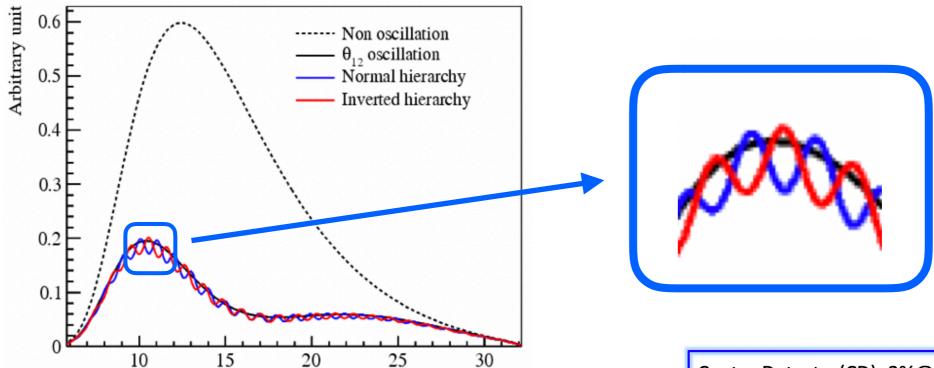


Summary WG4 (Experimental techniques)

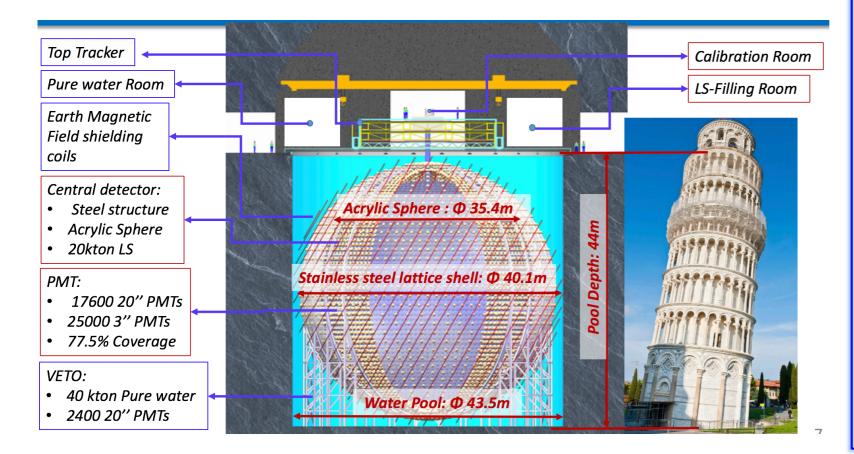
Davide Sgalaberna (CERN)
XVth Rencontres du Vietnam
9th August 2019

Cong Guo (IHEP)

JUNO



L/E (km/MeV)



Center Detector(CD): 3%@1MeV

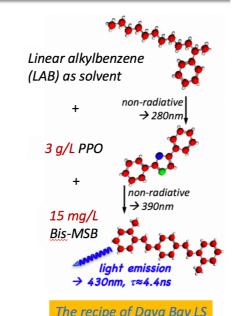
- Liquid Scintillator
 - Light yield >1100P.E./MeV;
 - Attenuation length > 22m@430mn;
- ➤ PMT
 - Full coverage;
 - High quantum & collection efficiency;
- > Acrylic sphere
 - High Transparency;
 - Good mechanical strength for >20 years;
- Low radioactivity
 - ²³⁸U,²³²Th≤10⁻¹⁵g/g, ⁴⁰K<10⁻¹⁶g/g for LS in reactor neutrino measurement ;
 - ²³⁸U,²³²Th, ⁴⁰K ≤ 1ppt for acrylic;

Cong Guo (IHEP)



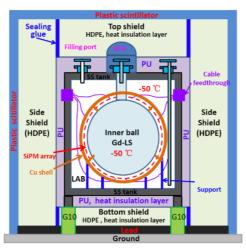
R&D status: Liquid Scintillator

- > Using a recipe optimized from Daya Bay's experience;
- > Tested and changed to be more suitable for JUNO;
- Higher light yield and more transparent;
 - 2.5g/L PPO;
 - 3mg/L Bis-MSB;
 - Filtration with Al₂O₃ column(Based on the "absorption" technique to remove the impurities and increase the attenuation length of LAB);
- Low background;
 - Distillation: Remove heavy metal and improve transparency;
 - Water extraction: Remove U/Th/K;
 - Gas stripping: Remove Ar/Kr/Rn;



JUNO-TAO

- Taishan Antineutrino Observatory (TAO), a ton-level, high energy resolution LS detector at 30 m from the core, a satellite exp. of JUNO.
- Measure reactor neutrino spectrum w/ sub-percent E resolution.
 - model-independent reference spectrum for JUNO
 - a benchmark for investigation of the nuclear database
- Ton-level Liquid Scintillator (Gd-LS)
- Full coverage of SiPM w/ PDE > 50%
- Operate at -50 °C (SiPM dark noise)
- 4500 p.e./MeV
- Taishan Nuclear Power Plant, 30-35 m from a 4.6 GW th core
- 2000 IBD/day (4000)
- Online in 2021



The recipe of Daya Bay LS

R&D status: Calibration system

> The calibration system need to accurately address both the nonuniformity and non-linearity in the detector energy response;

fully(~78%) cover CD;

> Implosion protection;

> Waterproof potting;

electronics.

> Custom-made divider and

> Stringent quality control;

- Energy scale uncertainty < 1%;</p>
- > Four complementary subsystems:
 - 1-D: Automated calibration unit(ACU);
 - → Scan the central axis;
 - 2-D: Cable loop system(CLS);
 - → Scan vertical planes; Guide tube calibration system(GTCS);
 - → Scan CD outer surface;
 - 3-D: Remotely operated vehicle(ROV);
 - → Full detector scan;
- > Radioactive Sources:
 - v, e+, n sources
 - 40K、54Mn、60Co、137Cs、22Na、68Ge、241Am-Be, 241Am-13C or 241Pu-13C、252Cf;

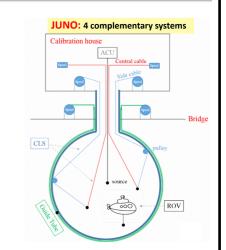
> Two sizes of PMTs will be used to • 17600 20" PMTs for CD (~75%) + 2400 20" PMTs for Veto; • 25000 3" PMTs(~2.5%); 6. Potting



