Based on: [arxiv 2203.11821]

A substandard Candle: The low-v Method

Rencontres du Vietnam: Neutrino 2022/07/22 Luke Pickering for the NUISANCE Collaboration





Motivation: Measuring the Neutrino Flux

$$N_{\text{near}}(E_{\text{obs}}) = \int dE_{\nu} \Phi_{\text{near}}(E_{\nu}) \cdot \sigma(E_{\nu}) \cdot \mathbf{D}_{\text{near}}$$
$$N_{\text{far}}(E_{\text{obs}}) = \int dE_{\nu} \Phi_{\text{far}}(E_{\nu}) \cdot P_{osc}(E_{\nu}) \cdot \sigma(E_{\nu}) \cdot \mathbf{D}_{\text{far}}$$
$$Can \ access \qquad \qquad Want \ to \ know$$

Extract oscillation probability from observed rate.

- Observed rate is convolution of flux, cross-section, and oscillations
- Current uncertainty on flux (~5-10%) and cross-section (~10%)
- Use near detector to constrain rate+shape = flux (X) cross-section
- Use constrained model to predict observation for a oscillation hypothesis to test



Motivation: Measuring the Neutrino Flux

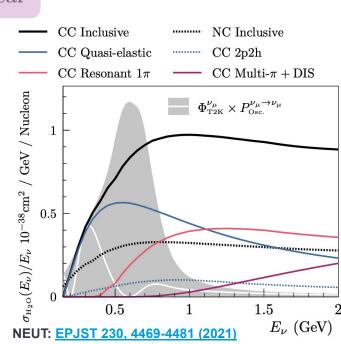
$$N_{\text{near}}\left(E_{\text{obs}}\right) = \int dE_{\nu} \Phi_{\text{near}}\left(E_{\nu}\right) \cdot \sigma\left(E_{\nu}\right) \cdot \mathbf{D}_{\text{near}}$$

Ideally would use near detector to constrain flux and cross-section separately

in situ measurements are difficult:

- Wide band beam
- Cross-sections changing rapidly
- Detector effects cause significant smearing

Have to get creative...





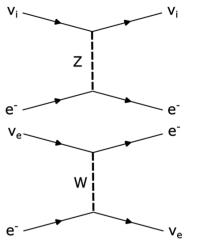
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Neutrino-electron elastic scattering:

- Simple interaction with no nuclear effects
- Rate is OOM lower than neutrino-nucleon scattering
- Future beams will produce sizeable sample
- Good for flux normalization

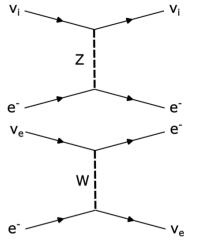




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Can try to isolate neutrino–nucleon events with a well-known cross-section:

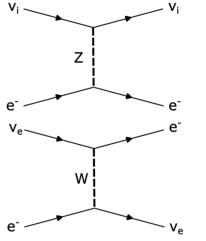
- Select very-elastic events
- \rightarrow Cross-section saturates at some Ev
- → Almost all energy to visible final-state charged lepton
- Minimal energy lost to nuclear response/hadron production



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PHYSICAL REVIEW D

VOLUME 38, NUMBER 9

1 NOVEMBER 1988

Neutrino reactions in the low-y region

R. Belusevic

Department of Physics, University College, University of London, London WC1E 6BT, United Kingdom

D. Rein

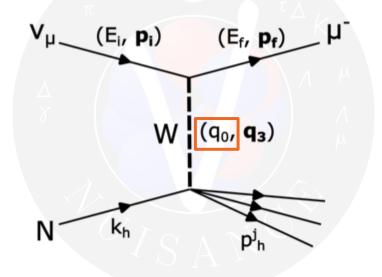
III. Physikalisches Institut, Rheinisch-Westfälische Technische Hochschule Aachen, D-5100 Aachen, Federal Republic of Germany (Received 10 June 1987; revised manuscript received 30 June 1988)

> The physics of nonscaling components in the region of low energy transfer (low-y region) is described. The following neutrino-induced processes were considered: resonance production, quasielastic scattering, and coherent meson production off nuclei or nuclear fragments. It is shown that the total exclusive cross section in a certain kinematical domain is energy independent at high energies (above 20 GeV). This fact can, in principle, be used for relative normalization of the neutrino flux.

> > Phys. Rev. D 38, 2753 (1988)



A Problem in Terminology

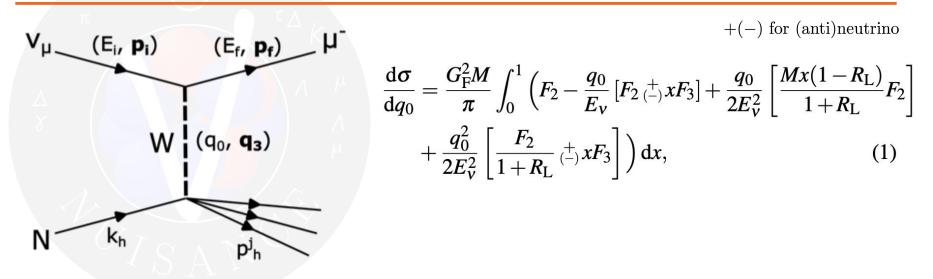


- Originally the "low-y" region
 - \circ Low inelasticity \rightarrow High elasticity
- "low-v" coined in terms of DIS formalism
- *v* is ambiguous in neutrino-scattering physics
- Has also been called "fixed v_0 " method
- Electron-scattering folks often use $\omega = E_i E_f$

• I will try to use q_o as in the oth component of the 4-momentum transfer q-vector.



