

# Search for Sterile Neutrino at NEOS Experiment

2017.07.21

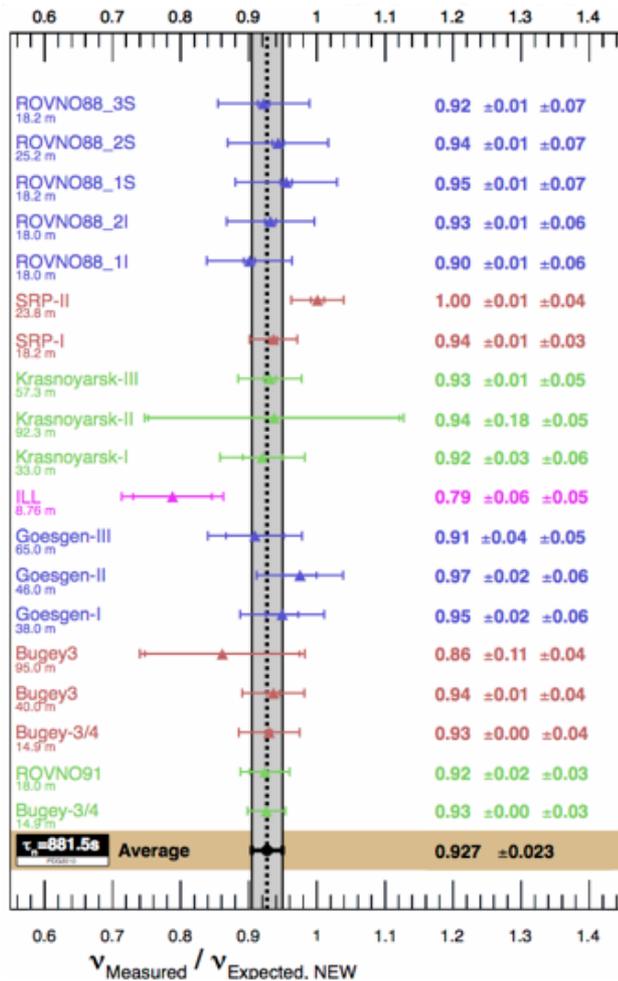
13<sup>th</sup> Rencontres du Vietnam

Hyunsoo Kim  
Sejong University



on behalf of NEOS Collaboration

# Reactor Antineutrino Anomaly (RAA)

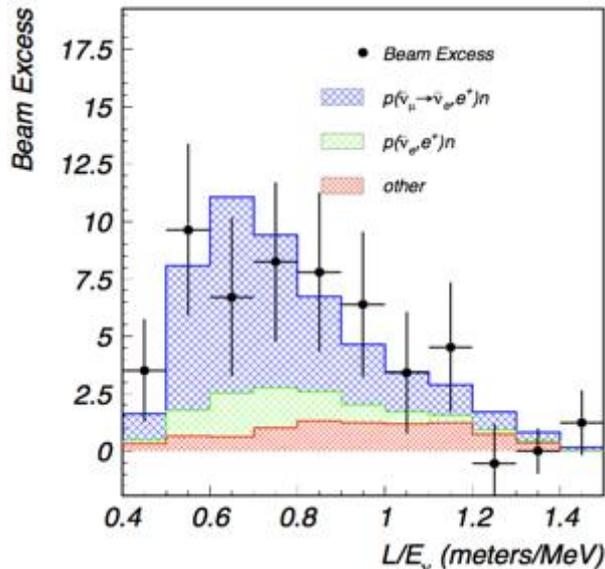


Short baseline reactor neutrino experiments

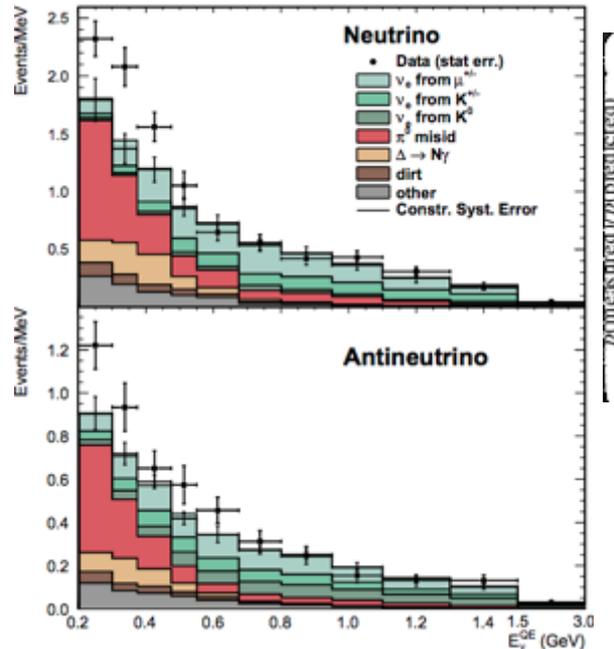
- Measured/Prediction= $0.927 \pm 0.023$

arXiv:1204.5379

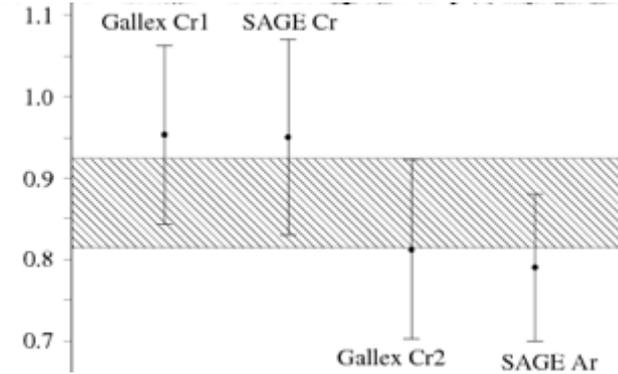
# Other Anomalies/Conflicts



PRD 64 (2001) 112007



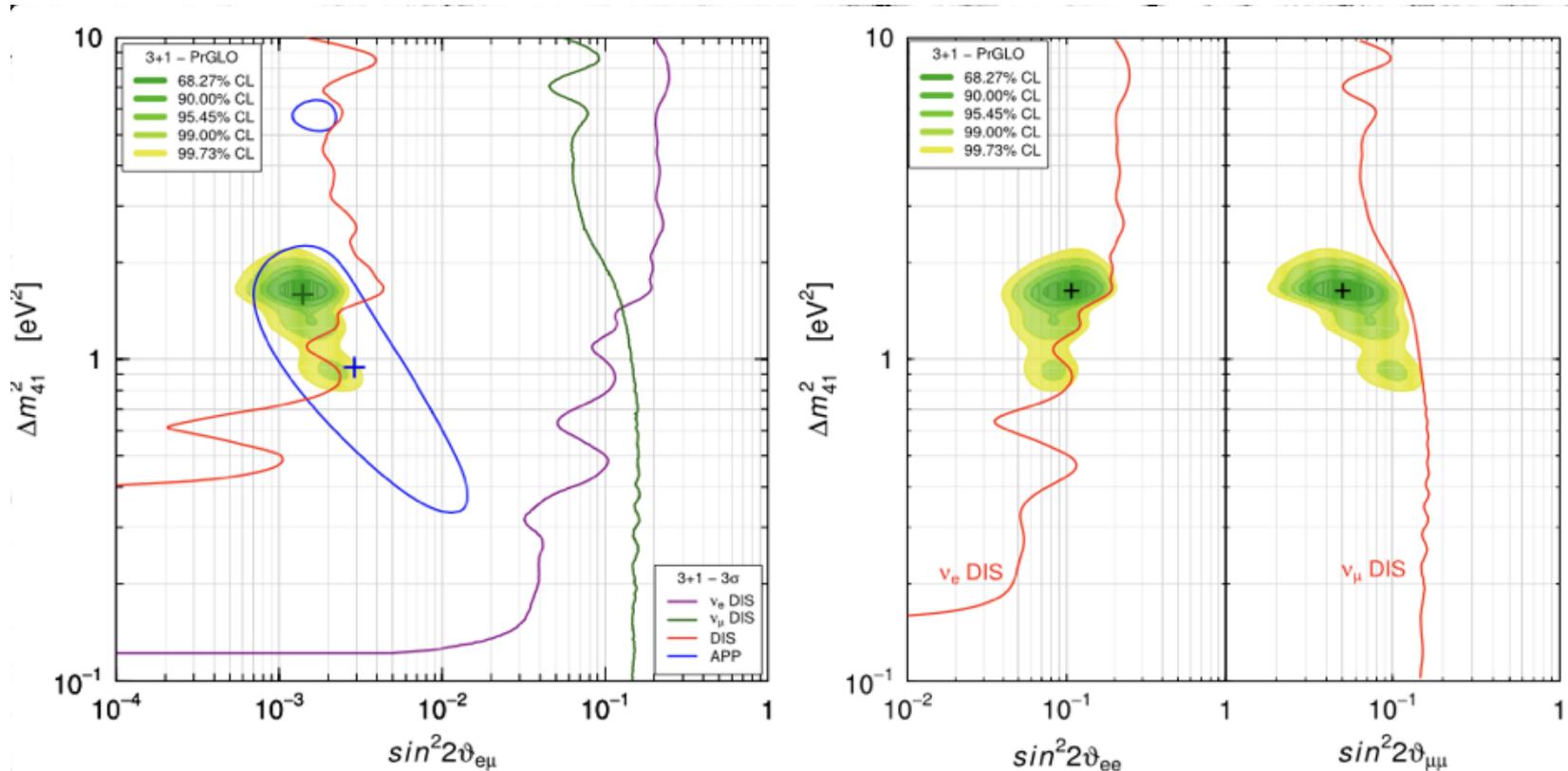
PRL 110 (2013) 161801



PRC 80 (2009) 015807

- LSND:  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance event excess with  $\Delta m^2 > 0.2 \text{ eV}^2$  ( $> 3\sigma$ )
- MiniBooNE:  $\nu$  mode disfavors /  $\bar{\nu}$  mode consistent with LSND result  $\Delta m^2 \sim 1 \text{ eV}^2$
- GALLEX / SAGE:  $2.9\sigma$  deficit from expected
- KARMEN, MINOS, IceCube, ... : negative results

