





Search for Charged Lepton Flavour Violation (CLFV)

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Rencontres de Vietnam - Neutrino
ICISE, Quy Nhon, Vietnam

Outline



Outline



- Physics motivation of charged lepton flavour violation (CLFV)
- Muon CLFV experiments
- Muon to electron conversion
 - Mu2e
 - COMET
 - COMET Phase-I
- Summary

Why CLFV?



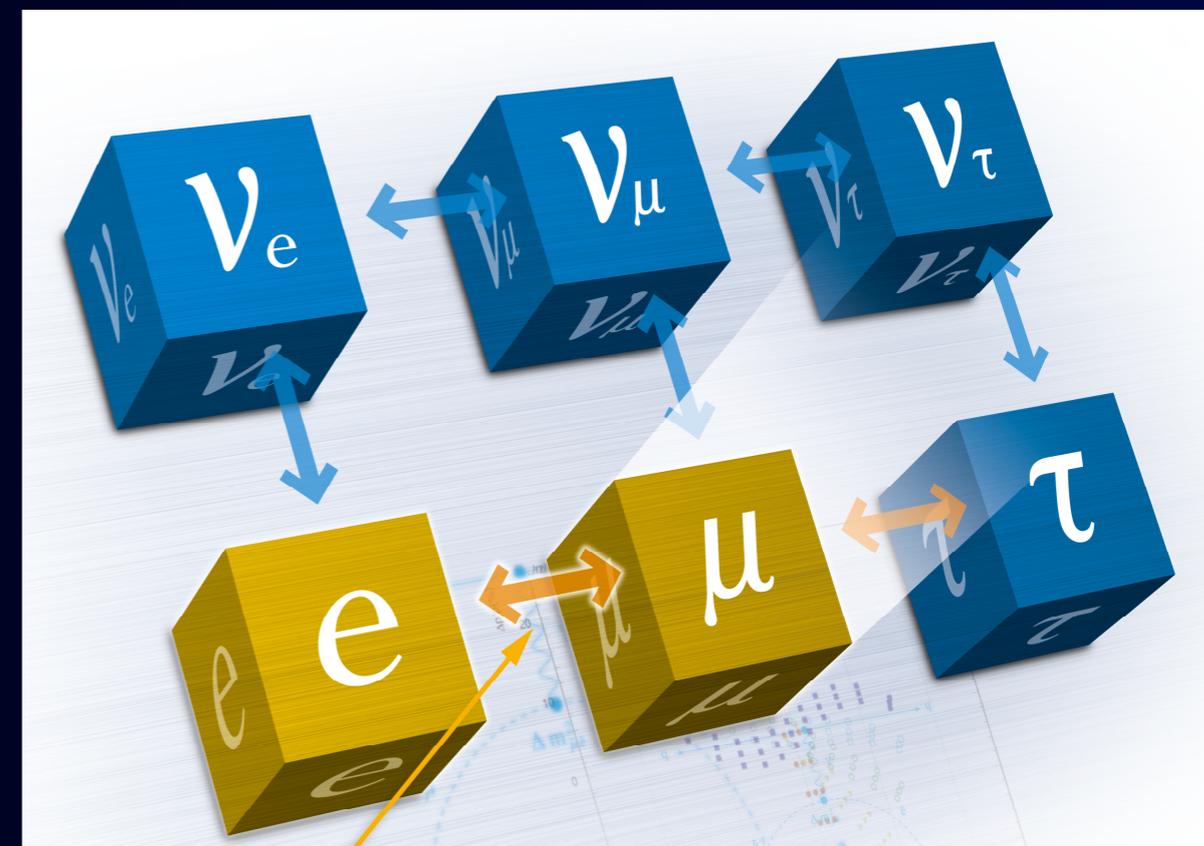
Neutrinos and Charged Leptons

PMNS mixing matrix

$$\nu_\ell = U_{\text{PMNS}} \nu_i, \quad \text{i.e.} \quad \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

CC current interaction

$$\mathcal{L}_{\text{cc}}^\ell \sim -\frac{g}{\sqrt{2}} U_{\text{PMNS}}^{\dagger ij} W_\mu^+ \bar{\nu}_{Li} \gamma^\mu e_{Lj}, \quad U_{\text{PMNS}} = R_L^\ell R_L^{\nu\dagger}$$



no oscillation of charged leptons
but transitions can occur.

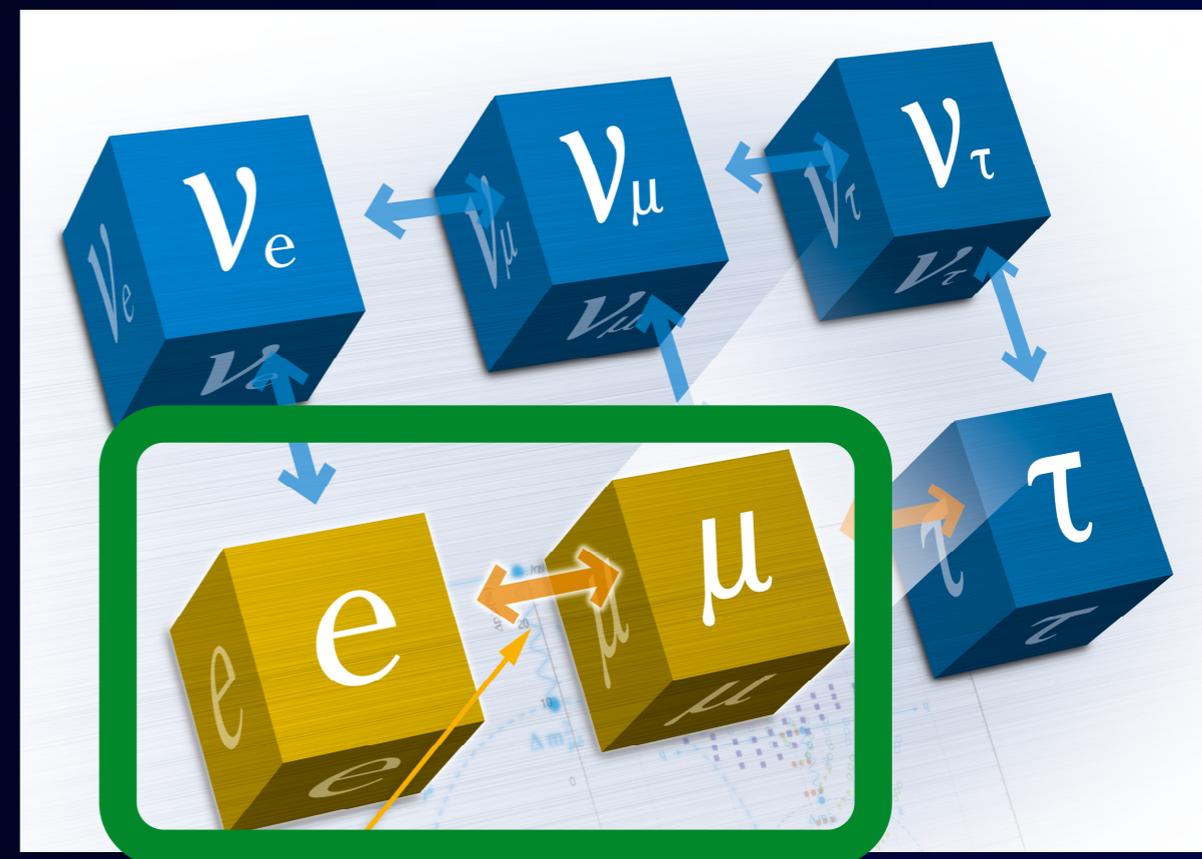
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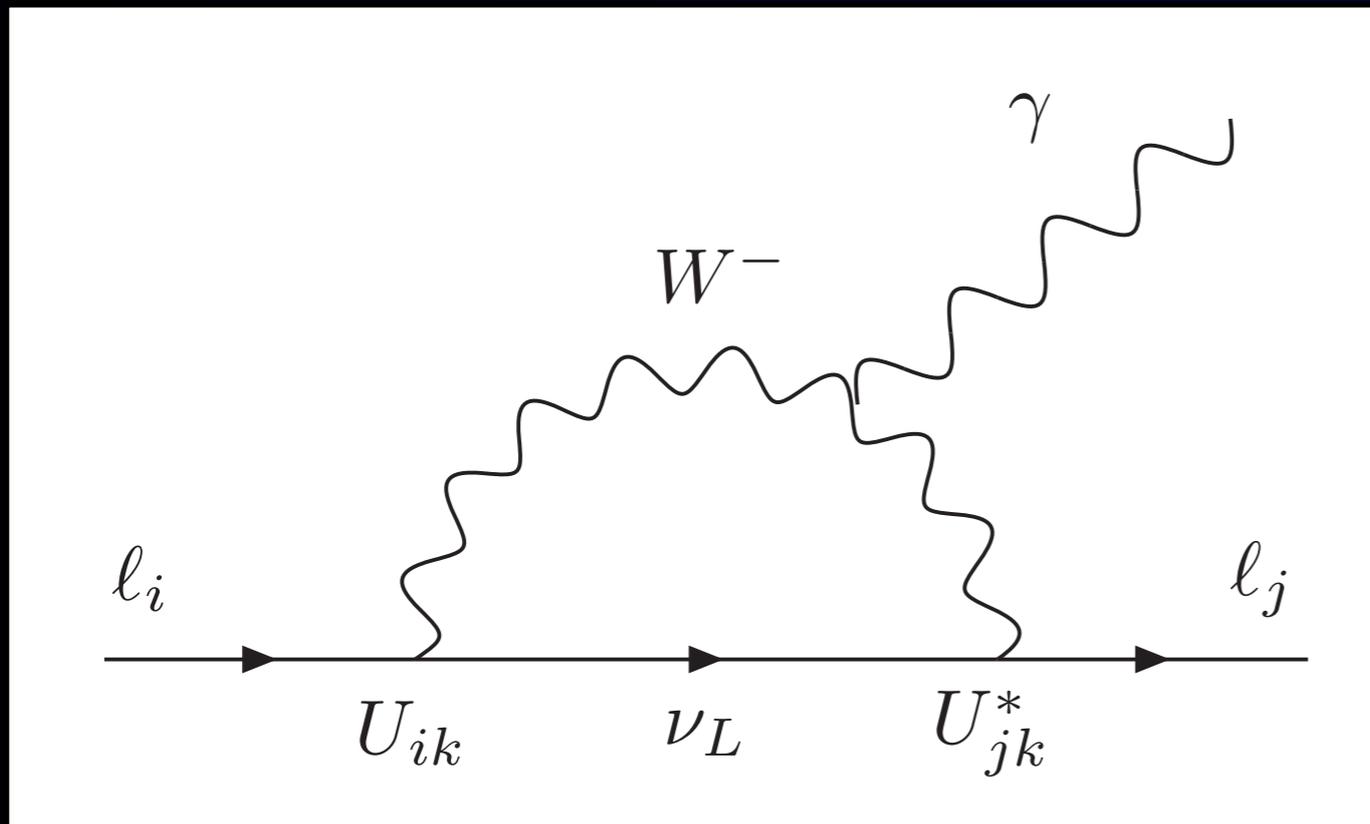


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SM Contribution to CLFV

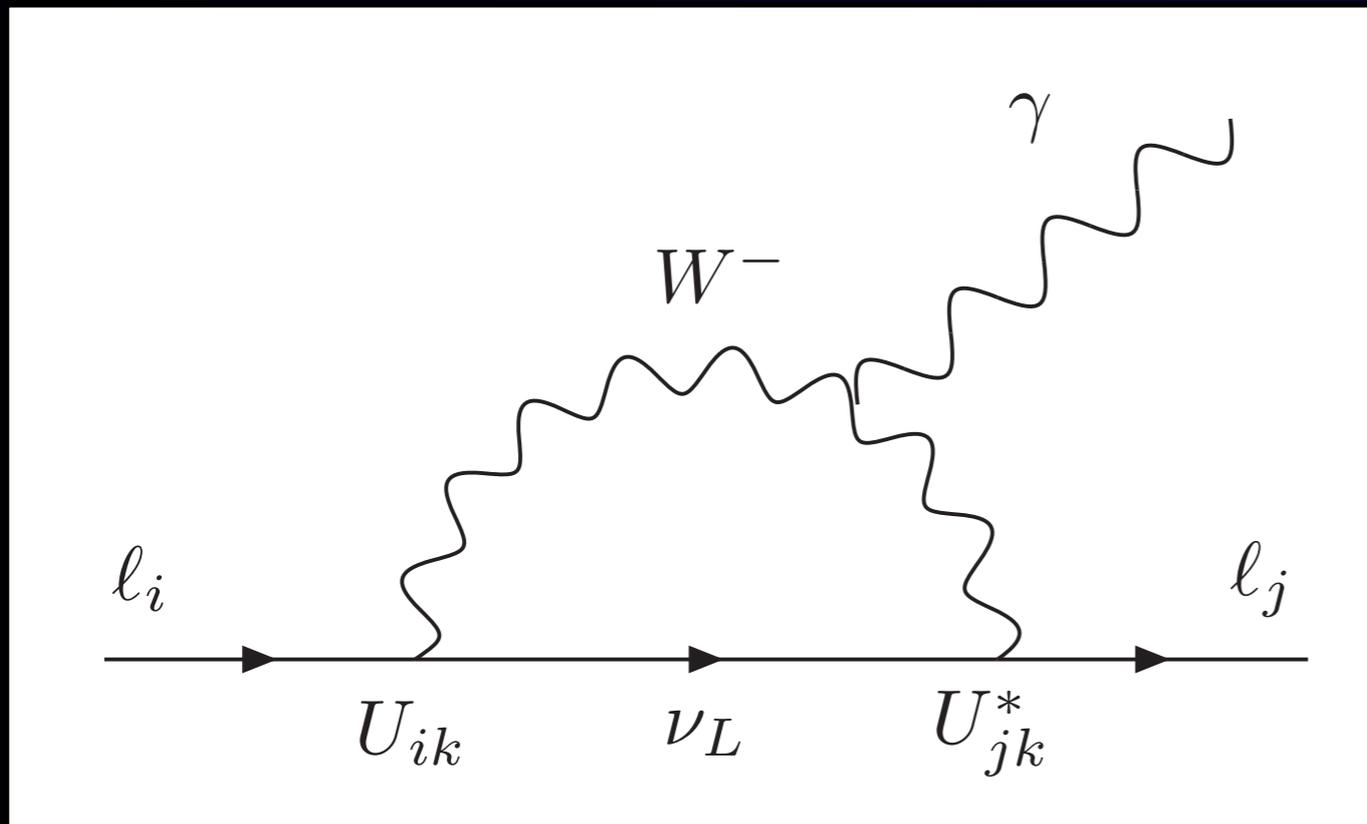


SM Contribution to CLFV



$$B(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_l (V_{MNS})_{\mu l}^* (V_{MNS})_{el} \frac{m_{\nu_l}^2}{M_W^2} \right|^2$$

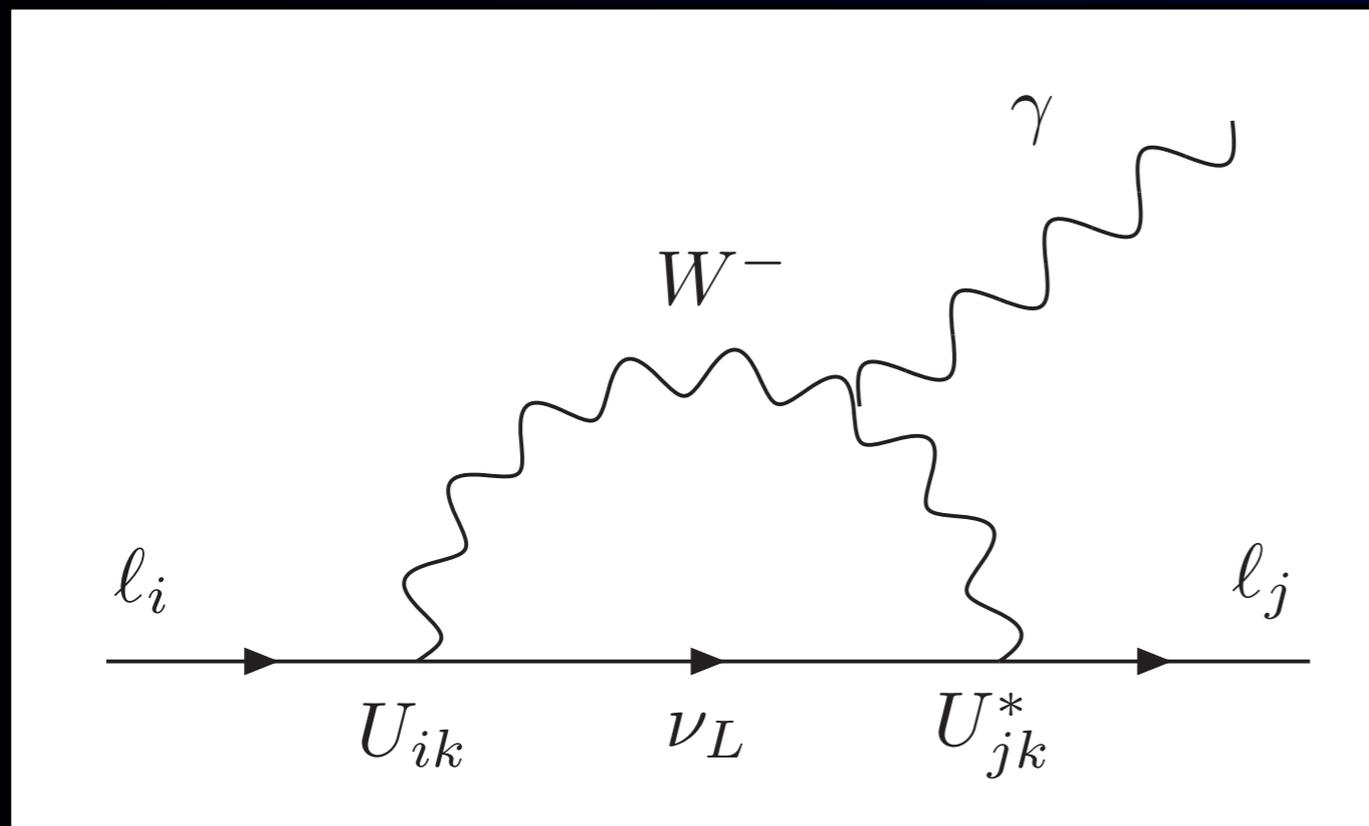
SM Contribution to CLFV



BR \sim O(10^{-54})

$$B(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_l (V_{MNS})_{\mu l}^* (V_{MNS})_{el} \frac{m_{\nu_l}^2}{M_W^2} \right|^2$$

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CLFV provides good opportunities to search for new physics beyond the Standard Model

New Physics Search with CLFV



Effective Lagrangian with New Physics



Effective Lagrangian with New Physics



$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)}$$

dimension 6

Λ is the energy scale of new physics ($\sim m_{\text{NP}}$)

C_{NP} is the coupling constant.

Effective Lagrangian with New Physics



$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)}$$

dimension 6

Λ is the energy scale of new physics ($\sim m_{\text{NP}}$)

C_{NP} is the coupling constant.

very high
energy scale Λ
with not-small
 C_{NP}

New Physics
could be....

or

very small C_{NP}
with not-high
energy scale Λ

Effective Lagrangian with New Physics



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ex: Charged lepton flavour violation (CLFV),
 $\mu \rightarrow e \gamma$ ($B < 4.2 \times 10^{-13}$ from MEG(2016))

$$\frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)} \rightarrow \frac{C_{\mu e}}{\Lambda^2} \bar{e}_L \sigma^{\rho\nu} \mu_R \Phi F_{\rho\nu}$$

$$\Lambda > 2 \times 10^5 \text{ TeV} \times (C_{\mu e})^{\frac{1}{2}} .$$

$$\Lambda > O(10^5) \text{ TeV with } C_{\mu e} \sim O(1)$$

or

$$C_{\mu e} \sim O(10^{-9}) \text{ with } \Lambda < O(1) \text{ TeV}$$

CLFV Rate



Lepton (SM forbidden)

rate

$$|A_{\text{SM}} + \varepsilon_{\text{NP}}|^2 \sim \cancel{|A_{\text{SM}}|^2} + \cancel{2\text{Re}(A_{\text{SM}}\varepsilon_{\text{NP}})} + \underline{|\varepsilon_{\text{N}}|^2}$$

$$R \propto \frac{1}{\Lambda^4}$$

CLFV Rate



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$$R \propto \frac{1}{\Lambda^4}$$

$$\Lambda \geq 10 \text{ TeV} \rightarrow R \leq 10^{-4}$$

Model Dependent CLFV



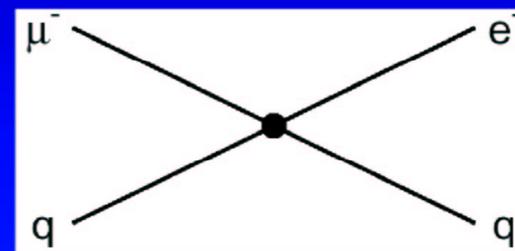
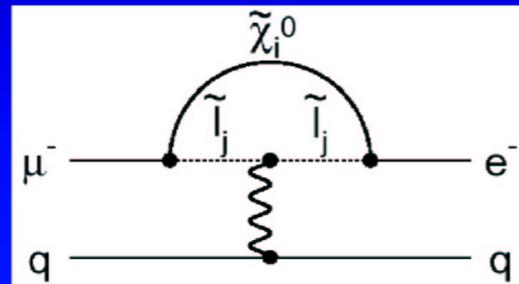
Model Dependent CLFV



Sensitivity to Different Muon Conversion Mechanisms

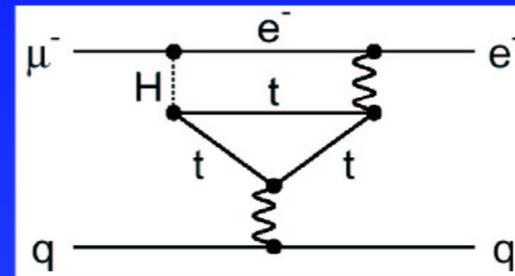
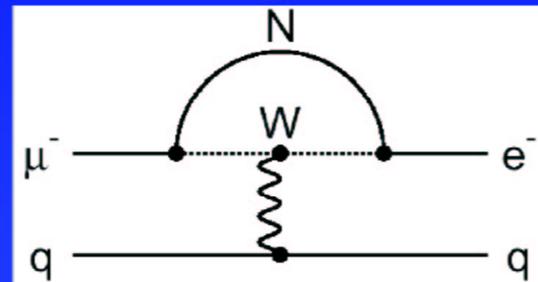


Supersymmetry
Predictions at 10^{-15}



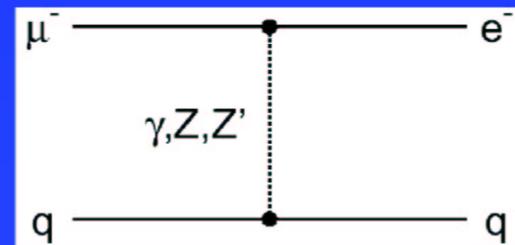
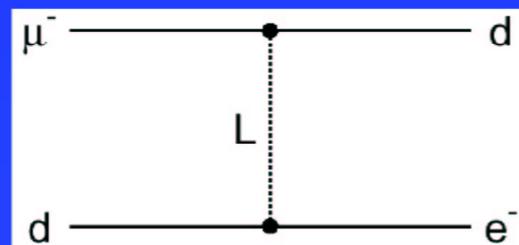
Compositeness
 $\Lambda_c = 3000 \text{ TeV}$

Heavy Neutrinos
 $|U_{\mu N}^* U_{eN}|^2 = 8 \times 10^{-13}$



Second Higgs doublet
 $g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$

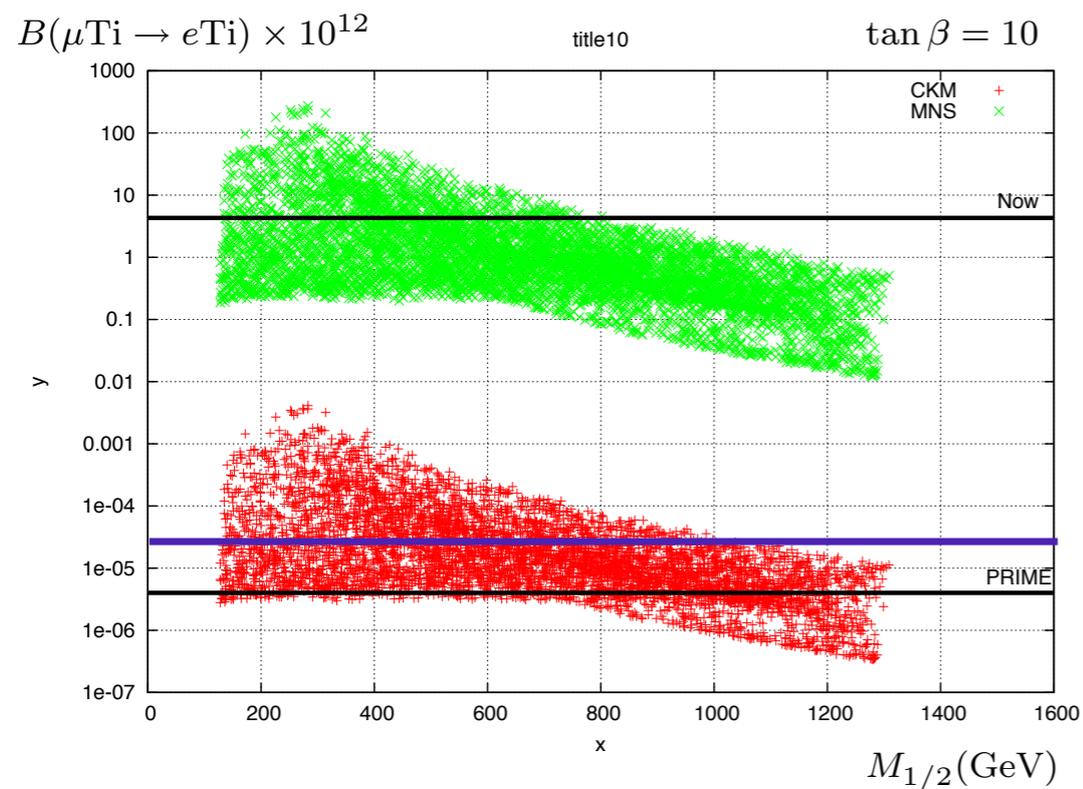
Leptoquarks
 $M_L = 3000 (\lambda_{\mu d} \lambda_{ed})^{1/2} \text{ TeV}/c^2$



