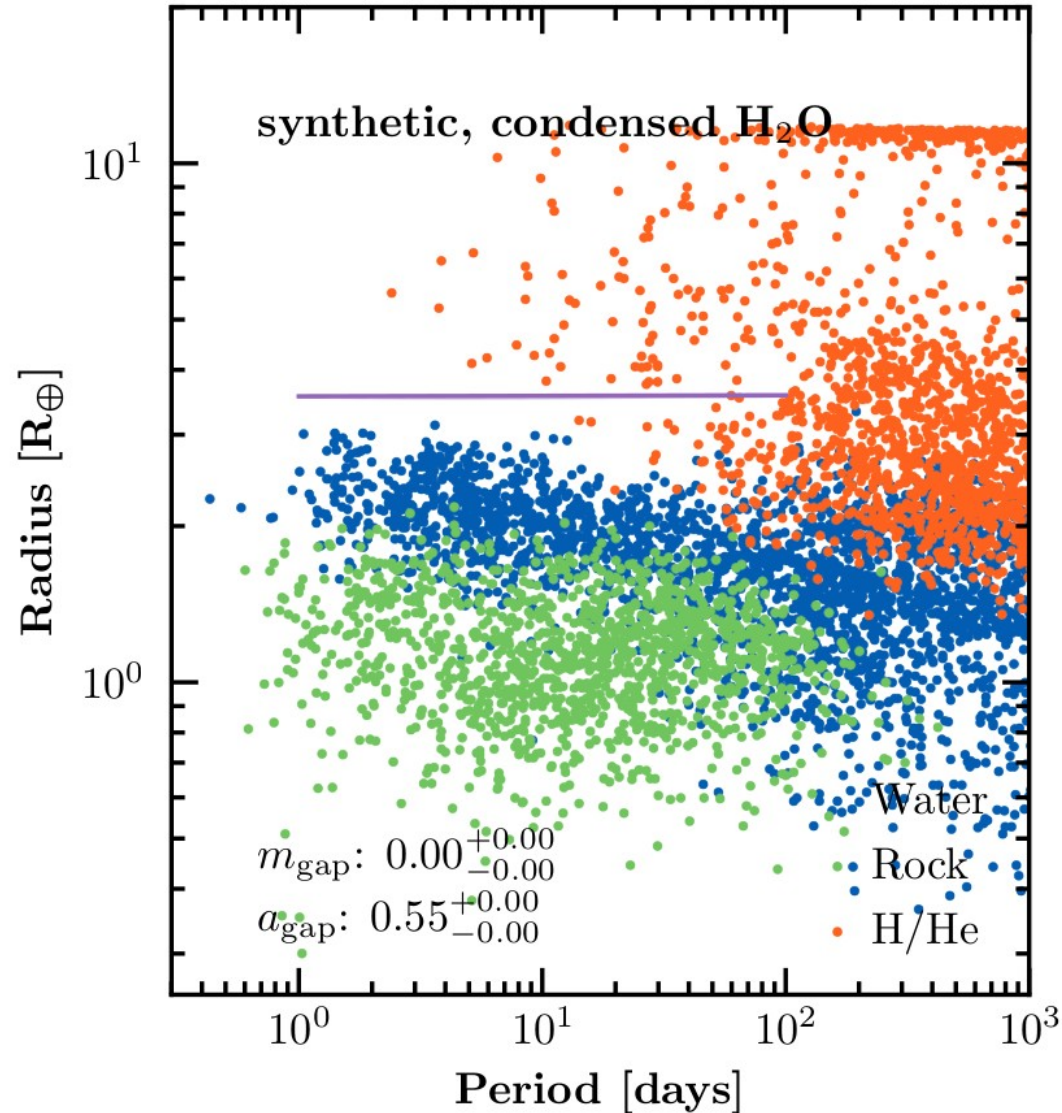


- **Solar-type star**
- **Formation (Emsenhuber+2021)**
 - planetesimal-based solid accretion
 - N-body interactions ($N_{\text{ini}}=100$, calculated for 100 Myr)
 - Giant impacts
 - Impact stripping
 - Orbital migration
 - Consistent gas accretion with 1D internal structure calculation
- **Evolution**
 - Cooling & contraction with initial stage given by formation
 - **Atmospheric escape of H/He mixture (Kubyskhina & Fossati 2021)**
 - Stellar evolution in X-ray and bolometric luminosity
 - **Interior structure with rock and condensed ice (Seager+2007)**
 - Tidal orbital decay
