

Sterile Neutrino Experiments at Accelerator

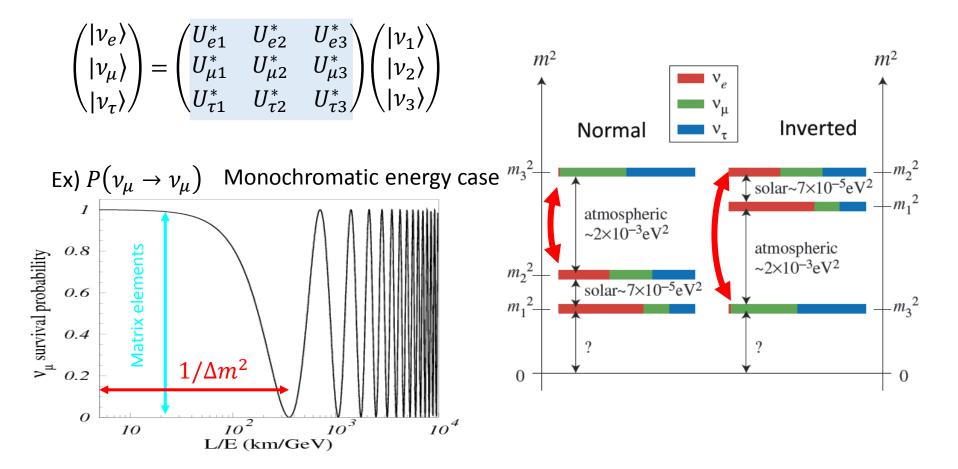
Takeo Kawasaki Kitasato University

Rencontres du Vietnam, Neutrinos, Quy Nhon, July 16-22,2017

Outline

- Neutrino oscillation
 - Sterile neutrino 3+1 model
 - Neutrino oscillation experiments at accelerator
- Indication of sterile neutrino
 - LSND
 - MiniBooNE
- Sterile neutrino search at accelerator experiment
 - JSNS²
 - SBN
 - NuPRISM
- Summary

Neutrino mixing matrix (3 flavors)



 ν oscillation is explained by mixing matrix elements and mass difference. Most Exp. Results are consistent with 3 neutrino mixing model.

Neutrino mixing matrix (4 or more flavors)

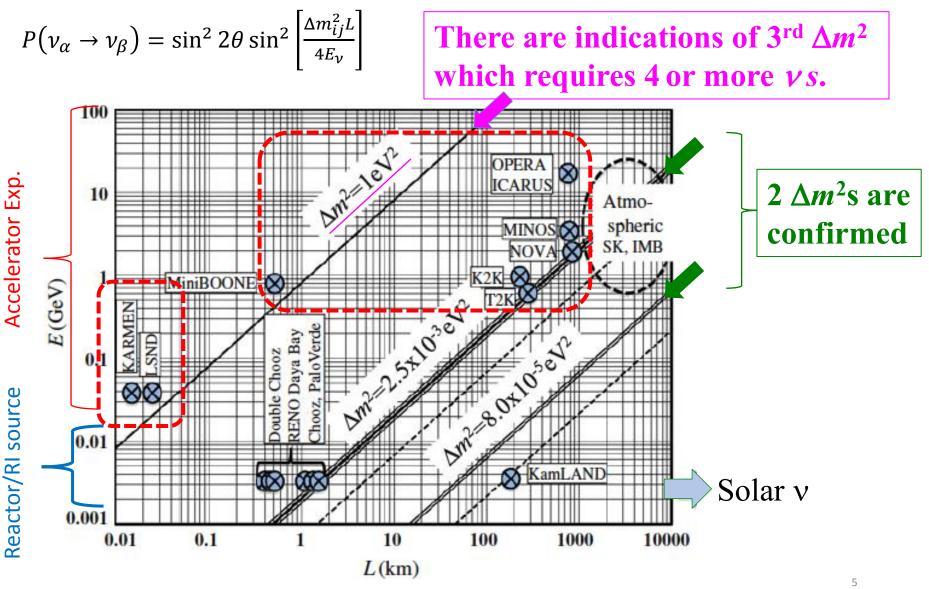
If there are more generations, it affect to the oscillation

$$\begin{pmatrix} |v_{e}\rangle\\|v_{\mu}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|v_{\tau}\rangle\\|$$

 v_s may be detected by v oscillations

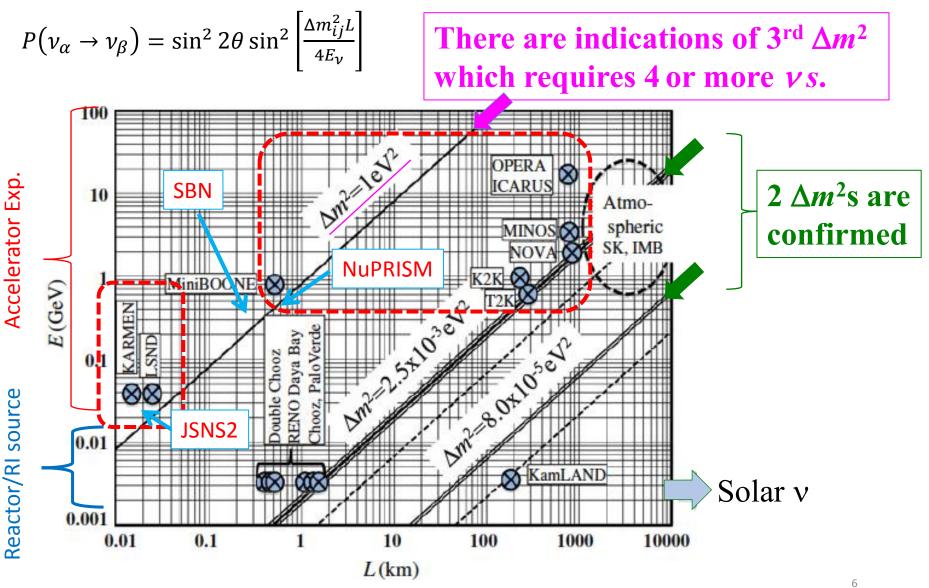
E-L relation of *v* oscillation experiments

