Anisotropy and the missing rest frame of the Universe

A challenge to the standard cosmological model

M. Rameez work in coll with: Secrest, von Hausegger, Colin, Mohayaee, Sarkar



19th Rencontres du Vietnam Quy Nhon, Socialist Republic of Vietnam

The cosmological principle

The Universe is (statistically) isotropic and homogenous (on large scales).



The real reason, though, for our adherence here to the Cosmological Principle is not that it is surely correct, but rather, that it allows us to make use of the extremely limited data provided to cosmology by observational astronomy.

If the data will not fit into this framework, we shall be able to conclude that either the Cosmological Principle or the Principle of Equivalence is wrong. Nothing could be more interesting.

Steven Weinberg, Gravitation and Cosmology (1972)

"Data from the Planck satellite show the Universe to be highly isotropic"



The CMB Dipole : Purely Kinematic?



Net motion of the Solar System barycentre: 369 +/- 2 km/s w.r.t 'CMB rest frame' towards

R.A = 168.0, DEC = -7.0

- Motion of the Sun around the Galaxy ~225 +/- 18 km/s
- The motion of the Local Group 627+/-22 km/s ApJ, 709, 483

Is this 'Purely Kinematic'?

What is the origin of this motion?

cooler

COBE Experiment, 1996 Planck 2015

 $\frac{\Delta T}{T} \simeq 10^{-3}$

A moving observer- Kinematic Dipole



On the expected anisotropy of radio source counts

G. F. R. Ellis^{*} and J. E. Baldwin[†] Orthodox Academy of Crete, Kolymbari, Crete

Received 1983 May 31; in original form 1983 March 31

Summary. If the standard interpretation of the dipole anisotropy in the microwave background radiation as being due to our peculiar velocity in a homogeneous isotropic universe is correct, then radio-source number counts must show a similar anisotropy. Conversely, determination of a dipole anisotropy in those counts determines our velocity relative to their rest frame; this velocity must agree with that determined from the microwave background radiation anisotropy. Present limits show reasonable agreement between these velocities.



4 Conclusion

Anisotropies in radio-source number counts can be used to determine a cosmological standard of rest. Current observations determine it to about $\pm 500 \text{ km s}^{-1}$, but accurate counts of fainter sources will reduce the error to a level comparable to that set by observations of the microwave background radiation. If the standards of rest determined by the MBR and the number counts were to be in serious disagreement, one would have to abandon either

(a) the idea that the radio sources are at cosmological distances, or

(b) the interpretation of the cosmic microwave radiation as relic radiation from the big bang, or

(c) the standard FRW Universe models.

Thus comparison of these standards of rest provides a powerful consistency test of our understanding of the Universe.

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Thus comparison of these standards of rest provides a powerful consistency test of our understanding of the Universe.

The situation anticipated by Ellis and Baldwin in 1984 now confronts us!



1.4 GHz survey of the Northern sky, by the National Radio Astronomy Observatory. Down to dec = -40.4°

1,773,488 sources above 2.5 mJy. But 'complete' with uniform sky exposure only above 10 mJy

Phys. Rev. D, 78, 043519

First seen by Singal, A. K. 2011, ApJL, 742, L23,

Sydney University Molonglo Sky Survey (SUMSS)



843 MHz survey of the Southern sky, by the Molonglo Observatory Synthesis telescope. Dec < -30.0°

211050 radio sources. Similar sensitivity and resolution to NVSS

The NVSUMSS-Combined All Sky catalog



- Rescale SUMSS fluxes by (843/1400)^{-0.75}
- Remove Galactic Plane at +/-10 degree in NVSS
- Remove NVSS sources below and SUMSS sources above dec -30 (or -40)
- Apply common threshold flux cut on both samples
- z~1



Velocity ~ 1355 \pm 351 km/s, Dir within 10° of CMB dipole direction.

Statistical significance, ~2.81 Sigma, with the 3D linear estimator, constrained mainly by the catalogue size

Bengaly et al 2018 JCAP 1804 (2018) no.04, 031 find a 5.1 sigma excess in TGSS !

SKA phase 1 measurement ~10%

Bengaly (et al) 2018 MNRAS, 486, Issue 1 (2019) 1350-1357

Siewert et al 2020, Astron. Astrophys. 653 (2021) A9

"We conclude that for all analysed surveys, the observed Cosmic Radio Dipole amplitudes **exceed the expectation**, derived from the CMB dipole."

