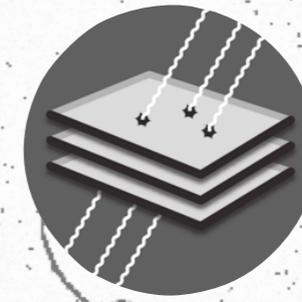


REAL DAMIC CCD IMAGE (ON SURFACE)



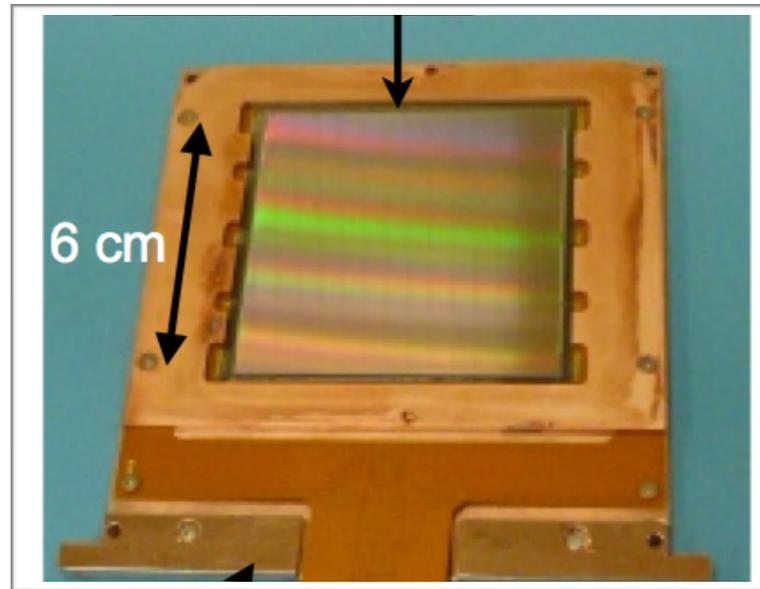
# Dark matter search in CCDs with DAMIC

Mariangela Settimo for the DAMIC and DAMIC-M collaborations  
Subatech, CNRS-IN2P3, Nantes (France)



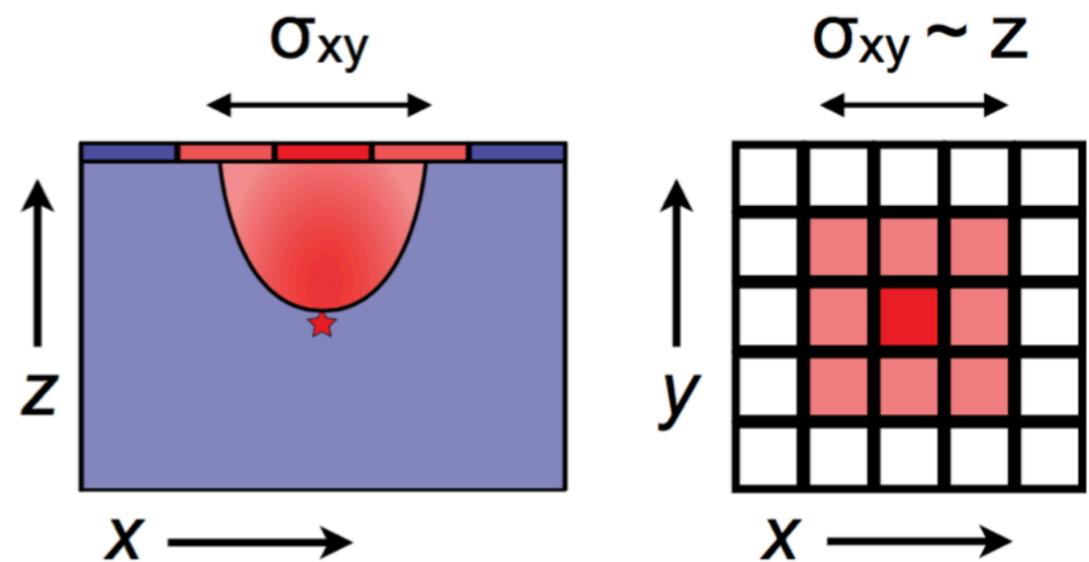
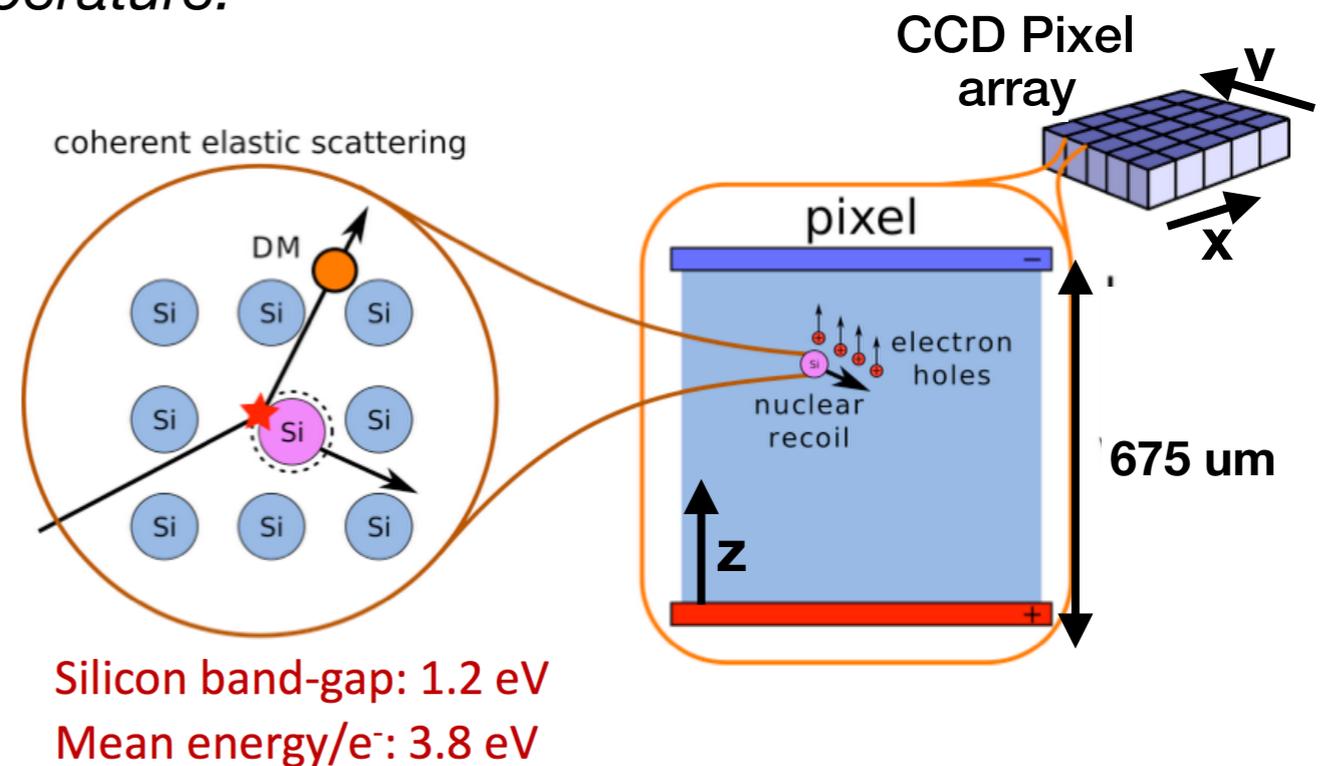
# Charge Coupled Devices (CCD) for dark matter

Detection of energy deposits from nuclear and electron recoils induced by DM in the Silicon bulk of CCDs, operating at low temperature.



**Unconventional use of CCDs :**  
 Fully depleted, high-resistivity  
 16Mpix, 15  $\mu\text{m}$  x15  $\mu\text{m}$ ,  
 650  $\mu\text{m}$  thick, 5.9 g mass

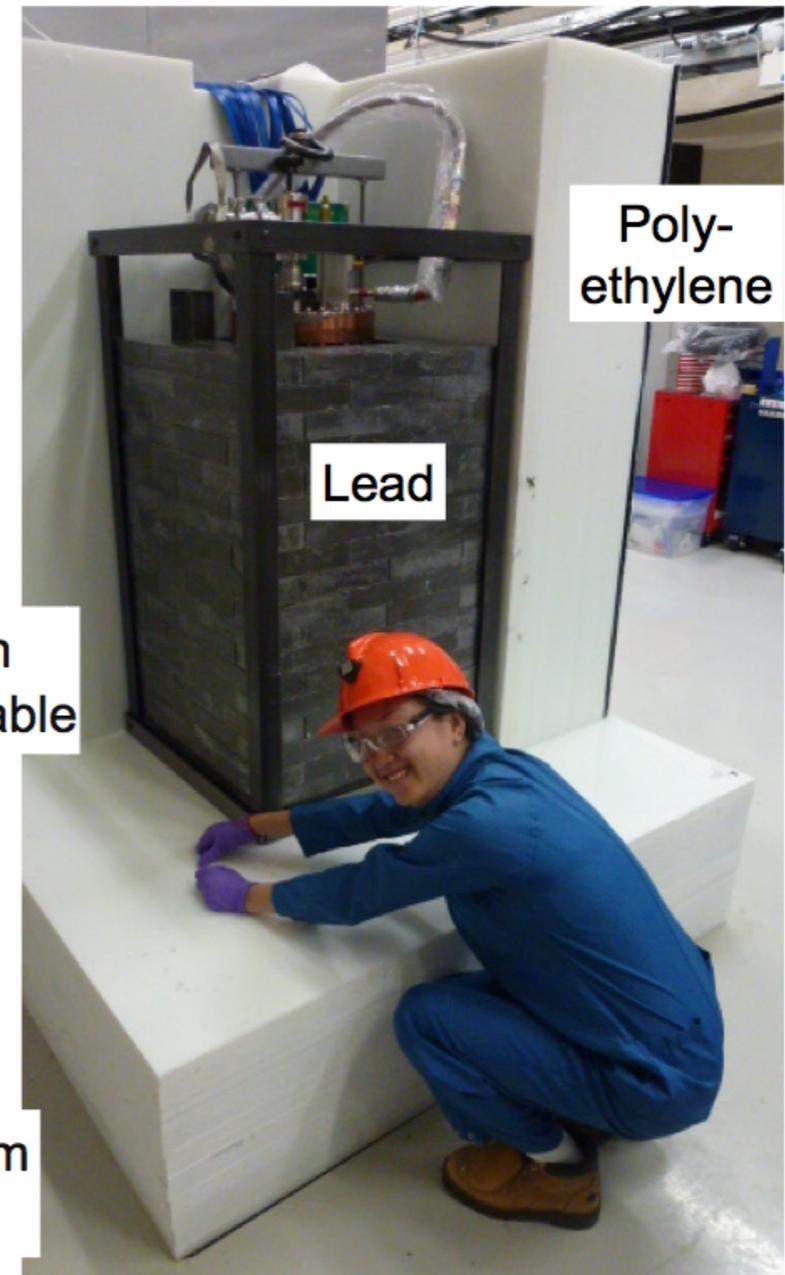
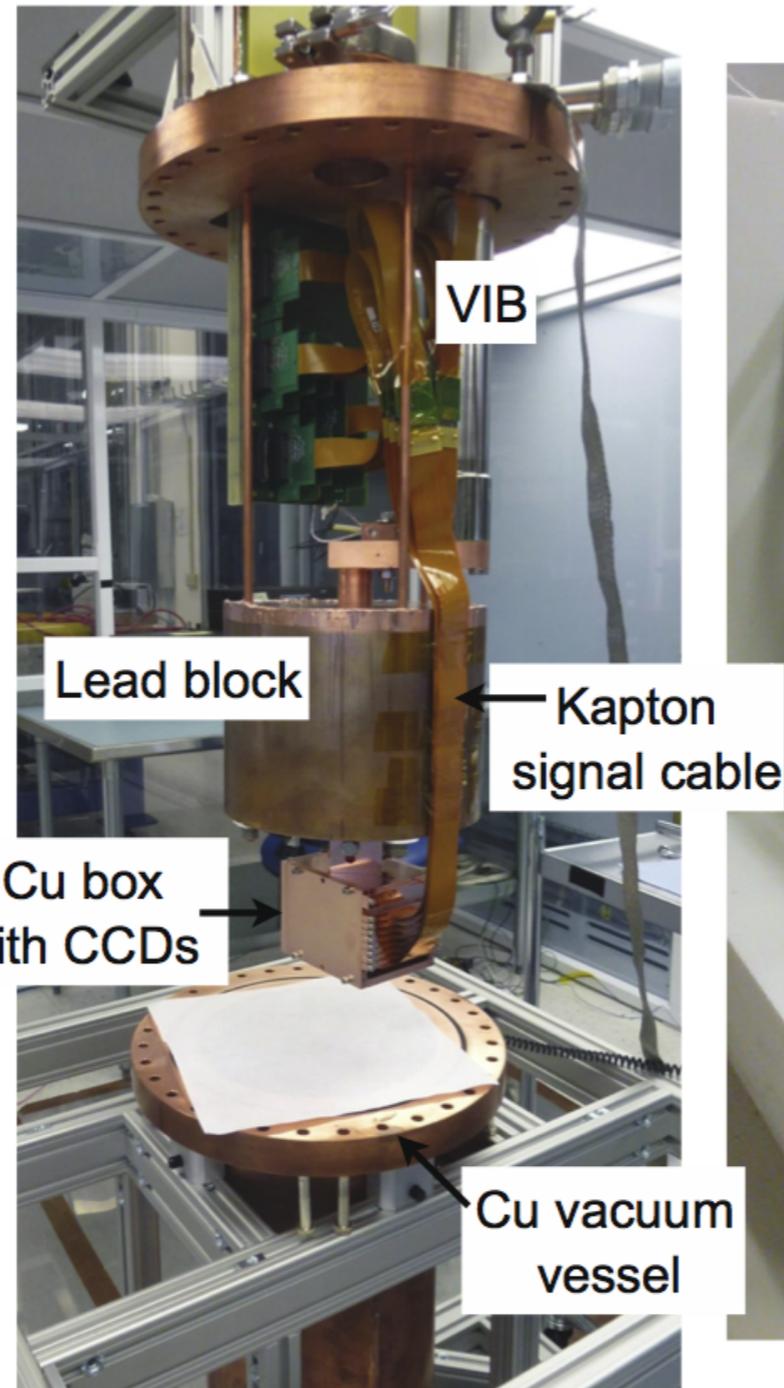
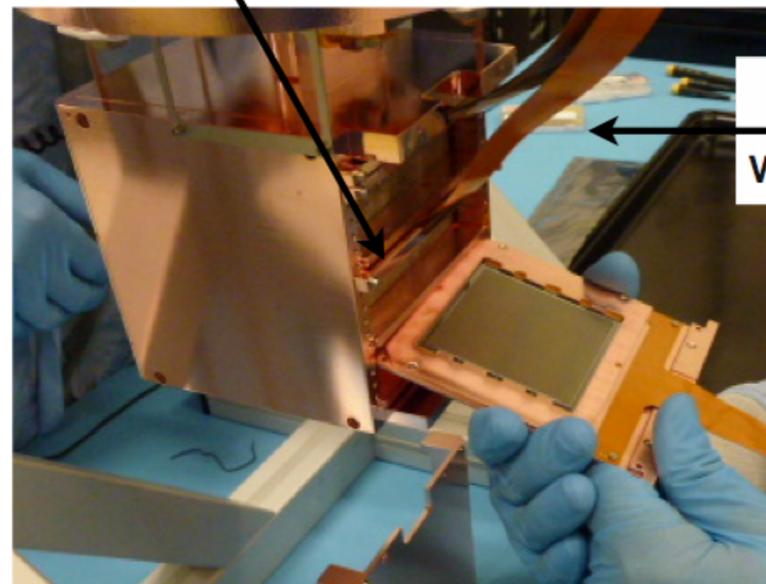
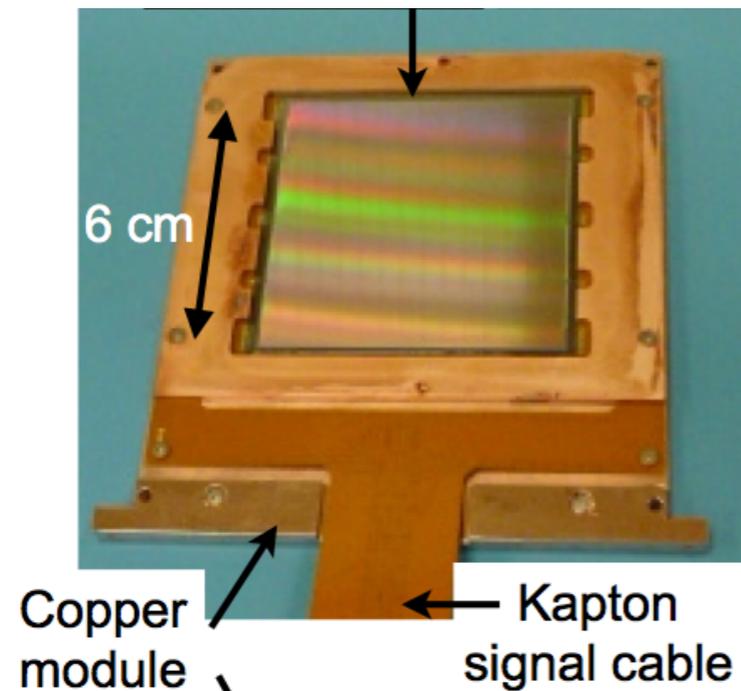
Sensitivity to  $< 10$  GeV WIMP masses  
 (recoils  $\sim$  keV), eV hidden photons and  
 DM electron-scattering



3D reconstruction (x, y, z)

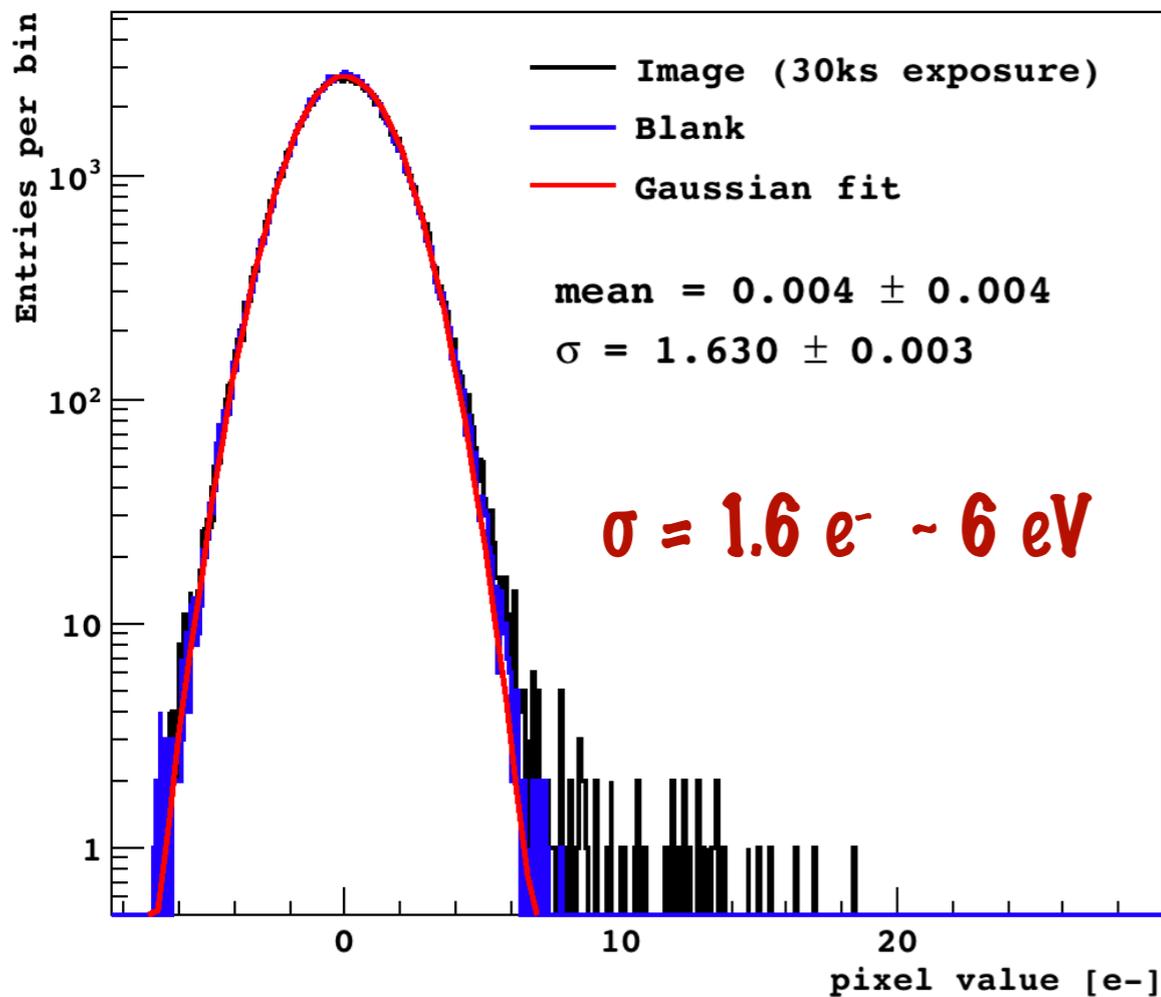
# DAMIC @ SNOLAB (2000 m underground lab)

