



## Low scale leptogenesis

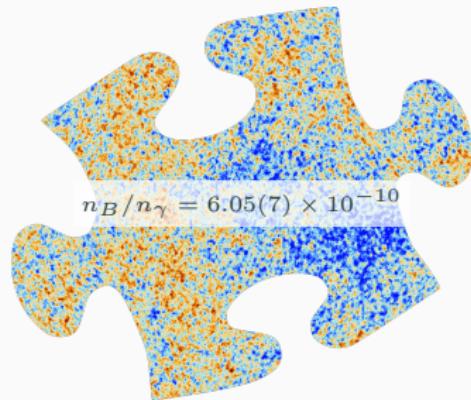
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Juraj Klarić

3 Neutrinos and beyond, XV. Rencontres du Vietnam, August 2019

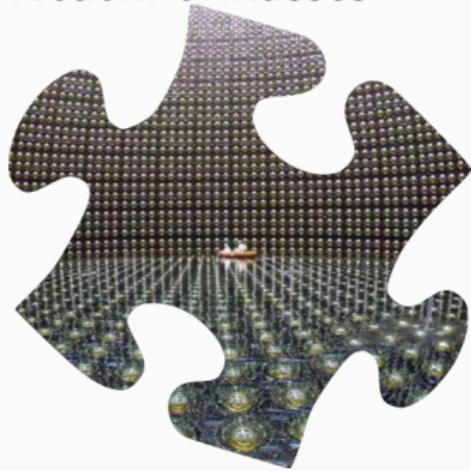
## Some of the shortcomings of the standard model:

BAU baryon asymmetry of the universe



[Planck collaboration]

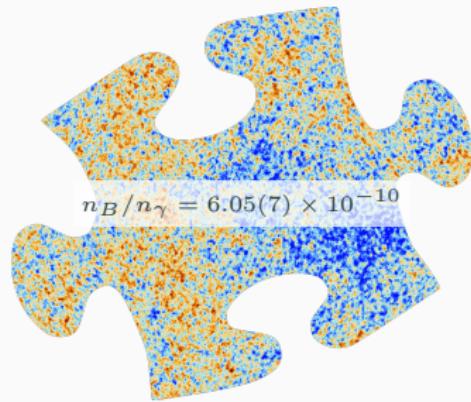
Neutrino masses



[Super-Kamiokande]

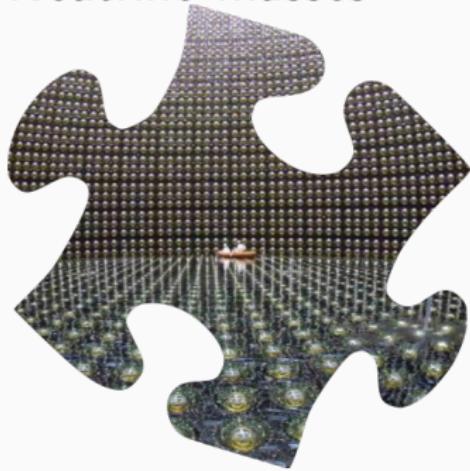
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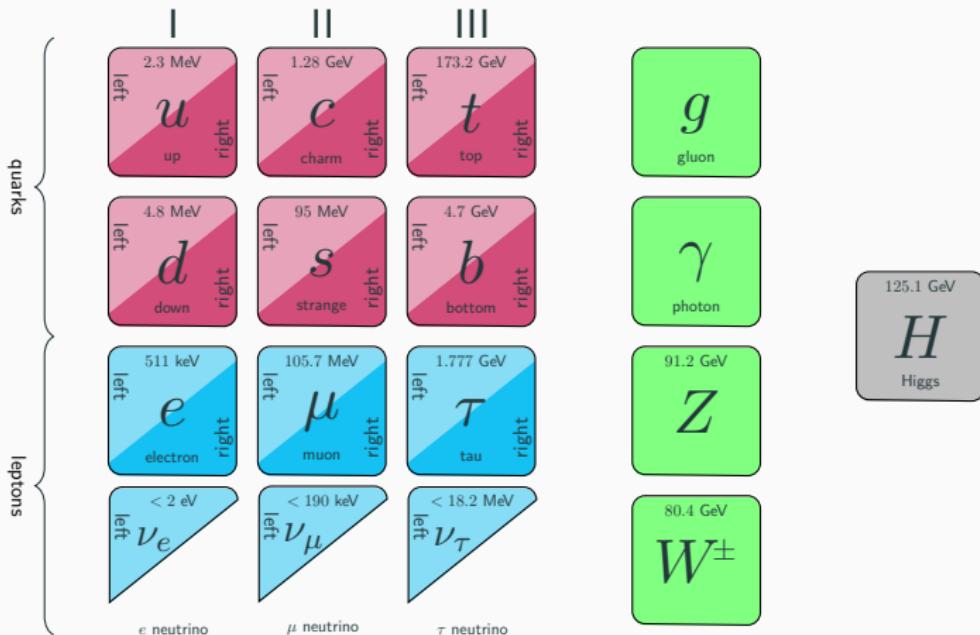
Neutrino masses



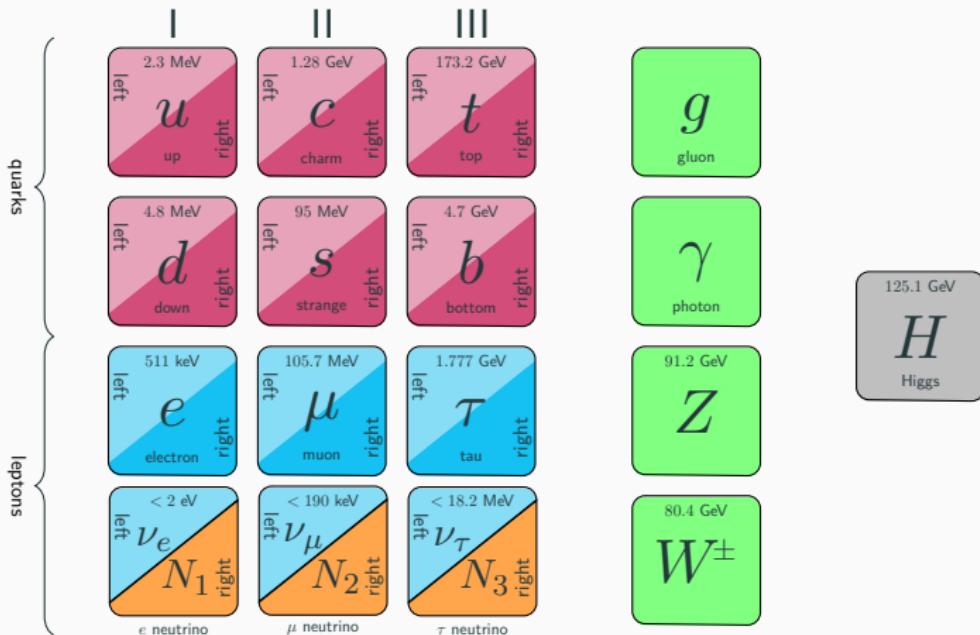
[Super-Kamiokande]