The Widefield Infrared Survey Explorer



All sky infrared survey over 10 months, in the bands 3.4, 4.6, 12 and 22 μ m using a 40 cm diameter telescope

Generated a catalog of 746 million+ objects, most of which are stars.

Directionally unbiased survey strategy, arc second angular resolution, multi band photometry.



CatWISE AGN 1355352 sources





Astrophys.J.Lett. 908 (2021) 2, L51

Rameez-Vietnam

Results



p = 5 × 10⁻⁷ (4.9 σ)

Obtained by scrambling the data itself, frequentist null hypothesis testing,

https://zenodo.org/record/4448512



Conservative Sample size weighted Z-scores : 5.1 σ

Astrophys.J.Lett. 937 (2022) L31 https://zenodo.org/record/6784602

Testing the Cosmological Principle with CatWISE Quasars: A Bayesian Analysis of the Number-Count Dipole

Lawrence Dam^{1,2}, * Geraint F. Lewis¹[†] & Brendon J. Brewer³

¹Sydney Institute for Astronomy, School of Physics, A28, The University of Sydney, NSW 2006, Australia

²Département de Physique Théorique and Center for Astroparticle Physics, Université de Genève, 24 quai Ernest-Ansermet, 1211 Genève 4, Switzerland ³Department of Statistics, The University of Auckland, Private Bag 92019, Auckland 1142, New Zealand



tions — quasars: general — galaxies: active

ABSTRACT

The Cosmological Principle, that the Universe is homogeneous and isotropic on sufficiently large scales, underpins the standard model of cosmology. However, a recent analysis of 1.36 million infrared-selected quasars has identified a significant tension in the amplitude of the number-count dipole compared to that derived from the CMB, thus challenging the Cosmological Principle. Here we present a Bayesian analysis of the same quasar sample, testing various hypotheses using the Bayesian evidence. We find unambiguous evidence for the presence of a dipole in the distribution of quasars with a direction that is consistent with the dipole identified in the CMB. However, the amplitude of the dipole is found to be 2.7 times larger than that expected from the conventional kinematic explanation of the CMB dipole, with a statistical significance of 5.7 σ . To compare these results with theoretical expectations, we sharpen the ACDM predictions for the probability distribution of the amplitude, taking into account a number of observational and theoretical systematics. In particular, we show that the presence of the galactic plane mask causes a considerable loss of dipole signal due to a leakage of power into higher multipoles, exacerbating the discrepancy in the amplitude. By contrast, we estimate using probabilistic arguments that the source evolution of guasars improves the discrepancy, but only mildly so. These results support the original findings of an anomalously large quasar dipole, independent of the statistical methodology used.



Rameez-Vietnam

AllWISE-Galaxies

following from MNRAS448,1305-1313 (2015)

- Magnitude cuts in different bands, Galactic plane cut at +/-15 degrees
 - Sample of 2.46 million Galaxies, 76% complete, with 1.8% star contamination



Cross correlate with deep surveys over a very narrow sky (SDSS, GAMA) to determine how many are stars and how many are Galaxies

The maximum is in the direction (AllWISE) 237.4° RA, -46.6 ° Dec 331.9° I 6.02° b

110 degrees from the CMB direction

Dipole magnitude ~0.049

Fully kinematic interpretation ~6000 km/s

in agreement with MNRAS 445 (2014) L60-L64

Getting rid of the stars



Apparent motion = parallax + proper motion

Stars in the Galaxy have higher apparent motions 400 mas/yr up to many arc seconds/ year

Cuts on apparent motion can bring star contamination down to 0.1%, while still keeping ~1.8 millin galaxies.

182.9° RA, -55.6° DEC, 50.1° from the CMB

Dipole magnitude reduces to 0.014

Star galaxy identification by cross correlating with SDSS

Suppressing local anisotropies

~200 Mpc

6.1" PSF

Remove extended sources and the supergalactic plane.

Further reduce z<0.03 sources by cross correlating with 2MRS and removing the correlated sources.



1192182 - AllWISE Galaxies



VELOCITY COMPONENTS OF THE OBSERVED CMB DIPOLE



Where is the cosmic 'rest frame'?