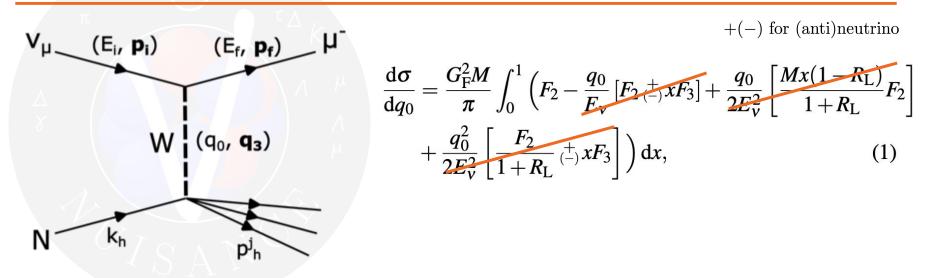
The "low-v" Method



- Write down inclusive DIS cross-section in terms of nucleon structure functions
- Notice that if $\frac{q_0}{E_{\nu}} << 1$ any neutrino energy dependence disappears



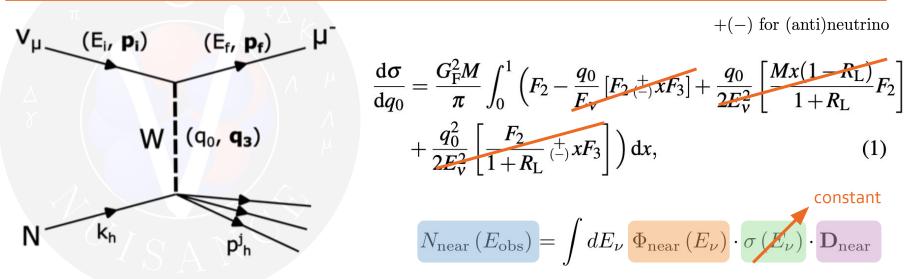
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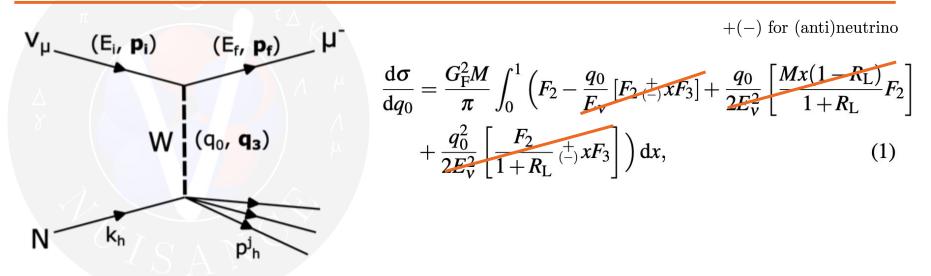
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The "low-v" Method



- Write down inclusive DIS cross-section in terms of nucleon structure functions
- Notice that if $\frac{q_0}{E_{\mu}} << 1$ any neutrino energy dependence disappears
- A flat cross-section can be used to access the shape of the flux from the event rate!



- 1. Eq. (1) describes the cross-section well in the region of interest
- 2. A sample with 'low enough' q can be unambiguously experimentally selected
- 3. The neutrino energy for events in the sample can be accurately reconstructed.



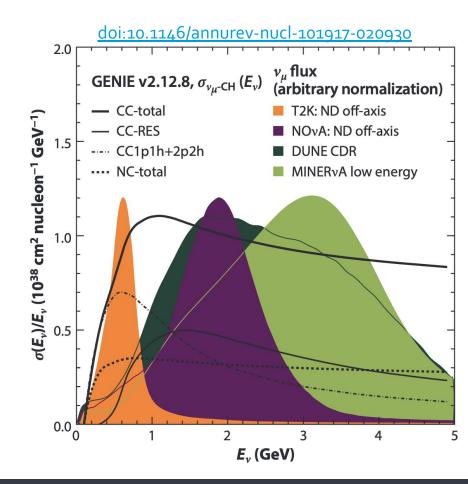
Utility for Near-future Oscillation Measurements



A substandard Candle: The Low-v Method

Experimental Fluxes

Relevant neutrino energy range: < 5 GeV





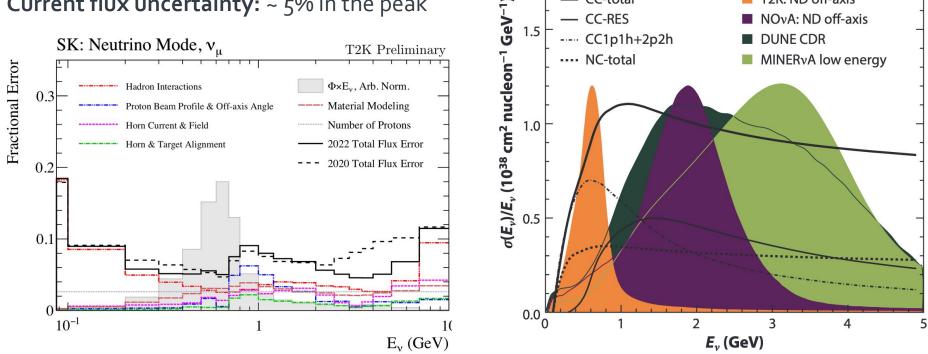
A substandard Candle: The Low-v Method



Experimental Fluxes

ROYAL HOLLOWAY

Relevant neutrino energy range: < 5GeV **Current flux uncertainty:** ~ 5% in the peak



