



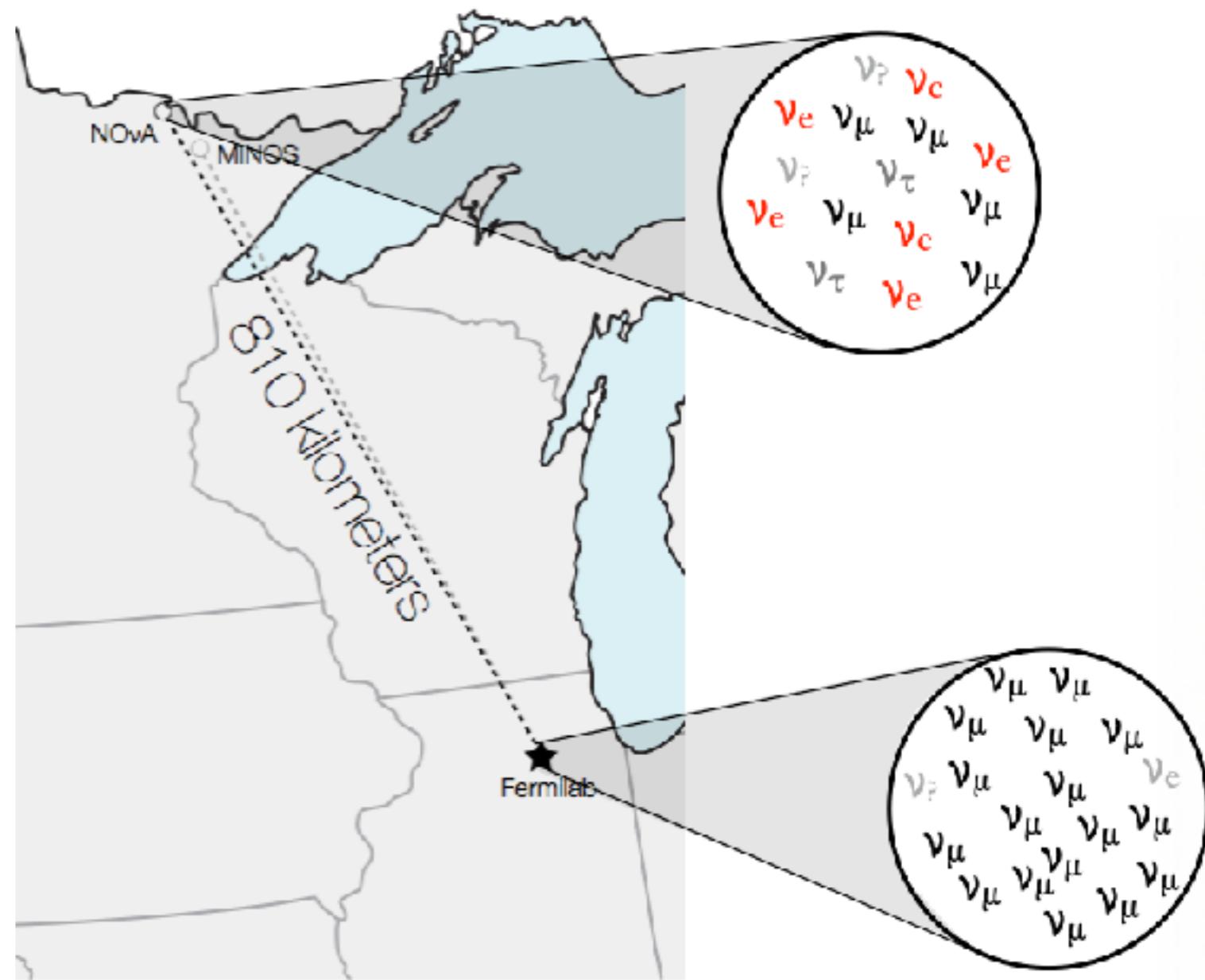
Results from NOvA

Jenny Thomas

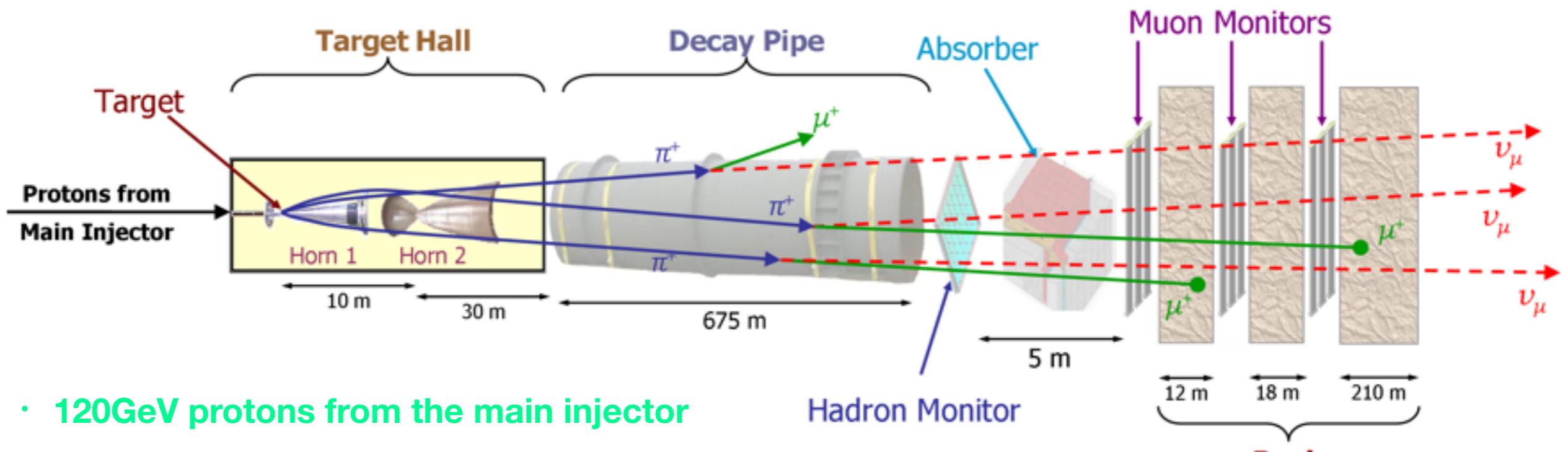


NOvA Experiment

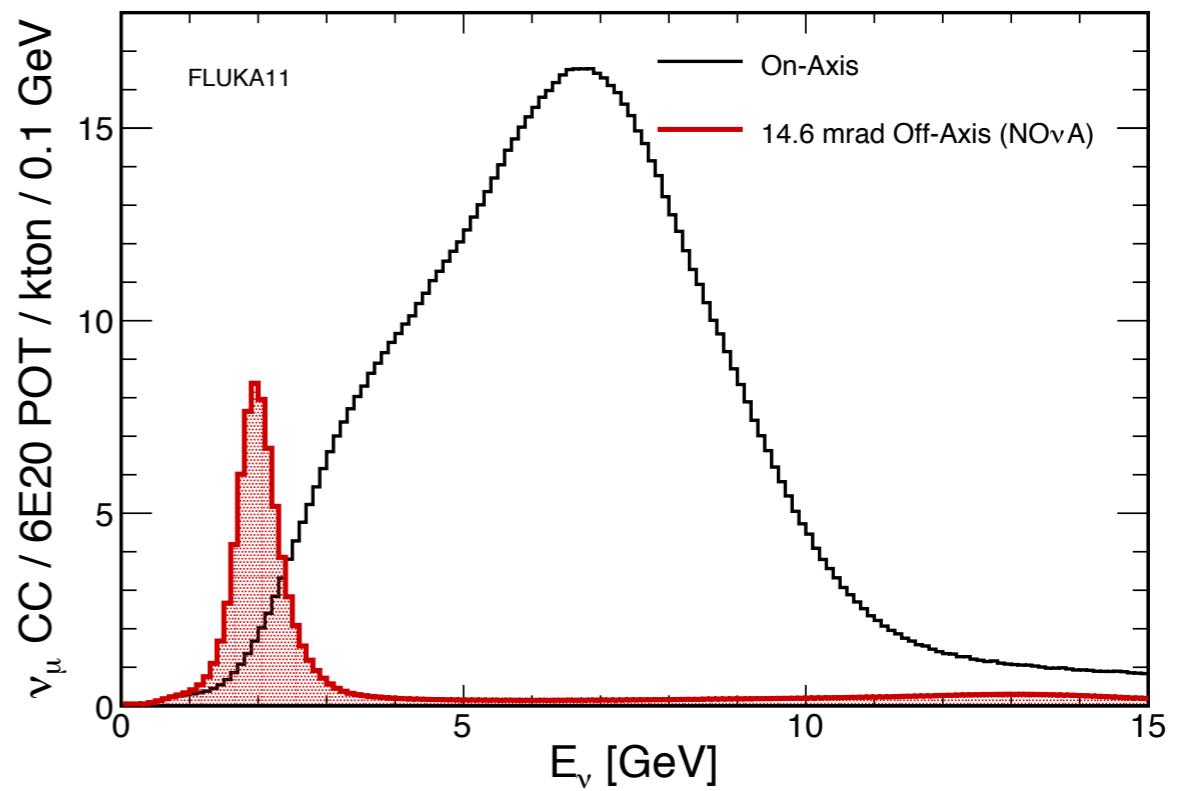
- Longest baseline accelerator neutrino search (810 km)
 - NuMI is a beam of mainly muon-neutrinos created at Fermilab
 - Two functionally identical detectors at 14.6 mrad (FD) and 14.6 mrad (equivalent) for smaller ND
- Measures muon-neutrino disappearance and electron-neutrino appearance
 - starting to do the same with anti-neutrinos
- Sensitive to PMNS matrix, mass hierarchy, CP violation