- **Apply observational Kepler bias**
 - KOBE, Mishra+2021

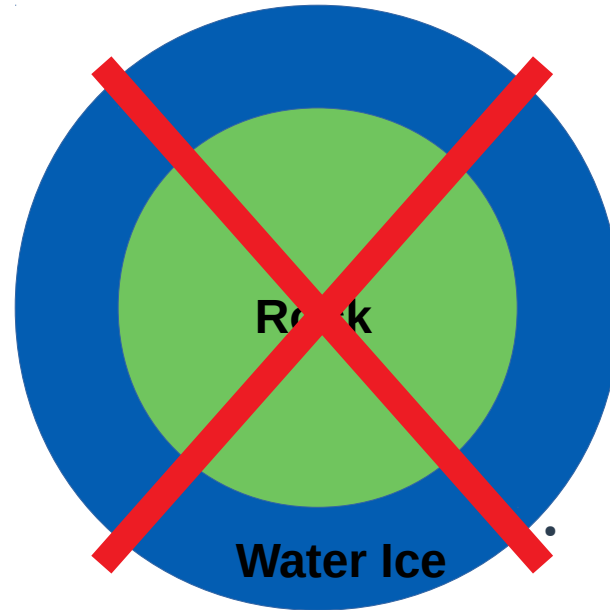


Results

- Condensed H₂O
- Valley at wrong location (3-4 R_e)
- Depends on mass distribution



Impossible sub-Neptune



(see also Jin&Mordasini 2018)

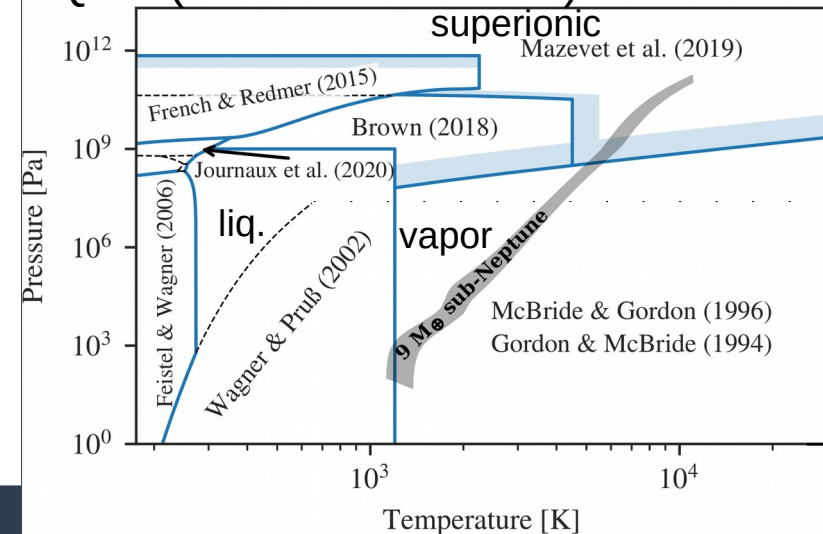


Phases of water

- **Expect vapor interior to runaway greenhouse limit (e.g. Boukrouche 2021)**
- **Water in close-in sub-Neptune is in supercritical vapor and superionic fluid state**

$$a_{\text{runaway}} = \sqrt{\frac{(1 - \alpha_{\text{al}})L_{\star}}{16\pi I_{\text{OLR}}}}$$
$$I_{\text{OLR}} = 281 \text{ W m}^{-2}$$