675  $\mu\text{m}$  thick, 16 Mpix CCD, 6 g



R&D program since 2011,  
40g demonstrator operating at SNOLAB since 2017

# Low noise and low dark current



lowest dark current ever measured in a Si detector :

**DC < 0.001 e/pix/day @ 140K**

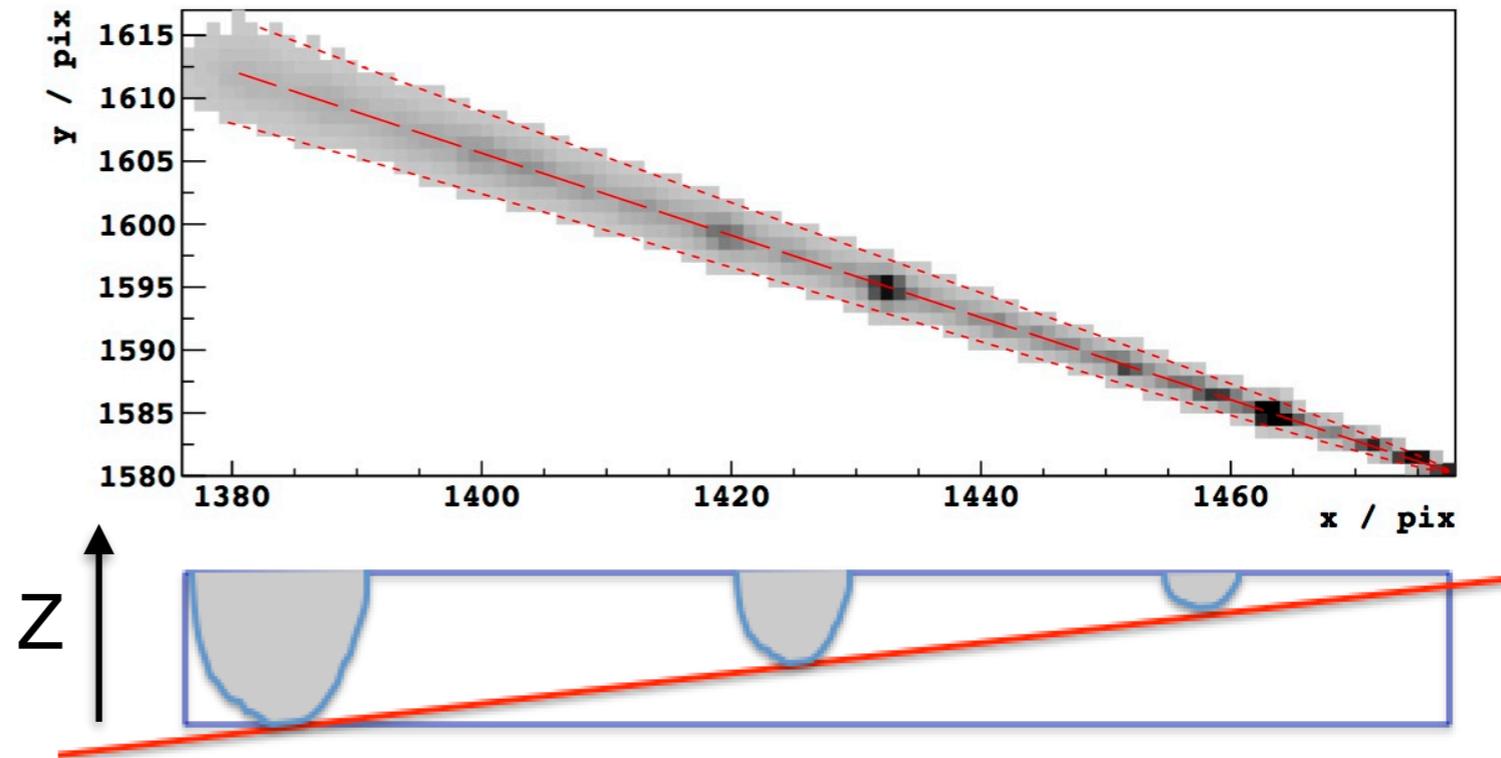
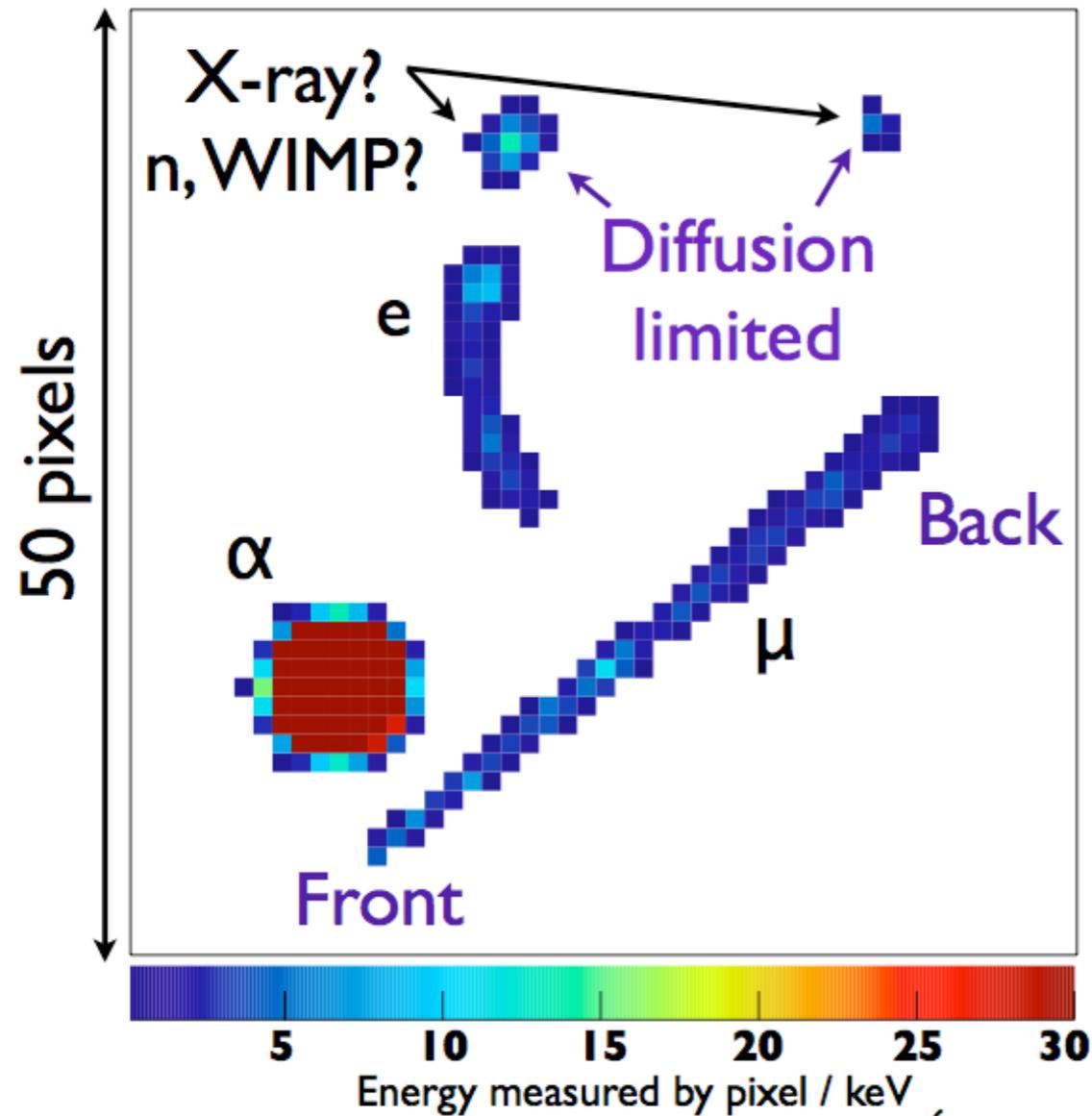
**Dominant noise from Readout ~ 1.6 e<sup>-</sup> (~ 6 eV)**



**Energy threshold ~ 40 eVee**

**Long exposures (30 ks and 100ks)**  
to optimise signal/noise

# Spatial resolution and particle identification



Diffusion parameters using events at the back of the CCD and muons passing through.

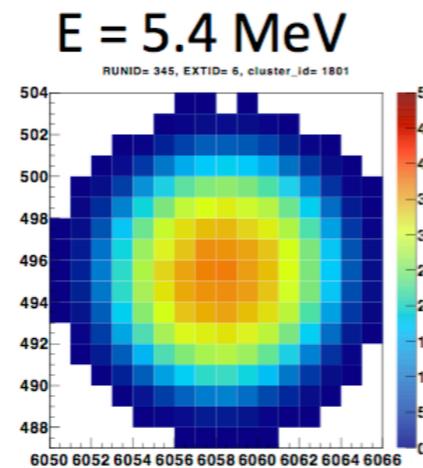
$(x,y)$  : crucial for particle and radioactive chains identification

$\sigma \approx z$  : provides information on the depth of the energy deposit  
—> reject surface background events

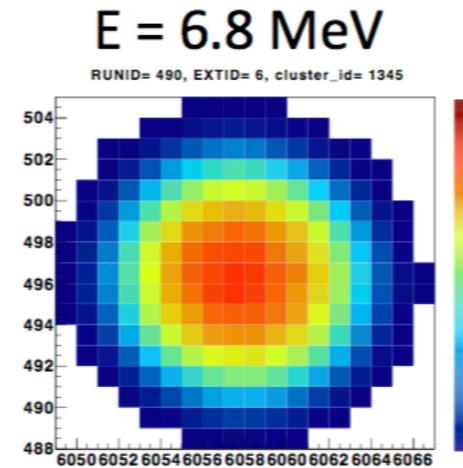
# Background characterization

**Spatial coincidence** of betas and/or alpha decays powerful tool to identify and reject surface and bulk background from radioactive decay.

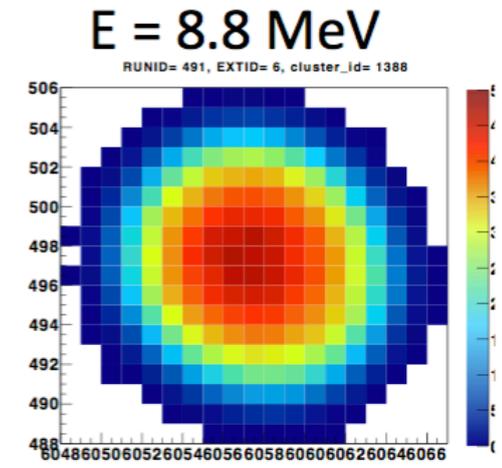
- **Surface background :**  
 $^{210}\text{Pb}$ , U/Th chain



1  $\Delta t = 17.8 \text{ d}$



2  $\Delta t = 5.5 \text{ h}$

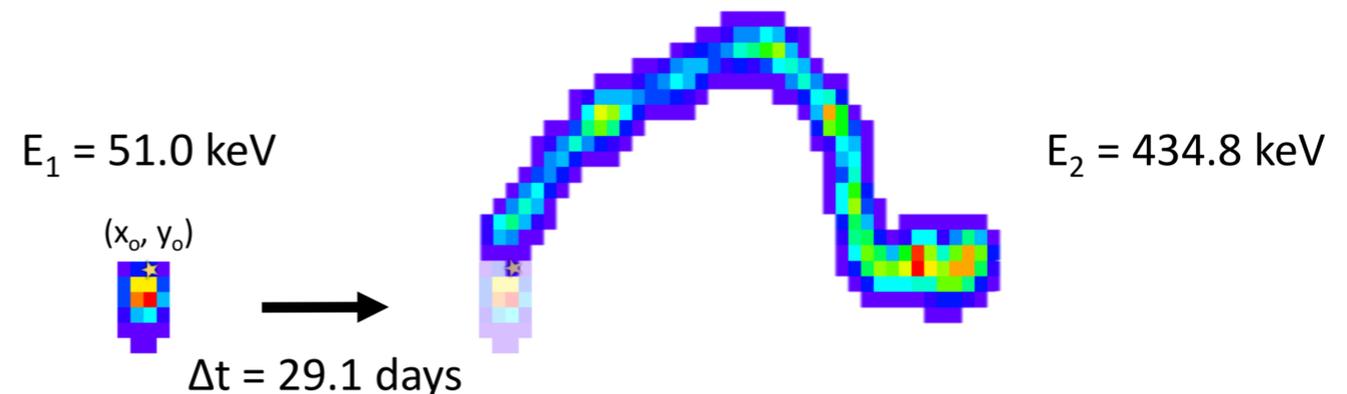
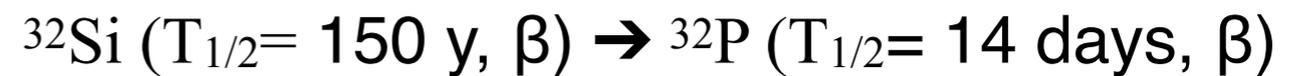


3

Three  $\alpha$  at the same location!

- **Bulk background :** Cosmogenic  $^{32}\text{Si}$   
rate measured in R&D  $< 80/\text{kg}/\text{day}$

UNAVOIDABLE BACKGROUND:  
essential measurement for next  
generation Si-detectors



# Surface and background contamination

Measured contamination of  $^{210}\text{Pb}$  and  $^{32}\text{Si}$   
limits are placed on  $^{238}\text{U}$  and  $^{232}\text{Th}$

## DAMIC 2019 analysis

$^{32}\text{Si}$

➤  $133.3 \pm 27.8 \mu\text{Bq/kg}$

$^{210}\text{Pb}$

➤  $83.1 \pm 11.8 \text{ nBq/cm}^2$

$^{238}\text{U}$

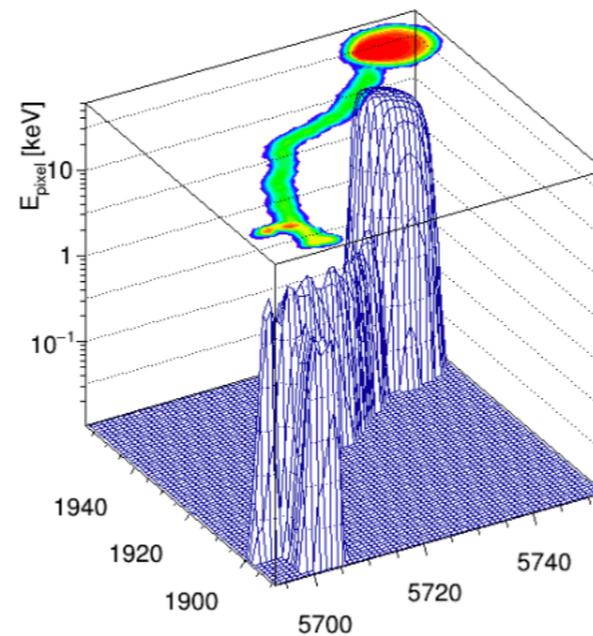
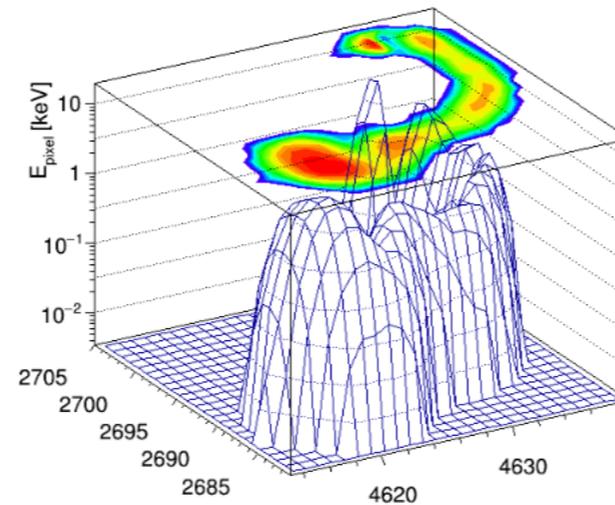
➤ No  $\alpha$ - $\beta$  sequences

➤ Upper limit:  
 $0.53/\text{kg/day}$  or  $1.5 \text{ ppt}$  [95%]

$^{232}\text{Th}$

➤ No  $\alpha$ 's with  $E = 18.7 \text{ MeV}$

➤ Upper limit:  
 $0.35/\text{kg/day}$  or  $1 \text{ ppt}$  [95%]



## DAMIC 2015 R&D result

$^{32}\text{Si}$

➤  $925.9 \pm 1273 / 752 \mu\text{Bq/kg}$

$^{210}\text{Pb}$

➤  $902.8 \pm 115.8 \text{ nBq/cm}^2$

$^{238}\text{U}$

➤ Upper limit:  $5/\text{kg/day}$  [95%]

