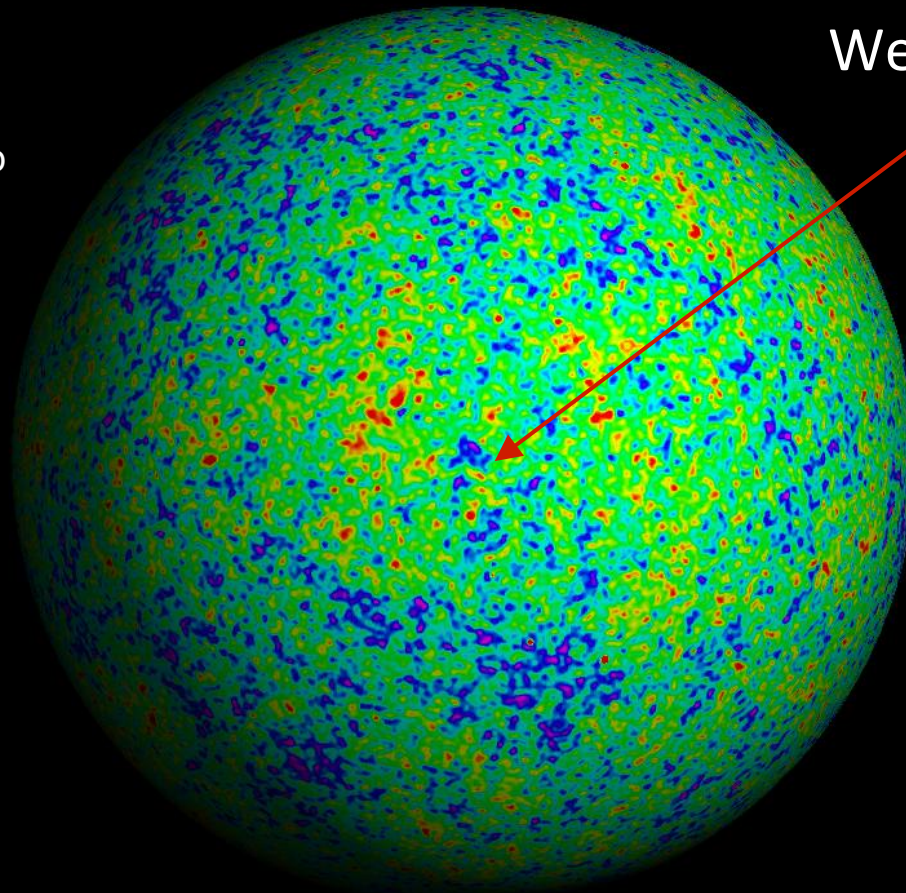




Pursuing the Shadows of Gravitational Waves from the Beginning of Time with BICEP/Keck and Future Experiments

Cosmic Microwave Background Surface of Last Scattering

All sky temperature map
projected on a sphere

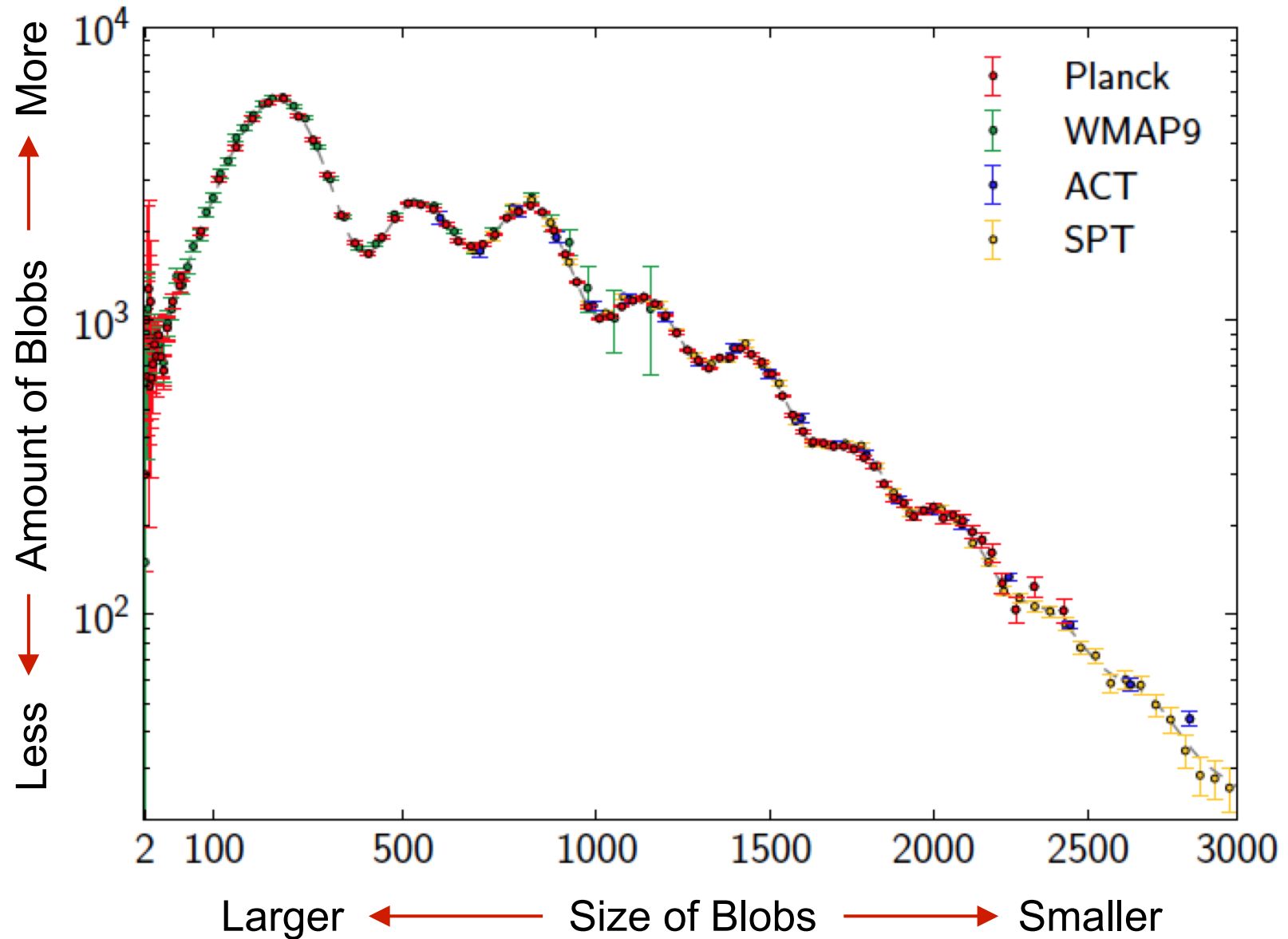


We are at the center

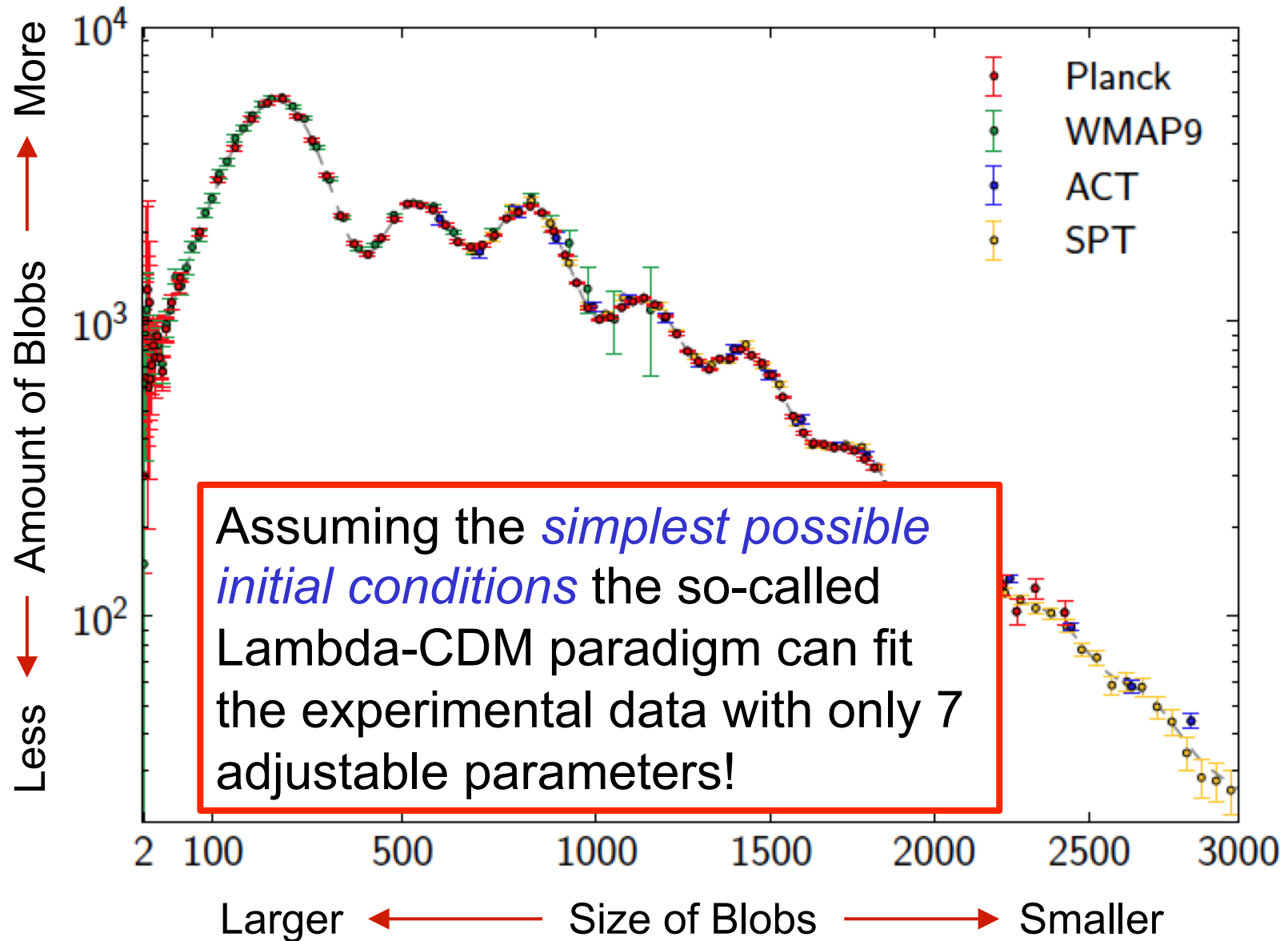
CMB temperature is a sample of the density structure on a shell cut
through the 380,000 year old Universe

Perturbations are one part in 10,000 at that time – and Gaussian!

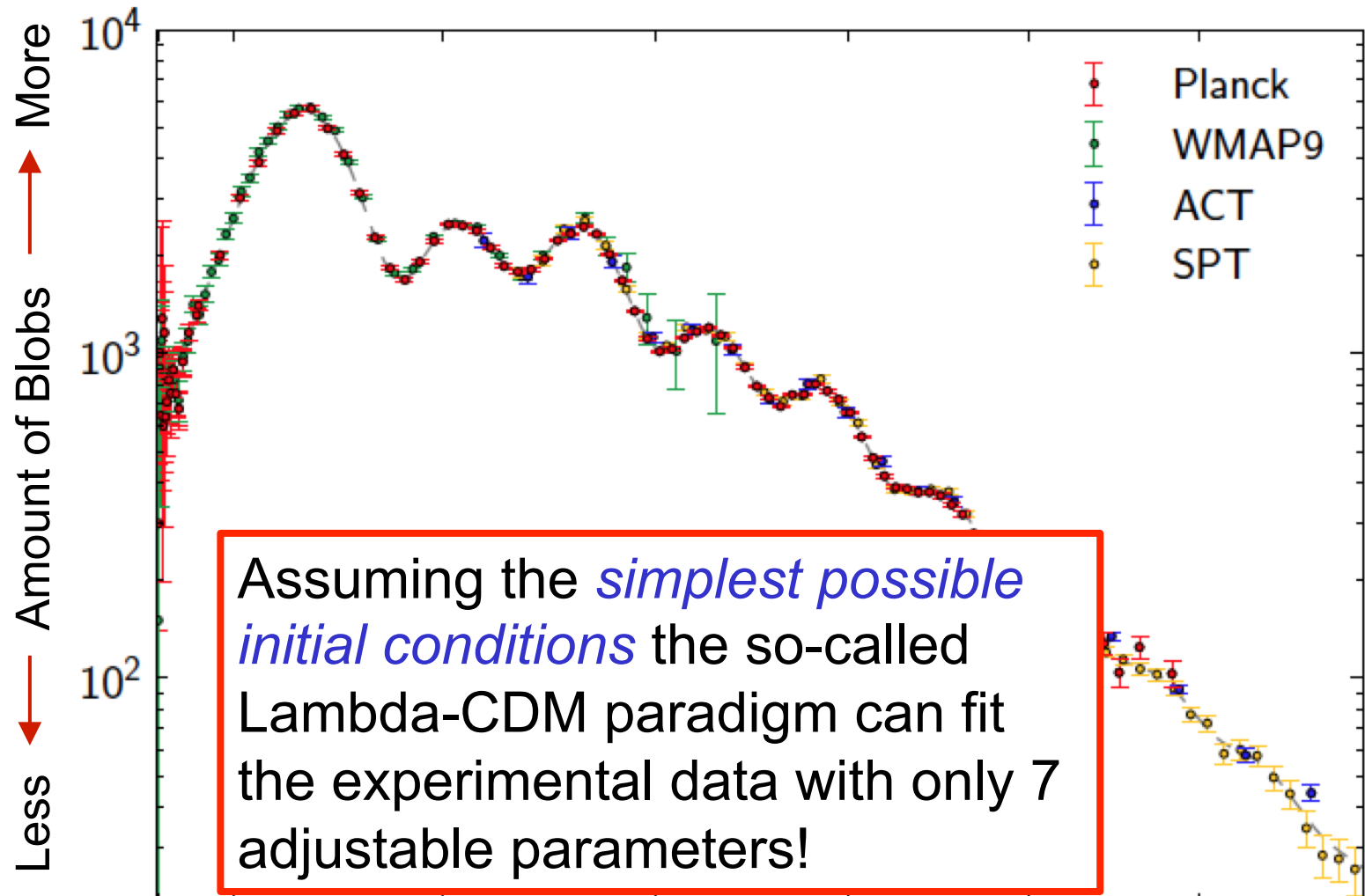
Power Spectrum (Blob size histogram)



Power Spectrum (Blob size histogram)



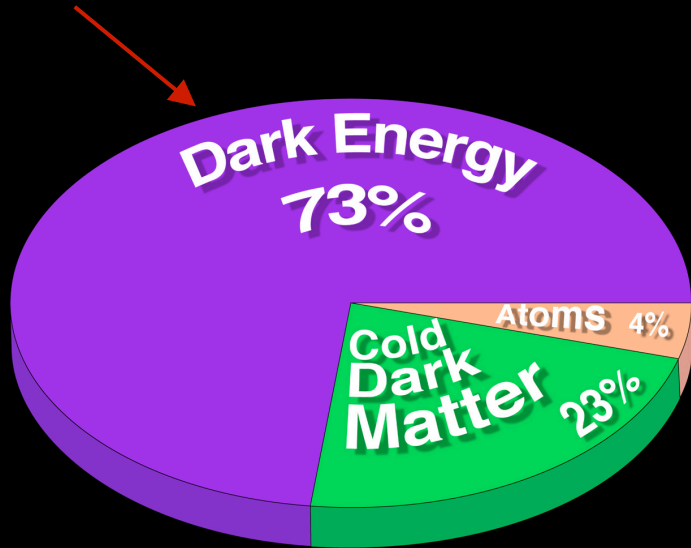
Power Spectrum (Blob size histogram)



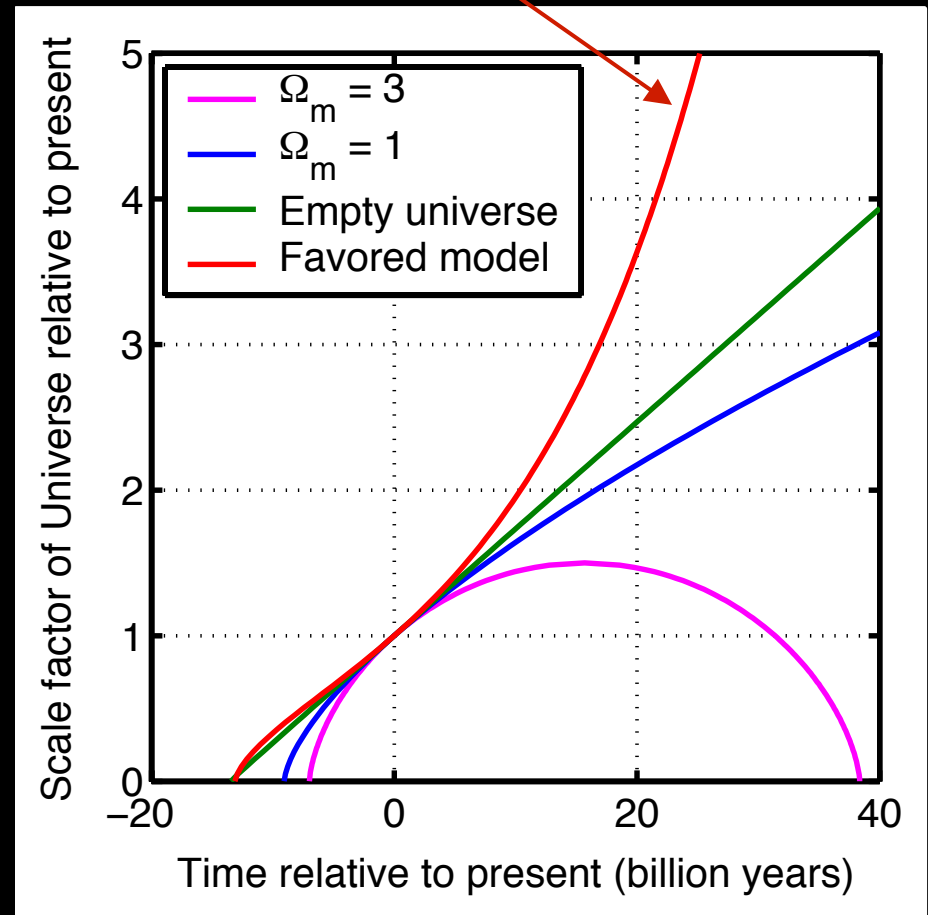
Compare that to the 26 parameters of the (extended) standard model of particle physics ;-)

Triumphant/Embarrassing LCDM

CMB and other data fits GR based LCDM model *beautifully*
– but it demands that 96% of the Universe is invisible to us



And it implies that the future is runaway expansion...