R&D status: JUNO PMTs

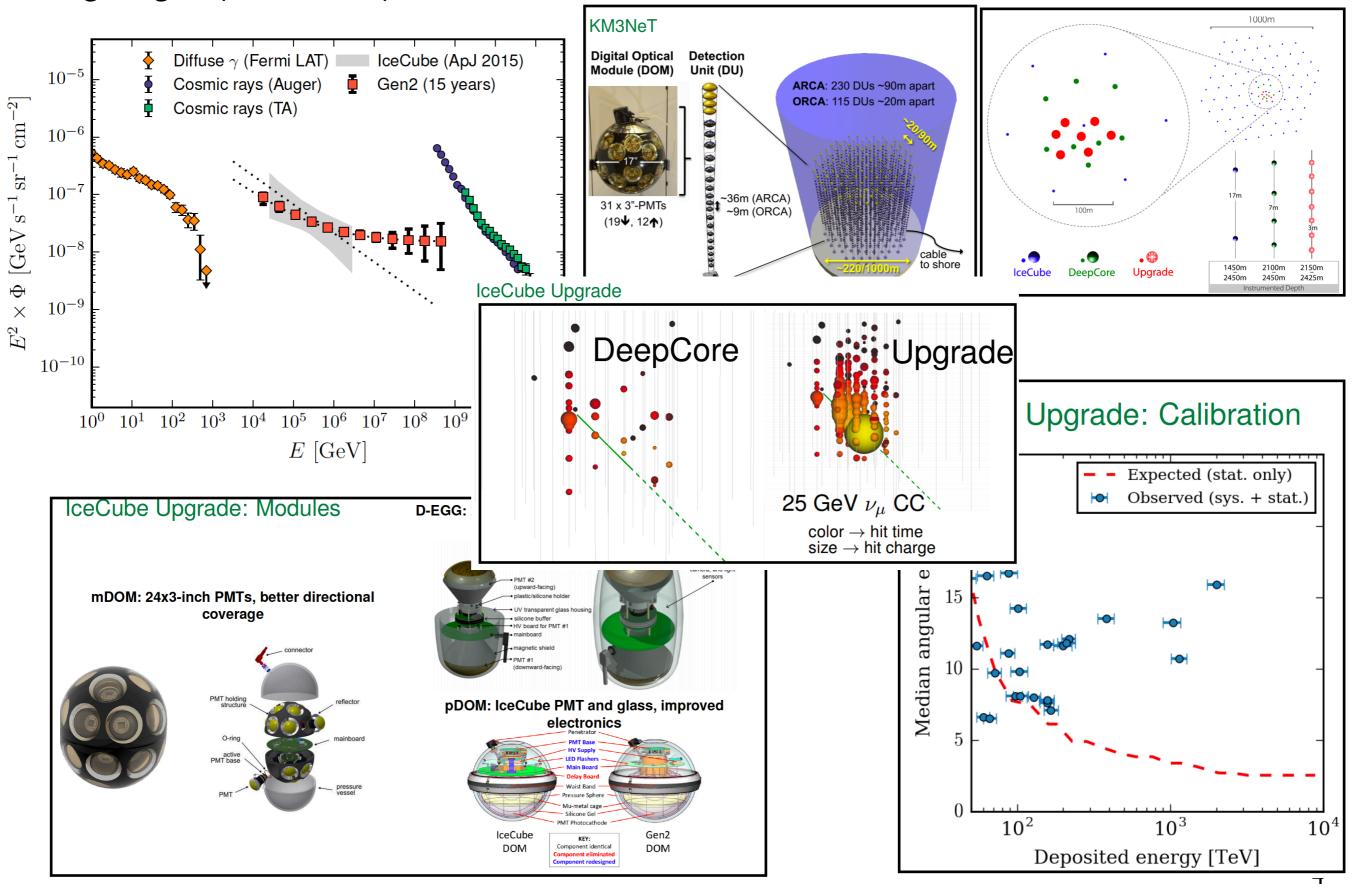
- Improve the energy scale precision, in particular, the coupling of non-linearity and non-uniformity.
 - SPMTs almost always work at SPE mode for IBD events and are expected to have almost zero dynamic range, hence virtually no non-linearity, thus providing a linear reference to LPMT;
- > 25000 3" PMTs contracted to HZC, ~19000 produced, ~15000 tested and accepted,





IceCube upgrade & Km3Net

J.Highnight (U.Alberta)



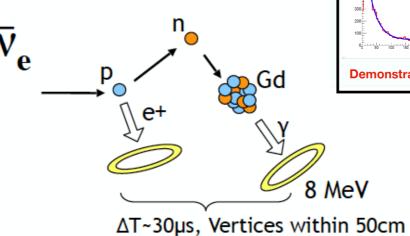
SuperK towards Gadolinium

Nakajima-san (ICRR, U.Tokyo)

Advantage of Gd:

- Large n-capture cross section:
 - 90% of Gd capture efficiency at 0.2% loading of Gd2(SO4)3 (corresponds to 100 ton/SK)
- Large released energy of ~8MeV
 - Well above most of natural radioactivity and the SK trigger threshold

Strongly tag electron antineutrinos by prompt (e+) and delayed (n-Gd) coincidences



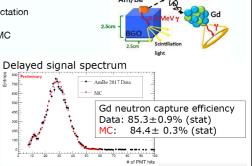
Neutron detection with Gd

- Test with "bulk" Gd loaded water in EGADS:
- AmBe neutron source w/ BGO crystal to detect 4.4 Me' "prompt" y signal

Astropart. Phys. 31, 320-328 (2009)

- Decay time constant consistent w/ expectation
- Energy distribution well reproduced by MC