*Is there a way to explain both?*

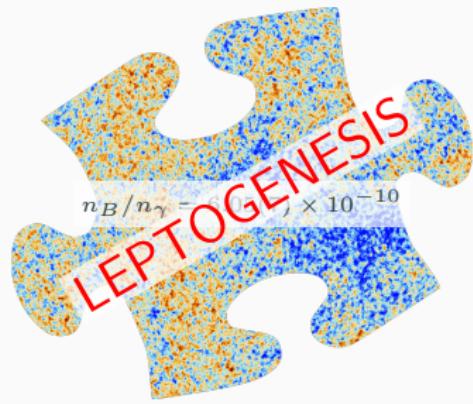
# Standard Model



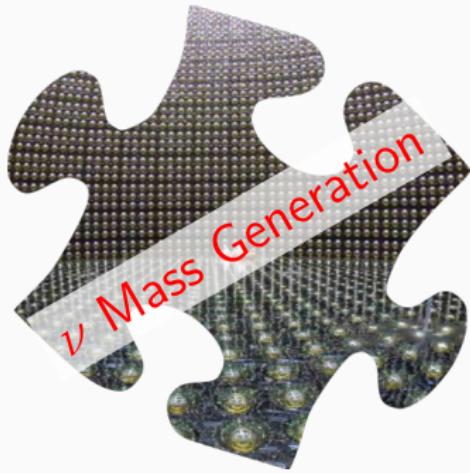
# Standard Model



## BAU baryon asymmetry of the universe



## Neutrino masses



# The “Seesaw” Relation

- Dirac Mass  $m_D = vY^\dagger$
- Right handed neutrino (RHN) Majorana mass  $M_M$

$$\mathcal{L} \supset \frac{1}{2} \begin{pmatrix} \overline{\nu_L} & \overline{N} \end{pmatrix} \begin{pmatrix} 0 & m_D \\ m_D^T & M_M \end{pmatrix} \begin{pmatrix} \nu_L \\ N \end{pmatrix}$$

## Active neutrino masses

$$m_\nu = -m_D M_M^{-1} m_D^T$$

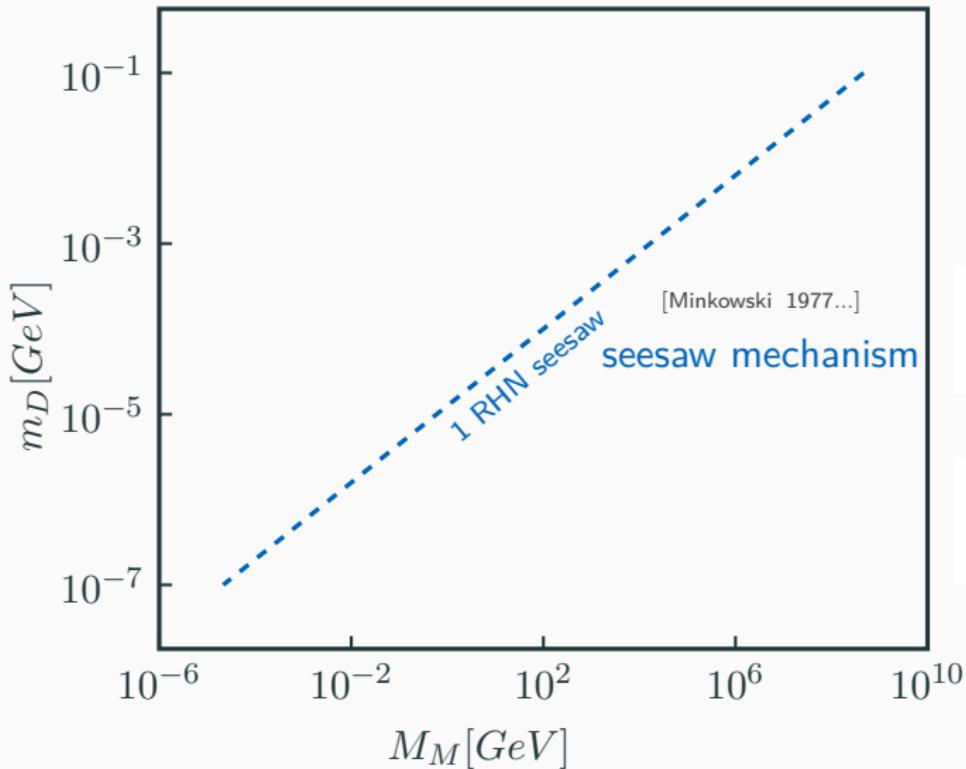
see talk by Sin Kyu Kang

for other mechanisms see talk by R. Ruiz

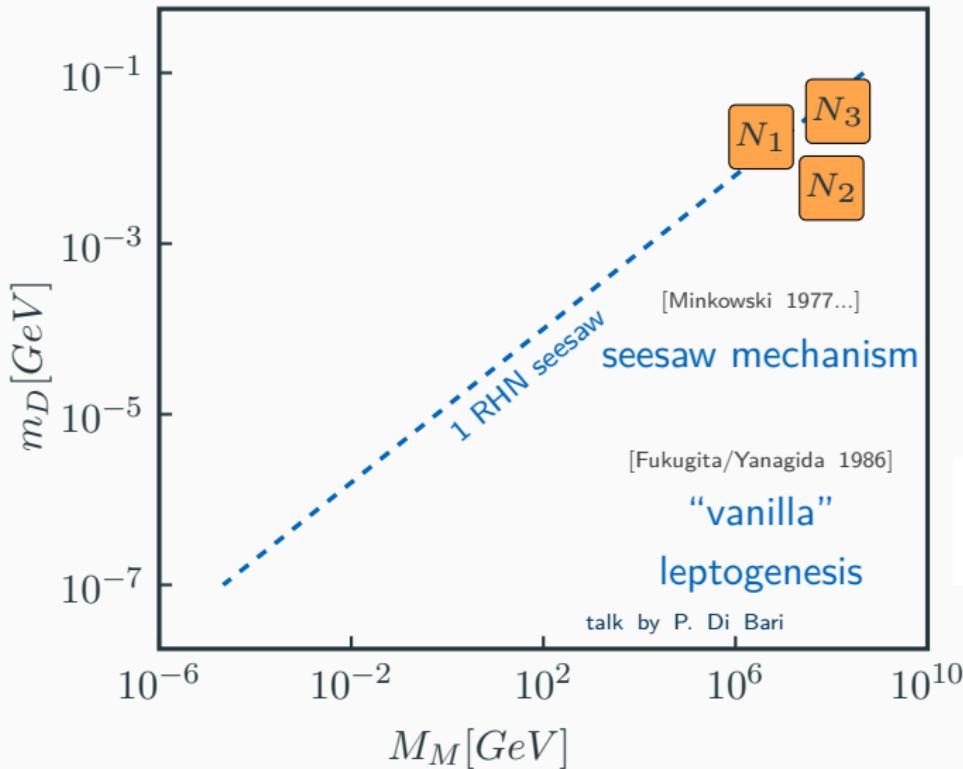
## Mixing with RHN

$$U_{ai}^2 \equiv \left| \left( m_D M_M^{-1} \right)_{ai} \right|^2$$
$$U^2 = \sum_{a,i} U_{ai}^2$$