Also it doesn't explain horizon/flatness etc...

History of the Universe

Inflation posits a pre-phase of exponential expansion before LCDM

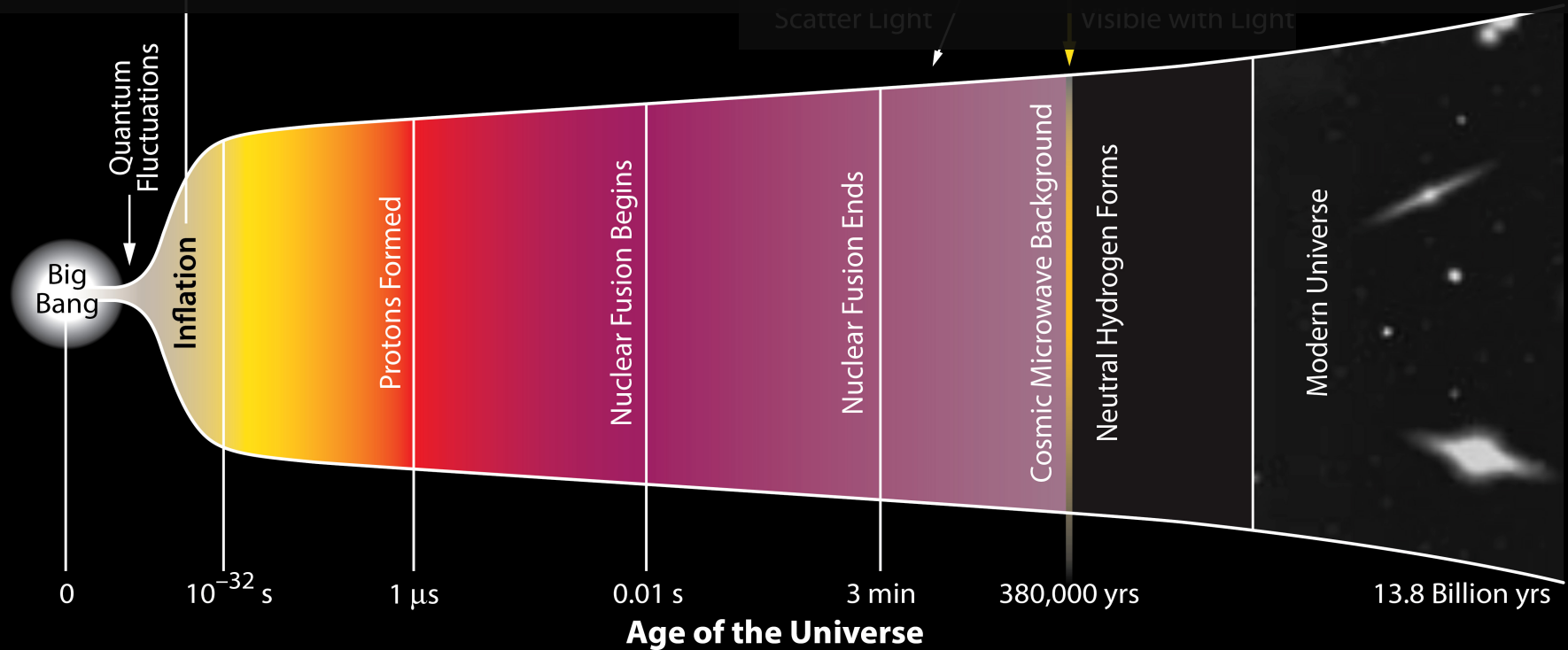


Alan Guth



Andrei Linde

Radius of the Visible Universe



What Does Inflation Do For Us?

Solves the horizon problem:
Why is the CMB nearly uniform?
How do apparently causally disconnected regions of space get set to the same temperature?



A volume much larger than our entire observable universe today was once a causally connected sub atomic speck.

Solves the flatness problem:
Why is the net spatial curvature close to zero?



Any initial spatial curvature is diluted away to undetectability by the hyper expansion.

Explains the initial perturbations:
Why Gaussian with close to flat power law spectrum?



Equal amounts of perturbations are injected by quantum fluctuations at each step in the exponential expansion.

Solves the monopole problem:
Why do we not observe magnetic monopoles in the Universe today?



Monopoles are diluted away to undetectability.

Inflation is controversial

Inflationary Paradigm after Planck 2013

Alan H. Guth,¹ David I. Kaiser,¹ and Yasunori Nomura²

¹*Center for Theoretical Physics, Laboratory for Nuclear Science, and Department of Physics, Massachusetts Institute of Technology, Cambridge, MA 02139, USA*

²*Berkeley Center for Theoretical Physics, Department of Physics, and Theoretical Physics Group, Lawrence Berkeley National Laboratory, University of California, Berkeley, CA 94720, USA*

(Dated: December 29, 2013, revised January 13, 2014)

[arxiv/1312.7619](https://arxiv.org/abs/1312.7619)



Inflationary schism after Planck2013

Anna Ijjas,^{1,2} Paul J. Steinhardt,³ and Abraham Loeb⁴

¹*Max-Planck-Institute for Gravitational Physics (Albert-Einstein-Institute), 14476 Potsdam, Germany*

²*Rutgers University, New Brunswick, NJ 08901, USA*

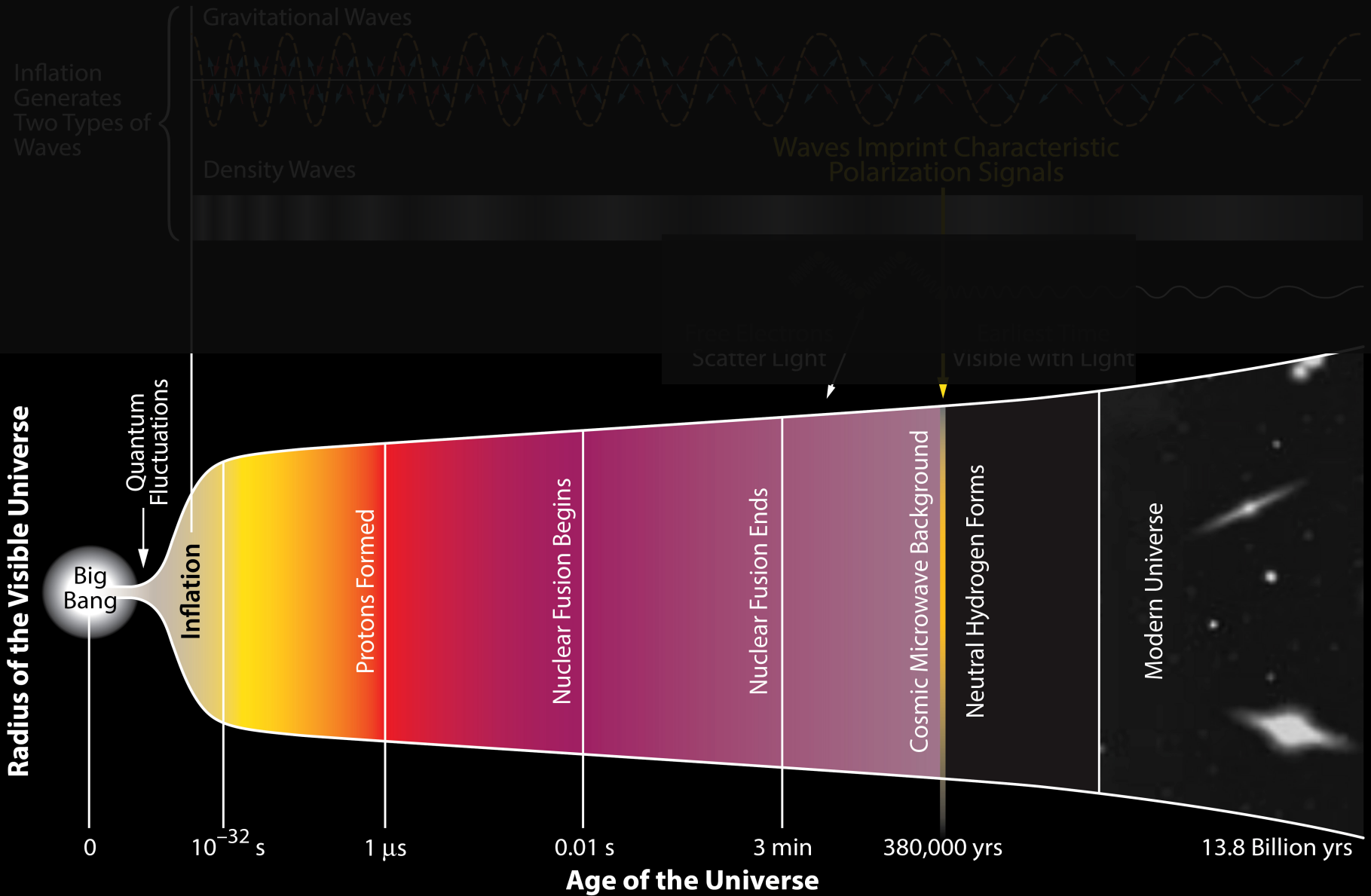
³*Department of Physics and Princeton Center for Theoretical Science, Princeton University, Princeton, NJ 08544, USA*

⁴*Harvard-Smithsonian Center for Astrophysics, Cambridge, MA 02138, USA*
(Dated: March 14, 2014)

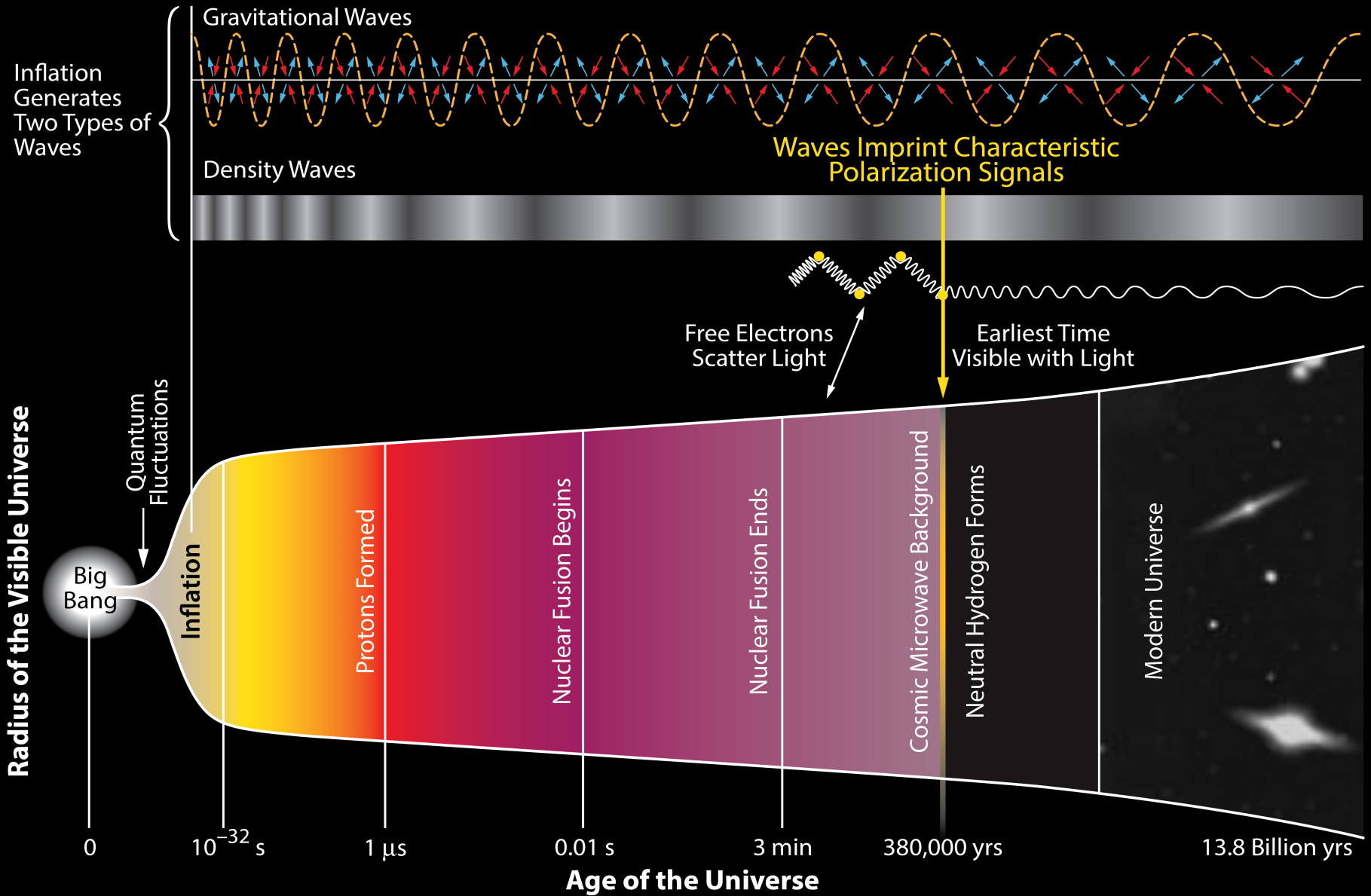
[arxiv/1402.6980](https://arxiv.org/abs/1402.6980)