Qin et al, <u>Astrophys. J.922:59,2021</u>

G. Lavaux, R.Brent Tully, R. Mohayaee, S. Colombi

•Astrophys.J. 709 (2010) 483-498

Gravity in the Local Universe: density and velocity fields using CosmicFlows-4

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Received A&A Oct 31, 2022 - AA/2022/45331; Accepted date

ABSTRACT

This article publicly releases three-dimensional reconstructions of the local Universe gravitational field below z=0.8 that were computed using the full catalogue CosmicFlows-4 of 56,000 galaxy distances and its sub-sample of 1,008 type Ia supernovae distances. The article also provides some first CF4 measurements of the growth rate of structure using the pairwise correlation of peculiar velocities $f\sigma_8 = 0.44(\pm 0.01)$ and of the bulk flow in the Local Universe of 200 ± 88 kms¹ at distance 300 h⁻¹₁₀₀ Mpc.

Key words. Cosmology: large-scale structure of Universe

Where is the cosmic 'rest frame'?

This is the model using which Pantheon and Pantheon+ compilations



Carrick, Turnbull, Lavaux, Hudson *MNRAS*, 450, 1, 11 2015, 317–332

"We find that an external bulk flow is preferred at the 5.1 σ level, and the best fit has a velocity of 159 ± 23 km s⁻¹ towards *I* = 304° ± 11°, *b* = 6° ± 13°" [beyond 200 h^{-1} Mpc radius] Infers the peculiar velocity field from a density contrast field derived from data – 2M++ compilation, using linear Newtonian perturbation theory.

$$v(\mathbf{r}) = \frac{H_0 f(\Omega_m)}{4\pi b_g} \int \delta_g(\mathbf{r}') \frac{(\mathbf{r}'-\mathbf{r})}{|\mathbf{r}'-\mathbf{r}|^3} d^3\mathbf{r}'.$$



The tilted Friedmann Universe



If we are inside a large local 'bulk flow'.

(Tsagas 2010, 2011, 2012; Tsagas & Kadiltzoglou 2015, Tsagas 2019, 2021)

The patch A has mean peculiar velocity \tilde{v}_a with $\vartheta = \tilde{D}^a v_a \ge 0$ and $\dot{\vartheta} \ge 0$ (the sign depending on whether the bulk flow is accelerating or decelerating)

Inside region B, the r.h.s. of the expression $1 + \tilde{q} = (1+q) \left(1 + \frac{\vartheta}{\Theta}\right)^{-2} - \frac{3\dot{\vartheta}}{\Theta^2} \left(1 + \frac{\vartheta}{\Theta}\right)^{-2}, \qquad \tilde{\Theta} = \Theta + \vartheta,$

drops below 1 and the observer 'measures' *negative* deceleration parameter in one direction of the sky - – i.e. towards the CMB dipole



This implies that **observers** experiencing locally accelerated expansion, as a result of their own drift motion, may also find that the acceleration is maximised in one direction and minimised in the opposite. We argue that, typically, such a dipole anisotropy should be relatively small and the axis should probably lie fairly close to the one seen in the spectrum of the Cosmic **Microwave Background.**

Testing this on a sample of 740 SN1e, JLA



The dipolar component of q is larger than the monopole, and dominates out to z>0.1, closely aligned to the CMB dipole

The significance of q_o being negative is $<1.4\sigma!$ Dipole Statistically significant at 3.9 σ level In agreement with the predictions by Tsagas,

Kicked off a debate with mainstream supernova cosmologists, about the data being corrected for 'peculiar velocities'

Cosmic acceleration may simply be an artefact of our being located inside a 'bulk flow'!

The question of peculiar velocity 'corrections'

-600

10-3

$$1 + z = (1 + \bar{z}) (1 + z_{pec}^{hel}) (1 + z_{pec}^{SN})$$
$$d_L(z) = \bar{d}_L(\bar{z}) (1 + z_{pec}^{hel}) (1 + z_{pec}^{SN})^2$$

Davis et. al. Astrophys.J. 741 (2011) 67 Ellis & Stoeger 1987

JLA (and Pantheon) redshifts and magnitudes have been 'corrected' to account for the local bulk flow.