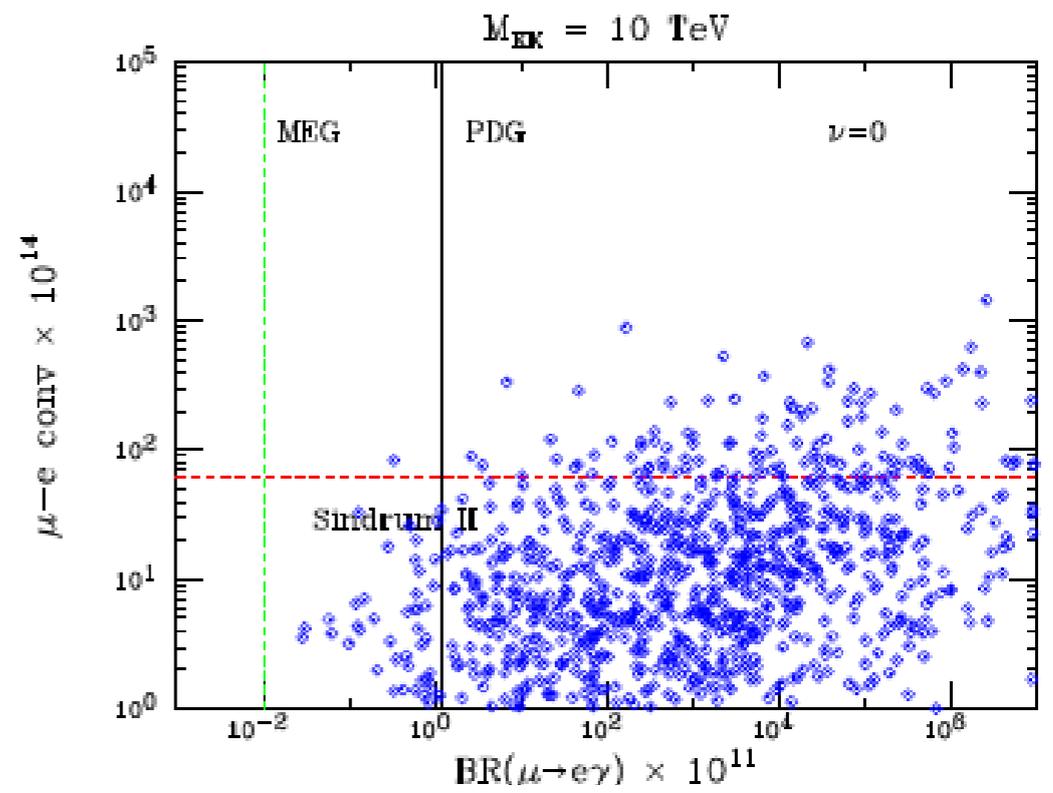
Heavy Z' ,
Anomalous Z
coupling
 $M_{Z'} = 3000 \text{ TeV}/c^2$
 $B(Z \rightarrow \mu e) < 10^{-17}$

After W. Marciano

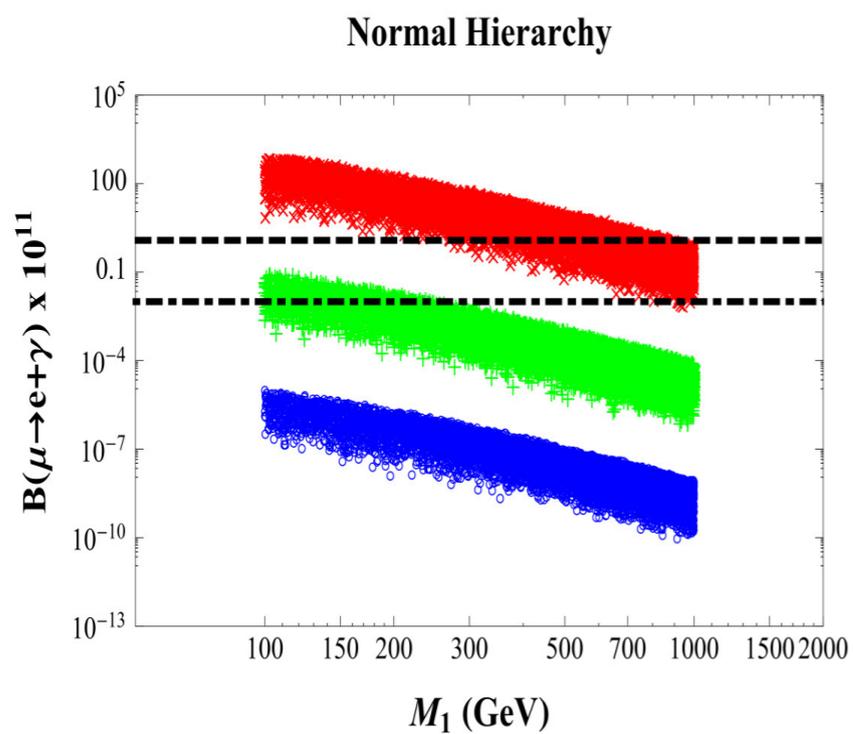
SUSY model



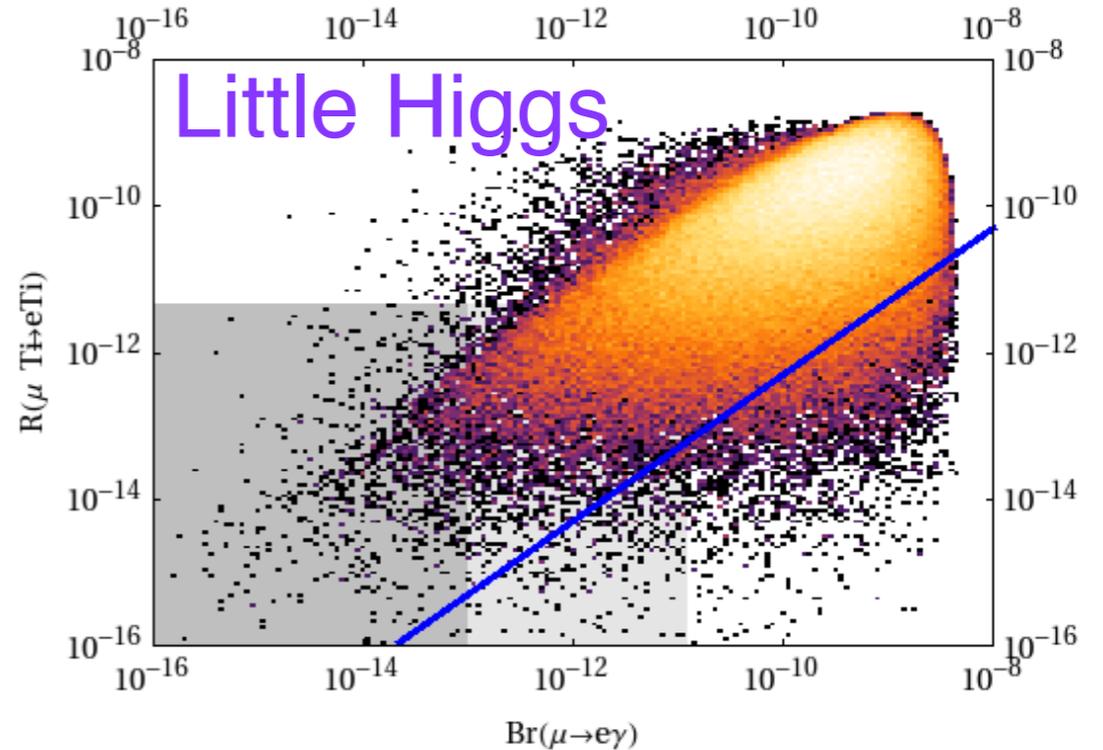
extra dimension model



low-energy seesaw model



little Higgs model



Muon CLFV



Experimental Limits at Present and in the Future



process	present limit	future	
$\mu \rightarrow e\gamma$	$<4.2 \times 10^{-13}$	$<10^{-14}$	MEG at PSI
$\mu \rightarrow eee$	$<1.0 \times 10^{-12}$	$<10^{-16}$	Mu3e at PSI
$\mu N \rightarrow eN$ (in Al)	none	$<10^{-16}$	Mu2e / COMET
$\mu N \rightarrow eN$ (in Ti)	$<6.1 \times 10^{-13}$	$<10^{-18}$	PRISM
$\tau \rightarrow e\gamma$	$<1.1 \times 10^{-7}$	$<10^{-9} - 10^{-10}$	superKEKB
$\tau \rightarrow eee$	$<3.6 \times 10^{-8}$	$<10^{-9} - 10^{-10}$	superKEKB
$\tau \rightarrow \mu\gamma$	$<4.5 \times 10^{-8}$	$<10^{-9} - 10^{-10}$	superKEKB
$\tau \rightarrow \mu\mu\mu$	$<3.2 \times 10^{-8}$	$<10^{-9} - 10^{-10}$	superKEKB/LHCb

Experimental Limits at Present and in the Future

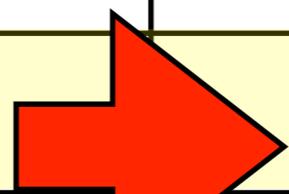


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$\times 10^{-4}$

$\mu \rightarrow e\gamma$



• **Event Signature**

- $E_e = E_\gamma = m_\mu/2$ (=52.8 MeV)
- angle $\theta_{\mu e}=180$ (back-to-back)
- time coincidence

• **Backgrounds**

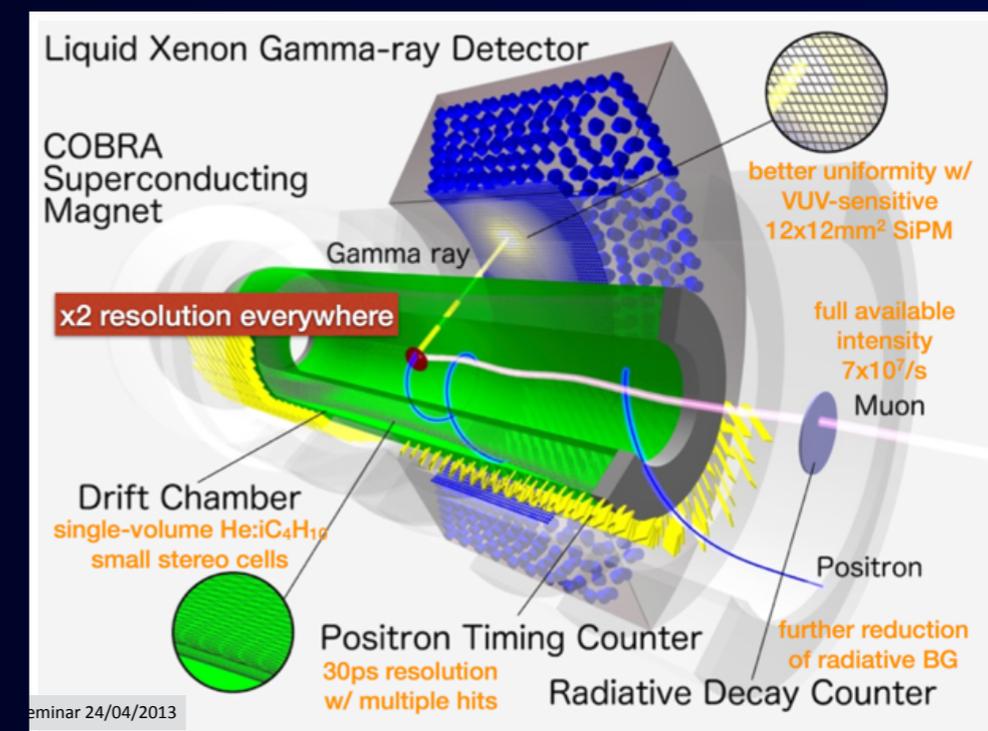
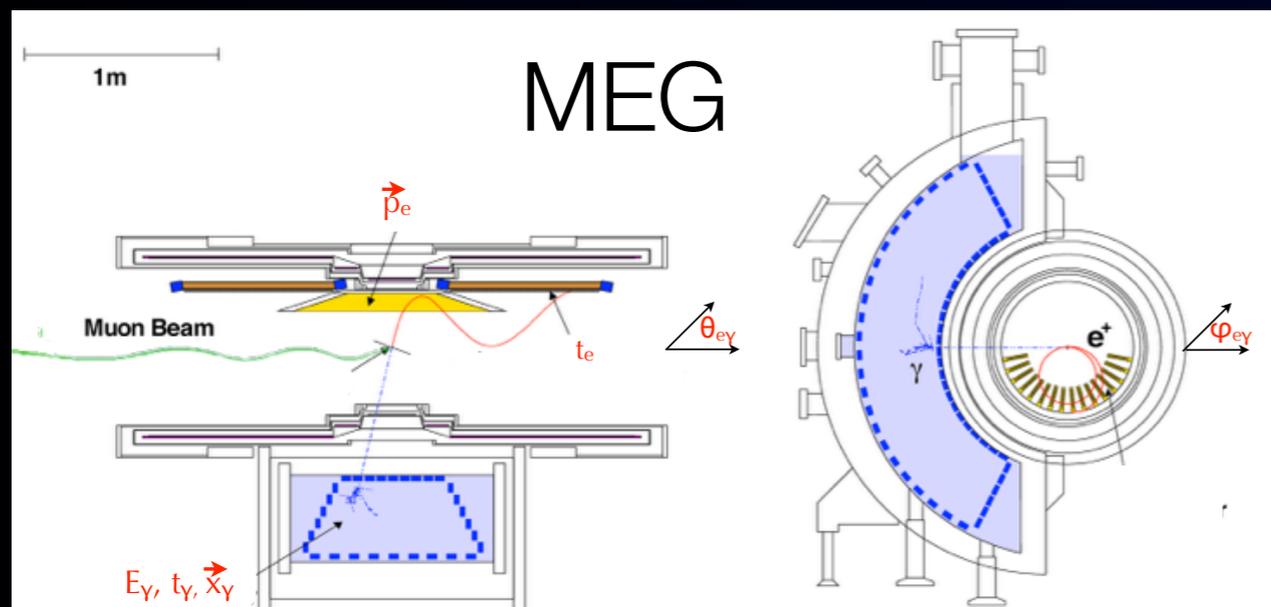
- prompt physics backgrounds
 - radiative muon decay $\mu \rightarrow e\nu\nu\gamma$
- accidental backgrounds

Final MEG result (2016)

$$B(\mu^+ \rightarrow e^+ \gamma) < 4.2 \times 10^{-13}$$

MEG II

goal $\sim 4 \times 10^{-14}$
2018-2020



$\mu \rightarrow eee$



- **Event Signature**

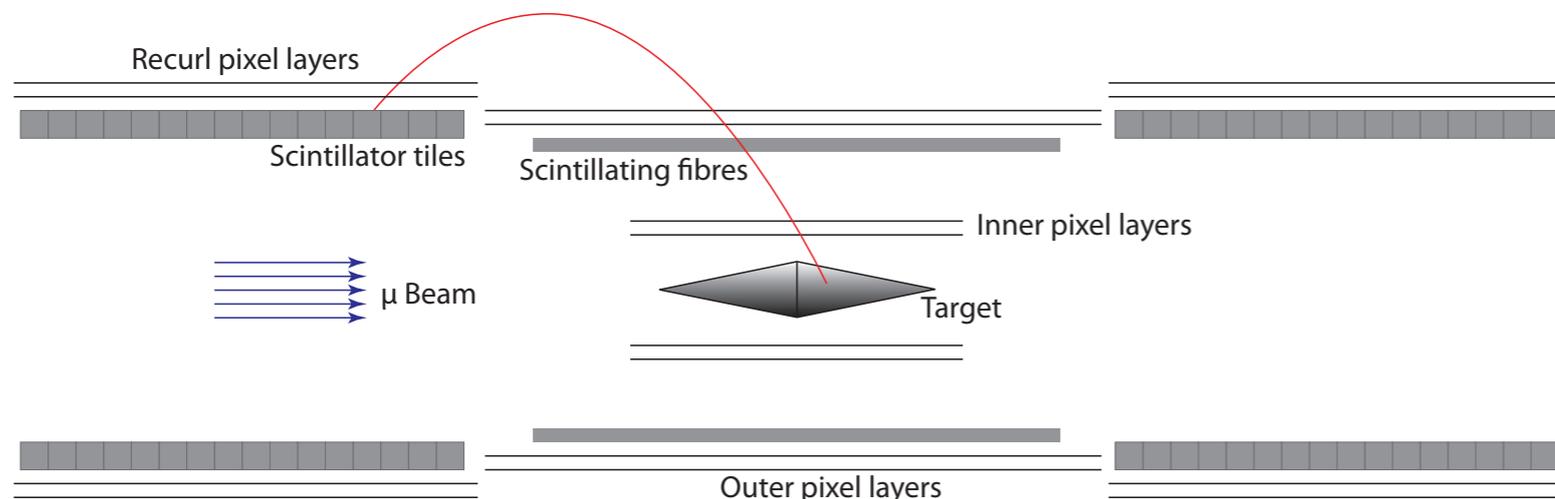
- $\Sigma E_e = m_\mu$
- $\Sigma P_e = 0$ (vector sum)
- common vertex
- time coincidence

- **Backgrounds**

- physics backgrounds
 - $\mu \rightarrow e\nu e\bar{\nu}$ decay
- accidental backgrounds

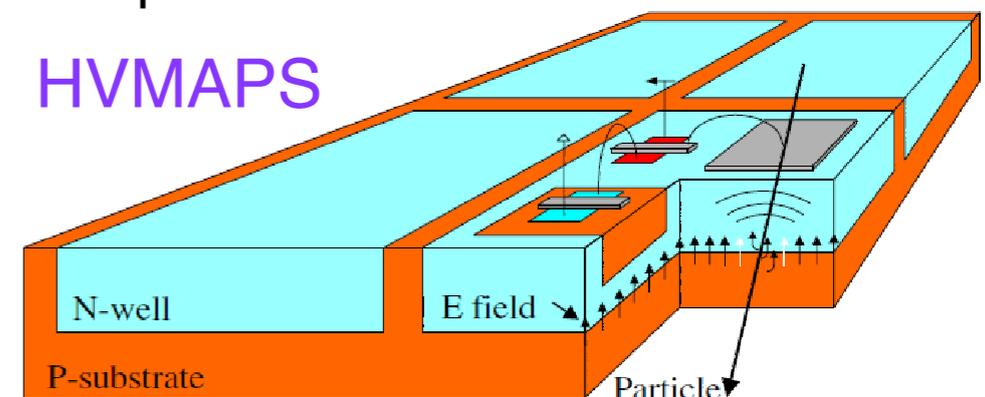
Mu3e (@PSI)

- Stage-I (2020 -)
 - $B \sim 10^{-15}$ at $\pi E5$
- Stage-2
 - $B < 10^{-16}$ at new muon source



50μm thickness

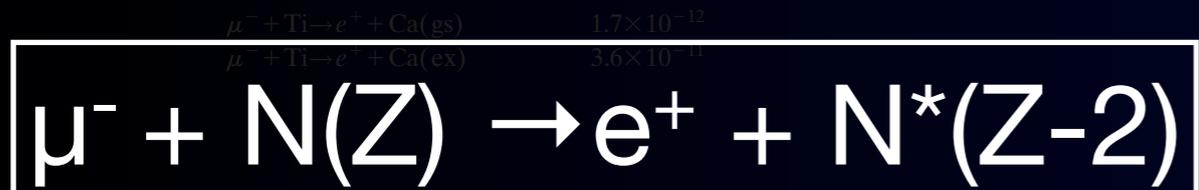
HVMAPS



μ^- to e^+ conversion



μ^- to e^+ conversion



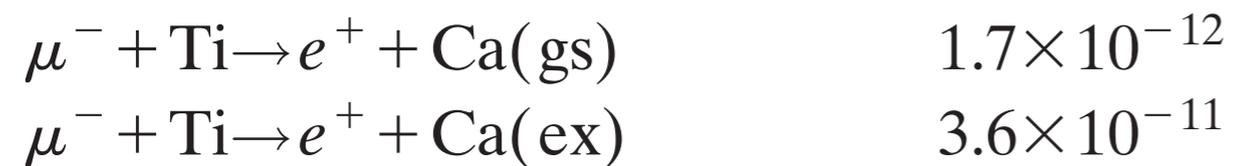
Lepton number violation (LNV)
and CLFV \Rightarrow CLNLFV

signal signature

$$E_{\mu e^+} = m_{\mu} - B_{\mu} - E_{rec} - \Delta_{Z-2}$$

backgrounds

positrons from photon conversion
after radiative muon/pion nuclear
capture



μ^- to e^+ conversion



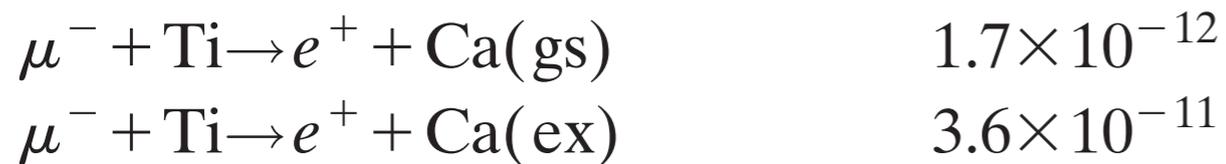
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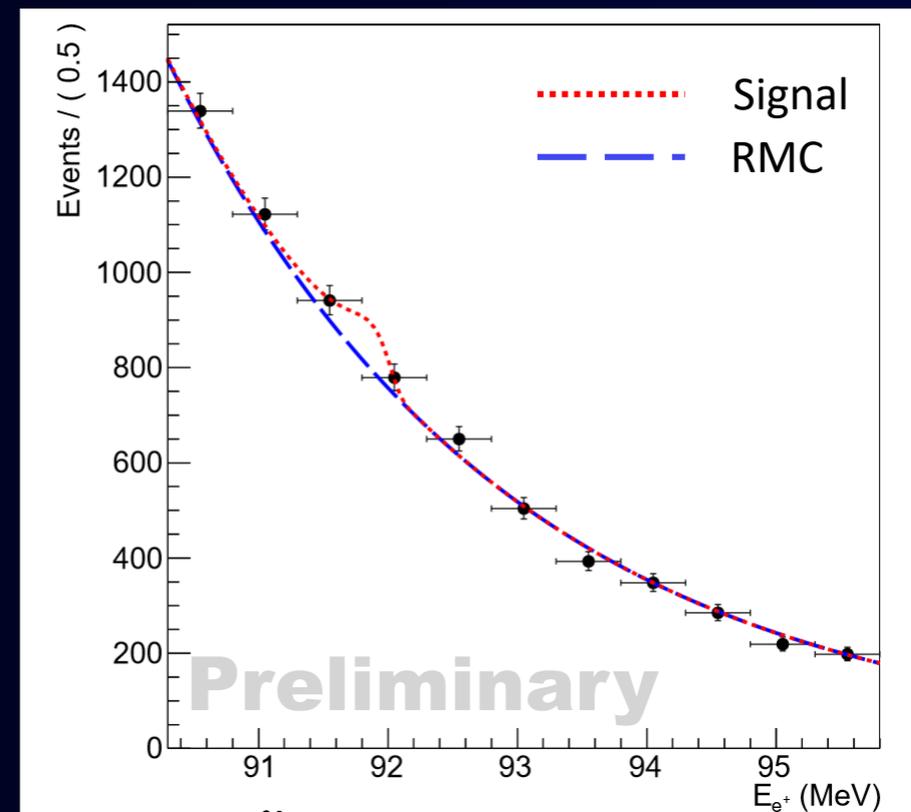
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showing that aluminum
is not a good target



μ^- to e^+ conversion



Lepton number violation (LNV)
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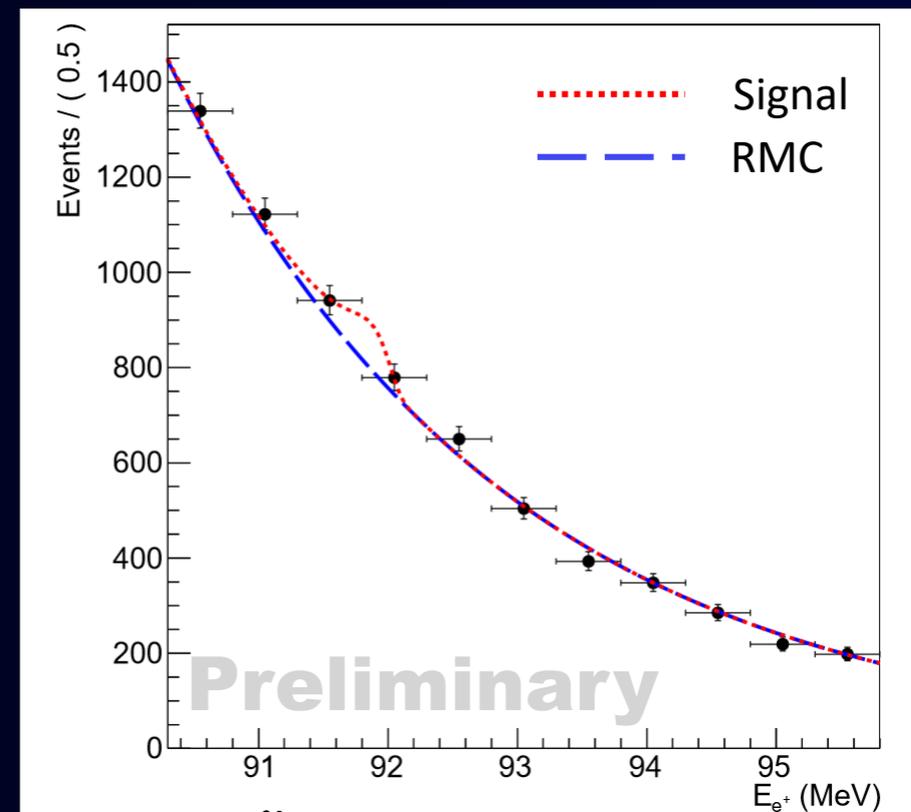
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backgrounds

positrons from photon conversion
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capture