How to make a neutrino beam

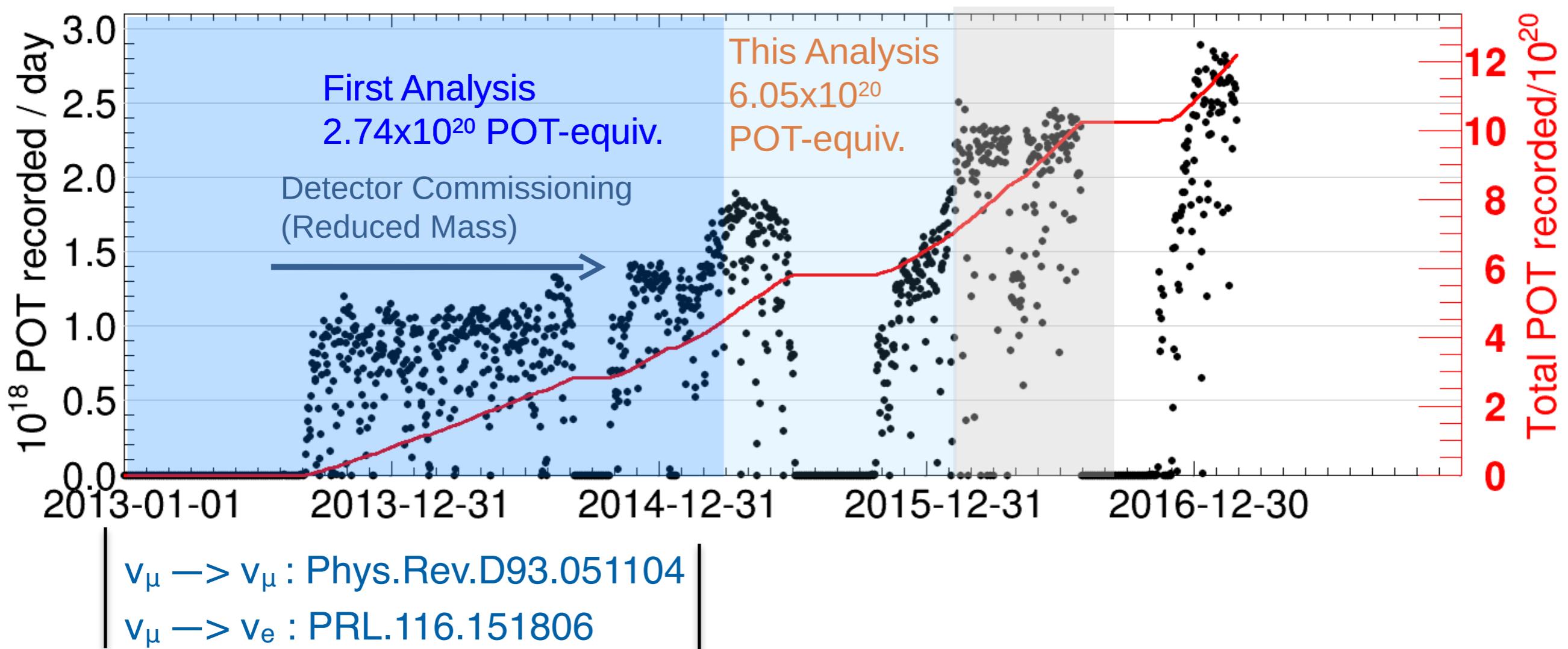


- 120GeV protons from the main injector
- Focus secondary pions using magnetic horns
 - Focus positive hadrons for neutrino beam, negative for antineutrino
- Pions decay to produce muon neutrinos
 - Decay kinematics mean a detector at 14.6mrad sees a narrowly peaked energy spectrum
- 97.5% muon-neutrino, only 0.7% beam electron-neutrino (remainder wrong-sign)