<pre>#name zcmb zhel</pre>	dz mb dmb x1 dx1 color dcolor
03D1au 0.503084	0.504300 0 23.001698 0.088031
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03D1ax 0.494795	0.496000 0 22.960139 0.088110
03D1bp 0.345928	0.347000 0 22.398137 0.087263
03D1co 0.677662	0.679000 0 24.078115 0.098356
03D1dt 0.610712	0.612000 0 23.285241 0.092877
03D1ew 0.866494	0.868000 0 24.353678 0.106037
03D1fc 0.330932	0.332000 0 21.861412 0.086437
03D1fg 0.798566	0 800000 0 24 510389 0 101777

$$C = [(1 + 2_{hel}) - (1 + 2_{cmb})(1 + 2_d)] \times C$$

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 $C = [(1 \pm 7) + (1 \pm 7) + (1 \pm 7)] \times C$

Flow model – SMAC has a ~600 km/s residual bulk flow

10⁻¹

SN1a at z>0.06 are assumed (arbitrarily) to be in the CMB

rest frame. (only uncorrelated 150 km/s in error budget)

10⁰

 10^{1}

 $z_{hel} \rightarrow measured$ $z_{cmb} \rightarrow inferred using a flow model$ 10⁻²

The Dipole of the Pantheon+SH0ES

Data Francesco Sorrenti, Ruth Durrer and Martin Kunz

Département de Physique Théorique and Center for Astroparticle Physics, Université de Genève, 24 quai Ernest Ansermet, 1211 Genève 4, Switzerland

E-mail: frances co.sorrenti@unige.ch, ruth.durrer@unige.ch, martin.kunz@unige.ch

Abstract. In this paper we determine the dipole in the Pantheon+ data. We find that, while its amplitude roughly agrees with the dipole found in the cosmic microwave background which is attributed to the motion of the solar system with respect to the cosmic rest frame, the direction is different at very high significance. While the amplitude depends on the lower redshift cutoff, the direction is quite stable. For redshift cuts of order $z_{\rm cut} \simeq 0.05$ and higher, the dipole is no longer detected with high statistical significant. An important rôle seems to be played by the redshift corrections for peculiar velocities.

4.4 The peculiar velocities in the Pantheon+ analysis

An important difference between our treatment and the Pantheon+ analysis [18], lies in the peculiar velocities of the SNe which we neglect in our analysis. The main reason we do this is that they should not contribute significantly to the dipole which is the main aim of this work. We also consider it problematic that the peculiar velocities inferred in [18] come purely from linear gravitational infall [24, 29], even though it is known from numerical simulations that at late times vorticity is as relevant as (if not larger than) the

No Hubble tension either

New Constraints on Anisotropic Expansion from Supernovae Type Ia

W. Rahman^{1*}, R. Trotta^{1,2,3}, S. S. Boruah⁴, M. J. Hudson^{5,6,7}, and D. A. van Dyk^{2,8}

The model used in Betoule et al. (2014) to estimate $z_{\text{pec}}^{\text{SN}}$ has been criticised by C19, who highlighted potential bulk flow velocity discontinuities at z = 0.04, pointed out that peculiar velocity corrections arbitrarily disappear beyond 200/h Mpc ($z \sim 0.067$, the limit of the galaxy density field measurements from which the peculiar velocities were derived) and that the residual uncorrelated velocity dispersion of $\sigma_v = 150$ km/s might be underestimated. While RH20 pointed out technical flaws with the analysis of C19, it is important in the light of this valid criticism to revisit the issue of low-redshift peculiar velocity corrections here.





Conclusion The Universe is anisotropic and the Cosmic Rest frame does not exist

- Ellis & Baldwin tests performed on Radio galaxy catalogues and WISE Quasars conclusively reject the exclusively kinematic interpretation of the CMB dipole at > 5 σ . CMB rest frame and matter rest frame are different. Cosmological principle has been falsified.
- SN1a data are better fit by a "tilted Friedmann model". Ensuing debate stultifies dark energy evidence.
- Strong hint towards the inhomogeneous cosmological models.
- Highlighted by Peebles in his review of anomalies in physical cosmology.
- 300+ citations, Quanta Magazine, New Scientist
- Stultifies the Hubble tension:
 - *M.R. and Subir Sarkar, Class.Quant.Grav.* 38 (2021) 15, 154005
 - A new cosmological tension!

Three projects in LSST DESC All who have data access are welcome to join

Reviews Mohayaee, Rameez & Sarkar *Eur.Phys.J.ST* 230 (2021) 9, 2067-2076

Subir Sarkar "Heart of Darkness" Inference: International Review of Science 6 (2022) 4

Dipole Cosmology: The Copernican Paradigm Beyond FLRW

QCD axion dark matter and the cosmic dipole anomaly

Chengcheng Han^{1,*}

¹School of Physics, Sun Yat-Sen University, Guangzhou 510275, China (Dated: November 29, 2022)

Chethan KRISHNAN^{a*}, Ranjini MONDOL^{$a\dagger$}, M. M. SHEIKH-JABBARI^{$b\ddagger$}