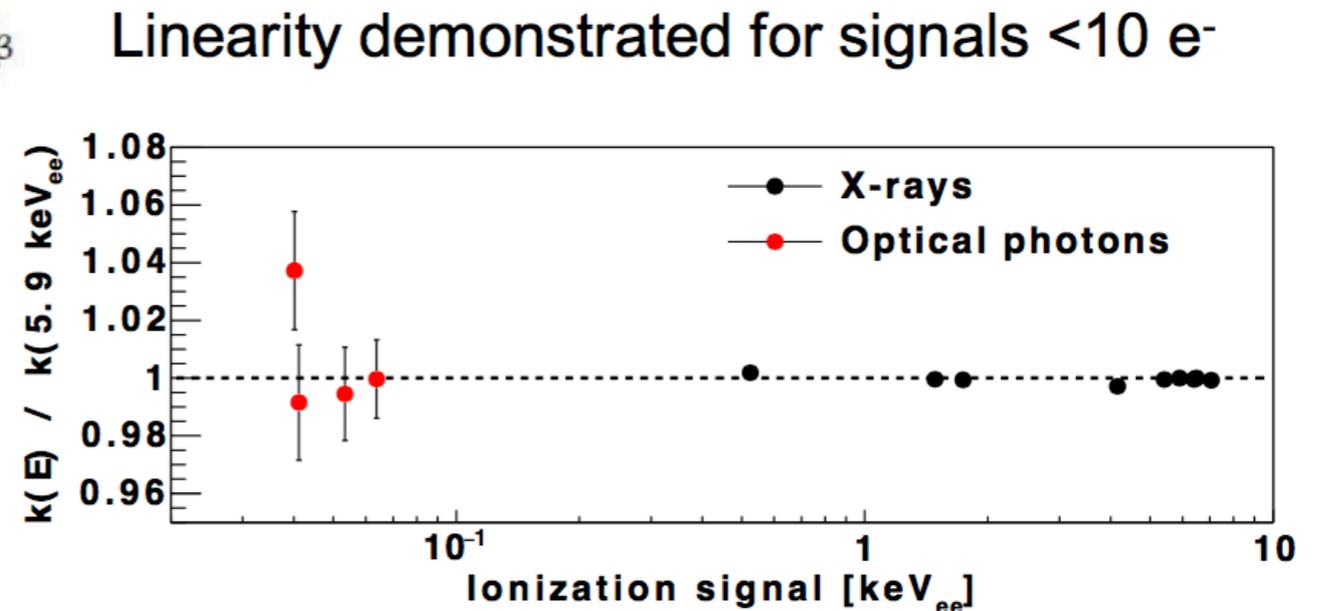
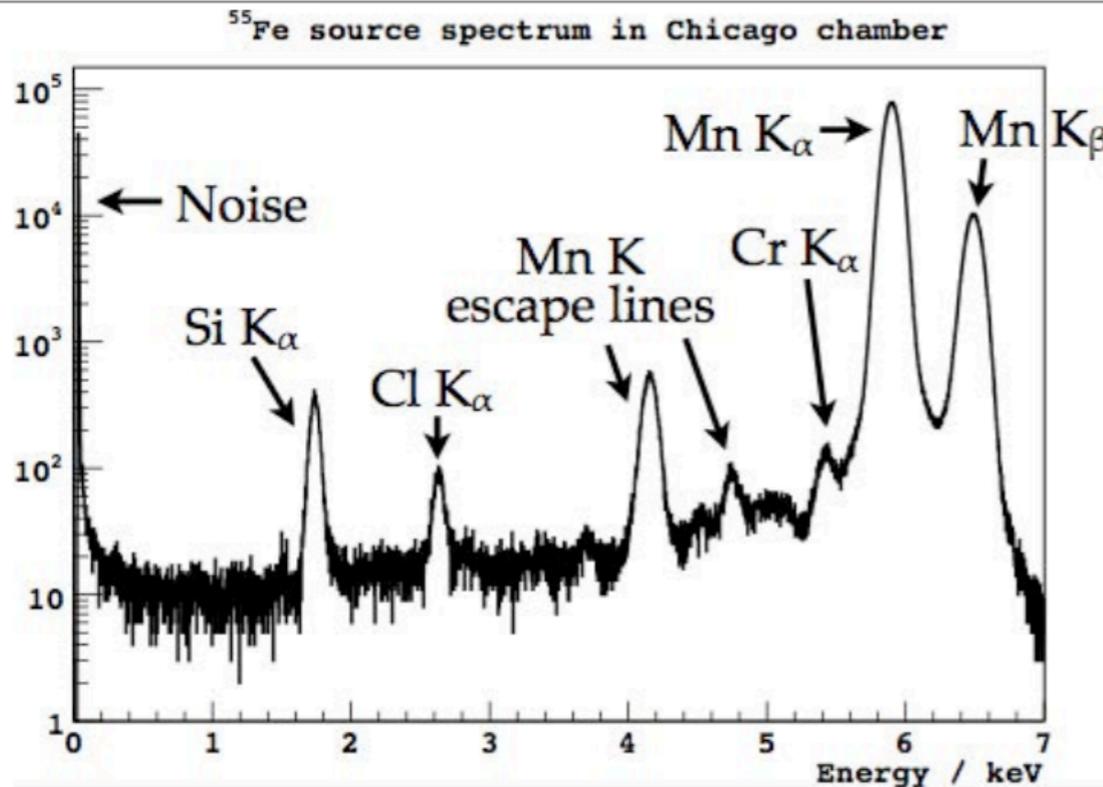
$^{232}\text{Th}$

➤ Upper limit:  $15/\text{kg/day}$  [95%]

*arXiv:1506.02562*

**Difference (factor x7) in  $^{32}\text{Si}$**

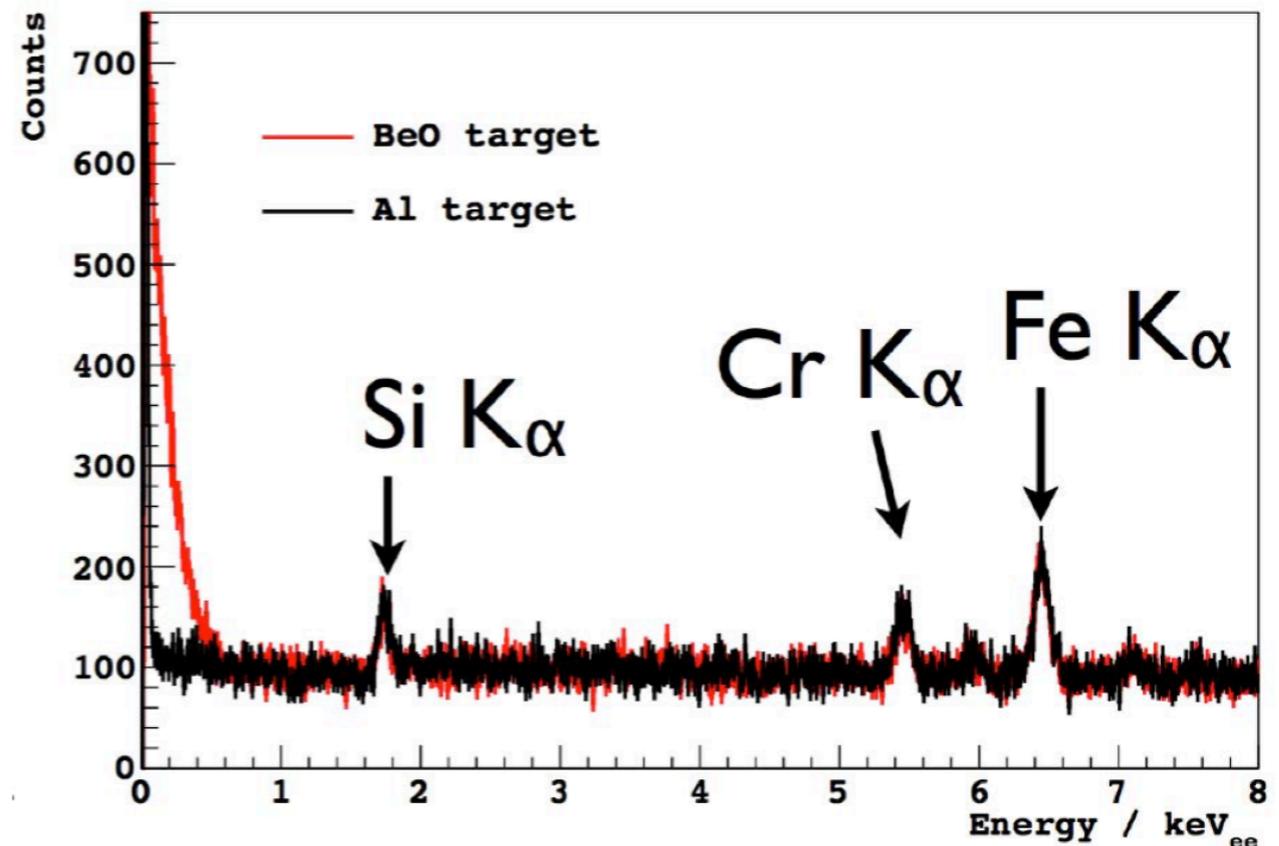
# CCD response : Energy linearity and neutron recoil



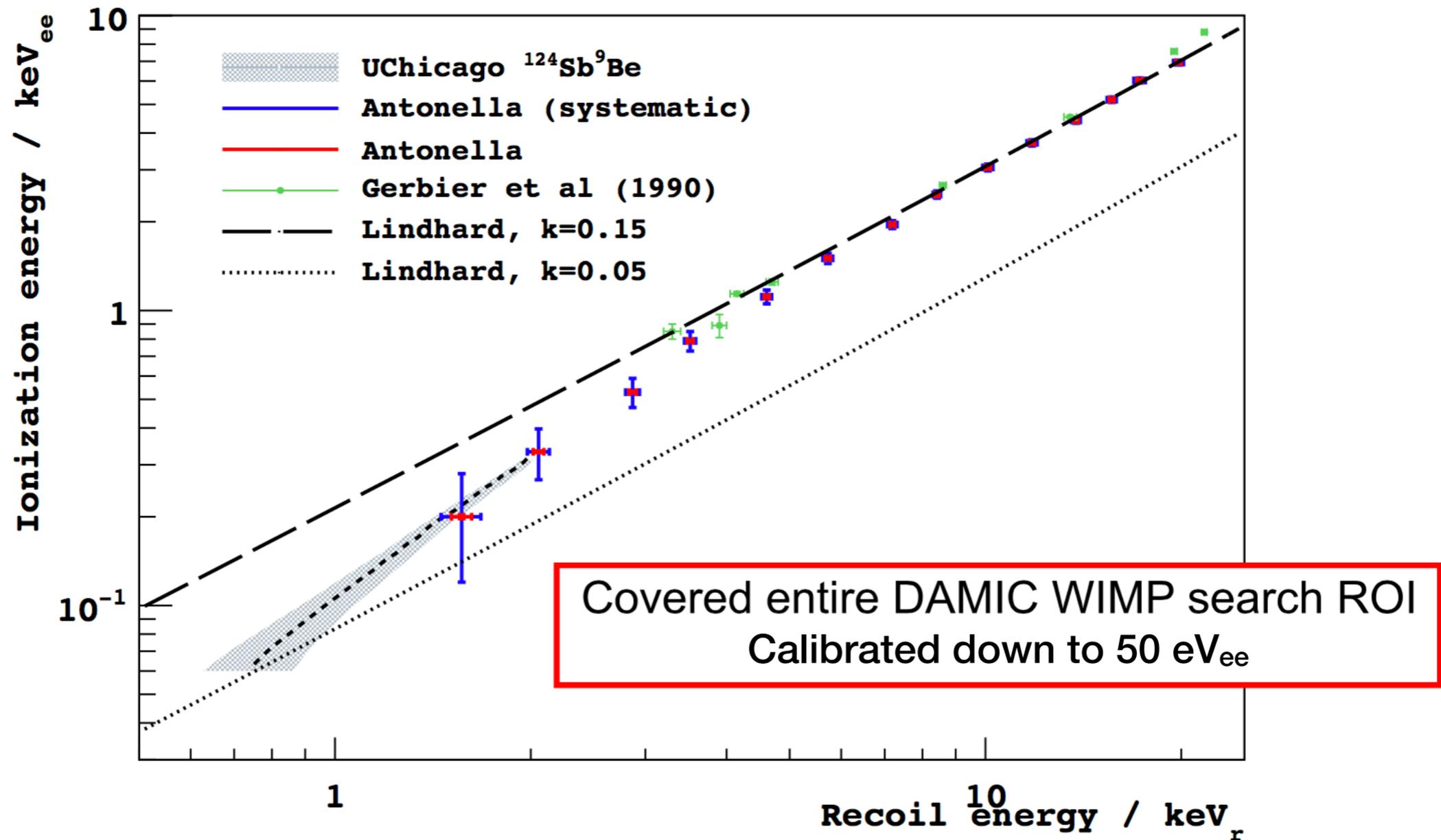
## Antimony source with Be / Al cap

“Neutron-on” with BeO (n+ $\gamma$ )  
 “neutron-off” with Al (only  $\gamma$ )  
 Clear signal from neutron-induced nuclear recoils

Nuclear recoil ionization efficiency from adjusting MC  $E_r$  to  $E_e$  spectrum



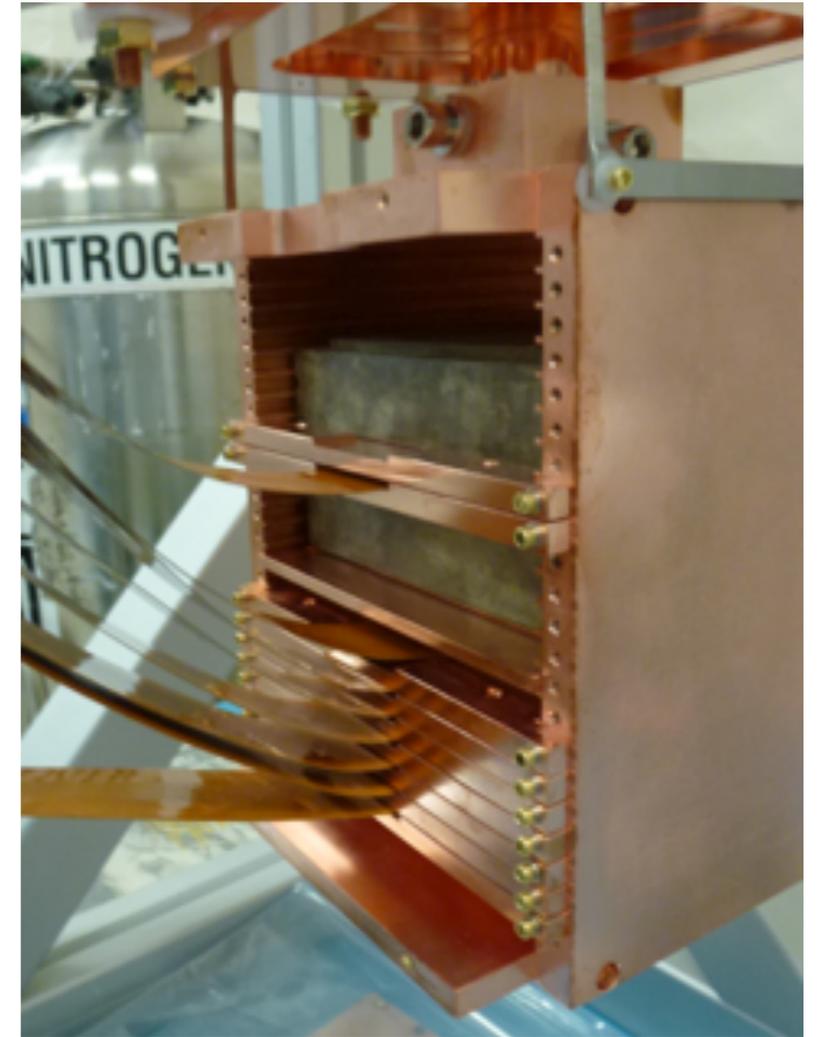
# Nuclear-recoil Ionisation efficiency



Deviation from Lindhard theory below 5.5 keV,  
crucial for low-mass WIMPs search in Si

# Current detector configuration

- ▶ **Seven 16 Mpix CCDs** in stable data taking since 2017 (1 CCD sandwiched in ancient lead)
- ▶ **40 g detector mass**
- ▶ Operating temperature down to **140K**
- ▶ **Exposure time for image : 8h and 24h** (each image acquisition is followed by a “blank” with zero exposure)
- ▶ **7.6 kg day of data acquired for background characterisation** and  $^{32}\text{Si}$ ,  $^{210}\text{Pb}$  measurements
- ▶ **13 kg day of data collected for DM search** (with  $1 \times 100$  hardware binning to enhance the signal/readout noise and lower the energy threshold)
- ▶ **Since 2019, resumed background runs and detector studies** for future DAMIC upgrade

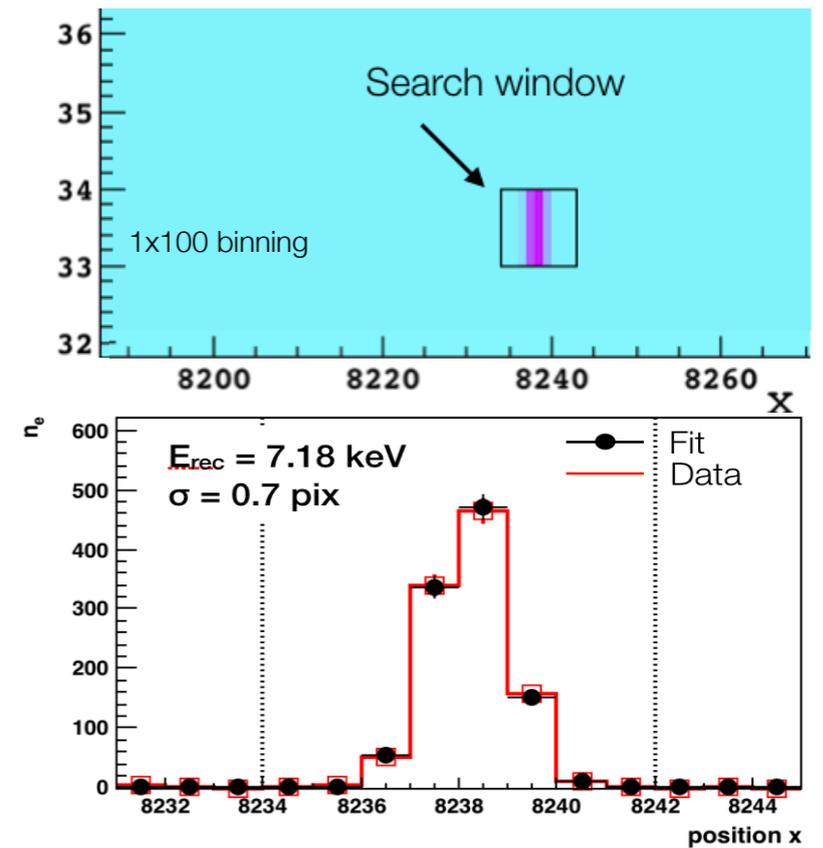


# WIMPs search strategy

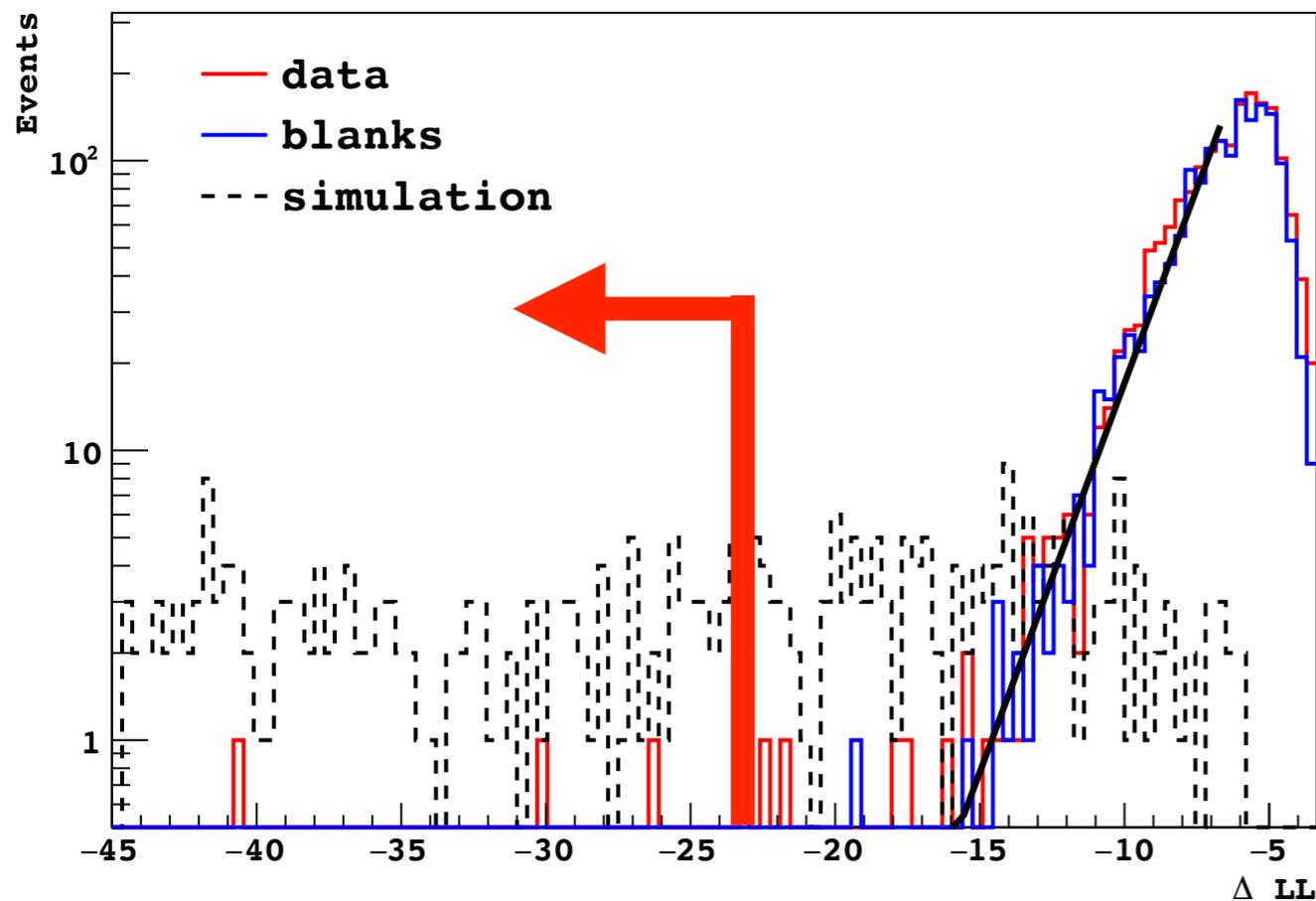
- ▶ Pedestal and correlated noise subtraction (hot pixels among several images masked)
- ▶ LL fit in a moving window across the image

$$\Delta LL = L_n - L_s$$

flat noise  $\curvearrowright$   $L_n$        $L_s$   $\curvearrowright$  Gaus signal + flat noise