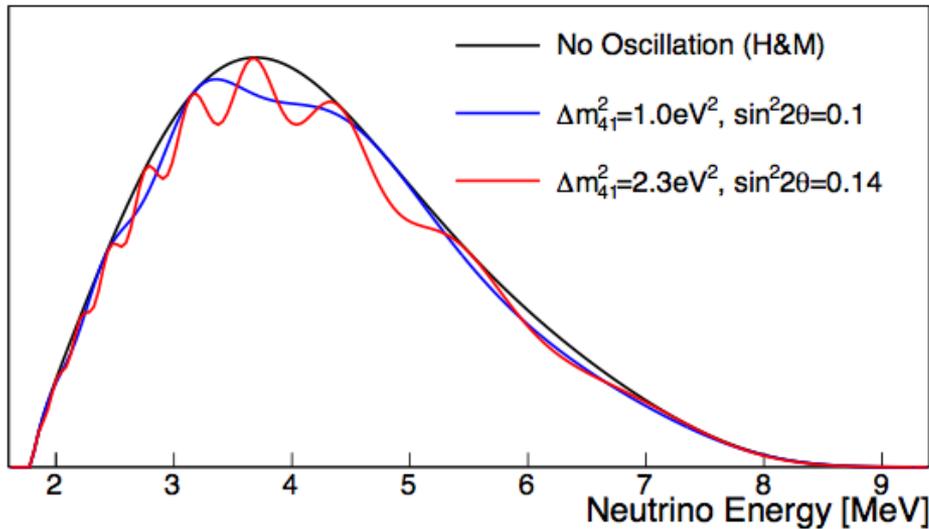
# Combined Results: (3+1)- $\nu$



Gariazzo *et al.*, JPG 43 (2016) 033001

All seem to point toward  $\sim 1$  eV sterile neutrino

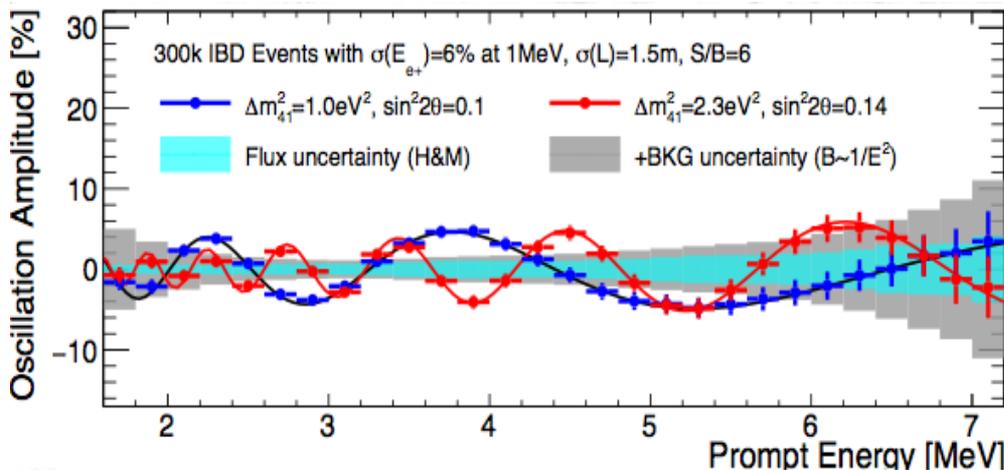
# (3+1) $\nu$ and SBL Oscillation



$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\theta_{14} \sin^2 \left( 1.27 \frac{\Delta m^2 L}{E_\nu} \left[ \frac{\text{eV}^2 \cdot \text{m}}{\text{MeV}} \right] \right)$$

Oscillation pattern washes out in the measured prompt energy spectrum

- Baseline (L): reactor/detector size
- Energy ( $E_\nu$ ): Energy loss/resolution
- Systematic errors
- Backgrounds



# NEOS Collaboration

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- **N**eutrino **E**xperiment for **O**scillation at **S**hort baseline
- 20 Researchers from 8 institutes in Korea
  - Chonnam National University
  - Chung-Ang University
  - Institute for Basic Science (IBS)
  - Korea Research Institute of Standard and Science (KRISS)
  - Korea Atomic Energy Research Institutes (KAERI)
  - Kyungpook National University
  - Sejong University
  - University of Science and Technology



- Project launched in 2012
- Data taking from Aug. 2015 to May 2016

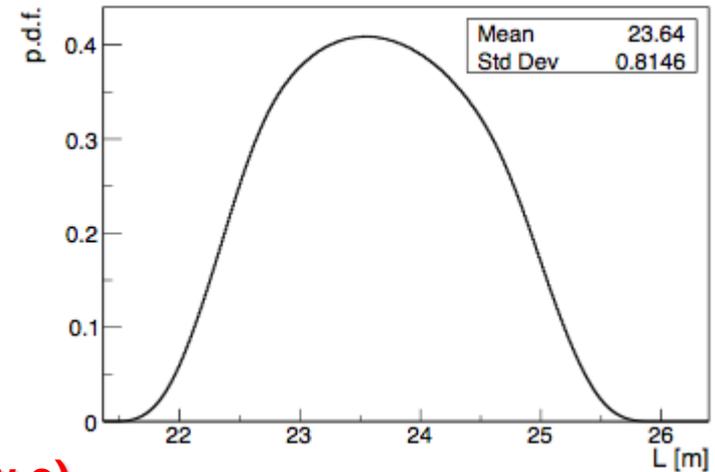
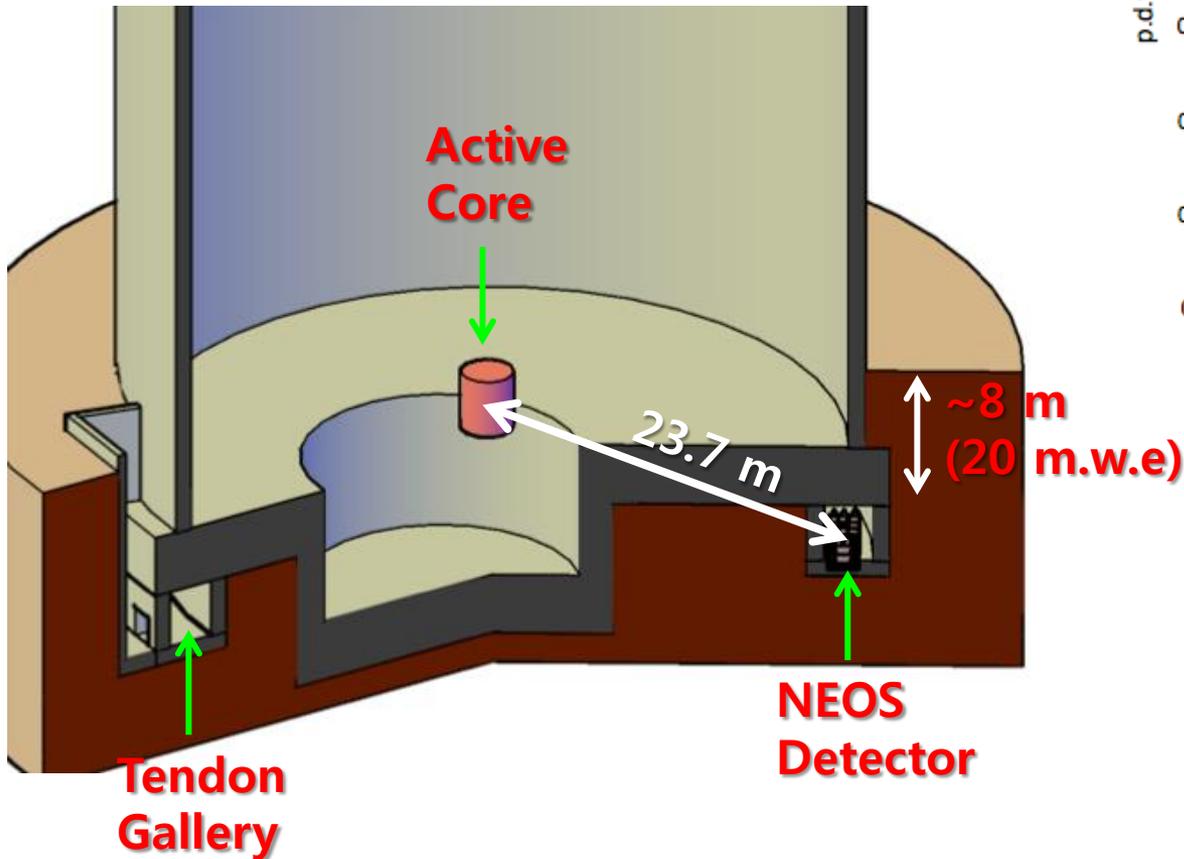
# Experimental Site

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- Reactor Unit 5 at Hanbit Nuclear Power Plant in Yeonggwang, Korea
- Same NPP for RENO Experiment.

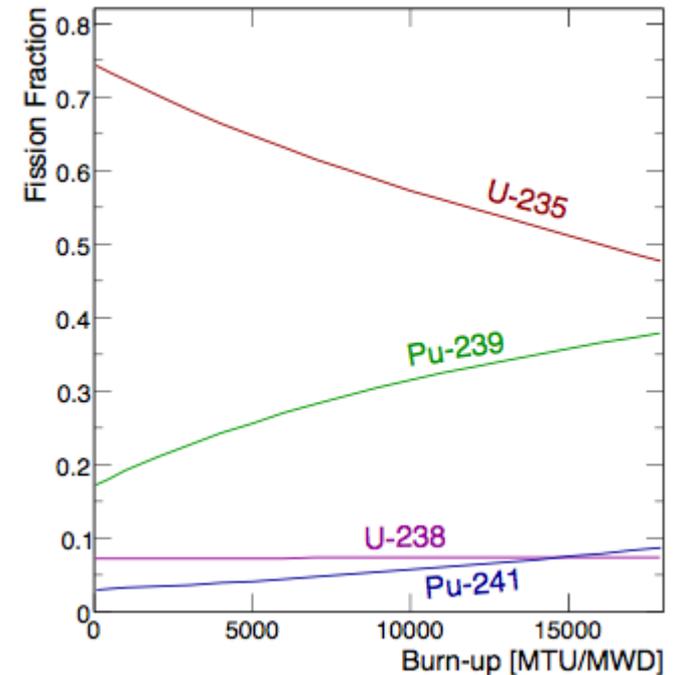
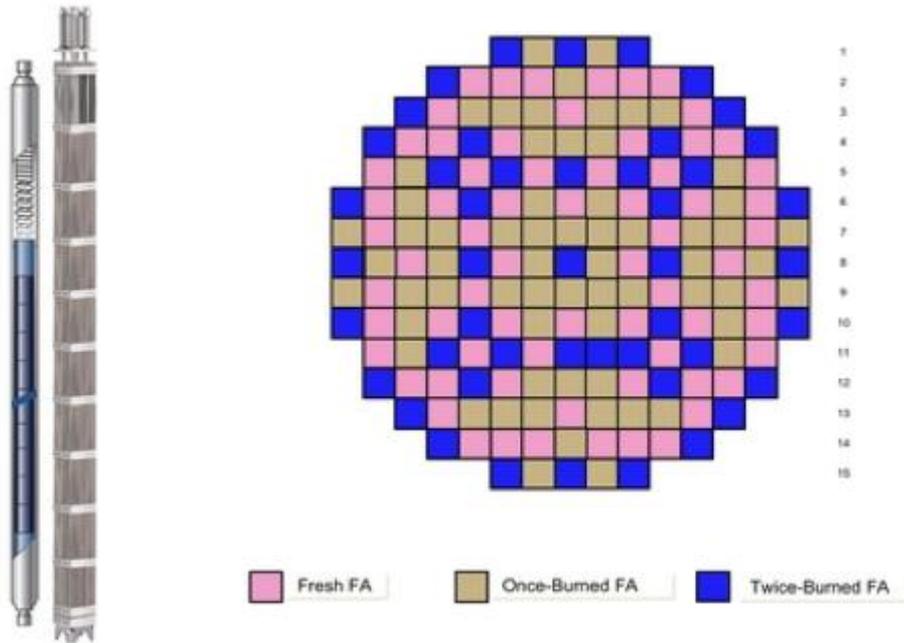
# Detector Location



Neutrino flight distance profile

- Neutrinos uniformly generated in cylindrical core
- IBD interaction  $\propto 1/L^2$

