



The **CYGNO** Experiment: A Directional Dark Matter Detector with Optical Readout

INITIUM: Innovative Negative Ion Time projection chamber for Underground dark Matter searches

A. Prajapati on behalf of CYGNO collaboration*

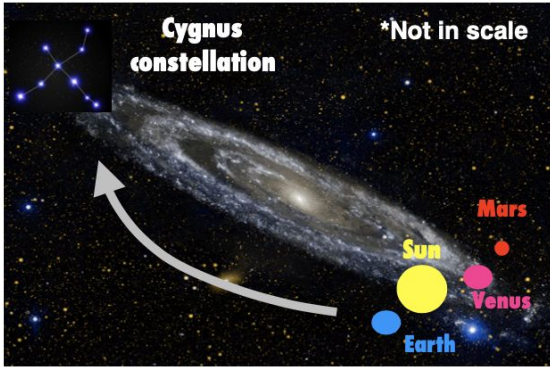
*Gran Sasso Science Institute, L'Aquila, Italy / Ph.D. student / Email: atul.prajapati@gssi.it

F. Amaro, R. Antonietti, E. Baracchini, L. Benussi, D. S. Cardoso, C. M. B. Monteiro, S. Bianco, C. Capocchia, M. Caponero, G. Cavoto, R. J. C. Roque, I. A. Costa, E. Di Marco, G. D'Imperio, G. Dho, F. Di Giambattista, R. R. M. Gregorio, F. Iacoangeli, H. P. L. Júnior, G. S. P. Lopes, G. Maccarrone, R. D. P. Mano, D. J. G. Marques, G. Mazzitelli, A.G. McLean, A. Messina, R. A. Nobrega, I. Pains, E. Paoletti, L. Passamonti, S. Pelosi, F. Petrucci, S. Piacentini, D. Piccolo, D. Pierluigi, D. Pinci, F. Renga, **A. Prajapati***, F. Rosatelli, A. Russo, G. Saviano, N. Spooner, R. Tesauro, S. Tomassini, S. Torelli, J. M. F. dos Santos

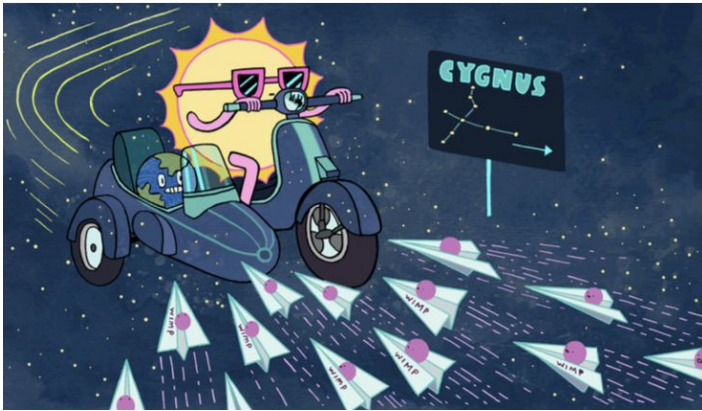


Part of this project has been funded by the European Union's Horizon 2020 research and innovation programme under the ERC Consolidator Grant Agreement No 818744





- ❖ Dark Matter forms a **halo** around our galaxy
- ❖ Our solar system rotates around galaxy towards Cygnus constellation
- ❖ Motion of our galaxy creates an **apparent wind** of DM coming from Cygnus constellation towards Earth



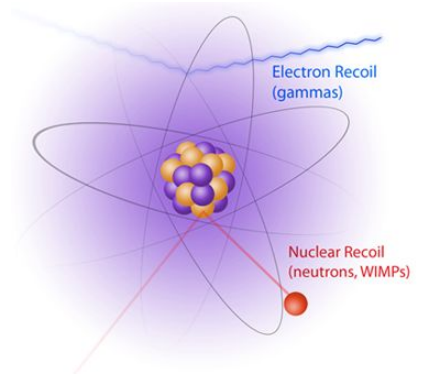
Wind of DM particles



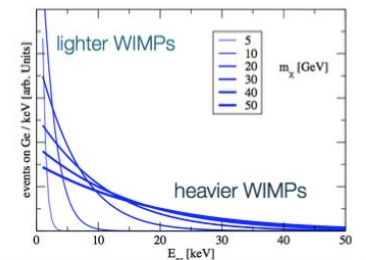
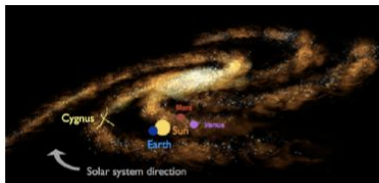
Elastic scattering with ordinary matter



Recoiling Nuclei

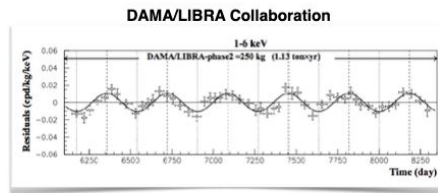
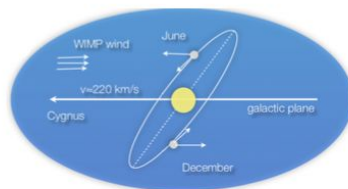