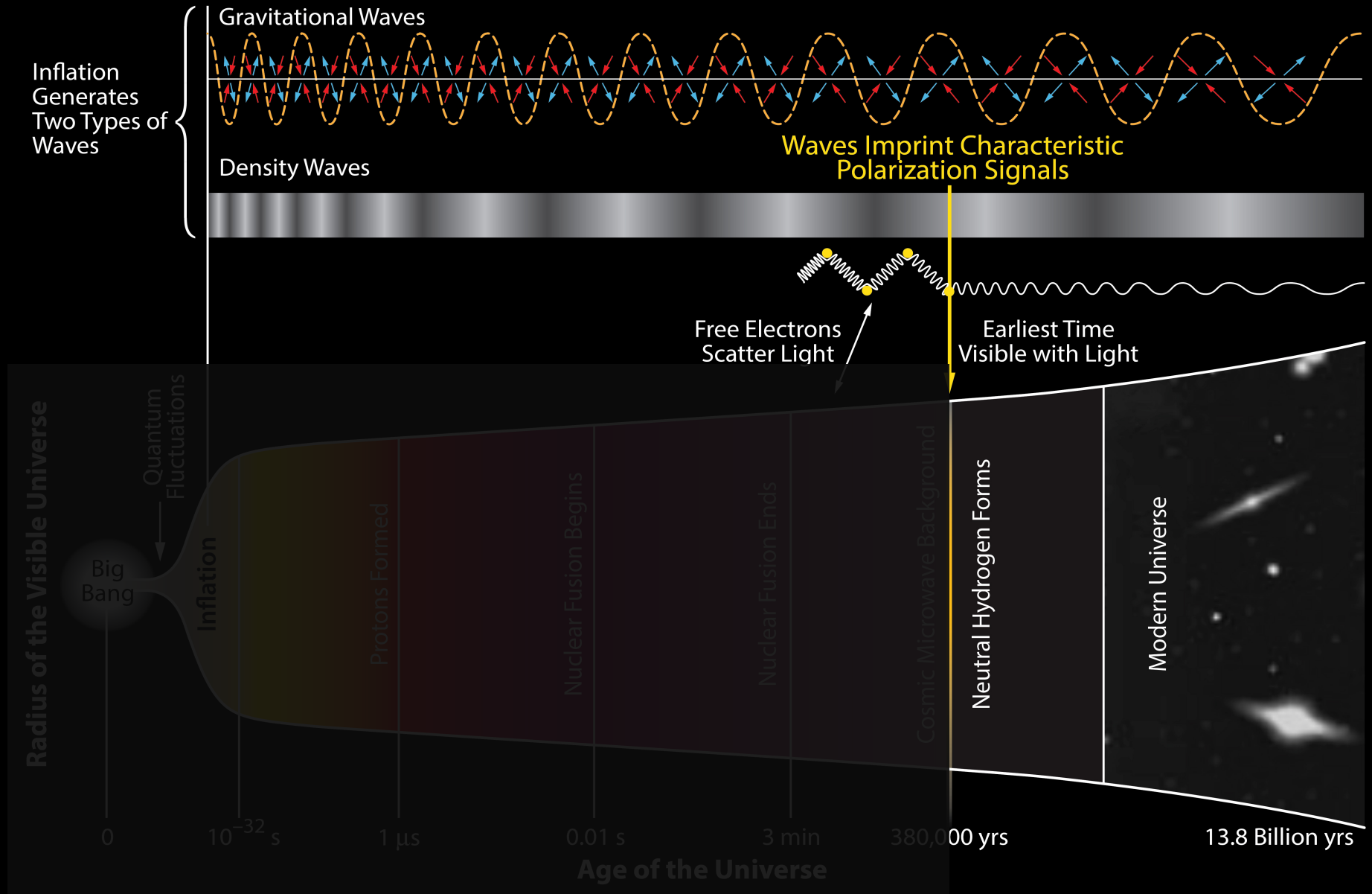
History of the Universe



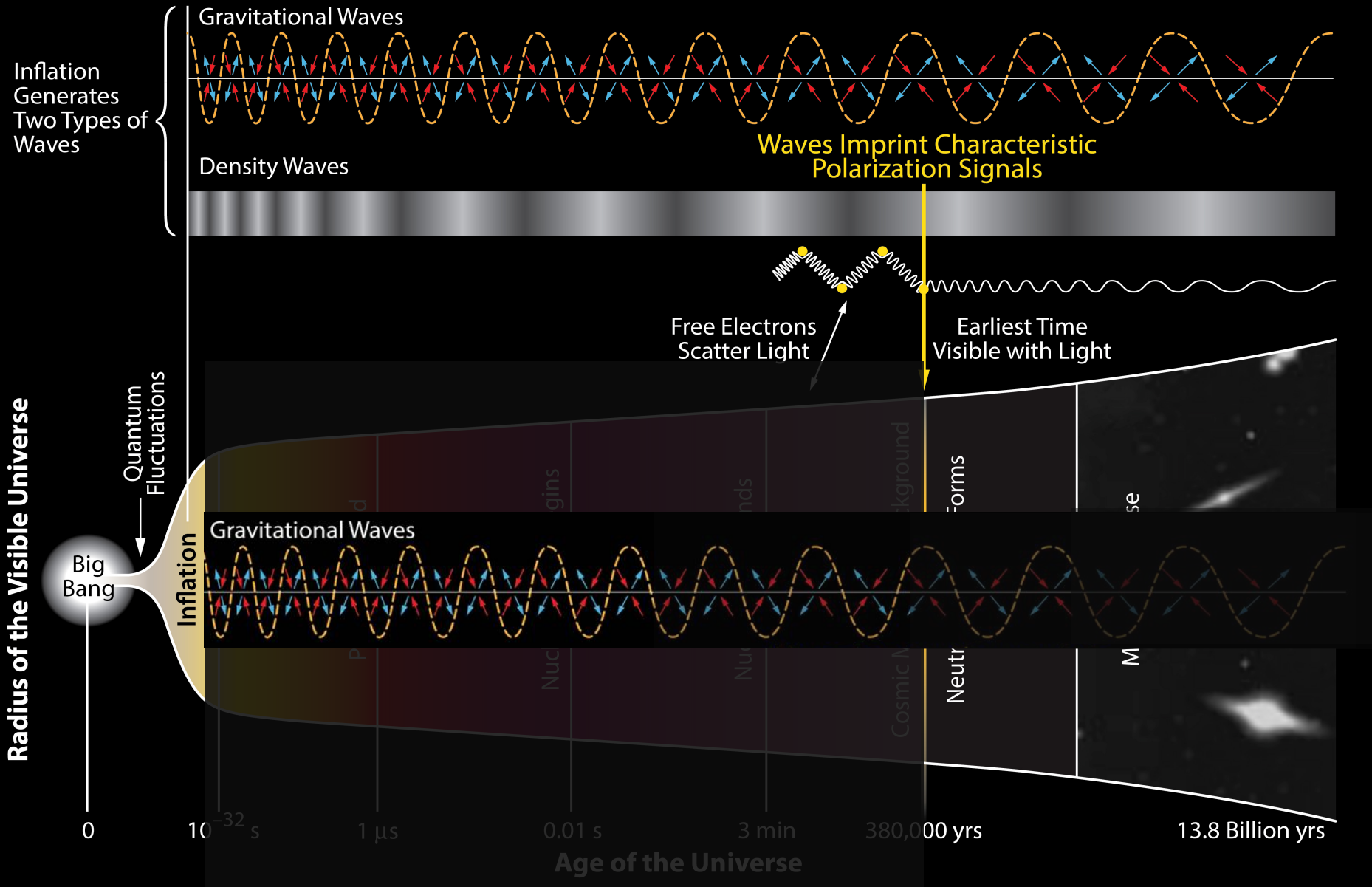
History of the Universe



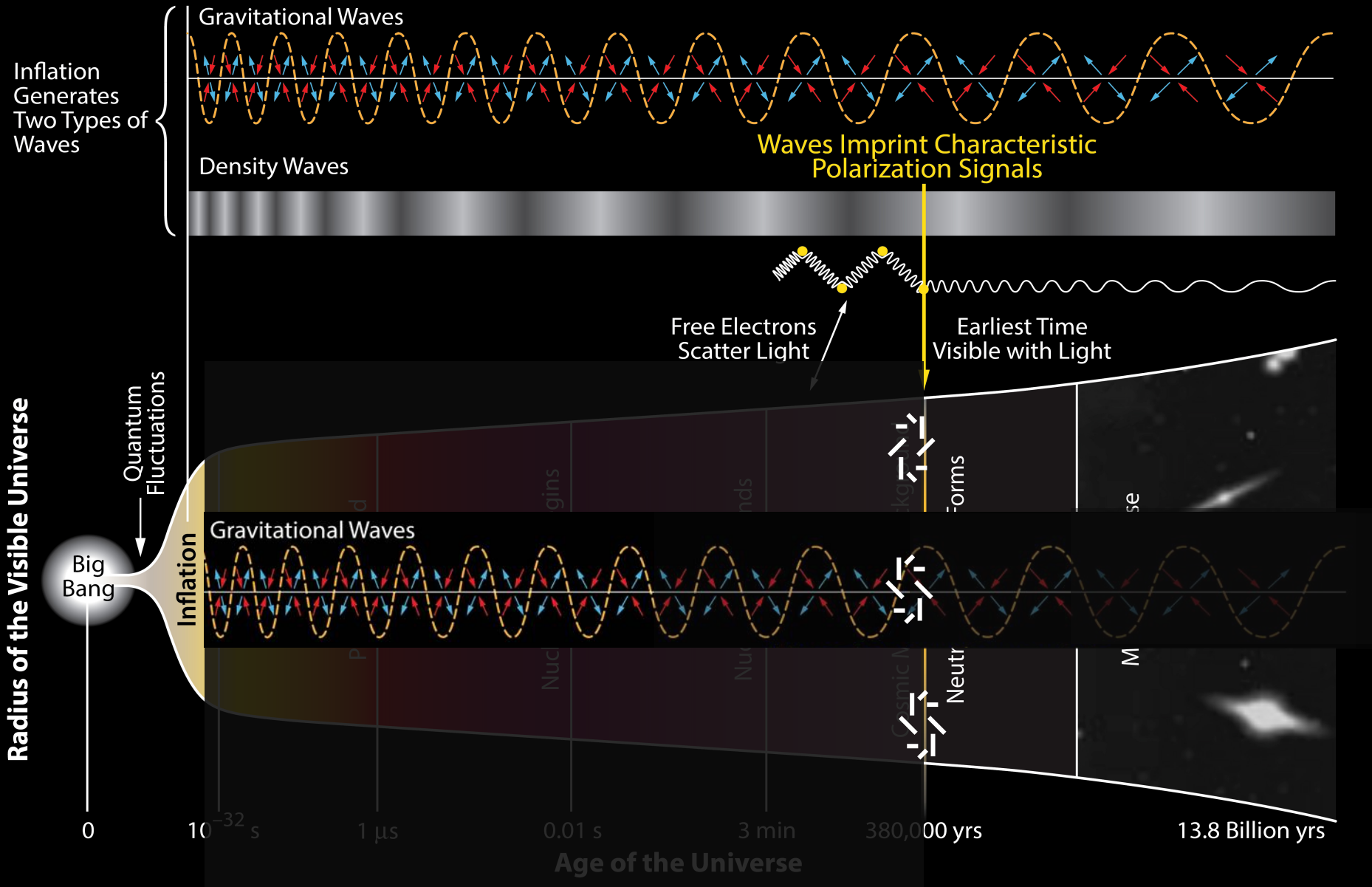
History of the Universe



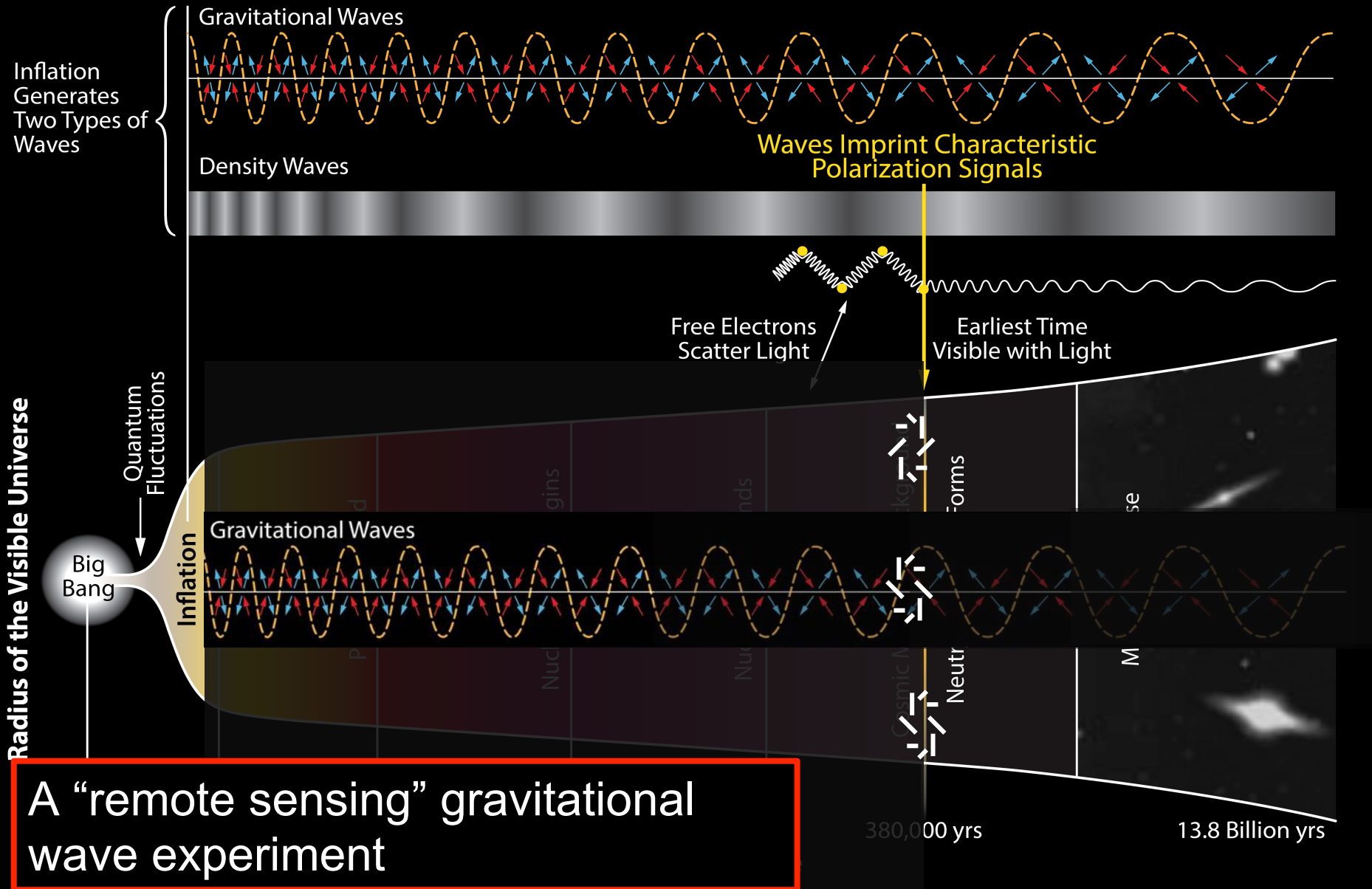
History of the Universe



History of the Universe



History of the Universe



CMB Polarization, B-modes and r

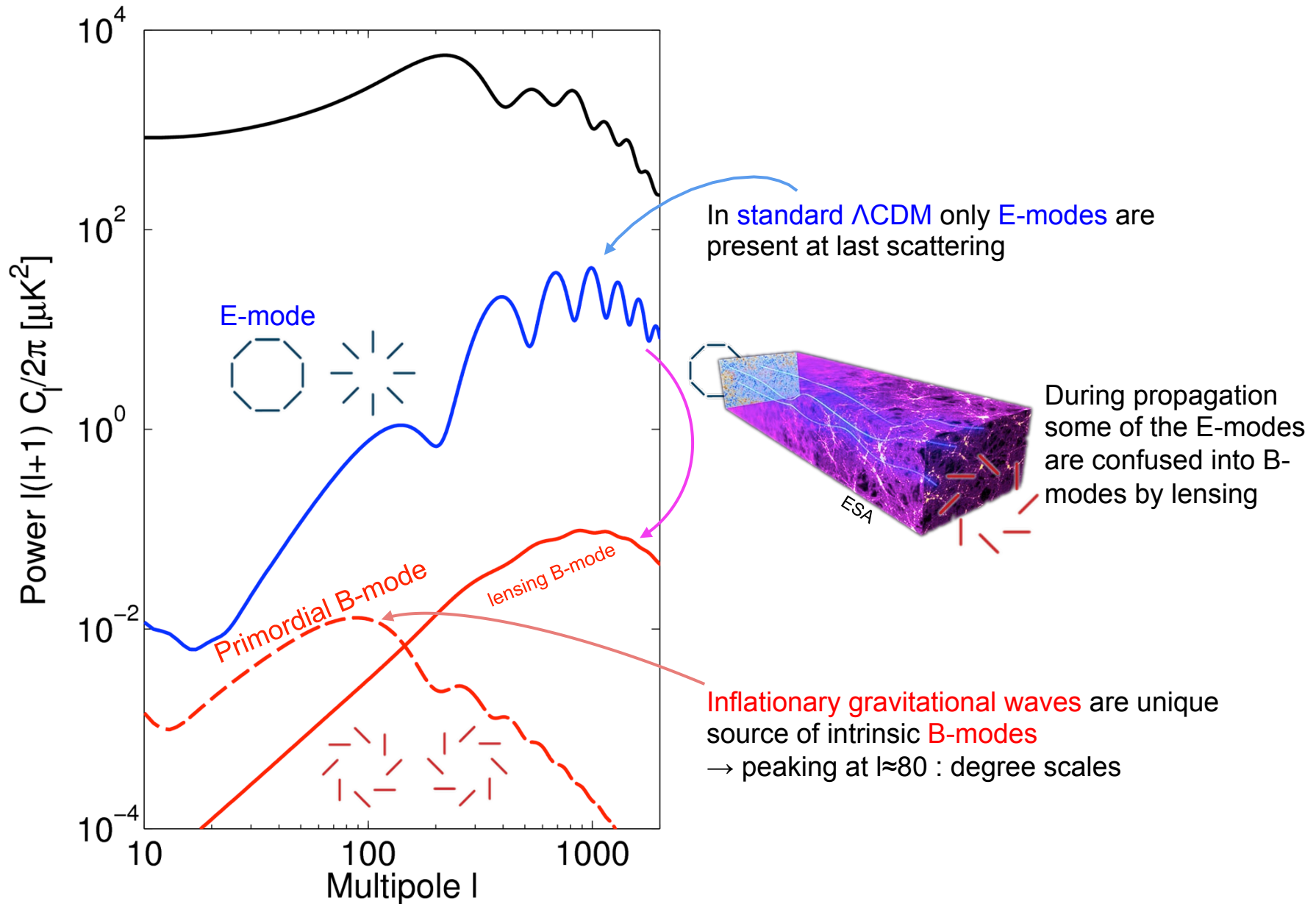
- The CMB is partially polarized (due to local radiation quadrupoles at last scattering)
- Any polarization pattern can be decomposed into E-modes (gradient modes) and B-modes (curl modes)
- Basic LCDM makes only E-modes at last scattering – although lensing deflections in flight produce a bit of a B-mode
- Primordial gravitational waves produce both E-modes and B-modes – but best to look for the B-modes since most distinct there
- Theory gives us a good template shape for the gravitational wave signal – but it does *not* tell us the amplitude
- The amplitude is parameterized by a single number r
- A wide range of inflation theories exist – the simplest are already ruled out – more complex ones can produce r which is undetectably small
- The experimental mission is to obtain the best possible sensitivity to r
- If we can detect r we determine the energy scale of inflation – if not we can rule out additional inflationary models

CMB Polarization, B-modes and r

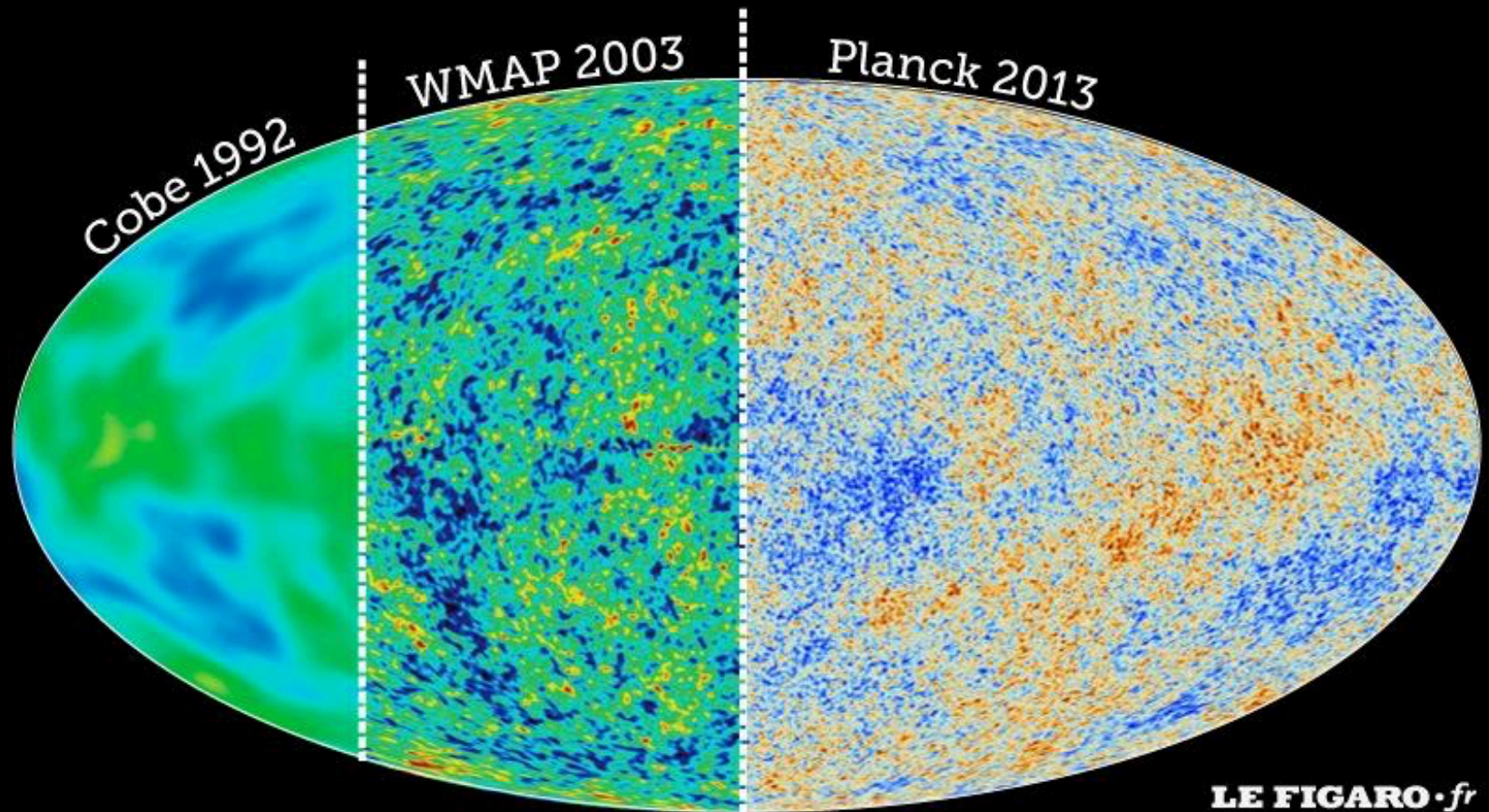
- The CMB is partially polarized (due to local radiation quadrupoles at last scattering)
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- The amplitude is parameterized by a single number r
- A wide range of inflation theories ruled out – more complex ones small
- The experimental mission is to
- If we can detect r we determine we can rule out additional inflation

Warning: It's a bit like the search for proton decay – a well motivated physics target to look for, but theories can be adjusted to make the amplitude arbitrarily small...