AQUA (Haldemann+2020)

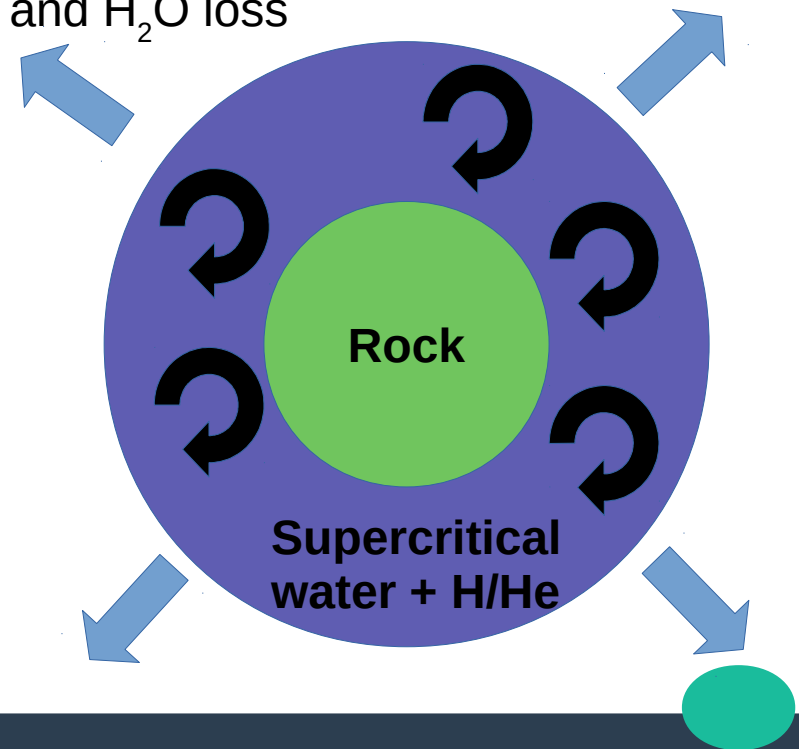


Radius Valley

- **Water should be in steam atmospheres (Turbet+2019)**
 - Could also explain the valley (Zeng+2019, Mousis+2020, Venturini+2020)
 - (Partial) Mixing with H/He is expected (Pierrehumbert 2022)
- **Can a uni-modal planet mass distribution reproduce the observed valley if there are steam atmospheres?**