$\mu^- + \text{Ti} \rightarrow e^+ + \text{Ca}(\text{gs})$	1.7×10^{-12}
$\mu^- + \text{Ti} \rightarrow e^+ + \text{Ca}(\text{ex})$	3.6×10^{-11}

showing that aluminum
is not a good target



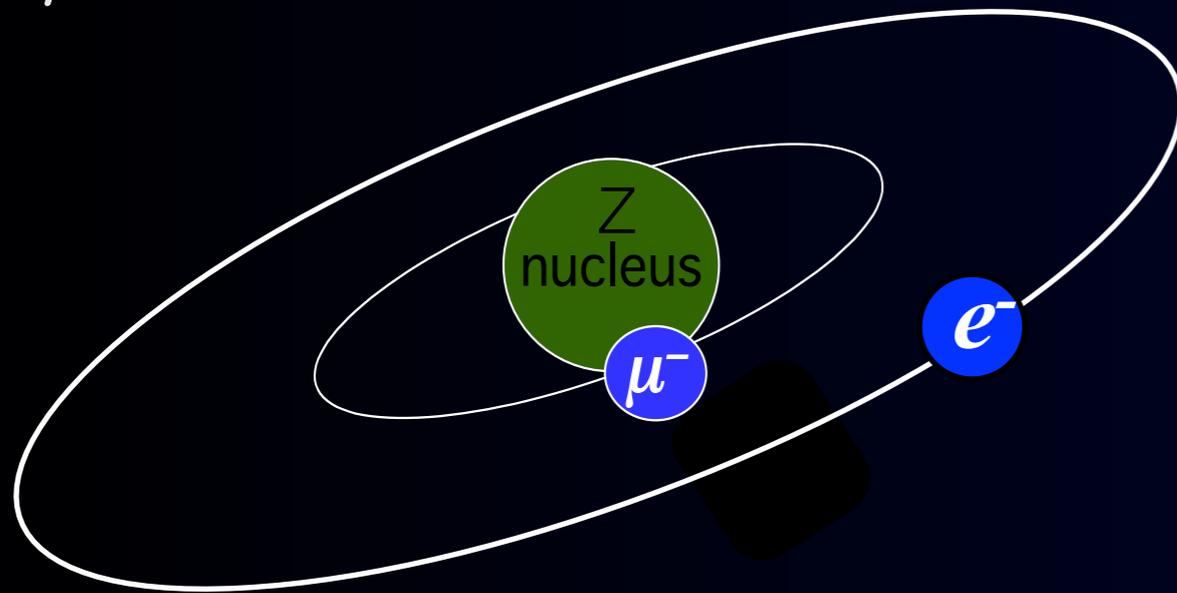
mass relation for target selection
 $M(A, Z - 1) > M(A, Z - 2)$
for $M(A, Z)$

$\mu^- + e^- \rightarrow e^- + e^-$ in a muonic atom



$\mu^- + e^- \rightarrow e^- + e^-$ in a muonic atom

$$\mu^- + e^- \rightarrow e^- + e^-$$



- $\mu^- e^- \rightarrow e^- e^-$ has the overlap of μ^- and e^- which is proportional to Z^3 .
- $\mu^- e^- \rightarrow e^- e^-$ has two-body final state.

New Process for Charged Lepton Flavor Violation Searches: $\mu^- e^- \rightarrow e^- e^-$ in a Muonic Atom

Masafumi Koike,^{1,*} Yoshitaka Kuno,^{2,†} Joe Sato,^{1,‡} and Masato Yamanaka^{3,§}

¹Physics Department, Saitama University, 255 Shimo-Okubo, Sakura-ku, Saitama, Saitama 338-8570, Japan

²Department of Physics, Osaka University, Toyonaka, Osaka 560-0043, Japan

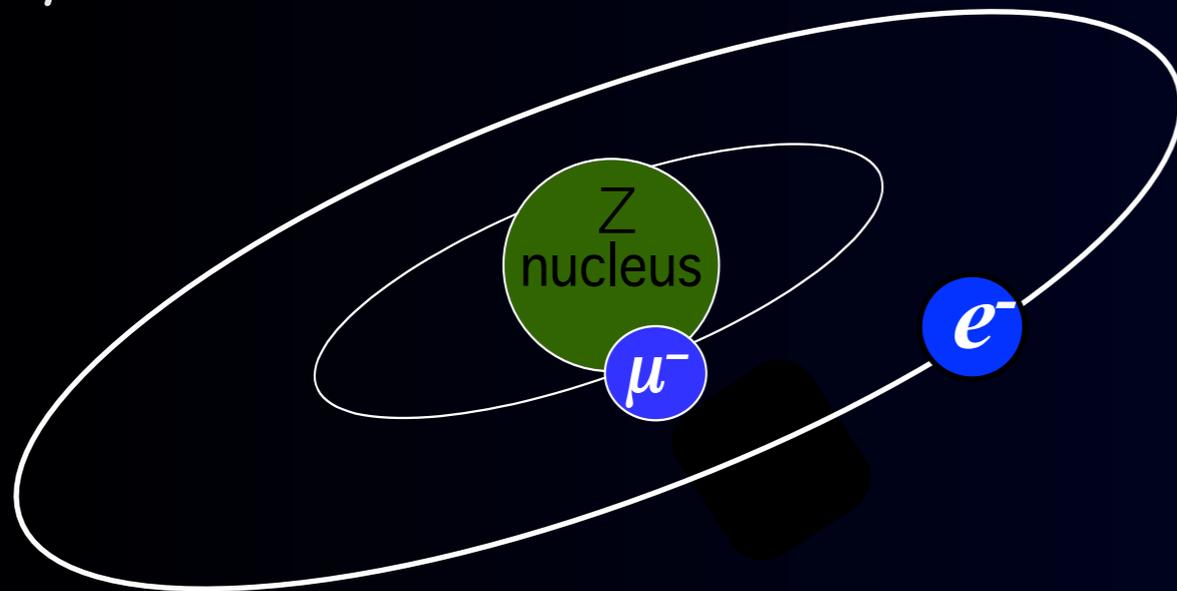
³Institute for Cosmic Ray Research, University of Tokyo, Kashiwa 277-8582, Japan

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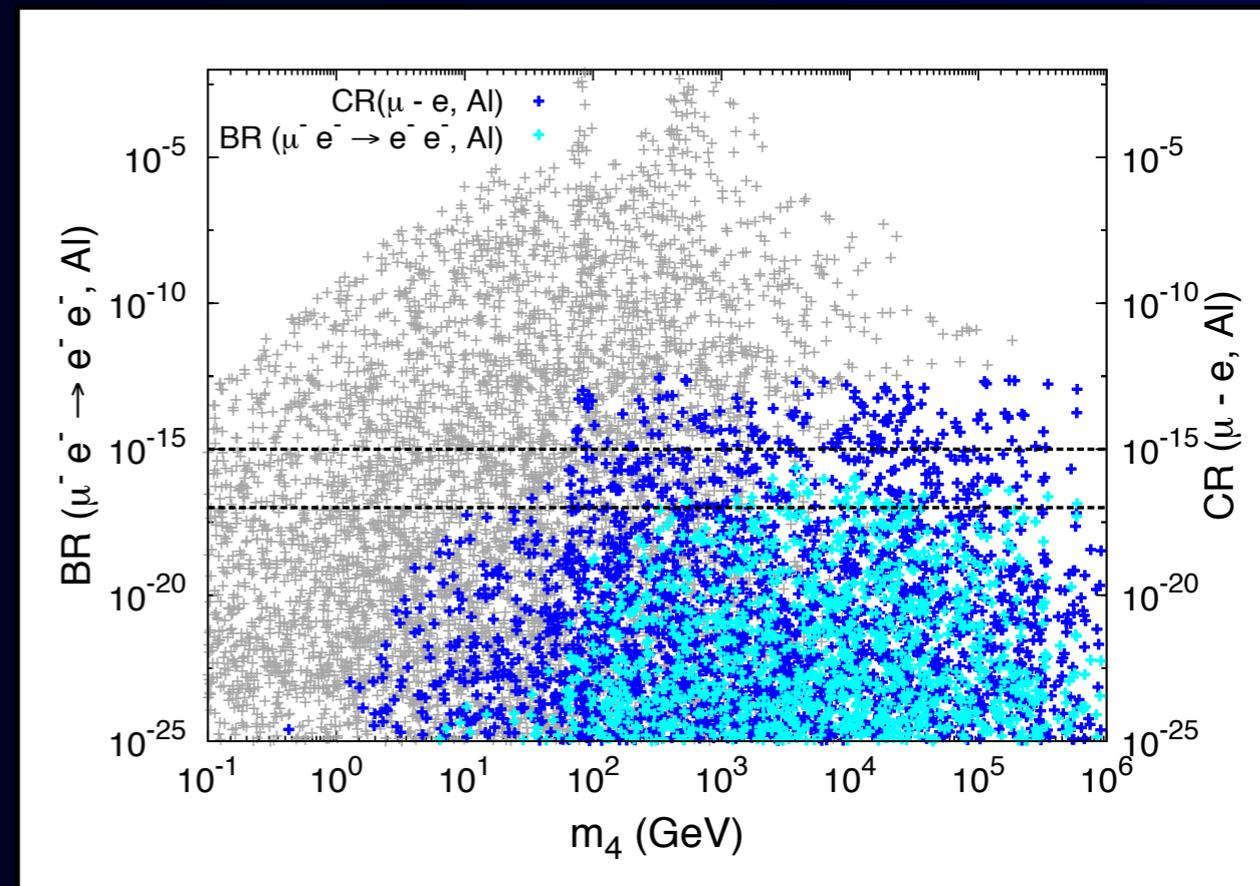
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3+1(sterile) model

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Experiments
for μ -e conversion



What is Muon to Electron Conversion?



$\mu \rightarrow e$ in vacuum

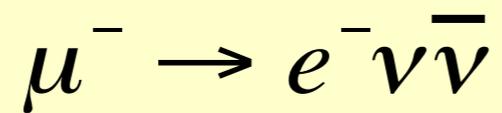
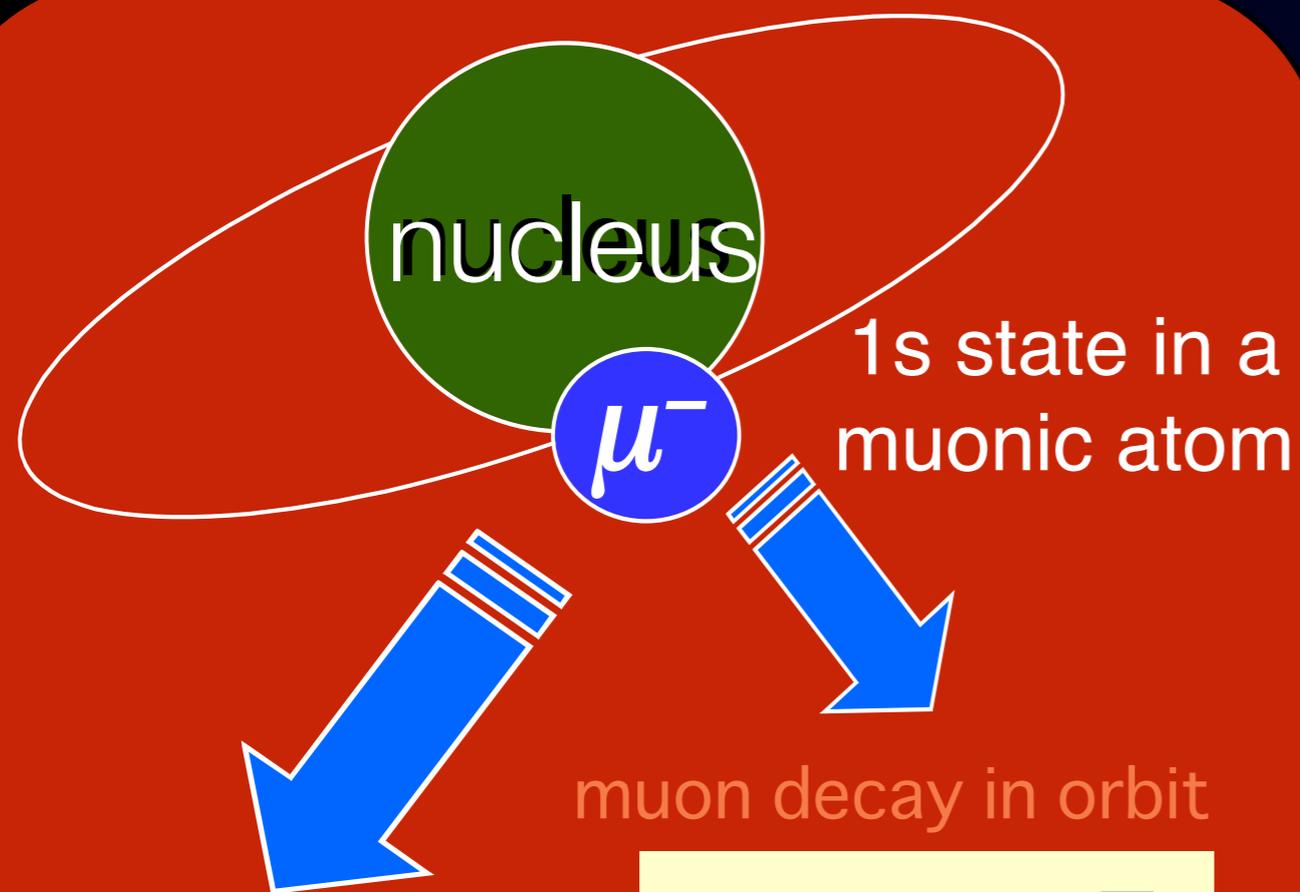
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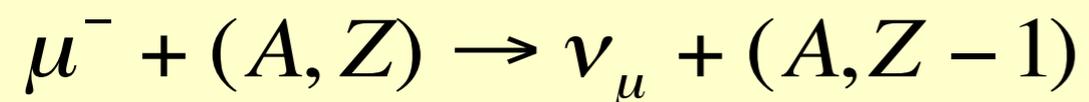
$\mu \rightarrow e$ ~~in vacuum~~ in matter

What is Muon to Electron Conversion?

$\mu \rightarrow e$ ~~in vacuum~~ in matter

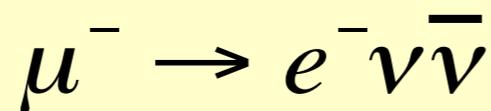
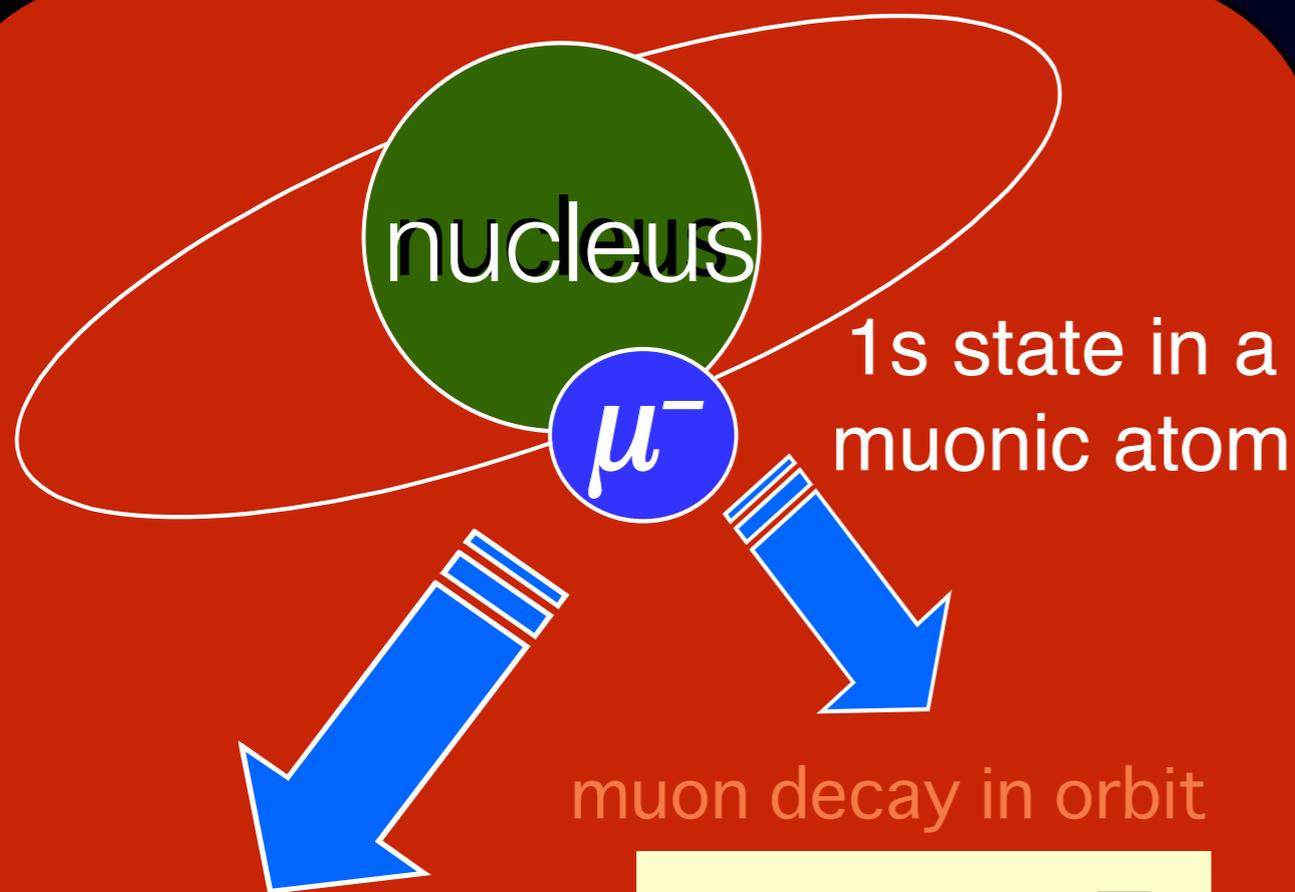


nuclear muon capture

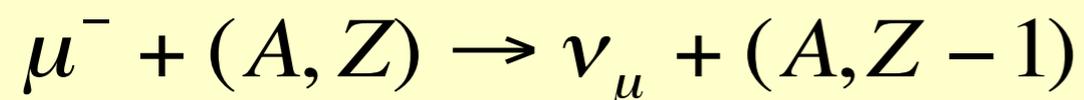


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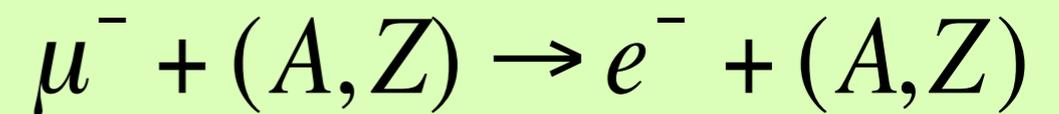
$\mu \rightarrow e$ ~~in vacuum~~ in matter



nuclear muon capture



muon to electron conversion



coherent process
to the ground state

$$\propto Z^5$$

Event Signature :

a single mono-energetic
electron of 105 MeV

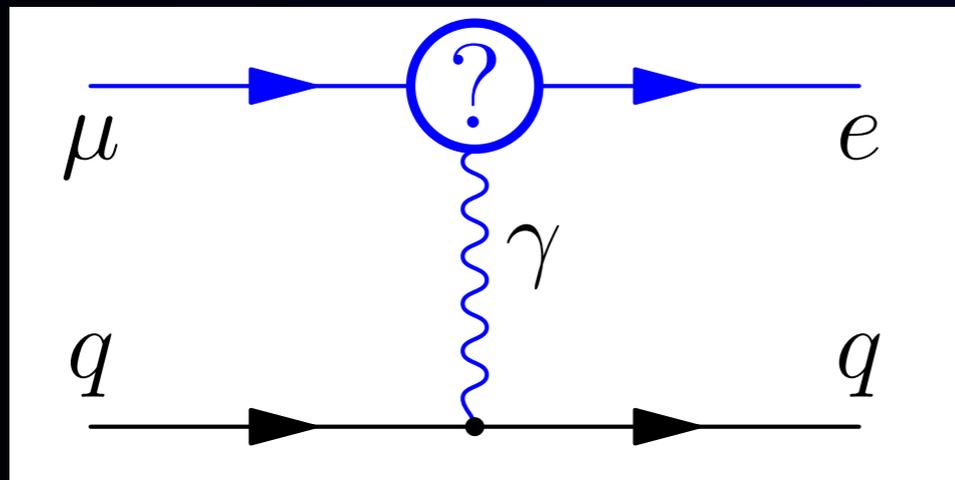
Backgrounds:

- (1) physics backgrounds
- (2) beam-related backgrounds
- (3) cosmic rays, false tracking

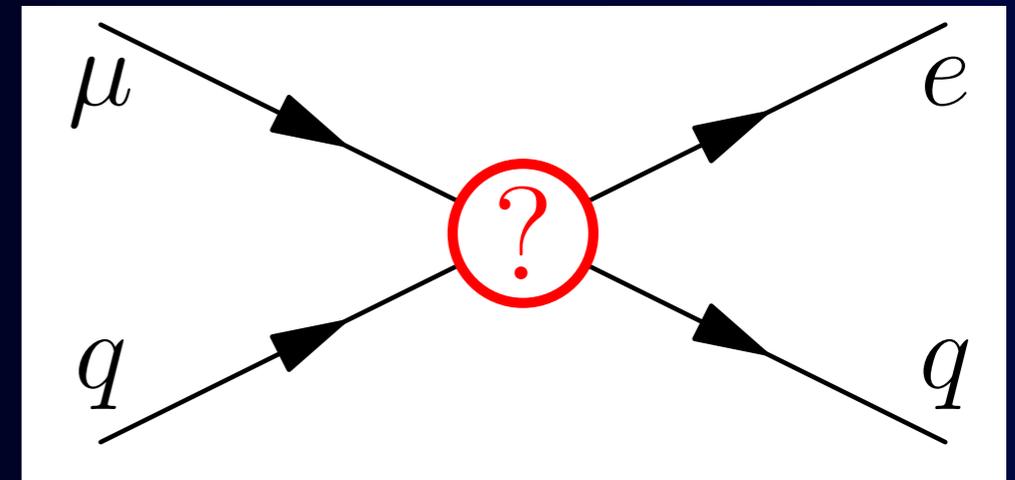
Physics Sensitivity Comparison : $\mu \rightarrow e\gamma$ vs. μ - e conversion



Photonic (dipole) interaction



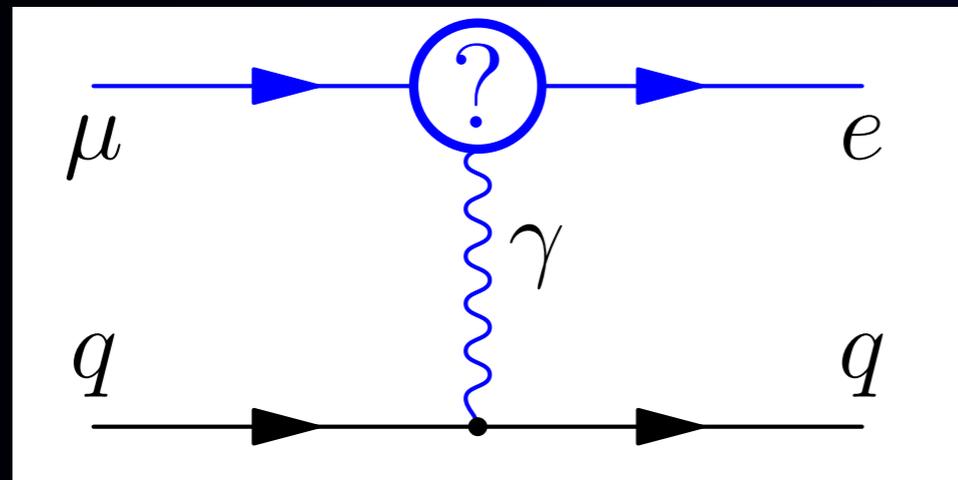
Contact interaction



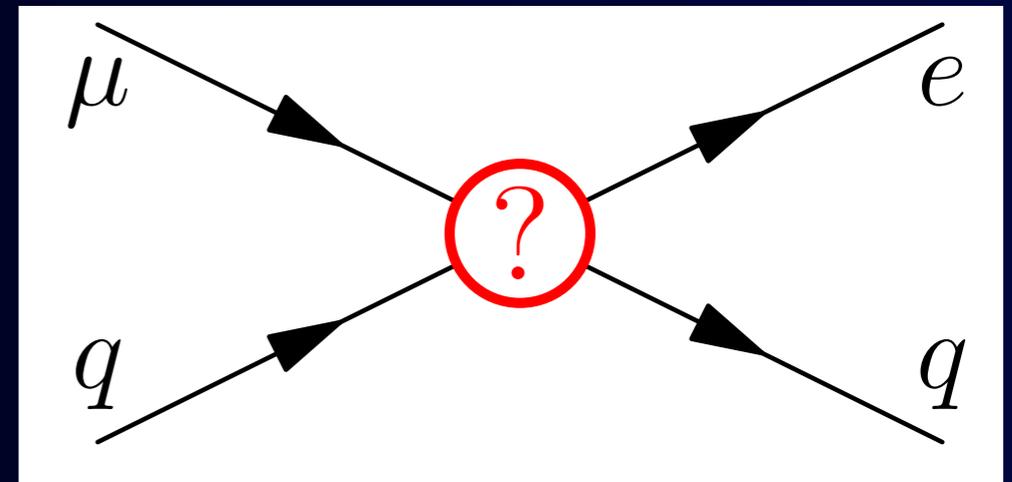
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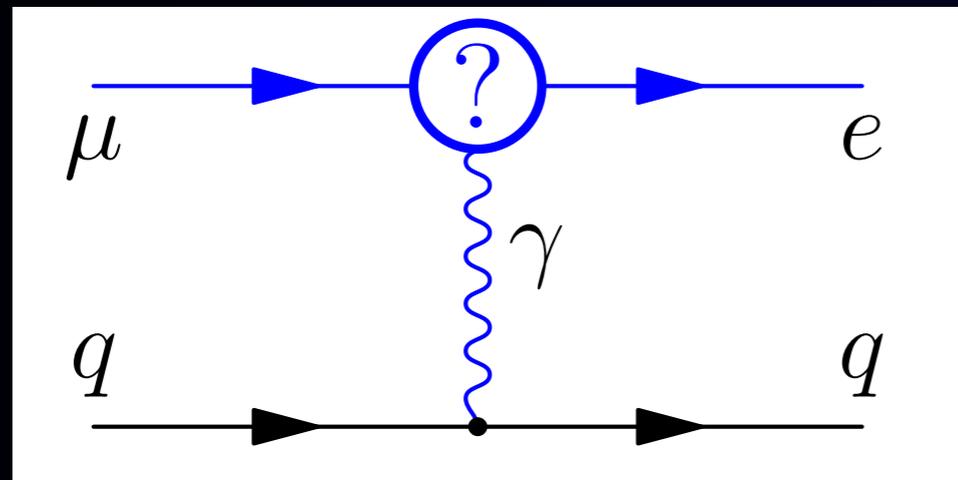
tree levels

$$L_{\mu N \rightarrow e N} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu} + \frac{\kappa}{1 + \kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L)$$