Energy Dependence



Energy dependence:
a falling exponential with
no peculiar features

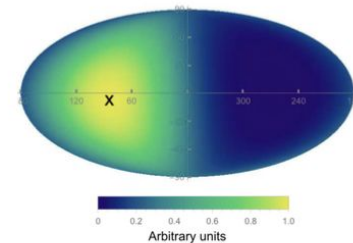
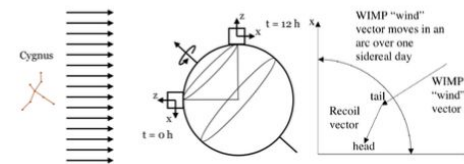
Time Dependence



Universe 4 (2018) no.11, 116

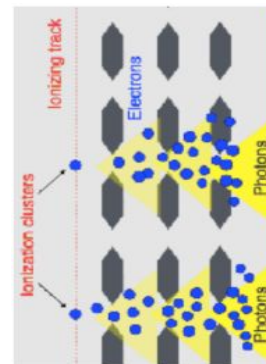
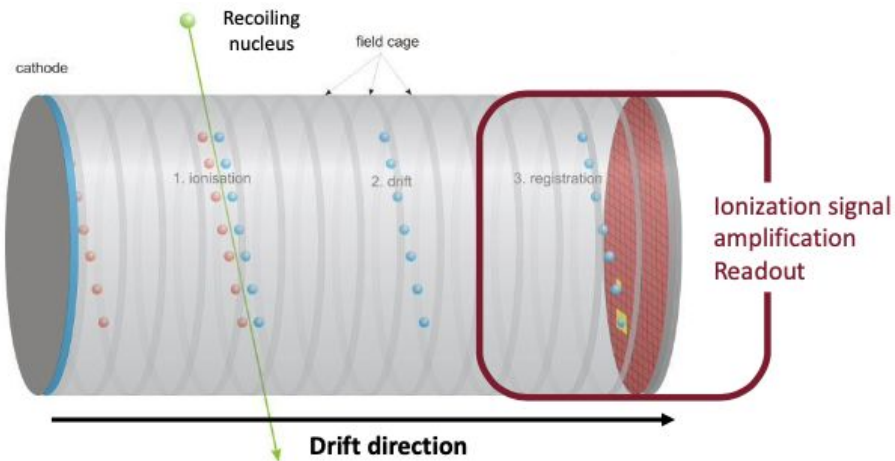
Temporal dependence:
a few % annual modulation

Directional Dependence



Directional dependence:
an $O(1)$ effect that no background
whatsoever can mimic

Increasing reliability but increasing difficulty in the experimental technique.



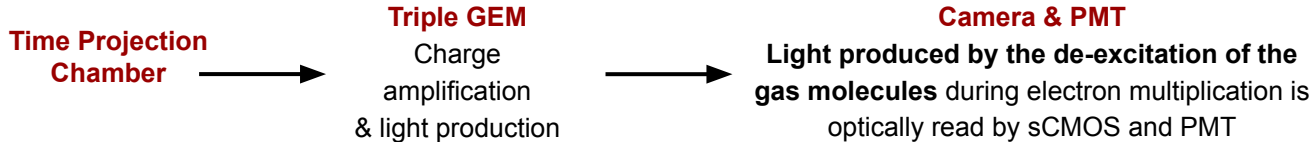
PMT



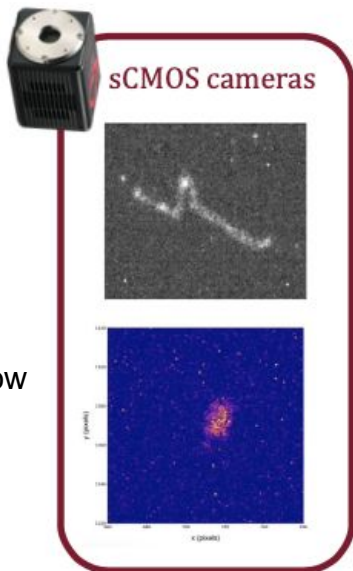
sCMOS

- ❖ CYGNO uses **He:CF₄** gas mixture at 1 atm
- ❖ **3 GEM** stack is used for charge amplification and light production

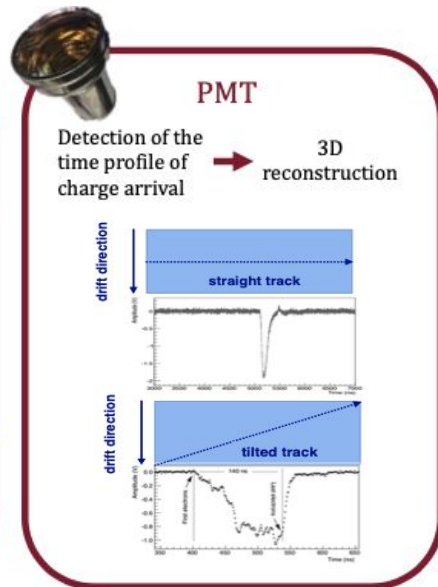
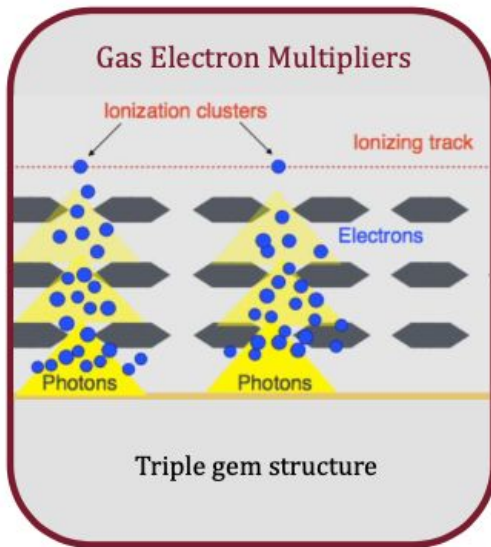
- ❖ Inherently a 3D detector
- ❖ Head/Tail recognition
- ❖ Background Rejection
- ❖ Particle Identification
- ❖ 3D fiducialization



We can measure **energy** and **X-Y coordinate** using sCMOS's high granularity and low readout noise.