Possible sub-Neptune

H/He and H₂O loss

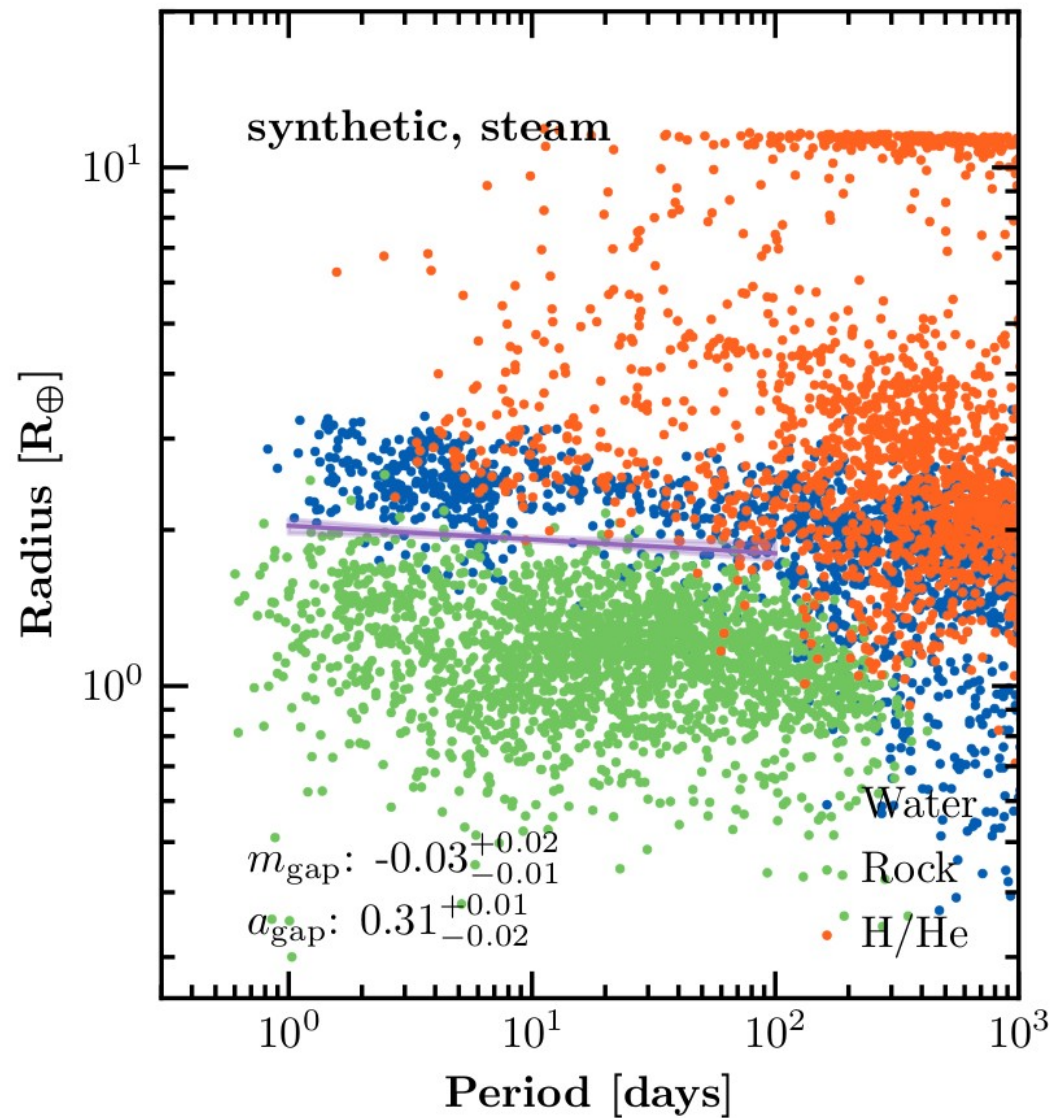


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 - Cooling & contraction with initial stage given by formation
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 - **H/He: Kubyshkina & Fossati 2021**
 - **Water: Johnstone 2020**
 - Stellar evolution in X-ray and bolometric luminosity
 - **Interior structure with new equation of state for water** (AQUA, Haldemann+2020)
 - Tidal orbital decay
- **Apply observational Kepler bias**
 - KOBE, Mishra+2021