2.0

 v_{μ} flux (arbitrary normalization)

T2K: ND off-axis

doi:10.1146/annurev-nucl-101917-020930

GENIE v2.12.8, $\sigma_{\nu_{\mu}}$ -CH (E_{ν})

- CC-total

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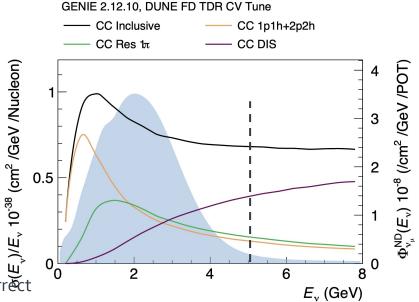
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- By 5 GeV the charged-current cross-section is *mostly* DIS.
- But we need $q_0/Ev \ll 1$
 - **DIS** formalism, but a selection of very-not **DIS** events...
- At what (q_0, Ev) does Eq. (1) break down?
 - Recent implementations ask the interaction model to correct for any non-constant behavior [EPJC 72, 1973 (2012)]

- 1. Eq. (1) describes the cross-section well in the region of interest
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Model Motivations

Need to investigate how well such a correction can be known with current simulation. Use a spread of reasonable model predictions:

- Not a replacement for well motivated theoretical uncertainties...
- ³ But they don't exist, so we have to make do with *semi*-sensible choices
- GENIE:
 - Version 2 Used by MINERvA in their Low-v constraint [Phys. Rev. D 94, 112007]
 - Version 3 Currently used by NOvA, MicroBooNE for their analyses
- **NEUT**: Used by T₂K
- **NuWro**: Performs well in comparisons to world cross-section data
- **GiBUU**: Sophisticated hadron-transport and quite different neutrino–nucleon model choices, also performs well in world data comparisons
- **SUSAv2 and CRPA**: State-of-the-art nuclear response modelling for pionless events

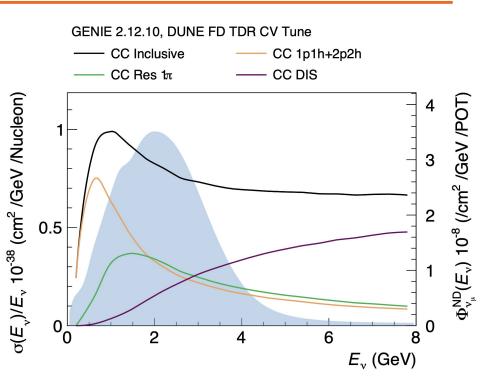


A View From The DUNEs

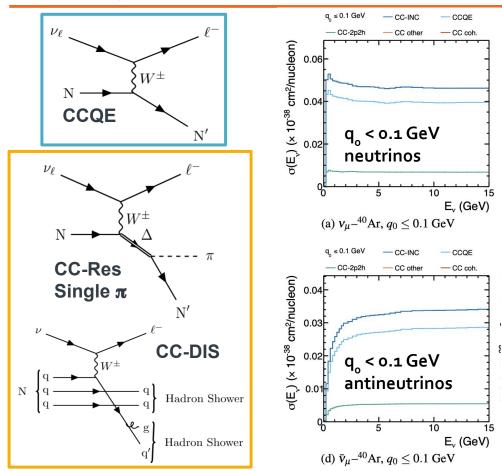
Study Configuration:

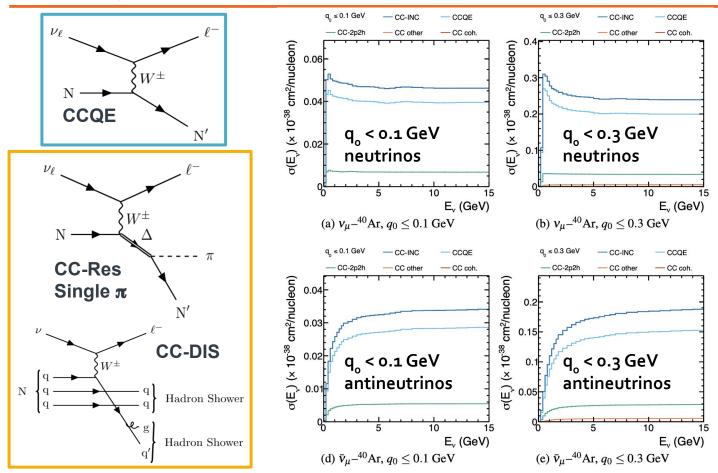
- Ar4o target
- Neutrinos and anti-neutrinos
- Spread of *reasonable* model choices