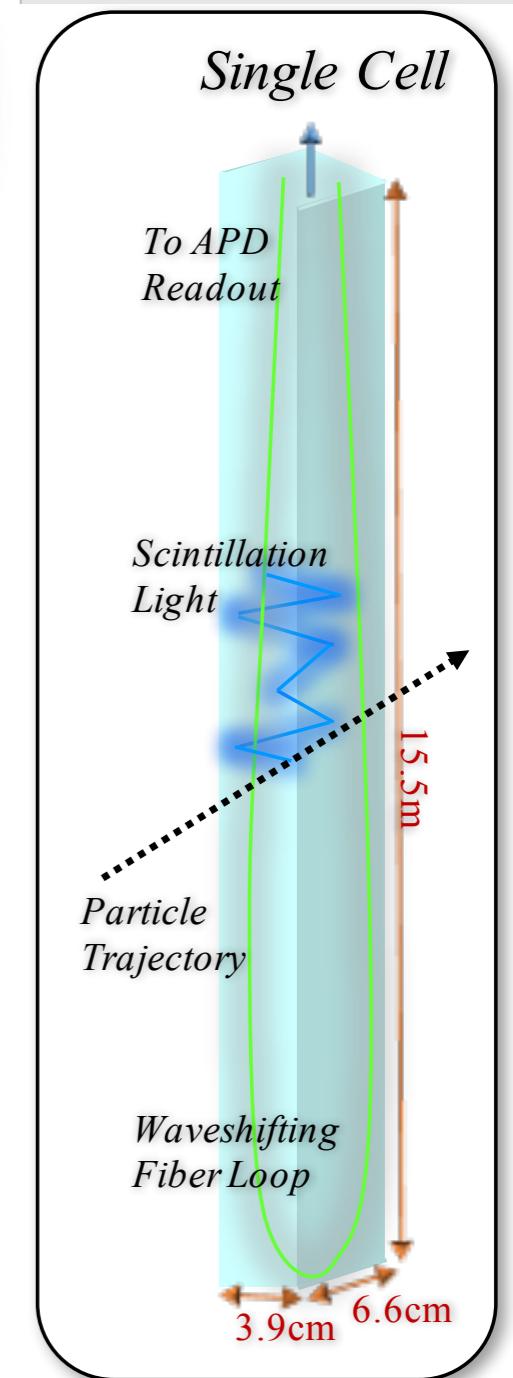
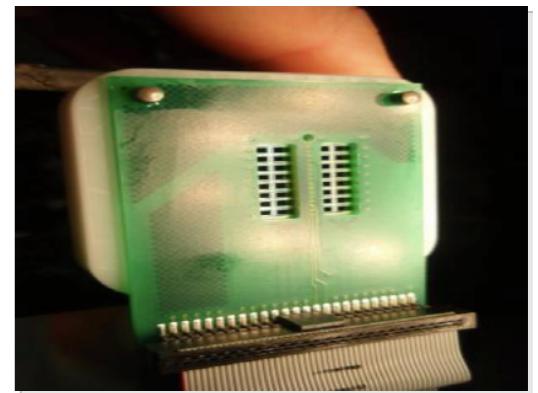