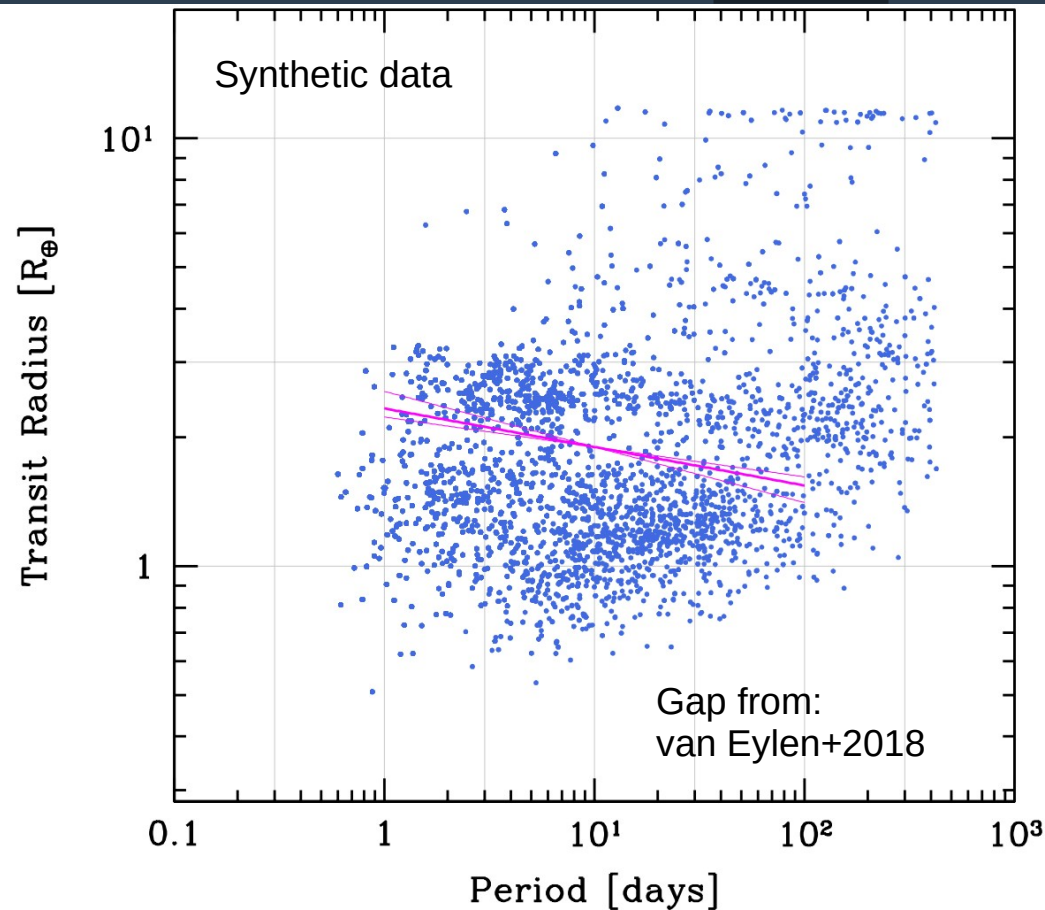
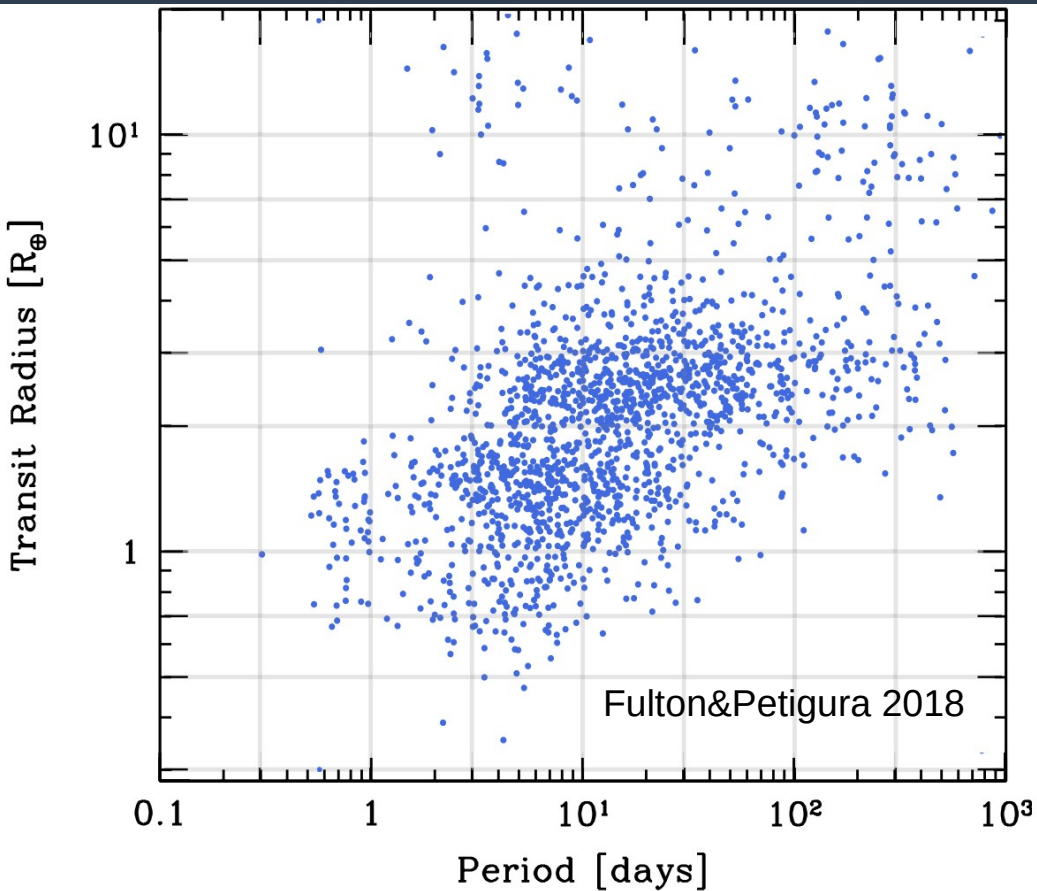