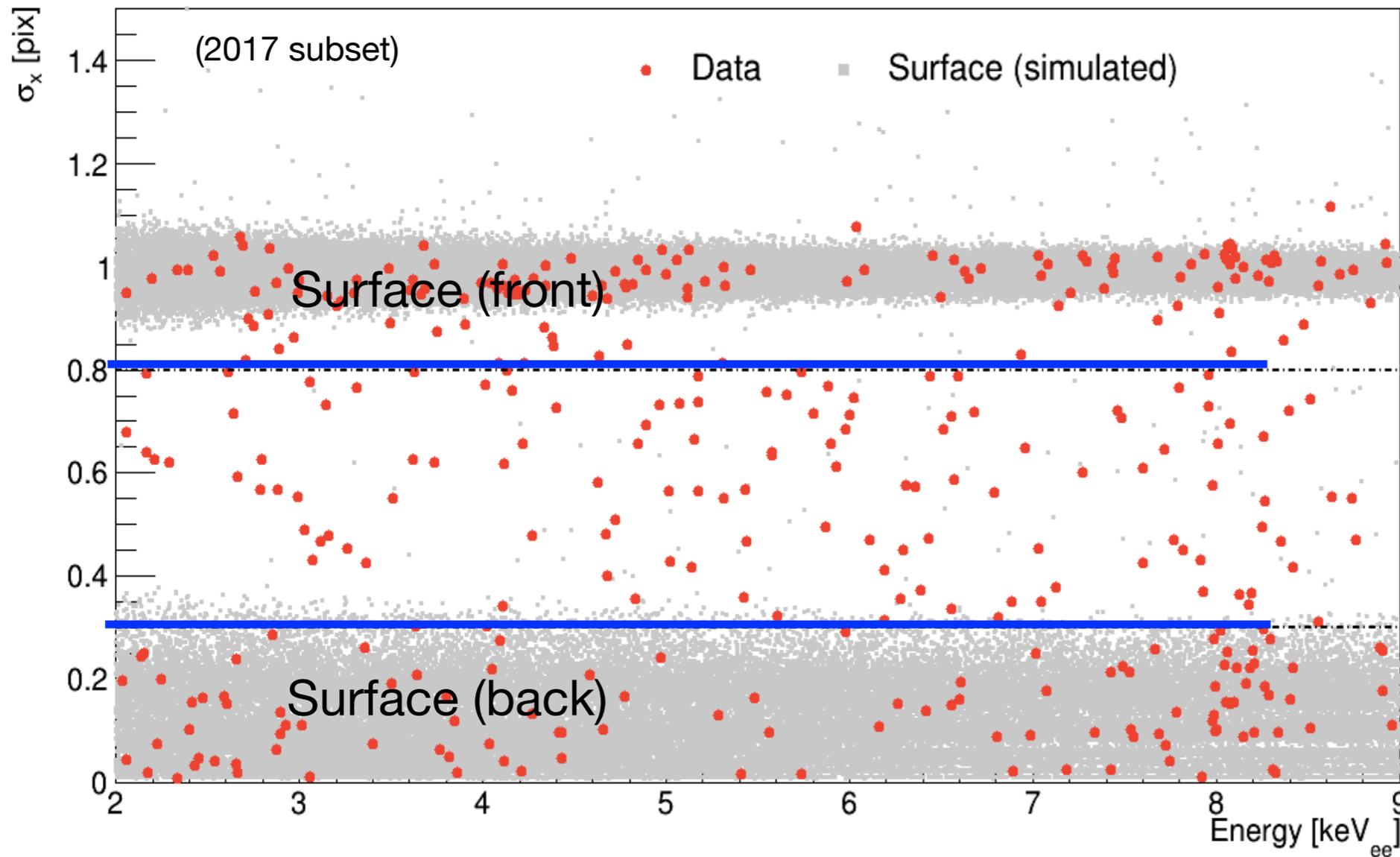


- ▶  $\Delta LL$  profile for noise clusters (in blanks)



- ▶ Define  $\Delta LL$  cut at  $< 0.001$  noise events;
- ▶ Selection efficiency  $\sim 10\%$  at  $50$   $eV_{ee}$  and  $> 80\%$  at  $100$   $eV_{ee}$

# Background handling



Fiducial volume

$$0.3 < \sigma < 0.8$$

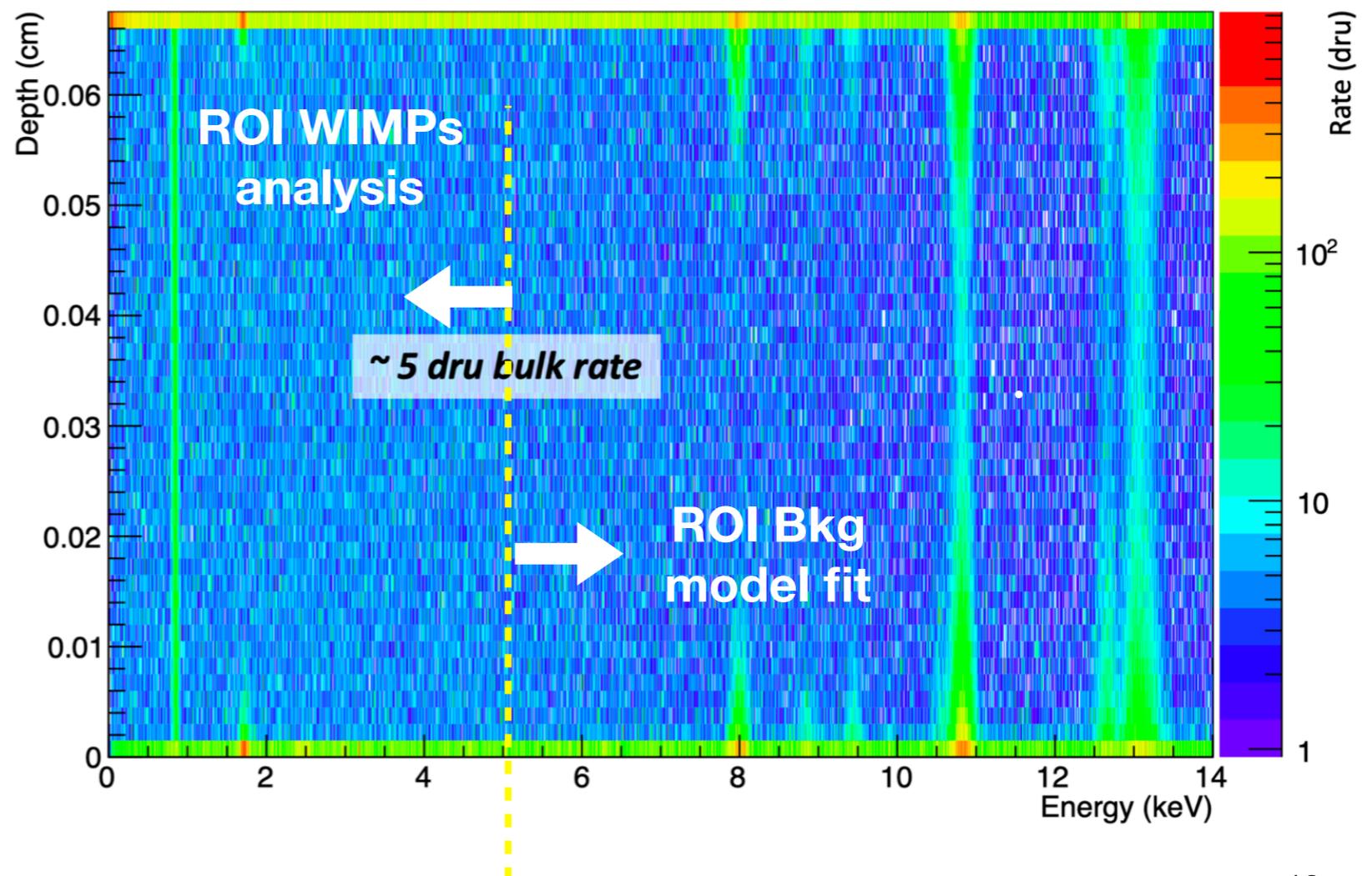
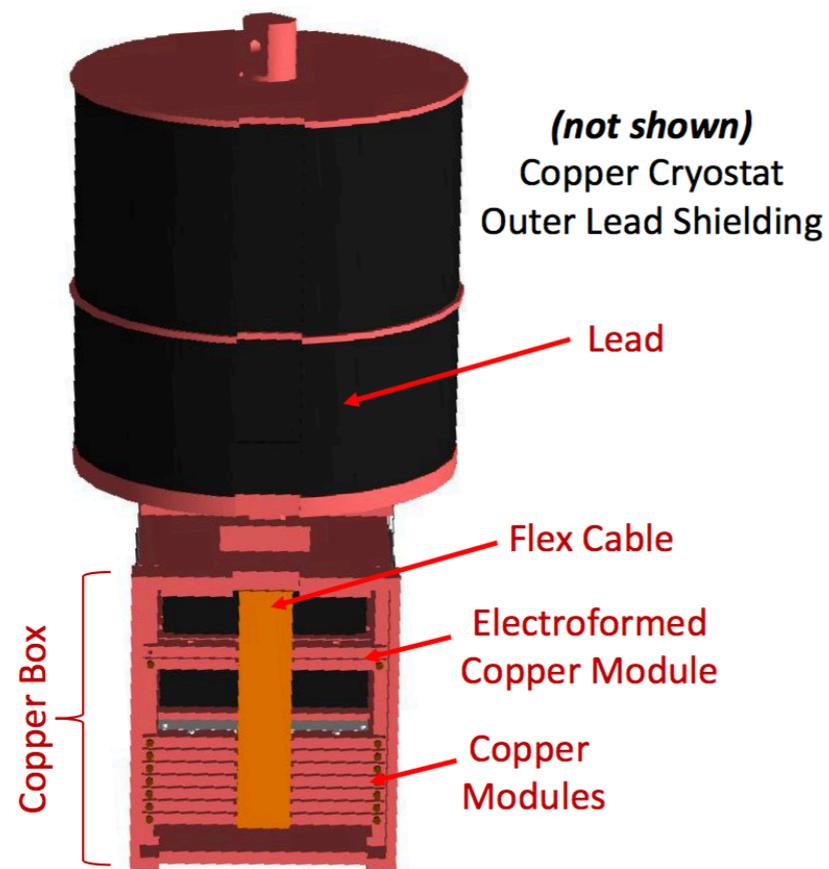
## Background handling approaches

**(a) Surface bkg rejection (2016)** : fiducial cut in depth to remove surface events (2016 results)

**(b) \*New 2019\* Bkg model** : based on Geant4 simulations of isotopes in the CCD bulk and surrounding materials and 2D binned LL fit to data

# Background model

- Geant4 simulation + detector readout of isotopes in detector material (assays to measure the activity of each component)
- Cluster reconstruction (energy, position and depth parameter  $\sigma$ ) as in data.
- **2D binned likelihood fit of  $(E, \sigma)$**  to data in the range 6-20 keV<sub>ee</sub> and we use it to construct the model at low energies

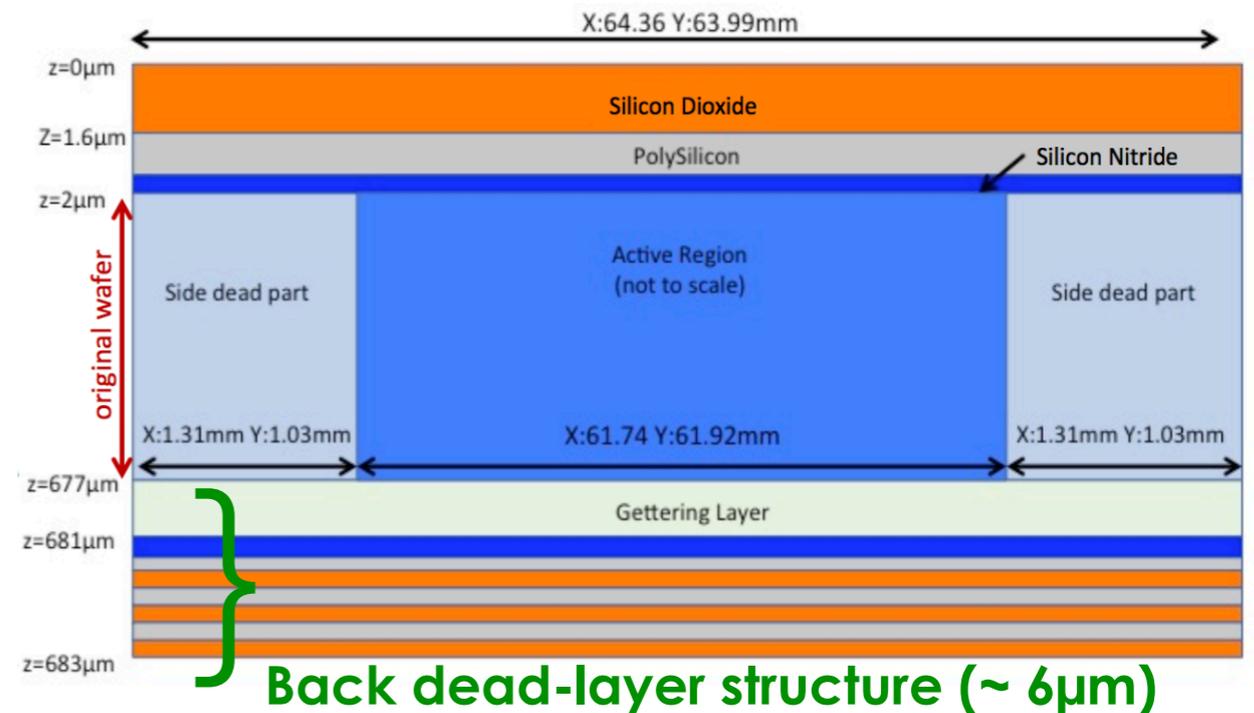
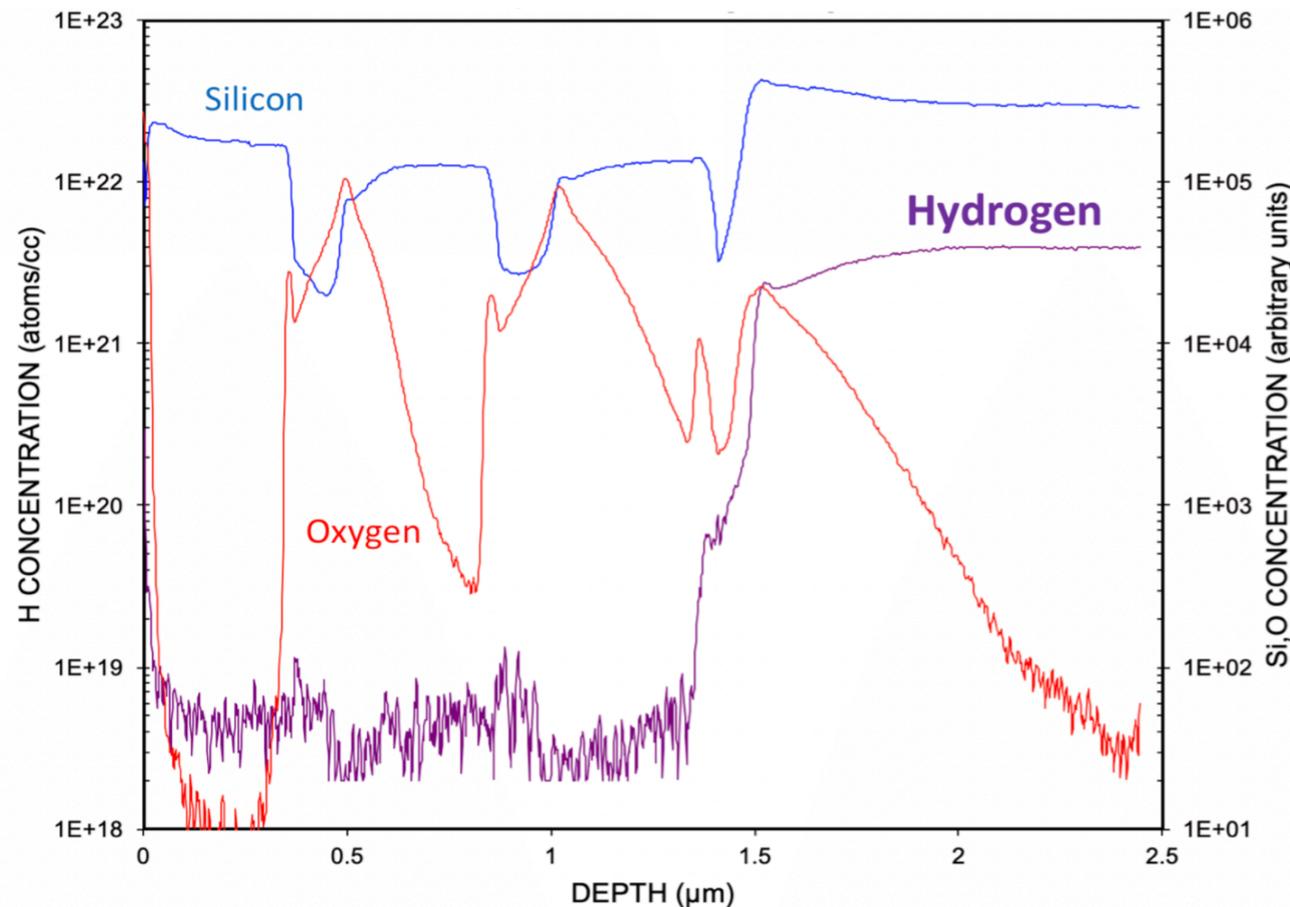
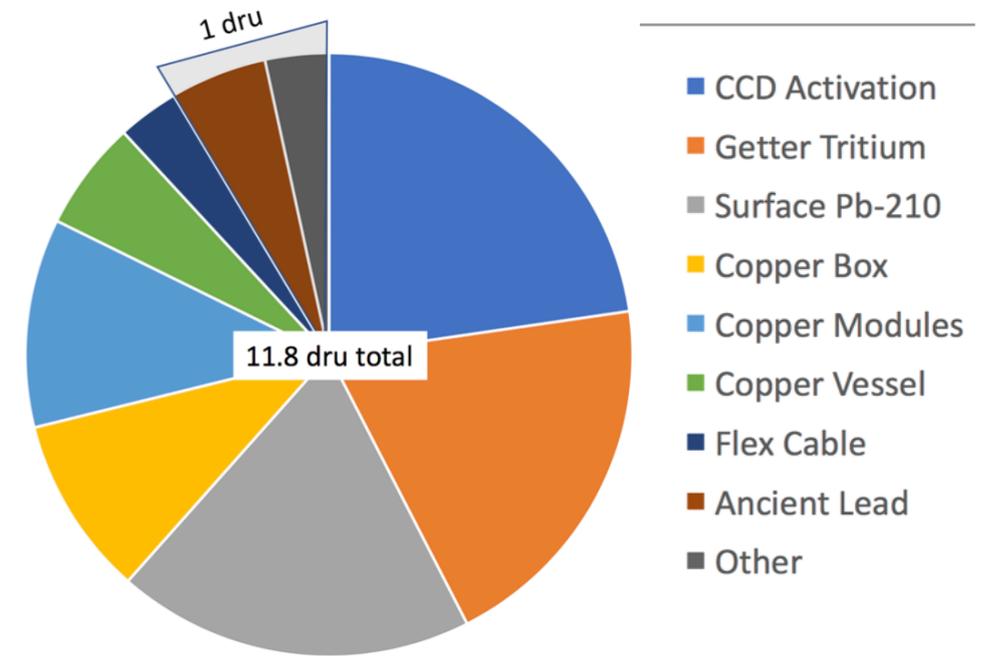