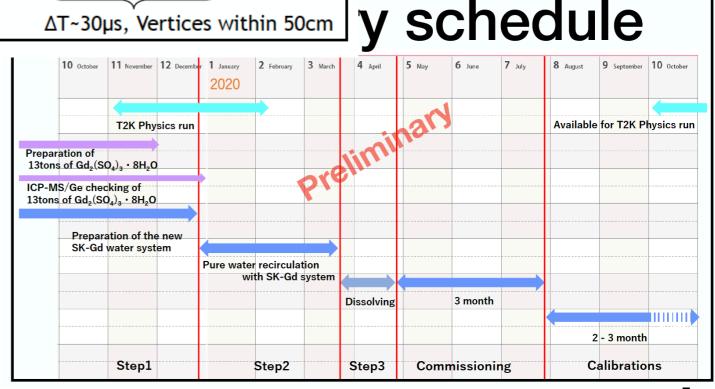
Gd₂(SO₄)₃8H₂O concentration



Demonstrated neutron detection with Gd in water Cherenkov detectors

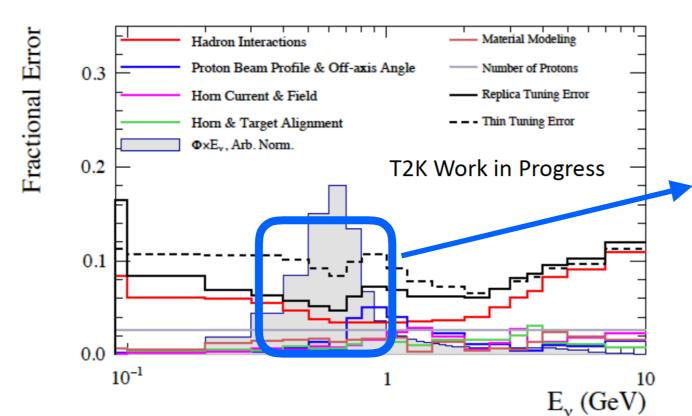
Goals of SK-Gd

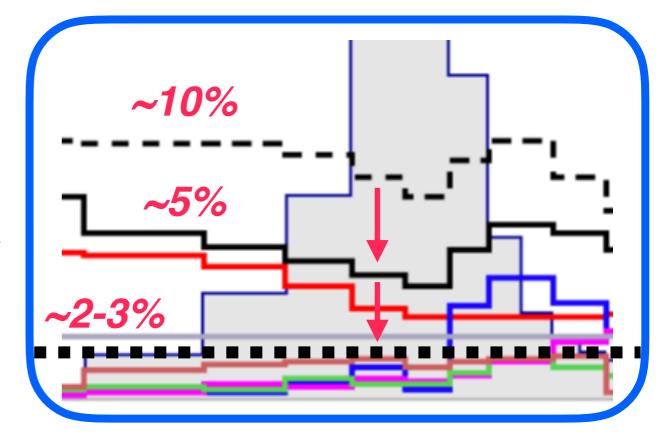
- First observation of Supernova Relic Neutrinos
- Improve pointing accuracy for galactic supernova
- Precursor of nearby supernova by Si-burning neutrinos
- Reduce proton decay background
- Neutrino/anti-neutrino discrimination (Long-baseline and atmospheric neutrinos)
- Reactor neutrinos

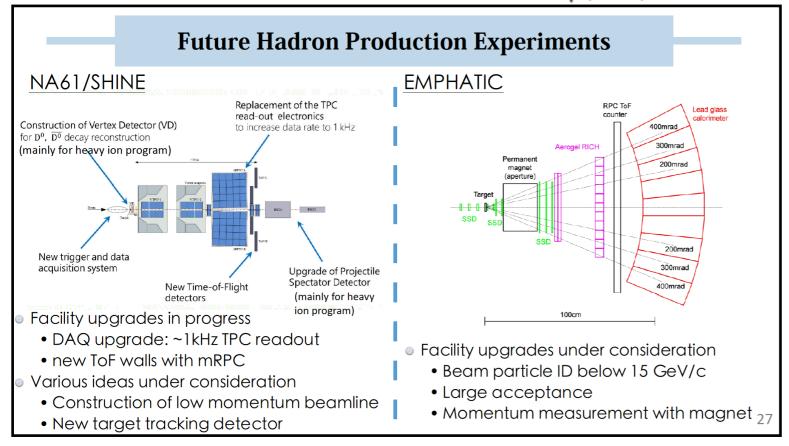


Neutrino flux measurement: Hadron production

Nagai-san (U.Colorado Boulder)



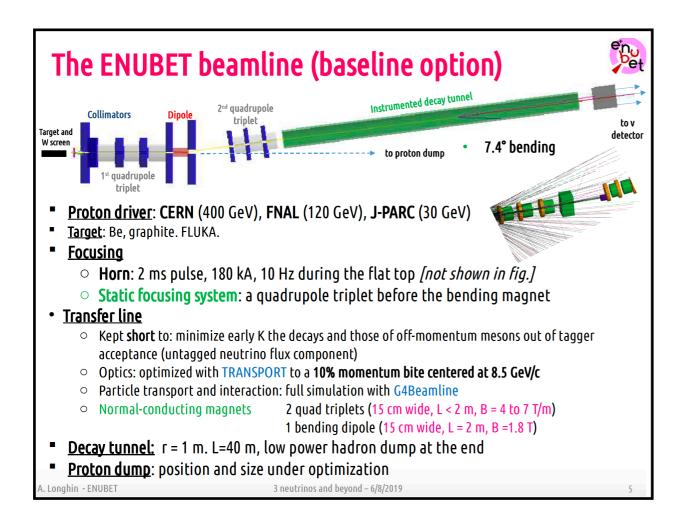


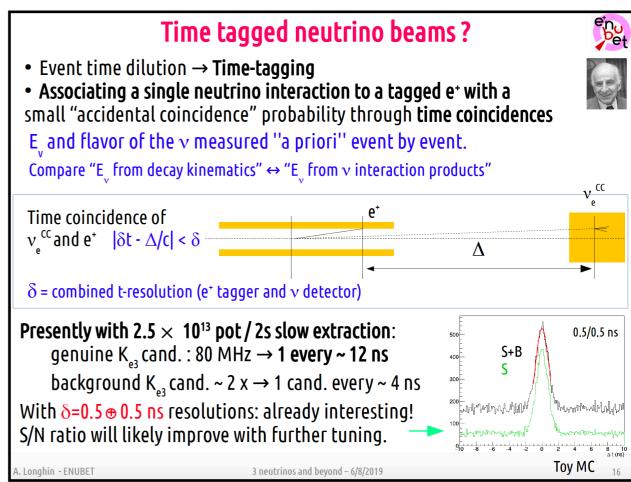


- Independent methods nu+e scattering, low-nu, momentum imbalance+neutrons, etc.) will be a good cross check
- Joint analysis of p-C/Ca/Be/... helfpul to better constrain the models?