CMB Polarization power spectra

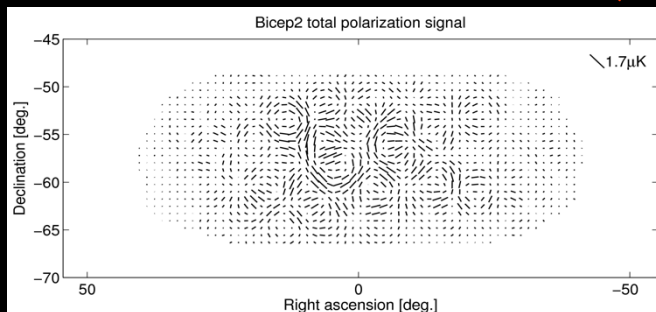
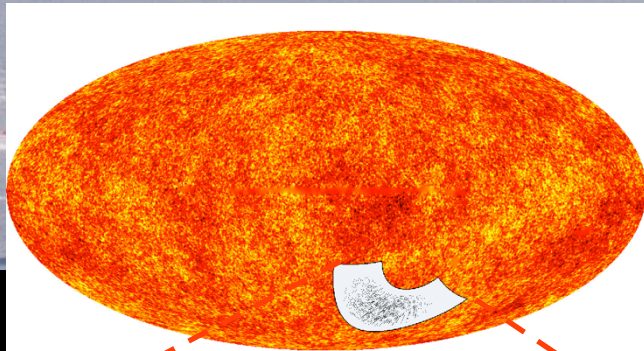


CMB space missions map the full sky

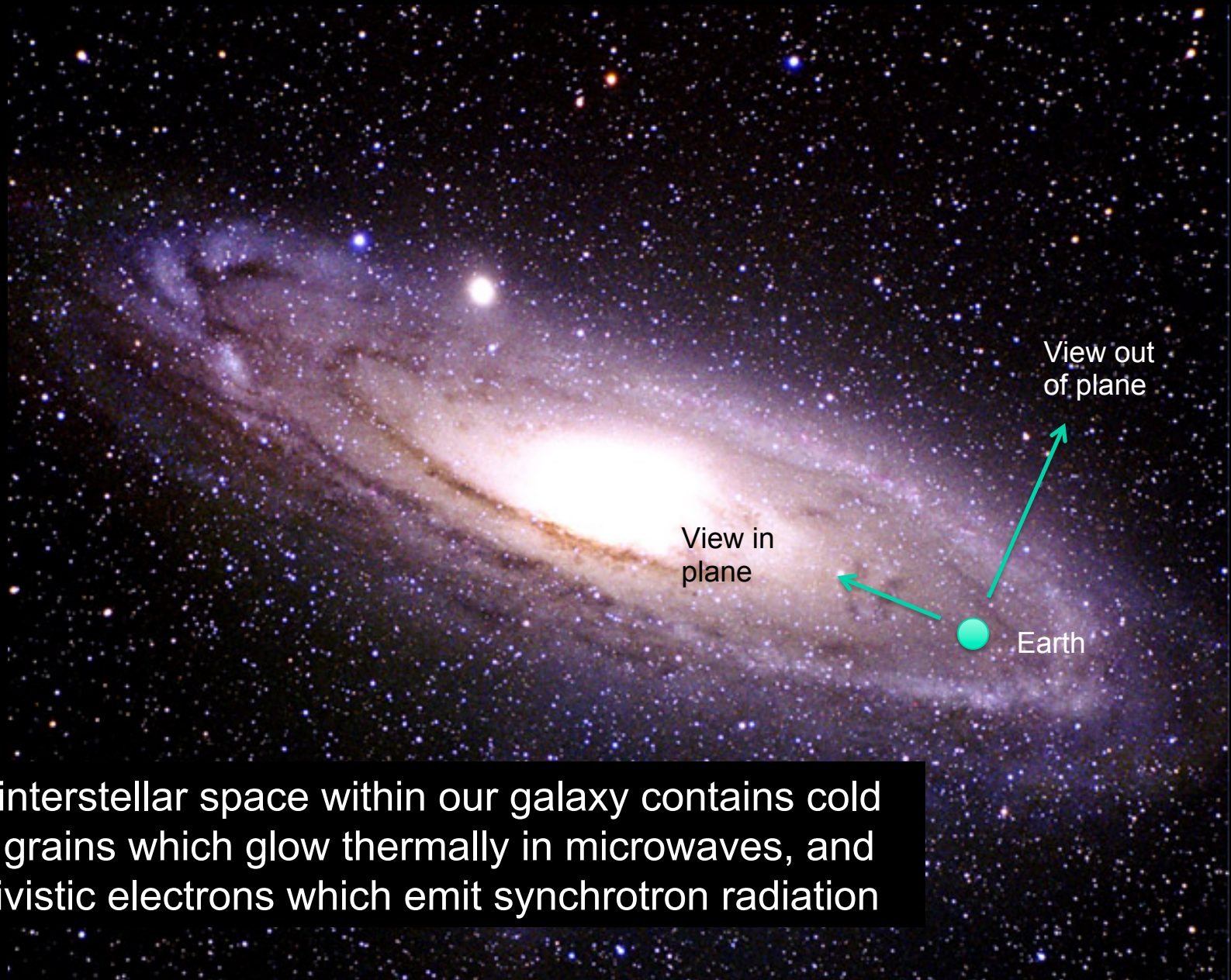


Ground based telescopes map part of the sky more deeply

South Pole CMB telescopes



Foreground emission from our galaxy



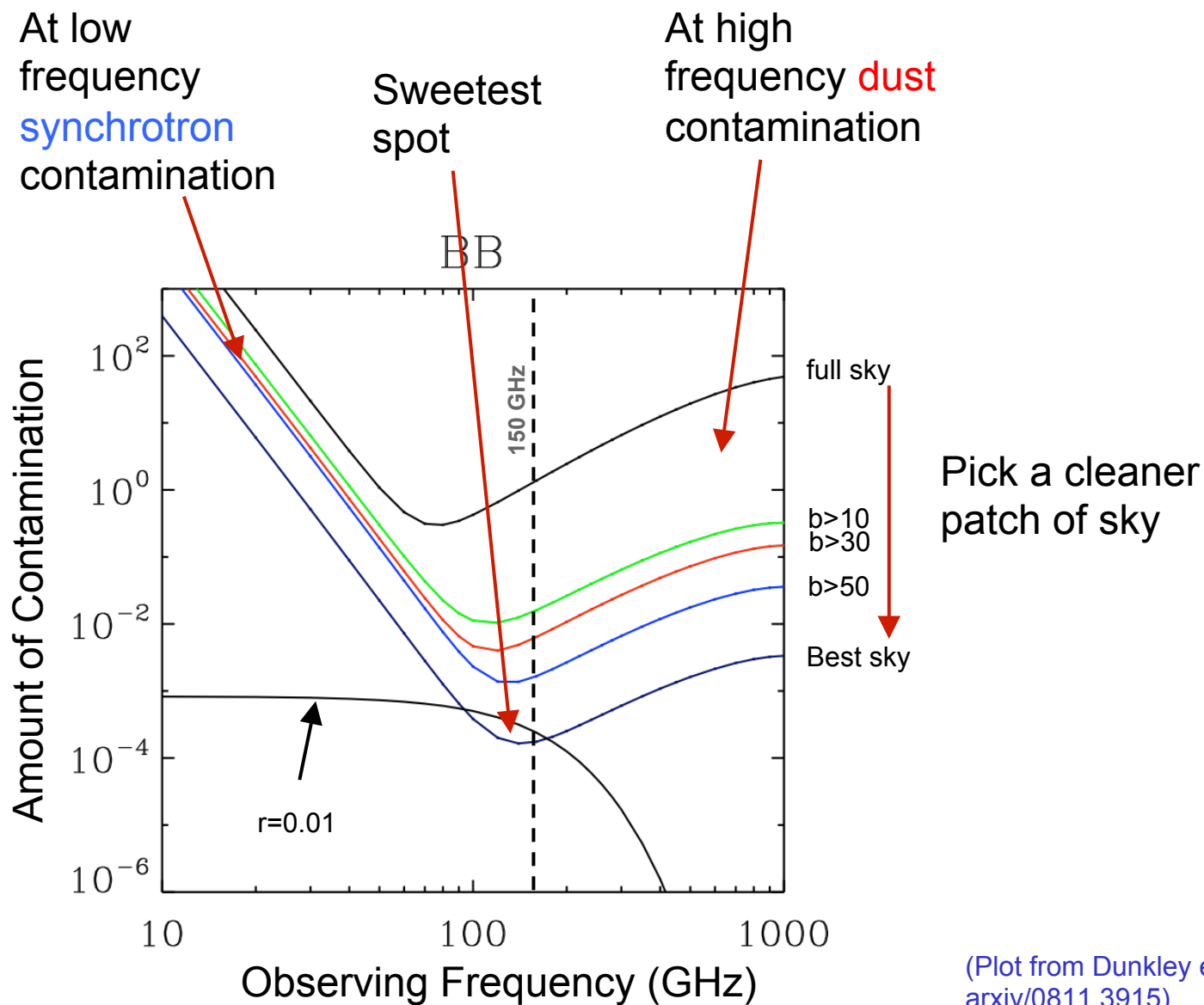
The interstellar space within our galaxy contains cold dust grains which glow thermally in microwaves, and relativistic electrons which emit synchrotron radiation

Foreground emission from our galaxy



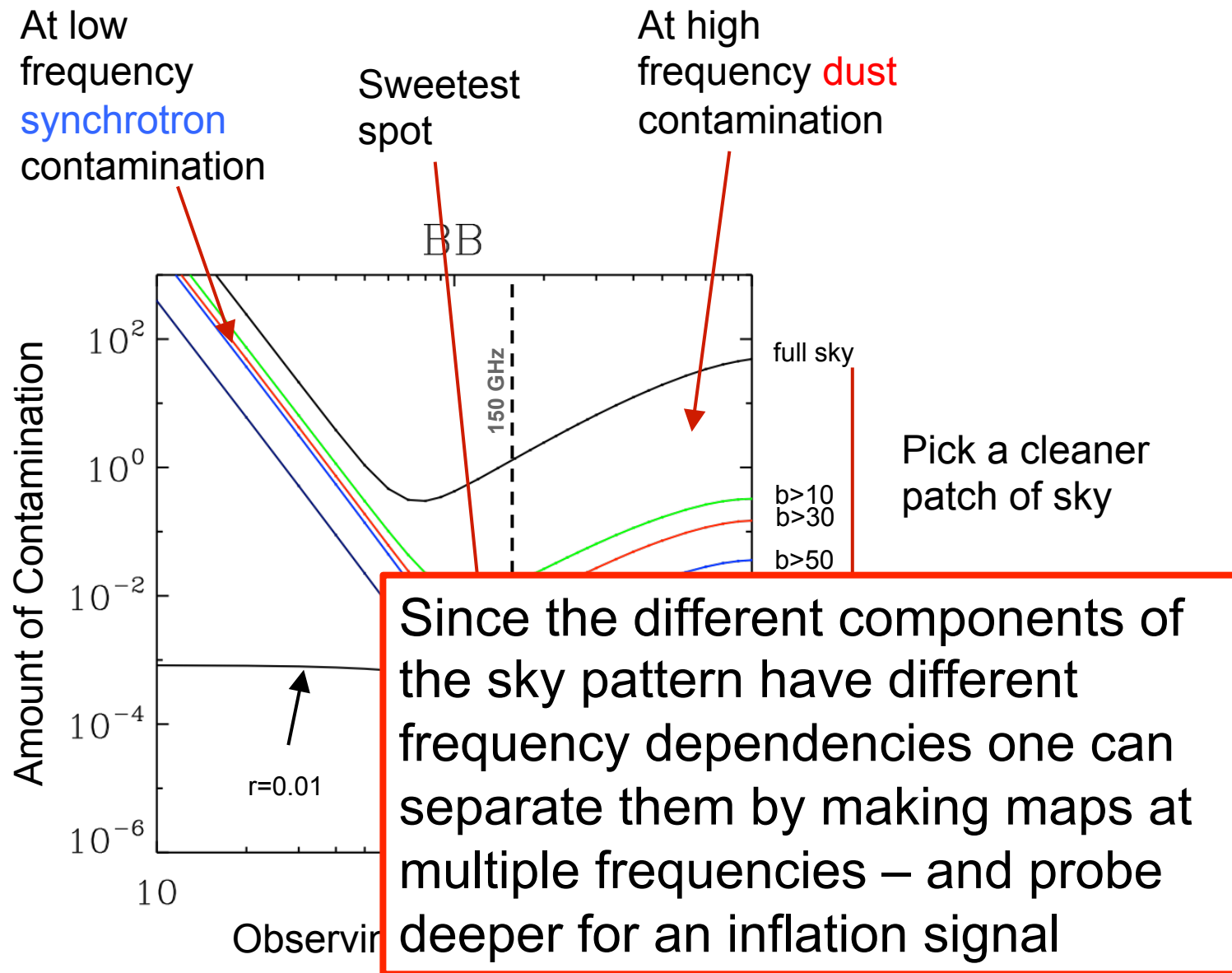
When CMB people talk about “foregrounds” it is analogous to what HEP people call “backgrounds” – something which gets in the way of the thing one is trying to measure.

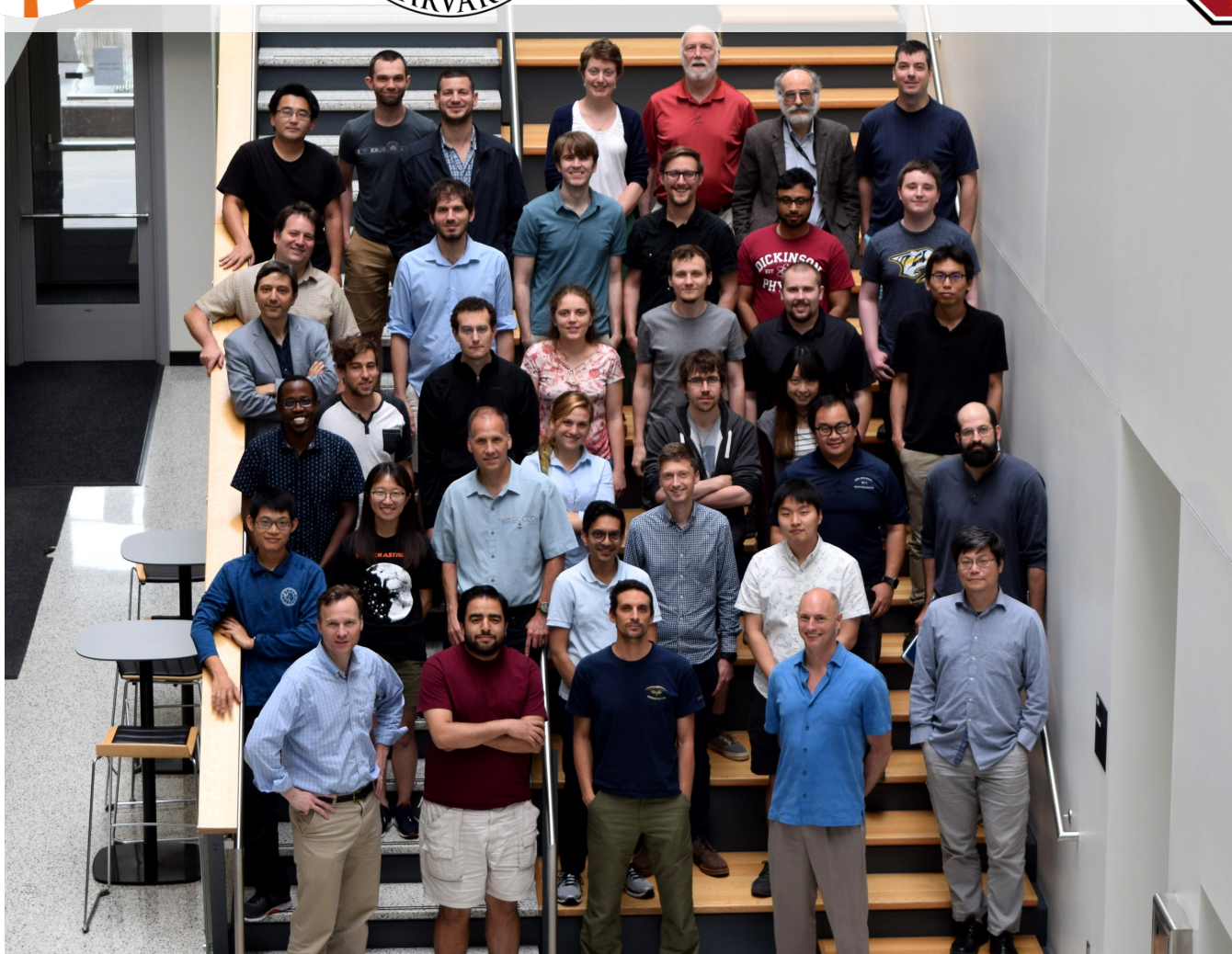
Polarized Foreground Contamination from Our Galaxy