^{*a*} Center for High Energy Physics, Indian Institute of Science, Bangalore 560012, India ^{*b*} School of Physics, Institute for Research in Fundamental Sciences (IPM), P. O. Box 19395-5531, Tehran, Iran SO(3) → U(1), tilted Bianchi V/VII_h - 4 Friedmann equations

Large-scale geometry of the Universe

Yassir Awwad^{\bigstar} and Tomislav Prokopec^{\diamond}

 \diamond Institute for Theoretical Physics, Spinoza Institute & EMME Φ Utrecht University, Princetonplein 5, 3584 CC Utrecht, The Netherlands

Thursten Perelman theorem -> anisotropic Thursten geometries should be considered on par with Friedmann geometry

Spatially Homogeneous Universes with Late-Time Anisotropy

Andrei Constantin[®],^{1,*} Thomas R. Harvey[®],^{1,†} Sebastian von Hausegger[®],^{2,‡} and Andre Lukas[®], [§] ¹Rudolf Peierls Centre for Theoretical Physics, University of Oxford, Parks Road, Oxford, UK ²Astrophysics, University of Oxford, Denys Wilkinson Building, Keble Road, Oxford, UK

This talk:

- Is the CMB dipole really 'purely kinematic'? Dipoles in number counts of flux limited catalogues:
 - High redshift Radio Galaxies (NVSS + SUMSS)
 - Low redshift infrared galaxies (AllWISE)
 - High Redshift Quasars (CatWISE)
 - Gaia UnWISE

MNRAS 471 (2017) no.1, 1045-1055 MNRAS 477 (2018) no.2, 1772-1781 arXiv: 2009.14826 in preparation

The situation that Ellis & Baldwin anticipated in 1984 has arrived.

- The bulk flow of the local Universe. Where is the cosmic rest frame?
- The tilted Friedmann Universe.
 - "Evidence for anisotropy of Cosmic Acceleration" : An amusing debate:

The issue of peculiar velocities and corrections.

- The Hubble tension makes no sense
- What exactly is going on in cosmology now.
- Backup

A historical review of Supernova cosmology and fitting.

A&A 631, L13 (2019) arXiv:1912.04257

arXiv: 1911.06456





Peculiar velocity impact on SN1a magnitude

-400

-600 10⁻³

 $1 + z = (1 + \bar{z})(1 + z_{pec}^{hel})(1 + z_{pec}^{SN})$ $d_L(z) = \bar{d}_L(\bar{z})(1 + z_{pec}^{hel})(1 + z_{pec}^{SN})^2$

Davis et. al. Astrophys.J. 741 (2011) 67

JLA (and Pantheon) redshifts and magnitudes have been 'corrected' to account for the local bulk flow.

<pre>#name zcmb zhel</pre>	dz mb dmb x1 dx1 color dcolor
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03D1fc 0.330932	0.332000 0 21.861412 0.086437
03D1fg 0_798566	0 800000 0 24 510389 0 101777

 $C = [(1 + z_{hel}) - (1 + z_{cmb})(1 + z_d)] \times c$ 1200 1000 800 600 700

SN1a at z>0.06 are assumed (arbitrarily) to be in the CMB rest frame. (only uncorrelated 150 km/s in error budget)

10⁻¹

10⁰

 10^{1}

Flow model – SMAC has a ~600 km/s residual bulk flow

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<pre>#name zcmb zhe</pre>	l dz mb dmb	x1 dx1 color	dcolor
03D1au 0.50308	4 0.504300 0	0 23.001698 0	.088031
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03D1ax 0.49479	5 0.496000 0	0 22.960139 0	.088110
03D1bp 0.34592	8 0.347000	0 22.398137 0	.087263
03D1co 0.67766	2 0.679000	0 24.078115 0	.098356
03D1dt 0.61071	2 0.612000	0 23.285241 0	.092877
03D1ew 0.86649	4 0.868000	0 24.353678 0	106037
03D1fc 0.33093	2 0.332000	0 21.861412 0	.086437
03D1fg 0_79856	6 0-800000 0	0 24 510389 0	101777

 $z_{hel} \rightarrow measured$ $z_{cmb} \rightarrow inferred using a flow model$

$$C = [(1 + z_{hel}) - (1 + z_{cmb})(1 + z_d)] \times c$$



SN1a at z>0.06 are assumed (arbitrarily) to be in the CMB rest frame. (only uncorrelated 150 km/s in error budget) Wrong 'correction' to SDSS2308 in JLA. Many such mistakes in Pantheon (eg : SN2246). Flow model – SMAC has a ~600 km/s residual bulk flow

Consequently, we use only z_{hel} and subtract out the corrections to m_B ³⁶

There is an arbitrary discontinuity within the data.