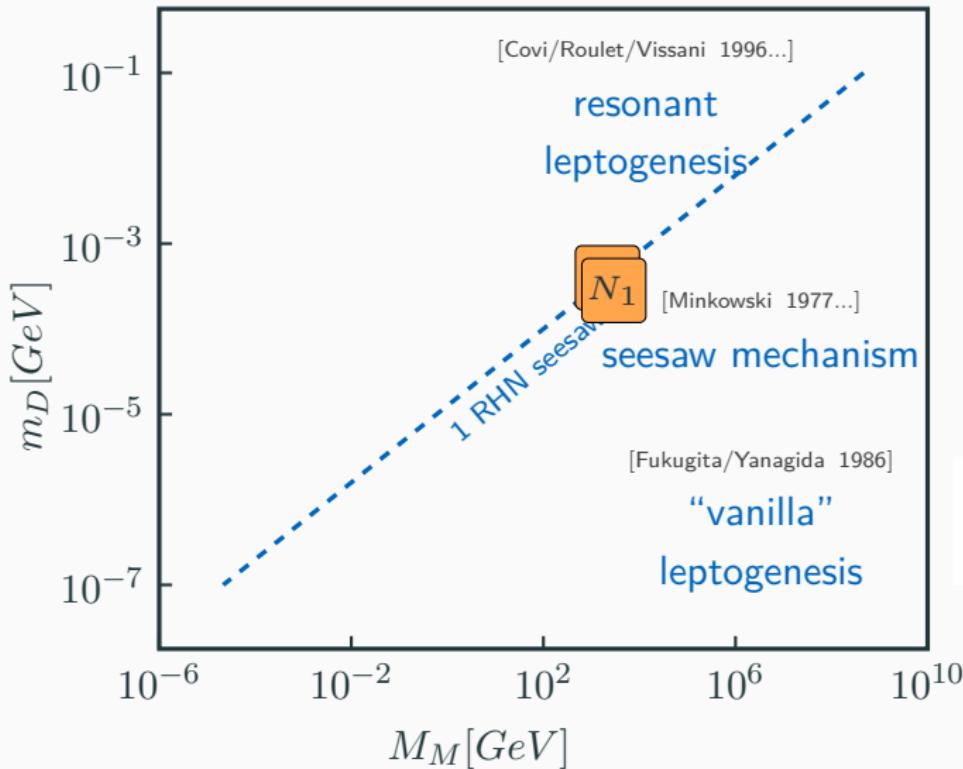
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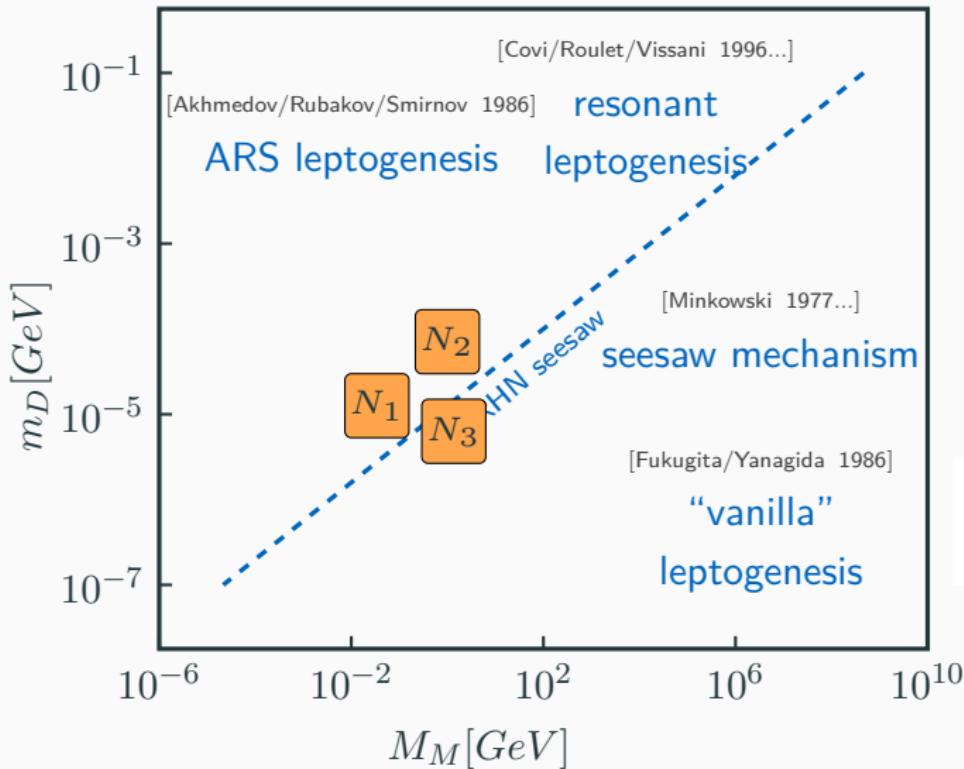
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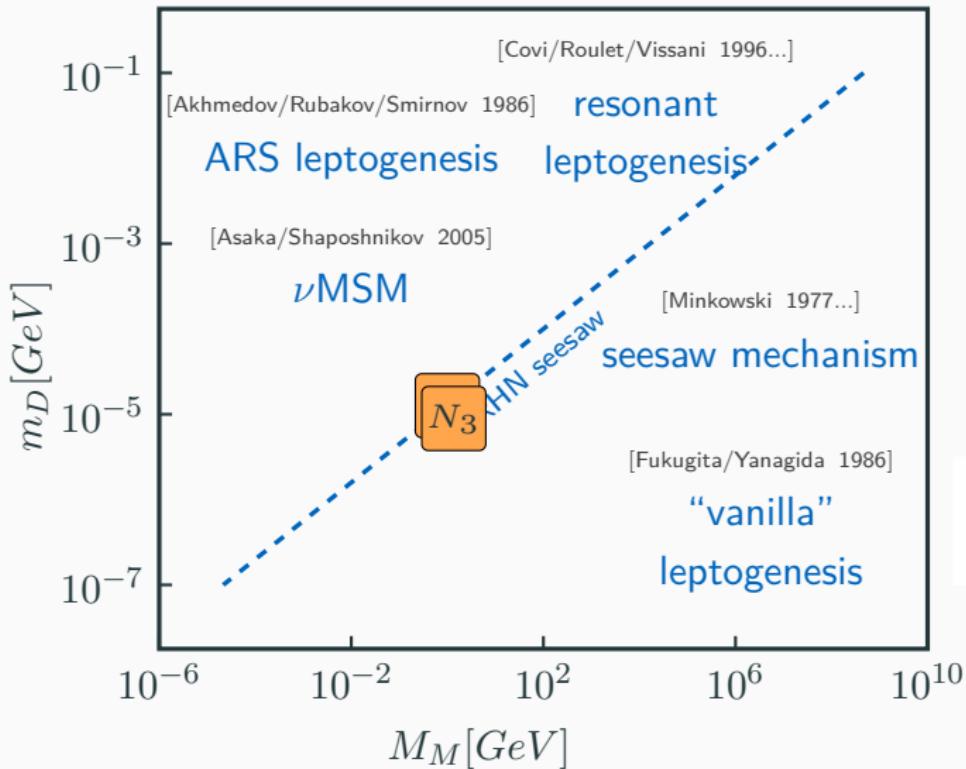
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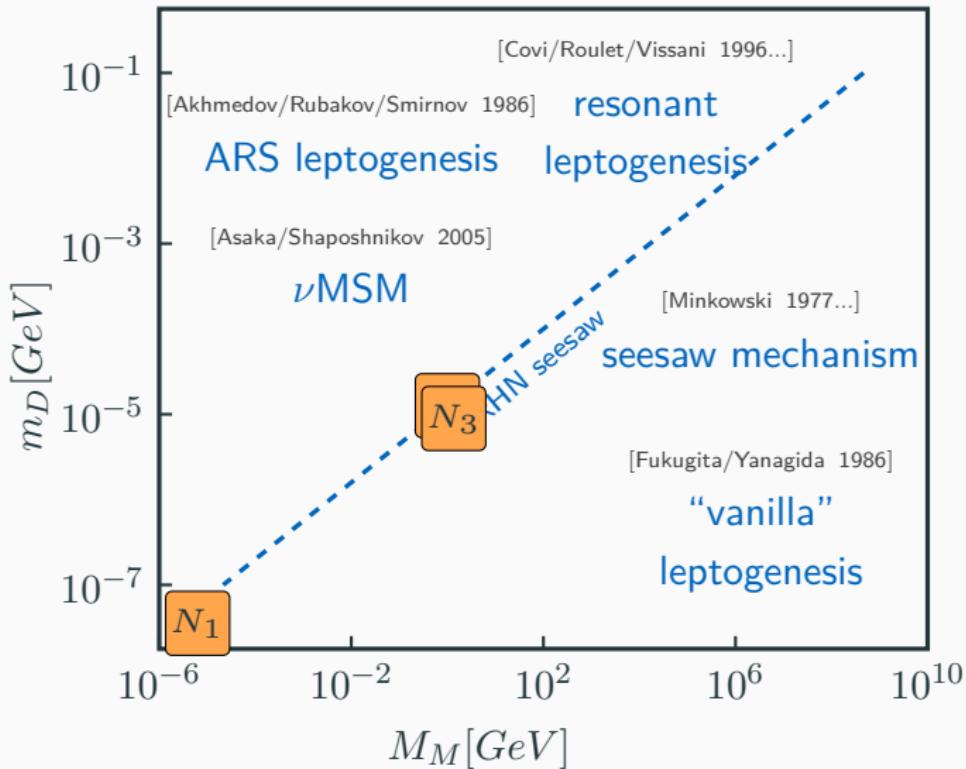
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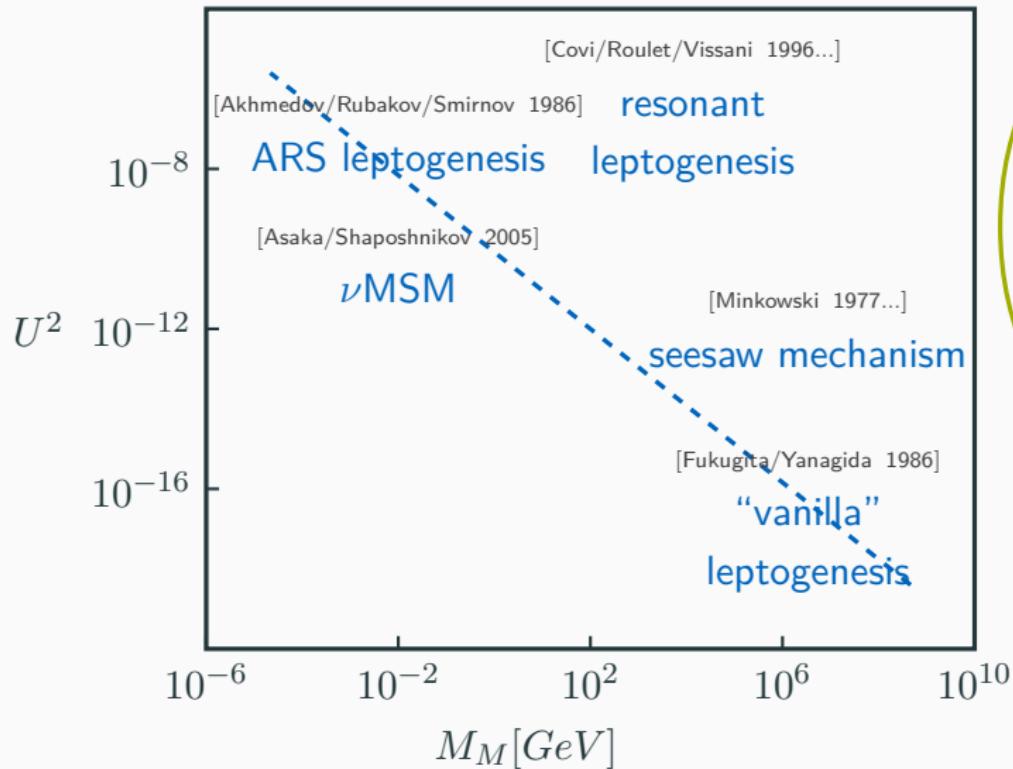
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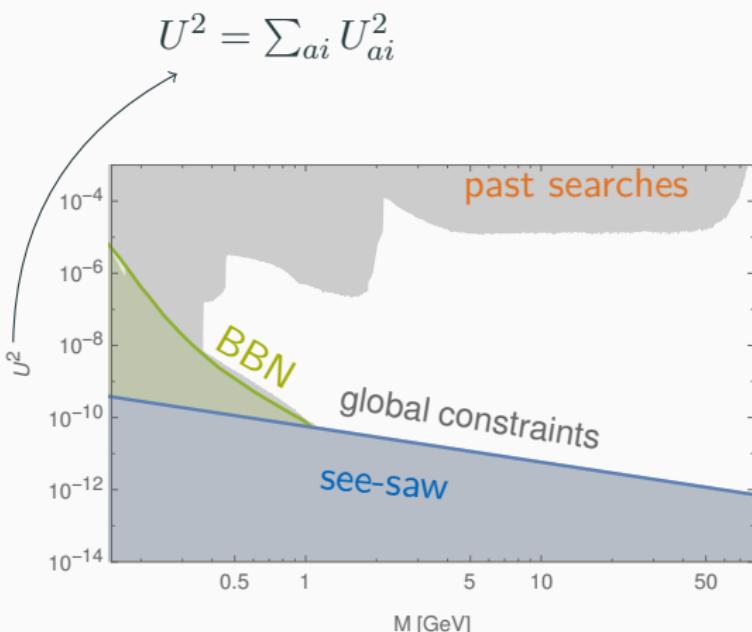
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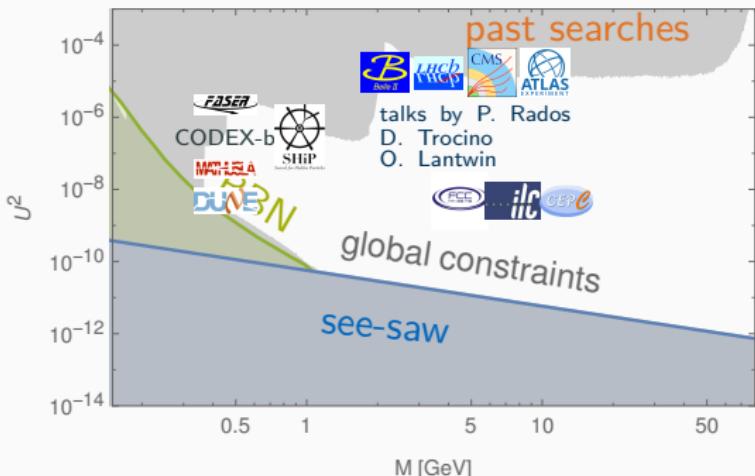
## Constraints on RHN properties: 2 RHN, Normal Ordering