Physics Sensitivity Comparison : $\mu \rightarrow e\gamma$ vs. μ - e conversion

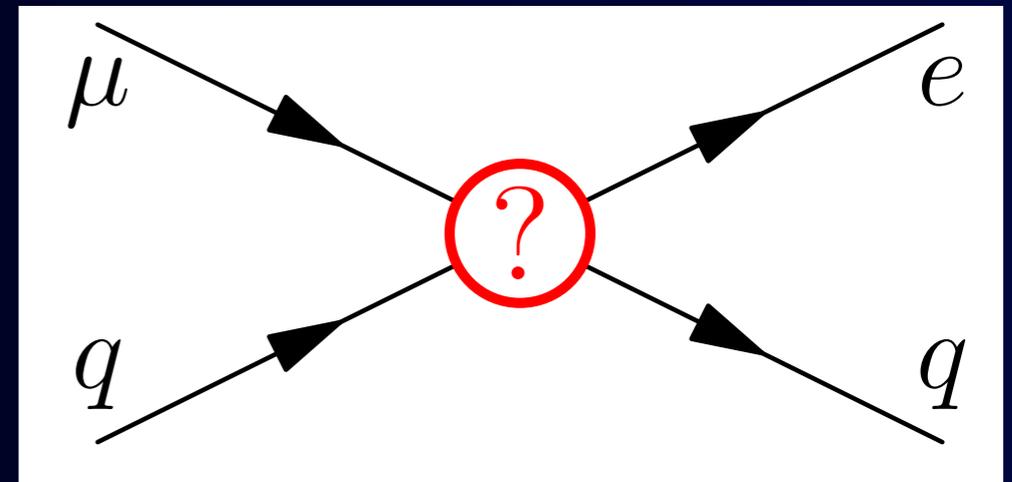


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$$L_{\mu N \rightarrow e N} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu} + \frac{\kappa}{1 + \kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L)$$

Contact interaction



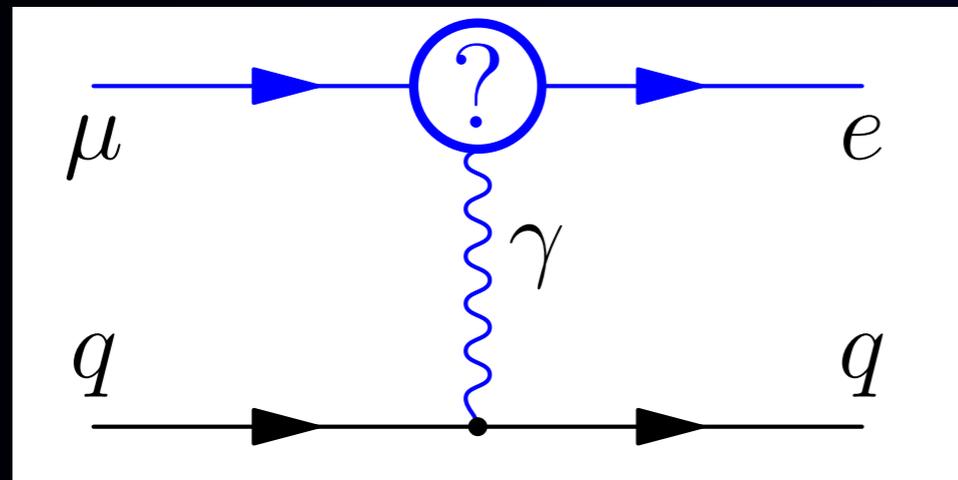
tree levels

$$L_{\mu \rightarrow e\gamma} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu}$$

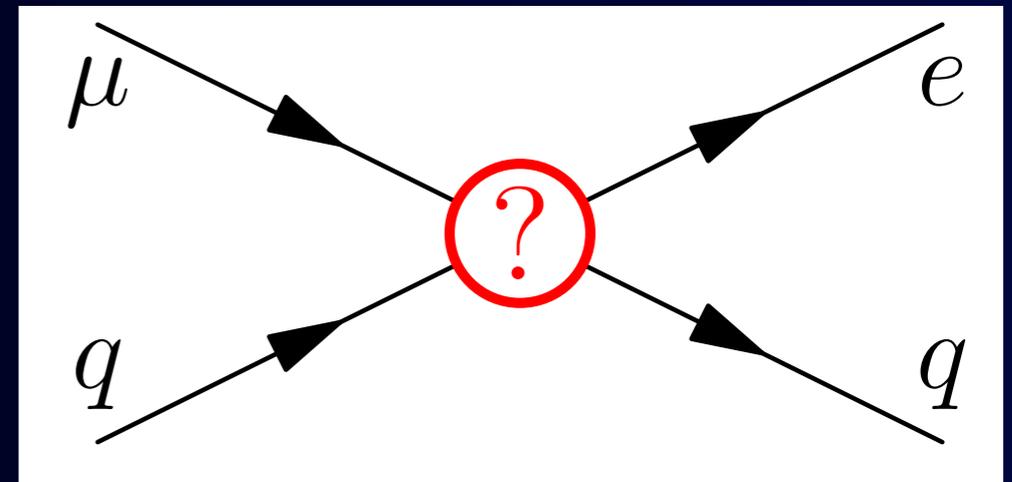
Physics Sensitivity Comparison : $\mu \rightarrow e\gamma$ vs. μ -e conversion



Photonic (dipole) interaction



Contact interaction



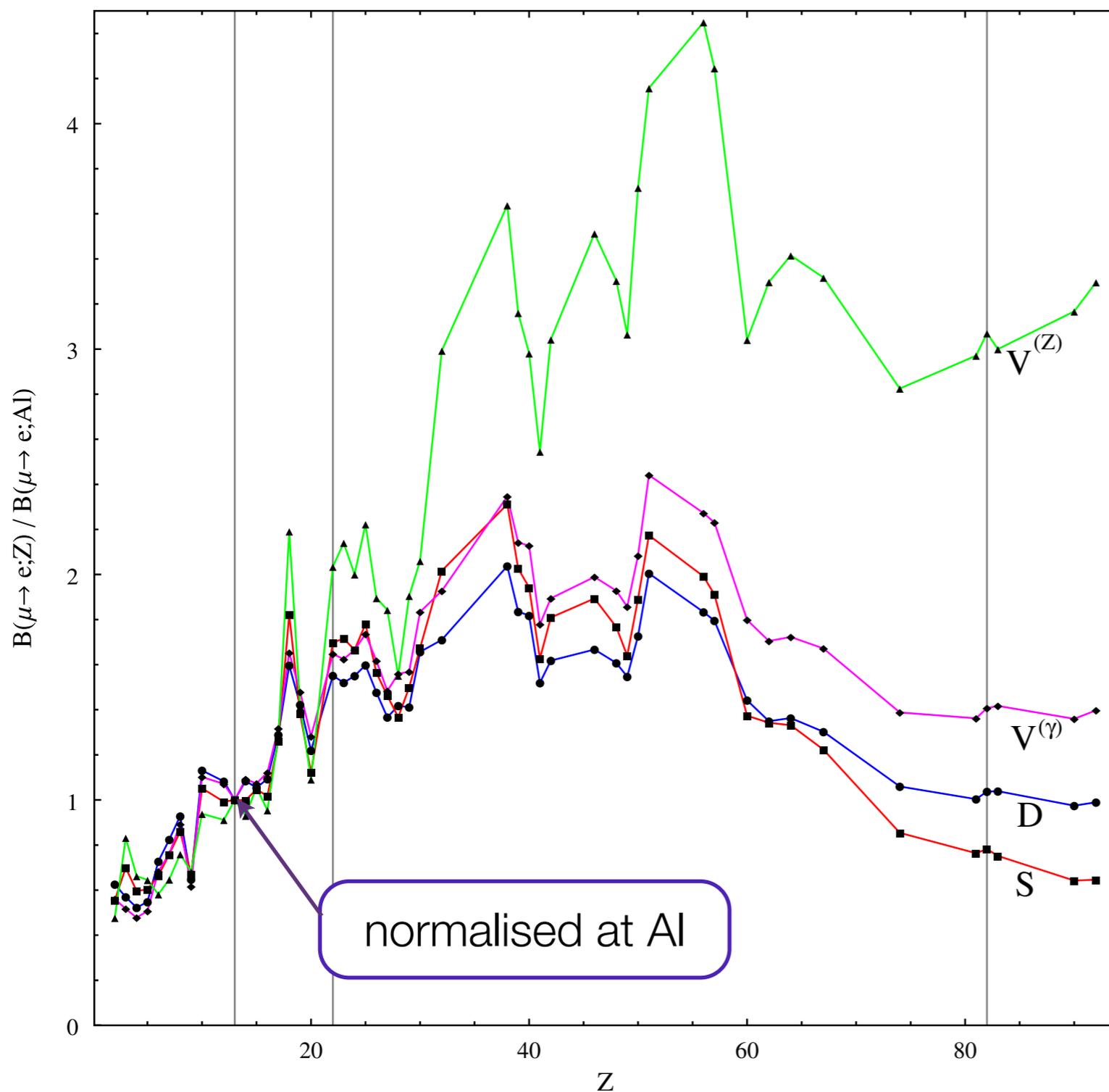
tree levels

$$L_{\mu N \rightarrow e N} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu} + \frac{\kappa}{1 + \kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L)$$

$$L_{\mu \rightarrow e\gamma} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu}$$

μ -e conversion sensitive to more new physics

μ -e Conversion : Target dependence (discriminating effective interaction)



V. Cirigliano, R. Kitano, Y. Okada,
and P. Tuzon, Phys. Rev. D80,
013002 (2009)

vector interaction
(with Z boson)

with Z
penguin

vector interaction
(with photon -
charge radius)

left-right
models

dipole interaction

SUSY-
GUT

scalar interaction

SUSY
seesaw

CLFV Effective Interactions



CLFV Effective Interactions



Dipole interaction

Contact interaction

dipole
interaction

vector
interaction

scalar
interaction

axial vector
interaction

tensor
interaction

CLFV Effective Interactions



Dipole interaction

Contact interaction

dipole
interaction

vector
interaction

scalar
interaction

Spin-independent
 μ -e Conversion
(coherent)

axial vector
interaction

tensor
interaction

Spin-dependent
 μ -e Conversion
(incoherent)

Spin dependent μ - e conversion



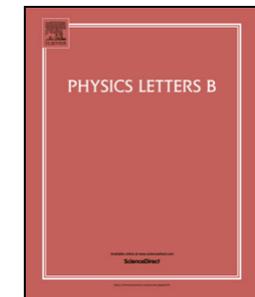
Physics Letters B 771 (2017) 242–246



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Spin-dependent $\mu \rightarrow e$ conversion



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ABSTRACT

The experimental sensitivity to $\mu \rightarrow e$ conversion on nuclei is expected to improve by four orders of magnitude in coming years. We consider the impact of $\mu \rightarrow e$ flavour-changing tensor and axial-vector four-fermion operators which couple to the spin of nucleons. Such operators, which have not previously been considered, contribute to $\mu \rightarrow e$ conversion in three ways: in nuclei with spin they mediate a spin-dependent transition; in all nuclei they contribute to the coherent (A^2 -enhanced) spin-independent conversion via finite recoil effects and via loop mixing with dipole, scalar, and vector operators. We estimate the spin-dependent rate in Aluminium (the target of the upcoming COMET and Mu2e experiments), show that the loop effects give the greatest sensitivity to tensor and axial-vector operators involving first-generation quarks, and discuss the complementarity of the spin-dependent and independent contributions to $\mu \rightarrow e$ conversion.

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Experimental Comparison : $\mu \rightarrow e\gamma$ and μ -e Conversion



Experimental Comparison : $\mu \rightarrow e\gamma$ and μ -e Conversion



	Beam	background	challenge	beam intensity
$\mu \rightarrow e\gamma$	continuous beam	accidentals	detector resolution	limited
$\mu \rightarrow eee$	continuous beam	accidentals	detector resolution	limited
μ -e conversion	pulsed beam	beam-related	beam background	no limitation

Experimental Comparison : $\mu \rightarrow e\gamma$ and μ -e Conversion



	Beam	background	challenge	beam intensity
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$\mu \rightarrow eee$	continuous beam	accidentals	detector resolution	limited
μ -e conversion	pulsed beam	beam-related	beam background	no limitation

μ -e conversion can go to higher sensitivity

Backgrounds for μ -e conversion



Backgrounds for μ -e conversion



intrinsic physics
backgrounds

Muon decay in orbit (DIO)
Radiative muon capture (RMC)
neutrons from muon nuclear capture
Protons from muon nuclear capture

Backgrounds for μ -e conversion



intrinsic physics
backgrounds

Muon decay in orbit (DIO)
Radiative muon capture (RMC)
neutrons from muon nuclear capture
Protons from muon nuclear capture

beam-related
backgrounds

Radiative pion capture (RPC)
Beam electrons
Muon decay in flights
Neutron background
Antiproton induced background

Backgrounds for μ -e conversion



intrinsic physics
backgrounds

Muon decay in orbit (DIO)
Radiative muon capture (RMC)
neutrons from muon nuclear capture
Protons from muon nuclear capture

beam-related
backgrounds

Radiative pion capture (RPC)
Beam electrons
Muon decay in flights
Neutron background
Antiproton induced background

cosmic-ray and other
backgrounds

Cosmic-ray induced background
False tracking

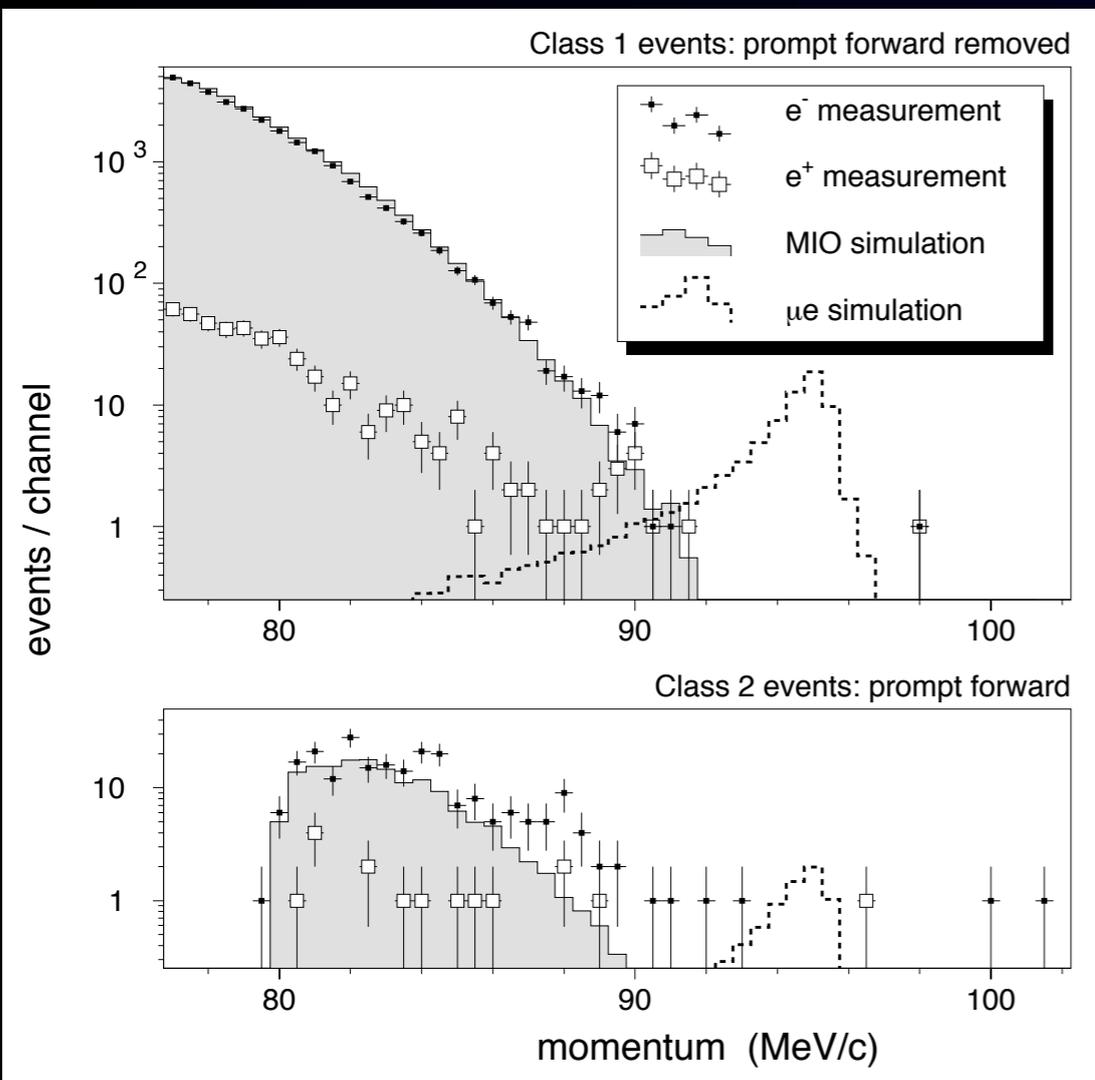
Previous and Future Measurements for Muon to Electron Conversion



Previous and Future Measurements for Muon to Electron Conversion



SINDRUM-II (PSI)



Published Results (2004)

$$B(\mu^- + Au \rightarrow e^- + Au) < 7 \times 10^{-13}$$

Previous and Future Measurements for Muon to Electron Conversion

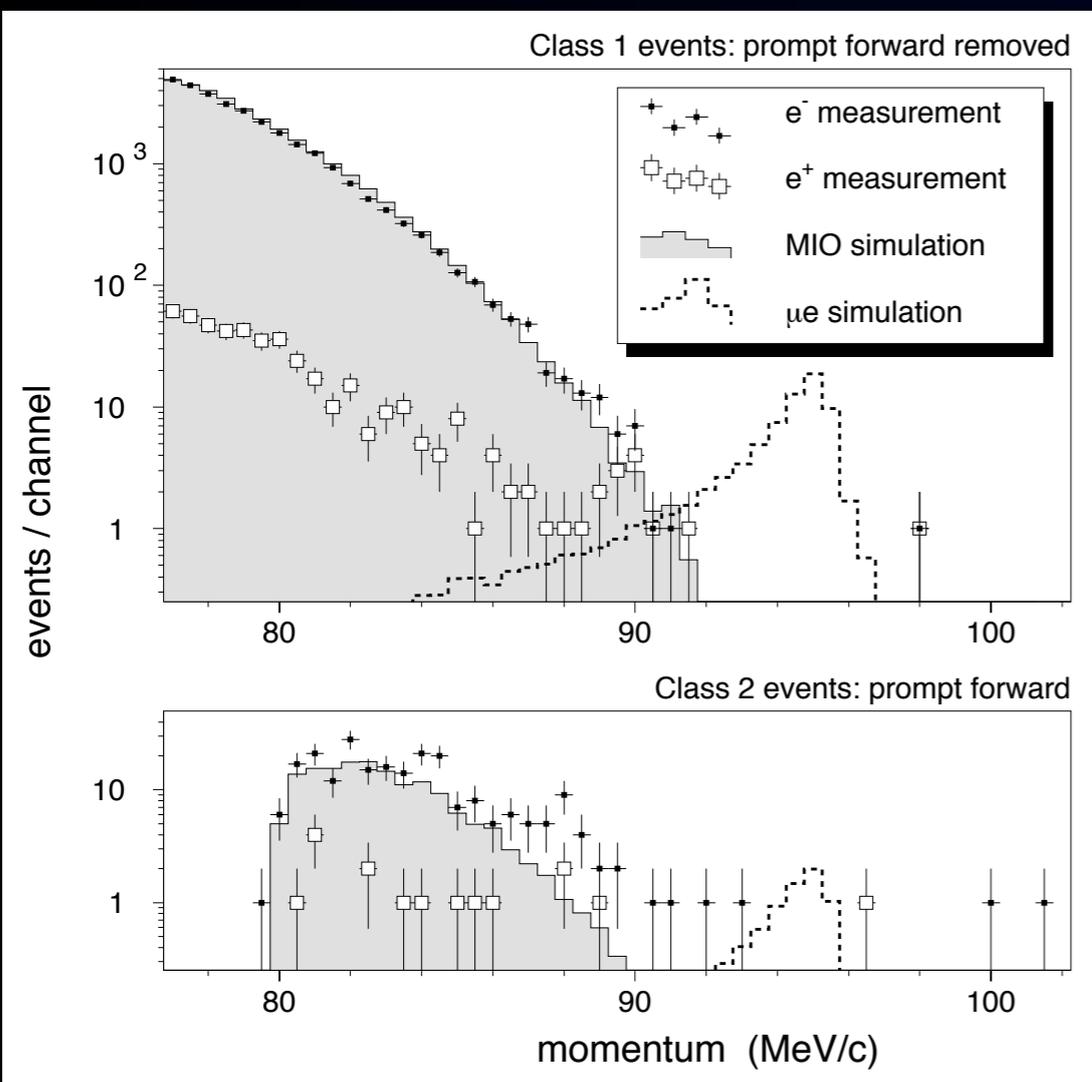


SINDRUM-II (PSI)

Future Measurements

A total number of muons is the key for success.

COMET : 10^{18} muons
(past exp. 10^{14} muons)



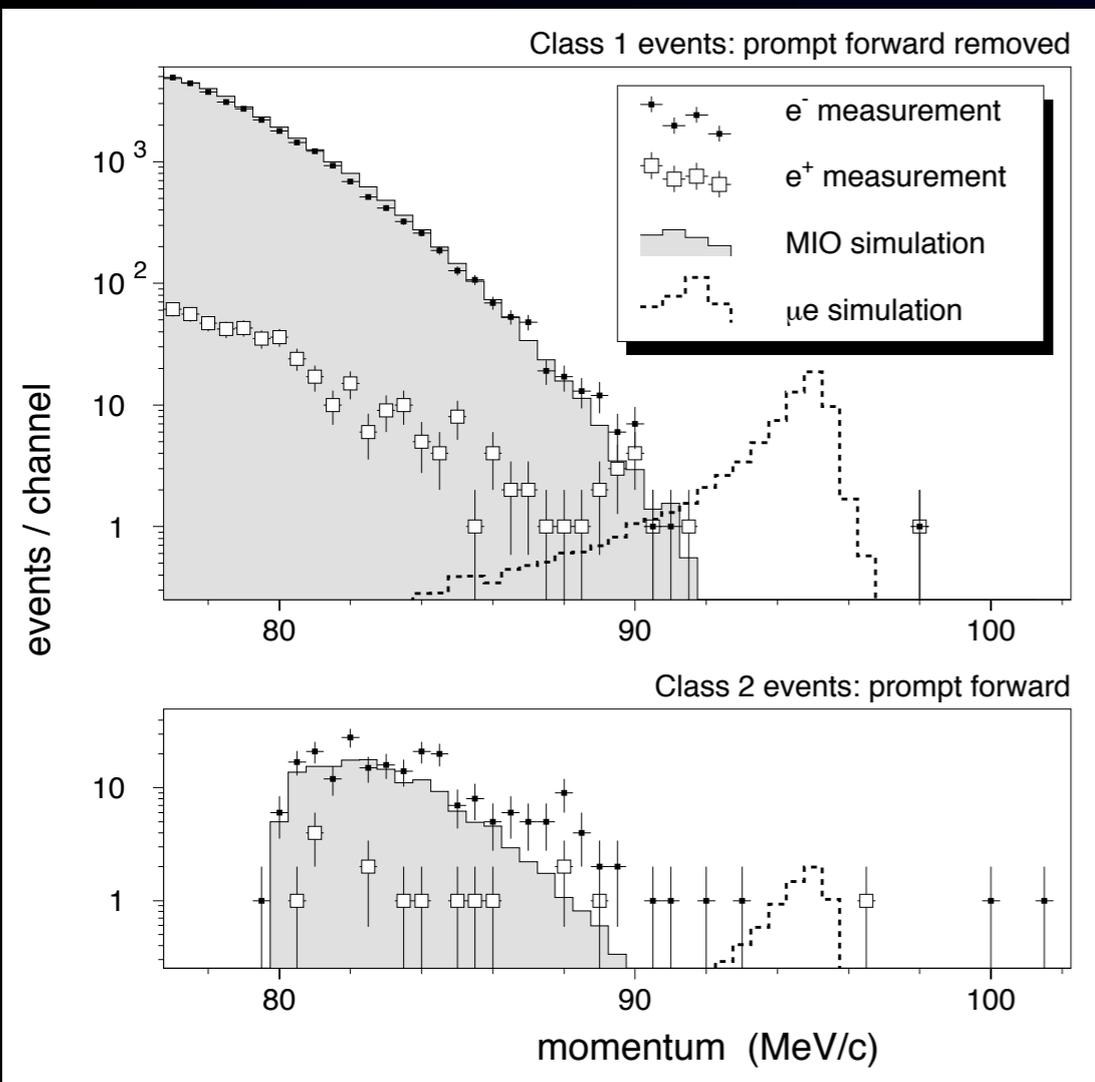
Published Results (2004)

$$B(\mu^- + Au \rightarrow e^- + Au) < 7 \times 10^{-13}$$

Previous and Future Measurements for Muon to Electron Conversion



SINDRUM-II (PSI)



Published Results (2004)

$$B(\mu^- + Au \rightarrow e^- + Au) < 7 \times 10^{-13}$$

Future Measurements

A total number of muons is the key for success.