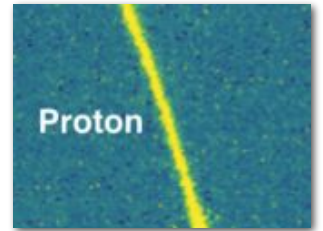
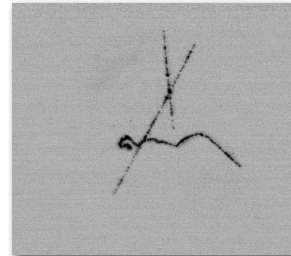
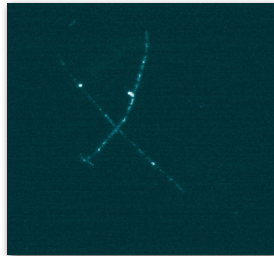
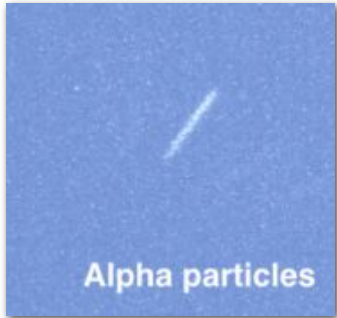
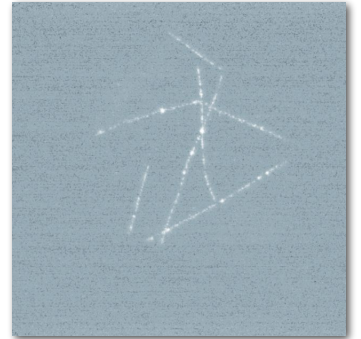
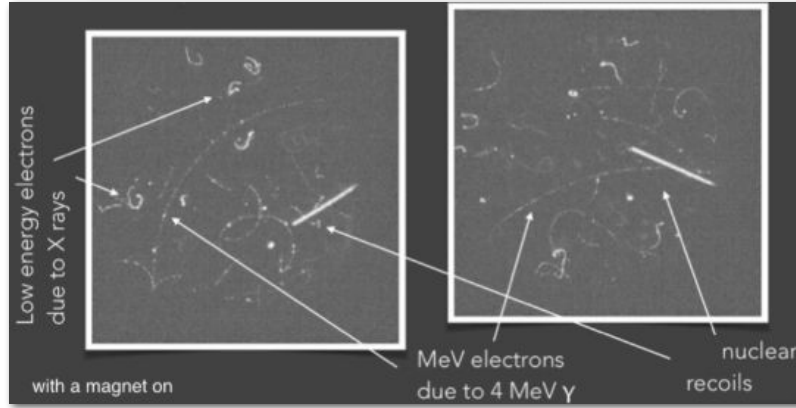
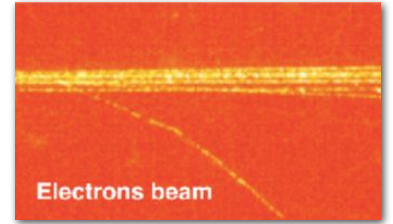
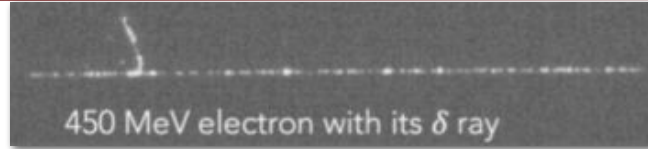
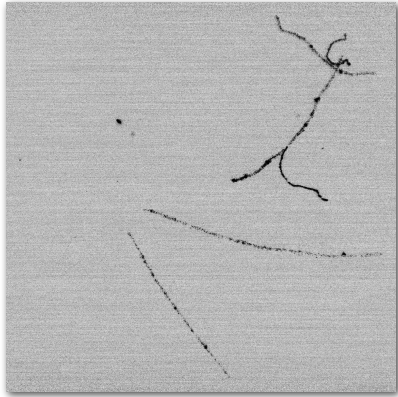