Results

- Vapor H₂O mixed with H/He
- Valley at right location (3-4 R_E)
- Separates rocky from H₂O-rich

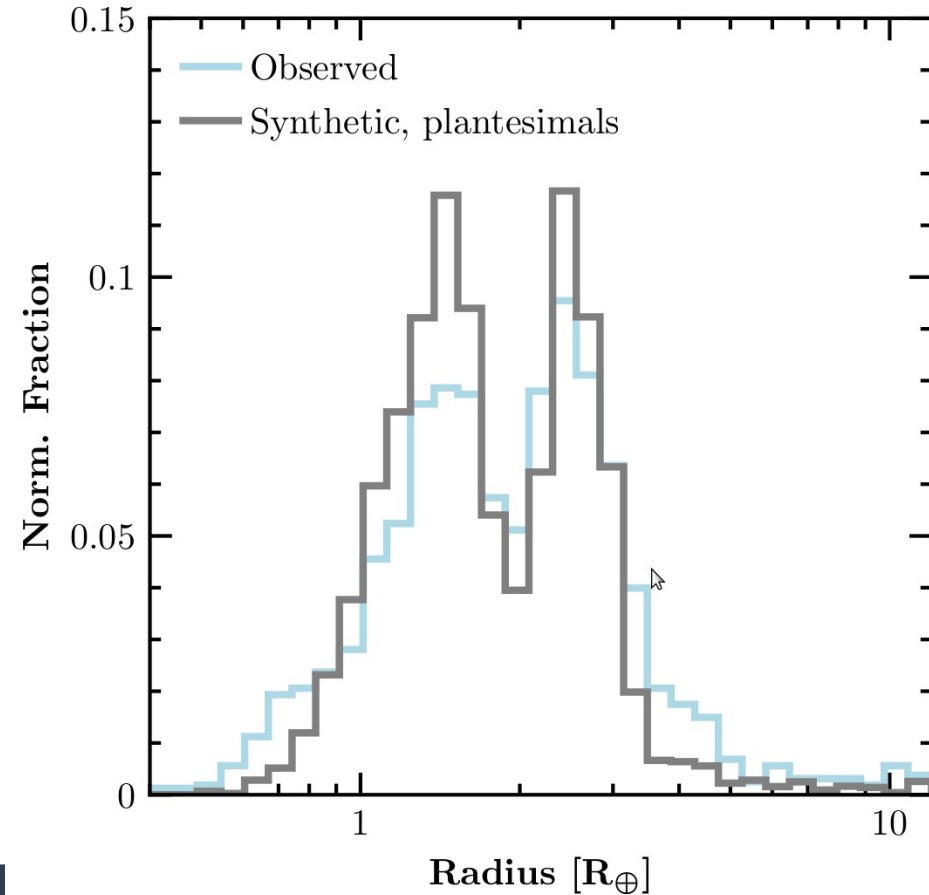


Results with KOBE (Kepler Bias)