Suekane@HINT2016

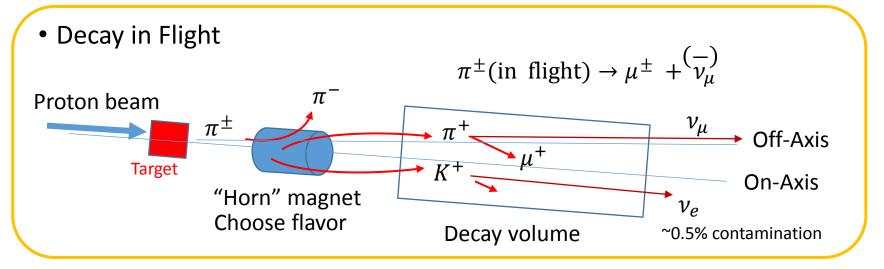


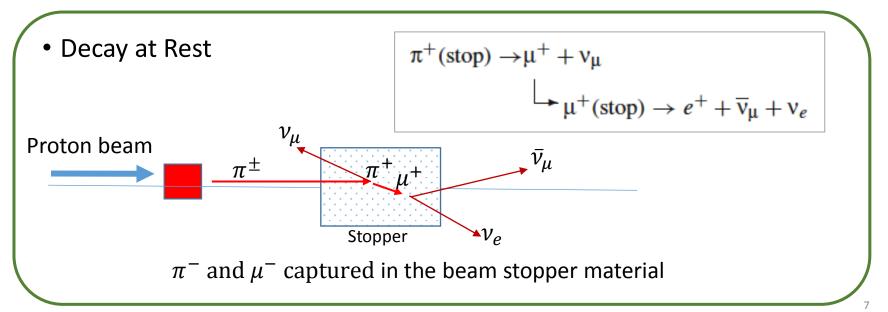
E-L relation of *v* oscillation experiments

Suekane@HINT2016



How to make neutrino beam





Indications of Sterile neutrinos

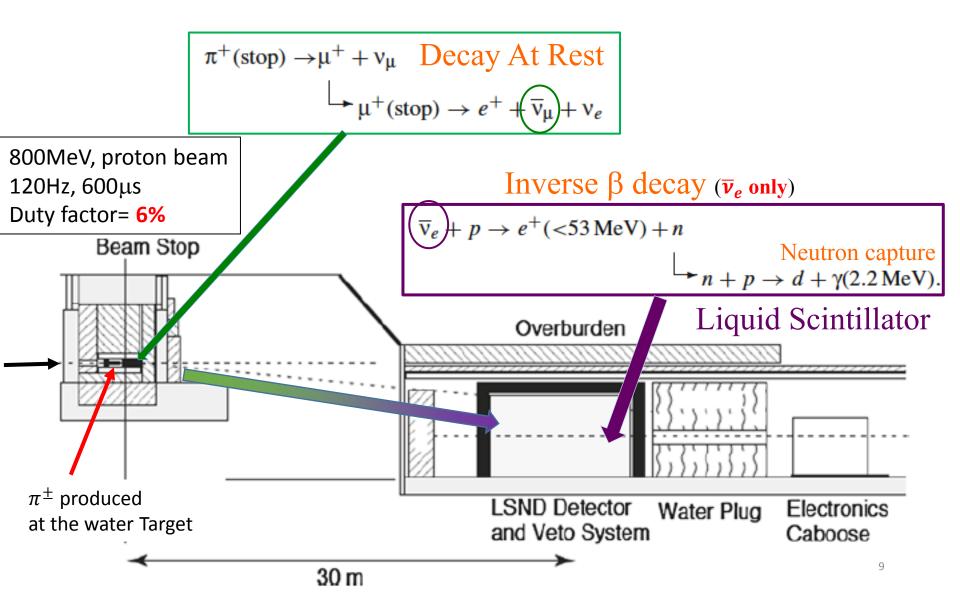
Experiment	v source	Mode	Significance
LSND	Decay-At-Rest	$ u_{\mu} ightarrow u_{e}$	3.8σ
MiniBooNE	Decay-In-Flight	$ u_{\mu} \rightarrow \nu_{e} $	3.4σ
		$\bar{ u}_{\mu} ightarrow \bar{ u}_{e}$	2.8σ
		combined	3.8σ
Ga-Solar	e capture	$v_e \rightarrow v_x$	2.7σ
Reactor	β-decay	$\overline{v_e} \rightarrow v_x$	3.0σ

Strong indication at Appearance modes at Accelerator experiment Reactor/source experiments show indication at Disappearance mode.

However, no positive result at Dis-appearance mode at Accelerator Exp.

LSND experiment

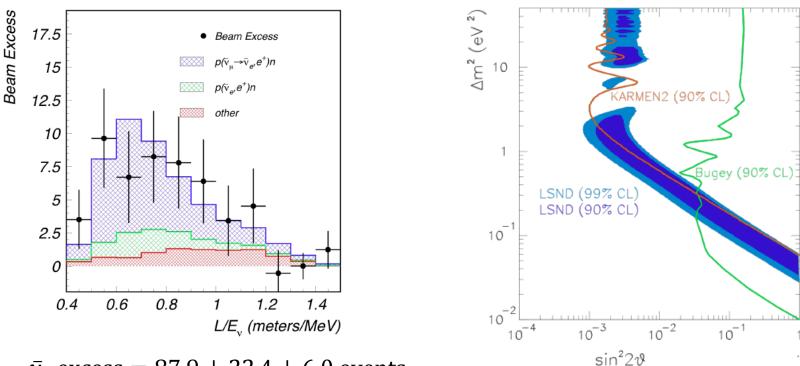
The most significant indication of the sterile $\boldsymbol{\nu}$