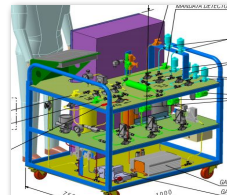
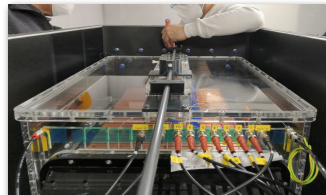
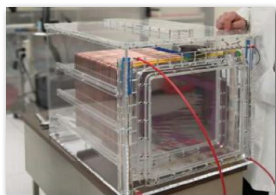
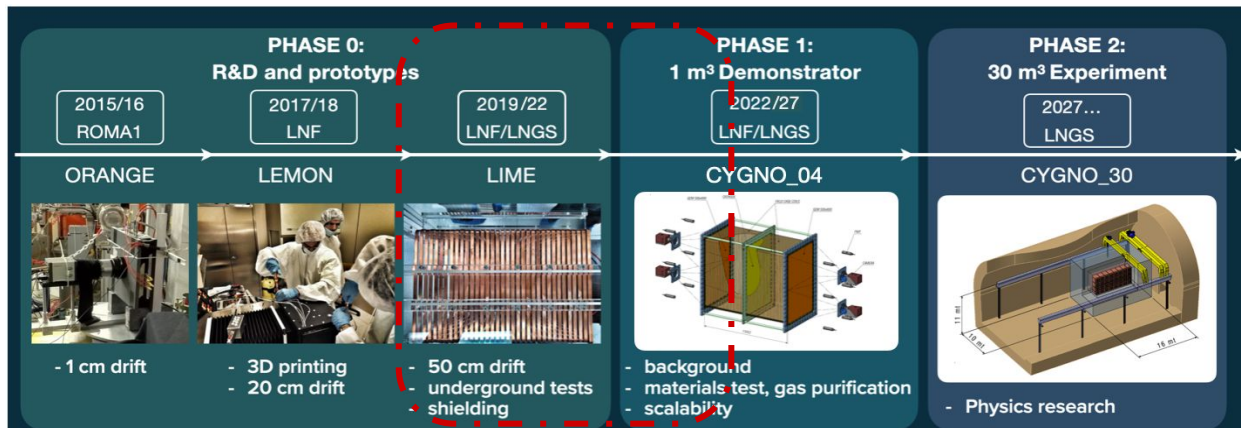
X-Y + Energy



Z + Energy

PMT measures the integrated **energy** and **time of arrival (dZ)** of charge carriers with high sampling rates.



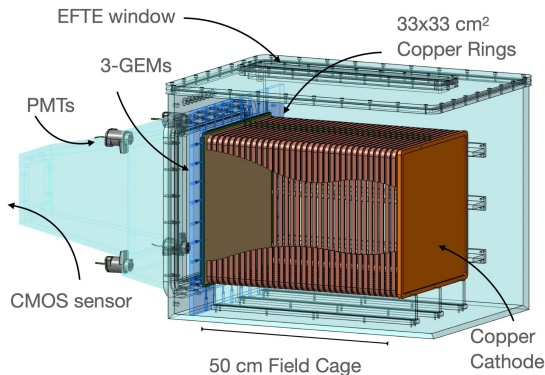


Ongoing studies:

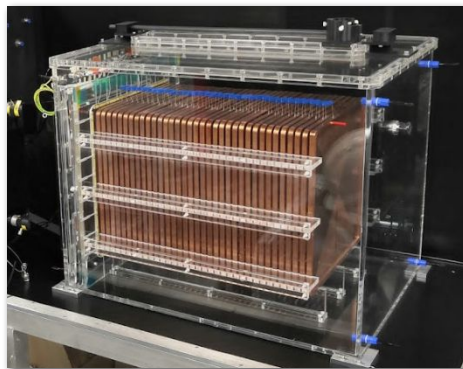
- ❖ Performance and stability test
- ❖ 3D reconstruction
- ❖ Directionality
- ❖ ER vs. NR discrimination
- ❖ Shielding materials
- ❖ Data/MC comparison

- ❖ Parallel research with MANGO detector for studying different GEM configuration, gas mixtures and Negative Ion Drift.

E. Baracchini et. al, JINST 13(2018) no.04, P04022



- ❖ **50 L** gaseous TPC with **50 cm drift**
- ❖ **He:CF₄ (60:40)** gas mixture at room temperature and atm pressure
- ❖ **Triple 33x33 cm² GEM** stack for amplification
- ❖ **Optical readout**
 - 4 PMTs
 - 1 sCMOS camera (Orca Fusion)



ORCA-Fusion CAMERA SPECS

LOW NOISE AND EXCEPTIONAL
READOUT NOISE UNIFORMITY

READOUT NOISE
0.7 electrons rms
Ultra-quiet Scan

DSNU
0.3 electrons rms

PRNU
0.06 % rms
At 7500 electrons

HIGH SPEED
100 frames/s
At 2304 × 2304

PIXEL SIZE
6.5 μm × 6.5 μm

DYNAMIC RANGE
21 400:1

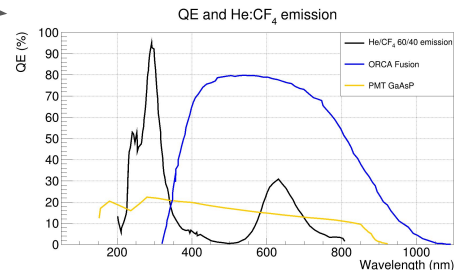
HIGH RESOLUTION
2304 × 2304
5.3 Megapixels