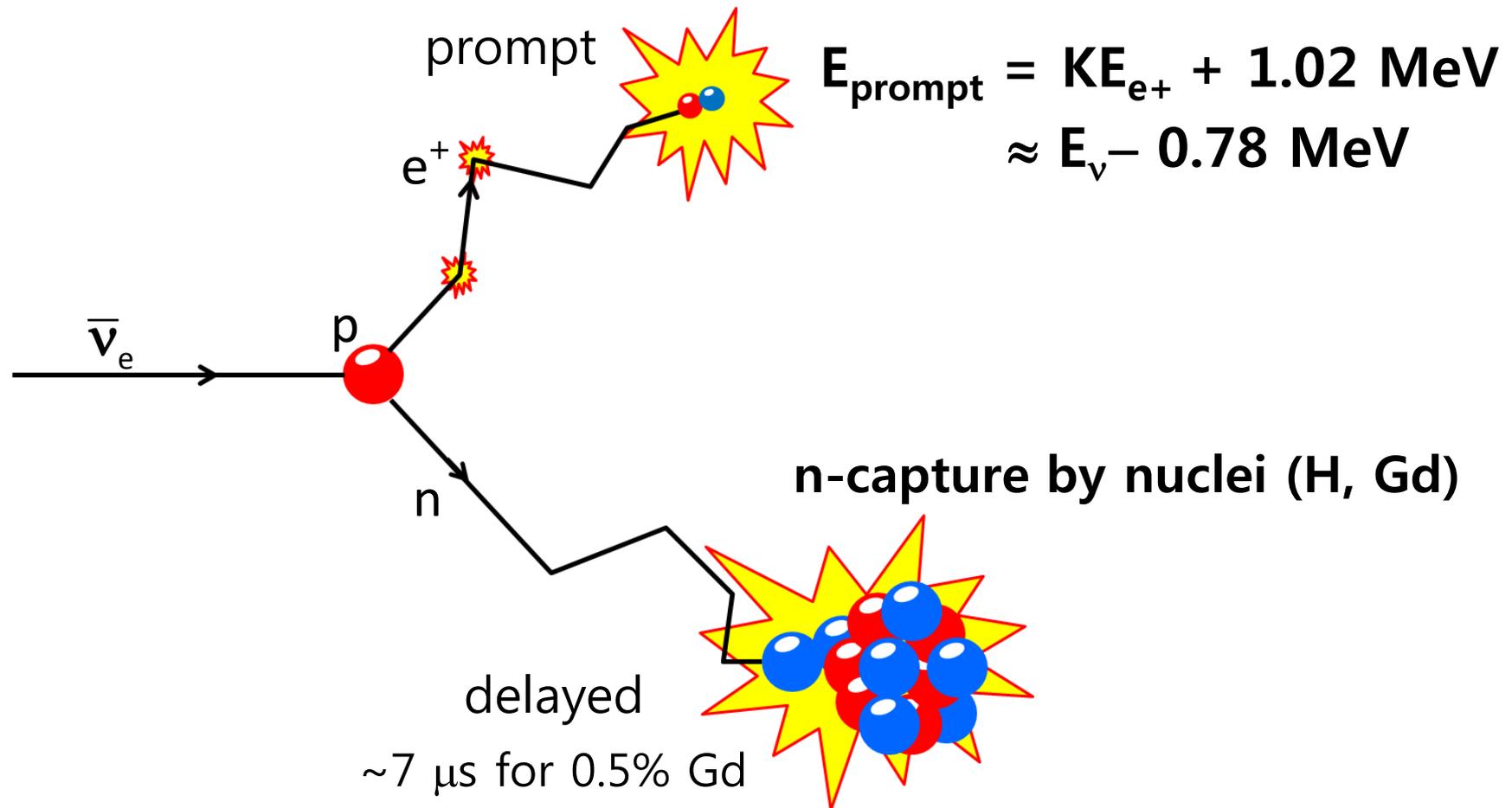
# Neutrino Source



- Thermal Output: 2.8 GW<sub>th</sub>
- Low enriched (4.65%) Uranium-235
- Refueling by swapping 1/3 of fuel rods for each burn-up cycle (~1.5 years)
- Active core size: 3.1 m (diameter), 3.8 m (height)

# Neutrino Detection

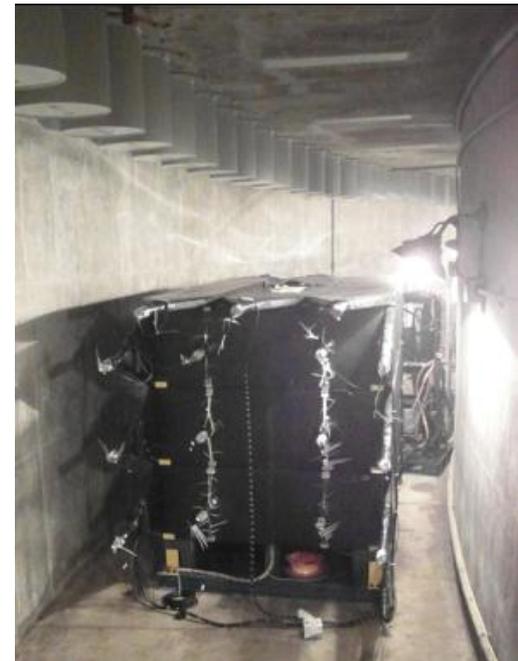
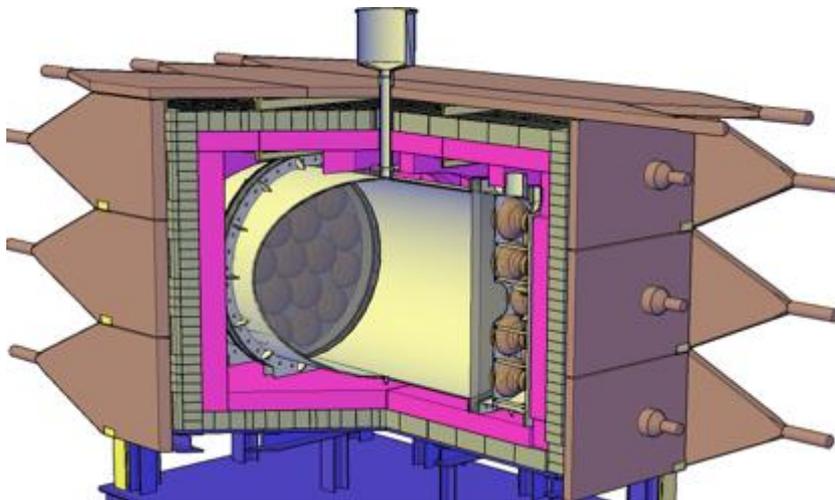
**Inverse Beta Decay**  $\bar{\nu}_e + p \rightarrow e^+ + n$



# NEOS Detector

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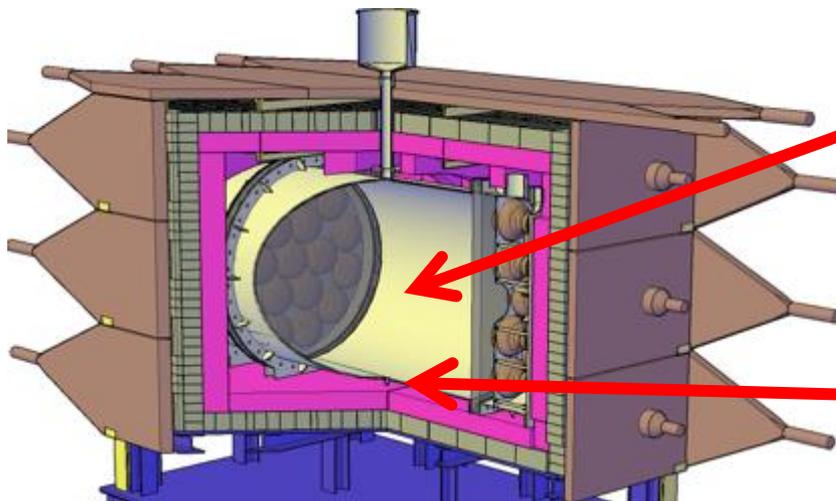
- Detector size determined by the available space in the tendon gallery [4 m(h) x 3 m(w)]
- Simple construction
  - Homogeneous liquid scintillator target with no gamma catcher and limited buffer coverage.



# NEOS Detector

## Homogeneous Liquid Scintillator Target

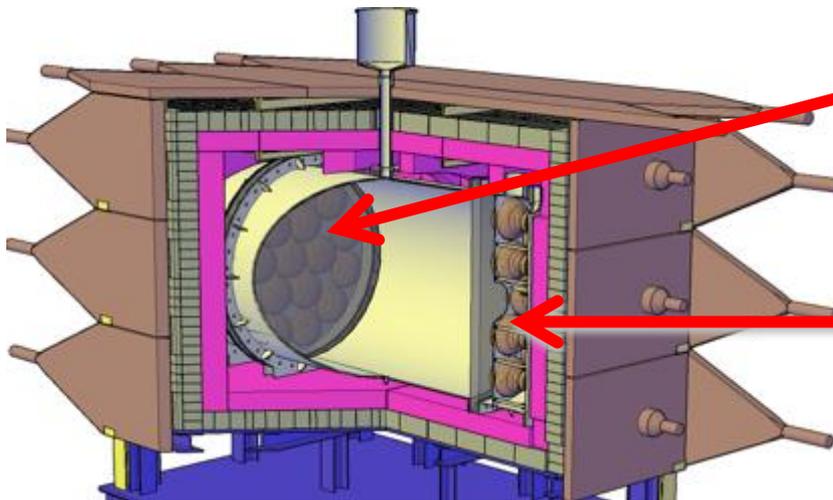
- ~1000 L in volume [1.2 m (L)/0.5 m (R)]
- LAB+DIN based LS (Ultima Gold-F) (9:1)
  - improve pulse shape discrimination over LAB only
- 0.5% Gd for high neutron capture efficiency
- Target vessel lined with teflon



# NEOS Detector

## PMTs

- 19 8-inch PMTs on each side of target
- PMTs are installed in buffer tanks filled with mineral oil
- Separated from target by a 6 cm thick transparent acrylic window.



# NEOS Detector

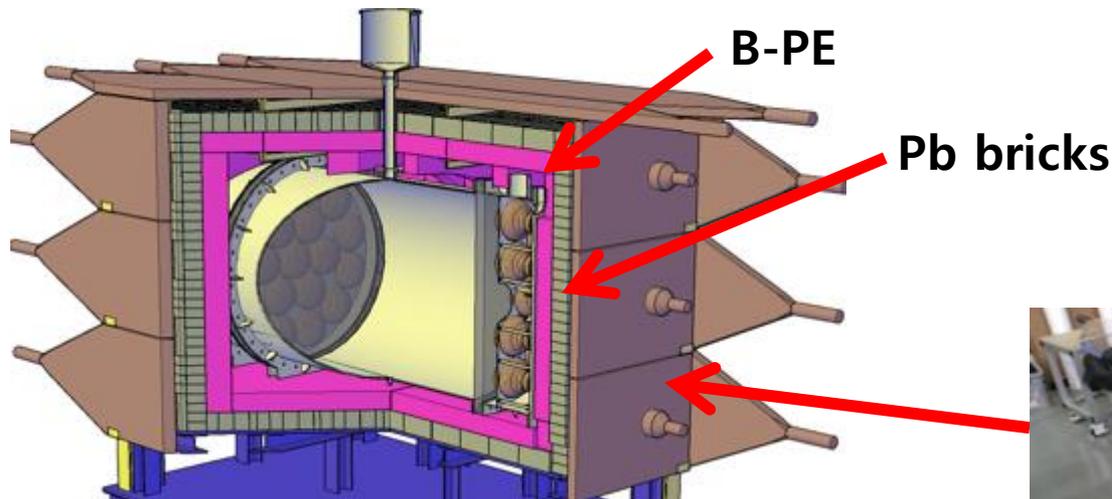
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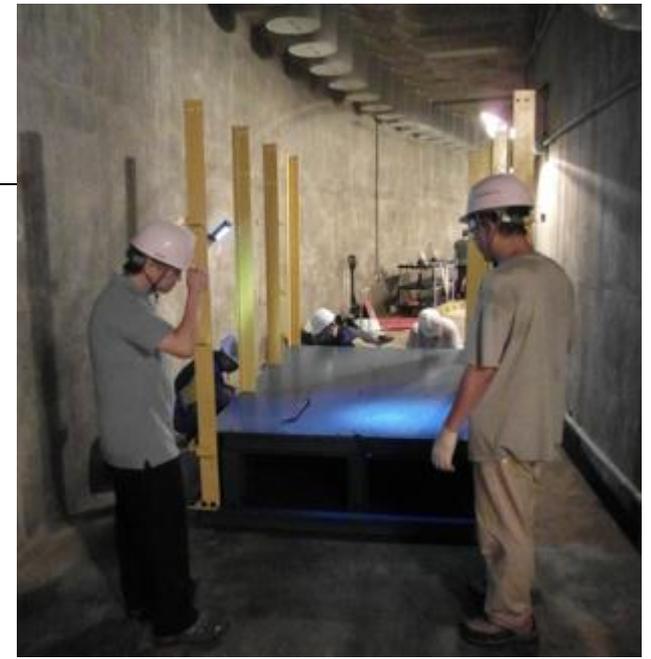
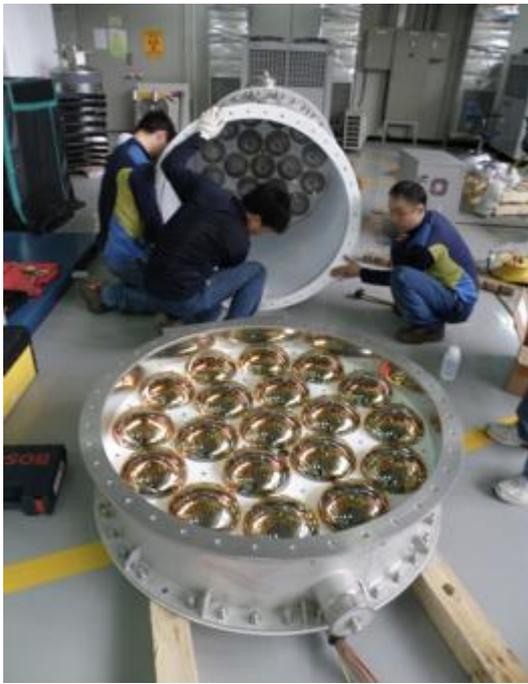
## Shieldings