(Plot from Dunkley et al
[arxiv/0811.3915](https://arxiv.org/abs/0811.3915))

Polarized Foreground Contamination from Our Galaxy

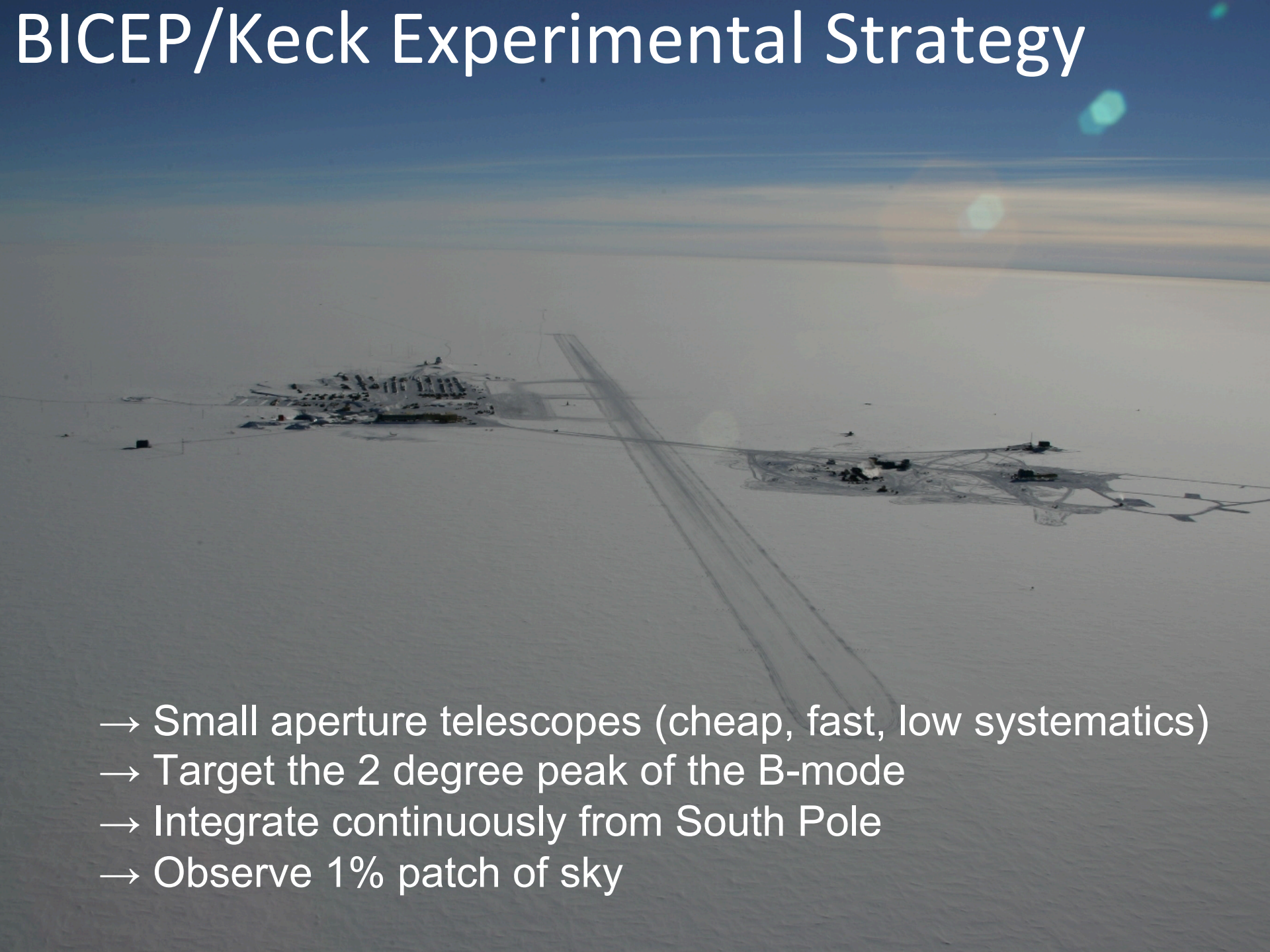




UNIVERSITY OF
TORONTO



BICEP/Keck Experimental Strategy



- Small aperture telescopes (cheap, fast, low systematics)
- Target the 2 degree peak of the B-mode
- Integrate continuously from South Pole
- Observe 1% patch of sky

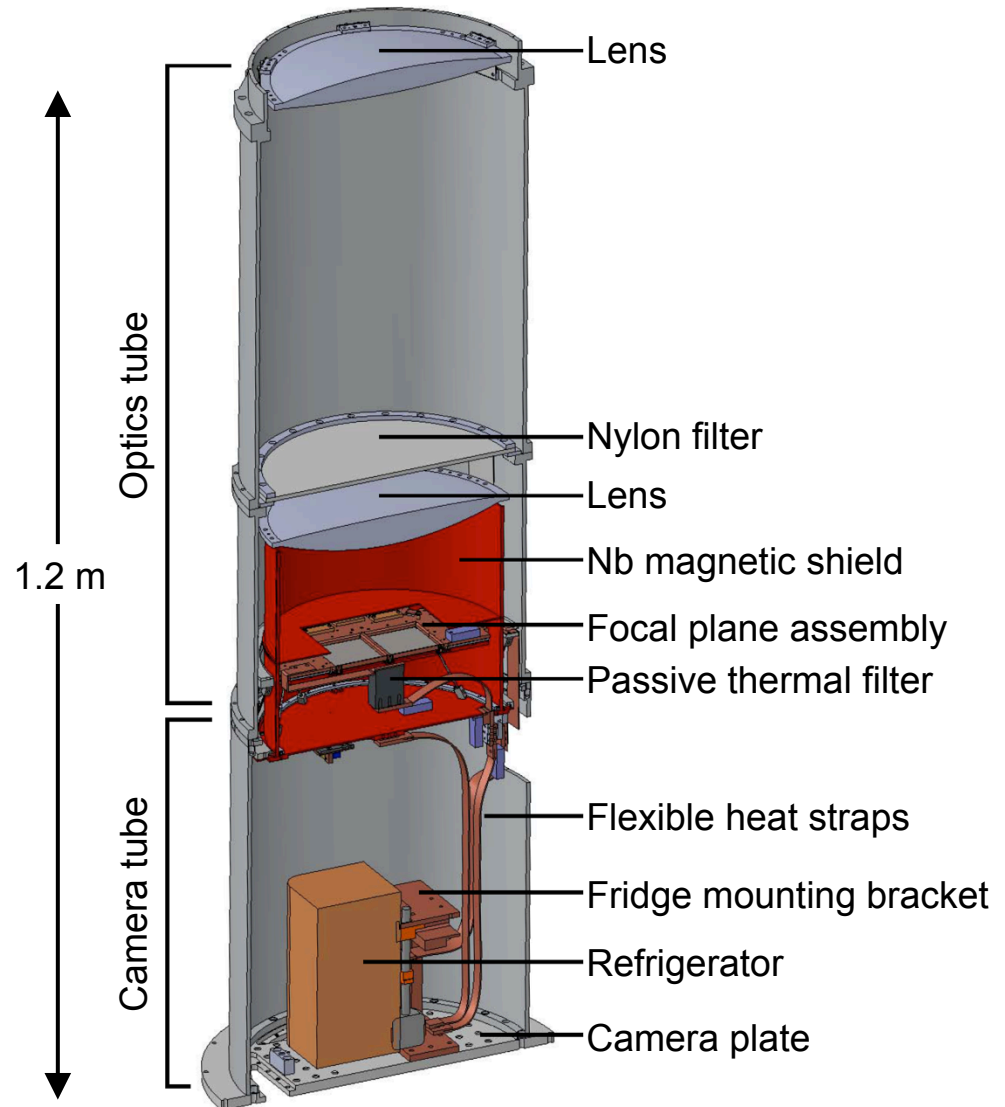
The BICEP/Keck Telescopes

Telescope as compact as possible while still having the angular resolution to observe degree-scale features

On-axis, refractive optics allow the entire telescope to rotate around boresight for polarization modulation

Liquid helium/pulse tube cools the optical elements to 4 K

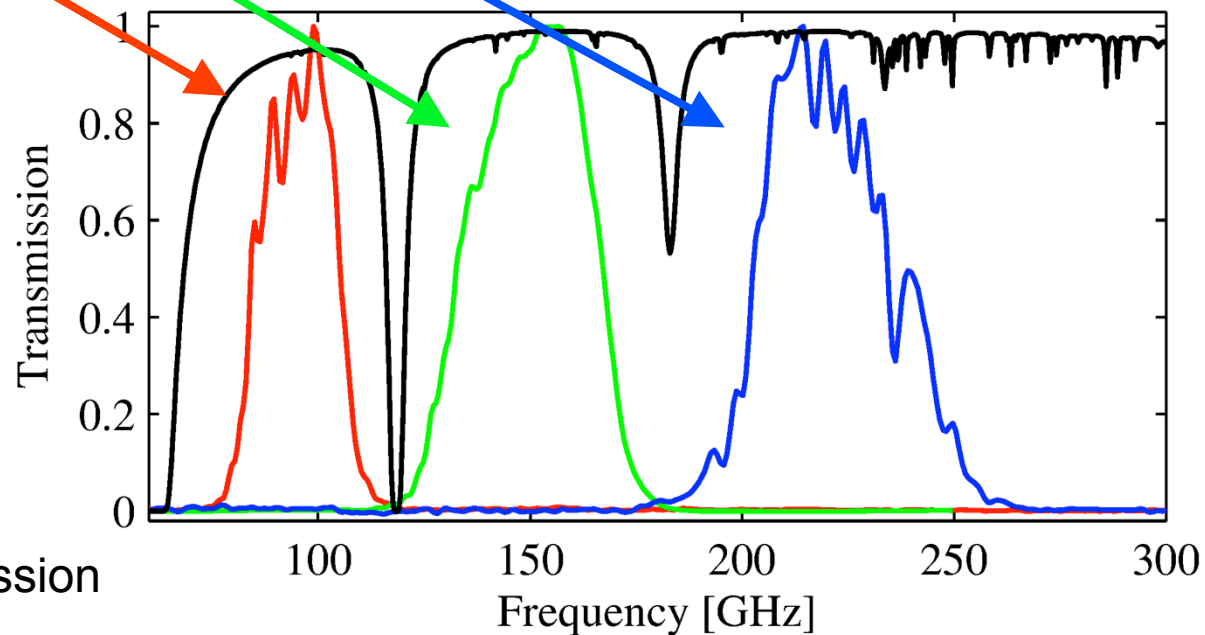
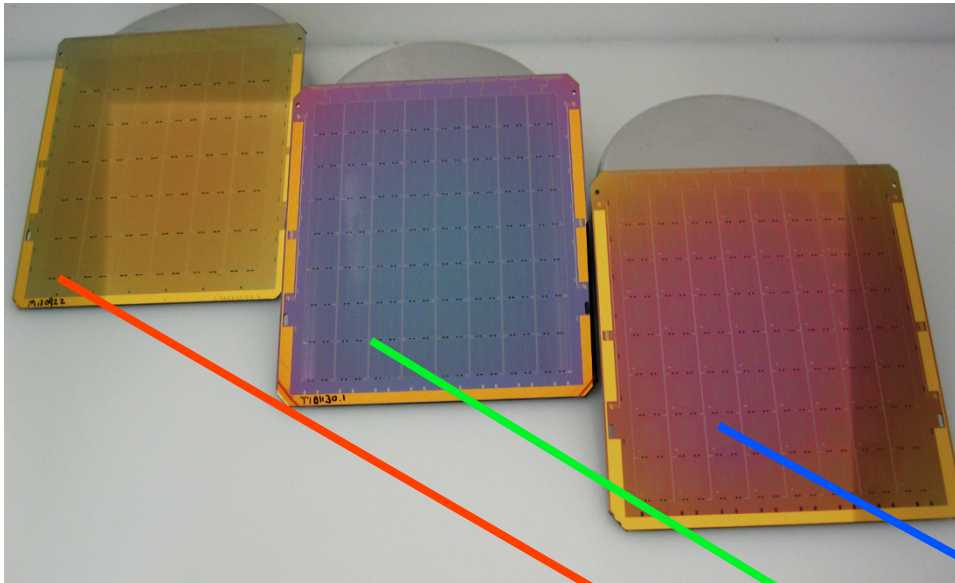
3-stage helium sorption refrigerator further cools the detectors to 0.27 K



Planar superconducting detector arrays

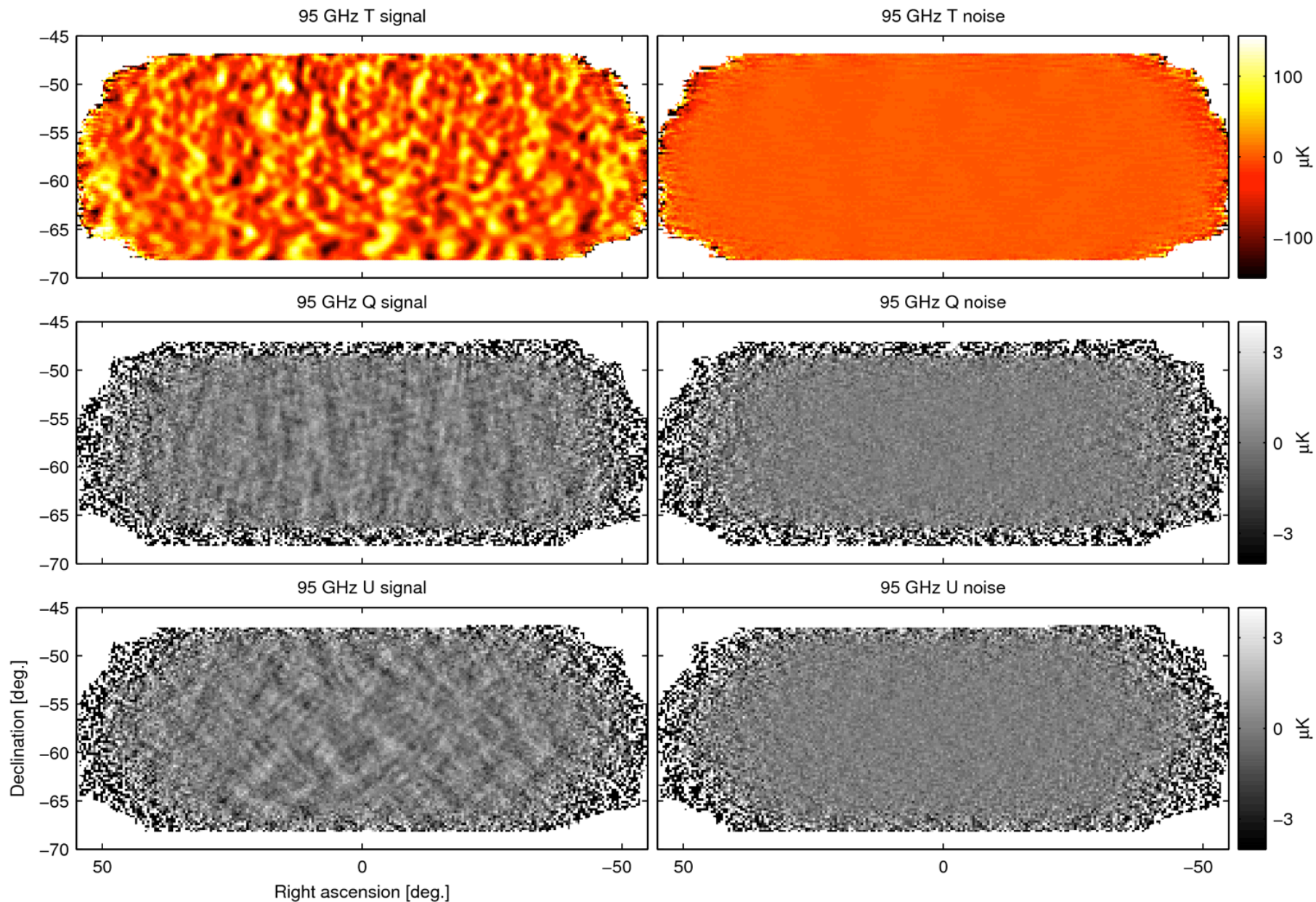
...designed to scale
in frequency

Up to 2013 – all 150GHz
2014 – 95/150GHz
2015 – 95/150/220GHz



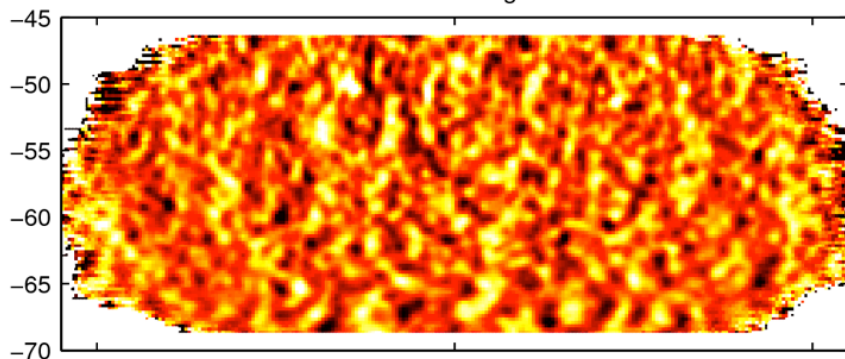
Typical South Pole
atmospheric transmission