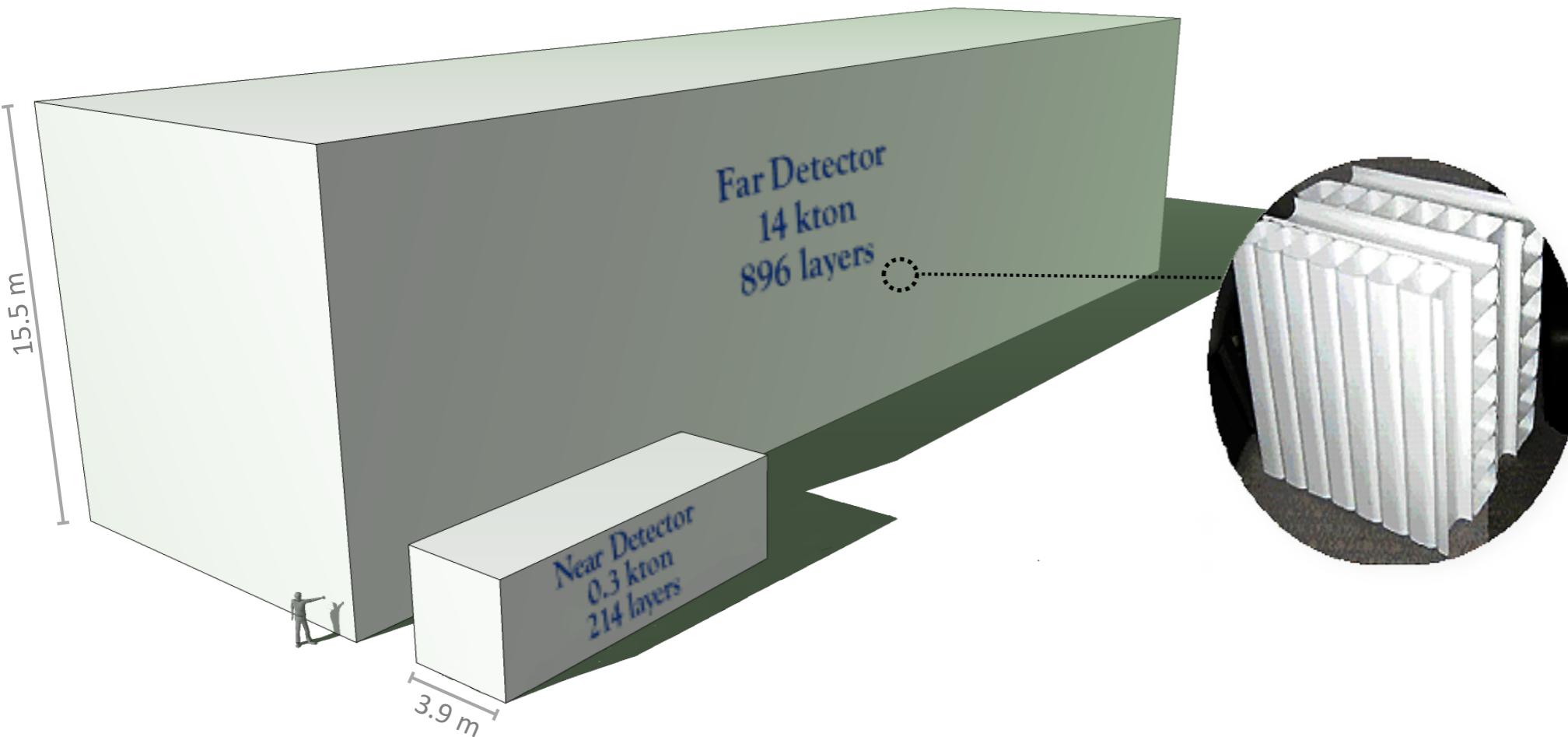