PEAK QE
80 %

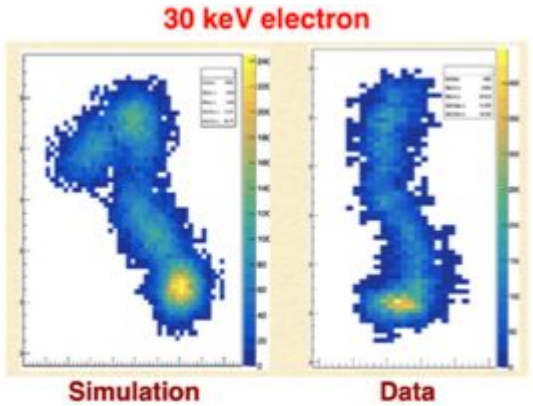
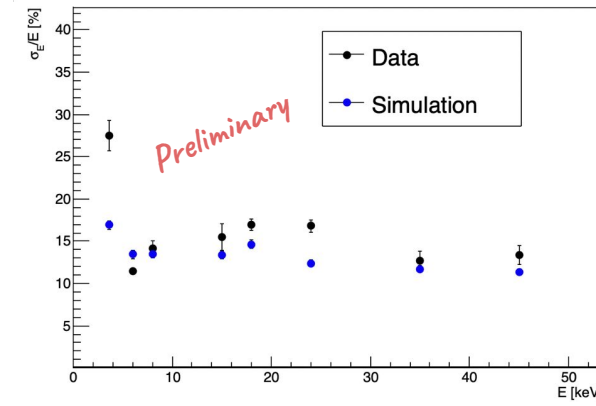
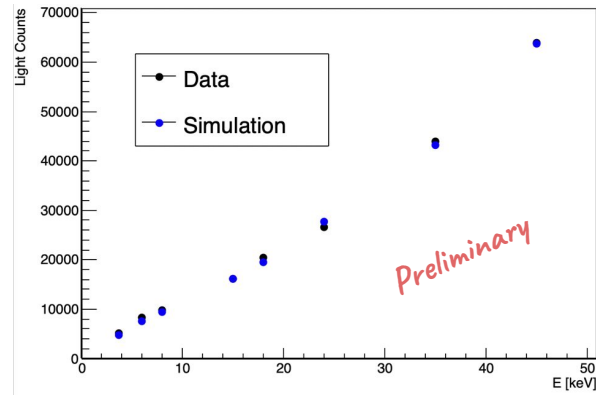


Carbon tetrafluoride (CF₄)

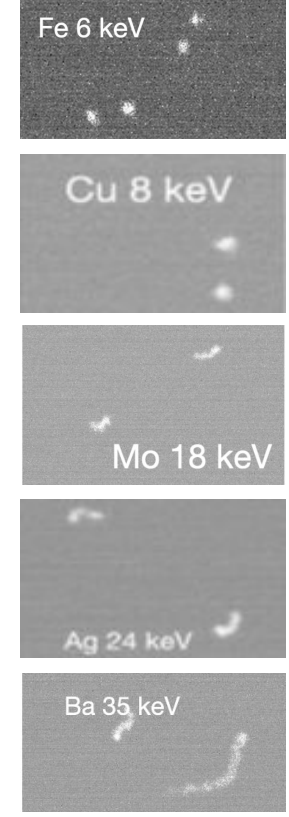
- Significant light yield at the camera's QE peak



- ❖ Multiple X-Rays sources were used to study linearity and energy resolution of the detector
- ❖ Data shows good linearity in [4-50] keV
- ❖ Energy resolution $\sim 13\%$ in [4-50] keV
- ❖ Data is in good agreement with simulation

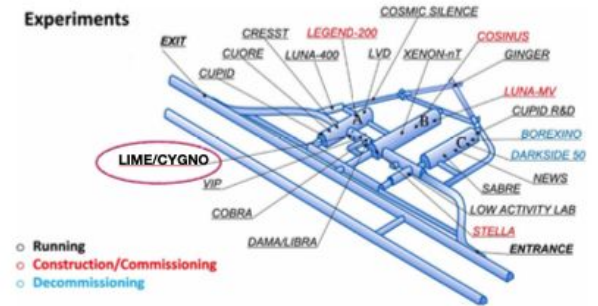
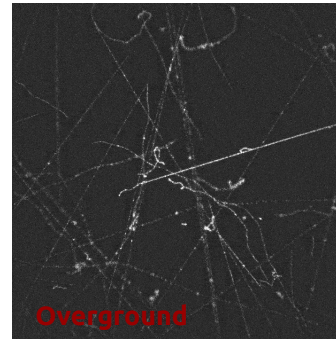
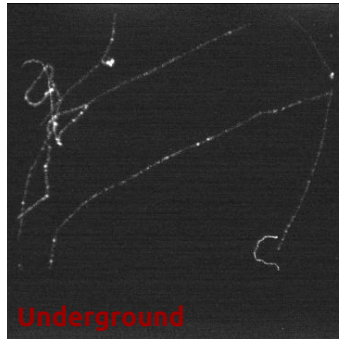
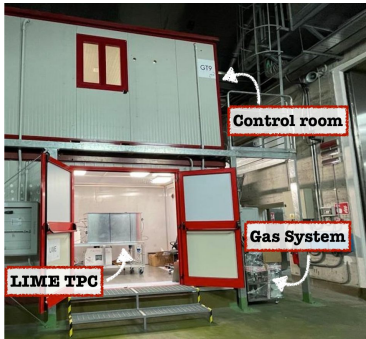


❖ Simulation developed taking into account the detector effects.