BK15 95GHz Maps

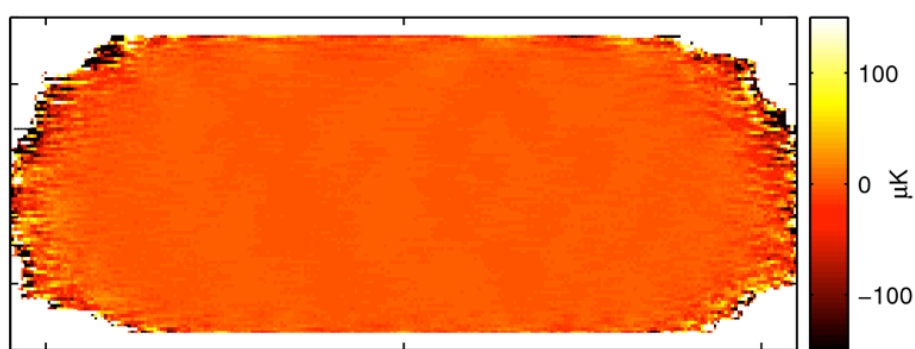


BK15 150GHz Maps

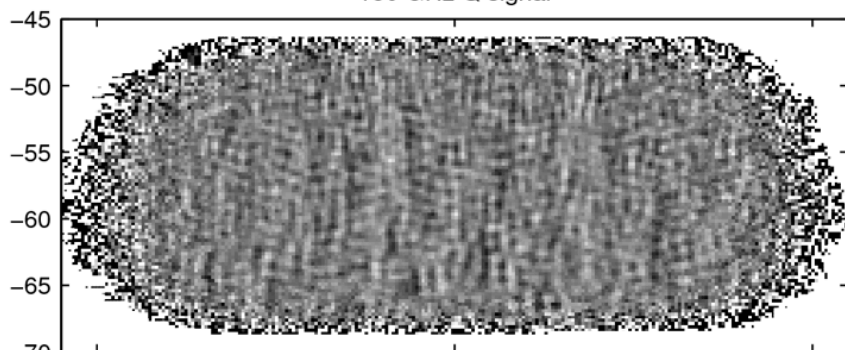
150 GHz T signal



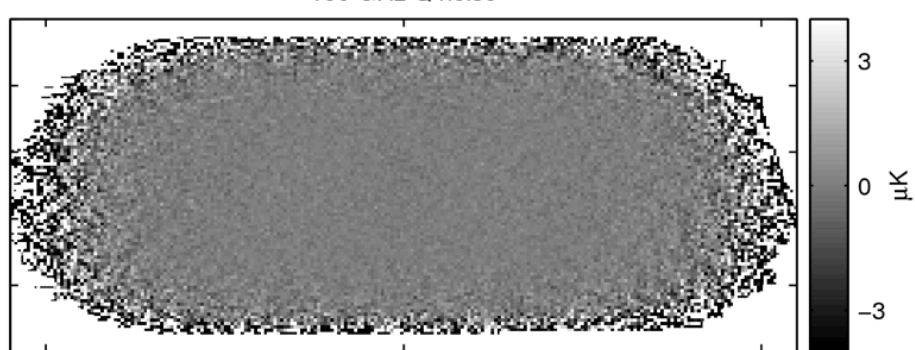
150 GHz T noise



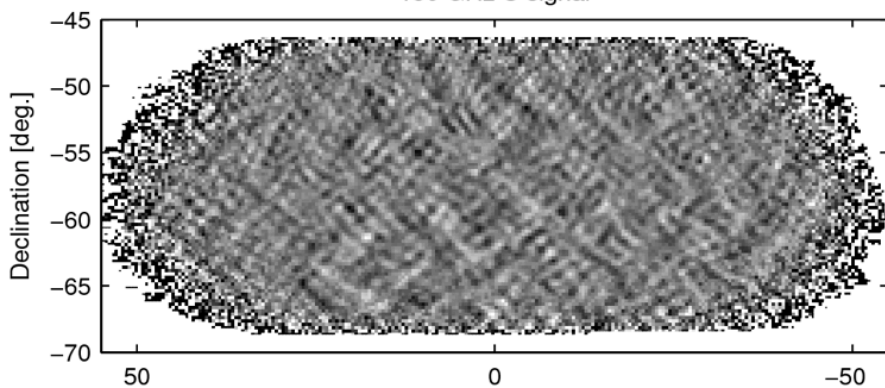
150 GHz Q signal



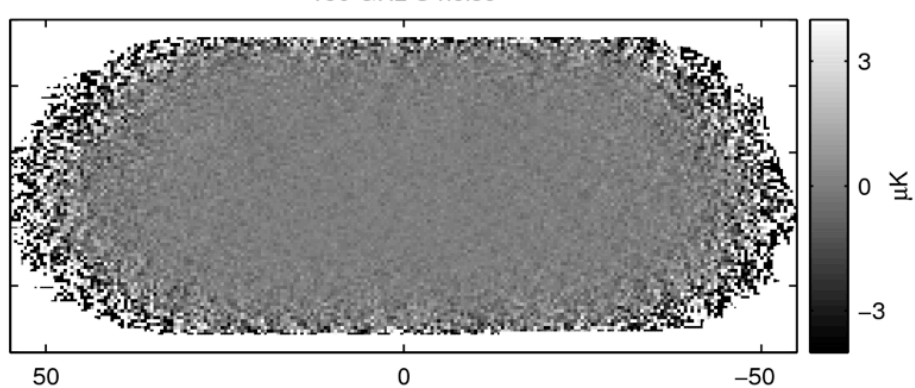
150 GHz Q noise



150 GHz U signal



150 GHz U noise

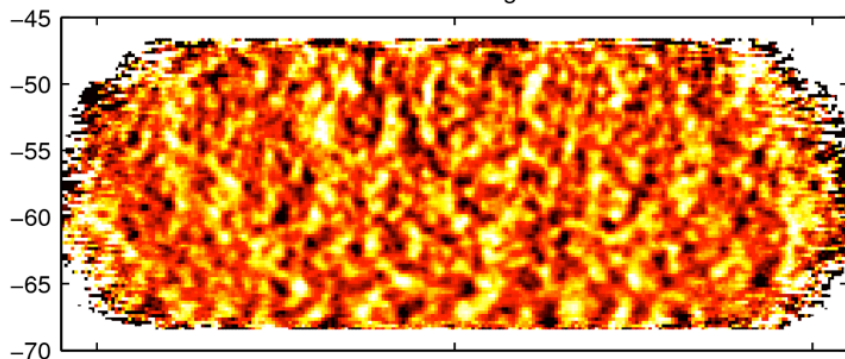


Right ascension [deg.]

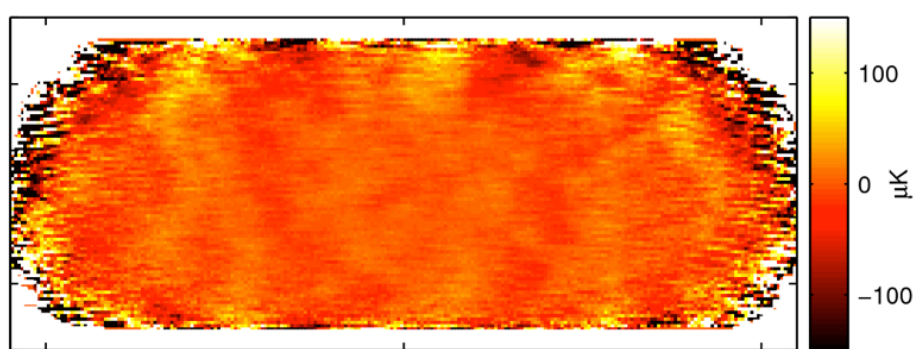
Declination [deg.]

BK15 220GHz Maps

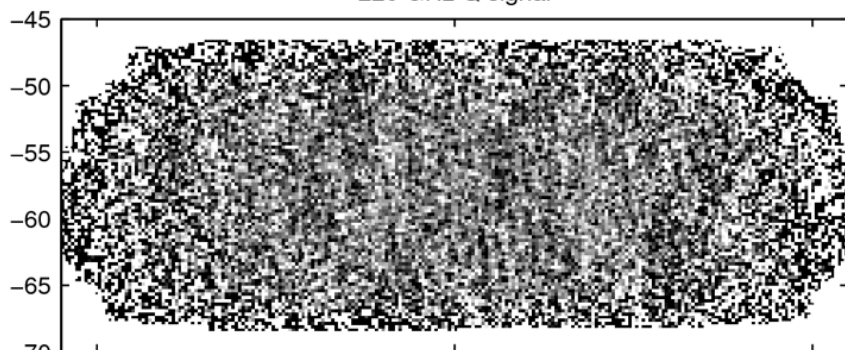
220 GHz T signal



220 GHz T noise



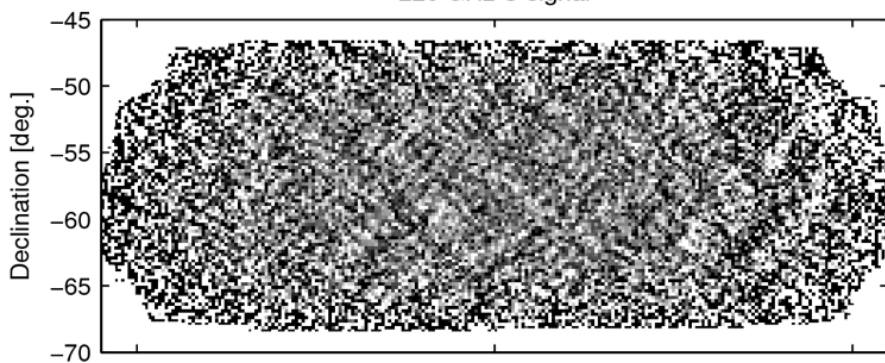
220 GHz Q signal



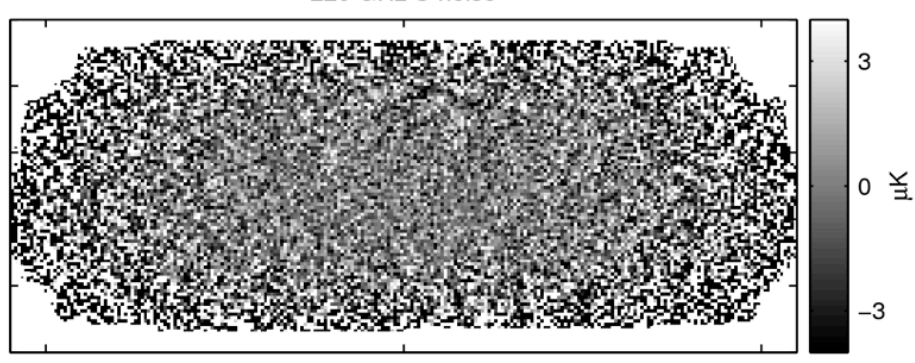
220 GHz Q noise



220 GHz U signal

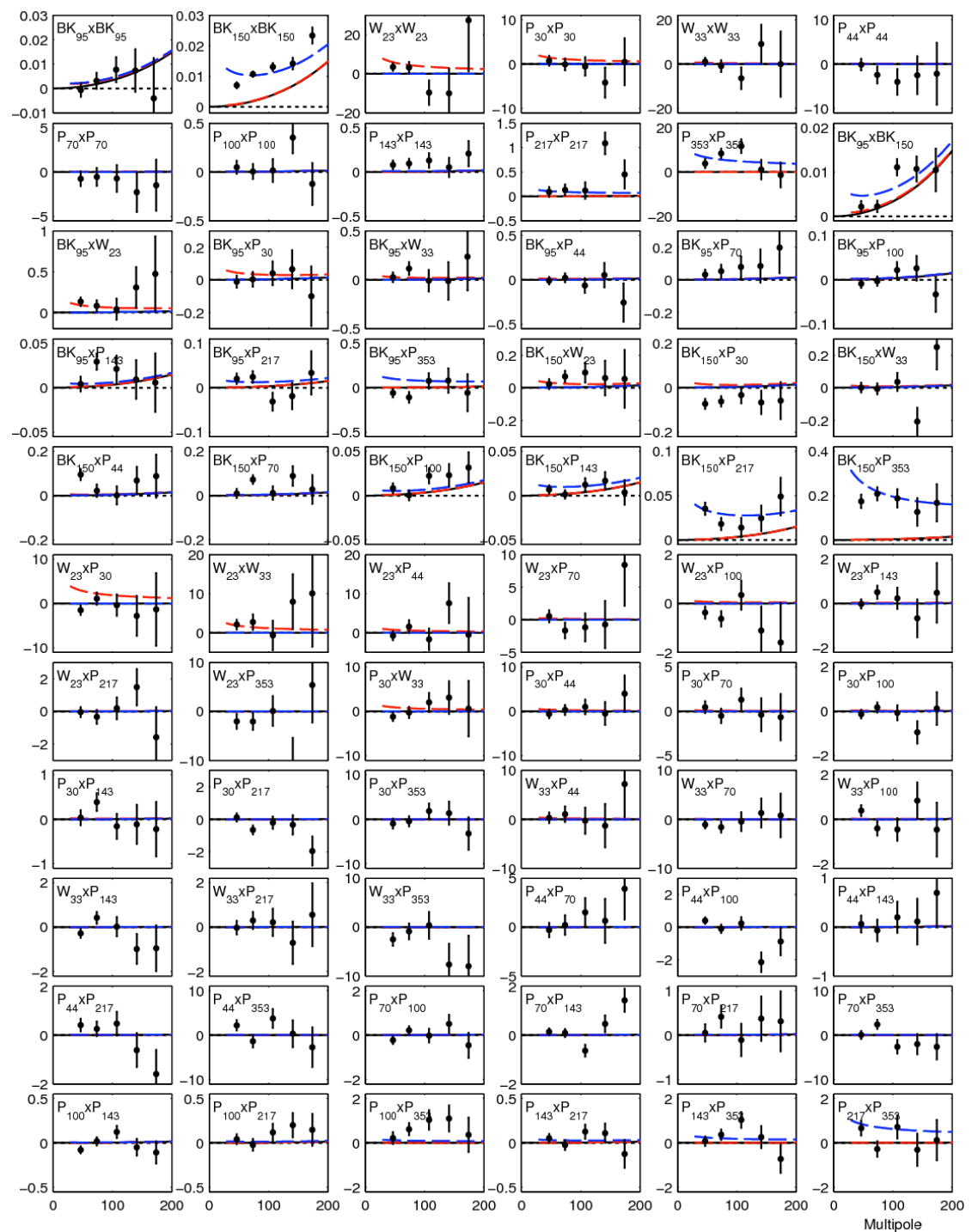


220 GHz U noise



Right ascension [deg.]

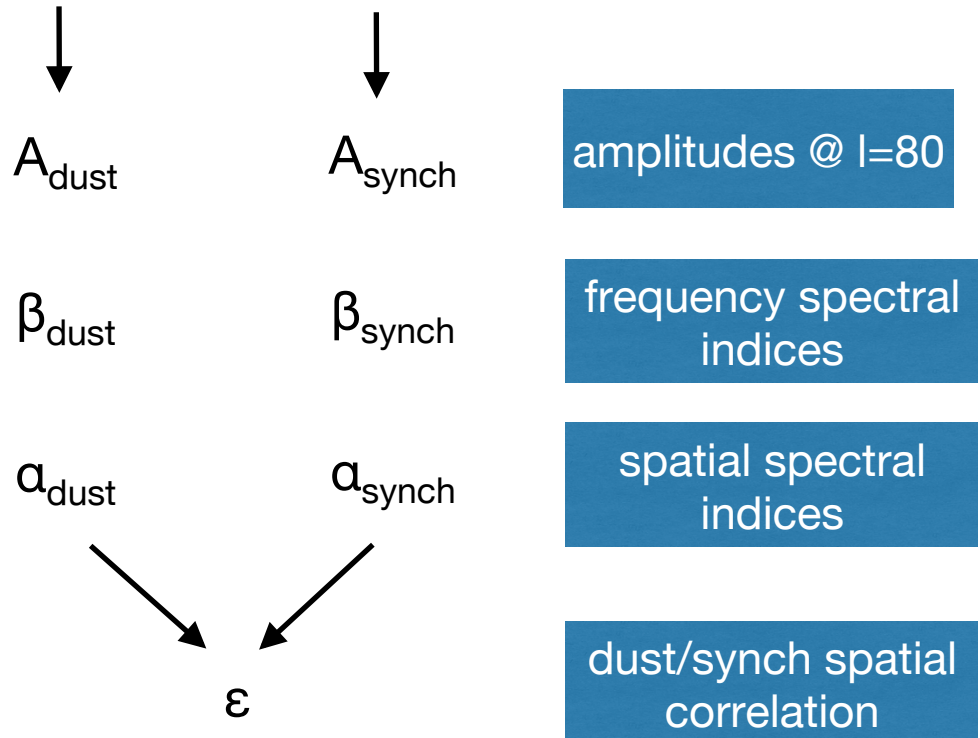
Take all possible
auto- and cross
spectra between
BICEP/Keck,
WMAP, and
Planck bands
(66 of them)



Multicomponent parametric likelihood analysis

Take the joint likelihood of all the spectra simultaneously vs. model for B-modes that is the Λ CDM lensing expectation + 7 parameter foreground model + r