- 10 cm thick borated polyethylene (B-PE) for neutrons
- 10 cm thick Pb against external gammas

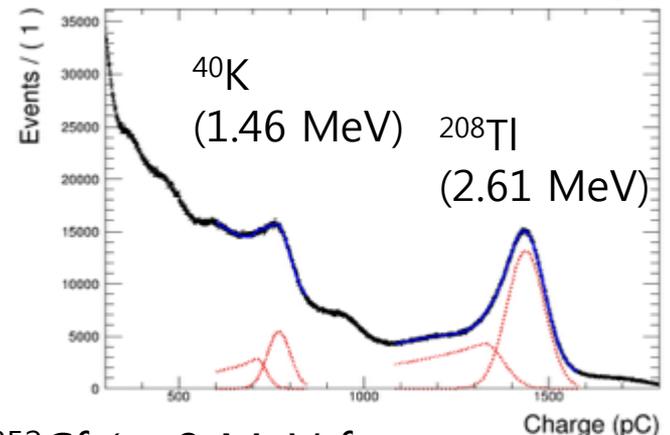
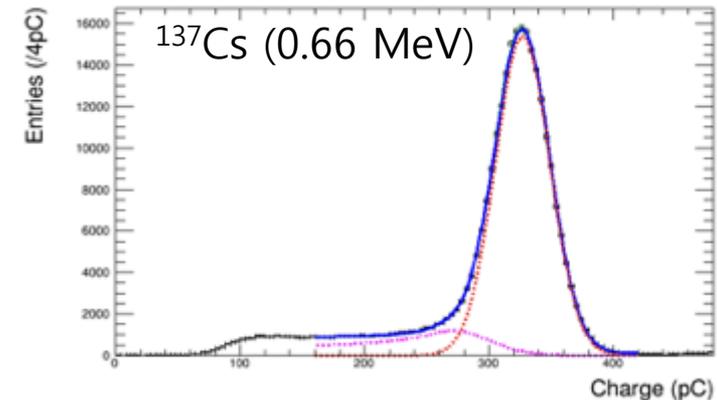
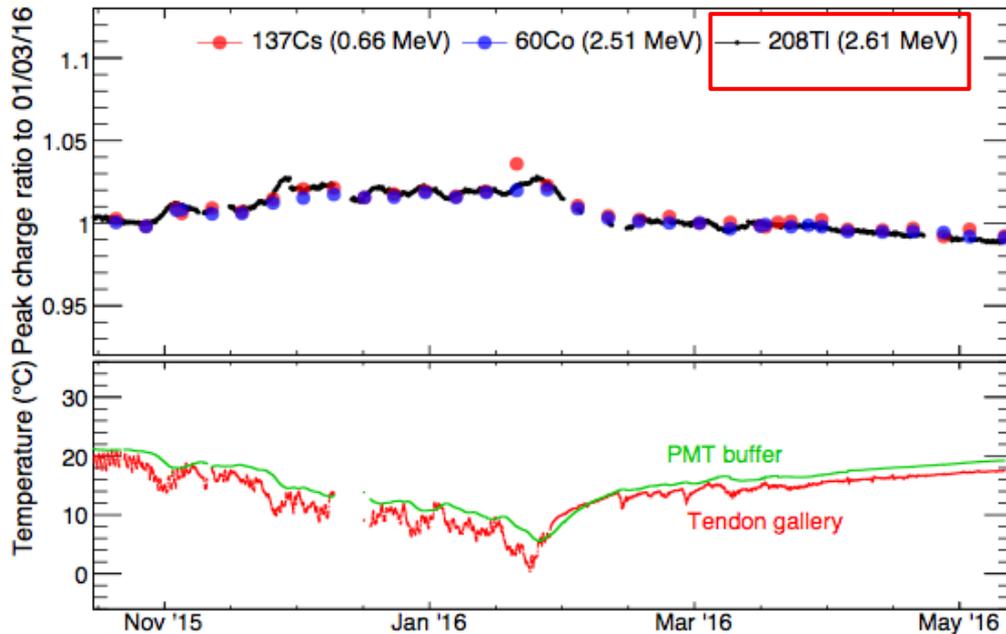
## Muon Veto

- Plastic scintillators covering top and sides.



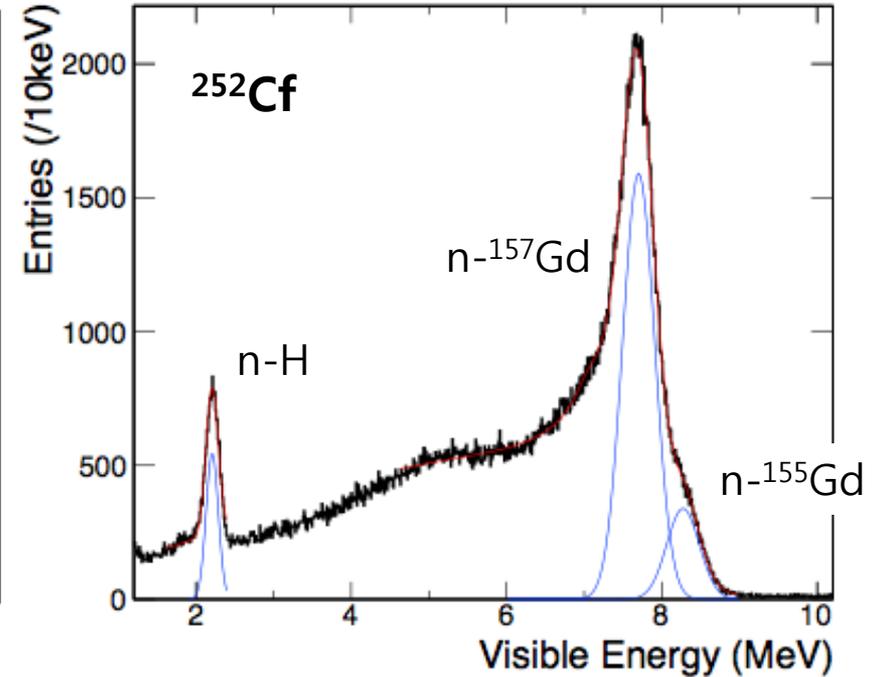
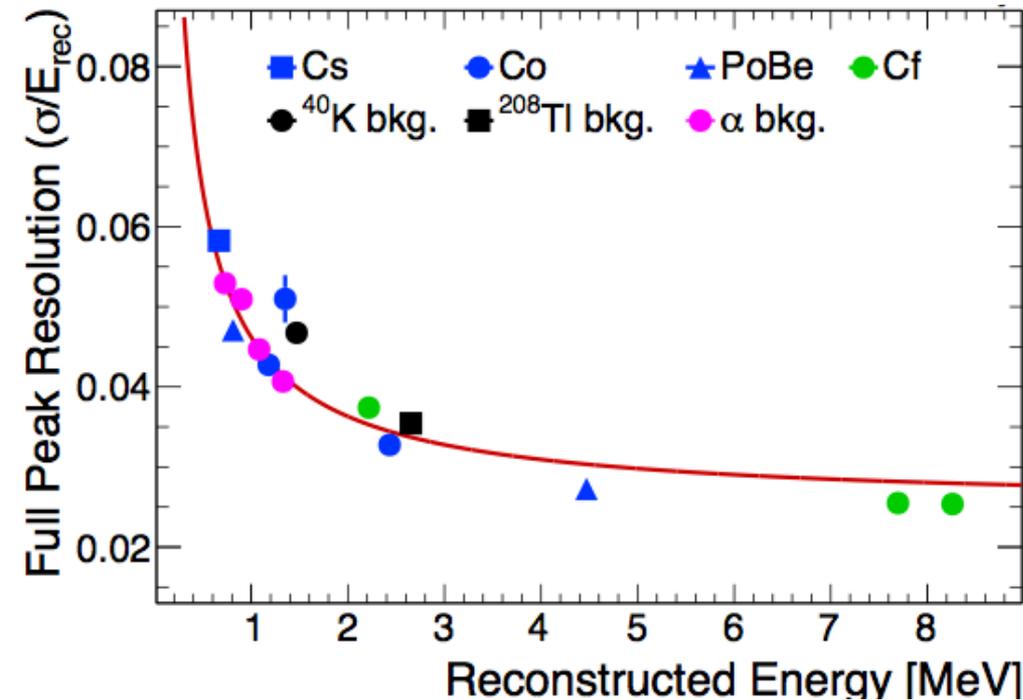