LSND Results

Phys. Rev. D64, 112007 (2001)

Oscillation Parameter



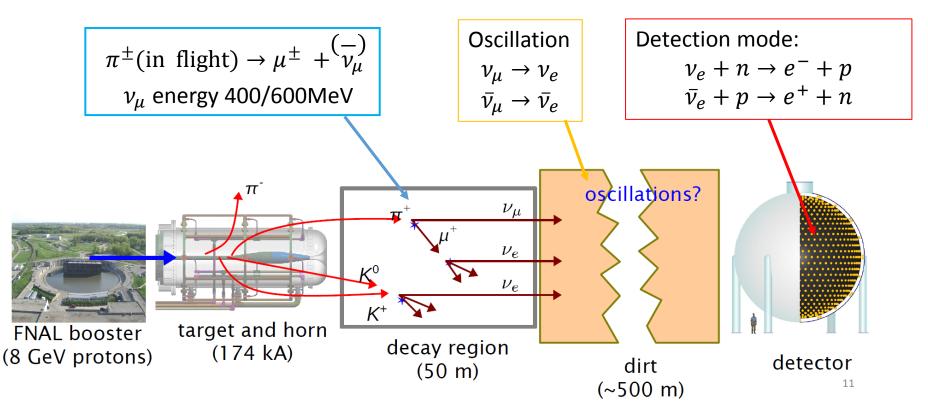
 $\bar{\nu}_e \text{ excess} = 87.9 \pm 22.4 \pm 6.0 \text{ events}$

To reject beam/cosmic related BG, many complicated cut were applied. Unknown systematics may have been introduced.

MiniBooNE experiment

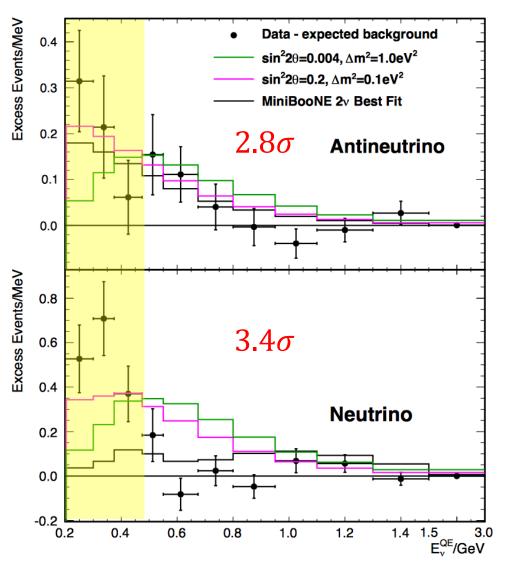
MiniBooNE Detector

- Placed at 500m to preserve LSND results with similar L/E
- 800-ton mineral oil
 - Cherenkov/Scintillation
 - Ring topology to separate e, μ and π^0
 - Can't distinguish e from γ



MiniBooNE results

Phys. Rev. Lett.110, 161801 (2013)



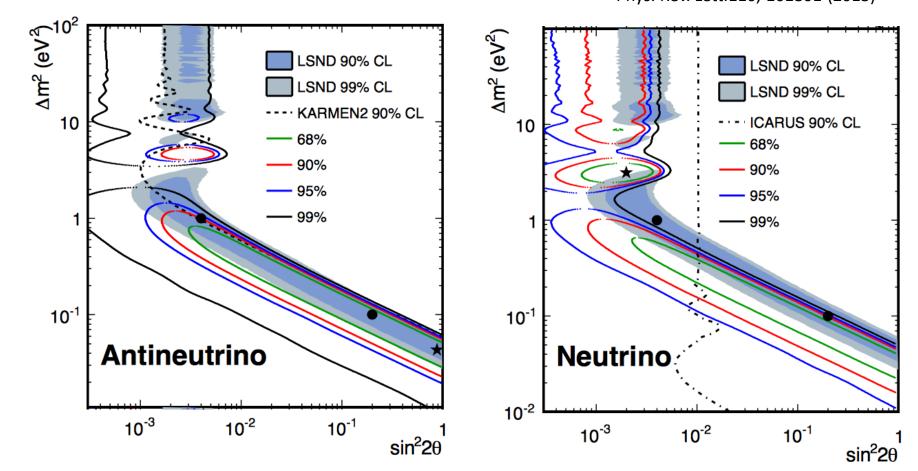
Event excess $\bar{\nu}$ mode: 78.4 ± 28.5 (2.8 σ) ν mode: 162.0 ± 47.8 (3.4 σ)

 ν mode case;

Large excess at low energy <475MeV But no excess at higher energy. The fit result doesn't explain it? \Rightarrow Due to miss-ID γ as electron?

MiniBooNE/LSND Results Comparison

Phys Rev. D80,073001 (2009) Phys. Rev. Lett.110, 161801 (2013)



MiniBooNE $\overline{\nu}$ mode is more compatible with LSND result compared with ν mode. How the difference is explained?

Short baseline accelerator experiments in preparation

LSND/MiniBooNE results should be checked by accelerator experiment for Appearance/Dis-Appearance mode.