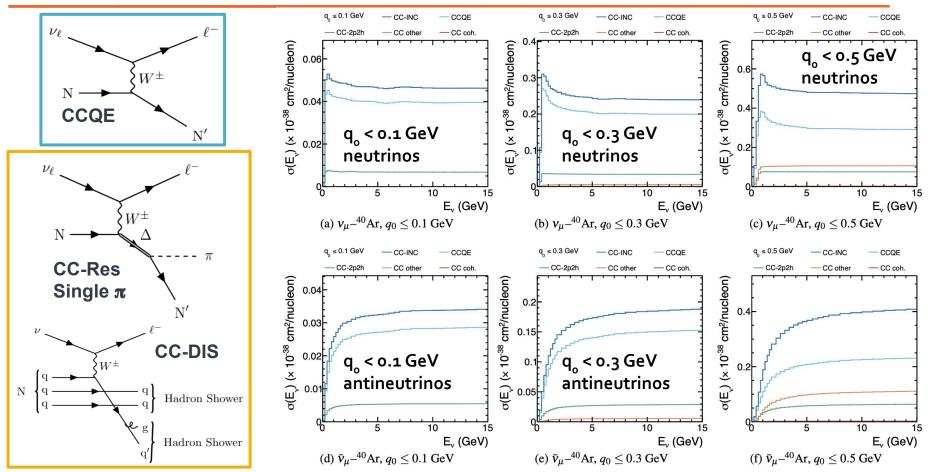
How well-understood is a perfect 'low-q_o' sample of events for a DUNE-like target?







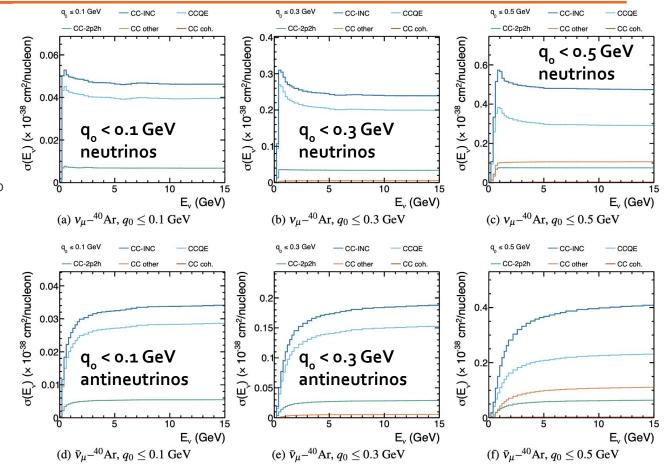




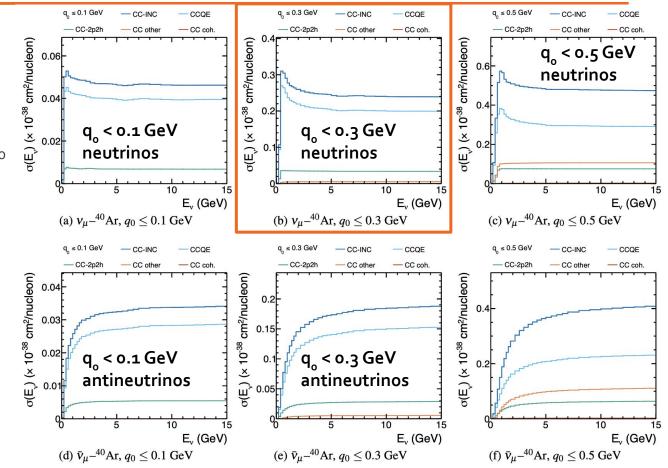
GENIE3: Total Cross Section

Observations:

- Cross-section becomes ~flat for neutrinos at reasonable values of q
- It doesn't become flat for antineutrinos except for q_o<0.1 GeV

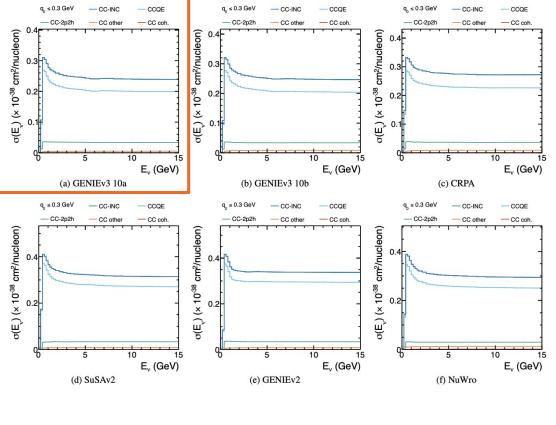


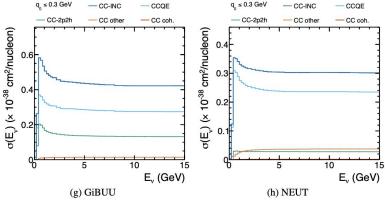
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Alternate Predictions

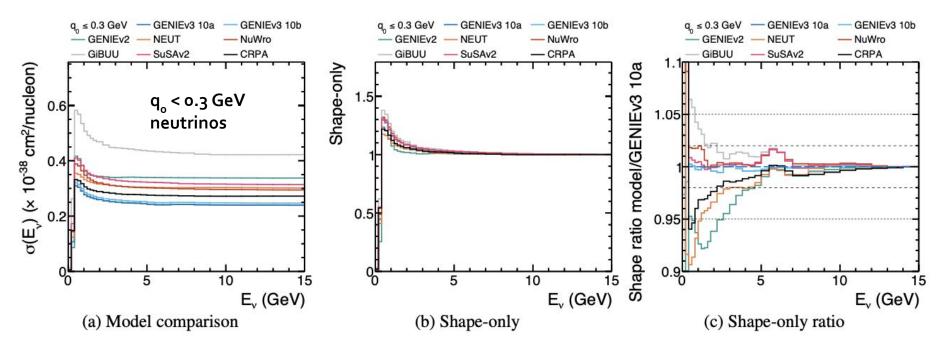
- General trends between predictions are similar by eye.
- GiBUU has very different 2p2h strength due to a model choice made for strength of correlated NN pairs.