COMET : 10^{18} muons
(past exp. 10^{14} muons)

1000 years needed at
PSI with intensity of
 10^{15} muons/year

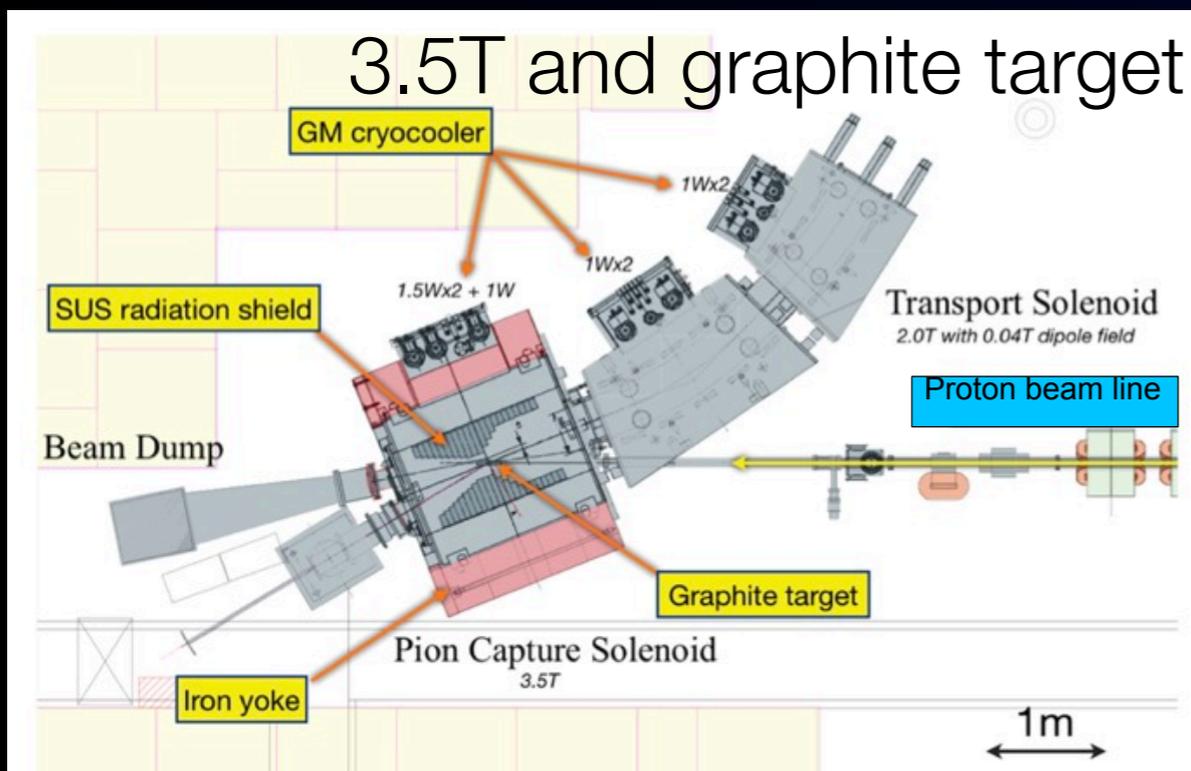
MuSIC at RCNP, Osaka University - Highly Intense Muon Source -



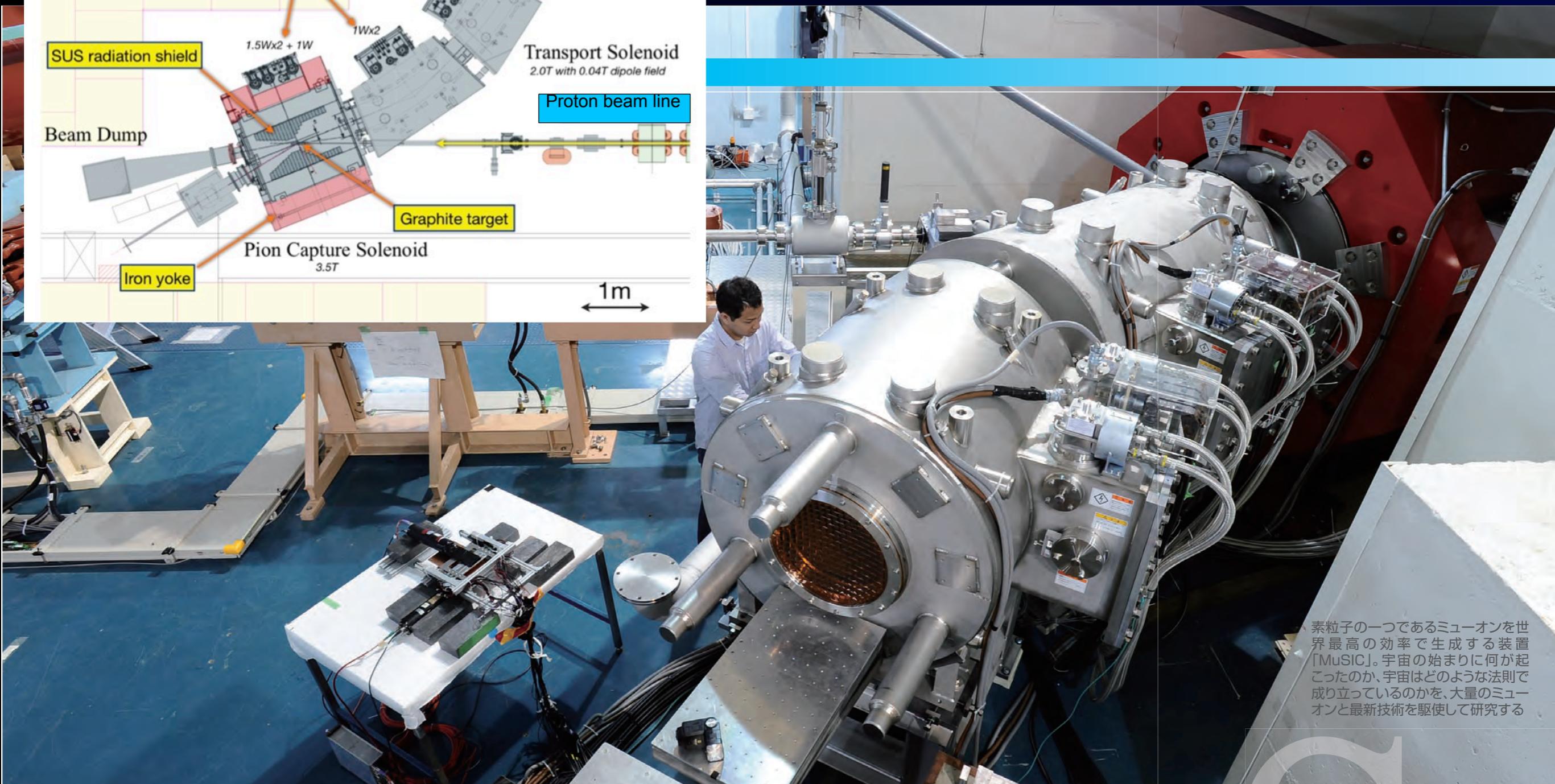
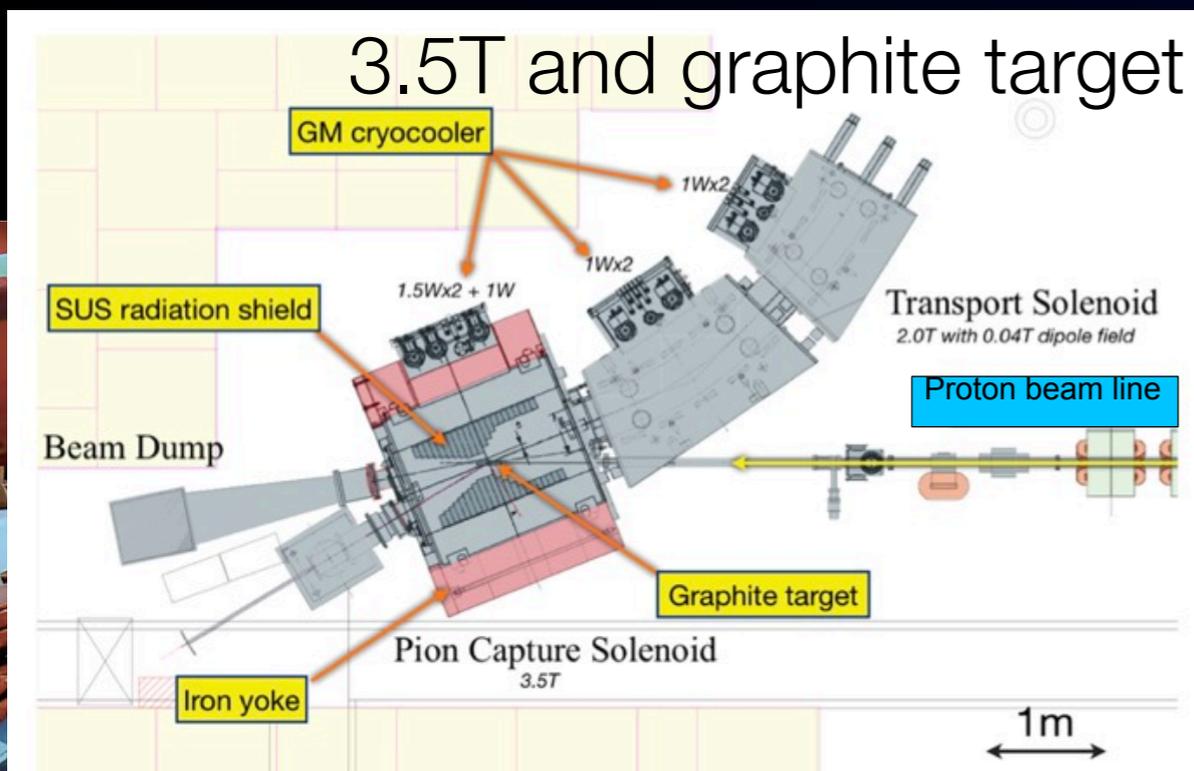
3.5T and graphite target

MuSIC at RCNP, Osaka University

- Highly Intense Muon Source -

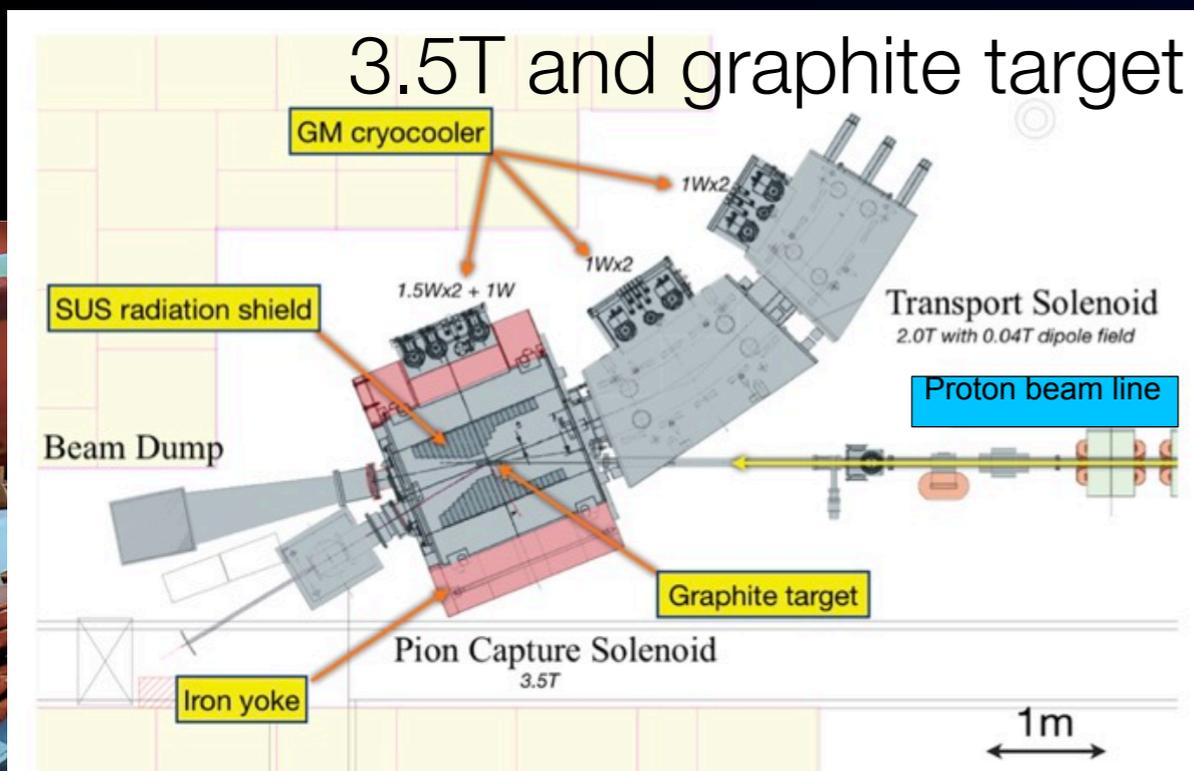


MuSIC at RCNP, Osaka University - Highly Intense Muon Source -

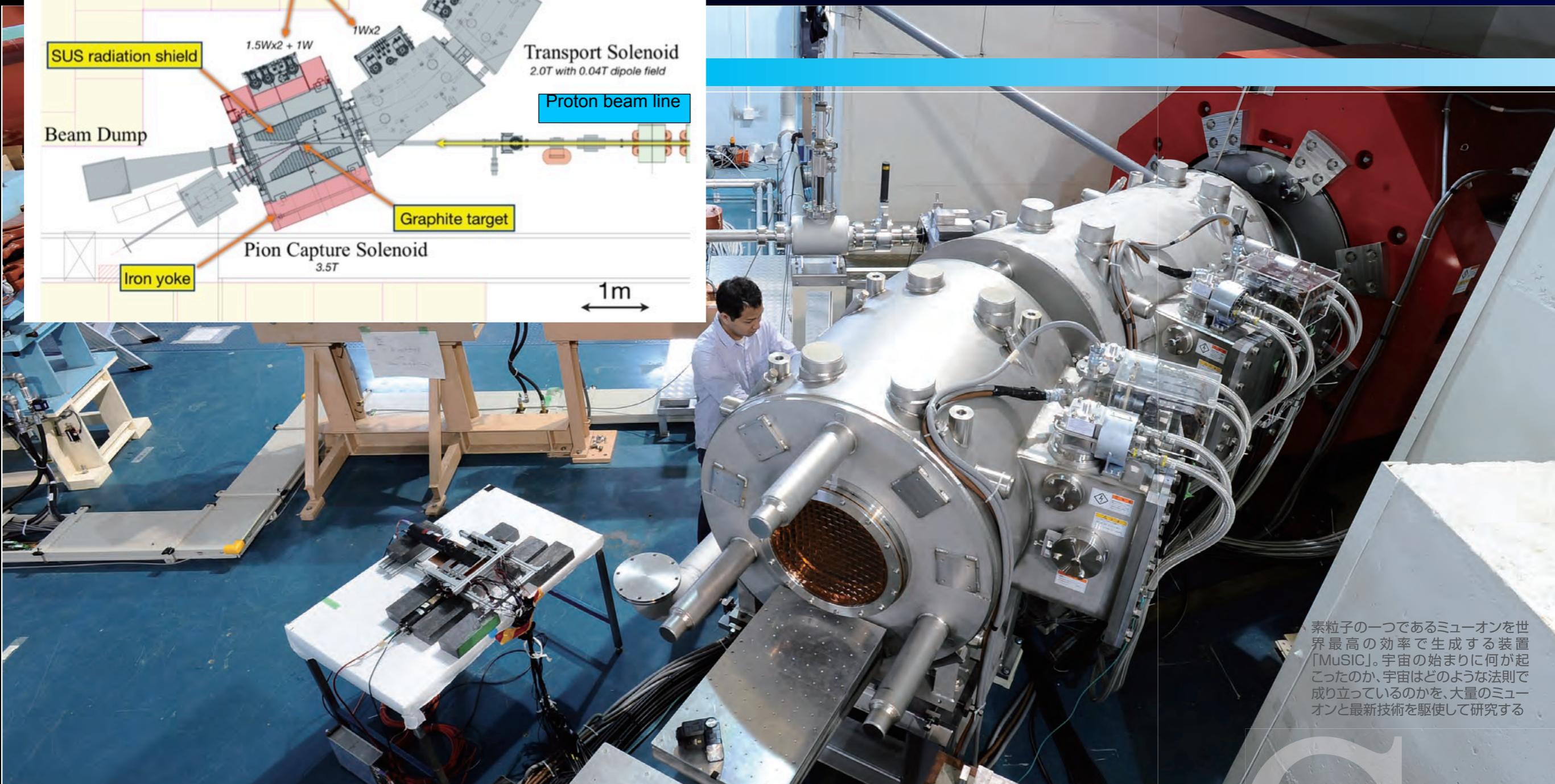


素粒子の一つであるミューオンを世界最高の効率で生成する装置「MuSIC」。宇宙の始まりに何が起こったのか、宇宙はどのような法則で成り立っているのかを、大量のミューオンと最新技術を駆使して研究する

MuSIC at RCNP, Osaka University - Highly Intense Muon Source -



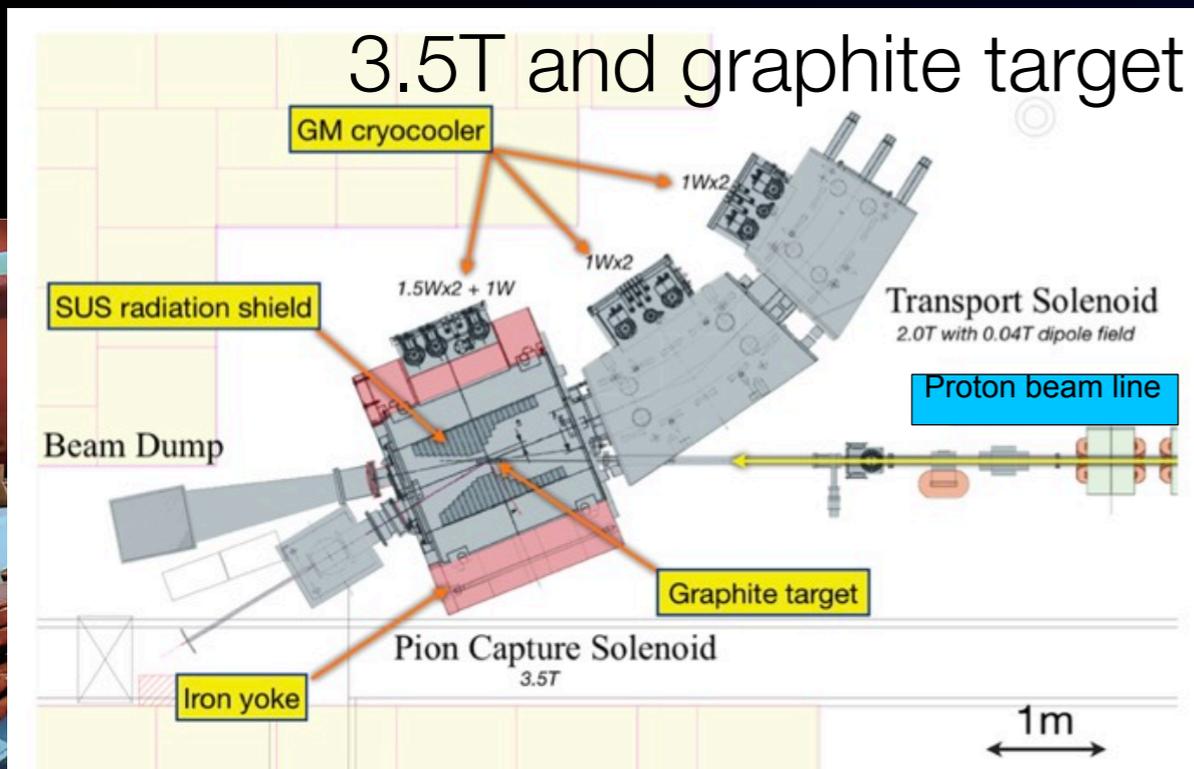
Muon Science Intense Channel (>2011)



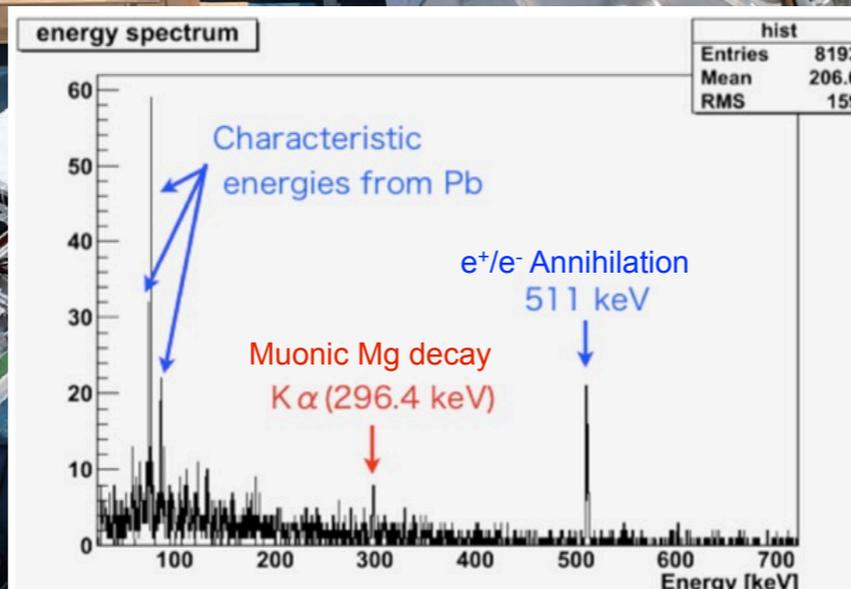
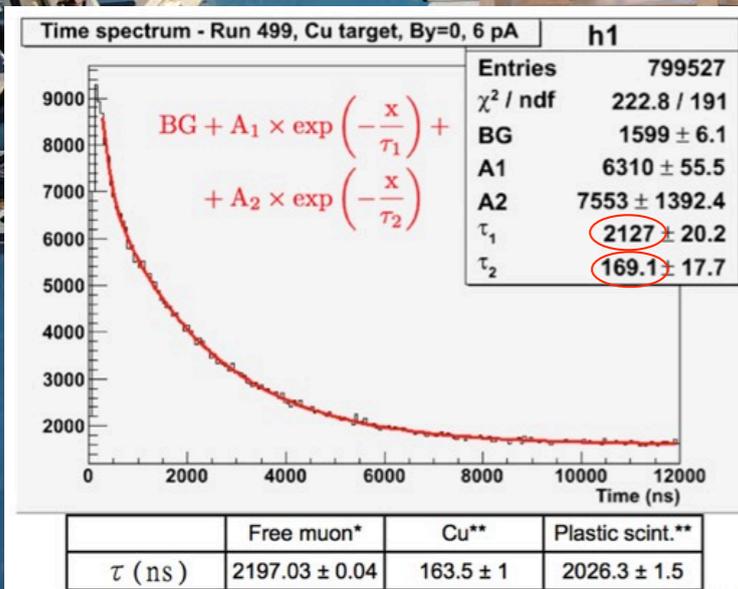
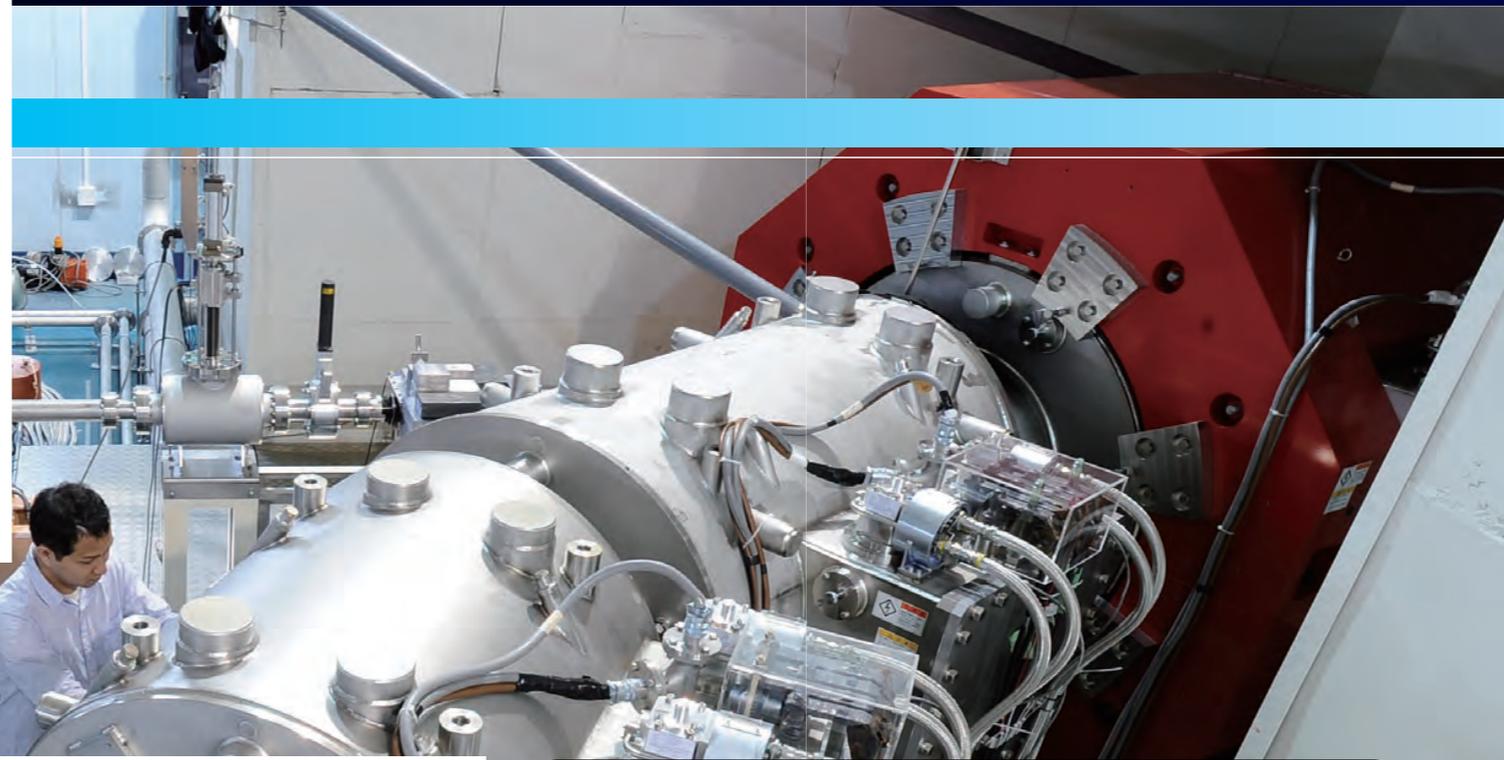
素粒子の一つであるミューオンを世界最高の効率で生成する装置「MuSIC」。宇宙の始まりに何が起こったのか、宇宙はどのような法則で成り立っているのかを、大量のミューオンと最新技術を駆使して研究する