# Main background contributions

Main contributions in (6 - 7.5keV):

- **$^{210}\text{Pb}$  on CCD and OFC Copper** surfaces
- **Tritium** on **bulk** (cosmogenic activation) and trapped in **back** dead-layers during CCD manufacturing

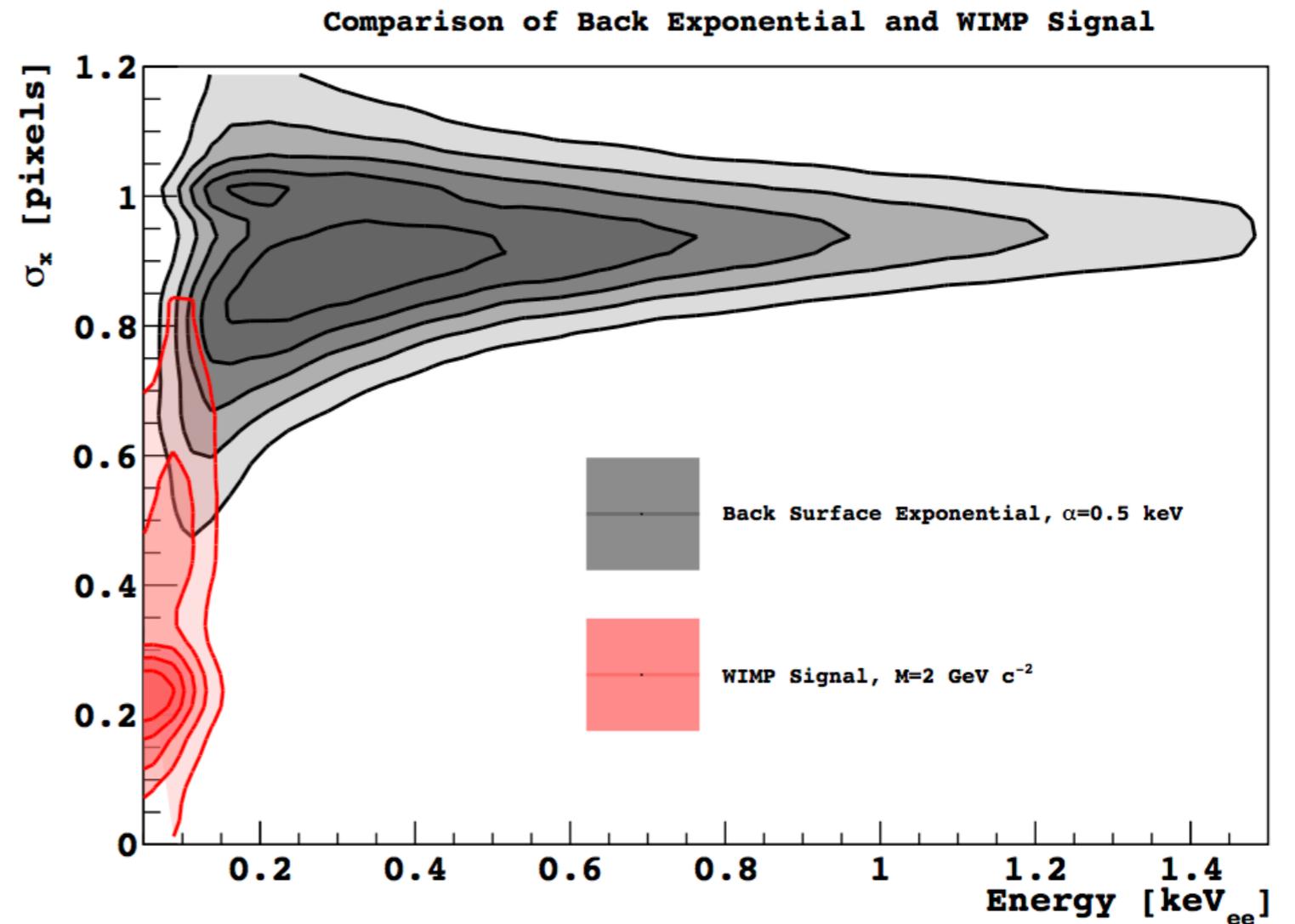
*SIMS measurements on the back of CCD confirmed the presence of  $10^{21}$  H atoms/cm<sup>3</sup> consistent with  $^3\text{H}$  decay (assuming  $^3\text{H}/\text{H}$  fraction of  $10^{-18}$  as in water)*



**New R&D programs using SNOLAB setup to improve bkg in CCD production**

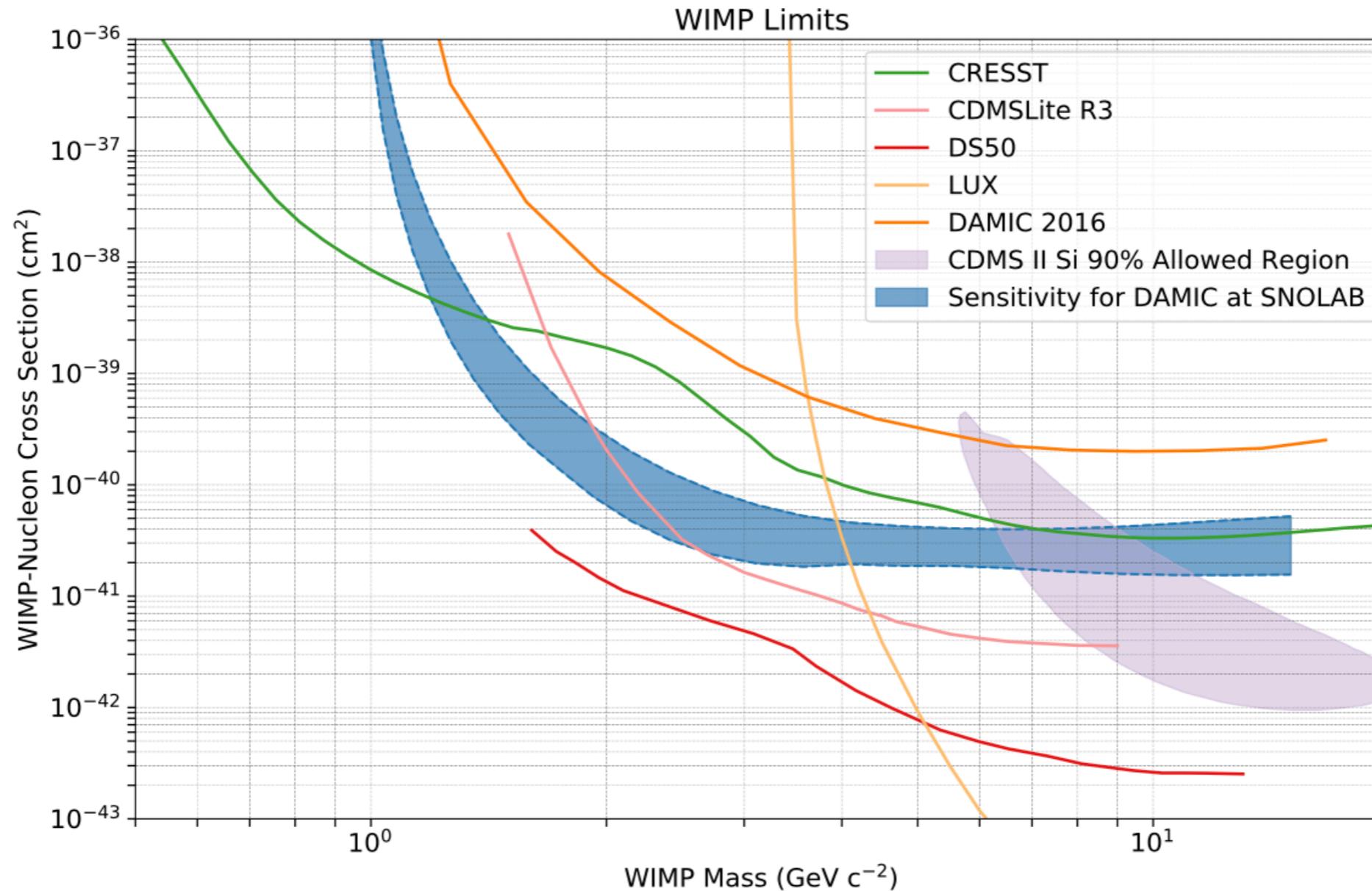
# Systematic uncertainties

- ▶ Dominant systematic uncertainty are radioactive contaminants on the back surface ( $^{210}\text{Pb}$ ,  $^3\text{H}$ ) from production and handling of CCDs
- ▶ The fit prefers  $^{210}\text{Pb}$  on the back of the active region
- ▶ By manually removing some of the  $^{210}\text{Pb}$  templates and repeating the fit, we are able to quantify the systematic uncertainty in the background model due to the location of  $^{210}\text{Pb}$ .
- ▶ Resulting difference is a back-side exponential component non-degenerate with an a low WIMPs signal



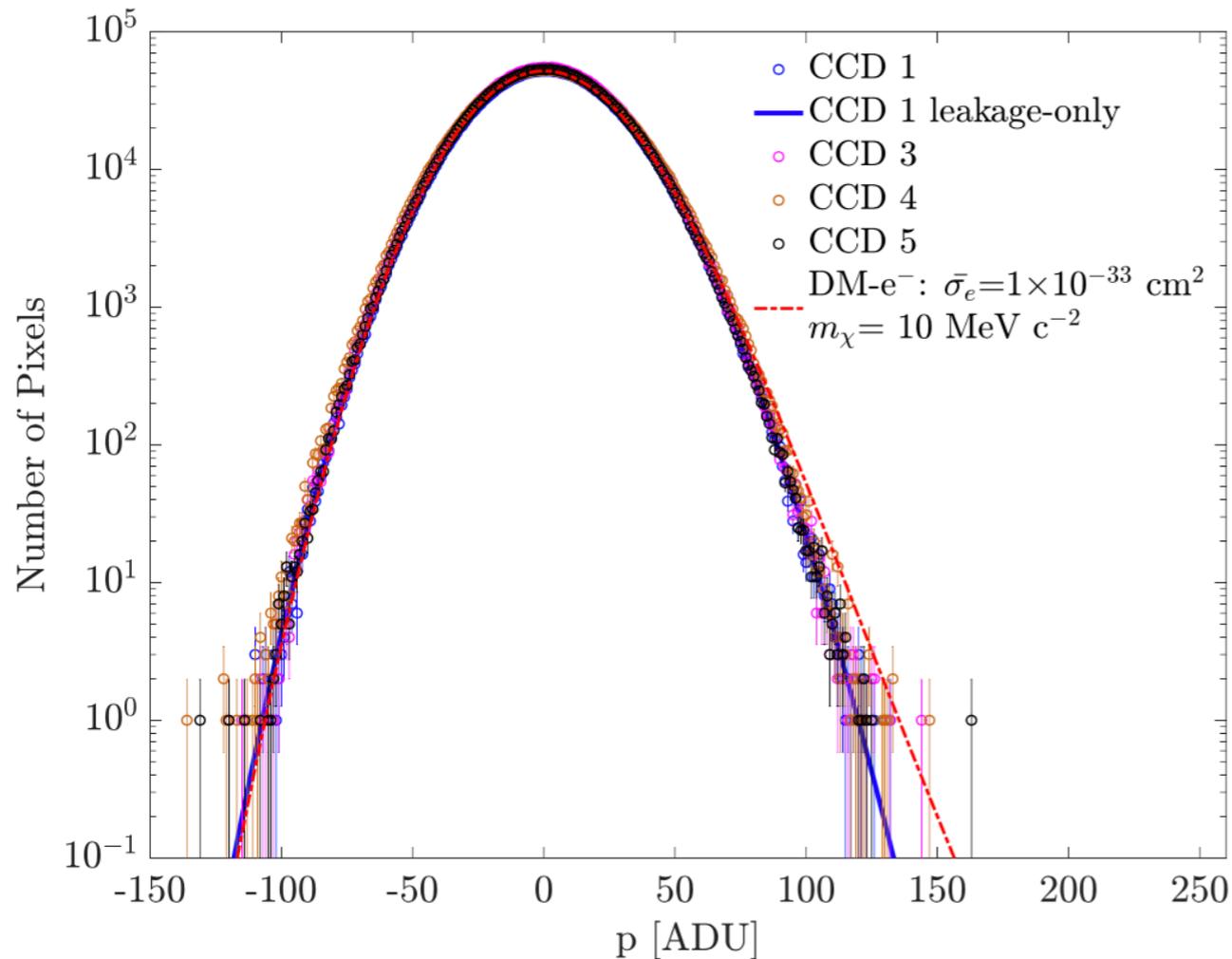
# Expected sensitivity

Sensitivity with 2017-2018 data.



Exploring for the 1st time the CDMS signal with the silicon target and a much lower energy threshold (0.6 keV<sub>nr</sub> ~ 0.05 keV<sub>ee</sub>)

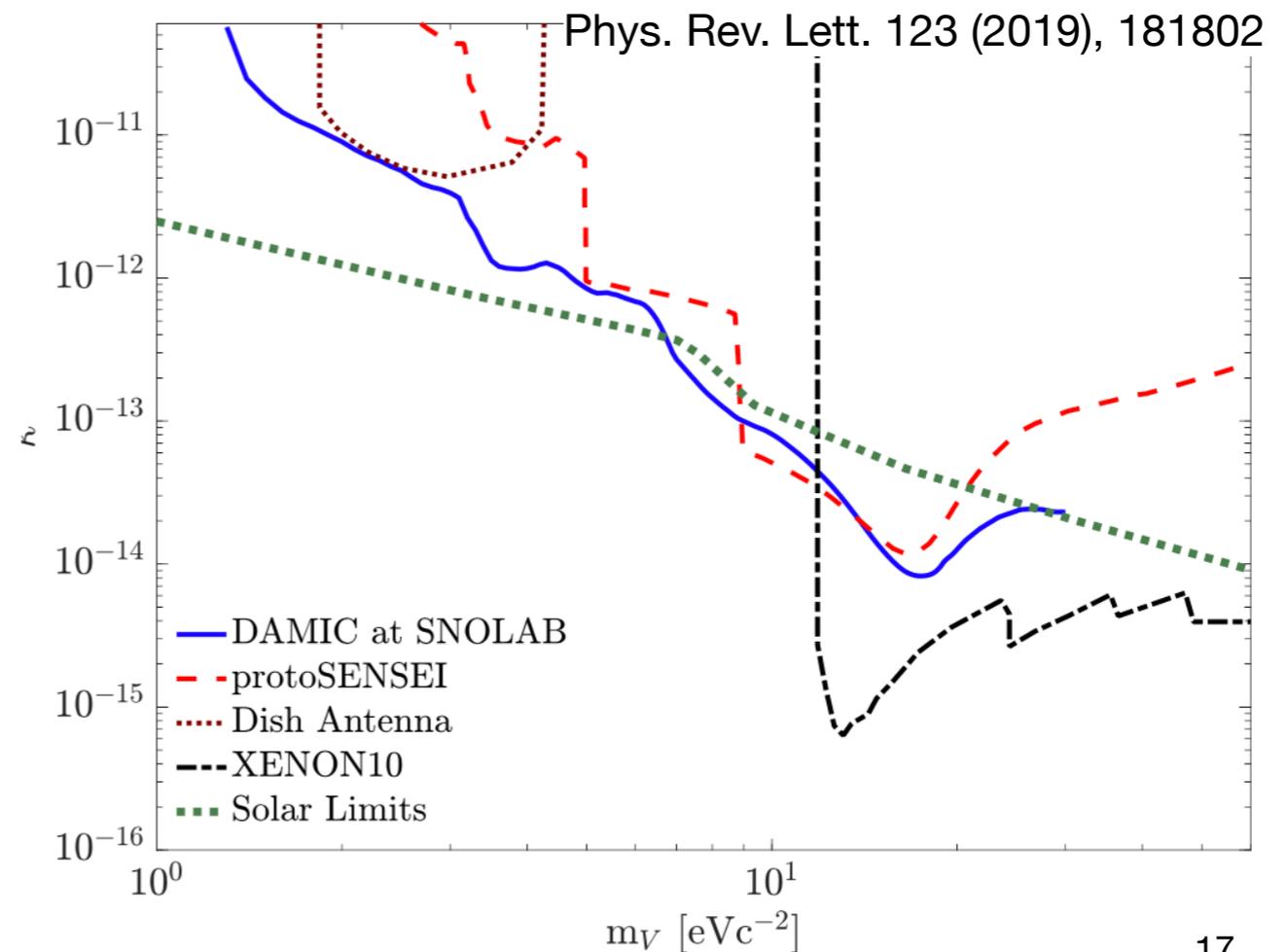
# Hidden photon search



LL fit to compare charge distribution to **leakage only model** and to **noise + signal** (hidden photon absorption / DM e-scattering signal)