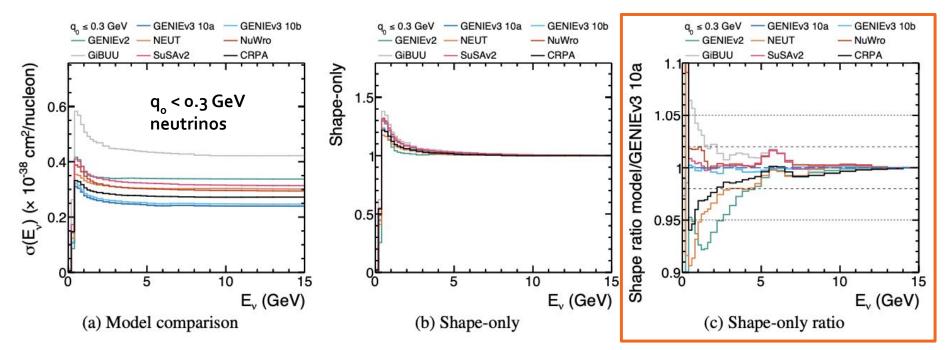
Comparing the Predictions

- Want to evaluate uncertainty on the relative constantness of the low-q cross-section
 - Take GENIEv₃ 10a as the reference model and examine variation of predicted low-q_o shapes
- For illustration have chosen $q_0 < 0.3$ GeV



Comparing the Predictions

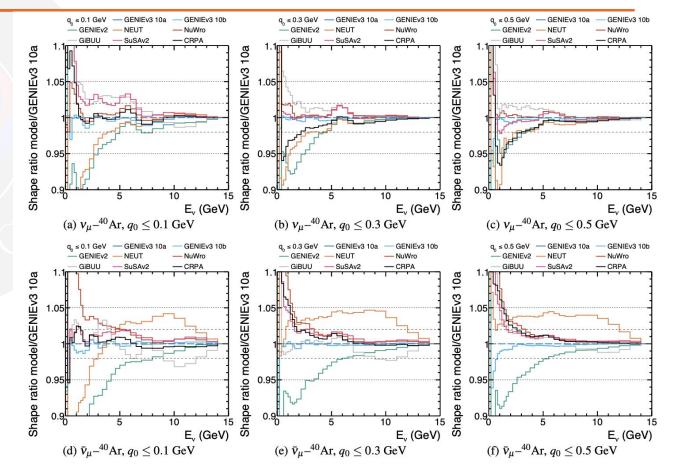
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Model Spread

Important LBL oscillation region 0.5–5 GeV:

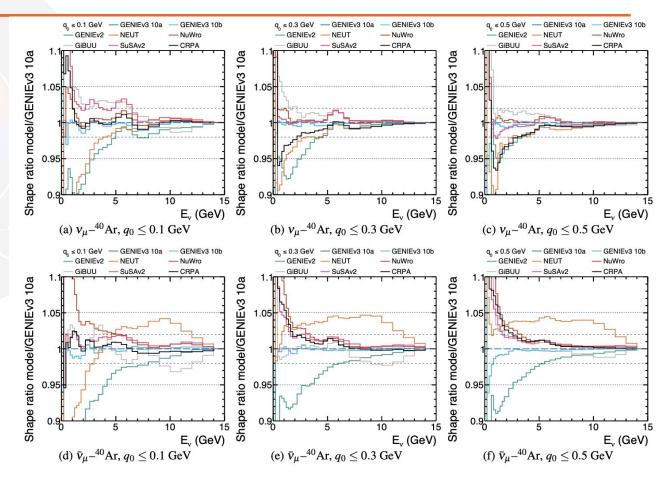
- Spread of 2–5% seen for neutrinos
- A bit larger for antineutrinos

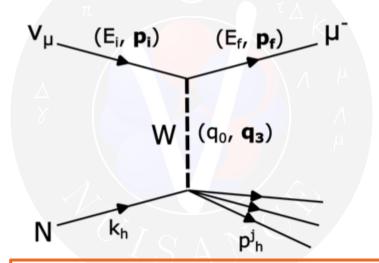


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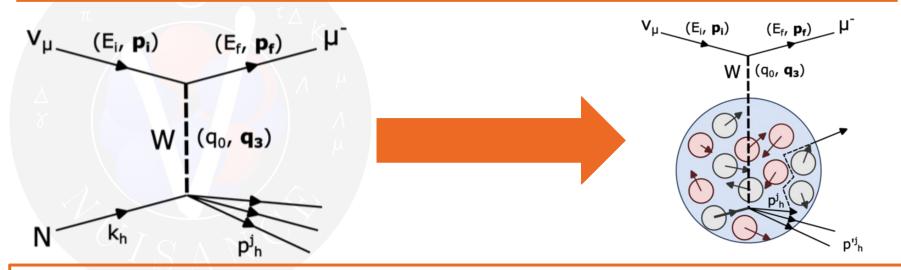
- Spread of 2–5% seen for neutrinos
- A bit larger for antineutrinos
- Counterintuitively the spread reduces for higher q_o cuts:
 - Opposite behavior than predicted by Eq. (1)
 - Likely due to nuclear effects that do not fit well into Eq. (1)
 - Pion-production and DIS models are more similar between generators