J-PARC(Japan Accelerator Research Complex)



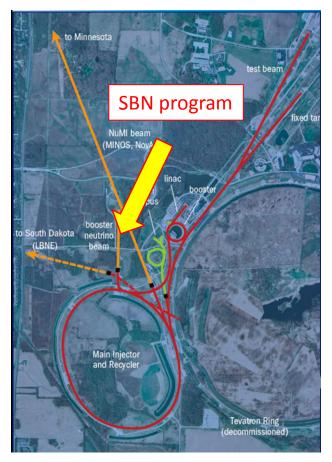
• JSNS² @J-PARC/MLF



- SBN program@FermiLAB BNB
- NuPrism@J-PARC /v-beam line

DIF/Appearance/Dis-appearance

Fermilab Booster ν beam



Direct test of LSND result

JSNS² @J-PARC/Material Life science Facility

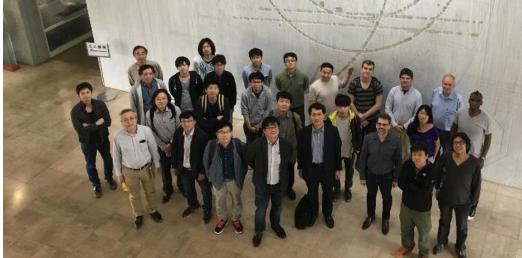
- Stage-1 approved.
- Funded for 1 detector

Members experienced at Reactor Exps. (Daya-Bay, DoubleChooz, RENO)

Consists of -21 institutes -53 collaborators -Japan/US/Korea Technical Design Report (TDR): Searching for a Sterile Neutrino at J-PARC MLF (E56, JSNS²)

S. Ajimura¹, M. K. Cheoun², J. H. Choi³, H. Furuta⁴, M. Harada⁵, S. Hasegawa⁵, Y. Hino⁴, T. Hiraiwa¹, E. Iwai⁶, S. Iwata⁷, J. S. Jang⁸, H. I. Jang⁹, K. K. Joo¹⁰, J. Jordan⁶, S. K. Kang¹¹, T. Kawasaki⁷, Y. Kasugai⁵, E. J. Kim¹², J. Y. Kim¹⁰, S. B. Kim¹³, W. Kim¹⁴, K. Kuwata⁴, E. Kwon¹³, I. T. Lim¹⁰, T. Maruyama^{*15}, T. Matsubara⁴, S. Meigo⁵, S. Monjushiro¹⁵, D. H. Moon¹⁰, T. Nakano¹, M. Niiyama¹⁶, K. Nishikawa¹⁵, M. Nomachi¹, M. Y. Pac³, J. S. Park¹⁵, H. Ray¹⁷, C. Rott¹⁸, K. Sakai⁵, S. Sakamoto⁵, H. Seo¹³, S. H. Seo¹³, A. Shibata⁷, T. Shima¹, J. Spitz⁶, I. Stancu¹⁹, F. Suekane⁴, Y. Sugaya¹, K. Suzuya⁵, M. Taira¹⁵, W. Toki²⁰, T. Torizawa⁷, M. Yeh²¹, and I. Yu¹⁸





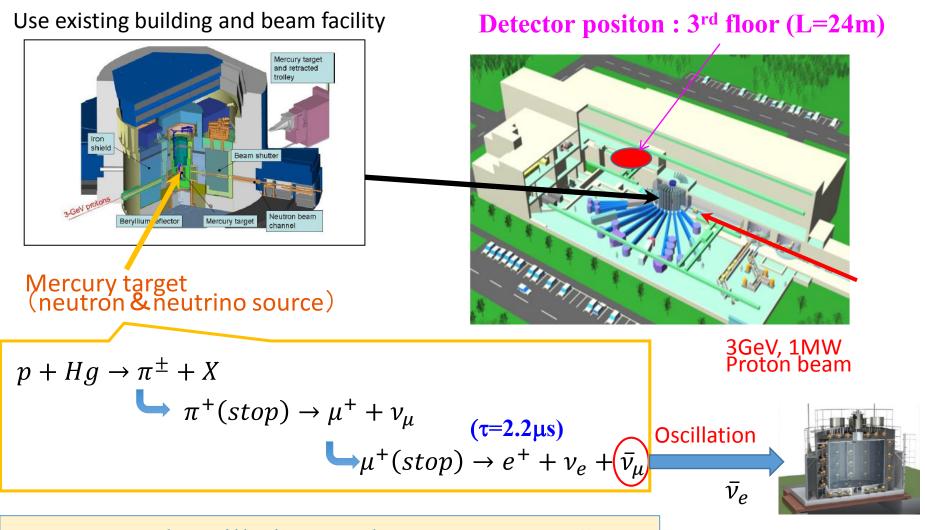
TDR released on May 2017

arXiv:1705.08629

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JSNS²: J-PARC Sterile Neutrino Search at J-PARC Spallation Neutron Source.

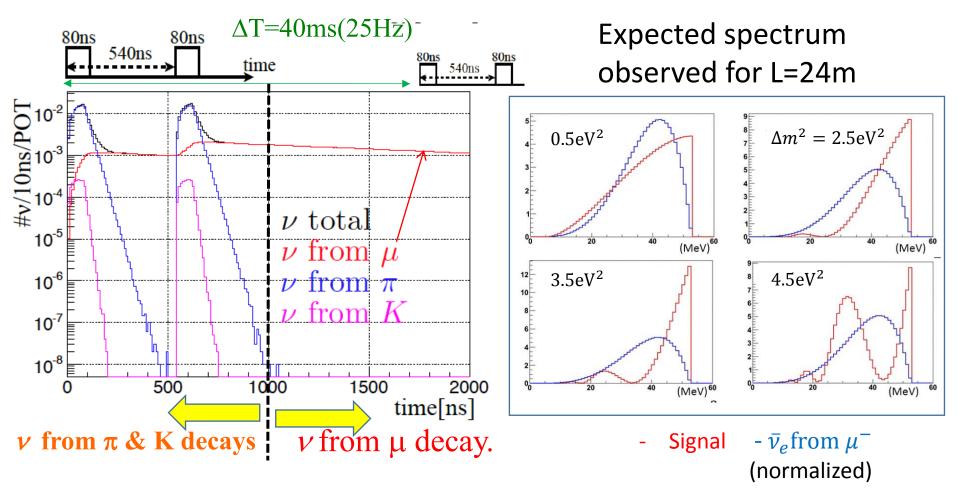
Experimental site: J-PARC MLF



v source and oscillation mode are same as LSND

Thanks to high-Z of Hg, $\bar{\nu}_e$ from μ^- decays are suppressed to 1/1,000.