ENUBET

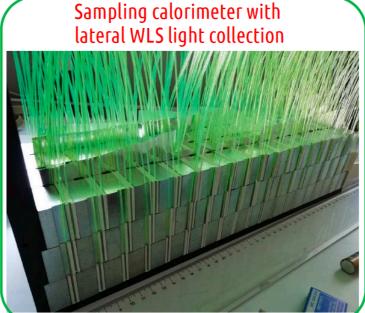
A.Longhin (U.Padova)





CERN PS test beam Nov 2016

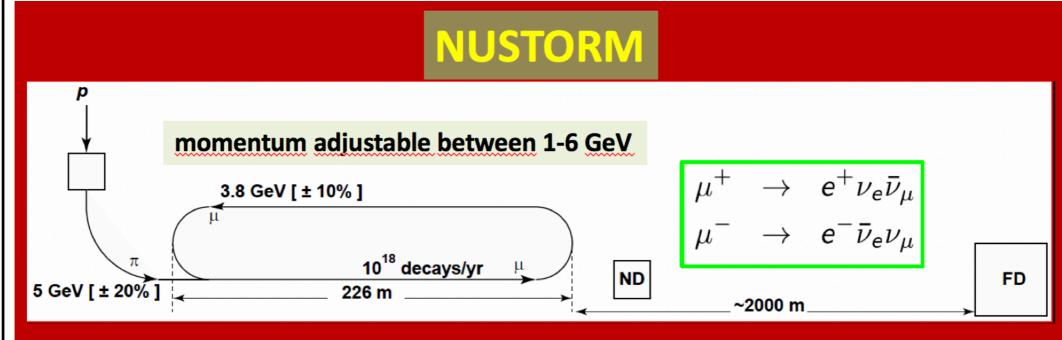






NuStorm

A.Blondel (LPNHE)



- -- first flash of muon (anti) neutrinos from the decay of injected pions.
- -- then muons from pion decay are stored in the ring.
- -- cross-sections to be measured in near detector similar to HyperK or DUNE ND (water, scintillator and Liq. Argon targets -- also probably iron)
 - → preferably magnetized for energy reconstruction?
- -- require precise instrumentation for number of stored muons and polarization
- -- storage time at 6 GeV is 5x144µs (fast extraction of short spill required)
- -- high field magnets would be highly beneficial by reducing the size of the facility.

Recommandations

A.Blondel (LPNHE)

submission to European Strategy for Particle Physics https://arxiv.org/abs/1812.06739

Future Opportunities in Accelerator-based Neutrino Physics

The Participants of the European Neutrino Town Meeting 22–24 October, 2018

CERN, 1 Esplanade des Particules, 1211 Geneva 23, Switzerland

Editors: Alain Blondel^a, Albert De Roeck^b, Joachim Kopp^c (full author list in the appendix)

(Dated: December 2018)

This document summarizes the conclusions of the Neutrino Town Meeting held at CERN in October 2018 to review the neutrino field at large with the aim of defining a strategy for accelerator-based neutrino physics in Europe. The importance of the field across its many complementary components is stressed. Recommendations are presented regarding the accelerator based neutrino physics, pertinent to the European Strategy for Particle Physics. We address in particular i) the role of CERN and its neutrino platform, ii) the importance of ancillary neutrino cross-section experiments, and iii) the capability of fixed target experiments as well as present and future high energy colliders to search for the possible manifestations of neutrino mass generation mechanisms.

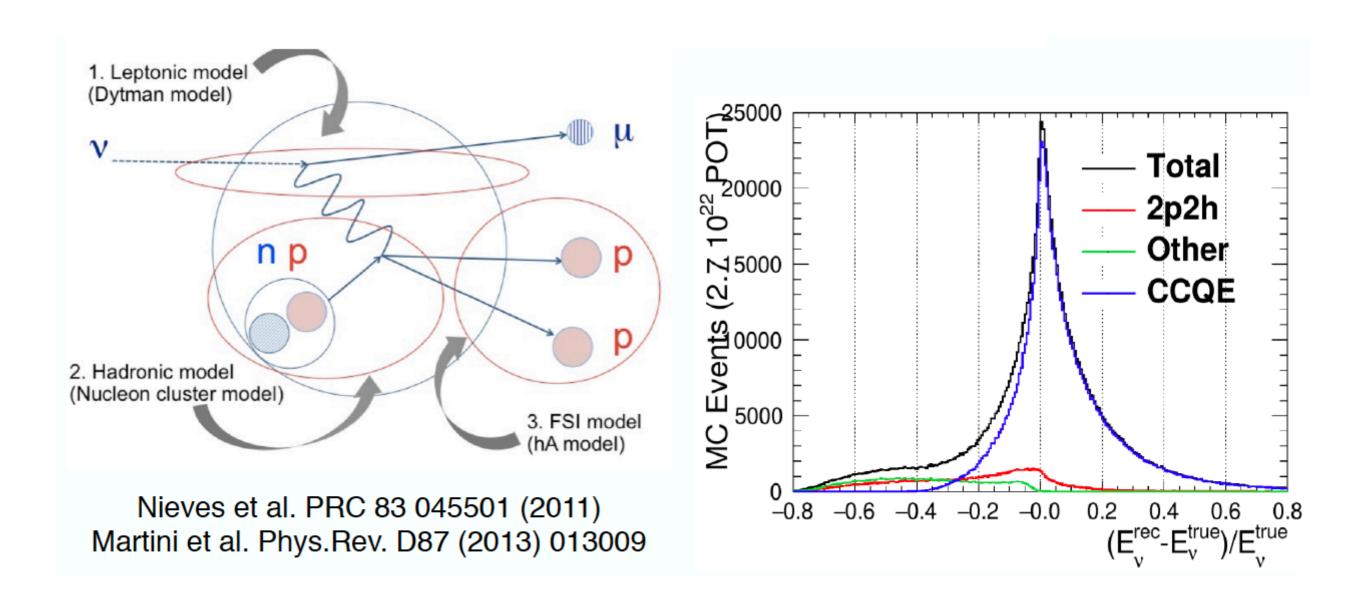
RECOMMANDATION:

Extracting the most physics out of DUNE and HyperK will require ancillary experiments:

- CERN should continue improving NA61/SHINE towards percent level flux determinations;
- 2) a study should be set-up to evaluate the possible implementation, performance and impact of a percent-level electron and muon neutrino cross-section measurement facility (based one.g. ENUBET or NuSTORM) with conclusion in a few years;

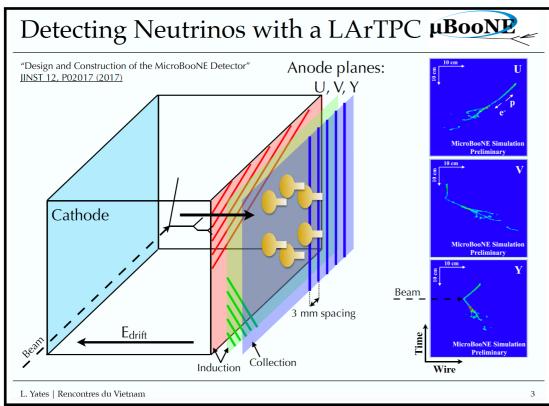
Complementarity between different techniques is very important

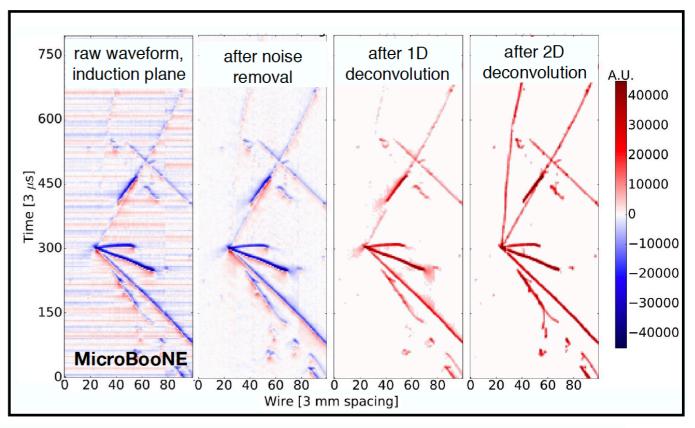
Precise detection of neutrino interactions

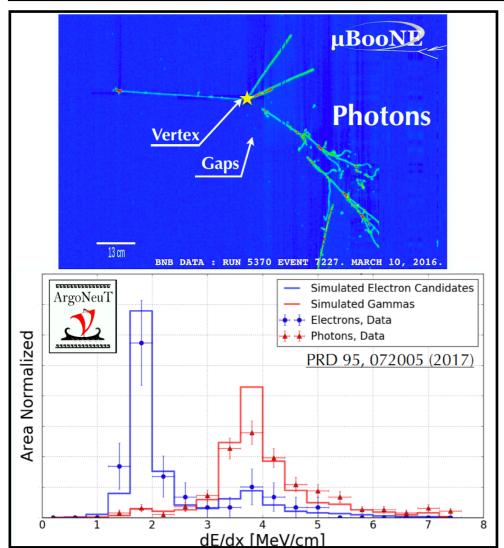


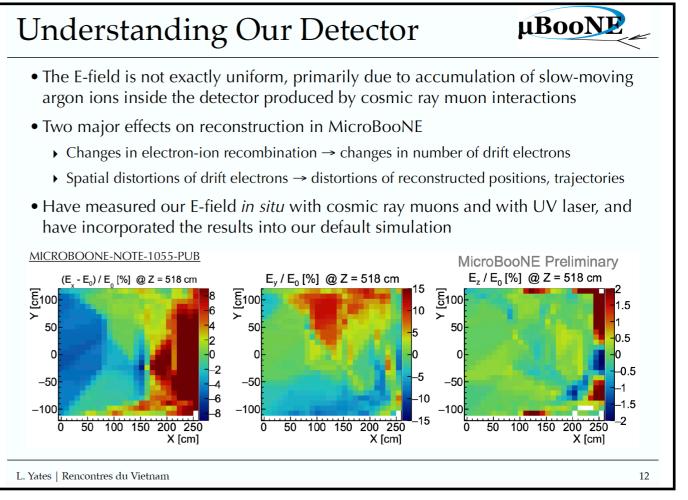
This is very tough!
That's why we need more precise (but big) detectors