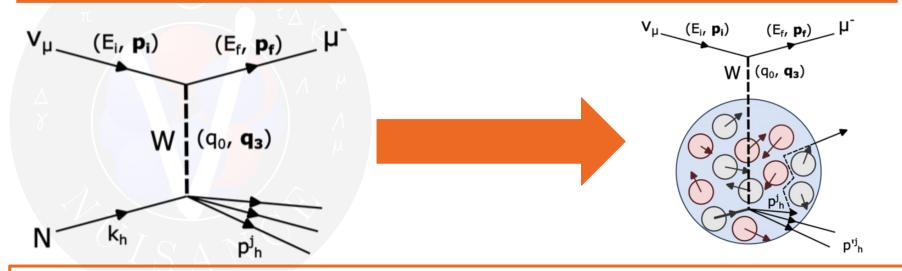
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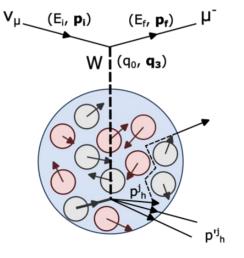


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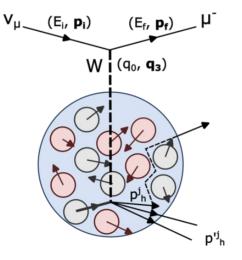
Selecting a Low-q Sample

- Cannot access true q_o because we cannot know incoming neutrino energy
 - Reconstruct from the hadronic side



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- Introduces significant dependence on nuclear response and detection capabilities:
 - / Missed-pion masses
 - Missed neutrons



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 - / Missed-pion masses
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- Define three q_o proxy variables
 - Perfect detector, measures all hadrons
 - Optimistic detector, can't see neutrons but tags every pion
 - Pessimistic detector, misses all neutrons and charged-pion masses

nnot know incoming
$$V_{\mu}$$
 (E, P) (E, P) (F_{i}, P_{i}) μ^{-}
w (q₀, q₃)
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 $E_{had}^{true} = \left(\sum_{i=n,p} E_{kin}^{i}\right) + \left(\sum_{i=\pi^{\pm},\pi^{0},\gamma} E_{total}^{i}\right)$
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$$E_{\text{avail}} = \left(\sum_{i=\pi^{\pm},p} E_{\text{kin}}^{i}\right) + \left(\sum_{i=\pi^{0},\gamma} E_{\text{total}}^{i}\right)$$

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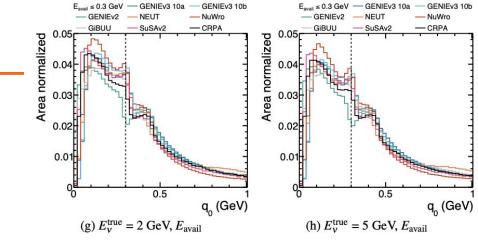
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E_{avail} Smearing

Shows the ranges of true qo values that pass E_{avail} < 0.3 cut for two neutrino energies.



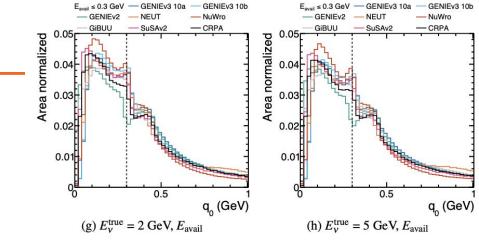


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- "feed down" predicted from higher q_o
- Significant model spread near the cut value



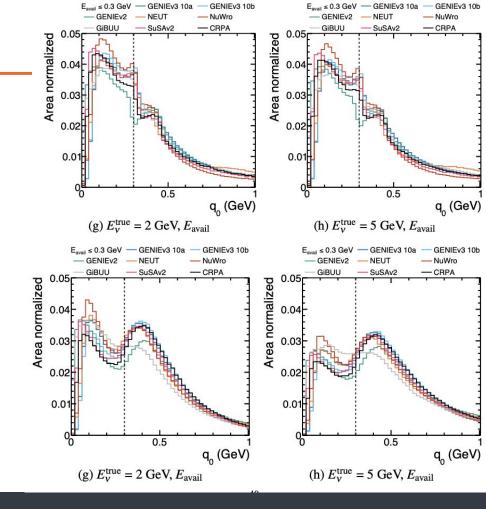


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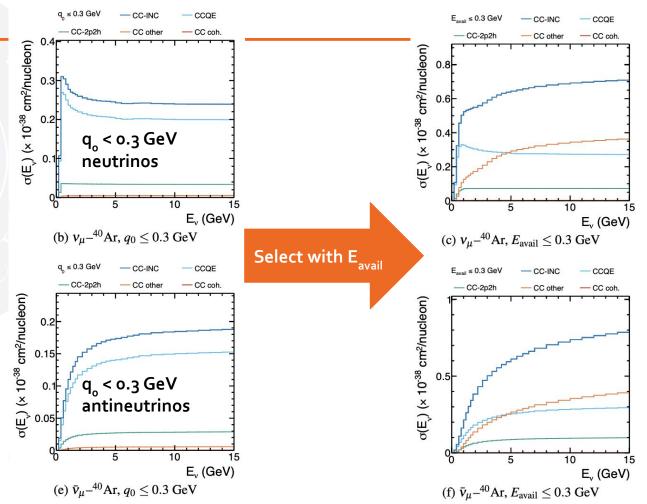
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- Significant model spread near the cut value
- Feed-down is worse for antineutrino