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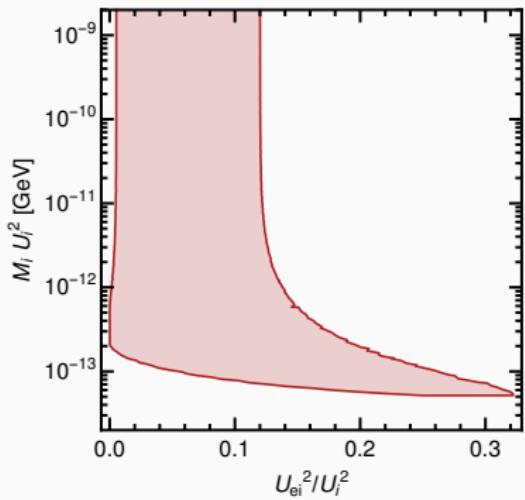
Indirect constraints:

- neutrinoless double  $\beta$  decay
- lepton universality
- CKM universality
- electroweak precision data
- $LFV$  in rare lepton decays:
  - $\mu \rightarrow e\gamma$
  - $\tau \rightarrow e\gamma$
  - $\tau \rightarrow \mu\gamma$



## Testing the $\nu$ mass generation: Normal Ordering

Low energy parameters fixed to best fit,  $\delta_{CP} \in [0, 2\pi]$  and Majorana phase  $\alpha \in [0, 4\pi]$