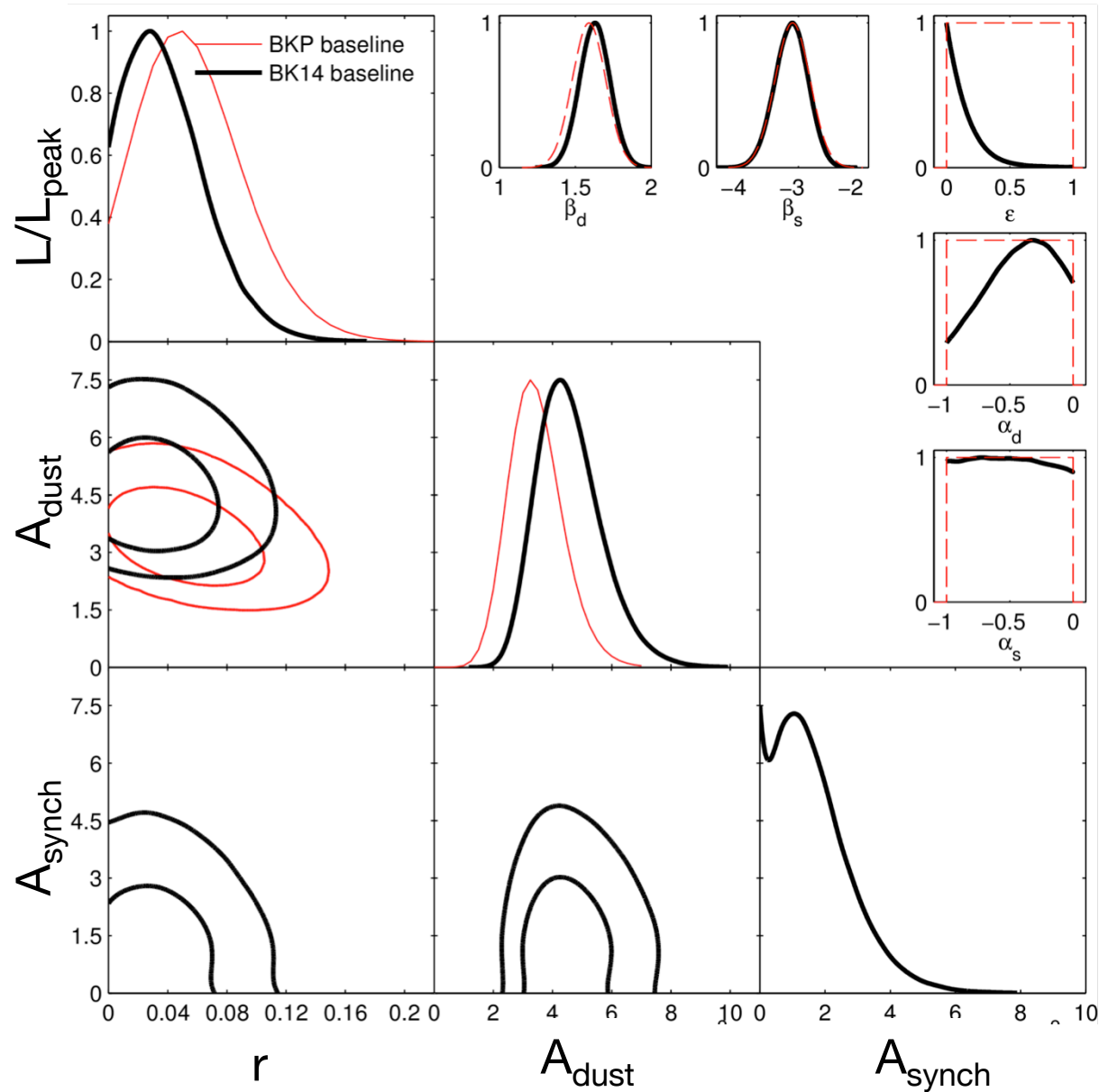
foreground model = dust + synchrotron



BK14 Results:

Put priors on the frequency spectral indices of dust & sync

Allow dust/sync correlation



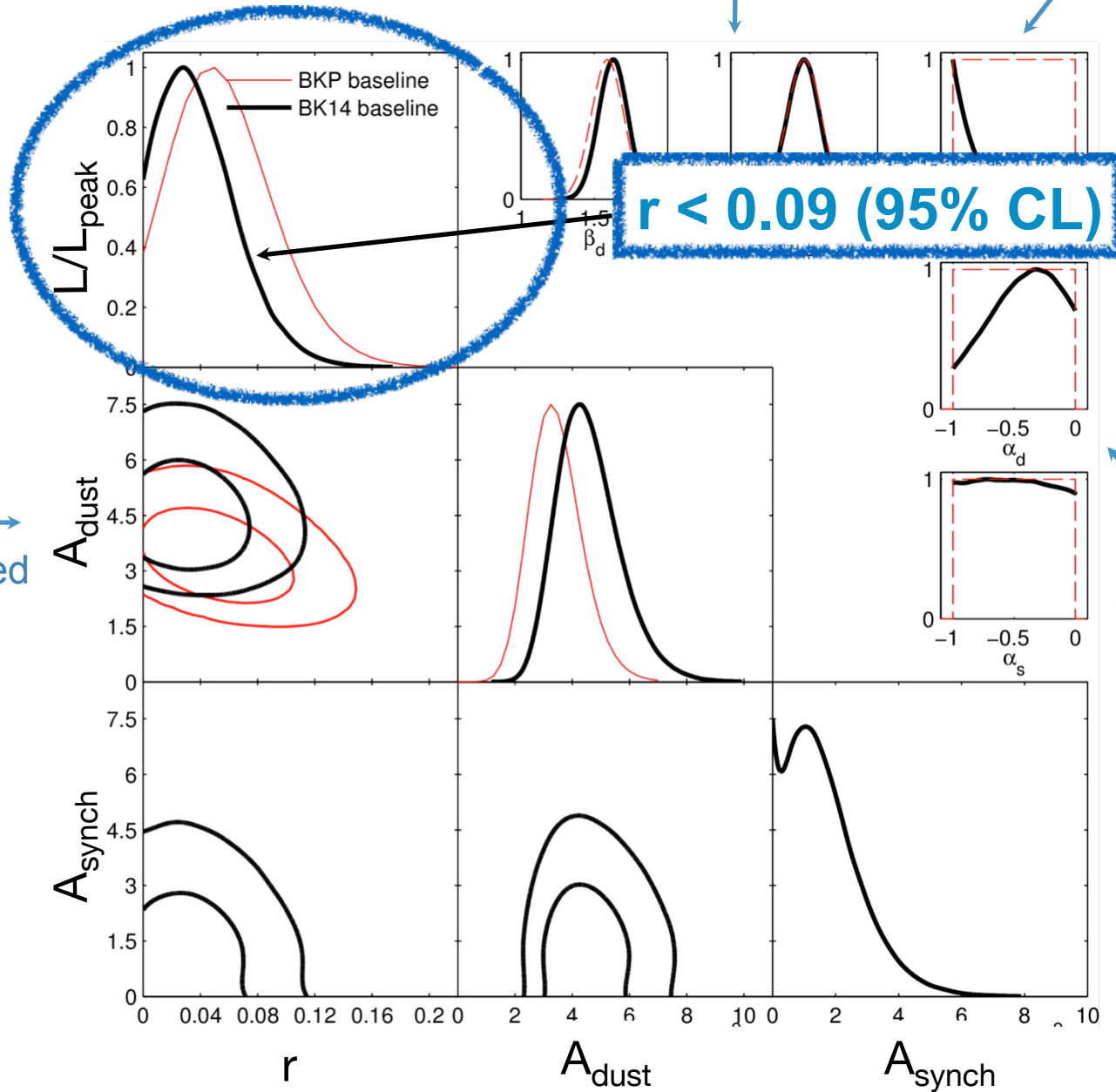
dust vs. r →
degeneracy lifted

Marginalize over
generous ranges in
spatial spectral
indices

BK14 Results:

Put priors on the frequency spectral indices of dust & sync

Allow dust/sync correlation

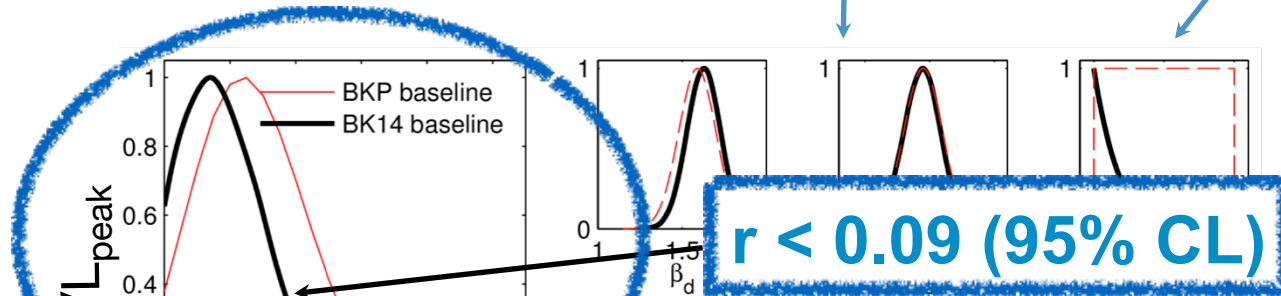


Marginalize over generous ranges in spatial spectral indices

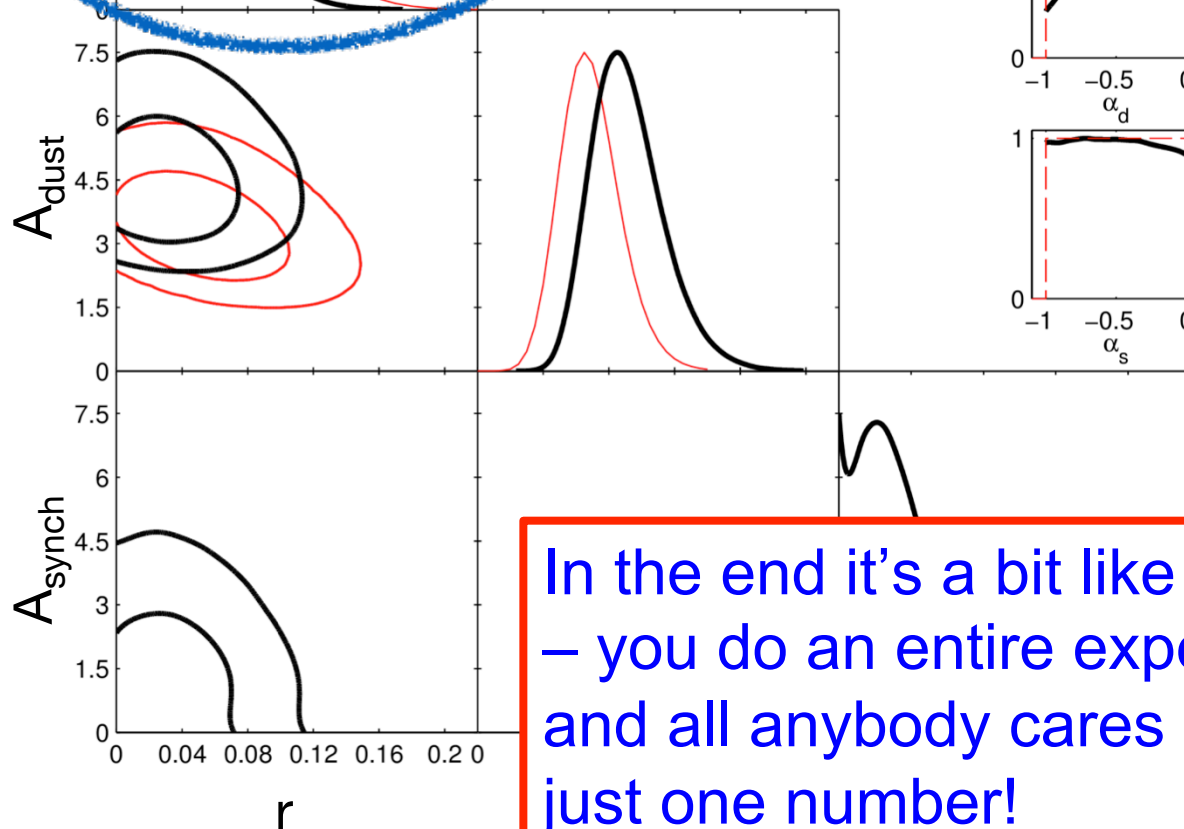
BK14 Results:

Put priors on the frequency spectral indices of dust & sync

Allow dust/sync correlation



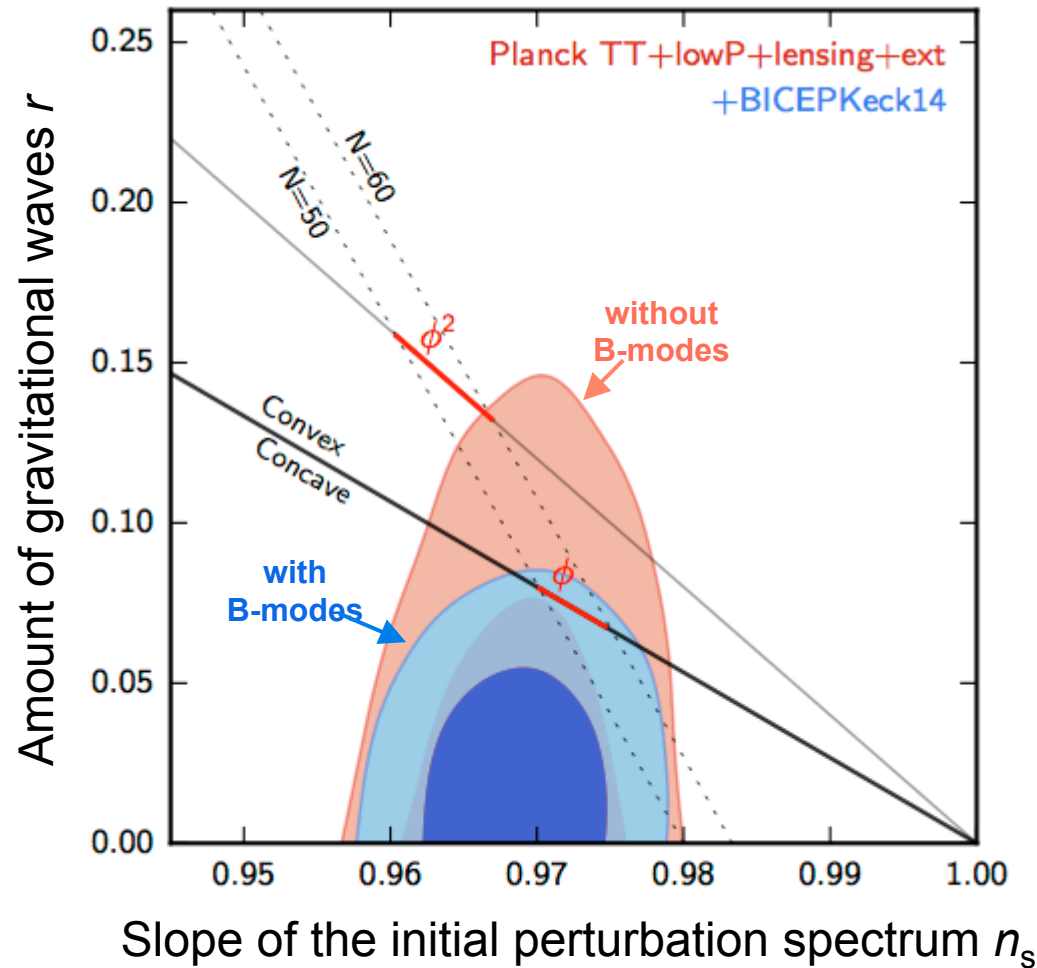
dust vs. r →
degeneracy lifted



Marginalize over
generous ranges in
spatial spectral
indices

In the end it's a bit like neutrinos
– you do an entire experiment
and all anybody cares about is
just one number!

Adding in Planck temperature measurements



**Steadily tightening the constraints on
inflationary models...**

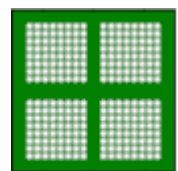
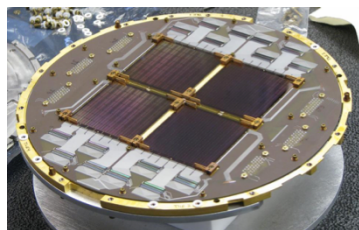
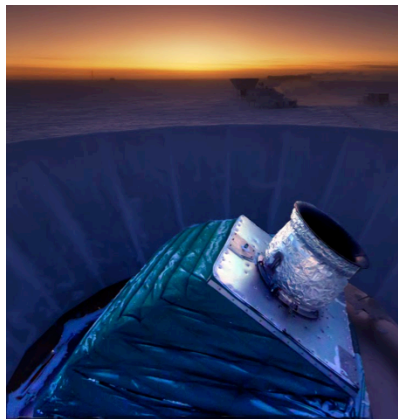
Telescope and Mount

Focal Plane

Beams on Sky

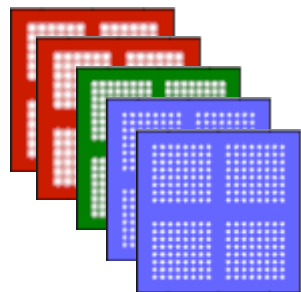
Stage 2

BICEP2
(2010-2012)



-5 0 5
Degrees on sky

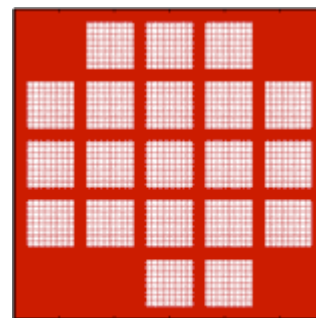
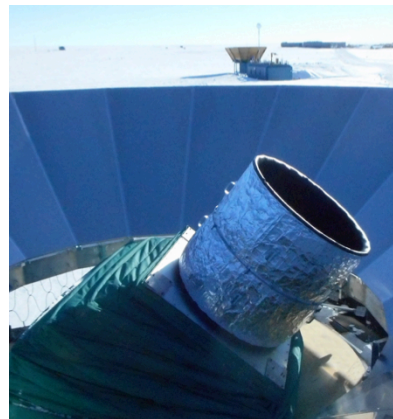
Keck Array
(2012-2019)



-5 0 5
Degrees on sky

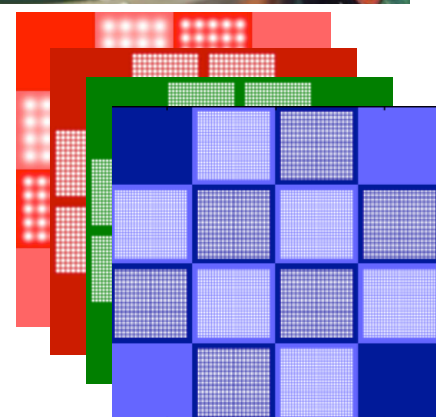
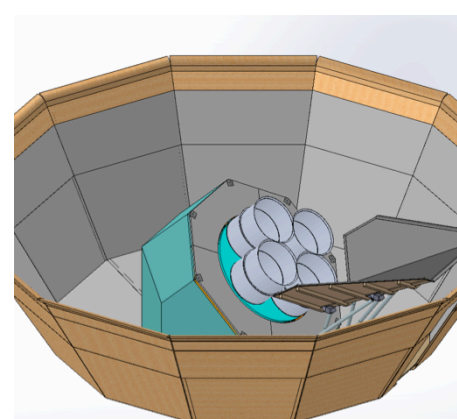
Stage 3

BICEP3
(2015-)