Spot like tracks
↓
Extended tracks

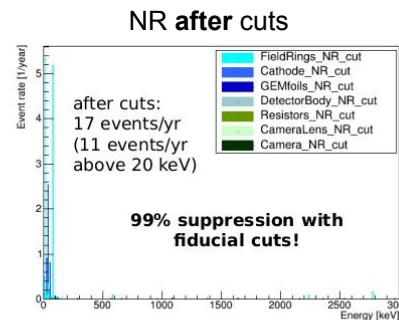
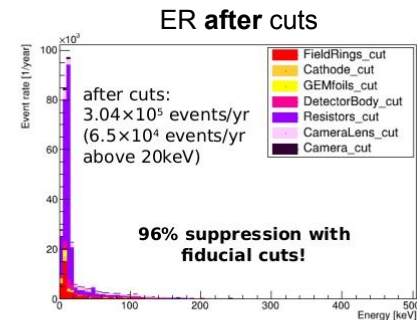
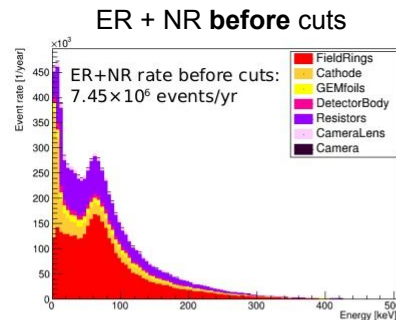
- ❖ LIME is currently installed at National Laboratory of Gran Sasso (LNGS) - INFN
- ❖ Continuously acquiring data for
 - Validation of simulated background model
 - Operating conditions optimization



- ❖ Study of expected internal bkg
- ❖ Radioactivity measured for all the detector components
- ❖ Main contribution is from resistors, GEMs and camera (lens and sensor)

Data taking program:

- ❖ **No shielding**
 - External bkg study and detector calibration
- ❖ **10 cm Cu**
 - Measurement of underground neutron flux
- ❖ **10 cm Cu + 40 cm water**
 - Study of internal bkg and validation with MC (reduction of ext. bkg. at a level less than internal bkg.)



Fiducial cuts* bring a 96% suppression of total number of recoils (ER + NR)

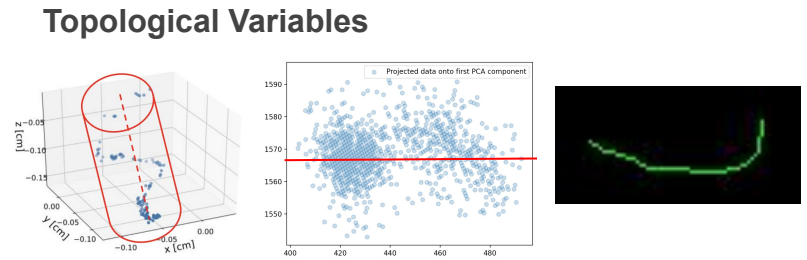
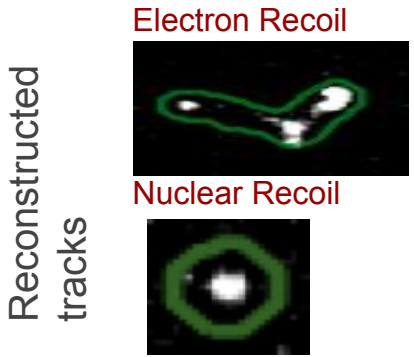
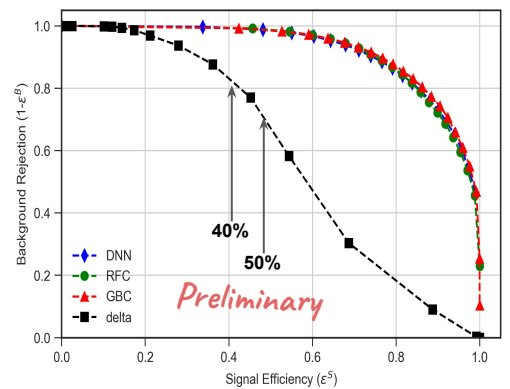
↓

After cuts and above 20 keV, we arrive at 6.5×10^4 ER/yr and **11 NR/yr**

*Cuts: 1 cm of image, 1 cm from GEMs, 4 cm from cathode

- ❖ Simulated ER and NR tracks are reconstructed in [2-36] keV range
- ❖ Topological variables are built from the reconstructed tracks
- ❖ 3 ML algorithms are trained on topological variables:
 - Random Forest Classifier (RFC)
 - Gradient Boosted Classifier (GBC)
 - Deep Neural Network (DNN)
- ❖ Development of convolutional neural networks (CNN) based model for track reconstruction and PID is ongoing