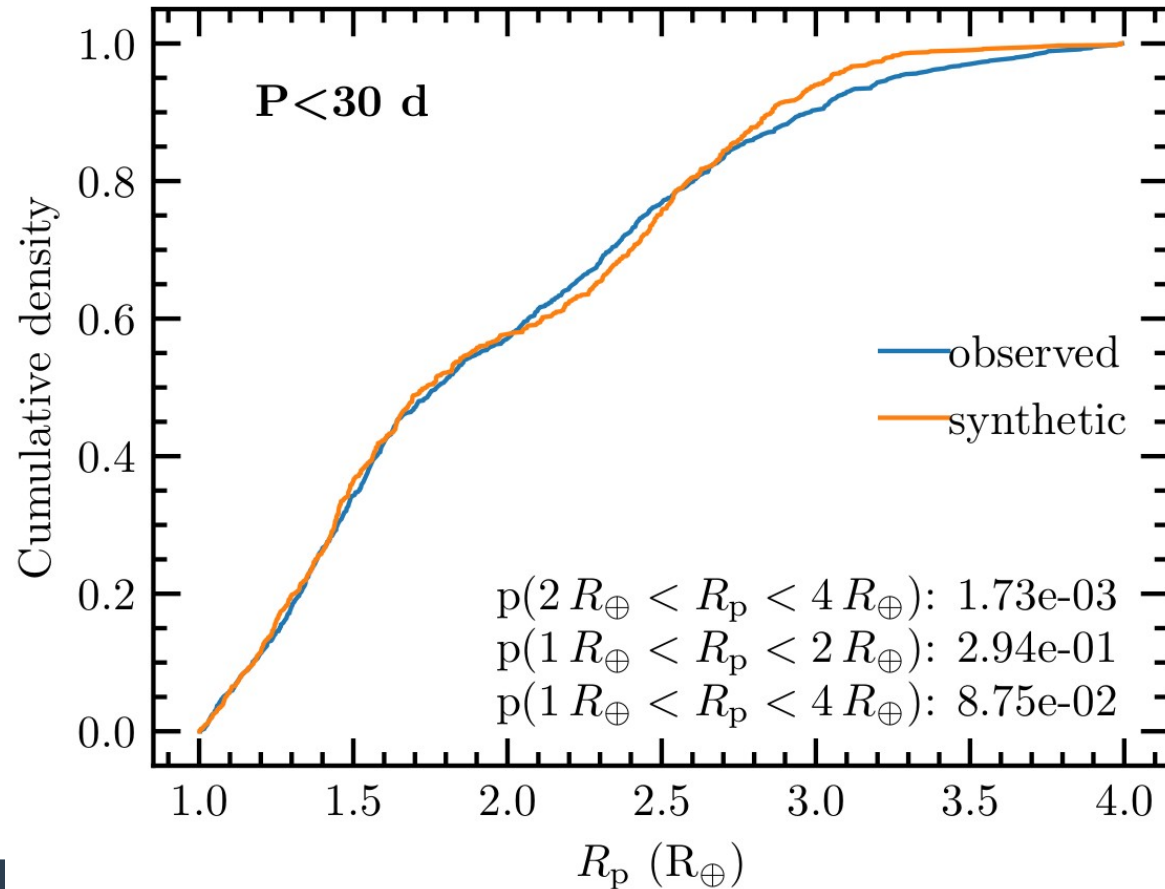
Results with KOBE (Kepler Bias)

