Neutral Naturalness: From Colliders to Cosmology and Astrophysics

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Based on 1803.03263, 1905.xxxx Chacko, DC, Geller, Tsai 1906.xxxx DC, Jack Setford ongoing work with Shayne Gryba

Cảm ơn ban to the organizers and the ICISE center





International Centre for Interdisciplinary Science and Education



Exploring Hidden Sectors

There has to be new physics...

The usual fundamental mysteries (Hierarchy Problem, DM, Baryogenesis, Neutrinos, ...) aren't going anywhere.

Higgs discoveries and DM astro measurements sharpen these questions!

Canonical solutions (SUSY, WIMP DM, ...) generally involve IRminimal models, where the new degree of freedom which solves the mystery has sizable direct coupling to the SM.

This leads to irreducible signatures that haven't shown up so far.

... where is it?

Hidden Sectors

Particles & forces hidden from us due to small coupling, not high mass.

Generically arise due to the grammar of QFT.

Confirmed examples: v's, DM

Give non-minimal IR spectra from minimal theory input (e.g. QCD cousins like Hidden Valleys)



Can couple to SM via small portal couplings, e.g. Heavy Mediators Higgs Portal Photon Portal

I. Exotic Higgs Decays as probes

LHC can probe tiny exotic branching ratios if decays spectacular. Sizable Higgs Portal couplings to new physics are generic.

2. Long Lived Particles (LLPs) are generic

Once produced, Hidden Sector states can only decay back to SM via small portal couplings, generically leading to long lifetimes. The LLP lifetime is (almost...) a free parameter!

3. Complementarity between Cosmology and Colliders



Models which avoid signatures in one will often show up in the other

(e.g. dark radiation, DM with structure, etc.)



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Big Picture I

Hidden sectors can give rise to "arbitrarily" rich cosmology and astrophysics.

Can we make this predictive?

Yes: make the hidden sector solve some of these fundamental mysteries.

 \rightarrow "signature generator" of complex hidden sector phenomena

Big Picture II



Neutral Naturalness

Neutral Naturalness

Solves the (little) Hierarchy Problem without colored top partners to explain LHC null results.

Example of a particularly motivated hidden sector.

Solution to the hierarchy problem that is discoverable via non-standard searches and demonstrates collider-cosmo complementarity: either get LLP signals

or

very rich cosmology and astrophysics

Minimal Twin Higgs (MTH)

 $SM_A \propto SM_B$ (mirror sector) particle content with Z_2 symmetry

Higgs sector: SU(4), broken by Gauge + Yukawa interactions to $SU(2)_A \propto SU(2)_B \propto Z_2$, which generate mass for goldstone boson.

$$\Delta V = \frac{3}{8\pi^2} \Lambda^2 \left(\lambda_A^2 H_A^{\dagger} H_A + \lambda_B^2 H_B^{\dagger} H_B \right) \qquad \qquad \lambda_A = \lambda_B \equiv \lambda + \Delta V = \frac{3\lambda^2}{8\pi^2} \Lambda^2 \left(H_A^{\dagger} H_A + H_B^{\dagger} H_B \right) = \frac{3\lambda^2}{8\pi^2} \Lambda^2 H^{\dagger} H_B$$

Z₂ symmetry of quadratically divergent contributions mimics full SU(4) symmetry, protects pNGB Higgs mass @ 1-loop.

This is an IR model up to few TeV. Have to UV complete. O(dozen) examples in literature

hep-ph/0506256 Chacko, Goh, Harnik 1411.3310 Burdman, Chacko, Harnik, de Lima, Verhaaren of SM [a complicated hidden valley!] Soft Z_2 breaking to make hidden higgs vev higher than SM to avoid Higgs bounds: $v_B/v_A > \sim 3$

SM \mathbf{Z}_{2} exchange particles and gauge groups

mirror sector

This requires tuning ~ $(v_B/v_A)^2$ ~ Br(h→mirror)

 Z_2 symmetry \rightarrow hidden sector copy

Strassler, Zurek 2006

Uncolored top partners.

Massless degrees of freedom: (twin photon, neutrinos) $\Rightarrow \Delta N_{eff} \sim 5$

Minimal model incompatible with cosmology.

mirror SM sector \mathbf{Z}_{2} exchange particles and gauge groups

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Soft Z₂ breaking to make hidden higgs vev higher than SM to avoid Higgs bounds: $v_B/v_A > \sim 3$

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Minimal model incompatible with cosmology.

Fix I: Hard Z₂ breakings e.g. Fraternal Twin Higgs

Craig, Katz, Strassler, Sundrum 1501.05310

→ mirror QCD gives rise to **light LLPs** produced via Higgs portal

SM	$\overline{Z_2}$	mirror sector
	exchange	
	particles	
	and	
	gauge	
	groups	

 Z_2 symmetry \rightarrow hidden sector copy of SM [a complicated hidden valley!]