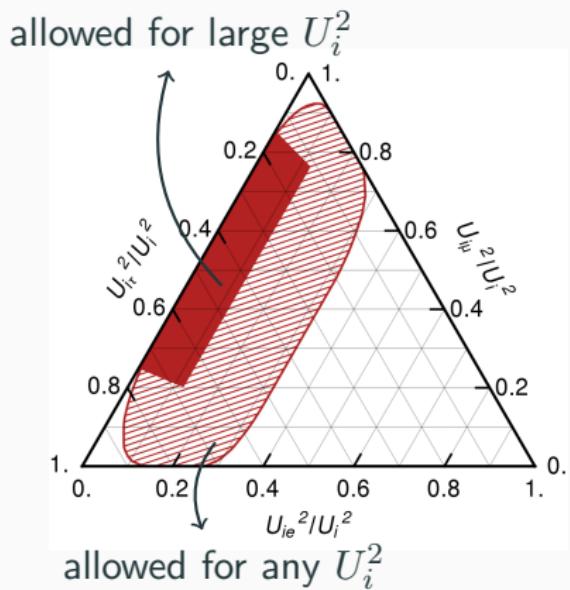
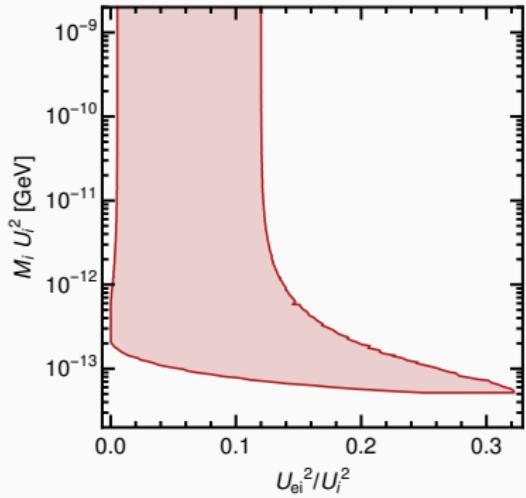


- the mixing angles  $U_{ai}^2$  are related to the measured light neutrino parameters
- this constrains ratios  $U_{ai}^2/U_i^2$
- constraints approach a limit for large  $U_i^2$
- if we can measure individual  $U_{ai}^2$  we have a consistency check for the seesaw mechanism

[Drewes/Hajer/JK/Lanfranchi 1801.04207]

[Drewes/Garbrecht/Gueter/JK 1609.09069] [Caputo/Hernandez/Lopez-Pavon/Salvado 1704.08721]

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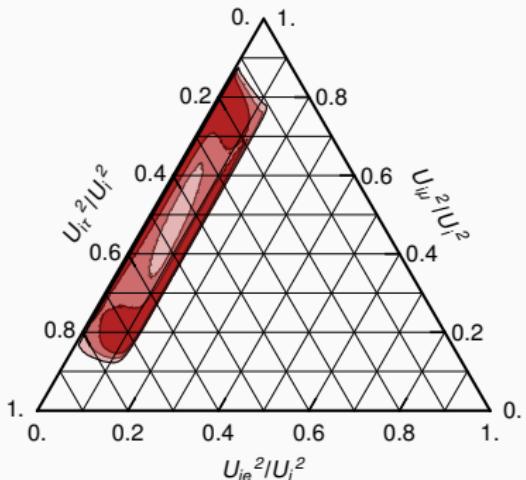
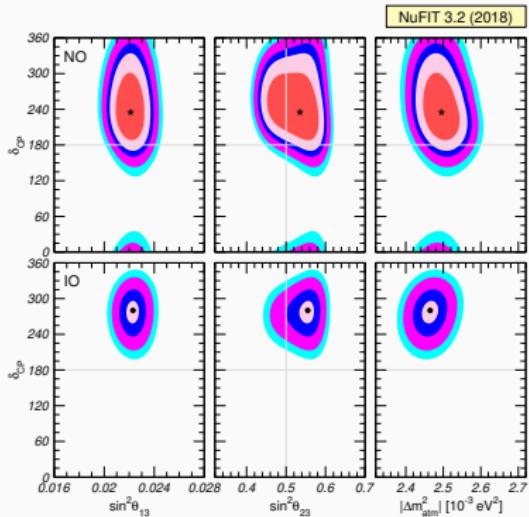


[Drewes/Hajer/JK/Lanfranchi 1801.04207]

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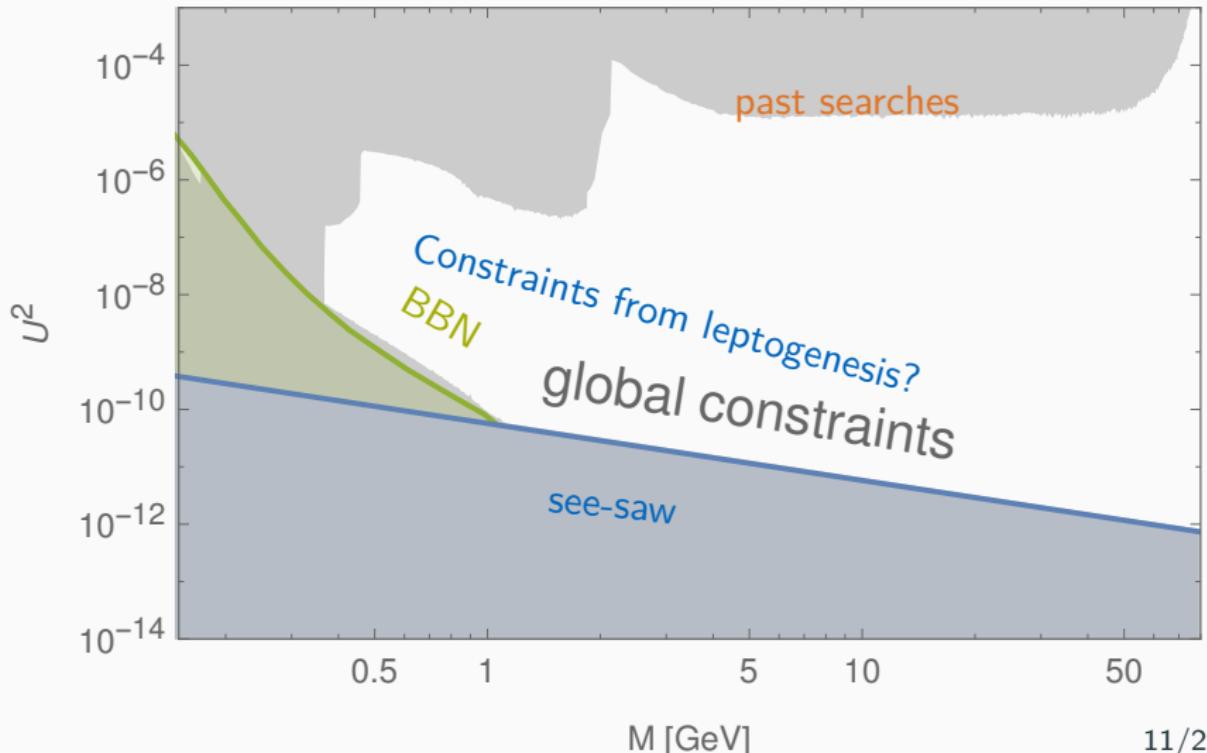
Measurement of  $\delta_{CP}$  improves the prediction:



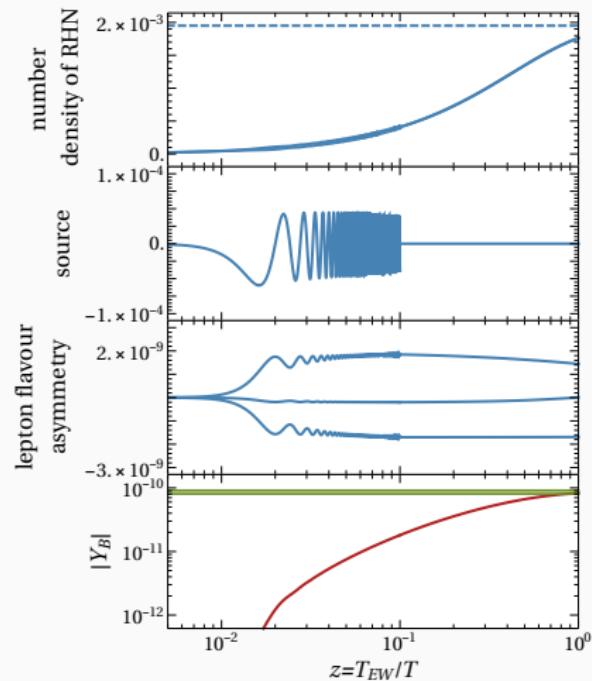
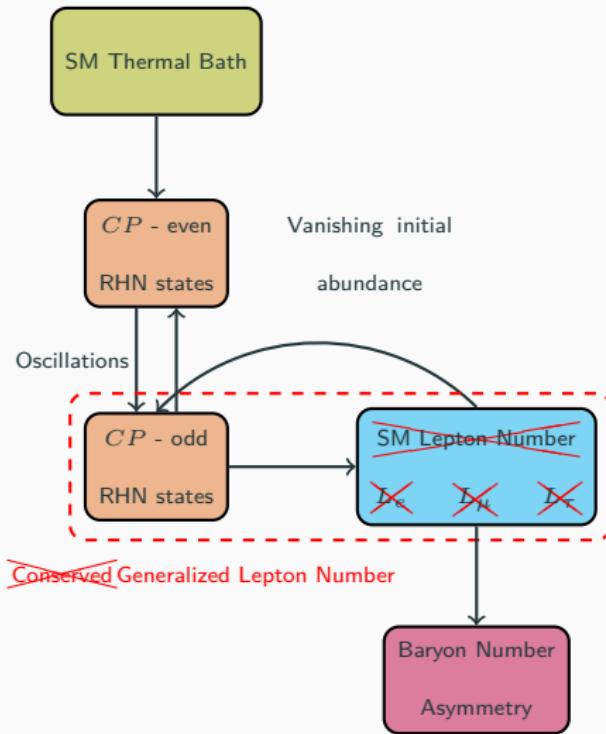
[Esteban/Gonzalez-Garcia/Maltoni/Martinez-Soler/Schwetz 1611.01514]

[Drewes/Hajer/JK/Lanfranchi 1801.04207]

## Constraints on RHN properties: 2 RHN, Normal Ordering



## Leptogenesis via Neutrino Oscillations

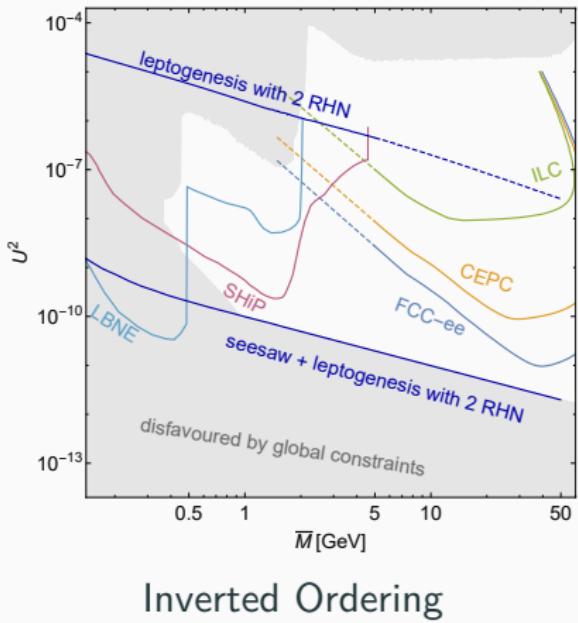


## Theoretical challenges and progress:

- derivation of the **density matrix** equations from **first principles**
- **rate calculations** for RHN production at finite  $T$
- **spectator effects**
- freezeout of the baryon number
- **analytical approximations** for different regimes
- violation of generalized lepton number
- momentum dependence
- systematic studies of the parameter space/ phenomenological implications

[ A. Abada, S. Antusch, E. K. Akhmedov, G. Arcadi, T. Asaka, S. Blanchet, I. Boiarska, K. Bondarenko, A. Boyarsky, L. Canetti, A. Caputo, E. Cazzato, V. Domcke, M. Drewes, S. Eijima, O. Fischer, T. Frossard, B. Garbrecht, J. Ghiglieri, D. Gueter, T. Hambye, P. Hernandez, H. Ishida, M. Kekic, J. K., M. Laine, J. Lopez-Pavon, M. Luente, M. Ovchinnikov, J. Racker, N. Rius, V. A. Rubakov, O. Ruchayskiy, J. Salvado, M. Shaposhnikov, B. Shuve, A. Y. Smirnov, D. Teresi, I. Timiryasov, A. Yavin... ]

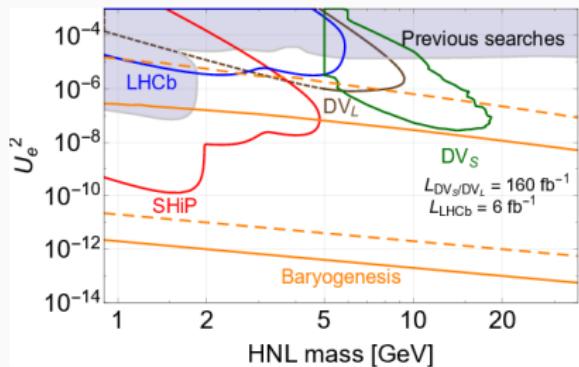
# Leptogenesis with 2 RHN



[Drewes/Garbrecht/Gueter/JK 1609.09069]

- large mixing angle  $U^2 \rightarrow$  large couplings  $\rightarrow$  BAU can be erased before EWPT
- flavour asymmetric washout can help “hide” the BAU in one of the flavours
- **seesaw** allows for  $U_\mu^2/U^2$  up to a factor  $\mathcal{O}(10^{-3})$  for IO

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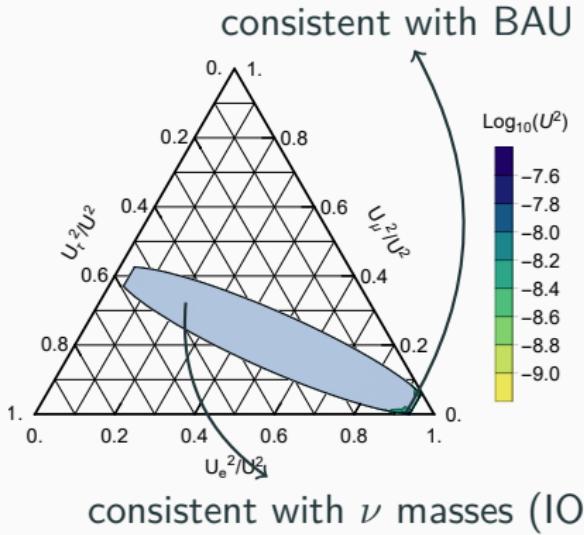


[Eijima/Shaposhnikov/Timiryasov 1808.10833]  
[Boiarska/Bondarenko/Boyarsky/Eijima/

Ovchynnikov/Ruchayskiy/Timiryasov 1902.04535]

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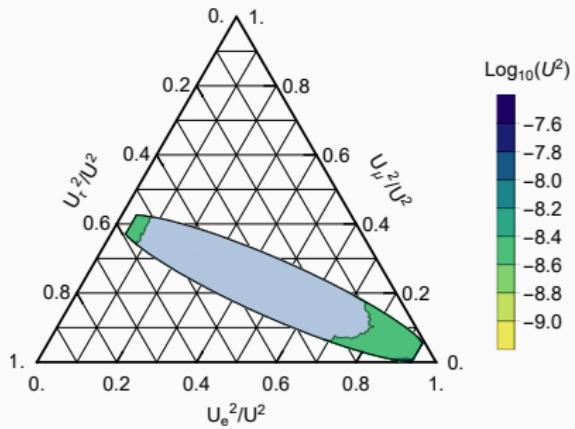


$$\bar{M} = 30 \text{ GeV}$$

$$U^2 = 4 \times 10^{-9}$$

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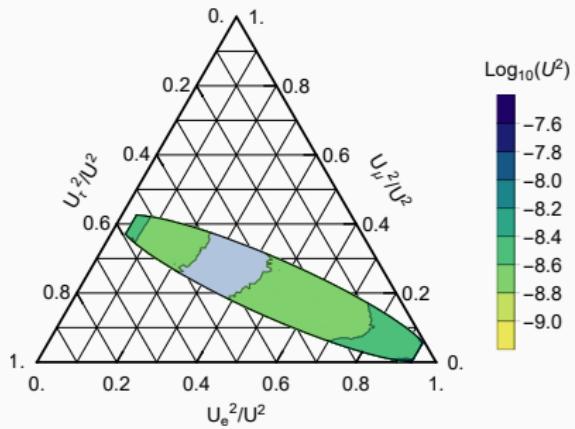
$$U^2 = 2.5 \times 10^{-9}$$