MuSIC at RCNP, Osaka University

- Highly Intense Muon Source -



Muon Science Intense Channel (>2011)



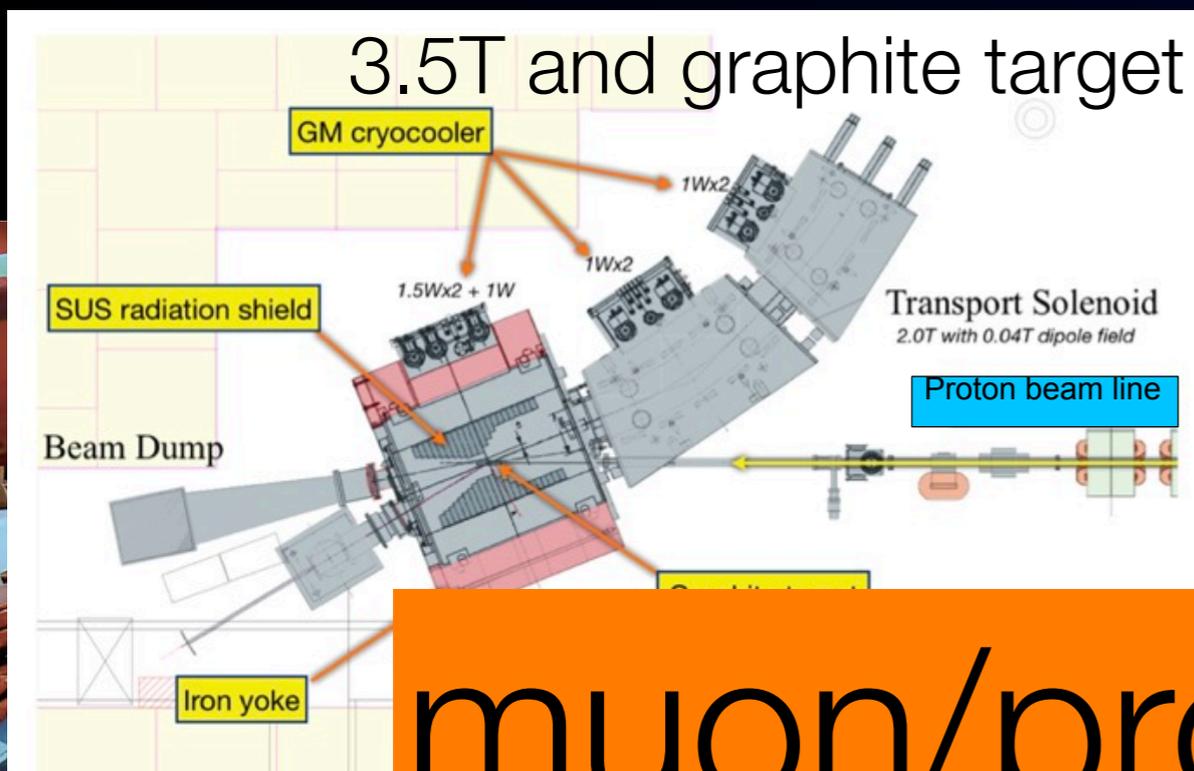
MuSIC muon yields

μ^+ : $3 \times 10^8 / \text{s}$ for 400W

μ^- : $1 \times 10^8 / \text{s}$ for 400W

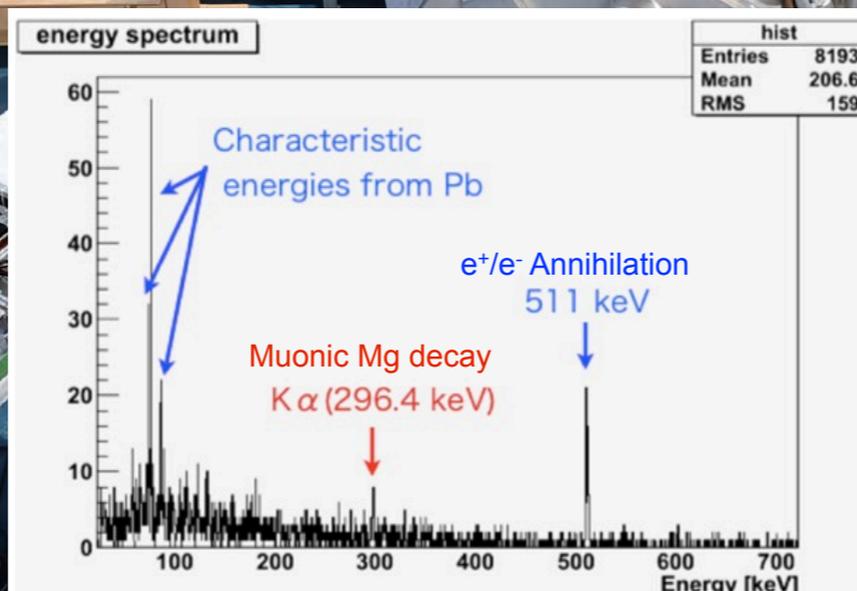
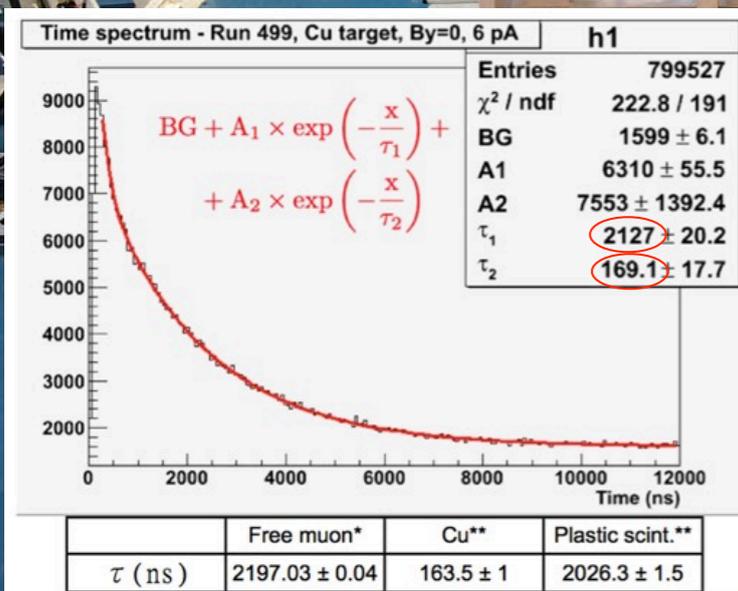
MuSIC at RCNP, Osaka University

- Highly Intense Muon Source -



Muon Science Intense Channel (>2011)

muon/proton ~ x1000



MuSIC muon yields

μ^+ : $3 \times 10^8 / \text{s}$ for 400W

μ^- : $1 \times 10^8 / \text{s}$ for 400W

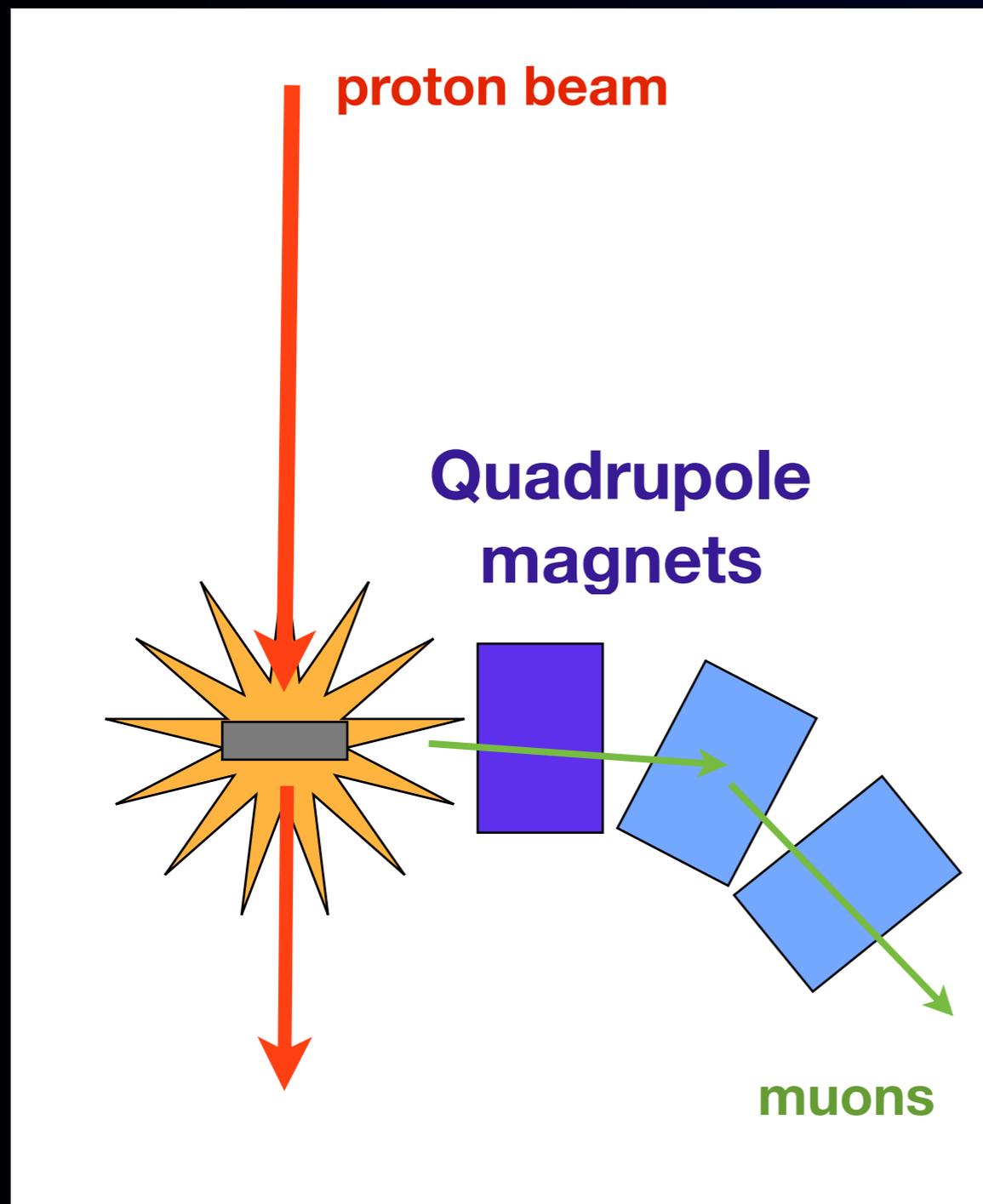
Production and Collection of Pions/Muons in Solenoidal Magnetic Fields



Production and Collection of Pions/Muons in Solenoidal Magnetic Fields



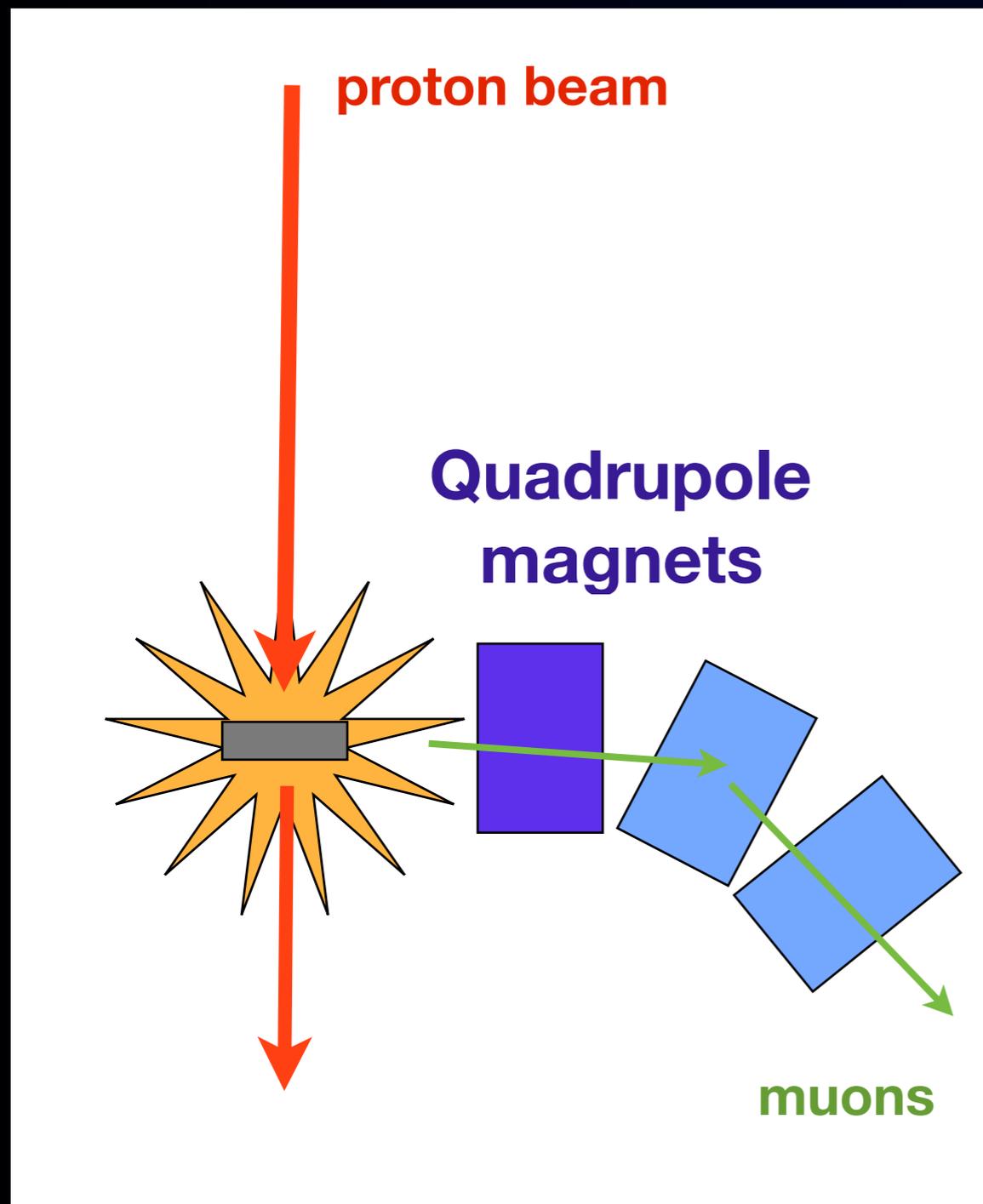
Conventional muon beam source



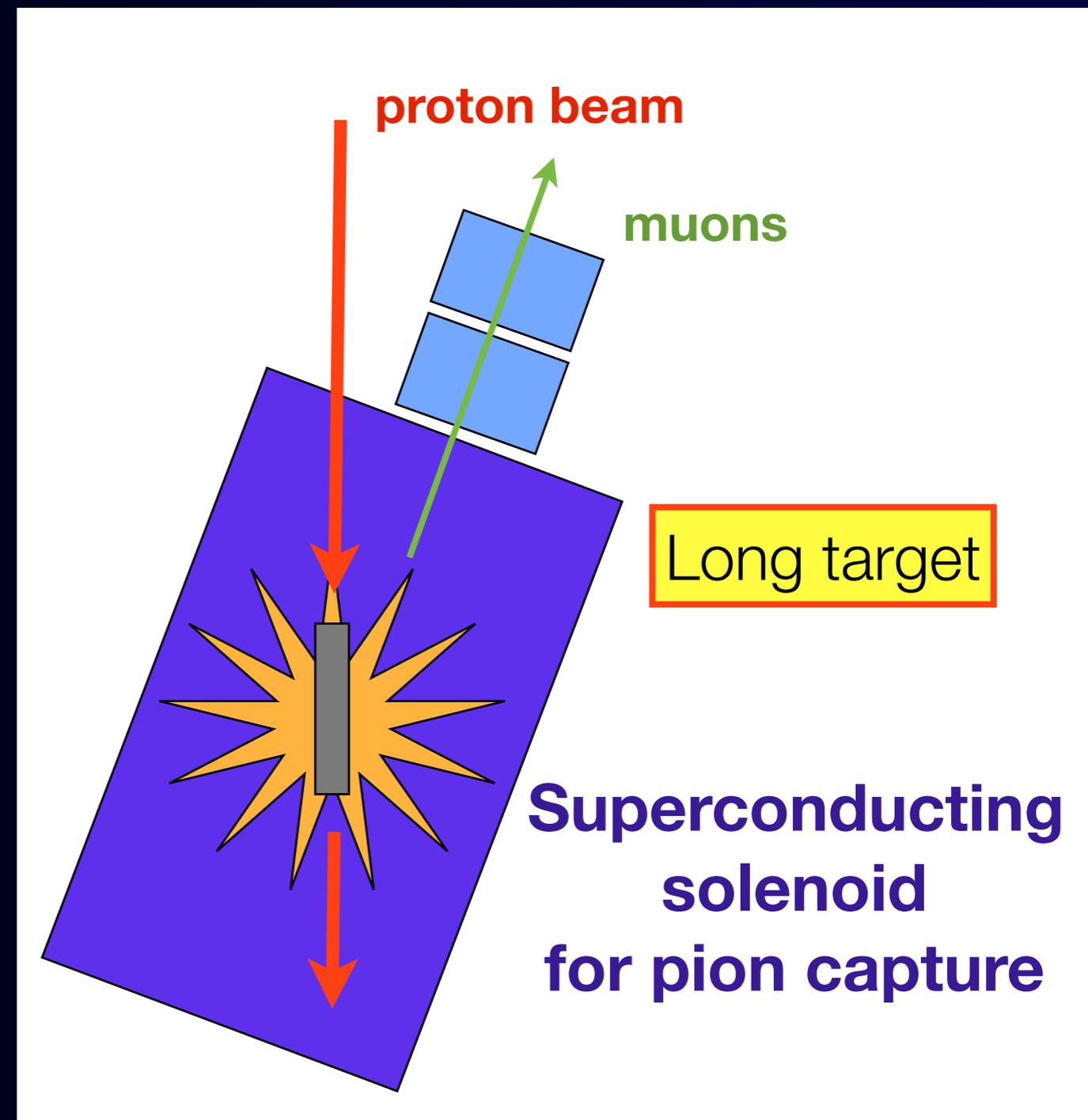
Production and Collection of Pions/Muons in Solenoidal Magnetic Fields



Conventional muon beam source



New muon beam sources



Article on the MuSIC facility...



Editors Suggestions

March, 2017, Editors Suggestions

PHYSICAL REVIEW ACCELERATORS AND BEAMS **20**, 030101 (2017)



Delivering the world's most intense muon beam

S. Cook,¹ R. D'Arcy,¹ A. Edmonds,¹ M. Fukuda,² K. Hatanaka,² Y. Hino,³ Y. Kuno,³
M. Lancaster,¹ Y. Mori,⁴ T. Ogitsu,⁵ H. Sakamoto,³ A. Sato,³ N. H. Tran,³ N. M. Truong,³
M. Wing,^{1,*} A. Yamamoto,⁵ and M. Yoshida⁵

¹*Department of Physics and Astronomy, UCL, Gower Street, London WC1E 6BT, United Kingdom*

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³*Department of Physics, Graduate School of Science, Osaka University, Osaka 569-0043, Japan*

⁴*Kyoto University Reactor Research Institute (KURRI), Kyoto 590-0494, Japan*

⁵*High Energy Accelerator Research Organization (KEK), Tsukuba 305-0801, Japan*

(Received 25 October 2016; published 15 March 2017)

A new muon beam line, the muon science innovative channel, was set up at the Research Center for Nuclear Physics, Osaka University, in Osaka, Japan, using the 392 MeV proton beam impinging on a target. The production of an intense muon beam relies on the efficient capture of pions, which subsequently decay to muons, using a novel superconducting solenoid magnet system. After the pion-capture solenoid, the first 36° of the curved muon transport line was commissioned and the muon flux was measured. In order to detect muons, a target of either copper or magnesium was placed to stop muons at the end of the muon

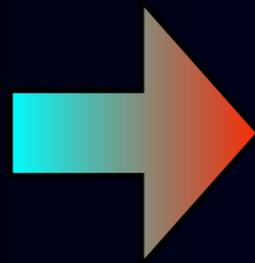
Improvements for Background Rejection



Improvements for Background Rejection



Beam-related
backgrounds



Beam pulsing with
separation of 1 μsec

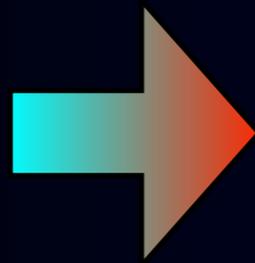
measured
between beam
pulses

proton extinction = #protons between pulses/#protons in a pulse $< 10^{-10}$

Improvements for Background Rejection



Beam-related backgrounds

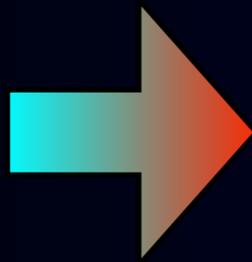


Beam pulsing with separation of 1 μsec

measured between beam pulses

proton extinction = #protons between pulses/#protons in a pulse $< 10^{-10}$

Muon DIO background



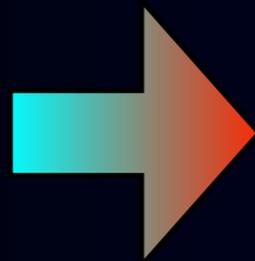
low-mass trackers in vacuum & thin target

improve electron energy resolution

Improvements for Background Rejection



Beam-related backgrounds

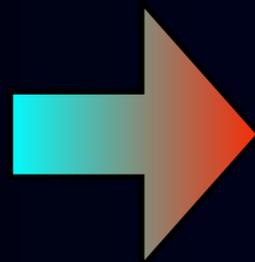


Beam pulsing with separation of 1 μsec

measured between beam pulses

proton extinction = #protons between pulses/#protons in a pulse $< 10^{-10}$

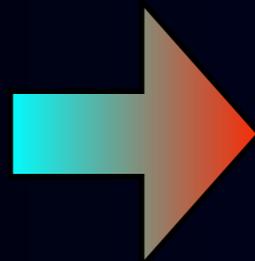
Muon DIO background



low-mass trackers in vacuum & thin target

improve electron energy resolution

Muon DIF background



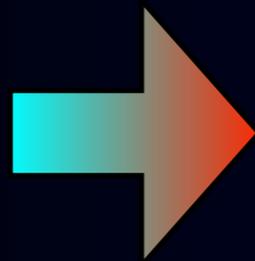
curved solenoids for momentum selection

eliminate energetic muons (>75 MeV/c)