- Radii agree qualitatively with observations



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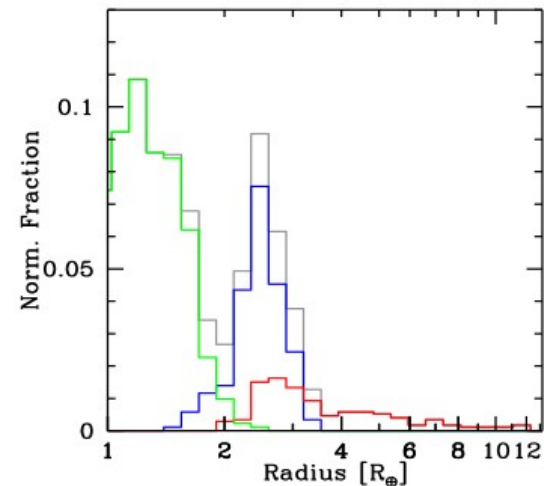
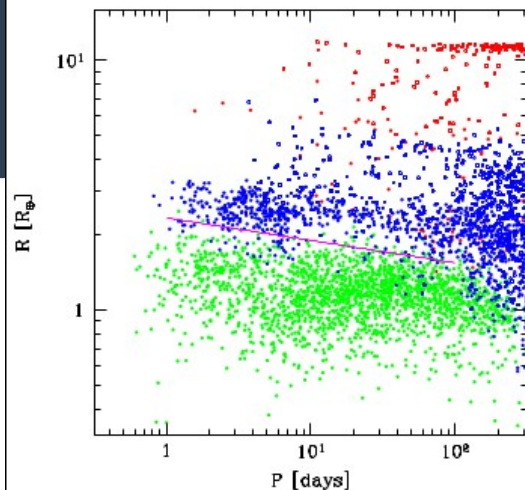
- Radii agree qualitatively with observations
- Within 30 d, even quantitatively
 - Perfect match for rocky planets



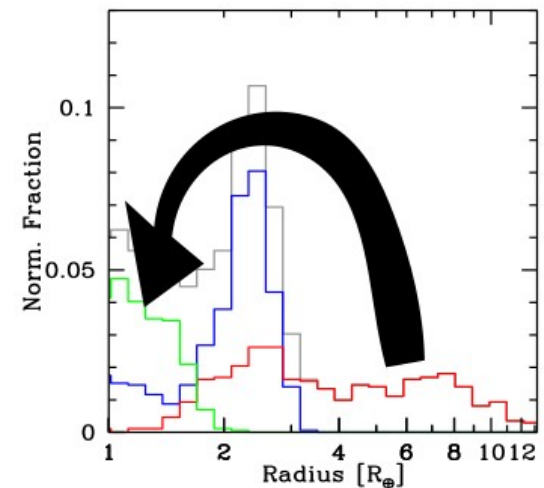
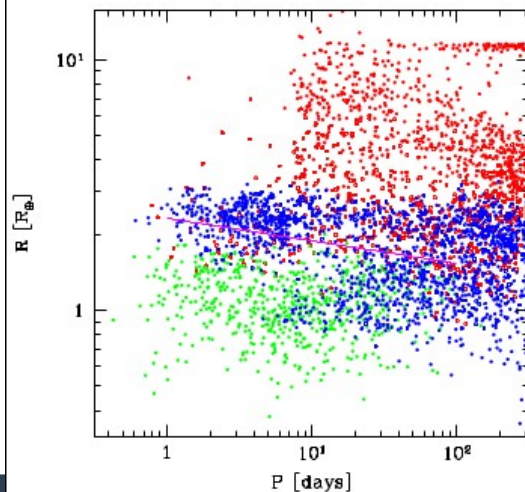
Variations

- **Without evaporation, H/He rich planets populate all radii**
 - Makes evaporation a necessary but not dominant process

with evaporation (nominal)



without evaporation



Conclusion

- **Updated planetary population synthesis models can be used for small planets**
 - Occurrence of habitable planets ranging from 0.1 to ~ 1 depending on size range
- **Reduction of giant planet formation increases habitable planet occurrence around intermediate mass stars**
- **Radius valley feature is not inconsistent with abundant water (steam) worlds**