# Calibration



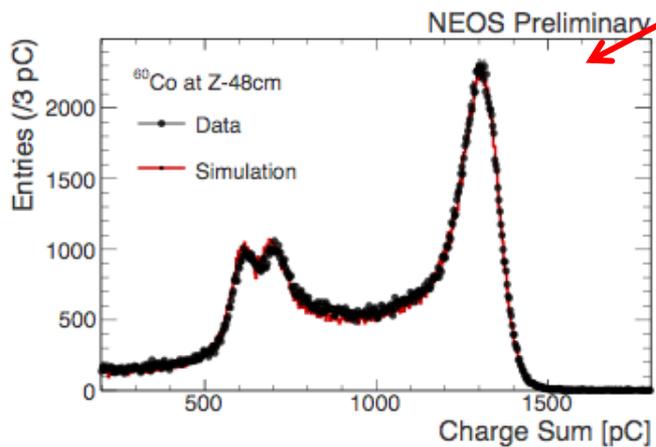
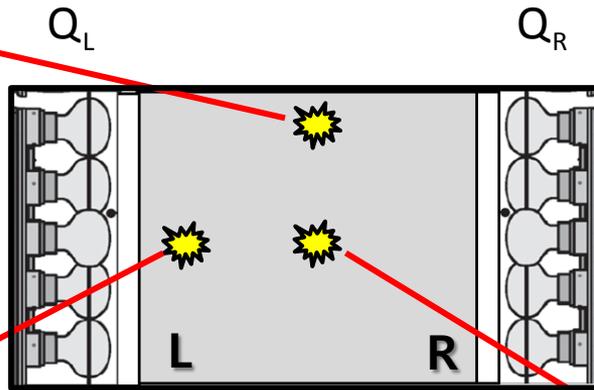
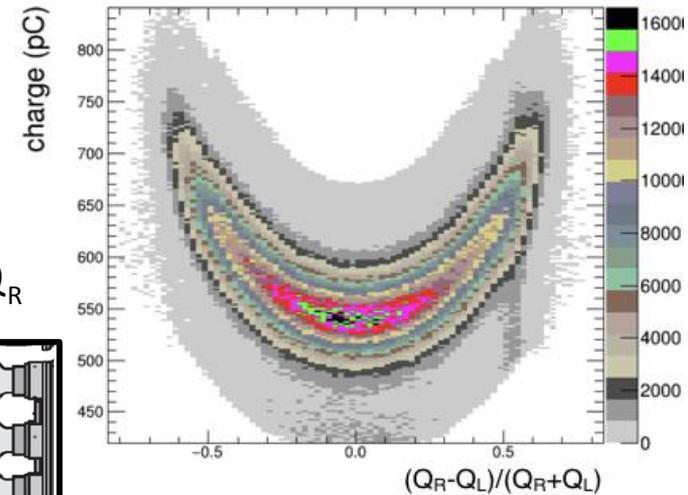
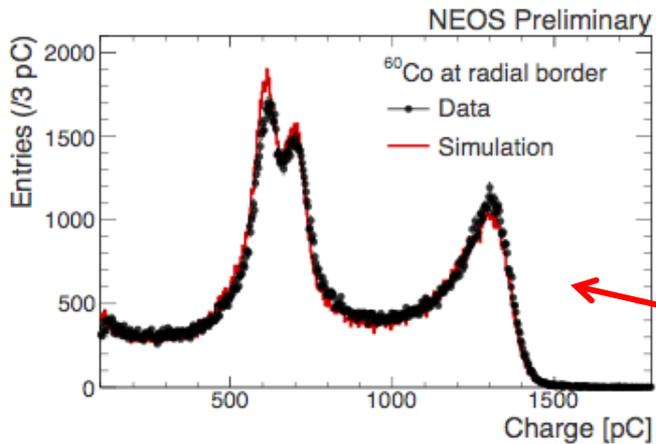
- Weekly calibration with point source
  - ✓  $^{137}\text{Cs}$  ( $\gamma$ :0.66 MeV),  $^{60}\text{Co}$  ( $\gamma$ : 1.17/1.33 MeV),  $^{252}\text{Cf}$  (n: 8 MeV from n-capture), PoBe (n,  $\gamma$ : 4.4 MeV)
- Internal/external backgrounds (continuous volume source)
  - ✓  $^{40}\text{K}$  ( $\gamma$ :1.46 MeV),  $^{208}\text{Tl}$  ( $\gamma$ :2.6 MeV),  $^{230}\text{Th}$  chain ( $\alpha$ ,  $\beta$ ,  $\gamma$ )

# Energy Resolution

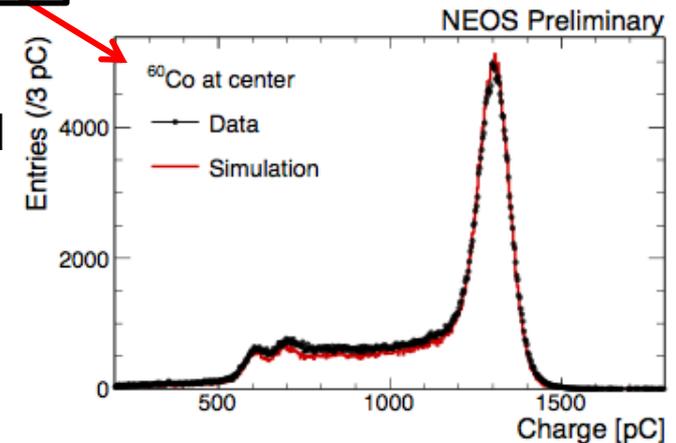


$\sigma/E=5\%$  for a full peak at 1 MeV

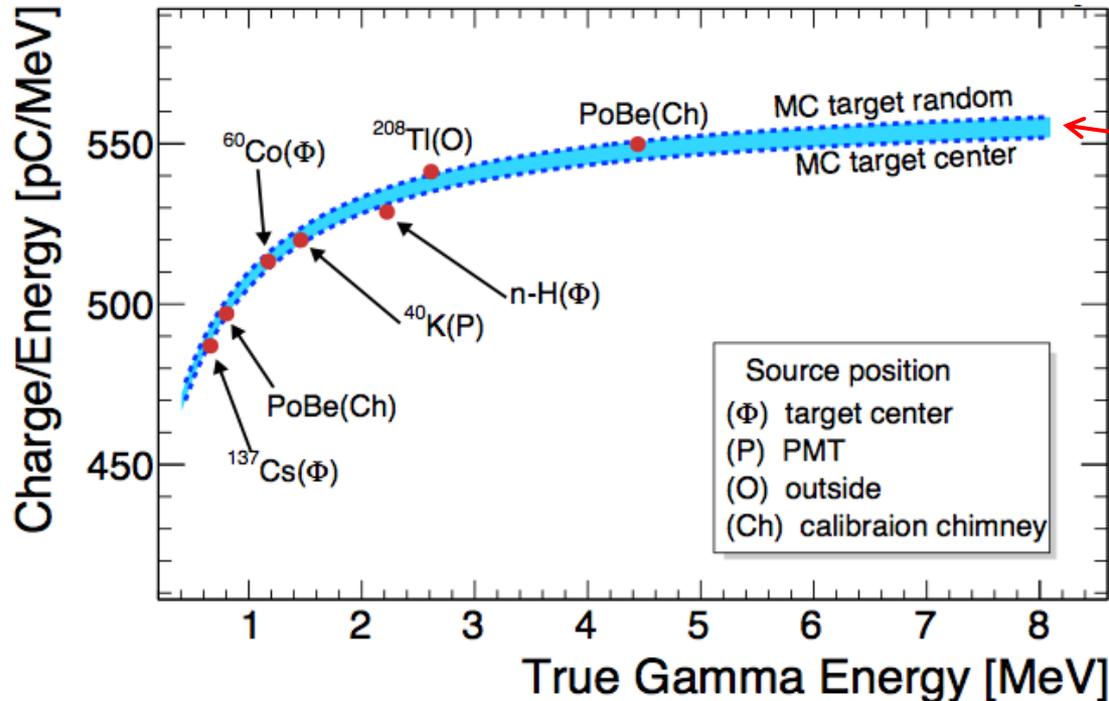
# Non-Uniform Response



Spatial correction applied



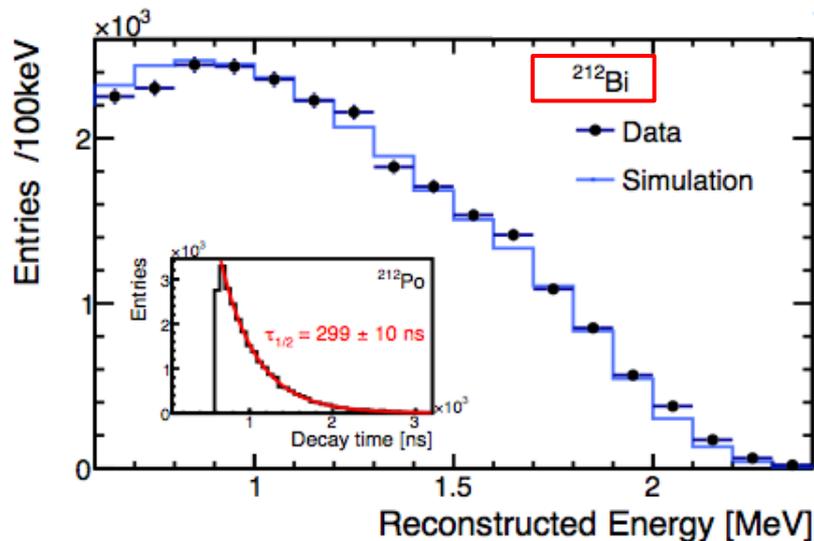
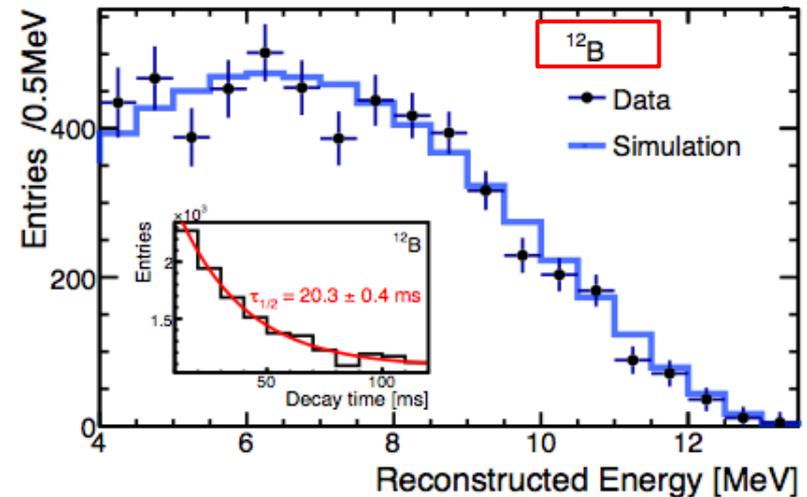
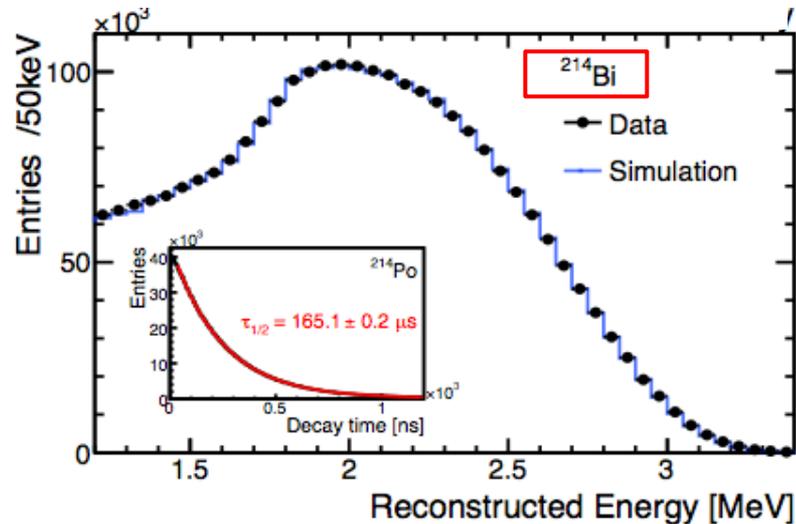
# Charge-to-Energy Conversion