My PhD thesis

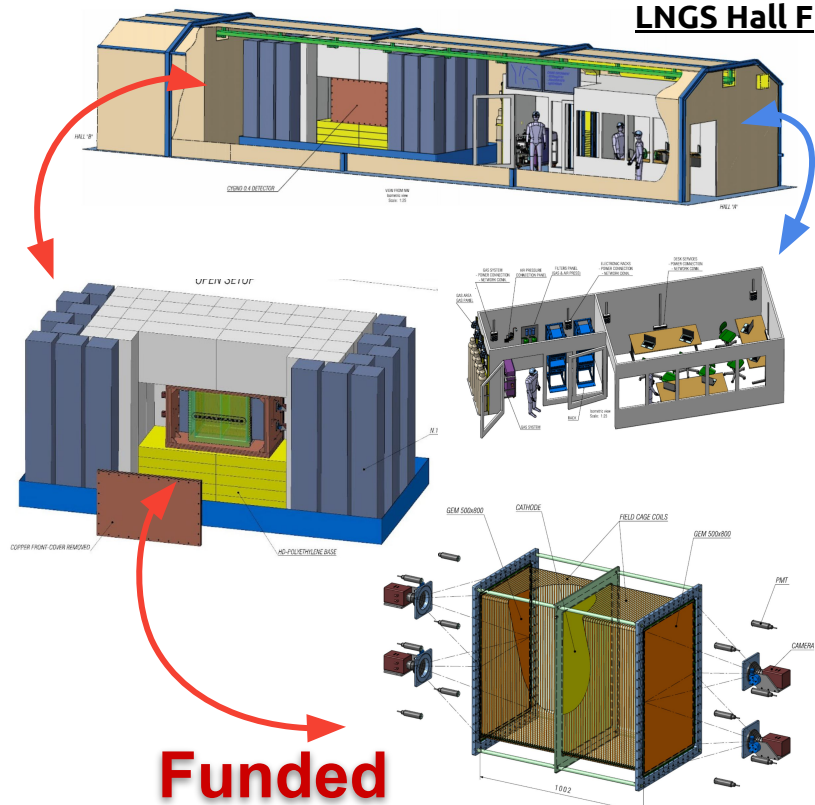


a) Cylindrical Thickness b) LAPA c) Skeleton

Models	Signal Eff. [ε ^S]%	Bkg. Rej. [1-ε ^B]%
RFC	40	99.54
	50	98.78
GBC	40	99.38
	50	98.55
DNN	40	99.43
	50	98.50
Cut-based	40	83.13
	50	67.20

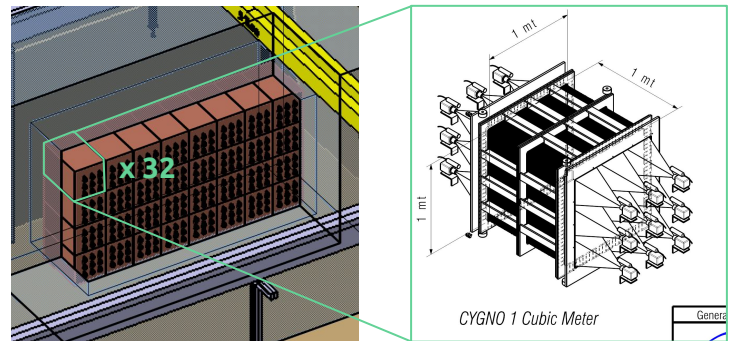
delta (light density): light integral/ no. of pixels
 cut-based* : Applying simple selection on the variable delta

- ❖ **0.4 m³** detector
- ❖ **Triple 50 x 80 cm²** GEMs
- ❖ Common central cathode
- ❖ Readout by **4 sCMOS** (ORCA Quest) and **12 PMTs**
- ❖ Low radioactivity acrylic glass vessel
- ❖ Field cage made by **copper strips on insulator support** (DRIFT like)
- ❖ Will be used to demonstrate the scalability/feasibility of detection technique towards **CYGNO_30** with O(30 m³)

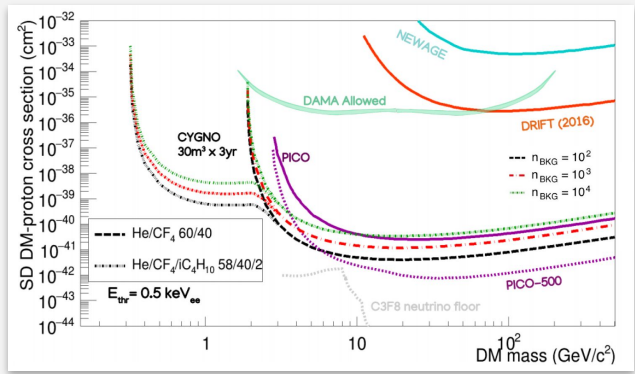
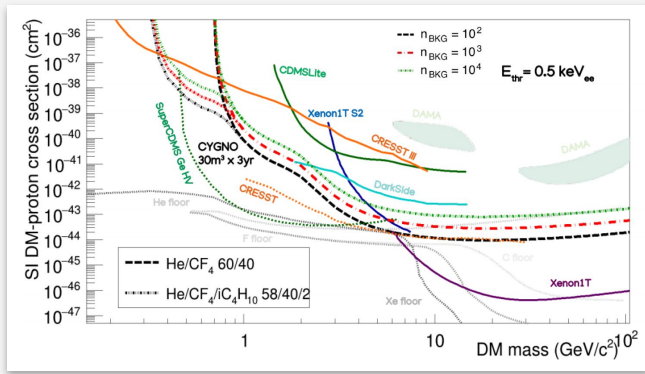


- ❖ Low mass (0.5 - 10 GeV) directional DM searches
- ❖ > 2027
- ❖ 30 - 100 m³ detector
- ❖ 0.5 - 1 keV_{ee} energy threshold
- ❖ 30° angular resolution

Amaro, F.D. et. al The CYGNO Experiment. Instruments 2022, 6, 6.



Expected **SI** and **SD** (90% CL)
interaction cross-section exclusion
 Quenching factor simulated
 with **SRIM** → Direct
 measurement incoming!
He allows us to explore very low DM
 masses!