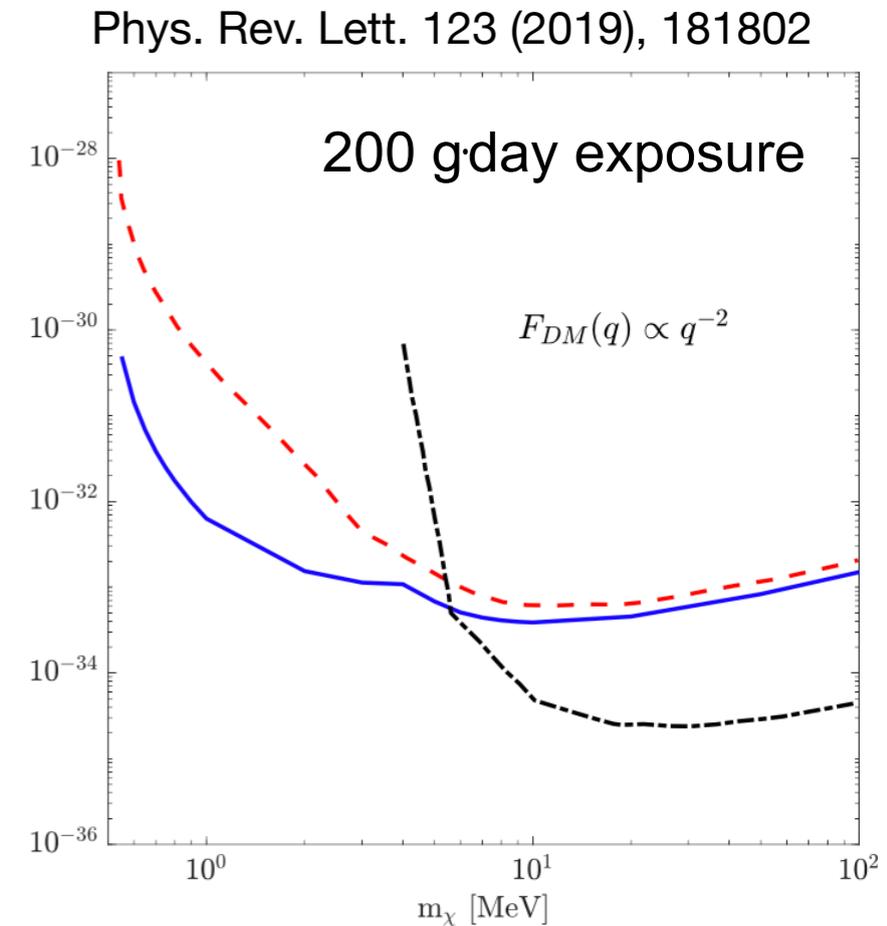
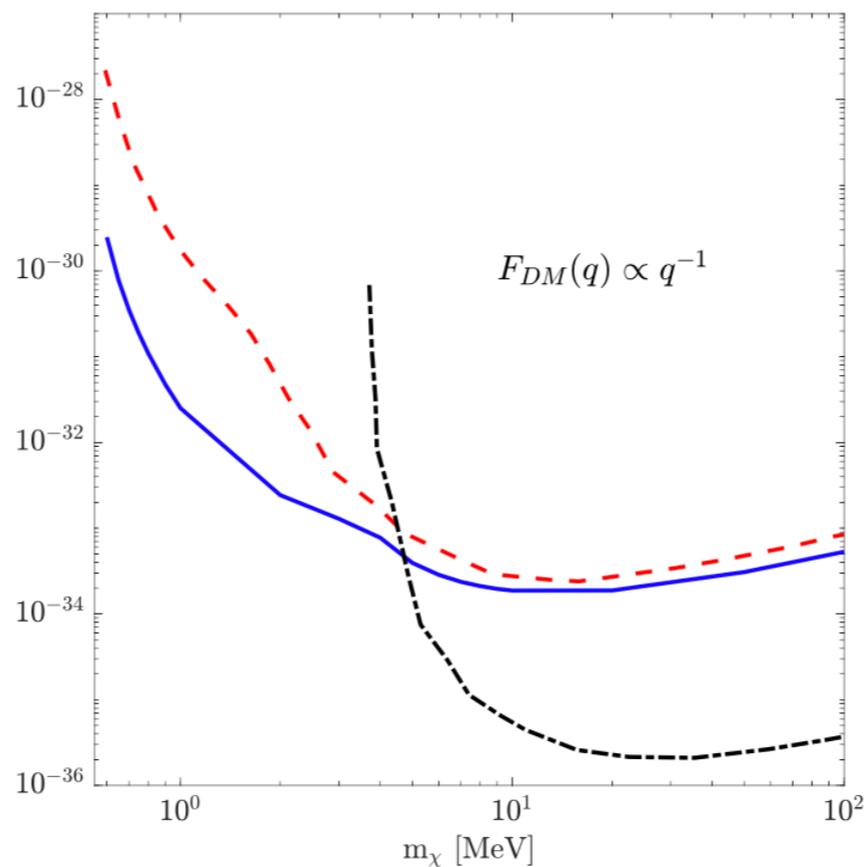
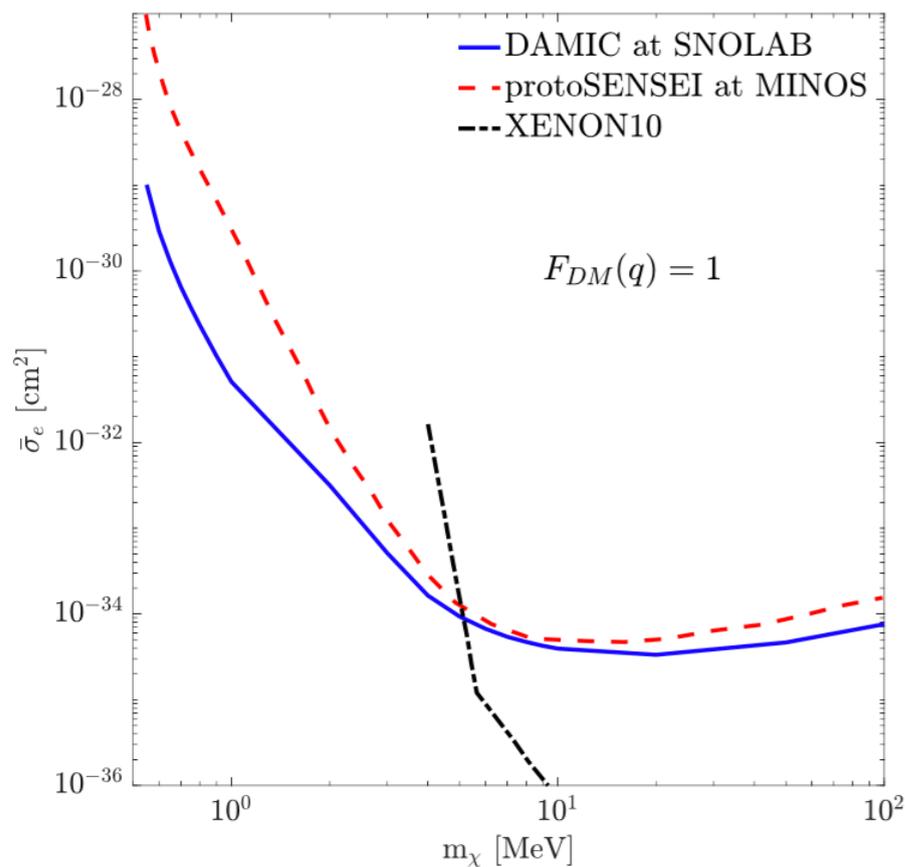
Best limit for masses 1-10 eV/c<sup>2</sup>

- Evaluate the charge distribution of pixels with  $< 10 e^-$  for contributions beyond the leakage current
- Select images with similar dark current. Data set: 200 g d exposure in 100ks images

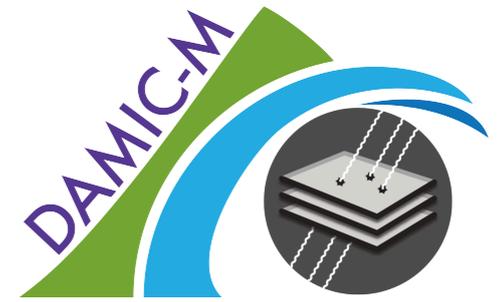


# DM - electron scattering

- ▶ Leakage current analysis as for hidden photon search
- ▶ Compare pixel distribution to **noise + signal from DM-e scattering**



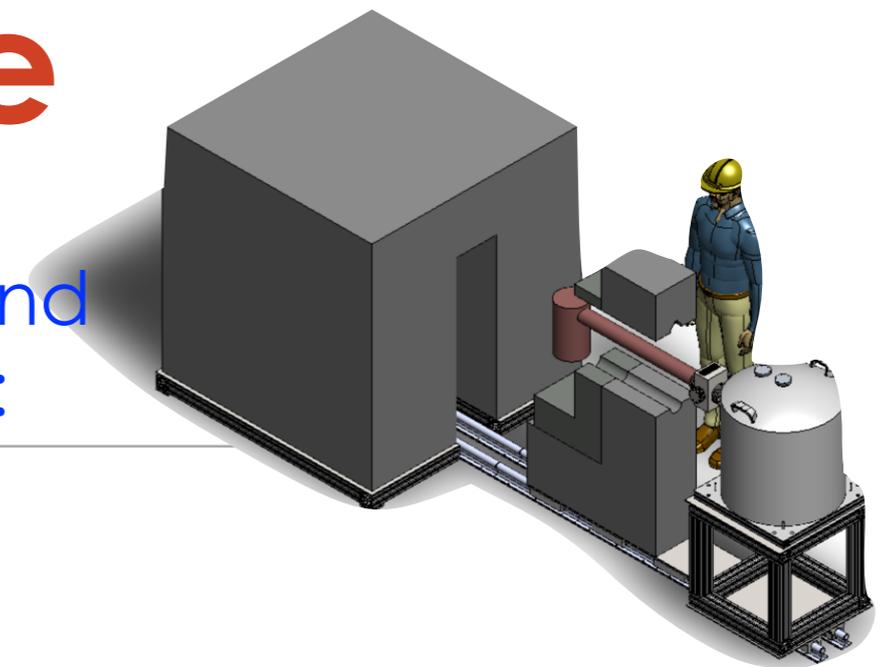
**Best exclusion limits for the scattering of DM particles  
with masses  $< 5 \text{ MeV}/c^2$**



**Next phase of DAMIC is started**

# DAMIC-M at Modane

A kg-size improved detector for WIMPs and dark-sector candidates at low-masses:

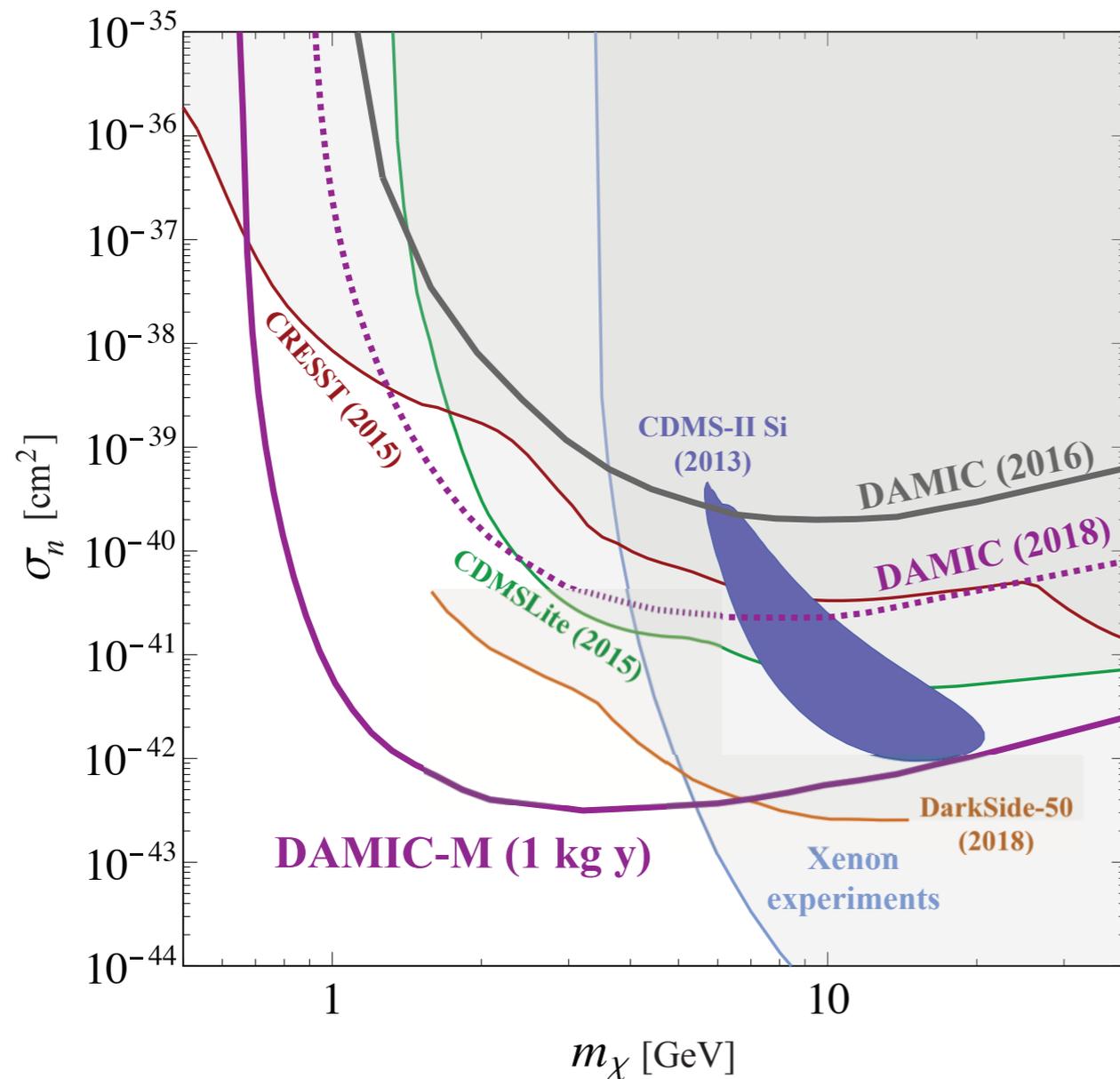


European Research Council  
Established by the European Commission

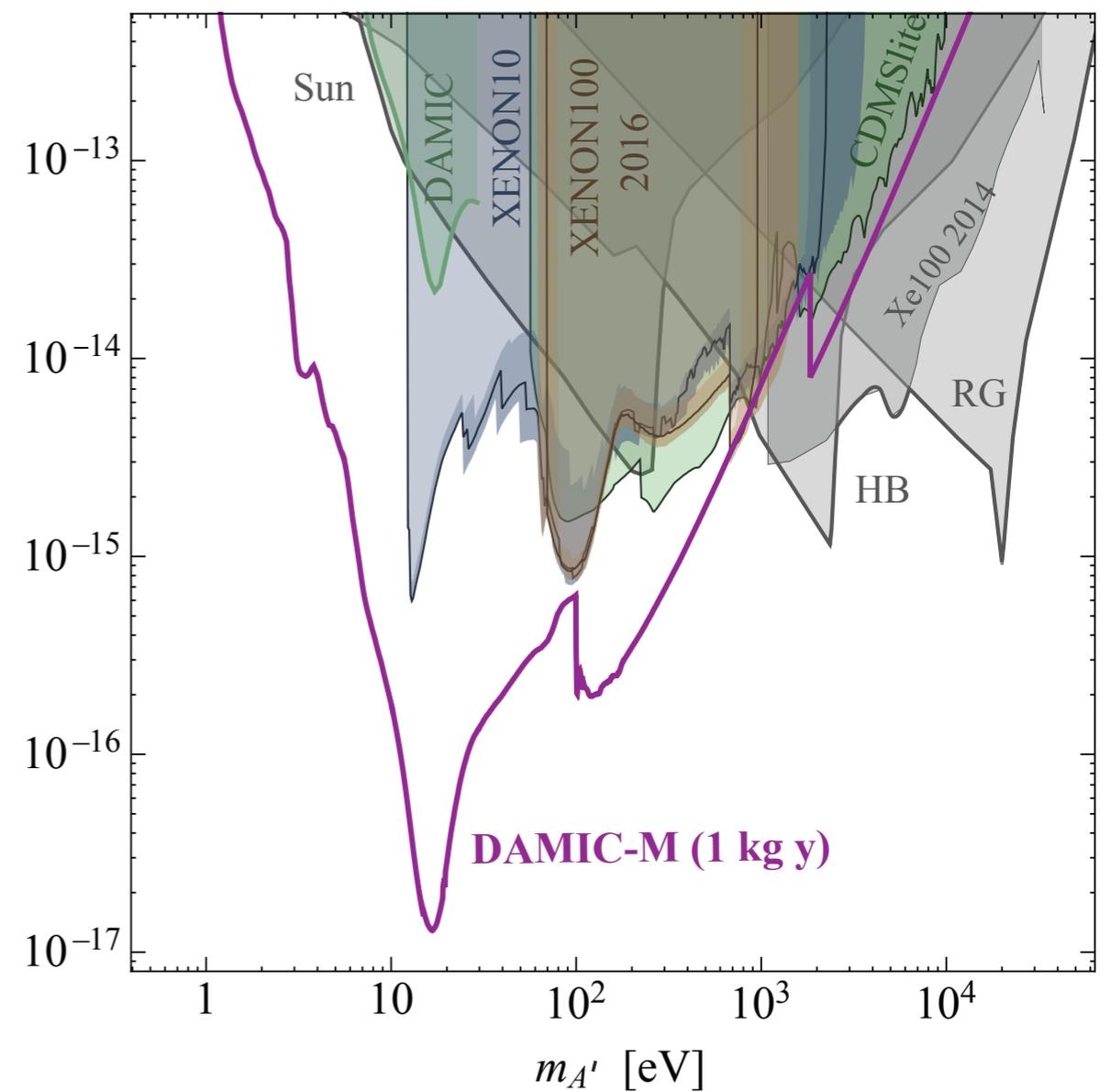


# Low mass WIMPS and Hidden photons

WIMP nuclear recoil search

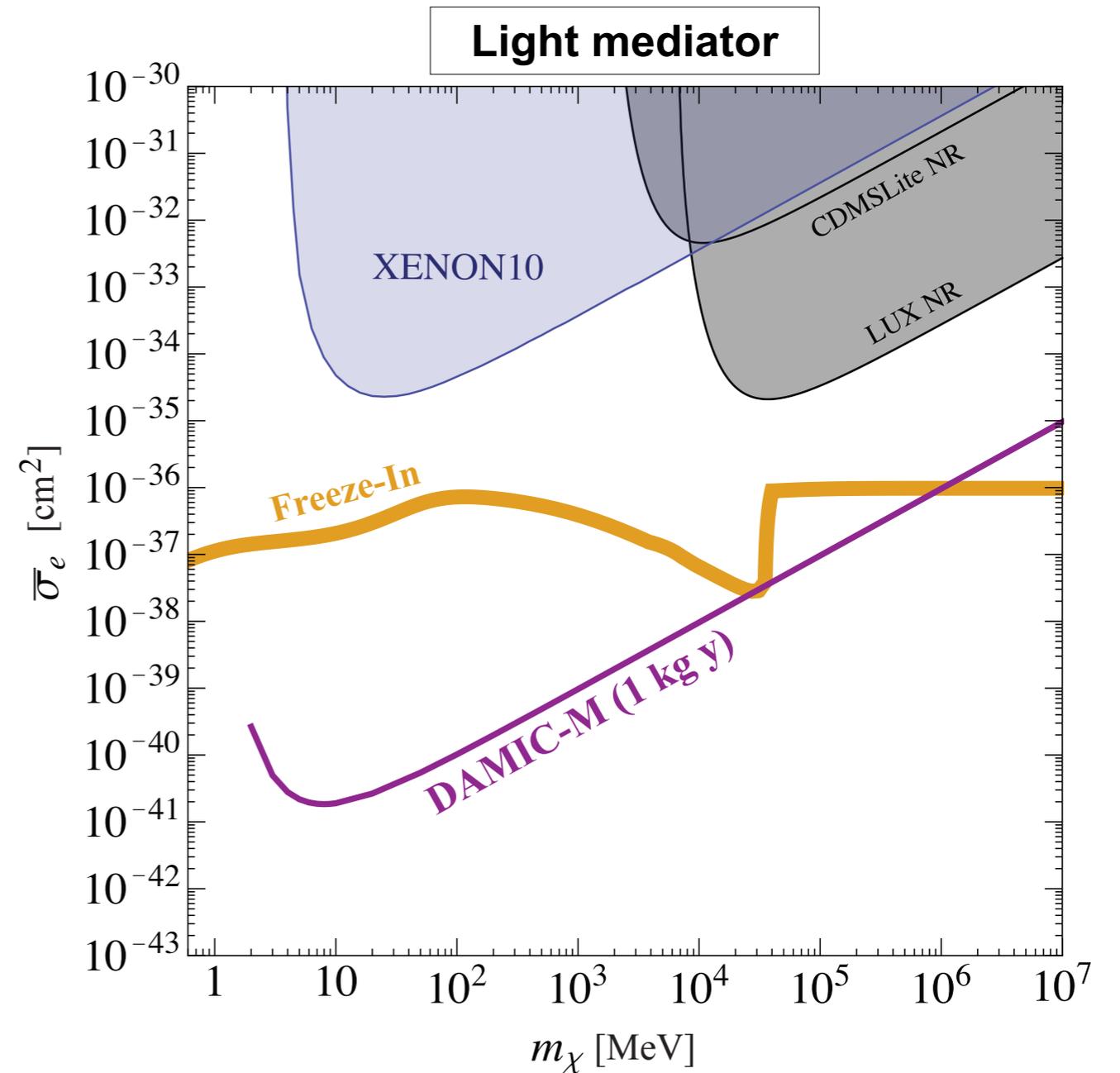
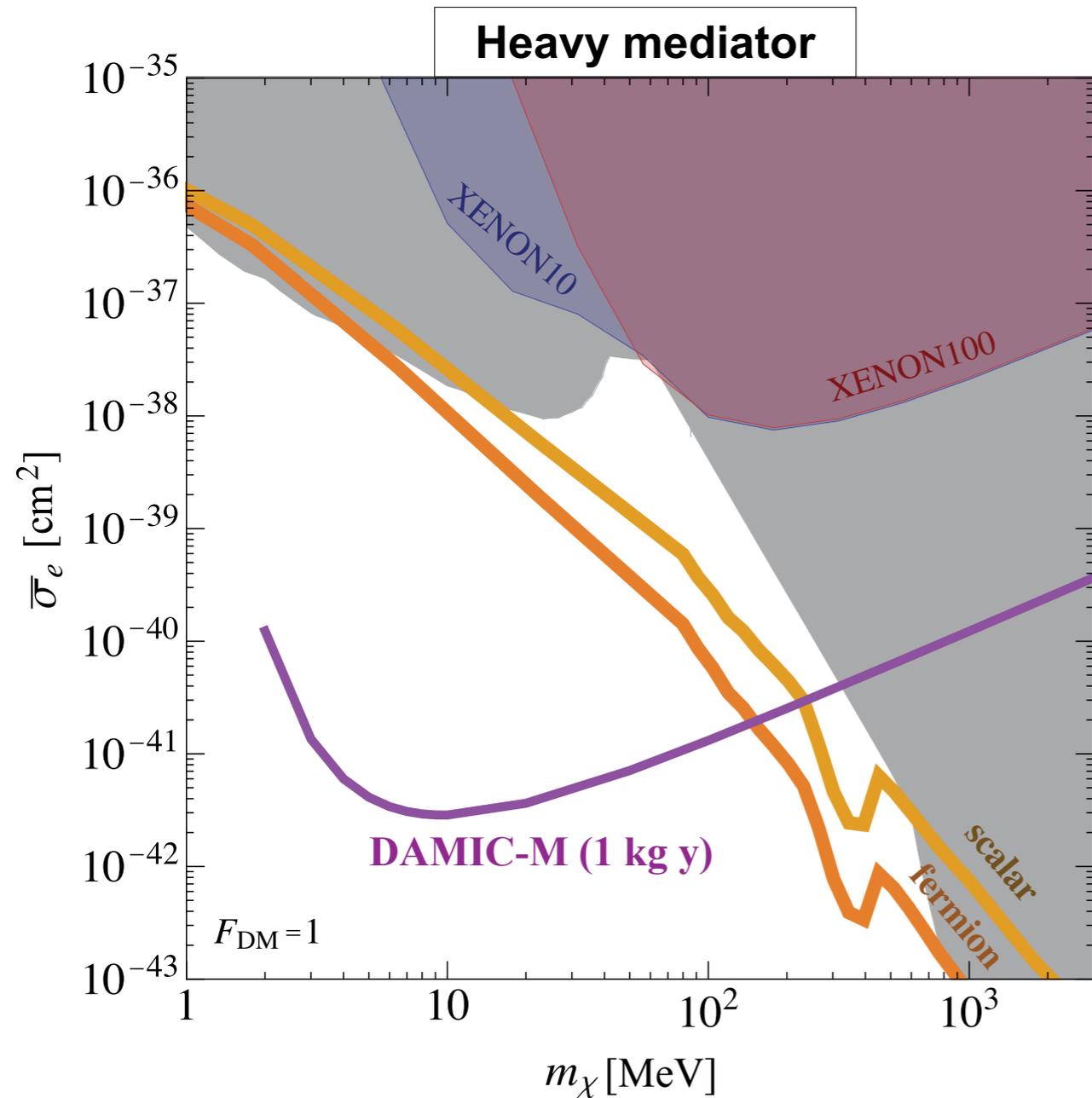