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[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/JK

1710.03744]

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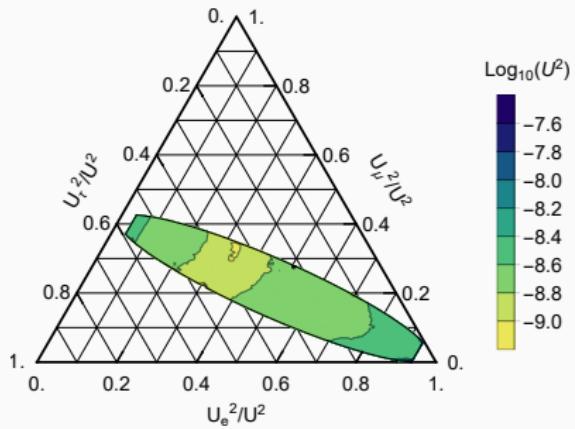
$$U^2 = 1.6 \times 10^{-9}$$

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[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/JK

1710.03744]

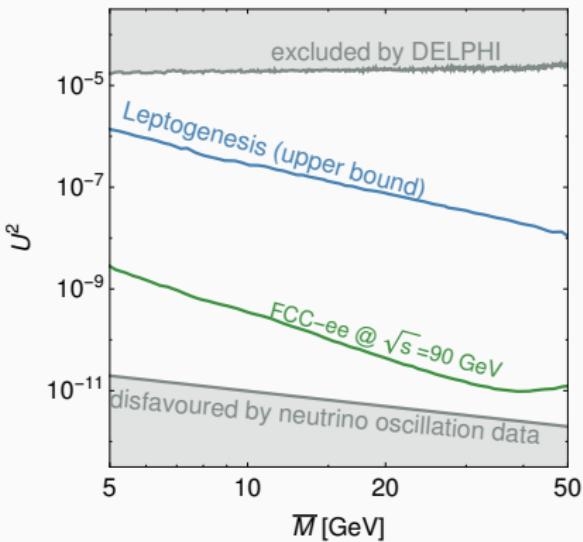
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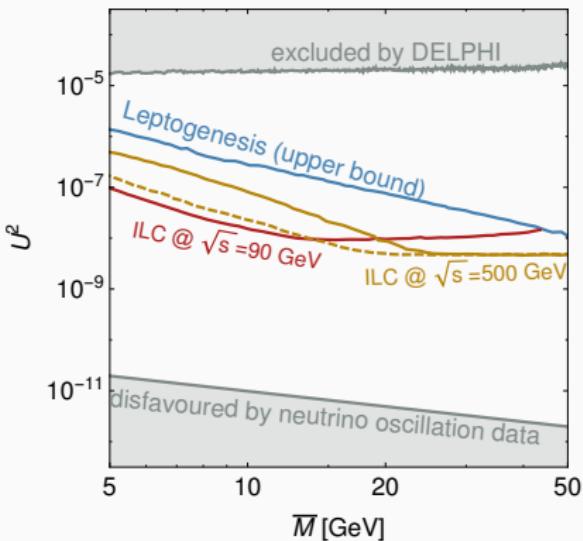
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# Sensitivity at Future lepton colliders: IO



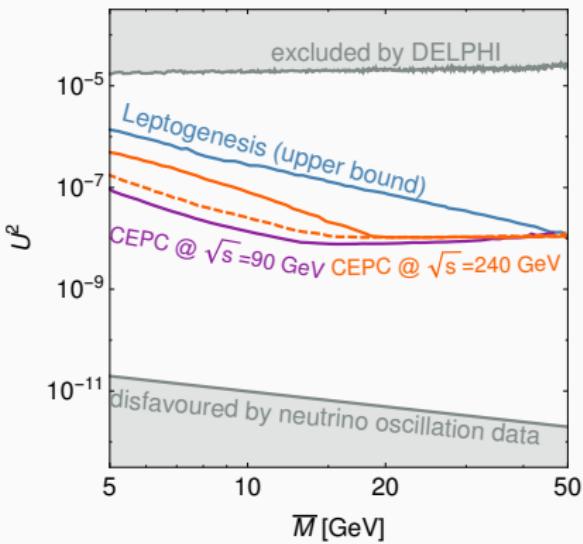
- displaced vertex searches
- sensitivity calculated for each point consistent with leptogenesis
- guaranteed discovery - all points in a given  $M-U^2$  bin have  $> 4$  events
- potential discovery - some points have  $> 4$  events
- for  $Z$ -pole run expected reach independent of flavour ratios
- away from the  $Z$ -pole, production through  $U_e^2$

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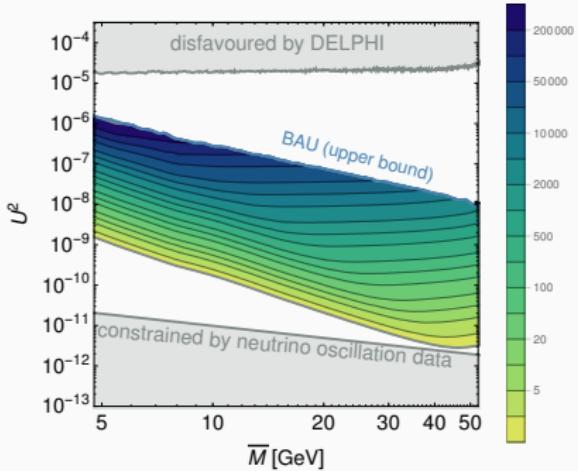
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# Measuring Flavour Ratios

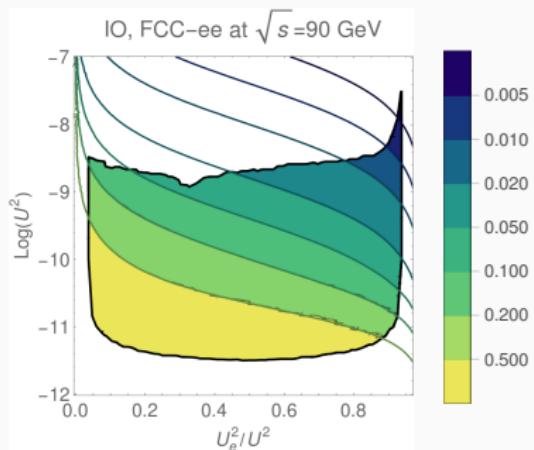


- large mixing angles - up to  $O(10^5)$  RHN can be produced at the FCC
- use this to measure the flavour ratios
- large  $M$  - fraction of semileptonic events with flavour  $a$  in the final state proportional to  $U_a^2/U^2$

[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/JK

1710.03744]