$$\frac{Q}{E_\gamma} = (p_0 + p_1 E_\gamma)(1 + p_2 e^{-p_3 \sqrt{E_\gamma}})$$

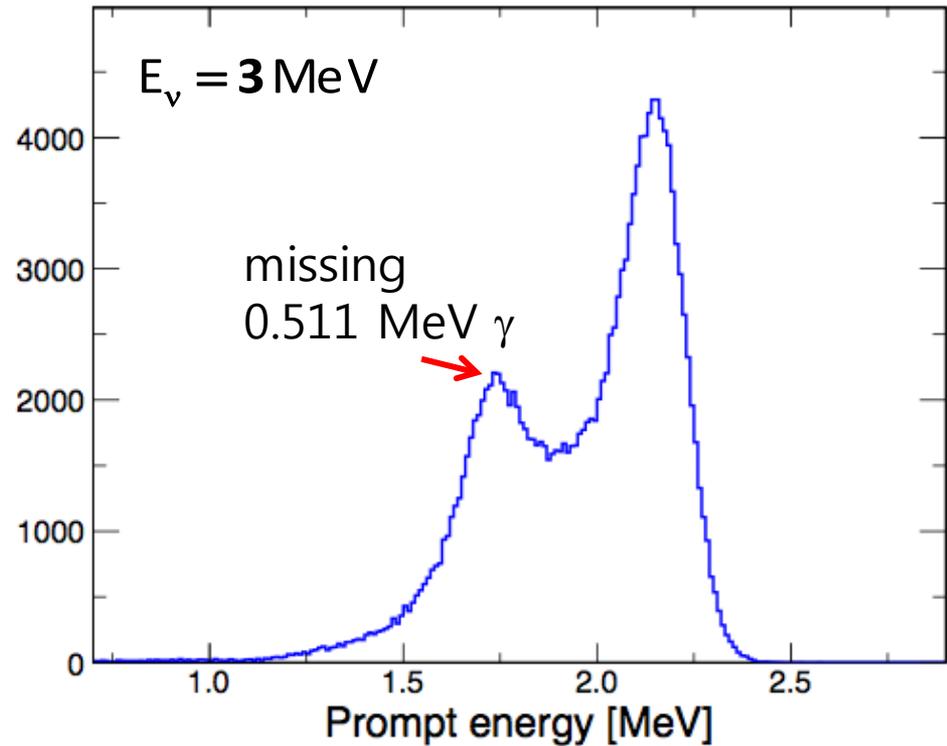
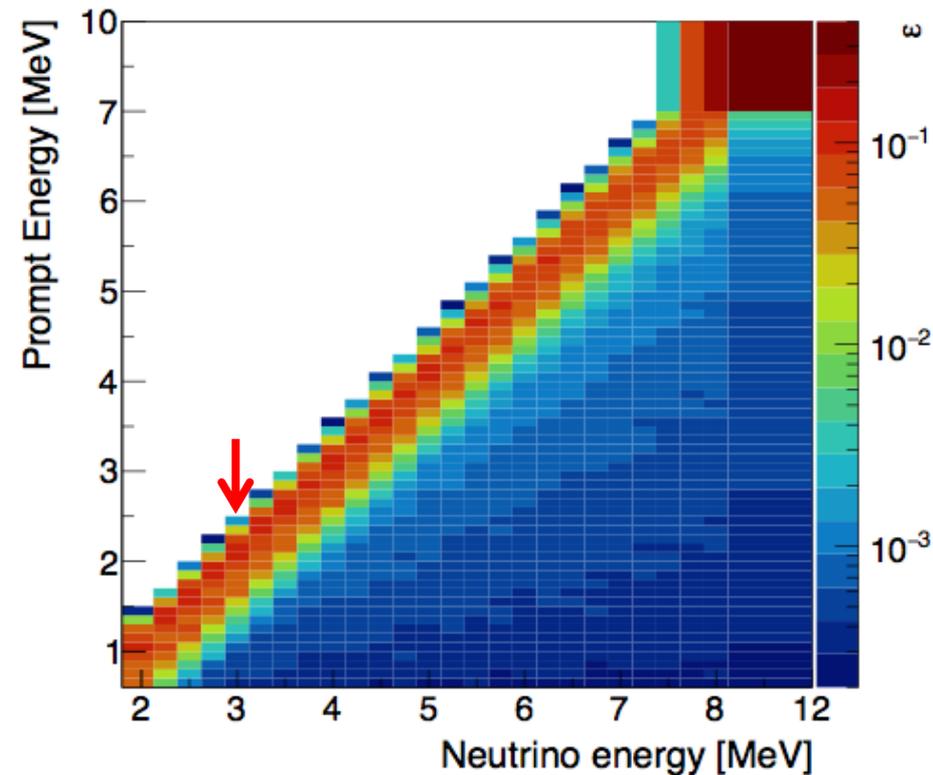
- Signal quenched at low energy
- Cerenkov light important at higher energy

# $\beta$ -decay Events and Simulation



- Geant4 based full MC simulation including trigger simulation
- Excellent agreement with data

# Neutrino Energy Response (MC)



Energy loss due to escaping  $\gamma$

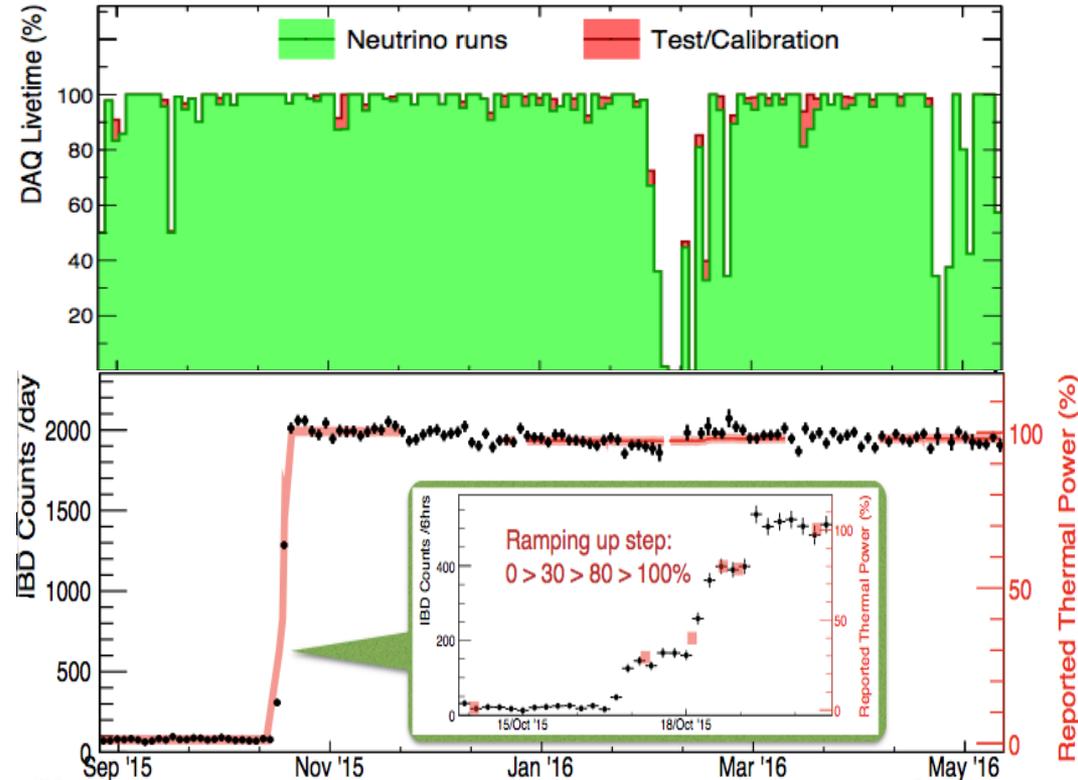
# IBD Candidate Event Rate

## Data taking period: Aug. 2015 - May 2016

- Data taking was terminated due to a scheduled maintenance of the tendon gallery (once every 5 years)
- Reactor-on: 180 days
- Reactor-off: 46 days
- DAQ over 90% efficient

## IBD Candidates

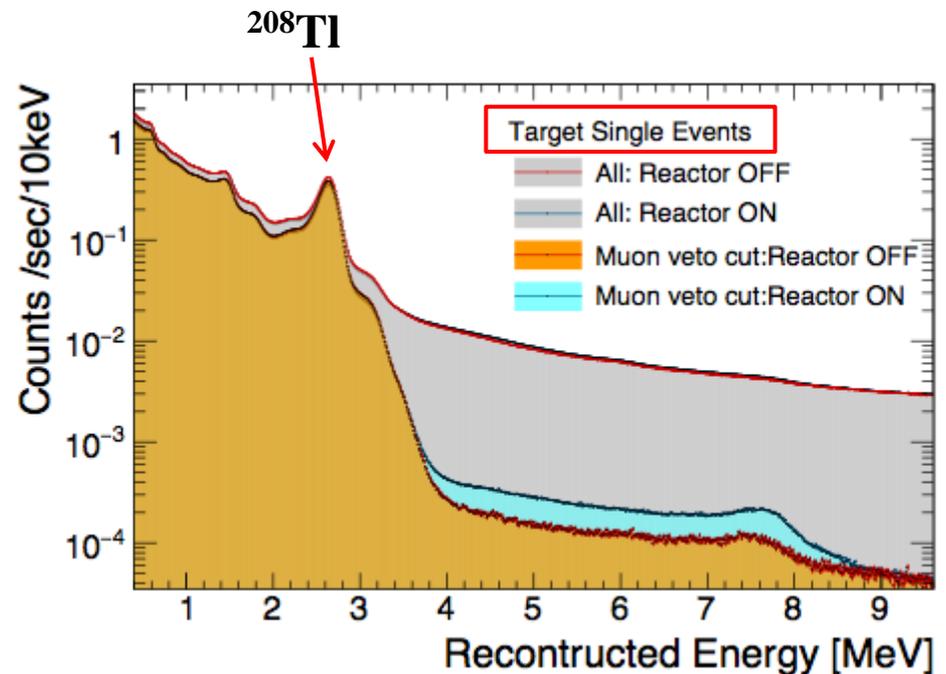
- Reactor-on: 2000/day
- Reactor-off: 80/day



# Single Event and Muon Veto

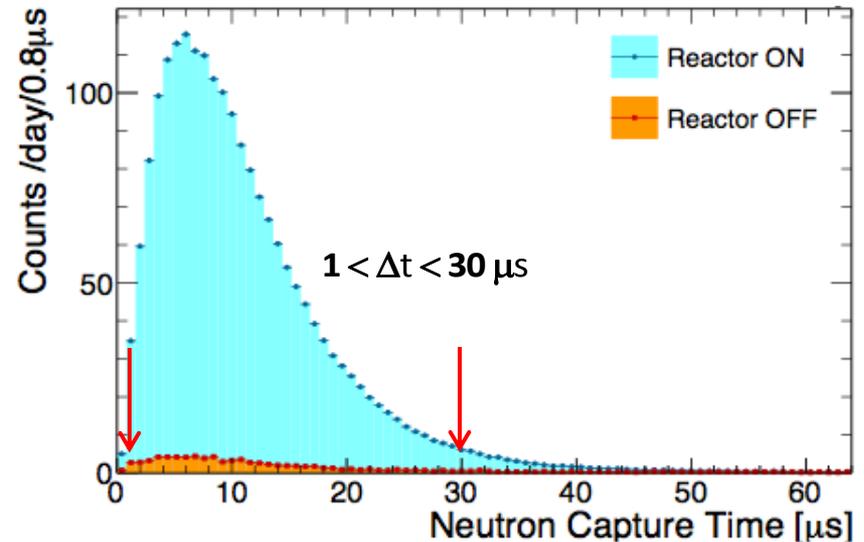
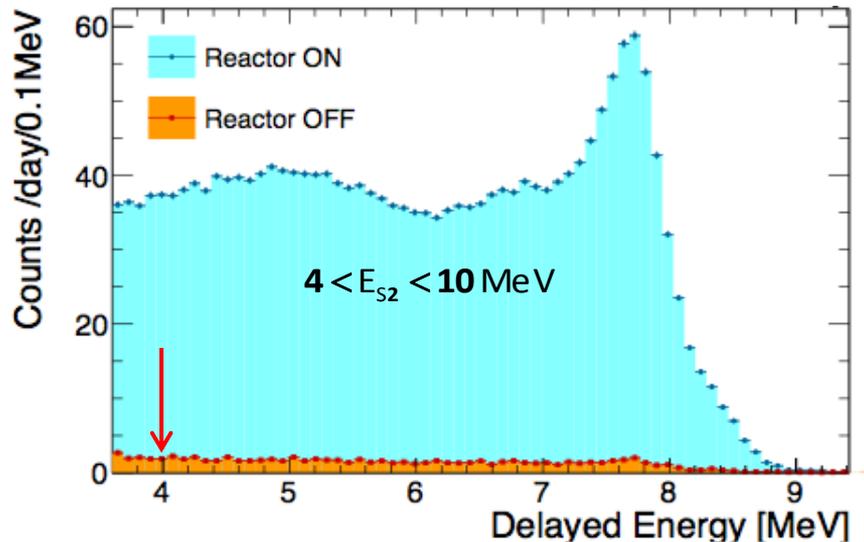
## Events before prompt-delayed event pairing