Hidden photon search



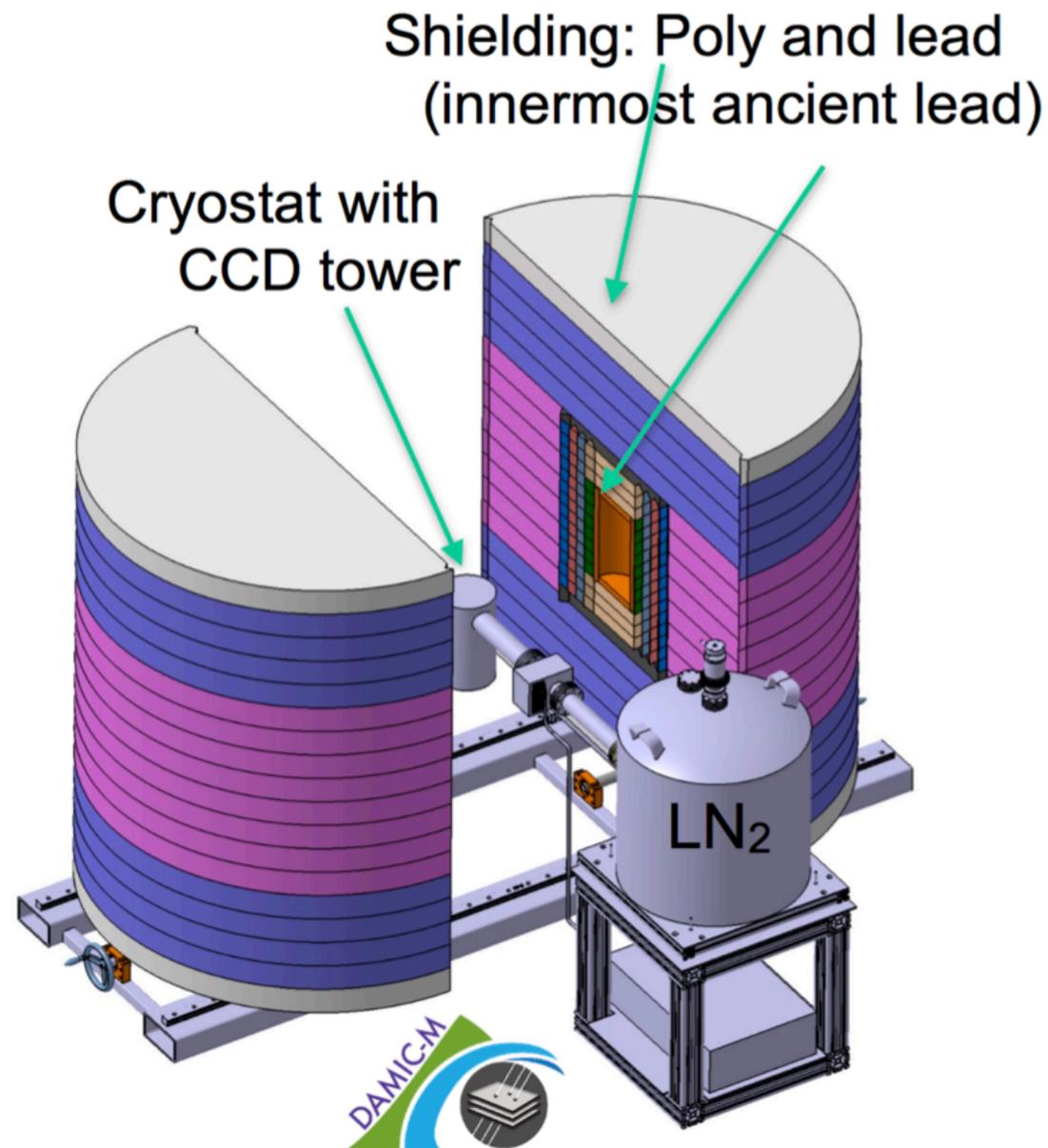
# Light dark matter (hidden sector)

## Light dark matter - electron scattering



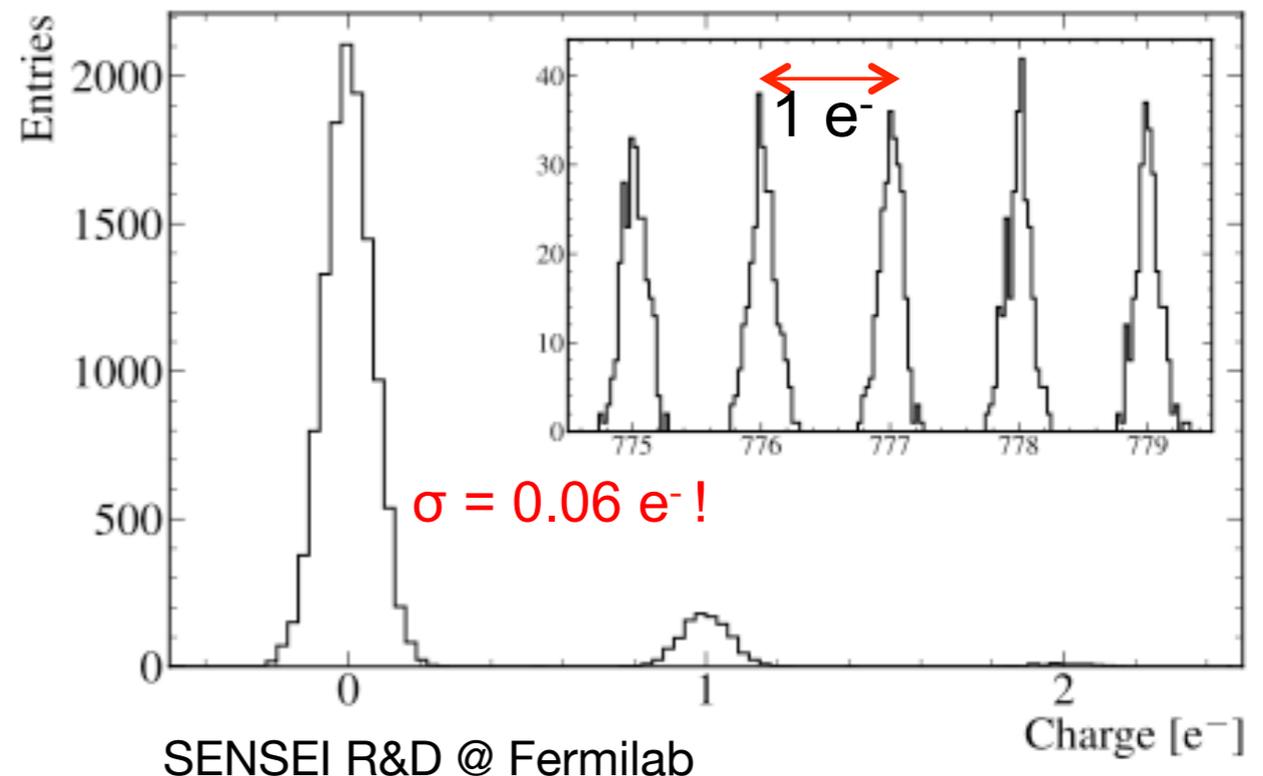
DAMIC-M will be sensitive to light dark matter even if these candidates constitute only a small fraction of the total DM in the universe

# DAMIC at Modane: what's new



- ▶ 50 CCDs (kg-size detector)
- ▶ CCD mass > 10 g
- ▶ Skipper readout for **single electron resolution (sub-eV noise)** demonstrated by SENSEI project
- ▶ **Bkg reduction to a fraction of dru** (improved design, materials, procedures)

R&D and design up to 2021  
Construction 2022  
Installation in 2023

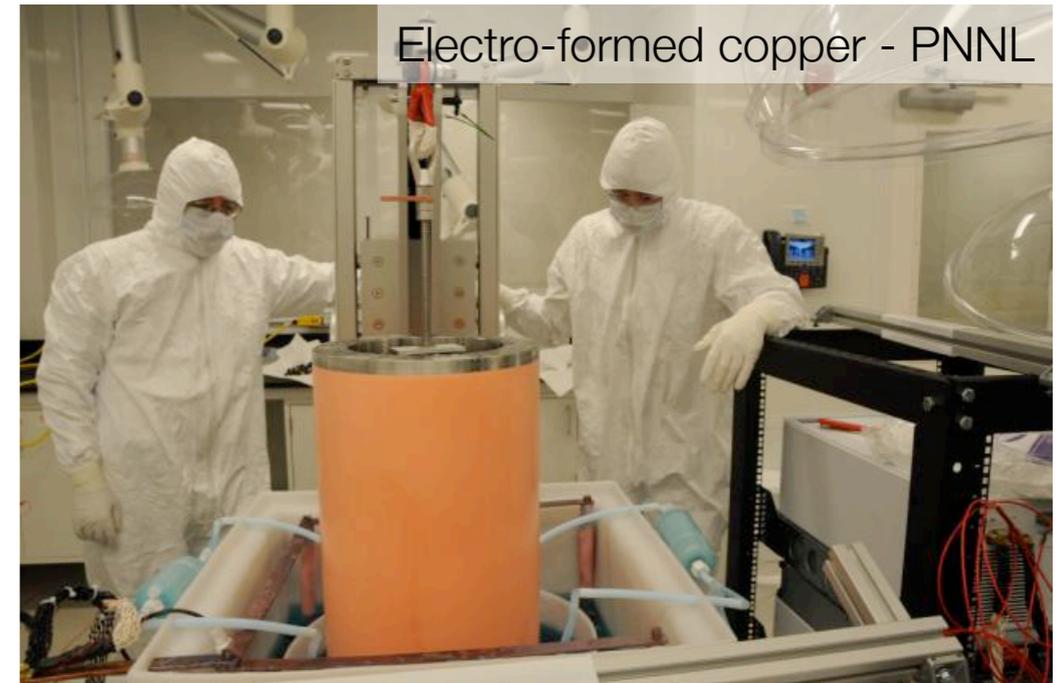


# Background improvements for DAMIC-M

## Background (to be reduced to fraction of DRU)

- **External background:** better material selection and handling (e.g. electroformed copper, surface contamination, Rdn)
- **Bulk background:**  $^{32}\text{Si}$  et tritium

Produced by cosmic rays on the Ar in atmosphere, it deposits on ground with precipitations (rain, snow,...)



## Tritium : expected to be the dominant background

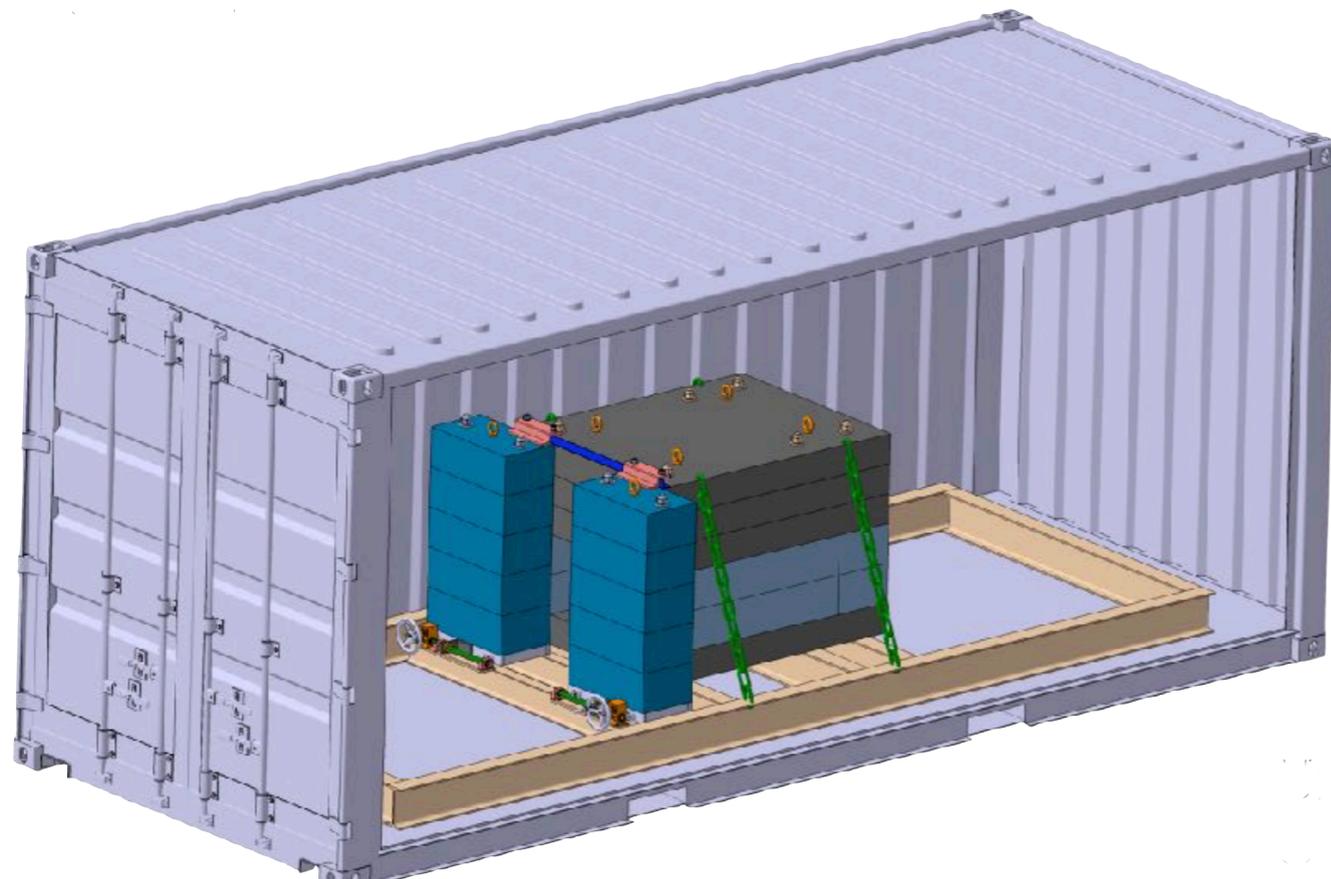
Produced by the cosmic neutrons and muon spallation in the Si bulk. production rate  $\sim 25 - 100/\text{Kg}/\text{day}$  (s.l.)

—> Minimize the time CCDs are exposed to cosmic rays (stock CCD underground, shielded container for transportation,...)



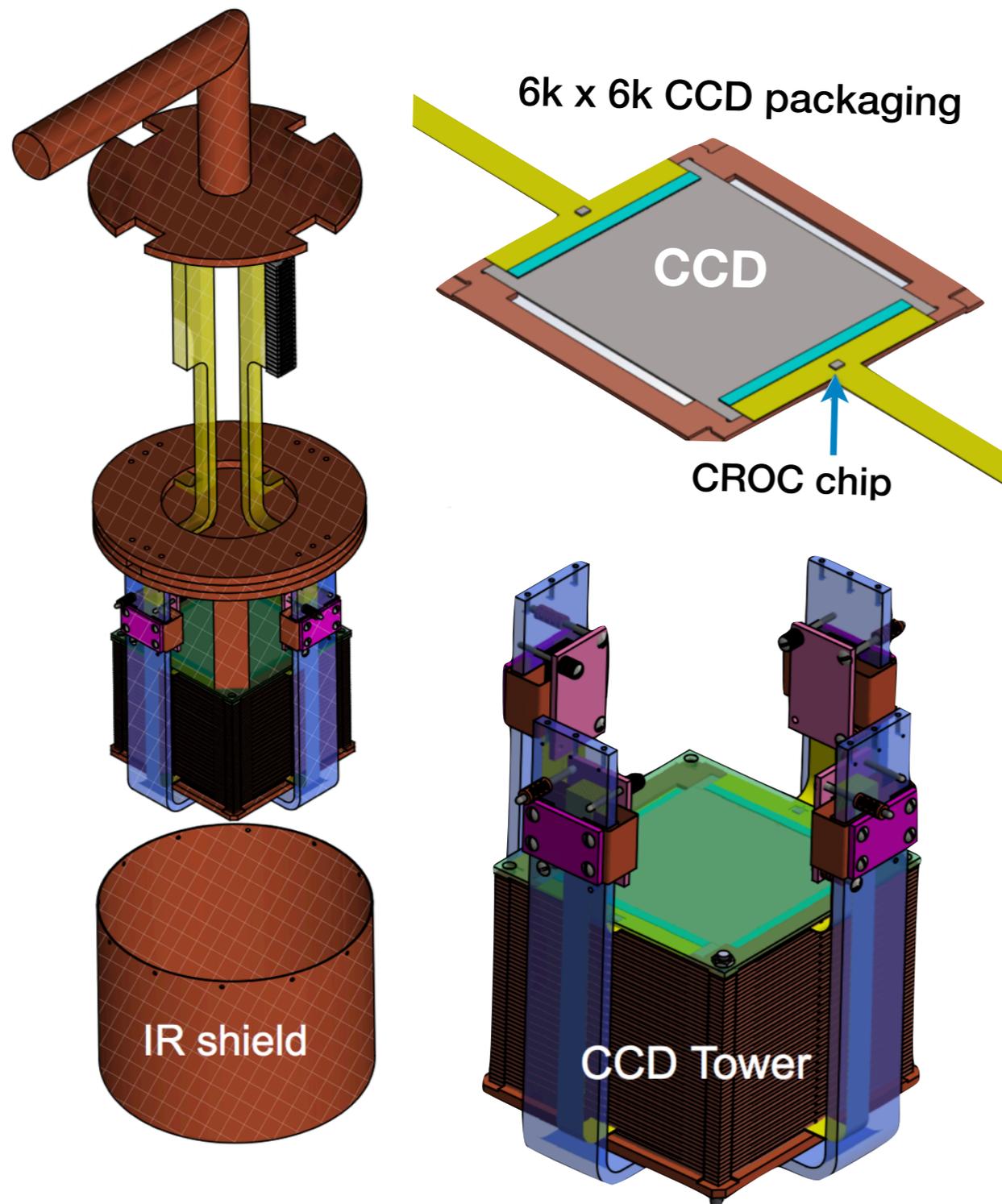
# Progresses on CCD production

- Silicon Ingot production by TOPSIL, Denmark
- Now at Boulby Underground Laboratory waiting for wafering by Shin-Etsu Handotai Europe (UK)