Improvements for Background Rejection



Beam-related backgrounds

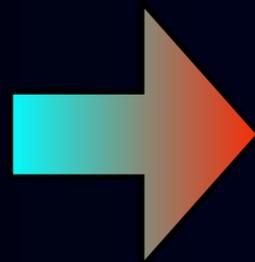


Beam pulsing with separation of 1 μsec

measured between beam pulses

proton extinction = #protons between pulses/#protons in a pulse $< 10^{-10}$

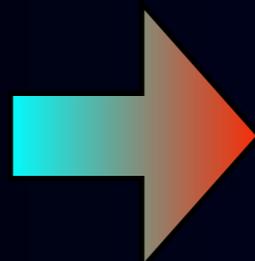
Muon DIO background



low-mass trackers in vacuum & thin target

improve electron energy resolution

Muon DIF background



curved solenoids for momentum selection

eliminate energetic muons (>75 MeV/c)

based on the MELC proposal at Moscow Meson Factory

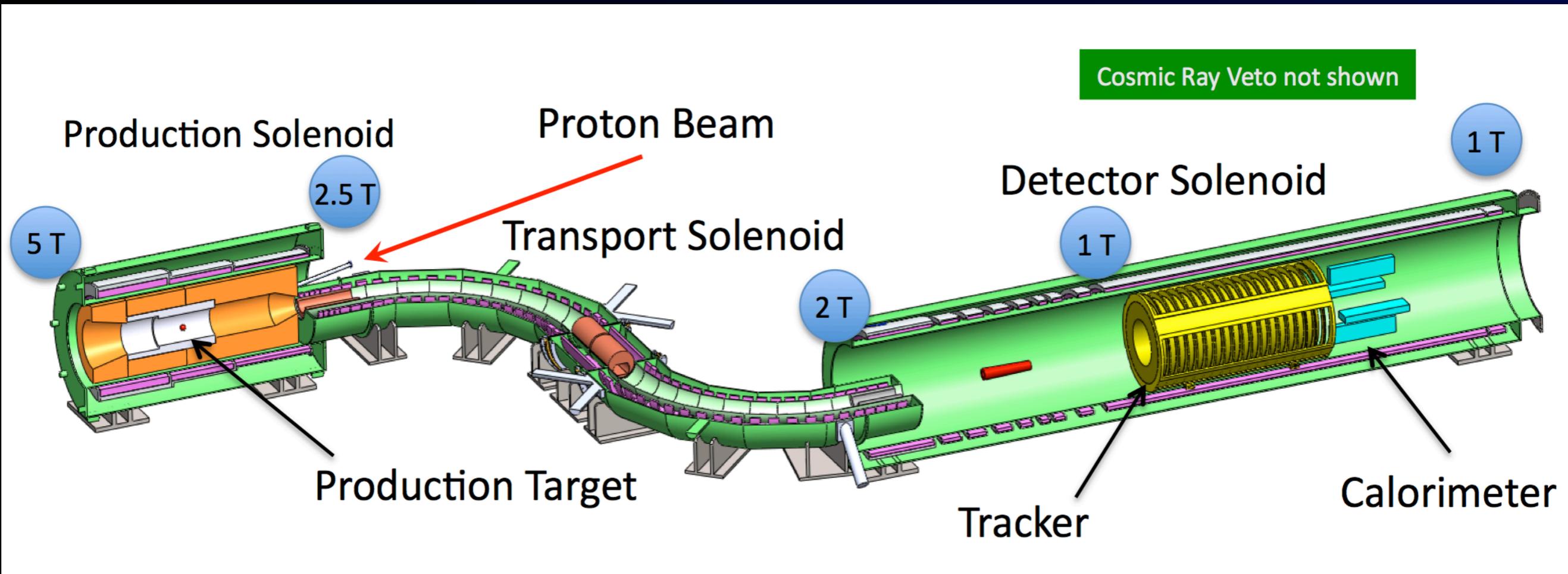
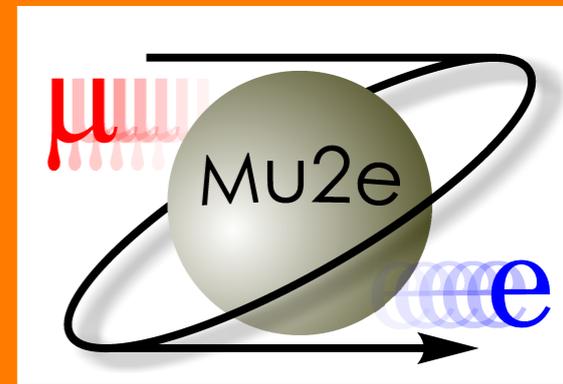
Experiments
for μ -e conversion



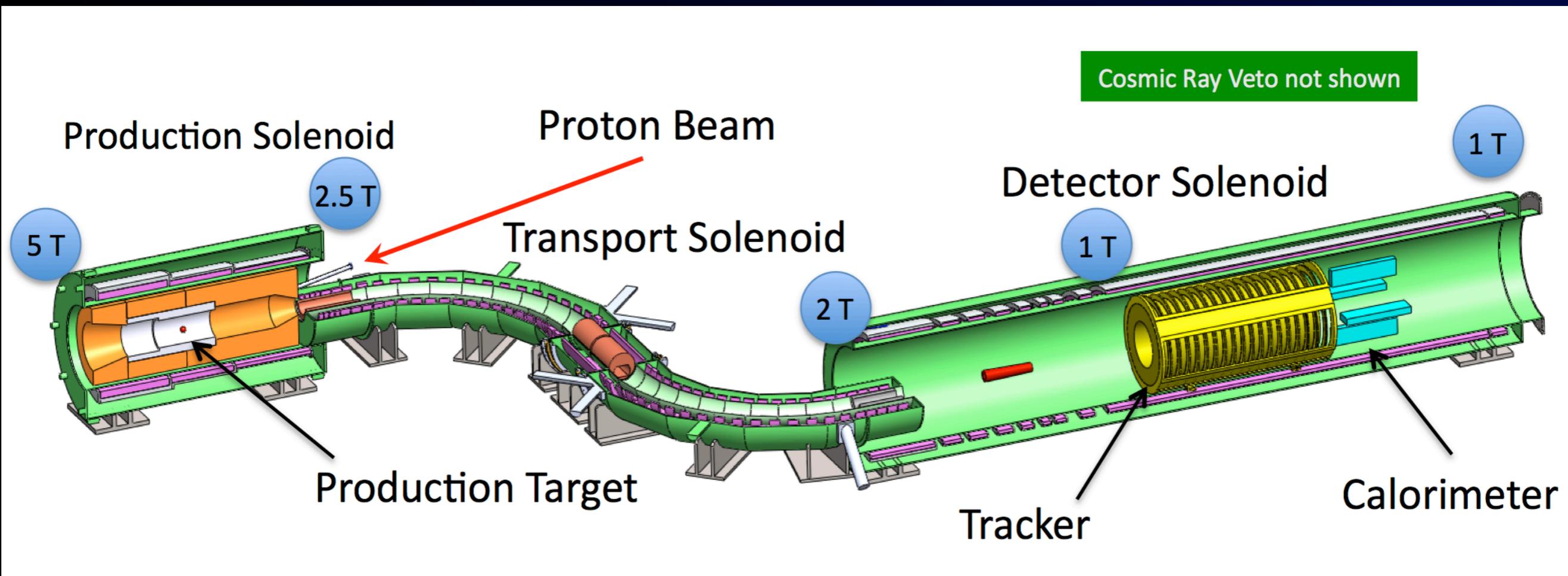
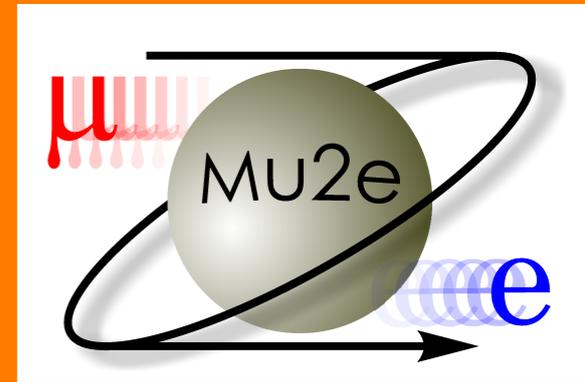
Mu2e at Fermilab



Mu2e at Fermilab



Mu2e at Fermilab



Single-event sensitivity : $(2.5 \pm 0.3) \times 10^{-17}$

Total background : (0.36 ± 0.10) events

Expected limits : $< 6 \times 10^{-17}$ @90%C.L.

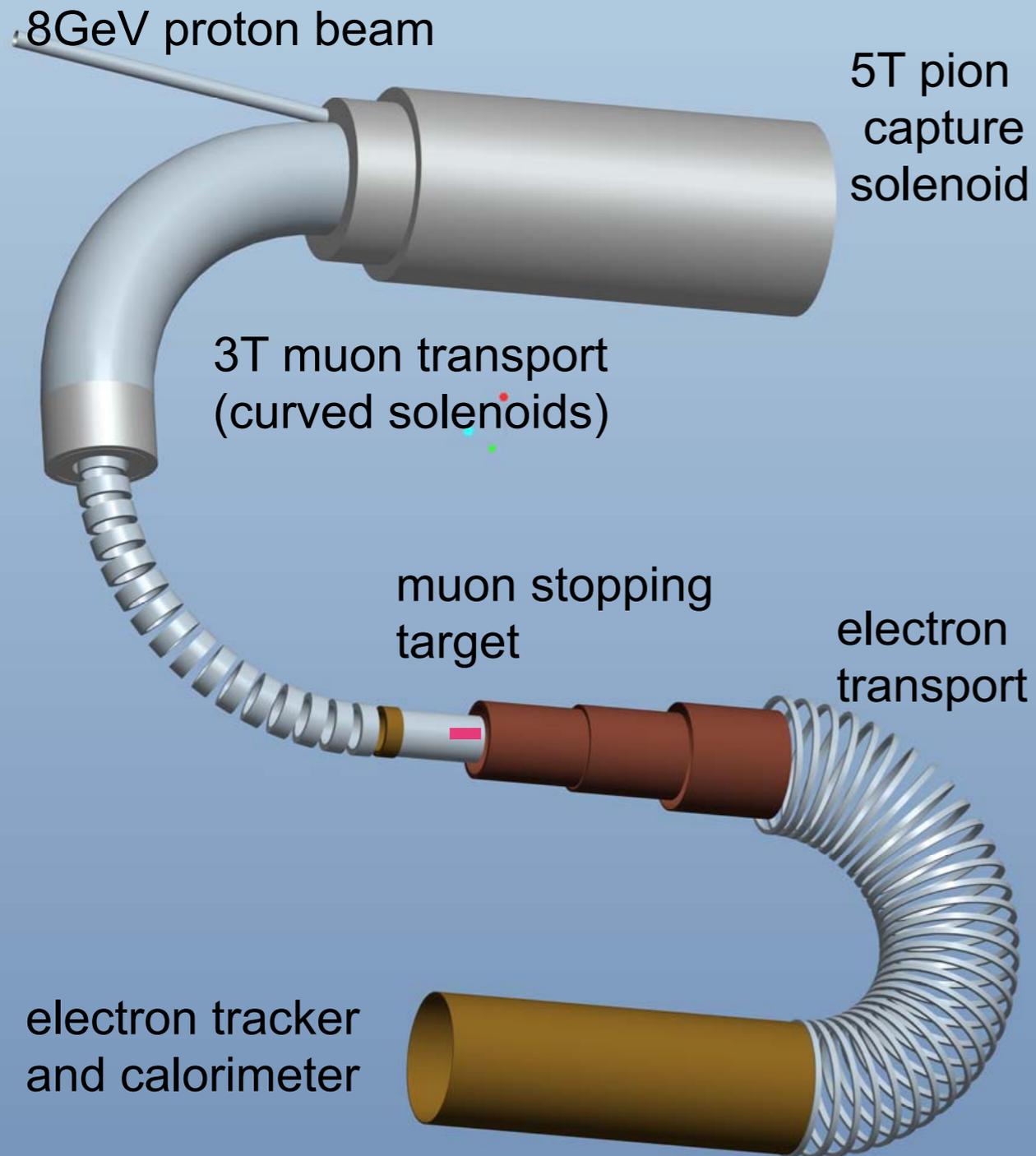
Running time: 3 years (2×10^7 sec/year)

proton beam power = 8 kW

COMET at J-PARC: E21



COMET at J-PARC: E21



Physics sensitivity : $(1.0-2.6) \times 10^{-17}$
Total background : 0.32 events
Expected limits : $< 6 \times 10^{-17}$ @90%CL
Running time: 1 years (2×10^7 sec)

proton beam power = 56 kW

COMET = COherent Muon
to Electron Transition







COMET Phase-I

COMET Staged Approach (2012~)

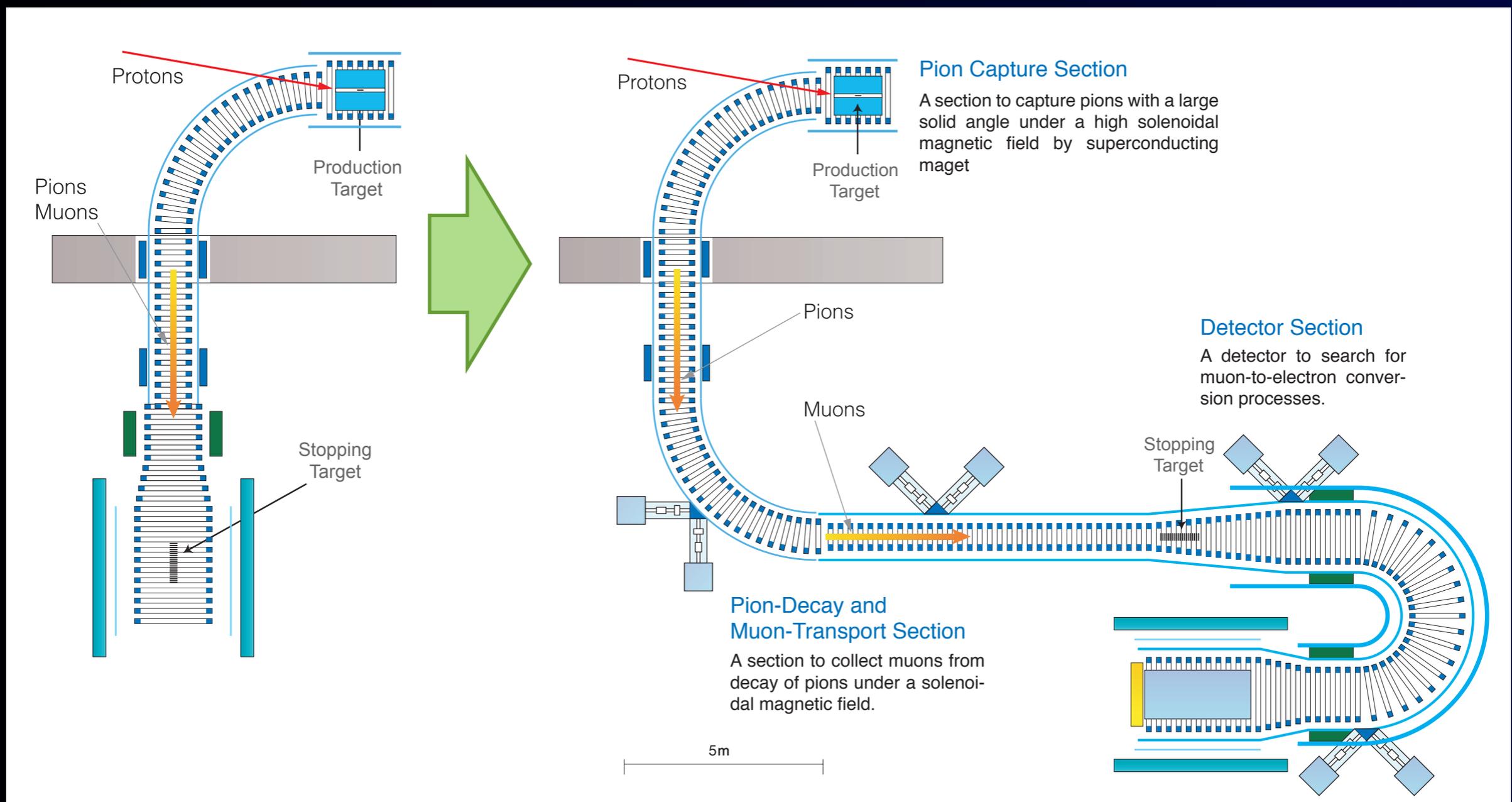


COMET Staged Approach (2012~)



COMET Phase-I

COMET Phase-II

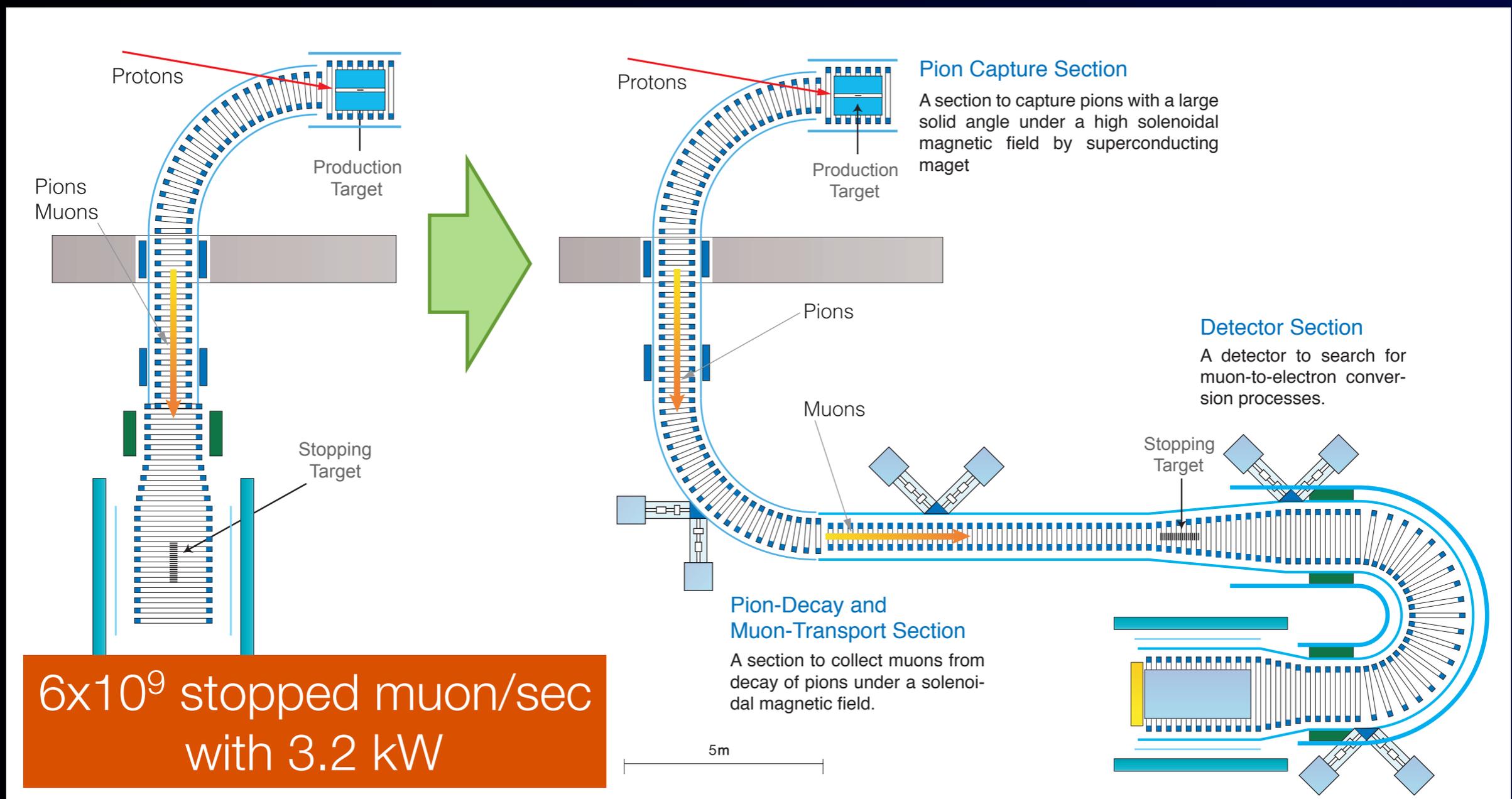


COMET Staged Approach (2012~)



COMET Phase-I

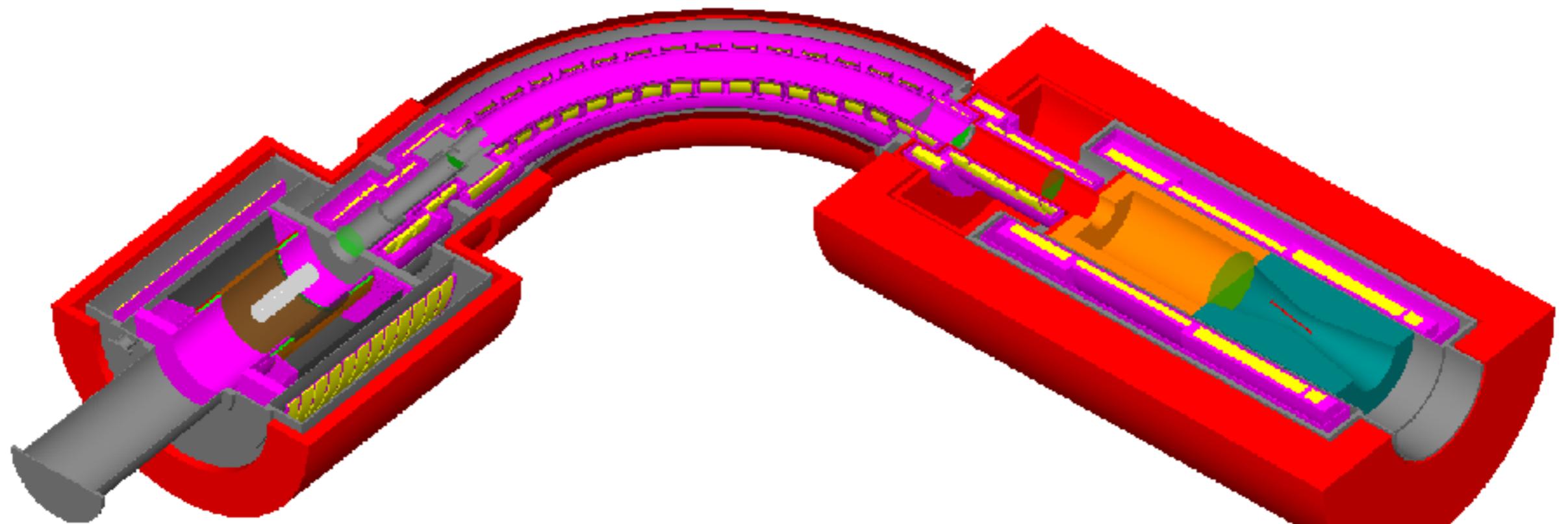
COMET Phase-II



COMET Phase-I



COMET Phase-I

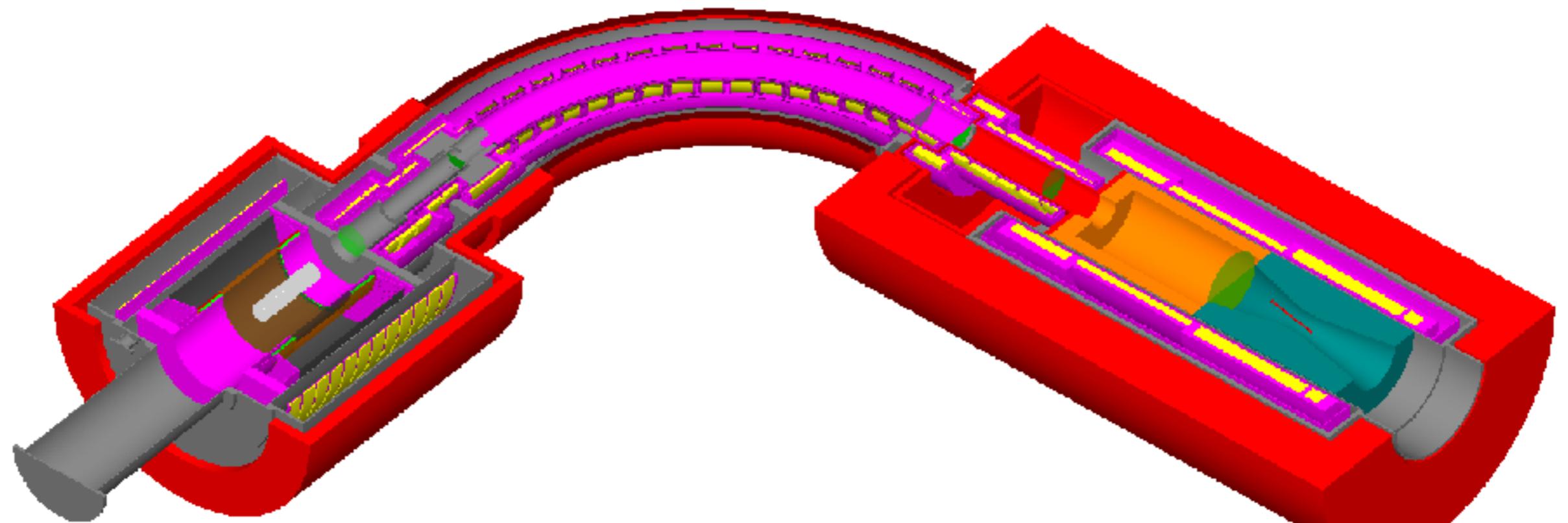


detector system

muon transport system

pion production system

COMET Phase-I



detector system

muon transport system

pion production system

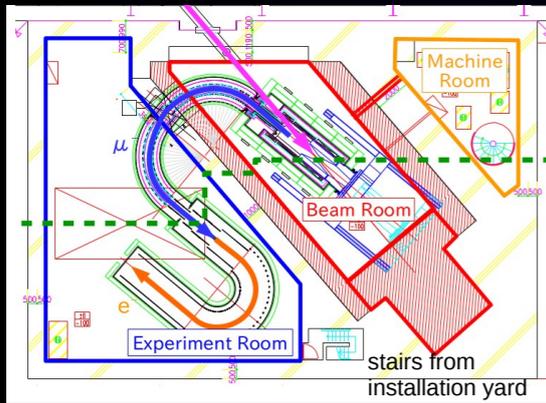
Single-event sensitivity : 3×10^{-15}

Total background : 0.2 events

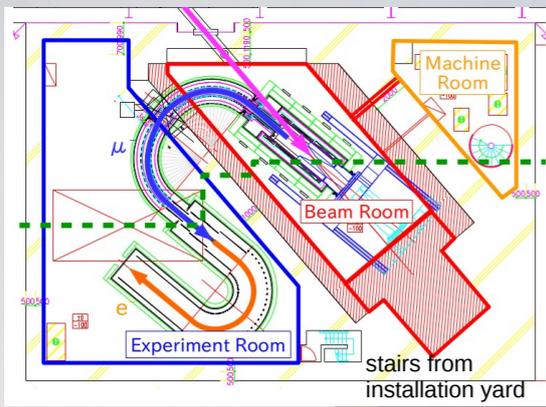
Expected limits : $< 6 \times 10^{-15}$ @90%CL