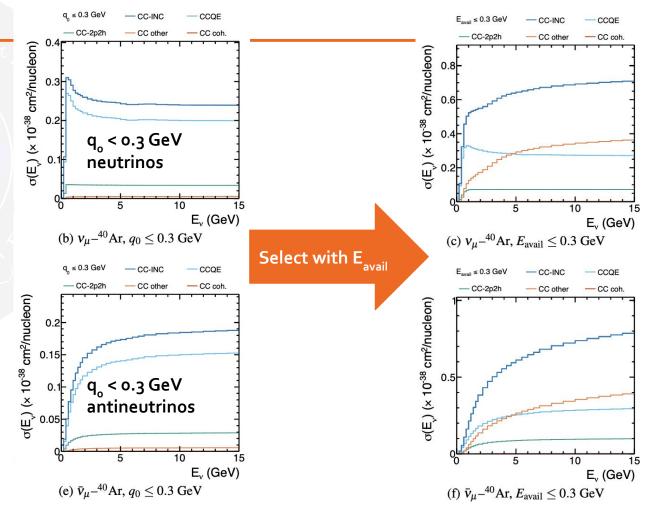
- much lower
- Cross-section is not predicted to become flat for neutrinos or antineutrinos





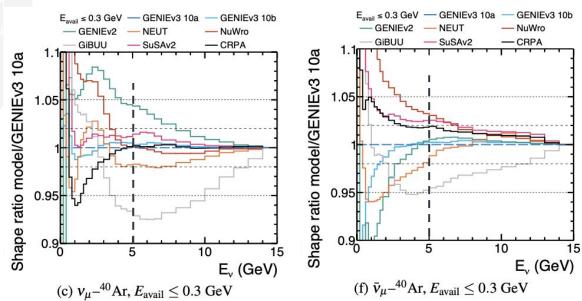
Observations:

- True QE fraction is much lower
- Cross-section is not predicted to become flat for neutrinos or antineutrinos
- → Rely on model-dependent corrections to use Low-v



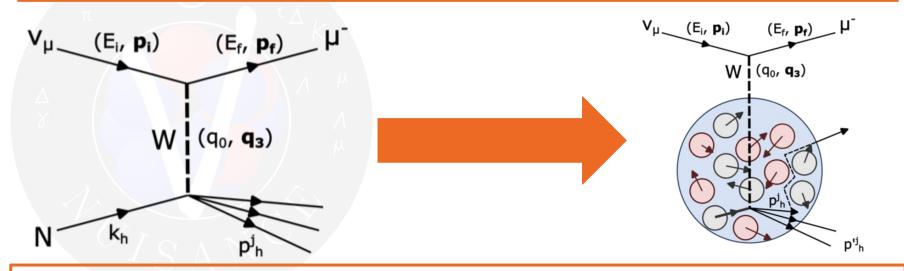
A More Realistic Energy-transfer Analysis

- In the important region, we see model-spread of 5%
 - Performs similarly to current indirect flux constraints
 - Hadron-production measurements are expected to continue to improve before DUNE/HK are systematics limited
- This is not a motivation of systematic uncertainty:
 - It is almost certainly an underestimate of the uncertainty
 - But it is a more comprehensive study of available predictions than has been published by any Low-v at few GeV Ev proponents





The Implicit Assumptions



The "low-v" method works well if:

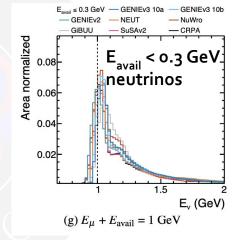
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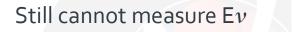
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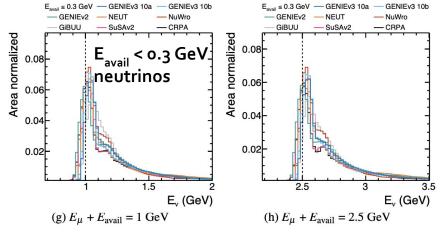
Neutrino Energy Smearing

Still cannot measure Ev



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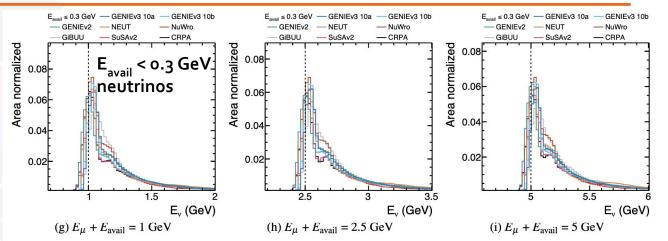