L. Yates (MIT) LAr detectors @SBN / MicroBooNE

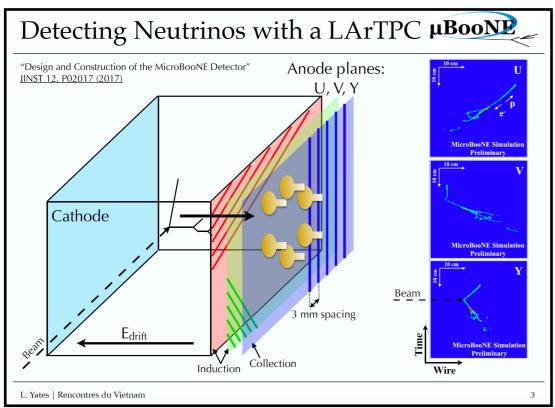


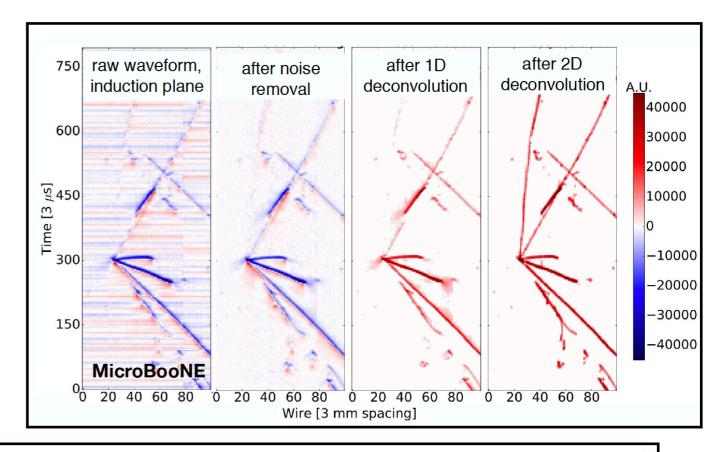


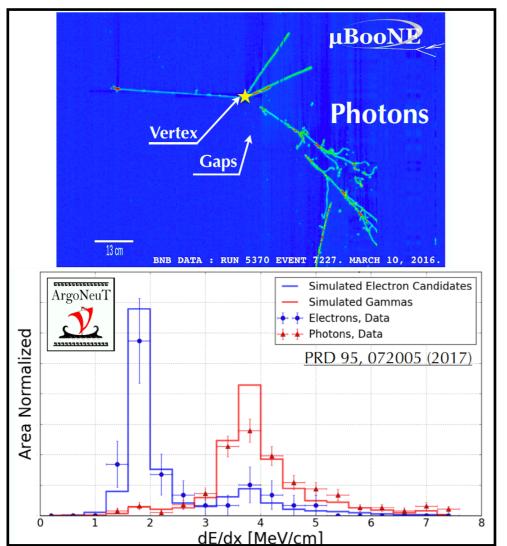




L. Yates (MIT) LAr detectors @SBN / MicroBooNE



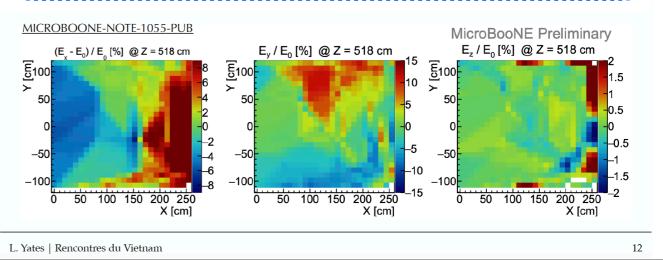




Understanding future detectors

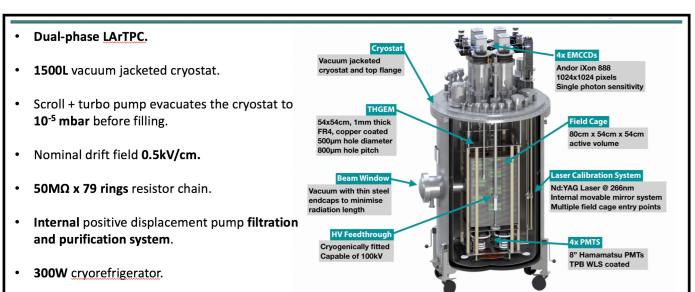
Issues understood in current and near-future LAr detectors will be useful for farther-future ones too.

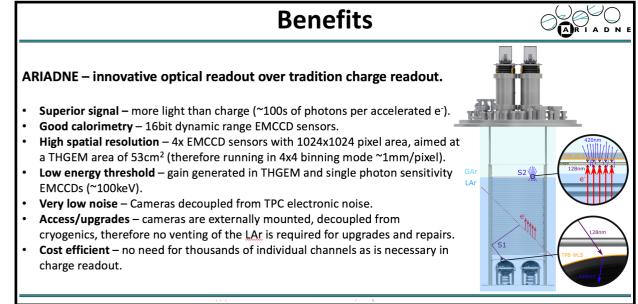
Development of reconstruction tools, understanding of common detector response issues, etc.

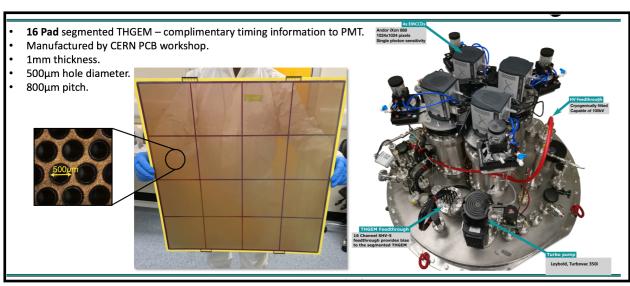


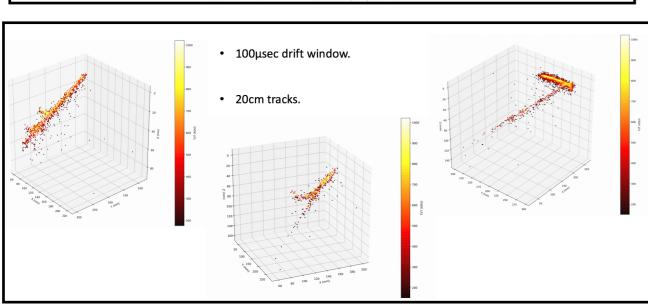
B.Philippou (U.Liverpool)

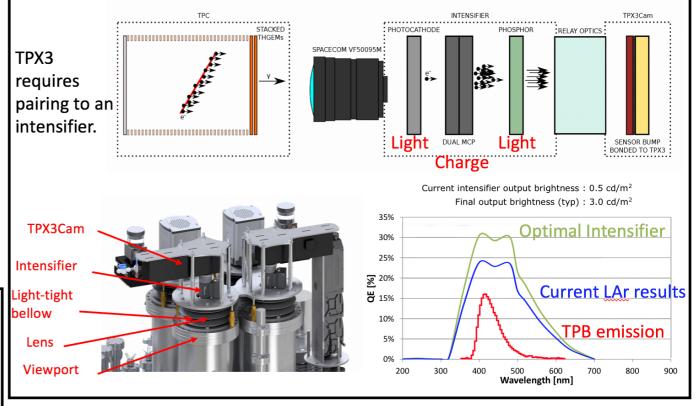
ARIADNE







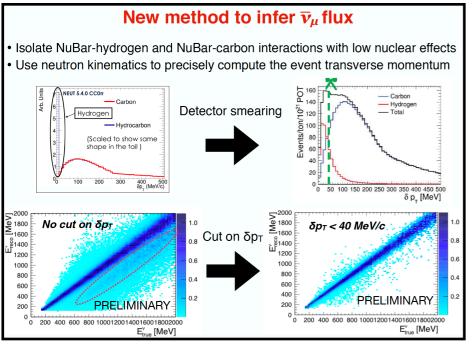


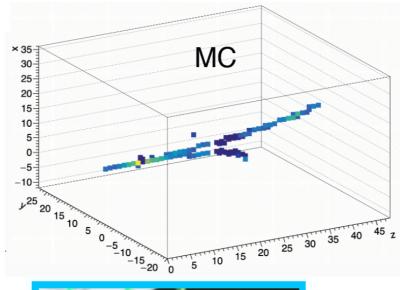


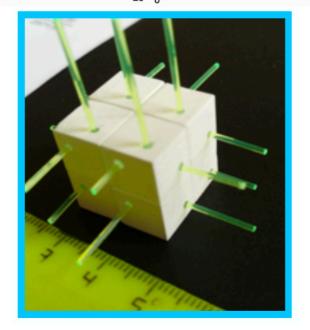
A completely new concept for LAr readout, that could simplify a lot the detector configuration