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LLPs @ LHC!



DC, Verhaaren 1506.06141

MATHUSLA Physics Case White Paper 1806.07396

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Fix 2: dilute mirror sector cosmological abundance: Asymmetric Reheating!



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Asymmetrically Reheated Mirror Twin Higgs

Example: vMTH

Let's also solve the Neutrino Mass problem: add RH neutrinos to MTH and implement type-I See-saw

Toy model with I RH neutrino without Z₂ breaking (can extend to 3 & various realistic flavor models):

$$\mathcal{L} \supset -y\left(L_AH_AN_A+L_BH_BN_B
ight) -rac{1}{2}M_N\left(N_A^2+N_B^2
ight) -M_{AB}N_AN_B+ ext{h.c.}$$

RH-neutrino mass eigenstates live in both sectors:

$$egin{aligned} N_+ &= rac{1}{\sqrt{2}} \left(N_A + N_B
ight) \ N_- &= rac{1}{\sqrt{2}} \left(N_A - N_B
ight) \end{aligned}$$

Example: vMTH

Only source of Z_2 breaking is larger mirror Higgs vev, but this causes lightest RH neutrino to decay preferentially to SM (heavier mirror W boson):

$$\Gamma_{N \to i} \propto \frac{m_{\nu_i}^2}{m_{W_i}^4} \longrightarrow \qquad \epsilon = \frac{\Gamma_{N \to B}}{\Gamma_N} \approx \frac{v^2}{f^2}$$

If the Neutrinos have mass at GeV scale, decay out of equilibrium AFTER the higgs portal freezes out (mirror & visible sector decoupled). → **Dilute mirror sector!**

$$M_N < 1 \text{ GeV} \left(\frac{0.01 \text{ eV}}{m_\nu}\right)^{1/2} \qquad \Delta N_{eff} \sim 5 \in = 5 \text{ (v/f)}^2$$

Phenomenology

In the vMTH, the dilution is dictated by $(v_A/v_B)^2$, which is the tuning of the model and also measurable at colliders via Br(h→invis).

Long-lived RH neutrino might also be detectable.

But let's focus on cosmology and astrophysics.

Choose a general parameterization of the Asymmetric Reheating mechanism within the MTH framework:

	$\Delta N_{eff},$	$v_B/v_A,$	$r_{ m all} = \Omega_{ m all\ mirror\ baryons} / \Omega_{ m DM}.$	
model like vMTH connects these two		l like I connects	any mirror-baryogenesis mechanism will give some	
		two	asymmetric mirror relic abundanc	

Three parameters determine a family of rich hidden sector dictated by the hierarchy problem.

 $\Delta N_{eff}, \quad v_B/v_A, \quad r_{all} = \Omega_{all \text{ mirror baryons}}/\Omega_{DM}.$

What does the cosmology and astrophysics look like?

We have to recalculate all of cosmological history...

*For now, no assumptions on what the majority of DM is made of... [work in progress with Shayne Gryba]

1803.03263 Chacko, DC, Geller, Tsai

Asymmetric MTH Cosmology

- BBN: predicts ~ 75% mirror Helium mass fraction in mirror sector (compare to 25% SM).
- Mirror-baryo-acoustic oscillations modify matter power spectrum, shows up in CMB & LSS:
 Current Ly-α constrains r_{all} < ~ 10%
 CMB Stage IV will probe r_{all} ~ 1%
- $\Delta Neff \sim 0.few$
 - same free-streaming vs scattering fraction as SM
- Mirror baryons part of our galaxy, but cool slower than SM baryons. Feedback is complicated.
 Distribution may be disk-like or halo-like.

Mirror Baryons Direct Detection

SuperCDMS (nuclear recoil) taking data, SENSEI (electron recoil) is approved.



If there is *any* ambient mirror ionization, fast mirror electrons provide excellent discovery channel! (Detection possible in other cases too) But complex mirror sectors can give rise to MUCH weirder astrophysical phenomena...

Mirror Stars

1907.xxxxx DC, Jack Setford

Hierarchy Problem → Mirror Stars?

Mirror DM (perfect SM copy) is an old idea.

Foot, Ignatiev, Volkas astro-ph/0011156 and more

Mirror star signatures never really studied.

Neutral Naturalness motivates *family* of mirror sectors that are *fundamentally* motivated and allow for mirror stars.

Similar to SM but different enough to change detailed stellar astrophysics.

Want to consider mirror stars as a *general* class of hidden sector signatures!

Factorize the Question



Factorize the Question



Figure this out first for general mirror star

1907.xxxxx DC, Jack Setford

Factorize the Question



Figure this out first for general mirror star

1907.xxxxx DC, Jack Setford

Neutral Naturalness is a great signature generator!

next paper

How to discover Mirror Stars?