 ν beam

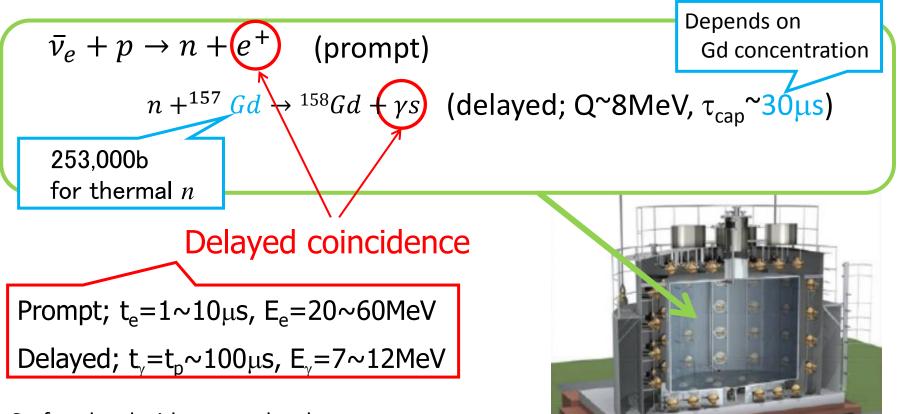


v timing window = $1 \sim 10 \mu s$

Pulsed beam can reduce beam BKG and accidental BKG drastically Cosmic-ray and accidental BG: suppression factor= 9ms/40ms~2x10⁻⁴

JSNS²:Detector

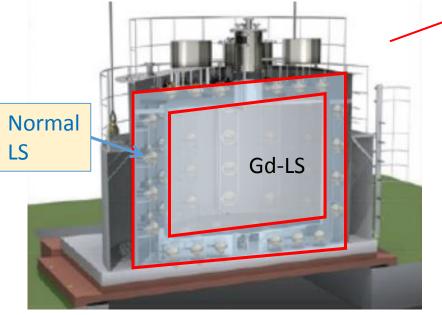
Gd-loaded Liq. Scintillator 17 tons × 2 (Established for reactor Exp.)



Surface level without overburden

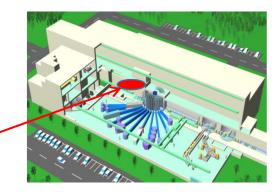
Further BG suppression with PSD (Pulse shape discrimination)

JSNS²:Detector



Use 2 of them. The detector size is limited by Facility requirement.

- Well established technology for reactor experiments.
 - Gd loaded LS in transparent Acrylic tank
 - Normal LS in Stainless Tank
- 200 8-inch PMTs to ensure good Energy resolution
- Huge cosmic induced BG at surface level.
 - LS region is divided to make Veto region.



3rd Floor for MLF maintenance work





Neutron beam lines

Comparison with LSND

Neutrino beam

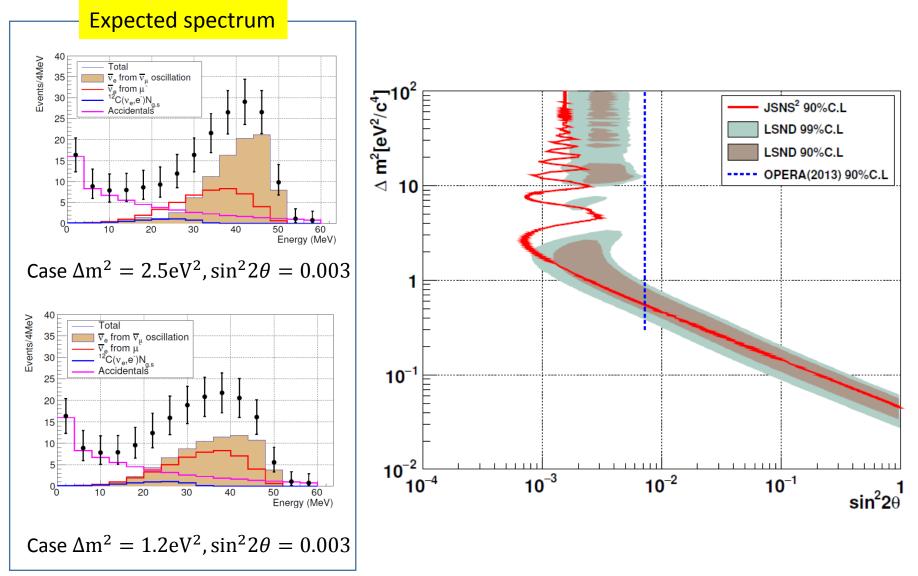
	Facility	Beam Pow. [MW]	Rep. Rate [Hz]	Pulse Width [ns]	Duty Factor
JSNS ²	J-PARC/MLF	1	25	620	1.55e-5
LSND	LANL/LAMPF	0.8	120	6e+5	0.072

Detector

	Туре	Mass [t]	L [m]
JSNS ²	Gd-LS with PSD	34	24
LSND	LS	167	30

$JSNS^2$: $S/N > 1,000 \times LSND$

JSNS² sensitivity (3y, 1 MW, 17 tons)



Schedule for the 1st Detector (17-ton)

Now

Japan Fiscal Year (April – March)

JSNS2 group is now preparing for the 1st detector construction

		-	-		
JFY	2017	2018	2019	2020	2021
S.S/Acrylic Tank					
PMT install	-				
LS production	-				
Dry run		-			
LS filling		-			
Start Exp. Run		-			

Continuous work for 2nd detector (17t) for increase of statistics and measurement with different baseline.

The bidding of Stainless tank was started on June, 2017 !