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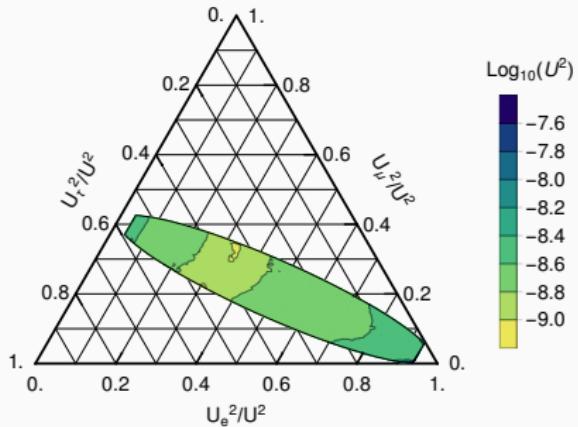
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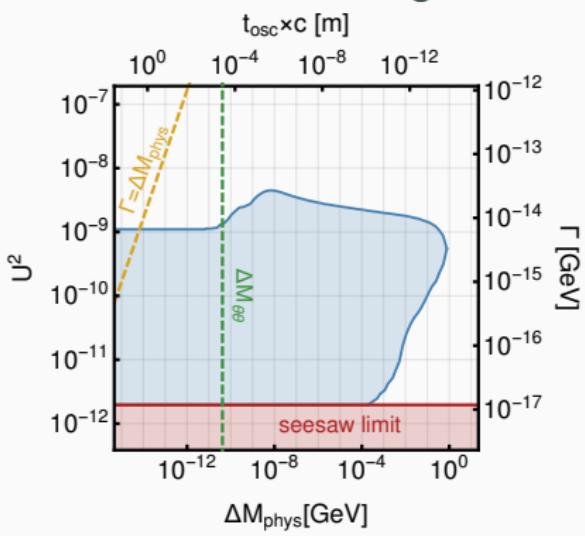
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# Measuring the mass splitting

Normal Ordering:



[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/JK

1710.03744]

- large range of viable  $\Delta M$  consistent with leptogenesis
- energy resolution of planned experiments -  
 $\Delta M/M \sim \mathcal{O}(\text{few}\%)$
- Higgs vev contribution to RHN mass difference  $\Delta M_{\theta\theta}$  practically implies lower limit on the mass splitting
- more information possible from LNV observables?

# LNV and the RHN mass splitting

$$\Gamma_{X \rightarrow \ell_a \ell_b} \sim \left| \sum_i \begin{array}{c} \nearrow \theta_{ai} \\ \text{---} \\ \text{---} \\ \text{---} \\ \text{---} \\ \nearrow \theta_{bi} \\ N_i \end{array} \right|^2$$

Same-to-opposite sign  
dilepton ratio

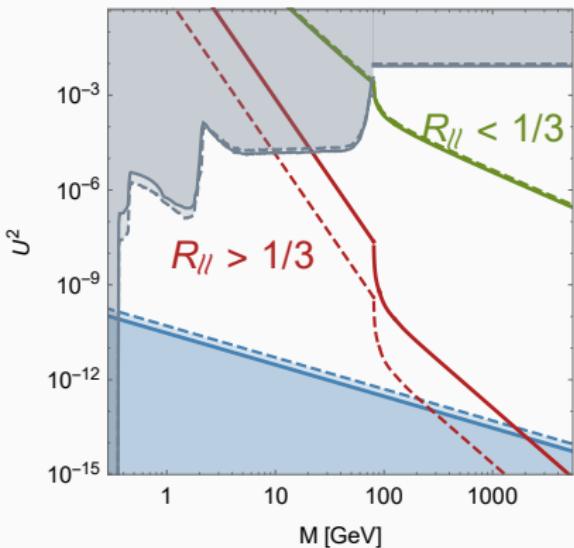
$$R_{\ell\ell} = \frac{\Delta M_{\text{phys}}^2}{\Delta M_{\text{phys}}^2 + \Gamma_N^2}$$

[Anamiati/Hirsch/Nardi 1607.05641]

see talk by R. Ruiz

- ratio between RHN lifetime and mass splitting determines the number of LNV decays
- non-vanishing neutrino masses imply a lower limit on  $\Delta M_{\text{phys}}$
- radiative stability of  $\nu$  masses implies an upper limit on  $\Delta M_{\text{phys}}$
- $R_{\ell\ell}$  can lead to information on the mass splitting

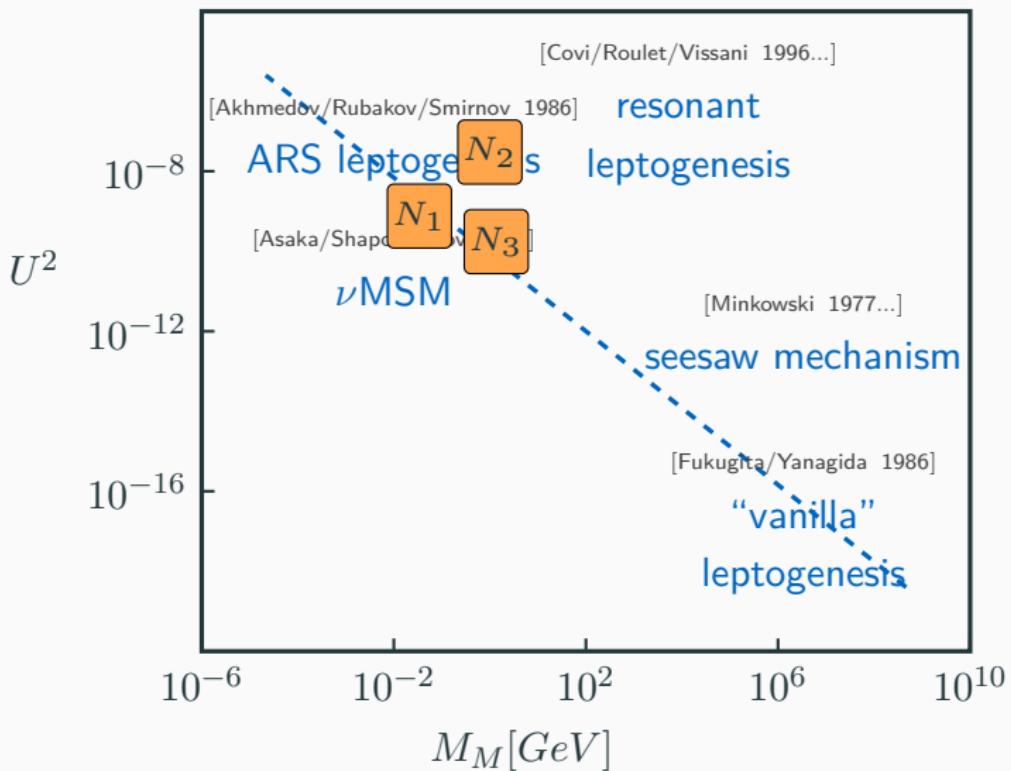
# LNV and the RHN mass splitting



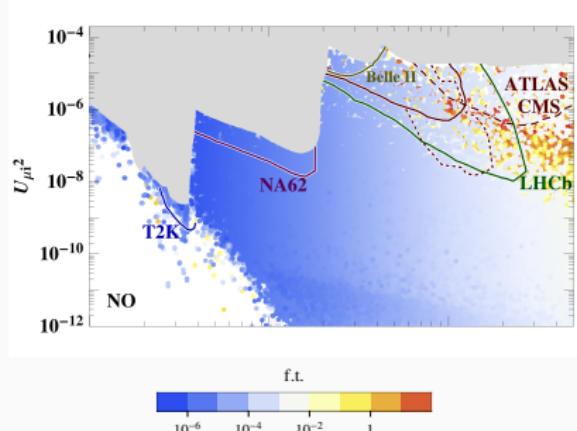
[Drewes/Klose/JK 1907.13034]

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# Leptogenesis with 3 RHN



# Leptogenesis with 3 RHN Normal Ordering



[Abada/Arcadi/Domcke/Drewes/JK/Lucente

1810.12463]

- no constraints on mixing angles from the seesaw mechanism, possible to avoid washout
- new sources of  $CP$  violation with 3 RHN
- possible level crossing between the RHN energies  
→ resonant enhancement!

## Conclusions

- GeV-scale RHN are a viable solution to the questions of the BAU and neutrino masses
- the minimal model with 2 RHN imposes strong constraints on the properties of the RHN
- 3 RHN - larger parameter space, weaker predictions
- measurements of the branching ratios  $N \rightarrow Xa$  are complementary with the low-energy neutrino oscillation experiments
- leptogenesis is within reach of existing experiments and future lepton colliders
- leptogenesis with large mixing angles implies strong constraints on the RHN branching ratios
- full testability is in principle possible, difficult in practice