Advantages

Reduced diffusion

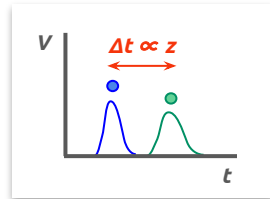
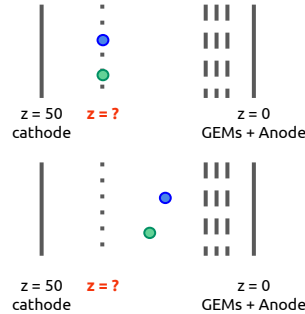
Longitudinal and transverse **diffusion** reduced to thermal limit

$$\sigma_D = \sqrt{\frac{4\varepsilon L}{eE}}$$

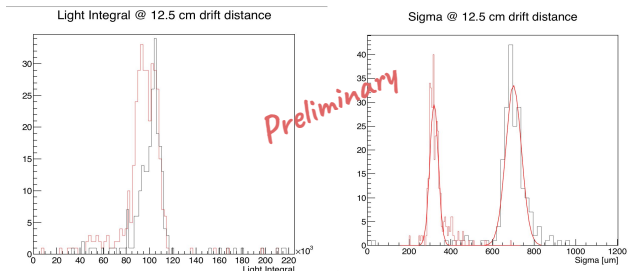
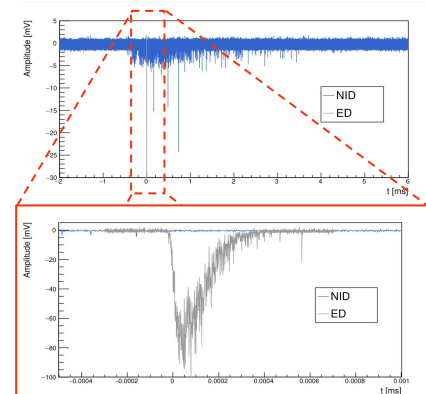
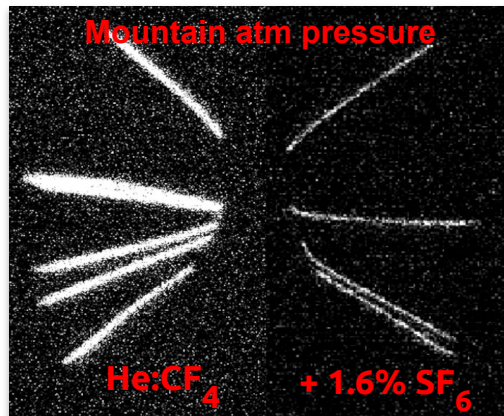
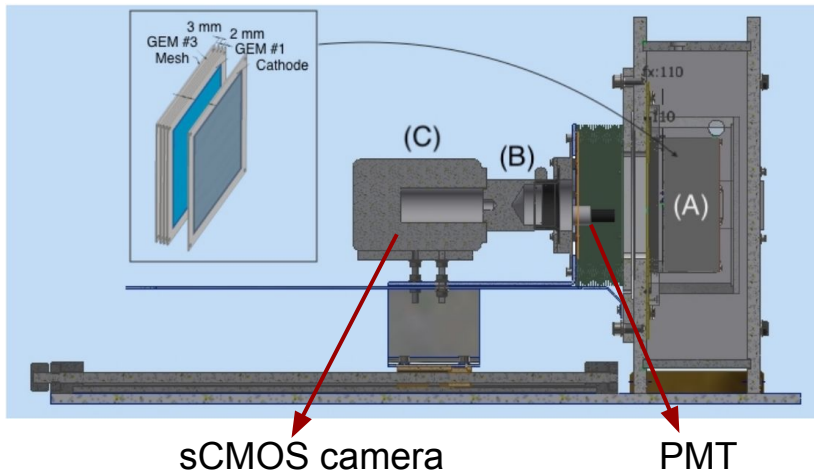
Better spatial resolution!

Absolute Z from Δt between minority charge carriers

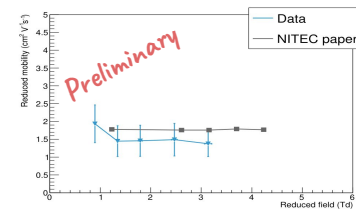
Multiple charge carriers



- ❖ Highly **electronegative dopant** is added to the gas (CS_2 , SF_6 , ...)
- ❖ Primary electrons are captured by the dopant at $O(100 \text{ um})$
- ❖ **Anions** are majority charge carriers instead of electrons
- ❖ σ_T and σ_L reduced to **thermal limit**
- ❖ **Lower drift** velocity of Negative ions **$O(\text{cm/ms})$** significantly improves the resolution along the drift direction



Same light ... smaller sigma



Reduced mobility compatible with SF₆

- ❖ The **CYGNO** collaboration is developing a **high-precision gaseous TPC** at atmospheric pressure with **optical readout**.
- ❖ The main focus is the **direct search of DM WIMP-like particles** in the **low mass range** (0.5-10 GeV).
- ❖ Through **directionality**, solar neutrinos can be discriminated and **unambiguous confirmation of DM** is possible.
- ❖ The **50L LIME prototype** was recently installed in the **underground LNGS** facilities.
 - The first **stability tests, background evaluations and measurements** are being carried out.
- ❖ **CYGNO_04**, already funded and with a TDR submitted, will allow us to test the experiment's **scalability**.
- ❖ **CYGNO_30** is under study, with its sensitivities looking promising.
- ❖ Several **R&D projects** are ongoing in order to find **optimal means of TPC operation**:
 - **Electroluminescence** observed in our conditions and its **potentialities** are under study
 - **Negative ion drift** observed for the first at atmospheric pressure and with PMTs.



...for more info:
**CYGN0 - Directional
Dark Matter Search**

[https://www.facebook.com/
cygn0.experiment](https://www.facebook.com/cygn0.experiment)

THANKS!