Neutrino Energy Smearing

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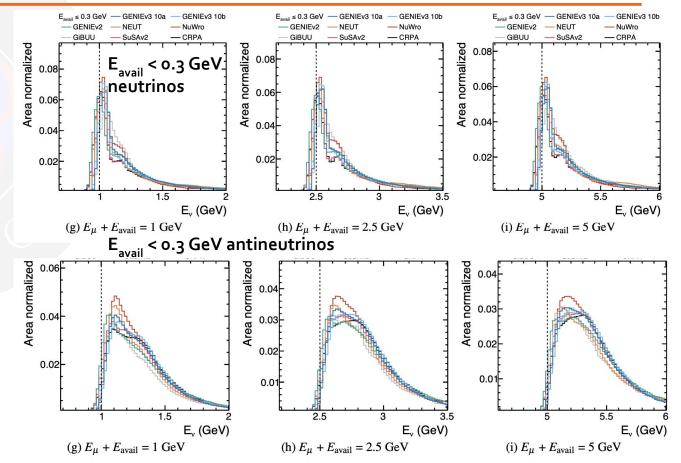
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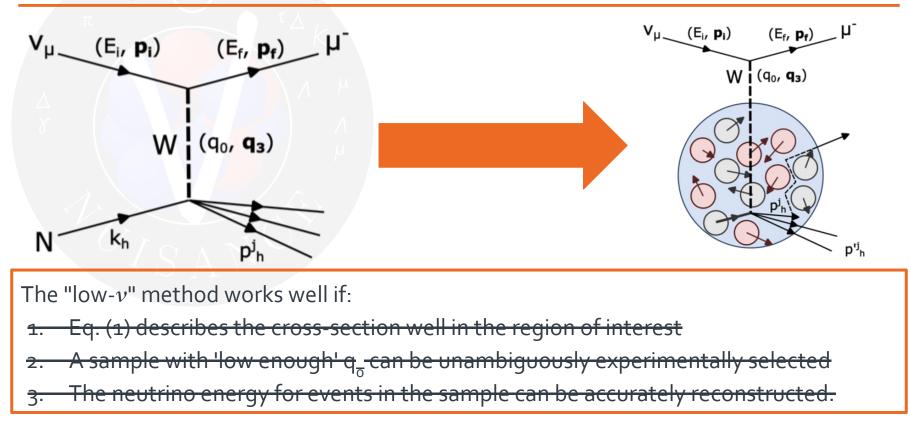
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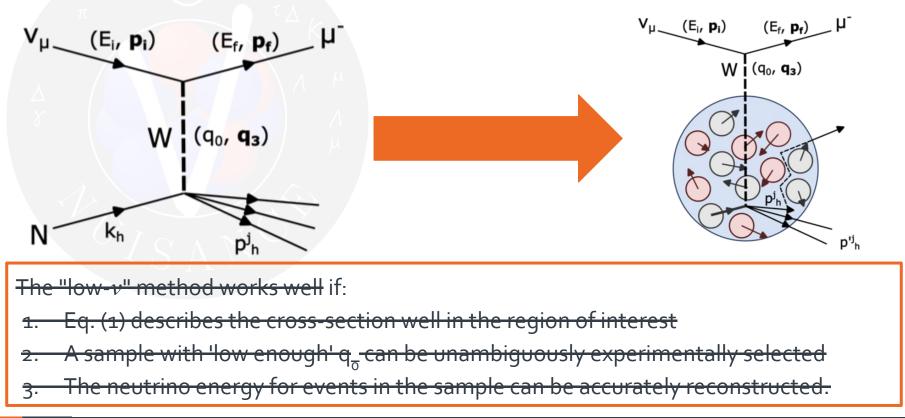
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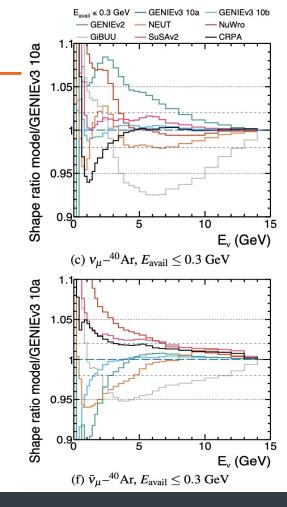




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Conclusions

 Far from relying on well-understood cross-section to detangle flux cross-section the "Low-v" method at low Ev is strongly neutrino interaction and nuclear response model dependent

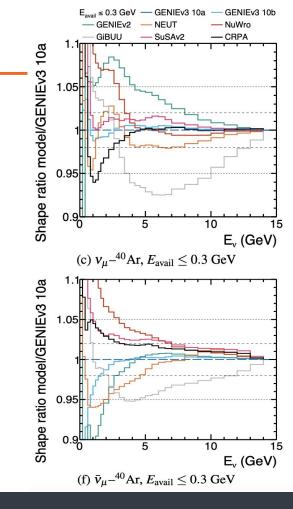






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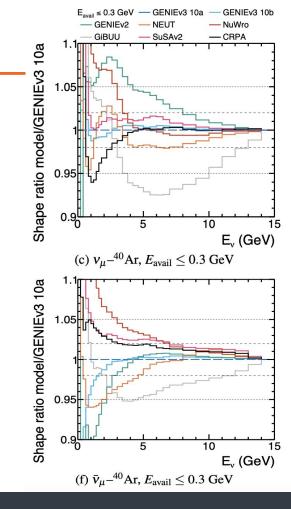
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- Even in simple MC truth studies, it is difficult to justify relying on "Low-v" for flux constraints below ~5 GeV for nu or 15 GeV for nubar
 - A real detector is likely to be even more messy.





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- Even in simple MC truth studies, it is difficult to justify relying on "Low-v" for flux constraints below ~5 GeV for nu or 15 GeV for nubar
 - A real detector is likely to be even more messy.
- Hadron-production measurements and other *in situ* techniques should be preferred as the main source of flux constraints





Thanks For Listening





Supported by URF\R1\211661



The Low-y Success Stories

- CCFR
- NuTeV
- ~NOMAD: Cut EHad < 20 GeV and Enu > 30 GeV