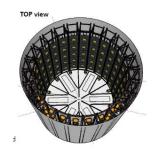
JSNS² status

Work for Detector construction (from TDR)

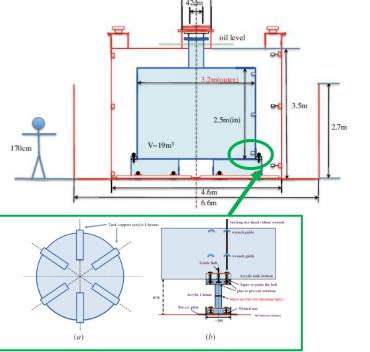
Stainless/Acrylic Tank design



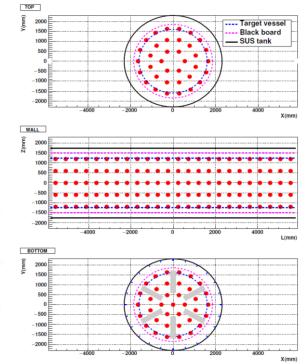
We have experiences at reactor Expetiments



(inner)PMT location



Ie box FINMET PMT support structure PMT PET PMT PET Black acrylic board



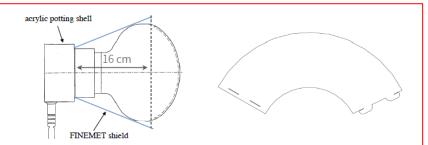
Consideration on Acrylic tank and PMTinstallation to S.S tank.

Compatibility tests of material inside Liquid Scintillators.

It is Important to chose the material installed into LS (LAB base).

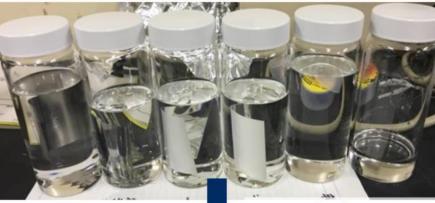
Check attenuation length of the liquids and the condition of the material surface/weight after long soak.

FINEMET for PMT magnetic shield





LUIREMIRROR for light reflection at veto region @Tohoku



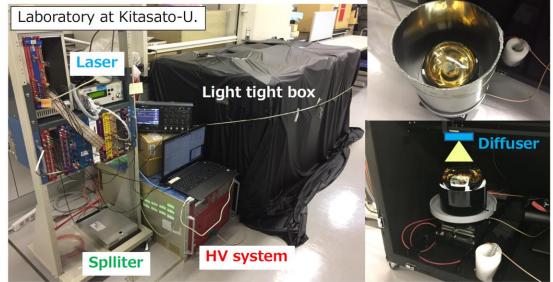
> 4 weeks later



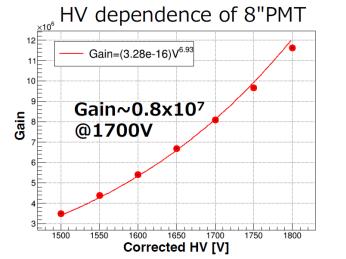
PMT Pre-Test/Calibration

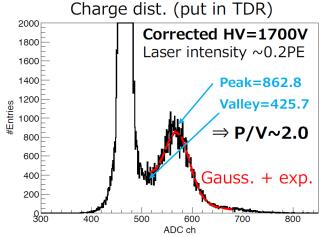
Development has been done. We are measuring a sample PMT.

8"PMT + µ-metal



• 004 5 4 20 • 004 5 4 20 • 004 5 4 20 • 004 5 4 20





Gain and P/V ratio were measured for PMT candidate.

@Kitasato-U



SBN program@Fermilab

Address MiniBooNE anomaly

- Funded
- Booster neutrino Beam ~700MeV
 - L/E ~ 1km/GeV
- Use 3 Liquid Argon TPCs(100-600t)
 - Good separation e/γ
 - Study v-LAr cross section
 - Further development for DUNE
- Sensitive to sterile oscillation
 - $v_{\mu} \rightarrow v_{e}$ appearance
 - $v_{\mu} \rightarrow v_{\mu}$ dis-appearance

I borrowed the slides from * D.Schnitz @ Neutrino2016

arXive: 1503.01520

A Proposal for a Three Detector Short-Baseline Neutrino Oscillation Program in the Fermilab Booster Neutrino Beam

The ICARUS-WA104 Collaboration

M. Antonello¹⁶, B. Baibussinov³¹, V. Bellini⁵, P. Benetti³², S. Bertolucci⁶, H. Bilokon¹⁵, F. Boffelli³², M. Bonesini¹⁷, J. Bremer⁶, E. Calligarich³², S. Centro³¹, A.G. Cocco¹⁹, A. Dermenev²⁰, A. Falcone³², C. Famese³¹, A. Fava³¹, A. Ferrar⁶, D. Gibin³¹, S. Gninenko²⁰, N. Golubev²⁰, A. Guglielmi³¹, A. Ivashkin²⁰, M. Kirsanov⁴⁰, J. Kistel³⁸, U. Kose⁶, F. Mammoliti⁴, G. Mannocchi¹⁵, A. Menegolli³², G. Mango²¹, D. Mladenov⁶, C. Montanari³², M. Nessi⁶, M. Nicotetta³¹, F. Noto⁶, P. Picchi¹⁵, F. Pietropaolo³⁷, P. Ploński⁴², R. Potenza⁵, A. Rappoldi³², G.L Raselli³², M. Rossella², C. Rubbia^{4,6,11,16}, P. Sala¹⁸, A. Scaramelli¹⁸, J. Sobczyk⁴⁴, M. Spanu³², D. Stefan¹⁸, R. Sulej⁴³, C.M. Sutera⁵, M. Torti³², F. Tortorici⁵, F. Varanini³¹, S. Ventura³¹, C. Vignoli¹⁶, T. Wachala¹², and A. Zani³²