Use SM stars as benchmark "mirror stars" to study signature in detail.

Mass / M_{sun}	He mass fraction	Radius / R_{sun}	$T_{core}~/~10^7~{ m K}$	Luminosity / L_{sun}	Lifetime / years
0.2	0.24	0.21	0.68	6.7×10^{-3}	$2.5 imes 10^{11}$
0.5	0.24	0.45	0.92	0.04	3×10^{10}
1	0.24	0.89	1.37	0.72	1×10^{10}
2	0.24	1.66	2.10	16.5	$9 imes 10^8$
5	0.24	2.71	2.73	578	$9 imes 10^7$
15	0.24	4.95	3.43	1.98×10^4	$5 imes 10^6$
50	0.24	9.82	4.05	3.72×10^5	$3 imes 10^5$
80	0.24	13.3	4.25	8.94×10^5	8×10^4
100	0.24	15.7	4.34	$1.31 imes 10^6$	$5 imes 10^4$
1	0.75	1.15	2.35	17.1	?
5	0.75	2.80	3.46	5700	?
15	0.75	5.15	4.13	1.05×10^5	?

Table 1: MESA benchmark stars used in the calculations.

Source of signal: mirror photon mixing!

 $\epsilon F_{\mu\nu}F'^{\mu\nu}$





I) Photons from surface $\sim \epsilon^2$



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Also straight from core? NO!





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SM Baryon Capture in Mirror Star



There is capture due to mirror-SM collisions...

$$\left(\frac{dN_i}{dt}\right)_{mirror} = n_i^{ISM} \sum_j \frac{4\pi N_j \epsilon^2 \alpha^2 Z_i^2 Z_j^2}{m_i m_j u^3}$$

SM ISM density

relative velocity

and SM-SM collisions

$$\left(\frac{dN_i}{dt}\right)_{geometric} = n_i^{ISM} \sqrt{\frac{3}{2}} \frac{\overline{v}_{esc}^2}{u} \pi R^2,$$

geometric limit due to unsuppressed interactions

Total amount captured \propto age of star!



I) Photons from surface $\sim \epsilon^2$

Also straight from core? NO!



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Also straight from core? NO!

Captured SM "Nugget"

Captured matter arranges itself in profile with $R \sim virial$ radius given by equilibrium temperature.

Equilibrium temperature given by SM ionization energy where **bremsstrahlung cooling** becomes efficient.

$$T_{eq} \sim 10^4 K$$

 $r_{SM} \sim 10^{-3} R_{star}$. Nugget is ~ planet size!

Nugget is **optically thin** for optical photons (bremsstrahlung cooling), but **optically thick** for ionization and X-ray photons.



I) Photons from surface $\sim \epsilon^2$

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I) Photons from surface $\sim \epsilon^2$

Also straight from core? NO!

2) Thermal emission from SM nugget: **Optical!**

SM "Nugget" Emission Spectrum



SM "Nugget" Emission Spectrum



Bremsstrahlung Cooling gives optical signal!

Given its low luminosity, MUCH too hot to be standard astrophysical object

SM "Nugget" Emission Spectrum



Given its low luminosity, MUCH too hot to be standard astrophysical object What's this... X-ray Black Body Signal!?



I) Photons from surface $\sim \epsilon^2$

Also straight from core? NO!

2) Thermal emission from SM nugget: **Optical!**





These mirror stars have a very bizarre double-signal:

optical bremsstrahlung photons from the captured SM matter

and

converted X-ray photons straight from the mirror star core

Mirror Stars in HR Diagram



Mirror Stars in HR Diagram



Some of these mirror stars could show up in GAIA data.

> Requires closer analysis.

Could then diagnose mirror star with X-ray observations



Upshot

Mirror stars have a *robust* and *highly distinctive* signal.

Hot! (~10,000K)
X-rays = Mirror Star Core Temperature!
Only visible if close → parallax → get absolute lumi → Faint!

Upon detailed examination, this will look nothing like anything from SM astrophysics!

Conclusions

Conclusions

Hidden sectors are motivated from bottom-up and top-down.

Collider (LLP) vs Cosmology (stable)

Top-down formulations give *rich, predictive hidden sectors* with rich cosmology and astrophysics that you can interrogate in detail.

Asymmetrically reheated MTH:

- solves the little hierarchy problem
- if no LLP signatures, can still measure $Br(h \rightarrow invis) \sim v_B^2/v_A^2$
- this parameter is correlated with many many cosmo + astro observables: ΔN_{eff} , LSS deviations, direct detection, mirror stars

MIRROR STARS can a rise in many hidden sectors. We found they produce a highly robust and distinctive signal!