SuperFGD for T2K/HK and DUNE

D.Sgalaberna (CERN)







Characterization of the SuperFGD concept

- Prototype 5x5x5 cm³, 1.3 m WLS fibers (Al-based paint at fiber end)
- \bullet Exposure to a 6 GeV π test beam at CERN
- Multi Pixel Photon Counter (MPPC)based readout



| Entries | 10661 | 1062 | 1062 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 1060 | 106

- Average light yield ~ 41 p.e. / fiber / cube (MIP)
 Light cross-talk between adjacent cubes ~ 3.7%
 Very good intrinsic time resolution (measured)
 - + $\sigma_t \sim 0.95$ ns (1 channel, 1 cube)

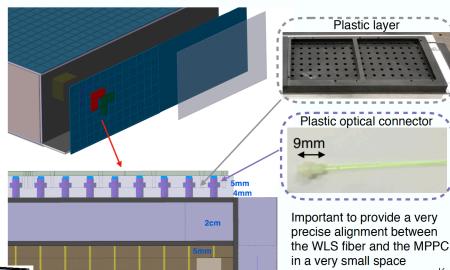
with a 5 GHz waveform digitizer)

+ $\sigma_t \sim 0.65$ ns (2 channels, 1 cube)

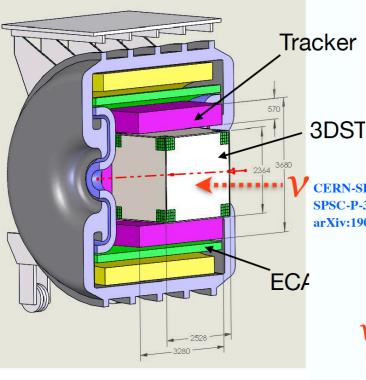
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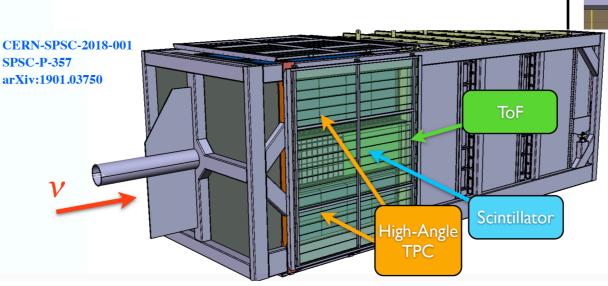
The SuperFGD detector

Very compact design increases HA-TPC acceptance to low-momentum particles



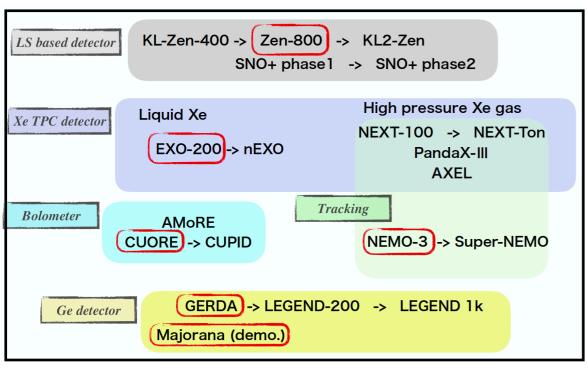
Neutron detection with kinematics measurement is a new paradigm to be demonstrated with data!

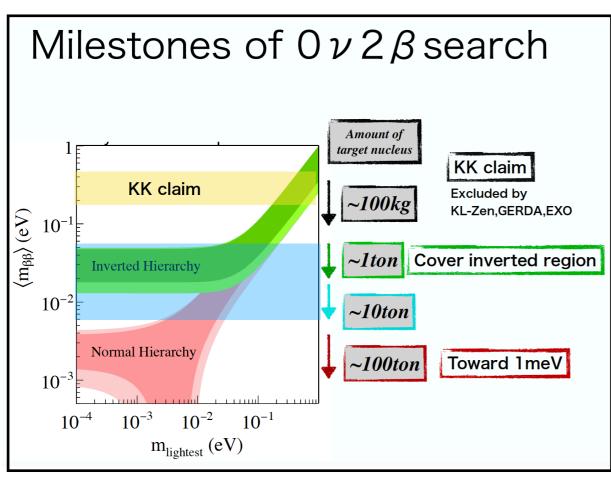


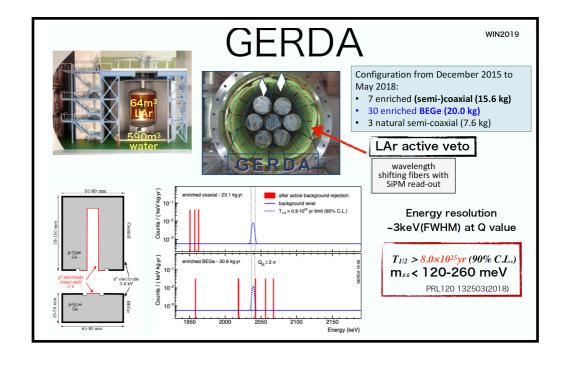


Nu-less ββ decay searches

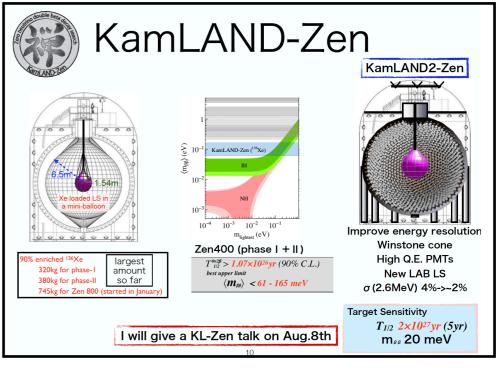
Ueshima-san (U.Tohoku)