- Muon veto: discards any events within 150  $\mu\text{s}$  after a muon event.
- a peak at 2.6 MeV due to  $^{208}\text{Tl}$
- Reactor-on and off difference after veto is due to IBD pairs



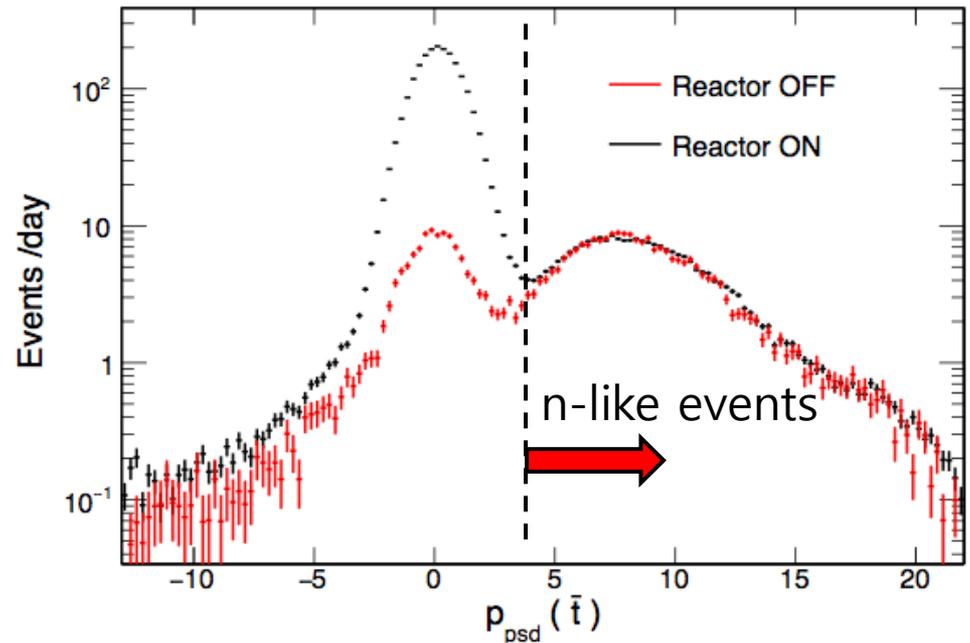
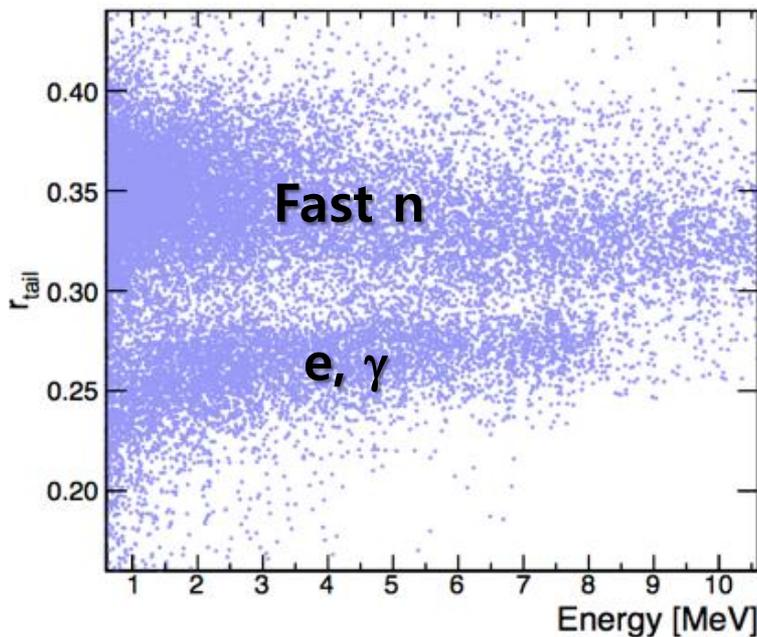
# IBD Reconstruction

- Prompt event:  $1 < E < 10$  MeV
- Delayed event:  $4 < E < 10$  MeV
- Delayed coincidence:  $1 < \Delta t < 30$   $\mu$ s
- Multiplicity veto: no event before 30, after 150  $\mu$ s from the prompt/delayed pair.

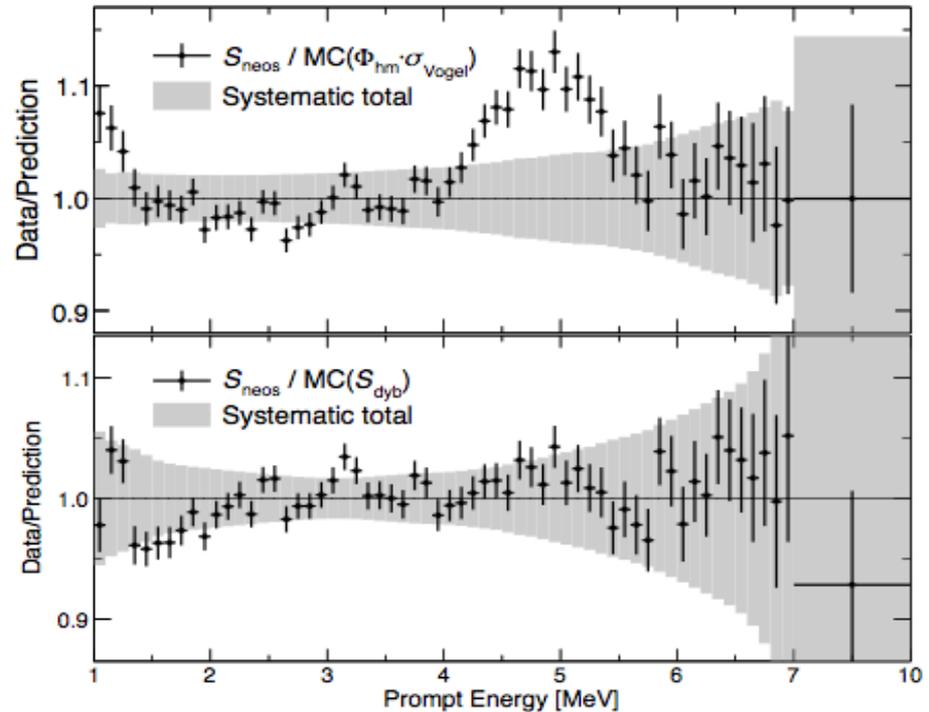
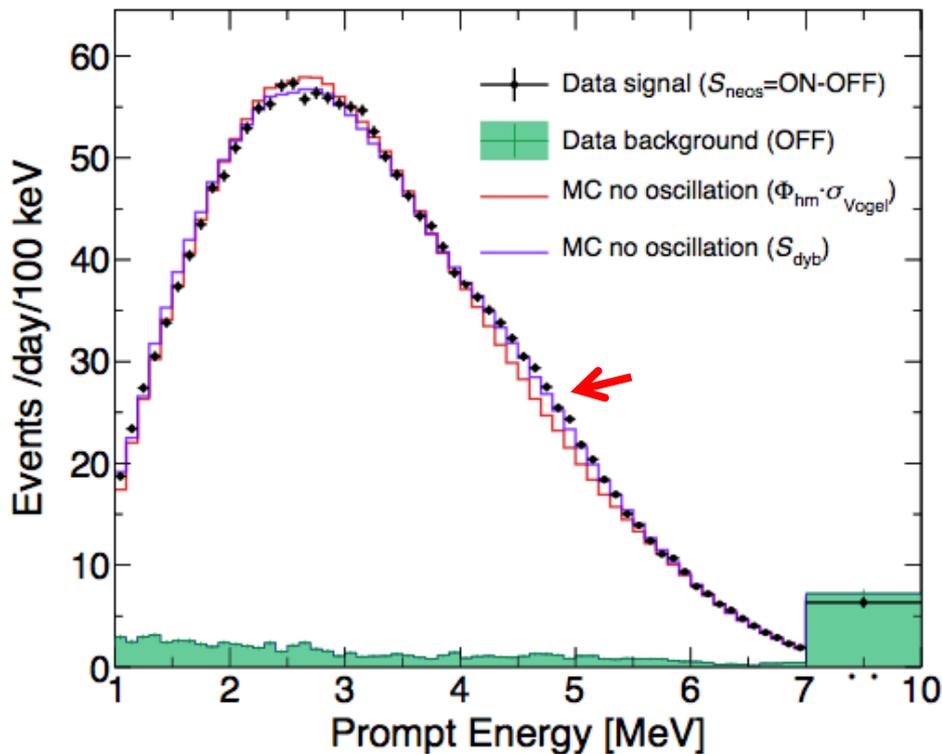


# Pulse Shape Discrimination

- Fast neutron scattering events have relatively slowly decaying pulse.
- Background is reduced by more than 70%.
- $\gamma$ -like event acceptance is 99.9%.



# Prompt Energy Spectrum



- S/N  $\sim 23$
- vs H-M: 5 MeV bump, not so practical to do the oscillation analysis
- vs Daya Bay: bump not totally disappeared, fission fraction difference
- Systematic uncertainties for the shape analysis:  
reference spectrum > energy recon. >> backgrounds, etc.