Also in the subsequent Pantheon compilation



https://github.com/dscolnic/Pantheon/issues/2

This is because in the absence of demonstrable convergence between the bulk flow of the local Universe and the 'CMB rest frame', there is no way to correct for it completely (one could fit it as a nuisance parameter). Key Hubble tension papers rely on these corrections or directly on the Pantheon compilation (for eg Kenworthy et al 2019)





What we mean by 'non Copernican observers' $R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$

The FLRW universe

The Real Universe



Can be described by one scale factor a(t) and Friedmann equations exactly.

 $\Omega_M + \Omega_K + \Omega_\Lambda = 1$ The cosmic sum rule $\dot{\Theta} = -\frac{\theta^2}{3} - 2\sigma^2 + 2\omega^2 - E\left[\vec{X}\right]_a^a + \dot{X}_{;a}^a + \Lambda$

Ellis, "On the Raychaudhury Equation" Pramana–J.Phys.,Vol. 69, No. 1, July 2007

Everything has a peculiar velocity of $\sim 10^{-3}$, they should be Maximal symmetry forbids peculiar velocities viewed as differences in the expansion rate of the Universe

Some existing debates in literature (inhomogeneous cosmology/backreactions) suggest that problems such as Dark Matter and Dark Energy can also be tackled be critically examining the tools and framework with which we do cosmology.

What we mean by 'non Copernican observers' $R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$

Table 1: Comparison of curvature properties within the FLRW class of cosmological models and for generic averaged globally hyperbolic spacetime models. Buchert and Heinesen 2020

		FLRW	Average within generic GR	and the second
	Topology	$\operatorname{sign}(\mathcal{R})$ determines the spatial topol-	$\langle \mathcal{R} \rangle_{\mathcal{D}}$ does not in general allow con-	
		ogy for simply-connected domains	clusions on topological properties	
	Integral constraint	local 'Newtonian' energy conserva-	general-relativistic coupling of $\langle \mathcal{R} \rangle_{\mathcal{D}}$	
		tion: $(\mathcal{R}a^2)^{\cdot} = 0$	to structure:	
Can be de Friedman			$\frac{1}{a_{\mathcal{D}}^6} \left(\mathcal{Q}_{\mathcal{D}} \ a_{\mathcal{D}}^6 \right)^{\cdot} + \frac{1}{a_{\mathcal{D}}^2} \left(\left\langle \mathcal{R} \right\rangle_{\mathcal{D}} a_{\mathcal{D}}^2 \right)^{\cdot} = 0$	$X^a_{;a} + \Lambda$
	Ciana a franca tama	(\mathcal{D}) is supported the set the	(\mathcal{P})	nury Equation"
	Sign of curvature	$\operatorname{sign}(\mathcal{K})$ is preserved throughout the	$\operatorname{sign}(\langle \mathcal{K} \rangle_{\mathcal{D}})$ can change in response	o 1 July 2007
		evolution of the Universe and on all	to structure in the spacetime and	0. 1, July 2007
		scales	may vary on different scales	
	Copernican principle	satisfied in its most strict interpreta-	can be satisfied in a weaker sense	
Maximal		tion. All fundamental observers are	than for FLRW. 'Distributional	Jniverse
		subject to the same local curvature	equivalence' between observers	ms

There is no Hubble constant, let alone a tension



of 9 km s₋₁ Mpc₋₁ is found to occur across the sky.

Migkas et al 2020

Conclusions

- Number counts of flux limited catalogues in radio and infrared all indicate somewhat significant (up to $\sim 3.9\sigma$) tensions with the 'purely kinematic' interpretation of the CMB dipole.
- Hopeful that SKA and EUCLID can set this to rest by testing.
 Convergence to the CMB rest frame has not been demonstrated.
 - There is a case for precision testing the CMB dipole.
 - The local Universe has a bulk flow out to ~400 Mpc. McClure and Dyer 2007 The CMB rest frame does not exist
- SN1a data pre ship with 'corrections' and are being continuously adjusted. The Hubble tension is manufactured using these corrections.
- Evidence 3.9 σ for a tilt in the local Universe. Isotropic acceleration compatible with 0 at < 1.4 sigma
- Since ΛCDM cosmology is dying, time to move to an anisotropic cosmology.