Running time: 150 days

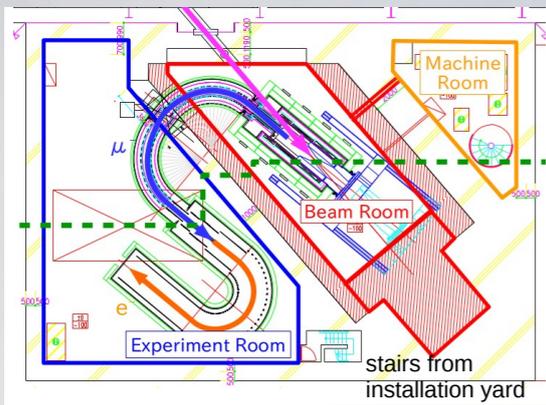
New COMET Building at J-PARC



New COMET Building at J-PARC

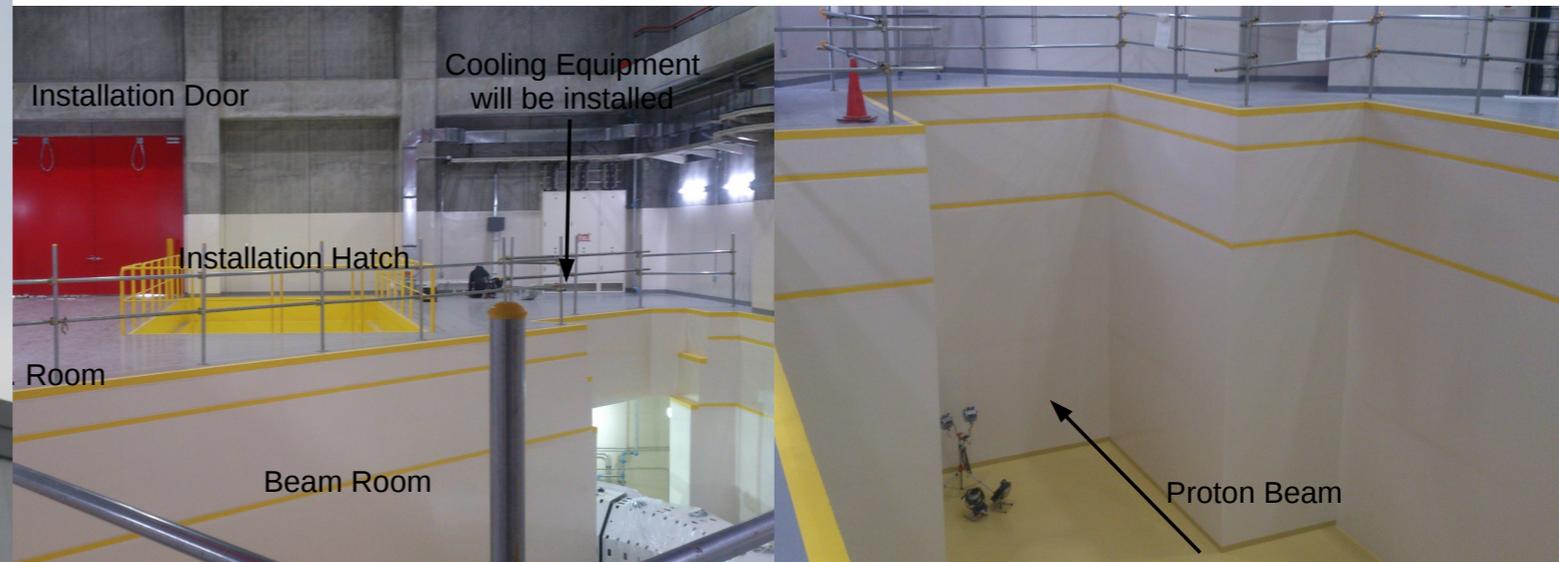


New COMET Building at J-PARC



Installation Yard

Beam Room



Experiment Room (View from Door)

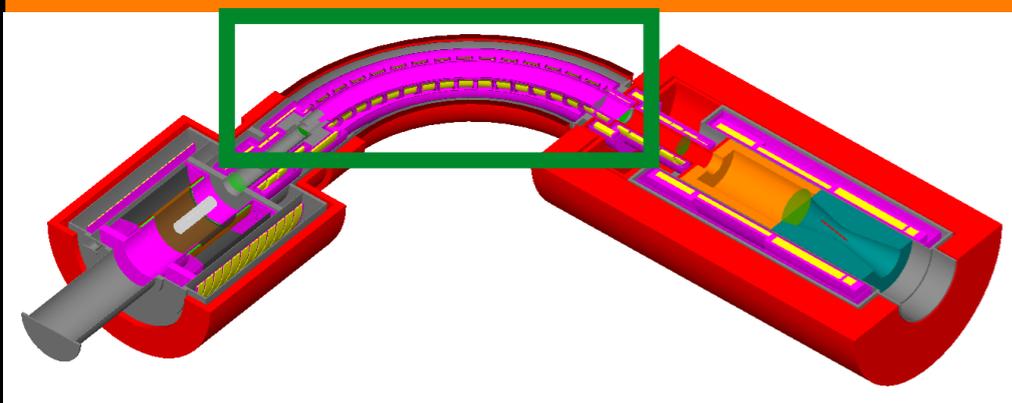
Beam Room (view from Installation Yard)



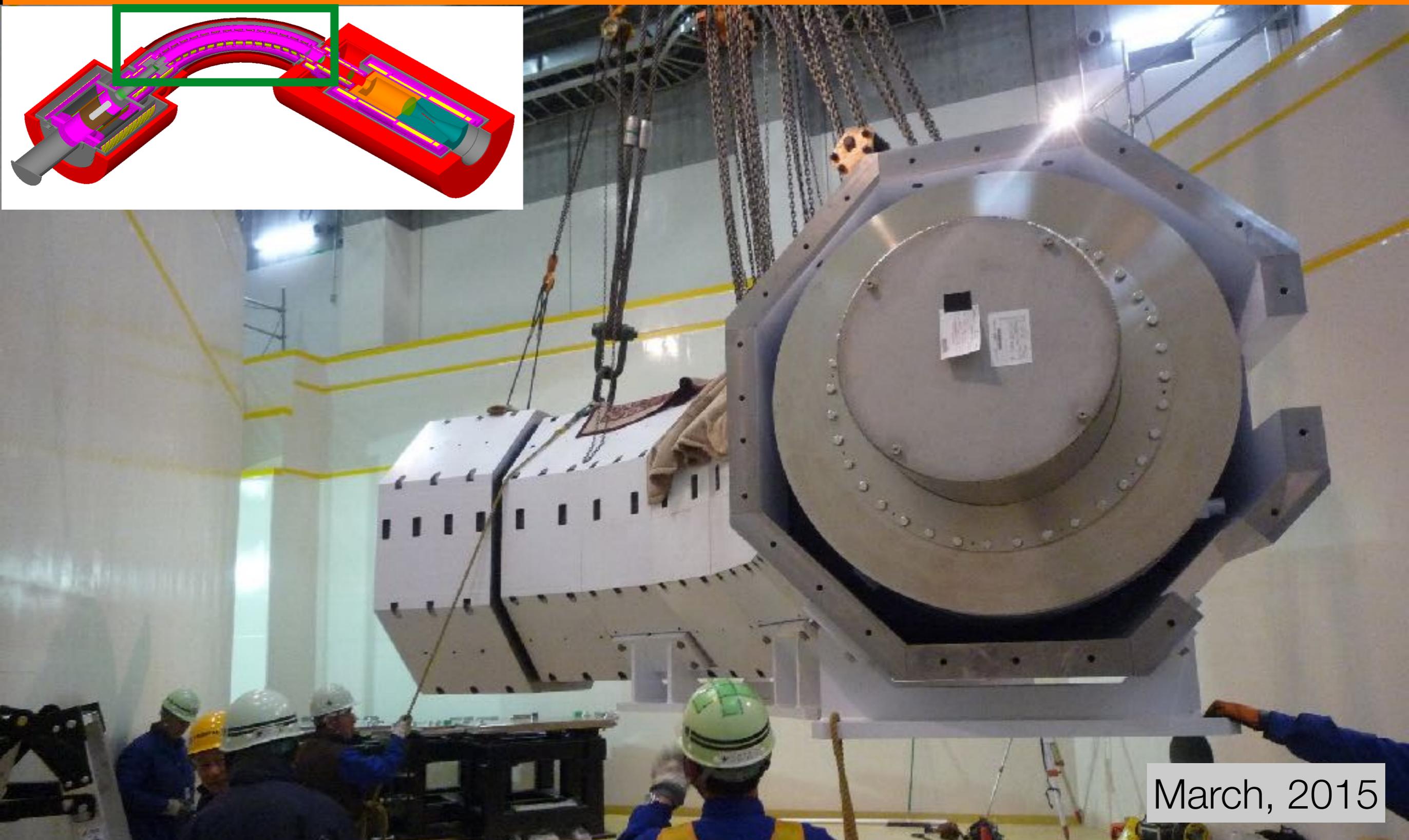
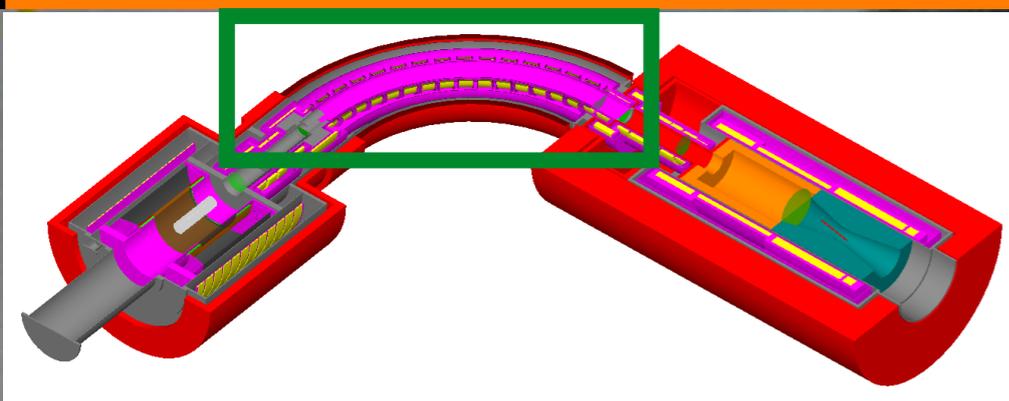
Curved Solenoids for Muon Transport Completed and Delivered!



Curved Solenoids for Muon Transport Completed and Delivered!

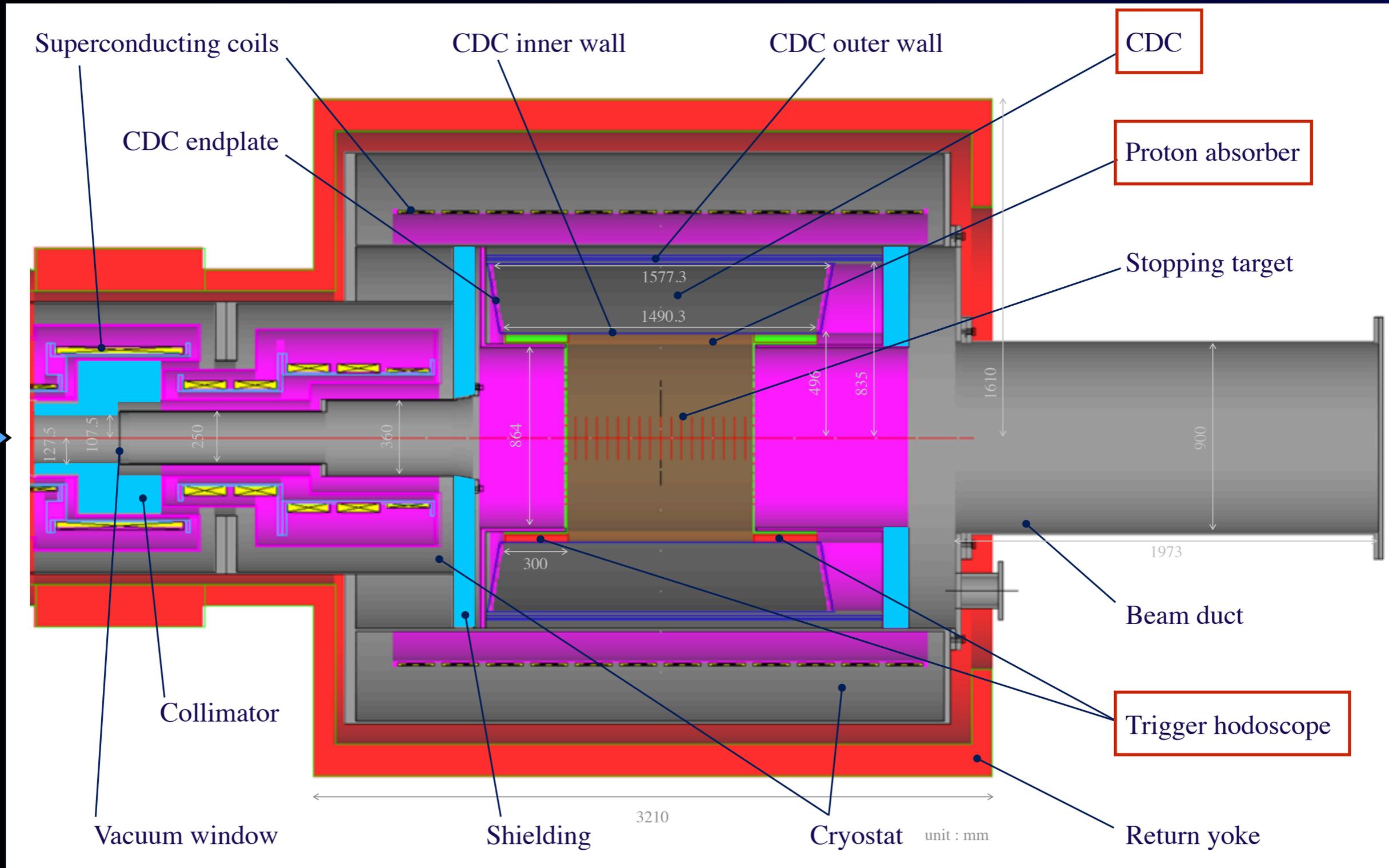
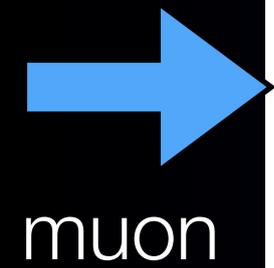


Curved Solenoids for Muon Transport Completed and Delivered!

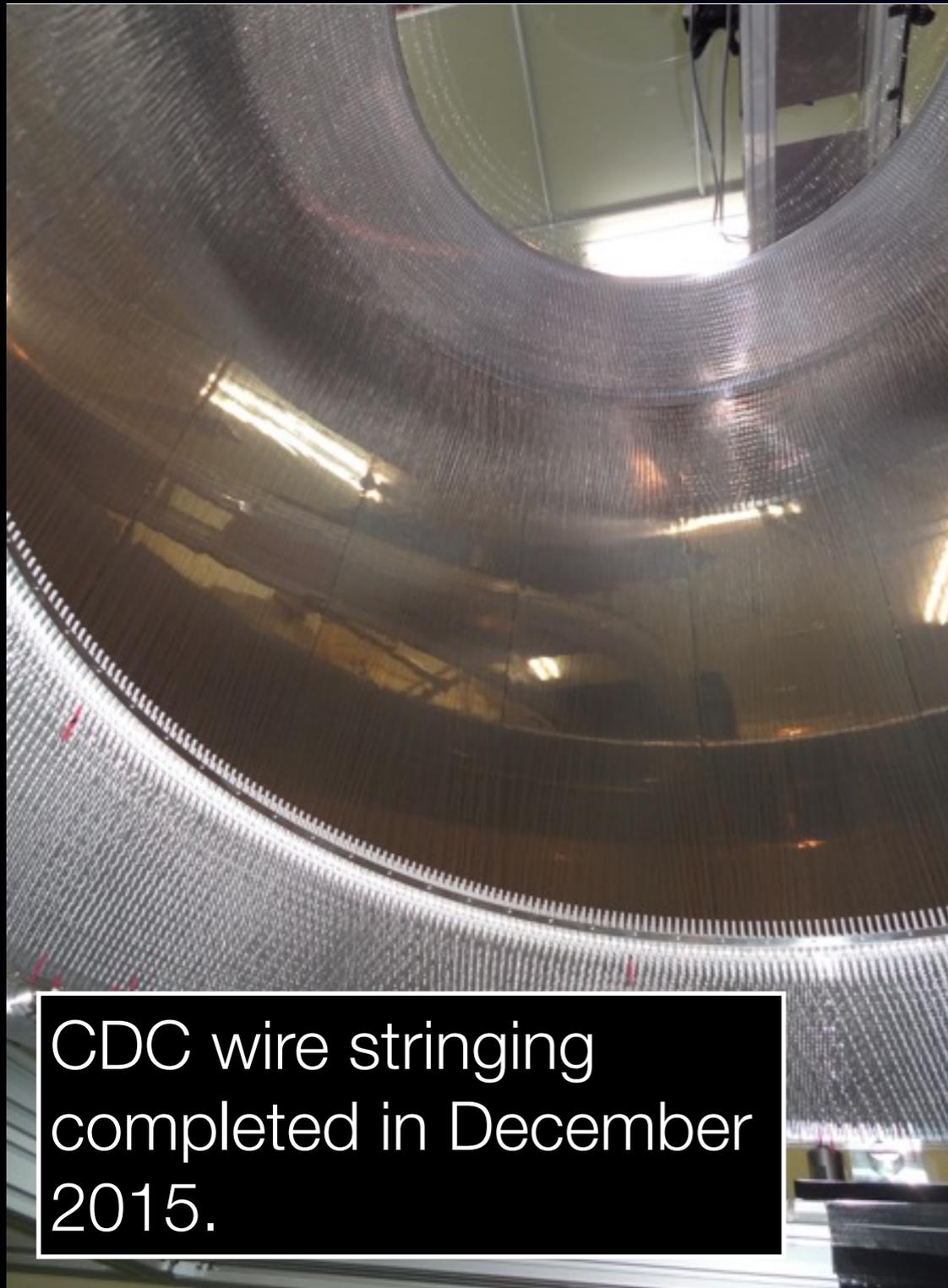


March, 2015

CyDet (Cylindrical Detector)



CDC Construction completed!

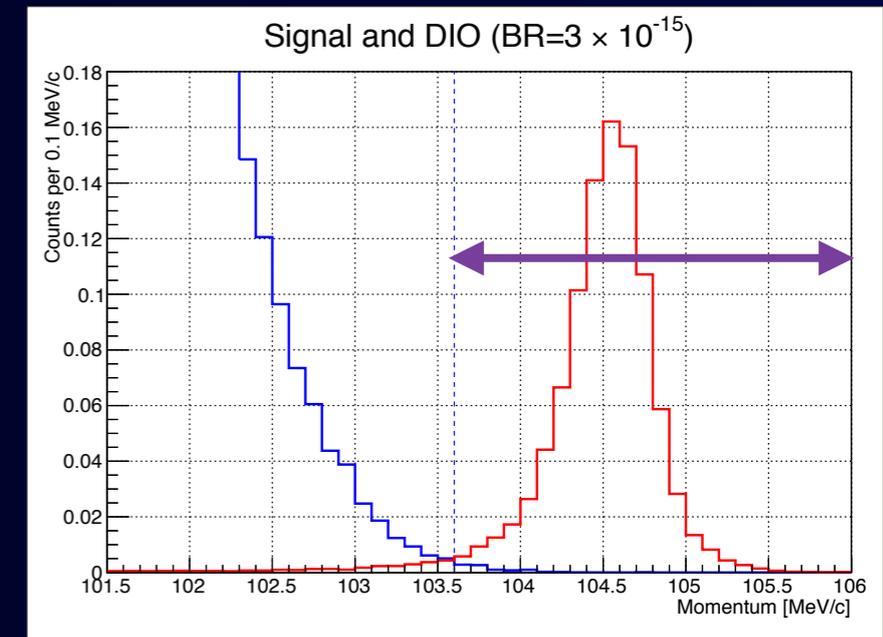


COMET Phase-I Signal Sensitivity

Signal Acceptance

Table 28: Breakdown of the $\mu^- N \rightarrow e^- N$ conversion signal acceptance.

Event selection	Value	Comments
Geometrical acceptance	0.37	
Track quality cuts	0.66	
Momentum selection	0.93	$103.6 \text{ MeV}/c < P_e < 106.0 \text{ MeV}/c$
Timing window	0.3	$700 \text{ ns} < t < 1100 \text{ ns}$
Trigger efficiency	0.8	
DAQ efficiency	0.8	
Track reconstruction efficiency	0.8	
Total	0.043	



Signal Sensitivity

$$B(\mu^- + Al \rightarrow e^- + Al) \sim \frac{1}{N_\mu \cdot f_{cap} \cdot A_e},$$

$$B(\mu^- + Al \rightarrow e^- + Al) = 3.1 \times 10^{-15}$$

$$B(\mu^- + Al \rightarrow e^- + Al) < 7 \times 10^{-15} \quad (90\% C.L.)$$

Muon intensity

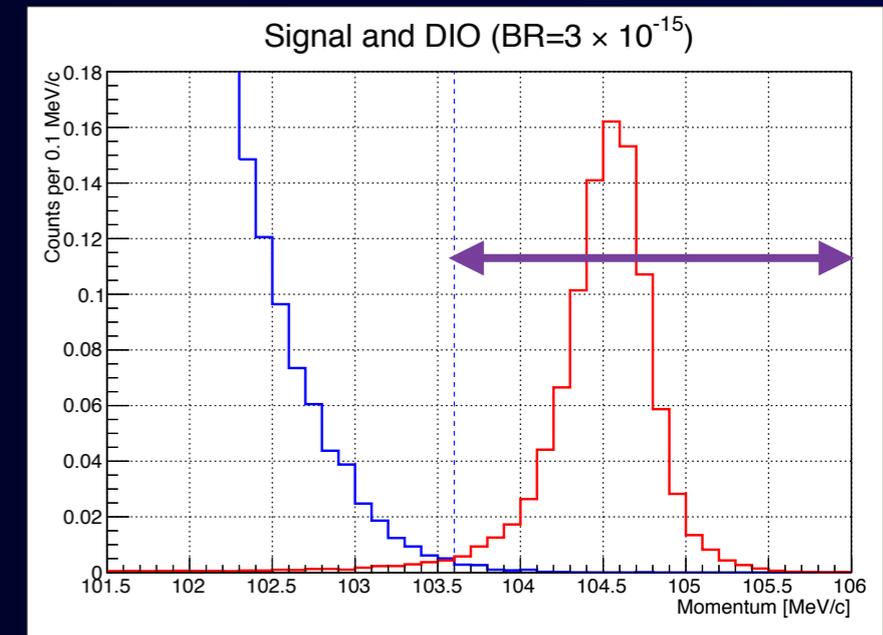
With 0.4 μA , a running time of about 150 days is needed.

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Signal Sensitivity

- $f_{\text{cap}} = 0.6$
- $A_e = 0.043$
- $N_\mu = 1.23 \times 10^{16}$ muons

$$B(\mu^- + Al \rightarrow e^- + Al) \sim \frac{1}{N_\mu \cdot f_{\text{cap}} \cdot A_e},$$

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With $0.4 \mu\text{A}$, a running time of about 150 days is needed.

Tentative Schedule of COMET Phase-I and Phase-II



	JFY	2015	2016	2017	2018	2019	2020	2021	2022	2023
COMET Phase-I	construction	[Grey bar spanning 2015-2018]								
	data taking					[Grey bar spanning 2019]				
COMET Phase-II	construction						[Grey bar spanning 2020-2022]			
	data taking								[Grey bar spanning 2022-2023]	

Tentative Schedule of COMET Phase-I and Phase-II



	JFY	2015	2016	2017	2018	2019	2020	2021	2022	2023
COMET Phase-I	construction	[Bar]								
	data taking					[Bar]				
COMET Phase-II	construction						[Bar]			
	data taking								[Bar]	

COMET Phase-I :
 2019 ~
 S.E.S. ~ 3×10^{-15}
 (for 150 days
 with 3.2 kW proton beam)

Tentative Schedule of COMET Phase-I and Phase-II



	JFY	2015	2016	2017	2018	2019	2020	2021	2022	2023
COMET Phase-I	construction	[Bar]								
	data taking					[Bar]				
COMET Phase-II	construction						[Bar]	[Bar]		
	data taking								[Bar]	[Bar]

COMET Phase-I :
 2019 ~
 S.E.S. ~ 3×10^{-15}
 (for 150 days
 with 3.2 kW proton beam)

COMET Phase-II :
 2022 ~
 S.E.S. ~ $(1.0-2.6) \times 10^{-17}$
 (for 2×10^7 sec
 with 56 kW proton beam)

Summary



Summary



- Charged lepton flavor violation (CLFV) would provide the best opportunity to search for new physics beyond the SM.
- Next generation experiments for CLFV with muons are coming.
 - MEG II for $\mu \rightarrow e \gamma$
 - Mu3e for $\mu \rightarrow eee$
 - COMET and Mu2e for muon to electron conversion.
- Stay tuned...

Summary



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my dog, IKU



Backup