# $\chi^2$ Construction

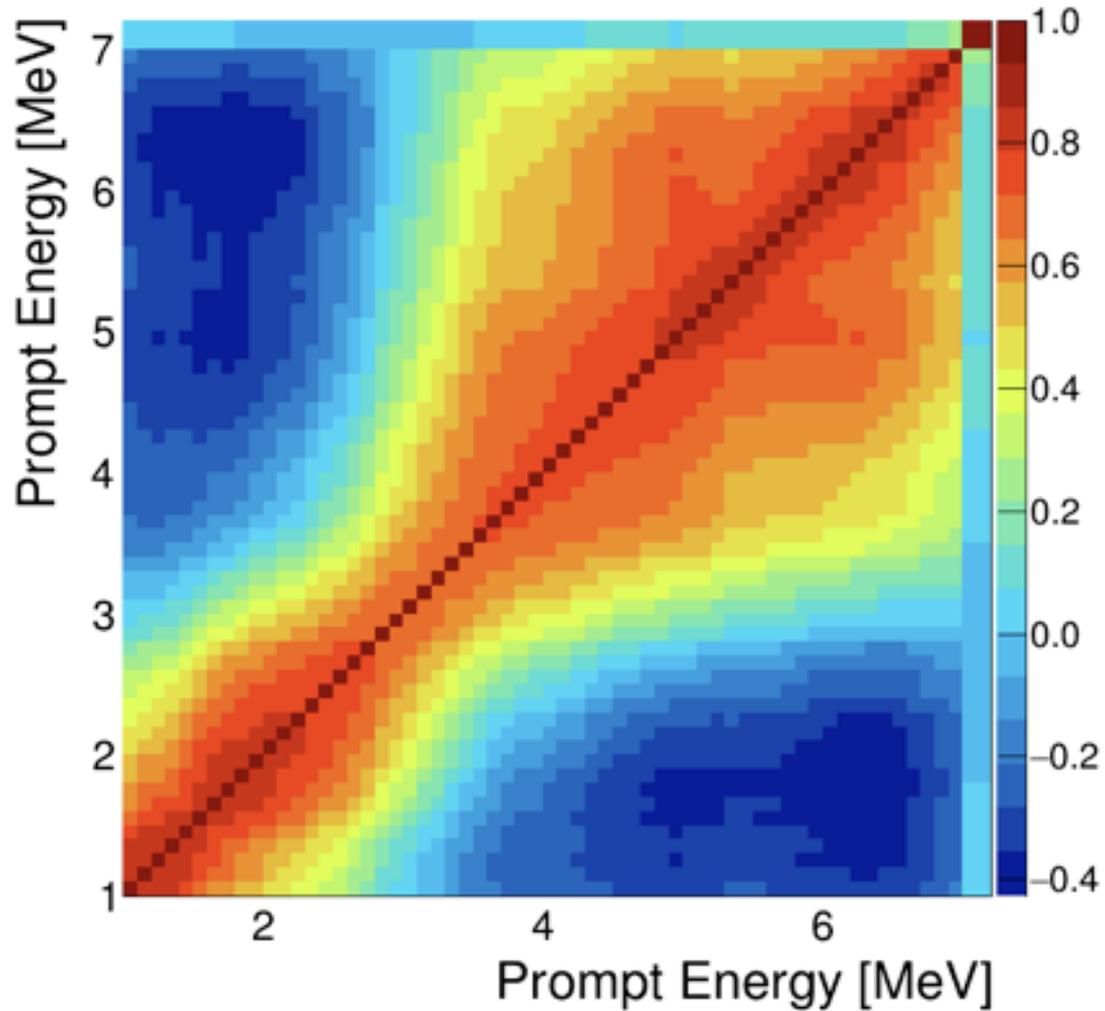
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$$\chi^2 = (N_i^{\text{on}} - (t_{\text{on}}/t_{\text{off}})N_i^{\text{off}} - T_i)V_{ij}^{-1}(N_j^{\text{on}} - (t_{\text{on}}/t_{\text{off}})N_j^{\text{off}} - T_j)$$

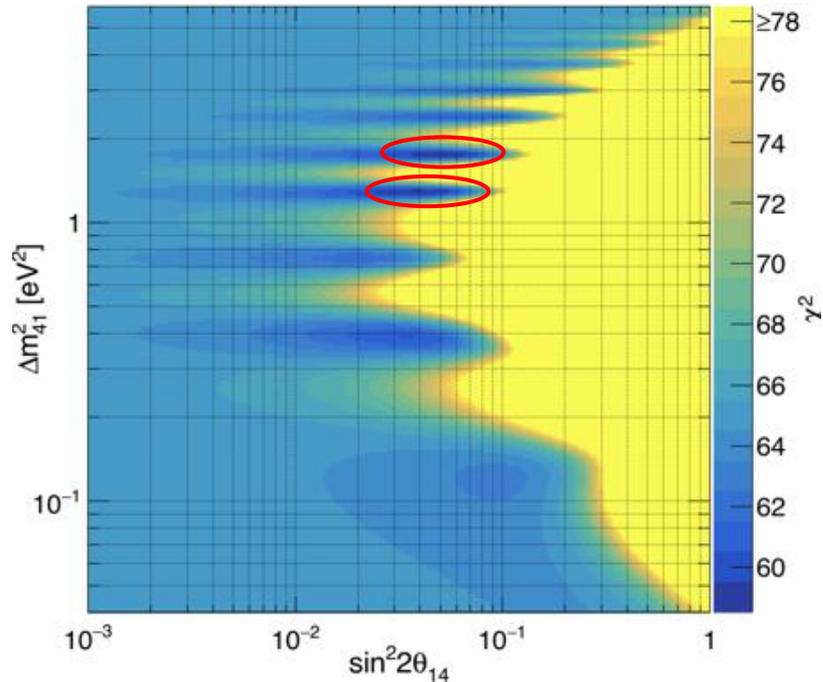
- $N_i^{\text{on}}$ : Number of events in the  $i^{\text{th}}$  energy bin in reactor on period
- $N_i^{\text{off}}$ : Number of events in the  $i^{\text{th}}$  bin in reactor-off period
- $t_{\text{on}}/t_{\text{off}}$ : time period of reactor on/off
- $T_i = T_i(\Delta m^2, \sin^2 2\theta)$ : expected spectrum for each oscillation parameter set generated from model spectrum and convoluted with the detector response (Daya Bay spectrum used)
- $V_{ij}$ : covariance matrix including statistical/systematic uncertainties and their correlation between energy bins

# Covariance Matrix

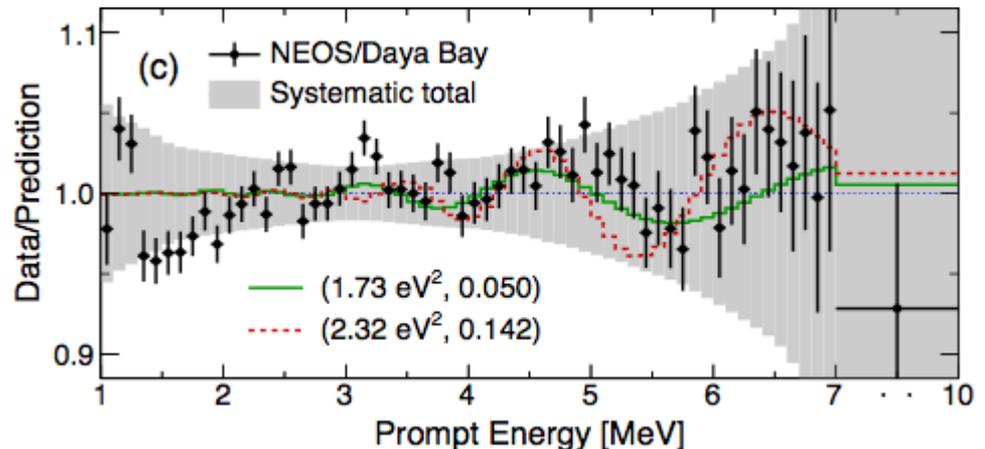
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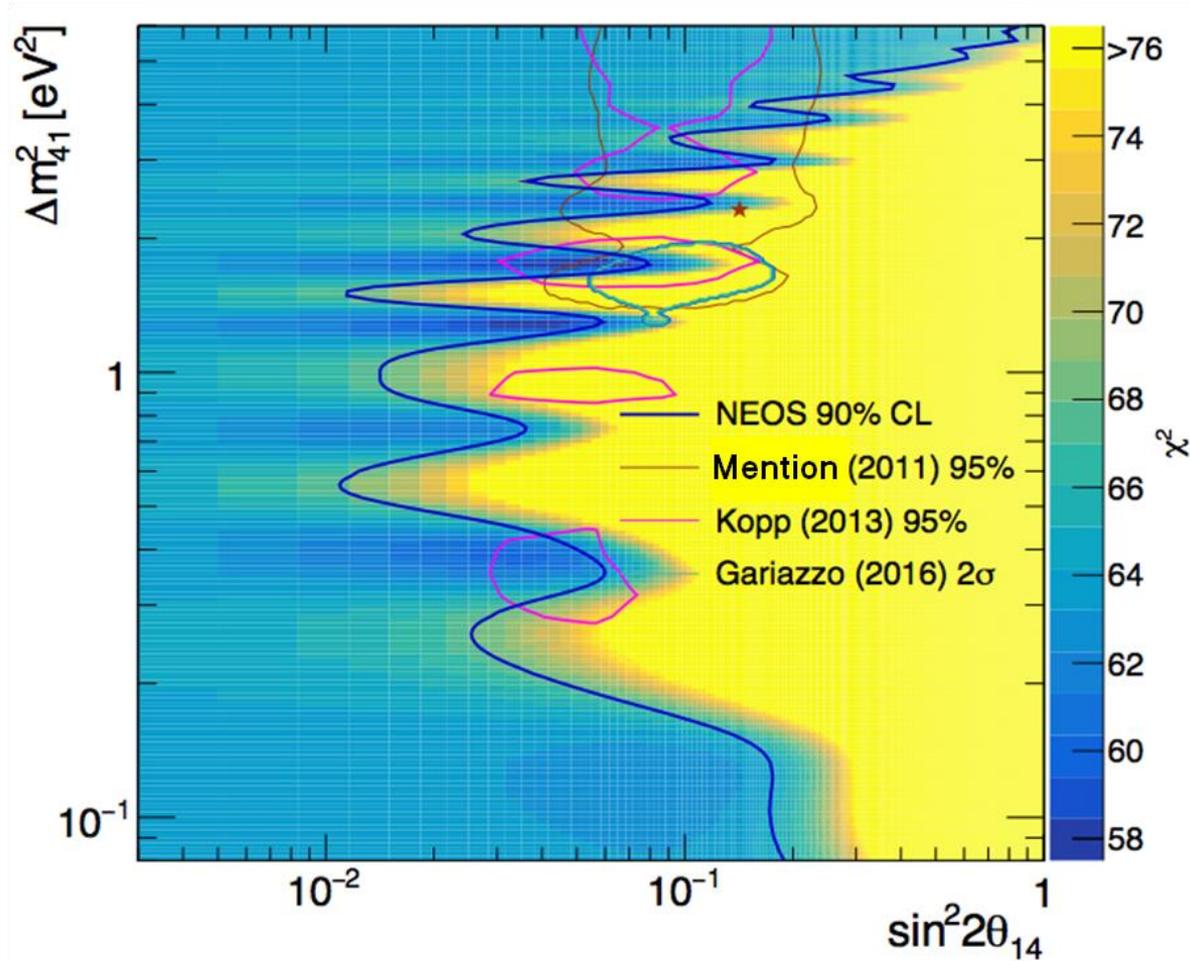
# Oscillation



- $\chi^2$  minima at  
 $(\sin^2 2\theta, \Delta m^2) = (0.04, 1.3 \text{ eV}^2)$   
 $(0.05, 1.73 \text{ eV}^2)$   
 with  $\Delta\chi^2 = \chi^2_{3\nu} - \chi^2_{4\nu} = 6.5$ .
- p-value  $\sim 22\%$
- No strong sign of 3+1 active-to-sterile  $\nu$  oscillation.



# Limits



Phys. Rev. Lett. 118, 121802 (2017)

# Future Works

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- “NEOS2” is in proposal stages
  - To install at a newly planned powerplant site.
    - New reactors would have  $\sim 4 \text{ GW}_{\text{th}}$  (Hanbit  $\sim 2.8 \text{ GW}_{\text{th}}$ )
    - We are writing a LoI for a dedicated space for various neutrino experiments.
  - Measuring one burn-up cycle at Hanbit NPP should help to understand the spectrum in details
- We are investigating a possible collaboration with DANSS.

# Summary

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- We examined neutrino spectrum shape at  $\sim 24$  m from a reactor.
- No strong evidence for oscillation from 3+1 neutrino ( $\Delta m^2 \sim 1 \text{ eV}^2$ ) mixing.
  - Best fits to  $(\sin^2 2\theta, \Delta m^2) = (0.04, 1.3 \text{ eV}^2) / (0.05, 1.73 \text{ eV}^2)$
  - RAA optimum is highly disfavored.