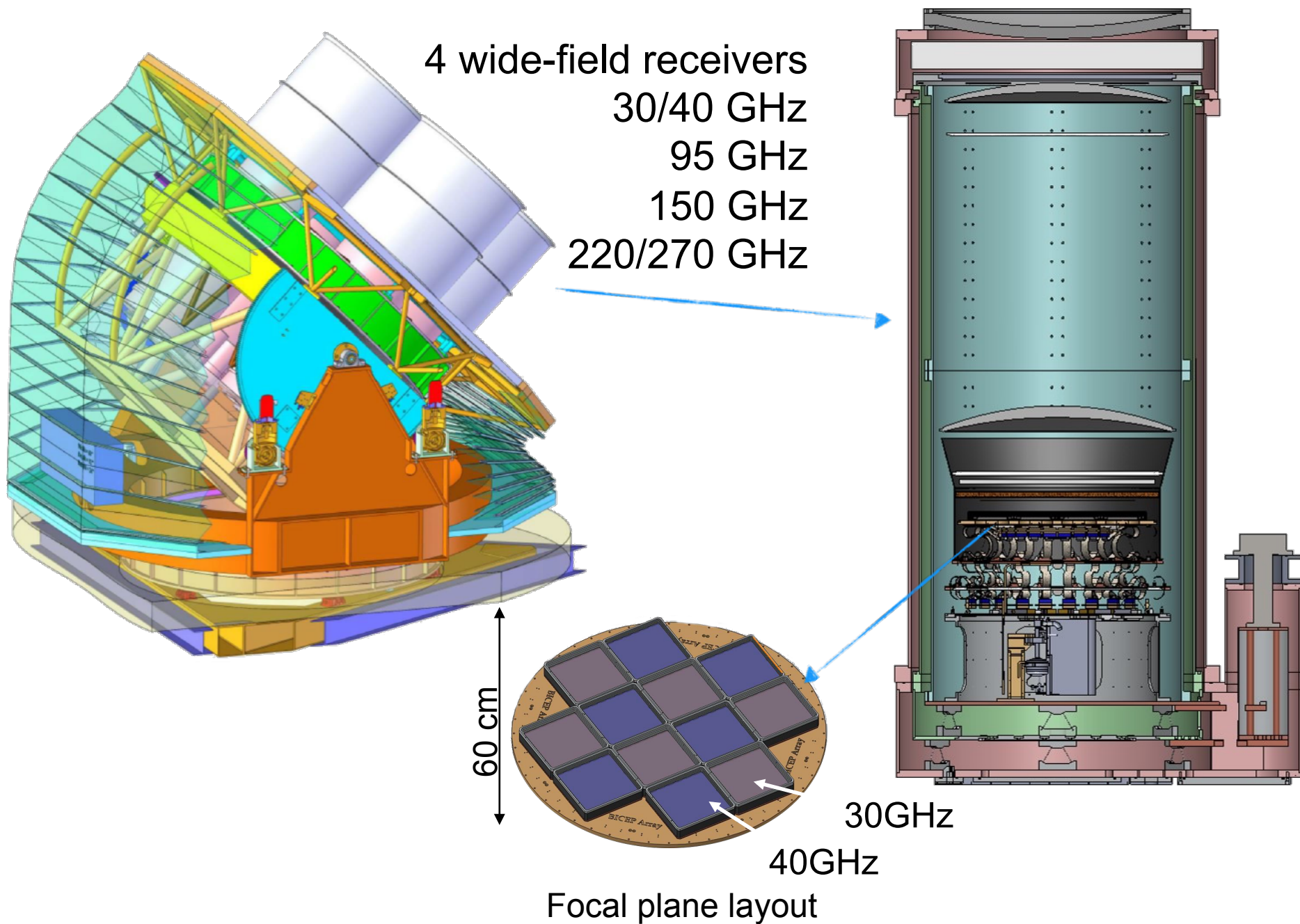
-10 -5 0 5 10
Degrees on sky

BICEP Array
(2020-)

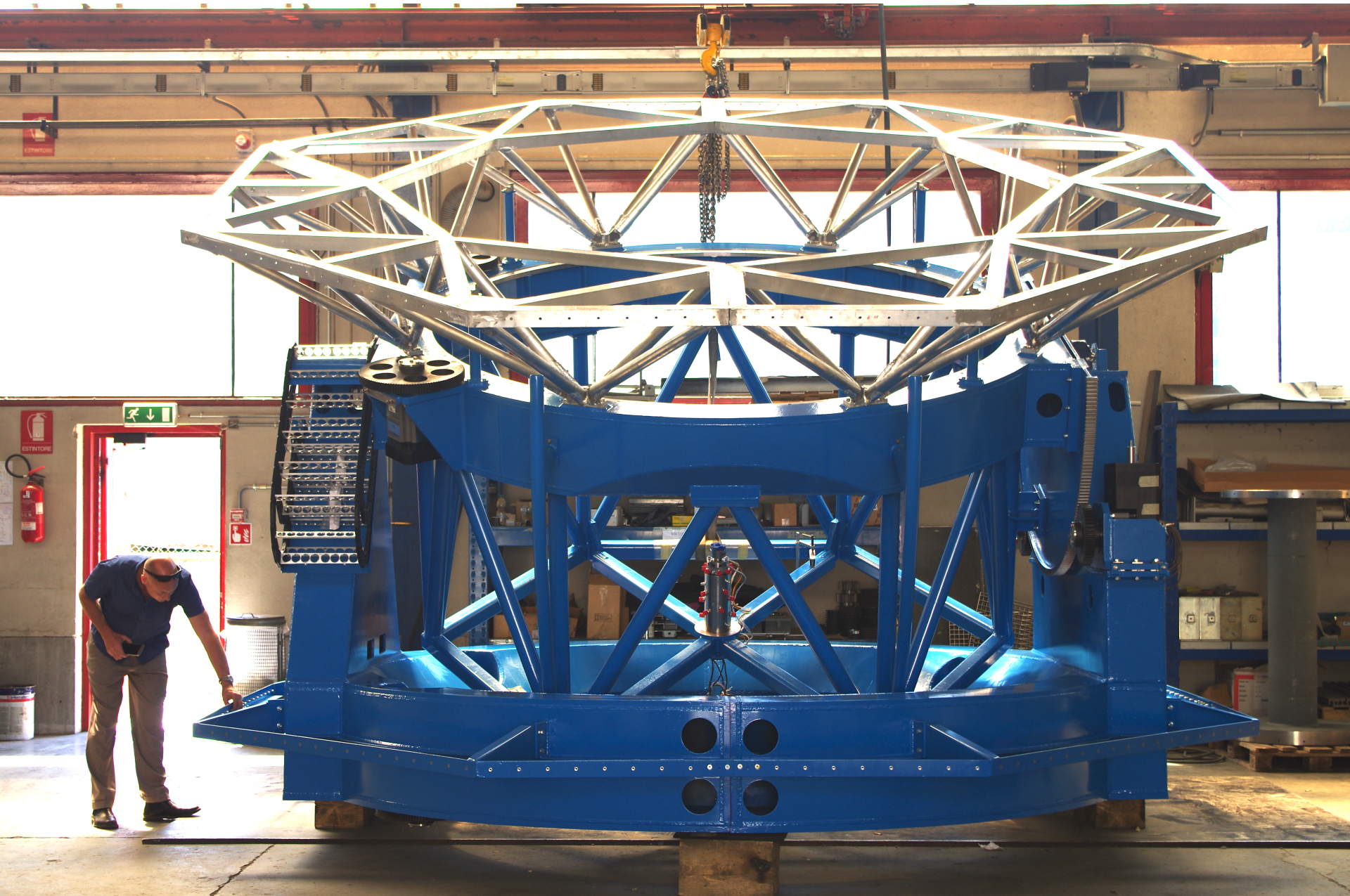


-10 -5 0 5 10
Degrees on sky

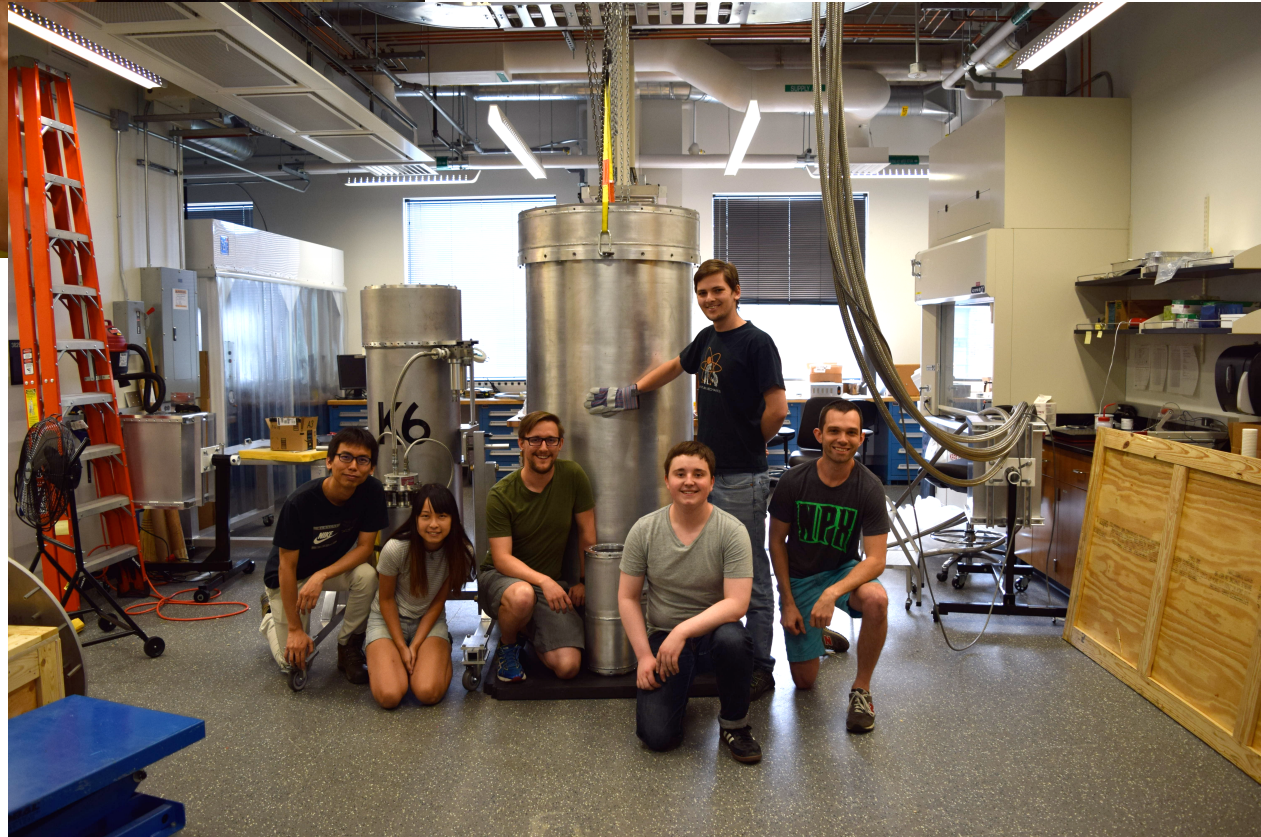
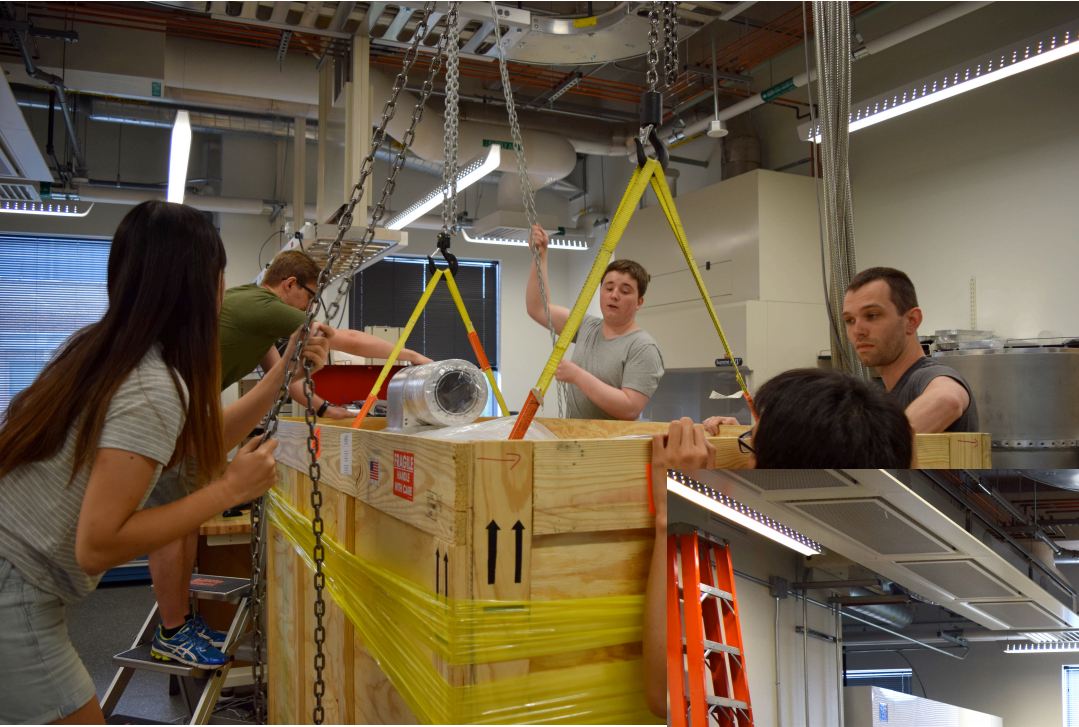
Next Gen Experiment BICEP Array Under Construction



BICEP Array Mount Completed

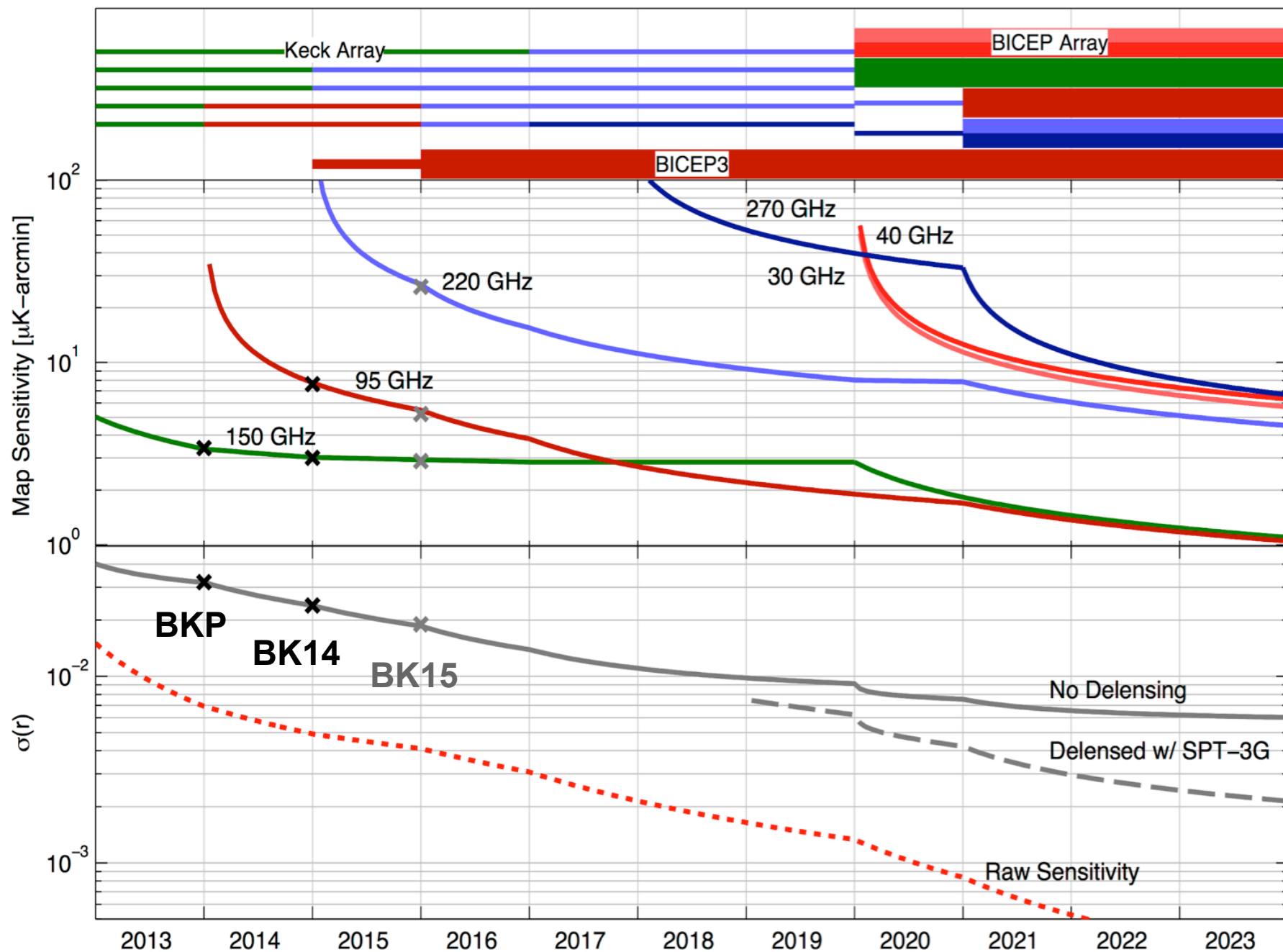


Cryostat Arrival at University of Minnesota



Stage 2

Stage 3



Conclusions

- By measuring CMB polarization we can look for primordial gravitational waves and either measure the energy scale of inflation, or rule out additional inflationary models
- BICEP/Keck leads the field in this endeavor:
- Published BK14 result sets $r_{0.05} < 0.09$ and $\sigma(r) = 0.024$
- (BK15 will shortly appear with $\sigma(r) = 0.020$)
- Galactic Foregrounds are a major issue but in principle can be overcome by multi-frequency measurements – we will only find out how well by trying
- BICEP Array is under construction and will go much further with next generation receivers in five bands
- Delensing in conjunction with SPT3G is under development
- BK23 result $\sigma(r) < 0.003$ by 2024
- We have competition from Simons Observatory who are building a BICEP Array like machine – and beyond that is merged mega experiment CMB-S4...