- To be shipped to DALSA (Canada) by sea: 8-15 days journey in a custom-made shielded container
- DAMIC-M Skipper CCDs  
Pre-production ~spring 2020

# Detector design and CCD tests

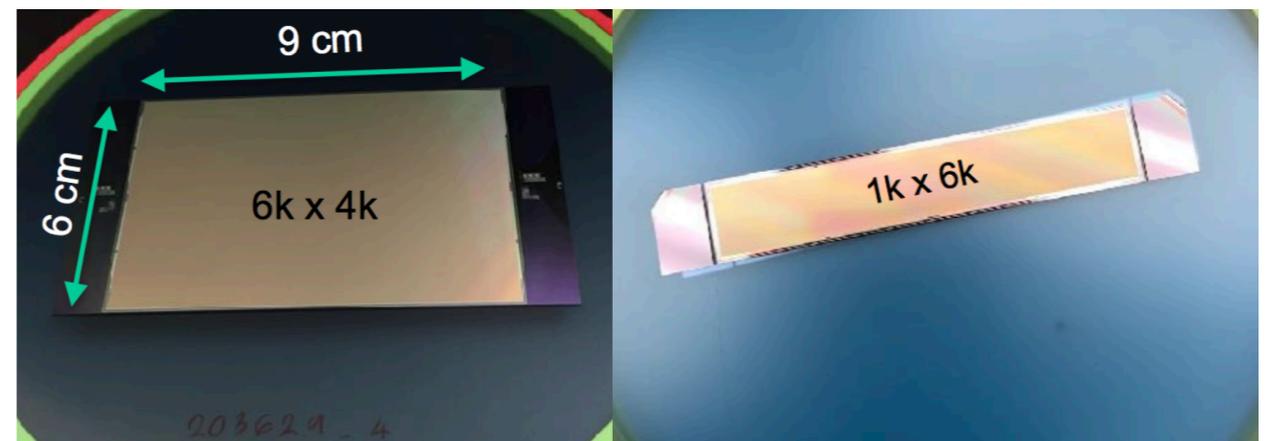
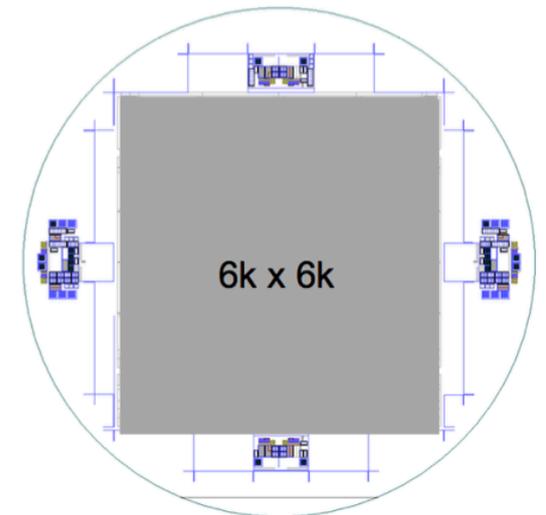


Design will be finalised by 2021

- ▶ Improvement of packaging procedures
- ▶ Evaluating new low bkg flex cables vs pico-coax (similar to Majorana's ones)
- ▶ Three prototype skipper-CCDs (675um thick) produced to test different readout amplifiers

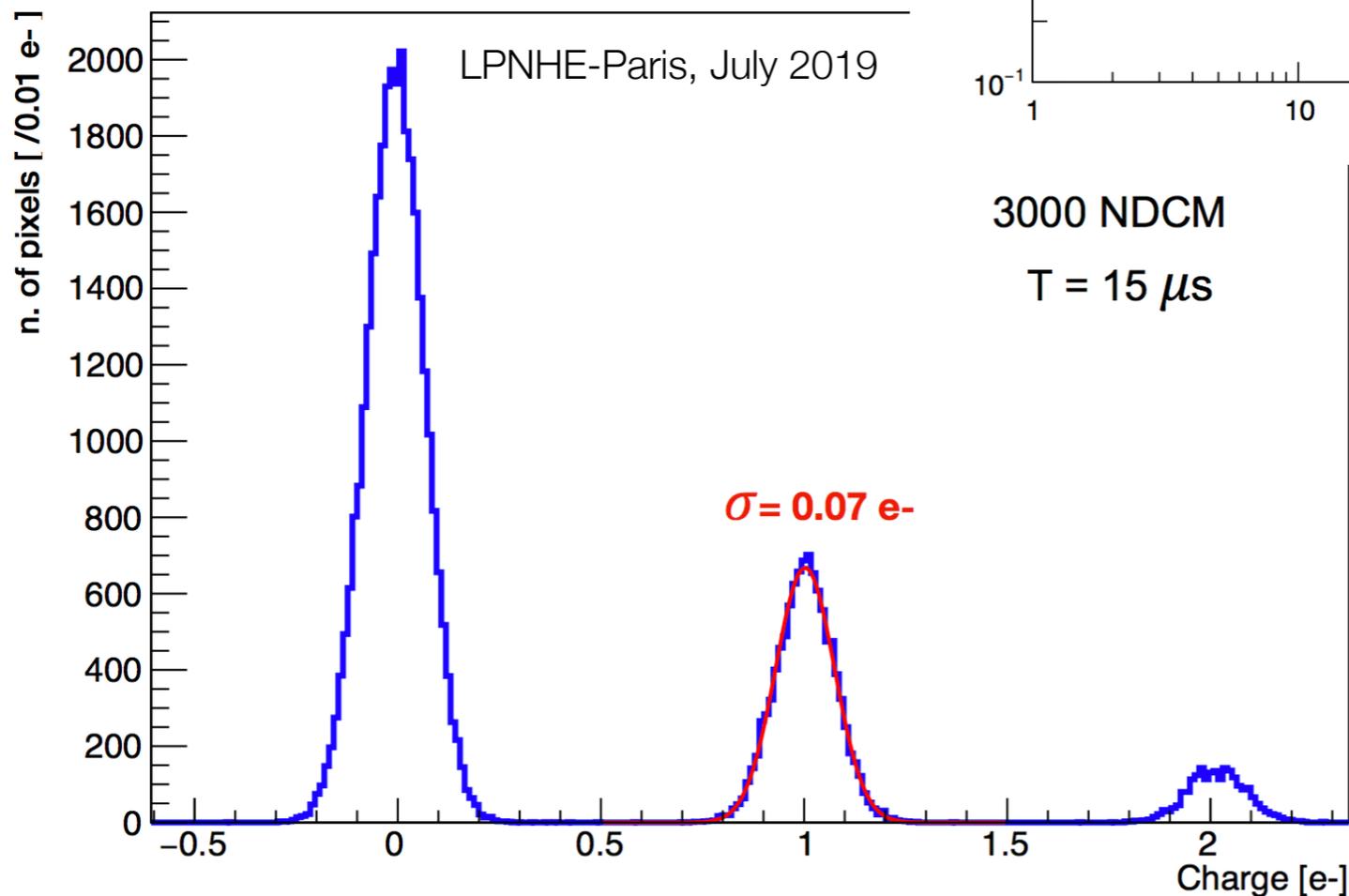
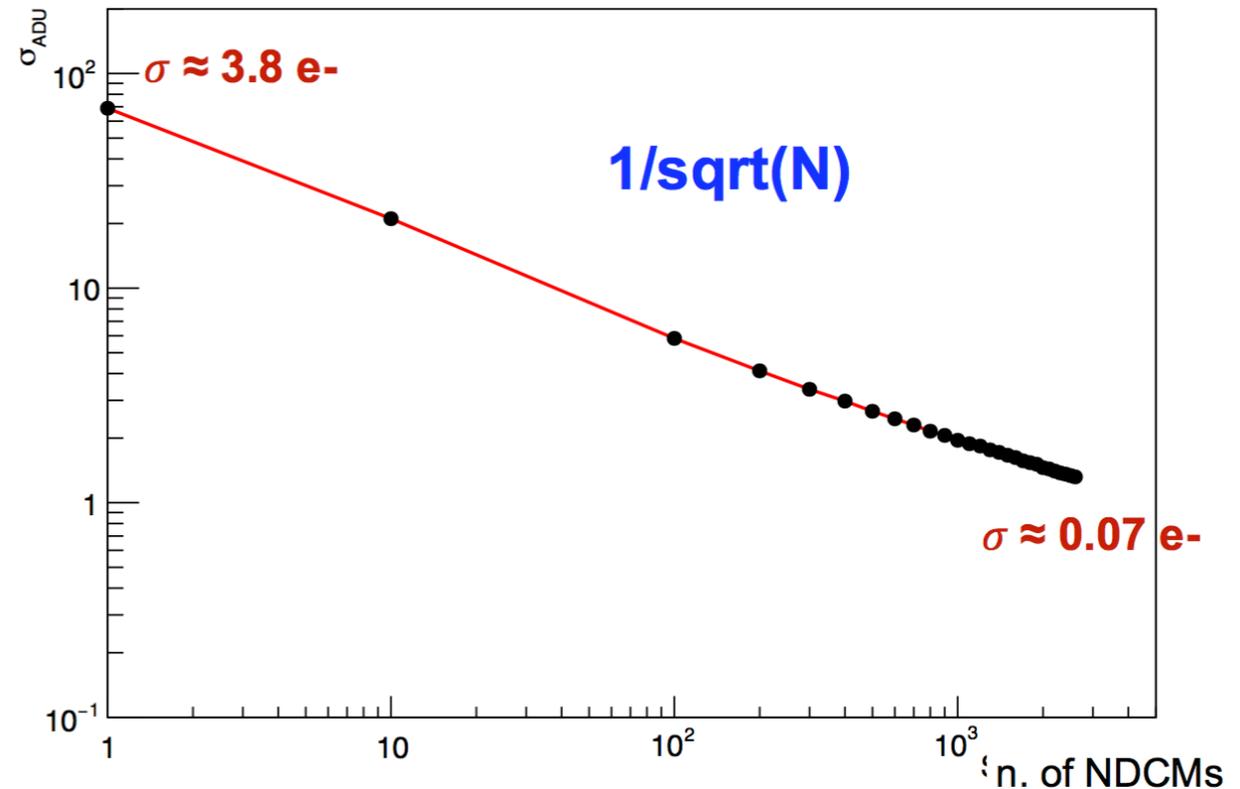
DAMIC-M skipper CCDs :  
6k x 6k pixels (i.e. 9cm x 9cm)

Design by S. Holland (LBNL),  
fabricated by Teledyne/DALSA

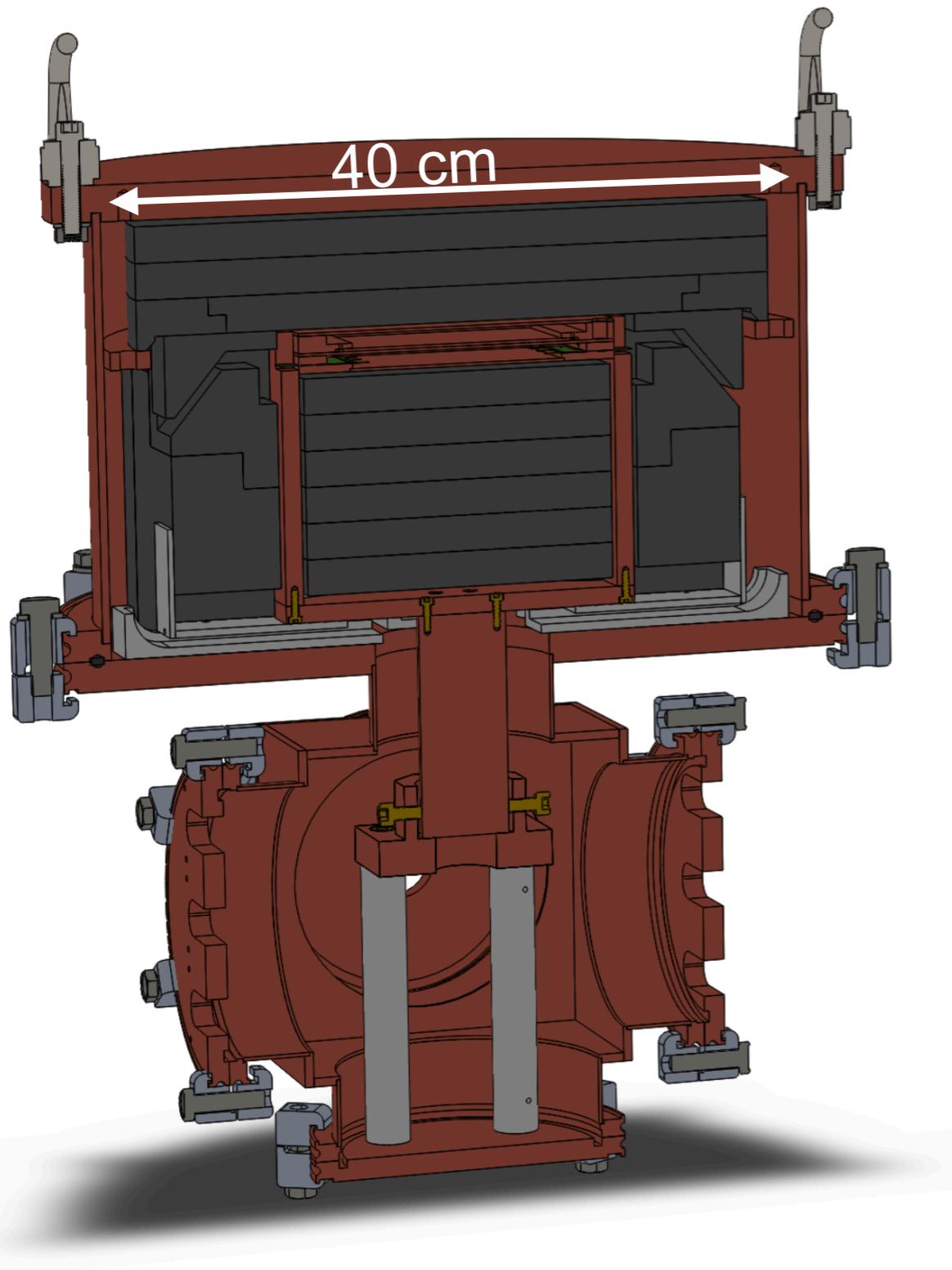


# DAMIC-M skipper CCD performance

Single electron resolution obtained for a wide range of integration times ( $T = 4 - 20 \mu\text{s}$ ) allowing optimization with respect to electronics noise



# “Low Background Chamber” prototype @ LSM



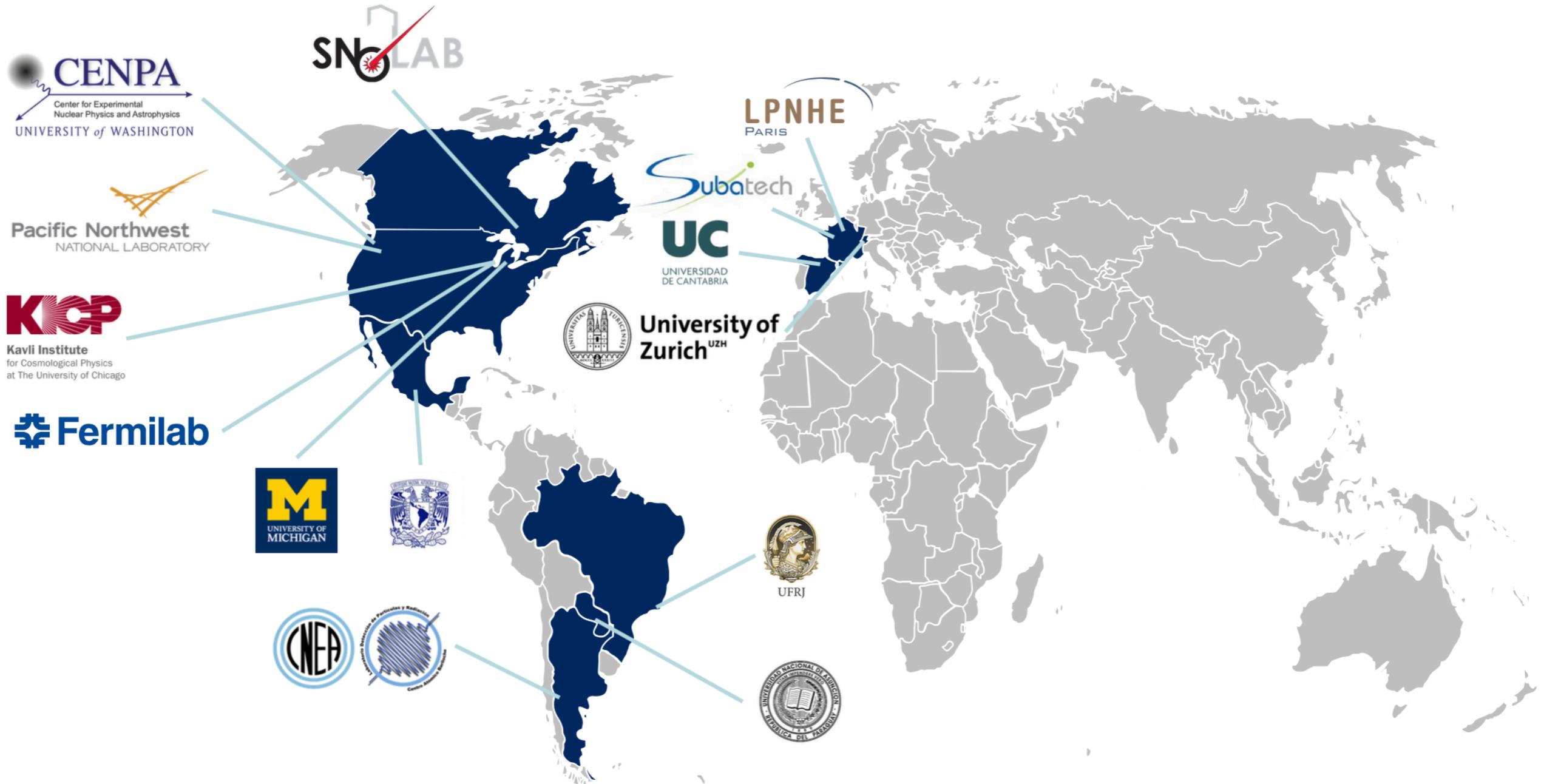
- ▶ Small detector hosting a 6k x 6k skipper-CCD
- ▶ Characterisation of DAMIC-M CCDs with low background (few dru)
- ▶ Study of  $^{32}\text{Si}$  background,  $^{210}\text{Pb}$  surface;
- ▶ Tests of new CCD packaging
- ▶ dark current and skipper CCD tests
- ▶ Installation at LSM beginning 2020

# Conclusions

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- ▶ **DAMIC has proved the CCD technology as a competitive technique for the search of low-mass Dark Matter candidates :**
  - ▶ Sensitivity for WIMPs masses below 10 GeV
  - ▶ Constrain Hidden photon absorption ( $m < 10$  eV) and electron scattering ( $< 5$  MeV)
- ▶ **Post-DAMIC R&D at SNOLAB:** background measurements for future Si-based experiments and CCD production and handling improvements under evaluation
- ▶ **Next phase: A kg-size detector at LSM (France)**
  - sub-electron resolution + increased detector mass + improved bkg to reach **eV energy thresholds** and **sensitivity to MeV DM candidates in the hidden sector**
  - Detector design, material procurements and CCD tests ongoing
  - A “low-background chamber” at the LSM taking data this year

# DAMIC Collaboration



**Thank you!**