The LAr1-ND Collaboration

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M. Bass³⁰, S. Bertolucci⁶, M. Bishai³, A. Bitadze²⁵, J. Bremer⁶, L. Bugel²⁶, L. Camillera⁷, F. Cavanna¹⁰a, H. Chen³, C. Chi⁹, E. Church^{10,45}, D. Cianci⁷, G.H. Collin²⁶, J.M. Conrad²⁶, G. De Geronimo³, R. Dharmapalan¹, Z. Djurcic¹, A. Ereditato², J. Esquivel¹⁰, J. Evans²⁵, B.T. Fleming⁴⁵,
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J. Spitz²⁶, N. Spooner³⁷, T. Strauss², A.M. Szelc⁴⁵, C. E. Taylor⁴⁴, K. Terao⁶, M. Thiesse³⁷, L. Thompson³⁷, M. Thomson⁴, C. Thora³, M. Tougs²⁶, C. Touramanis²³, R.G. Van de Water²⁴, M. Weber², D. Whitington¹⁴, T. Wongjirad⁴⁶, B. Yu³, G.P. Zeller¹⁰, and J. Zennamo⁷

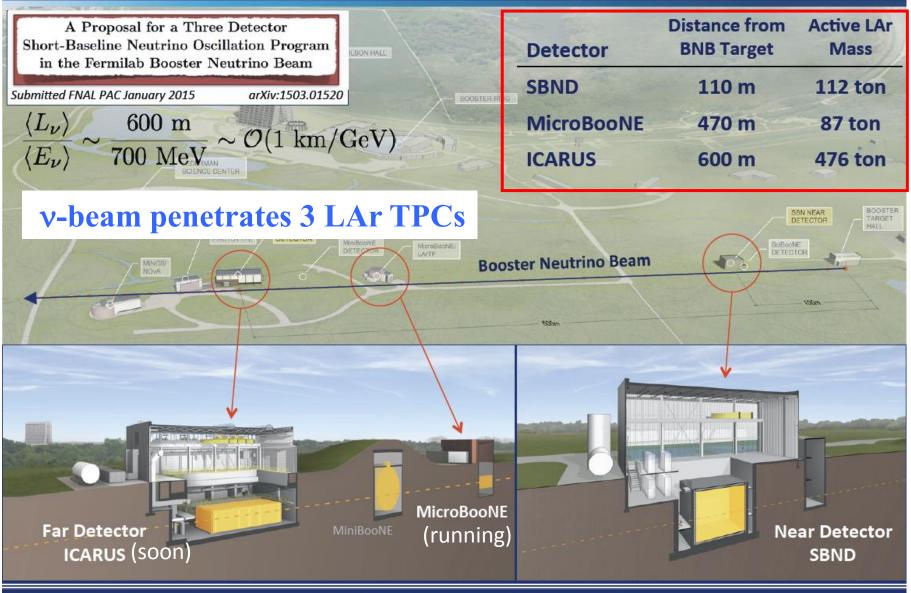
The MicroBooNE Collaboration

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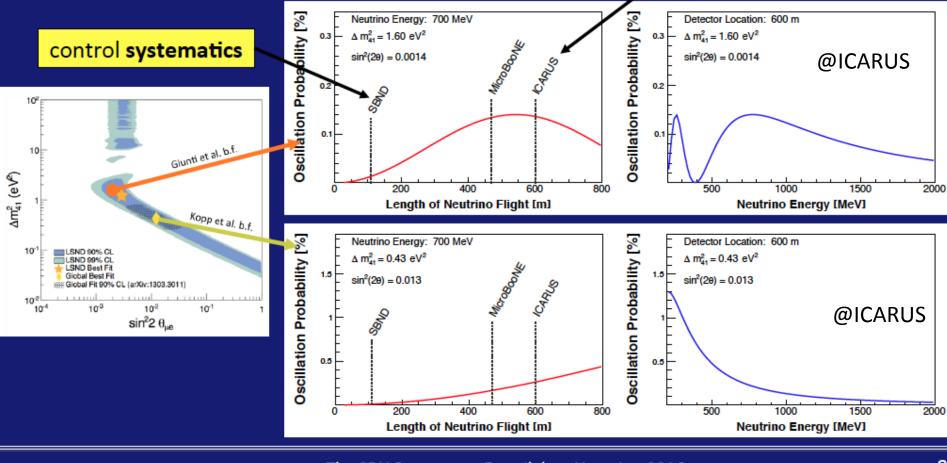
The Three LArTPC SBN Program



David Schmitz, UChicago

The SBN Program at Fermilab - Neutrino 2016

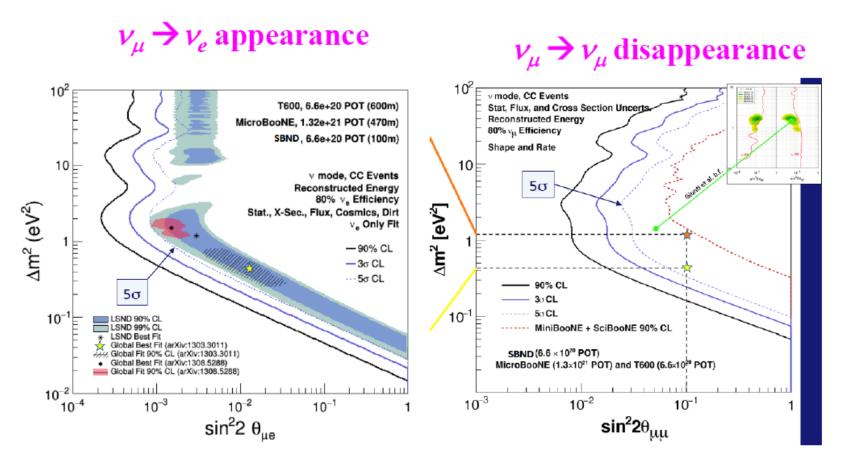
SBN will apply the advantages of the LArTPC technology and <u>multiple</u> detectors at different baselines to the question of high-∆m² <u>sterile neutrino oscillations</u> to definitively test currently allowed oscillation parameter regions at ≥5σ increase statistics of signal



David Schmitz, UChicago

The SBN Program at Fermilab - Neutrino 2016

Sensitivities for sterile ν oscillations(3yrs)