SuperNEMO
AMoRE
NEXT
PandaX-III
SNO+
KL2-Zen
LEGEND
CUPID
nEXO

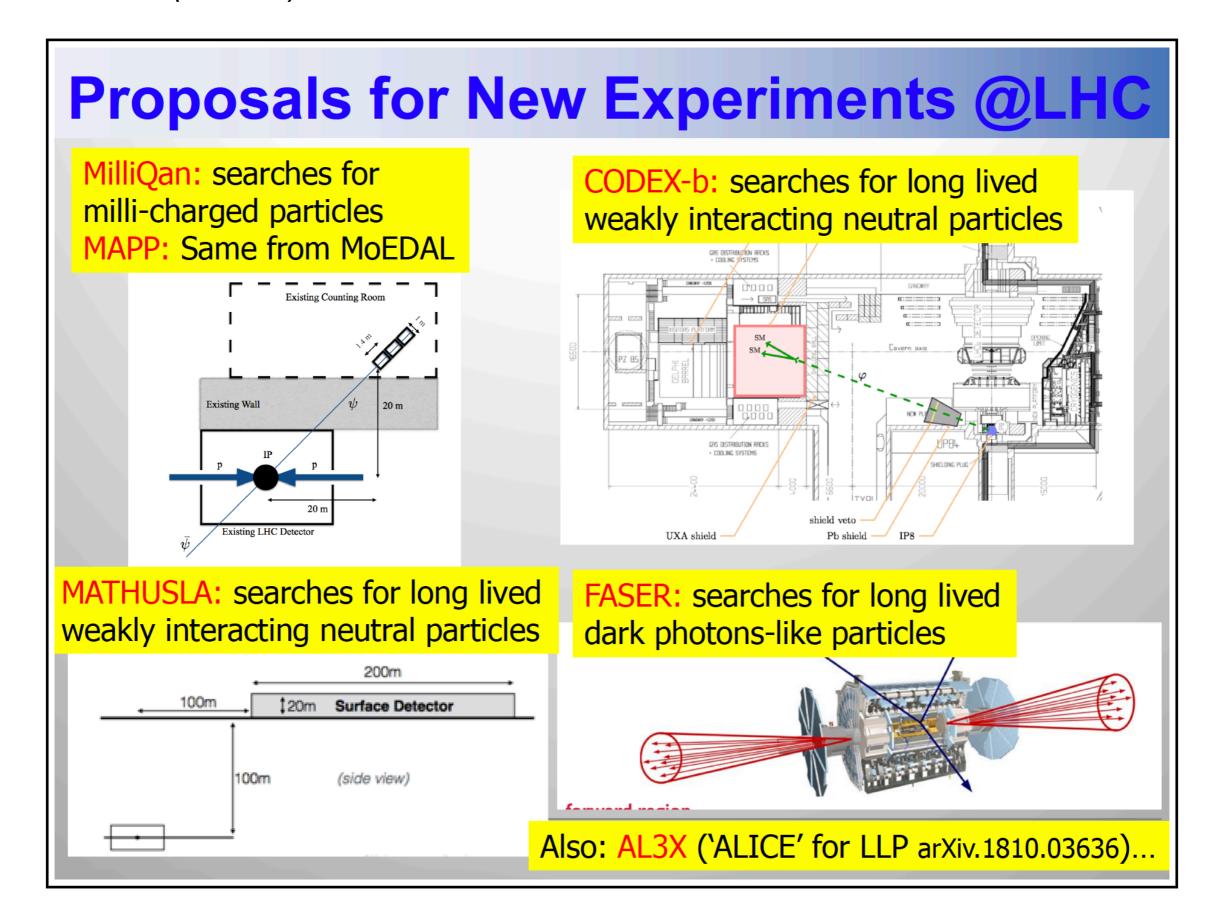


"How much could a ~100ton detector" cost?"

"Maybe like DUNE / HyperK"

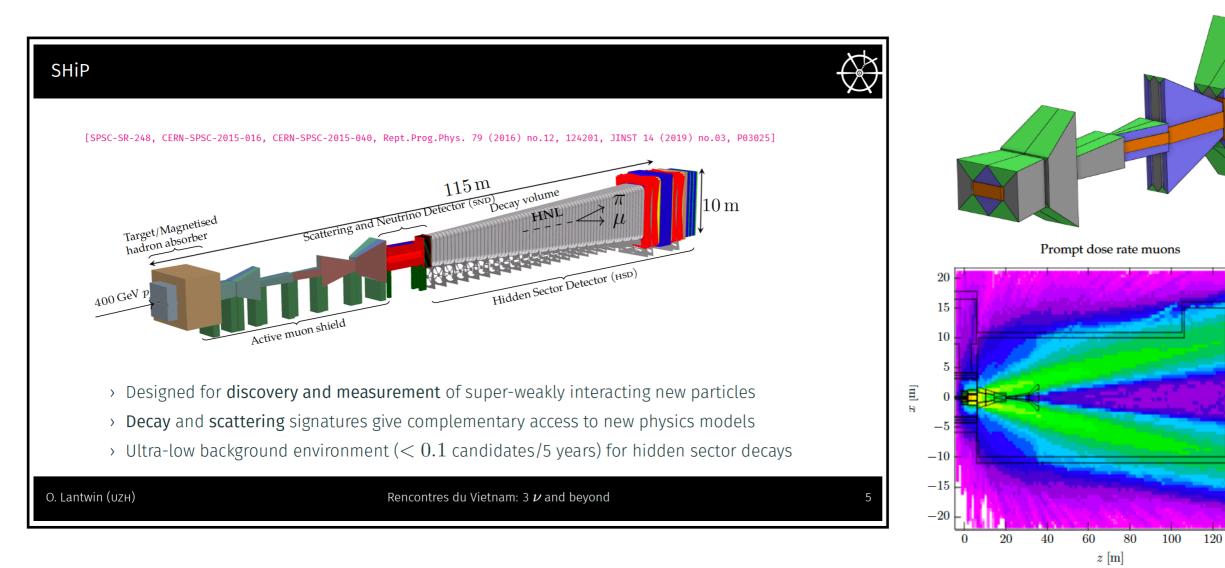
FASER & Mathusla

A. de Roeck (CERN)



HNL @SHiP

O. Lantwin (U.Zurich)



- > < 0.1 background events expected o 3 candidates sufficient for discovery
- > Can measure decay vertex, invariant mass, impact parameter of signal candidate
 - > mass, charge and flavour of new particles measurable
 - > redundant background rejection (+background tagger, timing,...)

A.Baldini (INFN Pisa)

µ→eγ @MEG