NuPrism@J-PARC neutrino beamline

Stage1 approved

Proposal for the NuPRISM Experiment in the J-PARC Neutrino Beamline

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R. Henderson,²⁶ T. Ishida,^{6,*} M. Ishitsuka,²³ C.K. Jung,^{14,†} A.C. Kaboth,⁸ H. Kakuno,²⁴ H. Kamano,¹⁶
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J. Vo,⁷ D. Wark,^{19,8} M.O. Wascko,⁸ M.J. Wilking,¹⁴ S. Yen,²⁶ M. Yokoyama,^{20,†} and M. Ziembicki²⁷

https://j-parc.jp/researcher/Hadron/en/pac_1507/pdf/P61_2015-5.pdf

- Use movable detector in a shaft hole
 - Water Cherenkov, ~1000 ton (depends on baseline)
 - Baseline :1-2km (not fixed yet)
 - Study ν nuclear interaction with various energy
- It is sensitive to sterile ν oscillation
 - Both appearance/dis-appearance mode

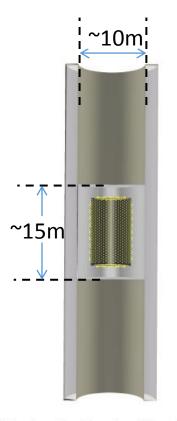
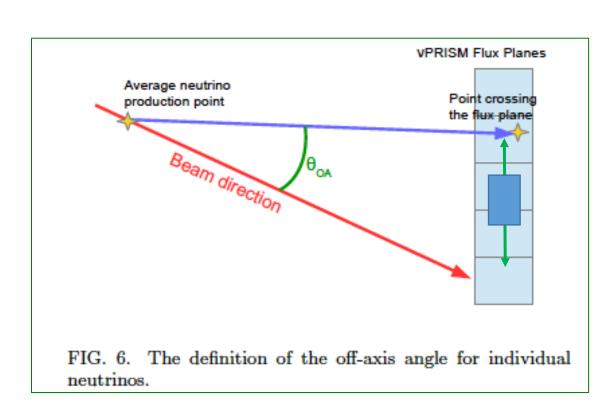
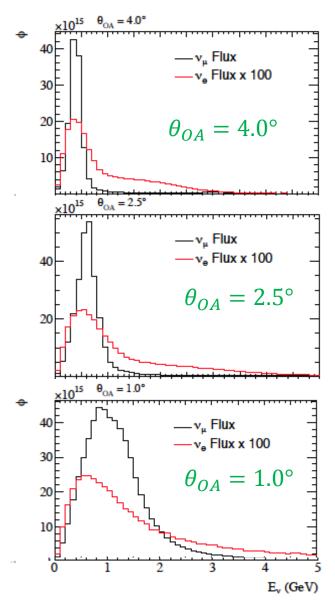


FIG. 39. The planned configuration of the nuPRISM detector within the water tank is shown. The instrumented portion of the tank moves vertically to sample different off-axis angle regions.

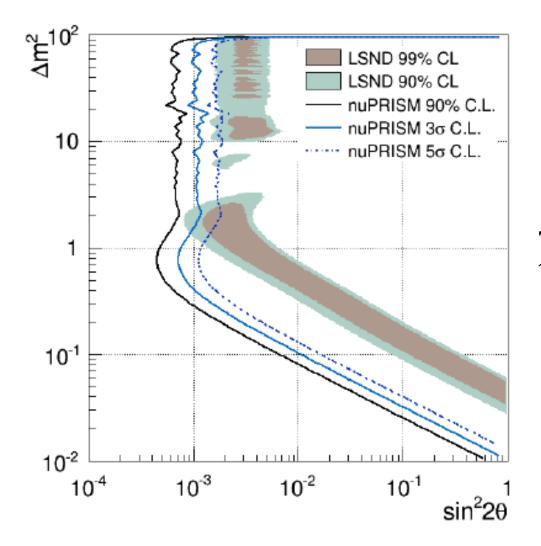


With Off-axis beam technique. Large angle gives sharp E_{ν} distribution.

Change Off-Axis angle by moving detector in a shaft,



NuPRISM: sterile ν sensitivity





T2K II exposure(2e21 POT) v_e appearance mode

suekane@HINT2016

Comparison of the future projects

	JSNS2	SBND/mBooN/ICARUS	nuPRISM	
Accelerator	J-PARC/MLF	Fermilab BvB	J-PARC/v	
E _v [MeV]	~50	~700	500~1500	
v source	$\mu^+\text{DAR}$	π^+ DIF	π^+ DIF	
L[m]	23	110/470/600	1000~2000	
Osc. mode	$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e}$	$v_{\mu} \rightarrow v_{e}, v_{\mu} \rightarrow v_{\mu}$	$V_{\mu} \rightarrow V_{e}, V_{\mu} \rightarrow V_{\mu}$	
Detector	Gd-LS	LAr-TPC	Water Cherenkov	
Fid. Mass(t)	17(phase0)	112/87/476	1000(depnds on L)	
Characteristic	Direct Test of LSND	Full v-Ar interaction reconstruction. Near/Far cancellation	Wide range semi- monochromatic v	
Status	funded for 1detector	funded	stage-1 approval	
Commissioning	2018~9	2018~9/2015/2018	- v	
Cost	\$	\$\$\$ SBN	\$\$\$ P R I S 3	
		Program	M S 3	

Summary

- There are some indications of sterile neutrinos
 - Significant at LSND/MiniBooNE: $\Delta m^2 \sim O(eV)$ at v_e appearance mode
 - Reactor/Source Exp. also show it in v_e dis-appearance mode
- SBL Accelerator experiments are sensitive to $\nu_{\mu} \rightarrow \nu_{e}$ and $\nu_{\mu} \rightarrow \nu_{x}$
 - Experiments are in preparation
 - JSNS2@J-PARC is funded for 1 detector
 - SBN@FermiLAB is funded and will start soon
 - Those are complementary to the reactor/source experiments
- The interesting results will appear in a few years.