NuMI Beam Performance

- Results today from data collected between February 6, 2014 and May 2, 2016
- Data equivalent to 6.05×10^{20} protons-on-target in a full 14 kT detector
- Achieved 700 kW design goal, most powerful neutrino beam in the world
- Switched to antineutrino beam, shut down first week of July

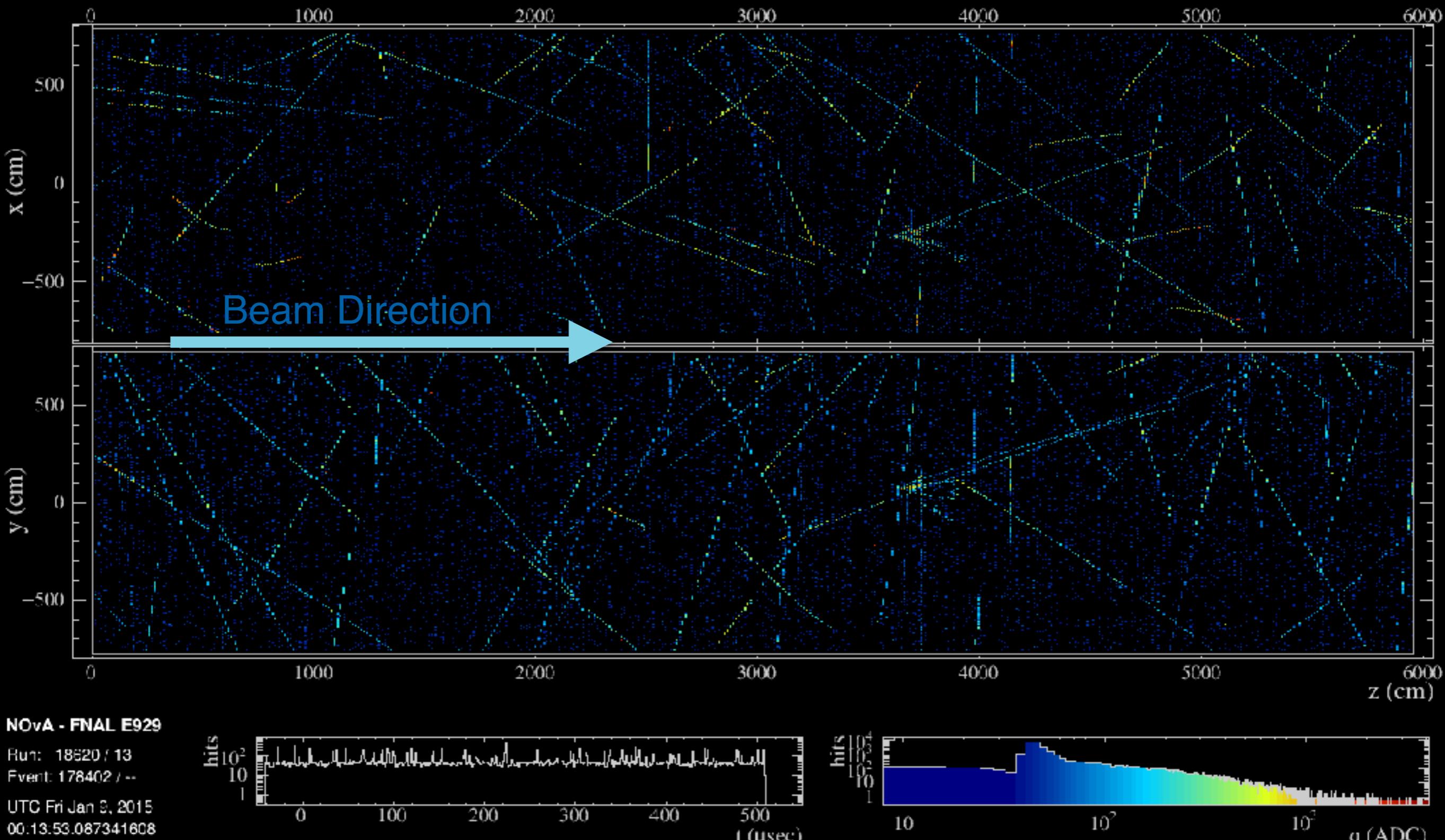


NOvA Detectors

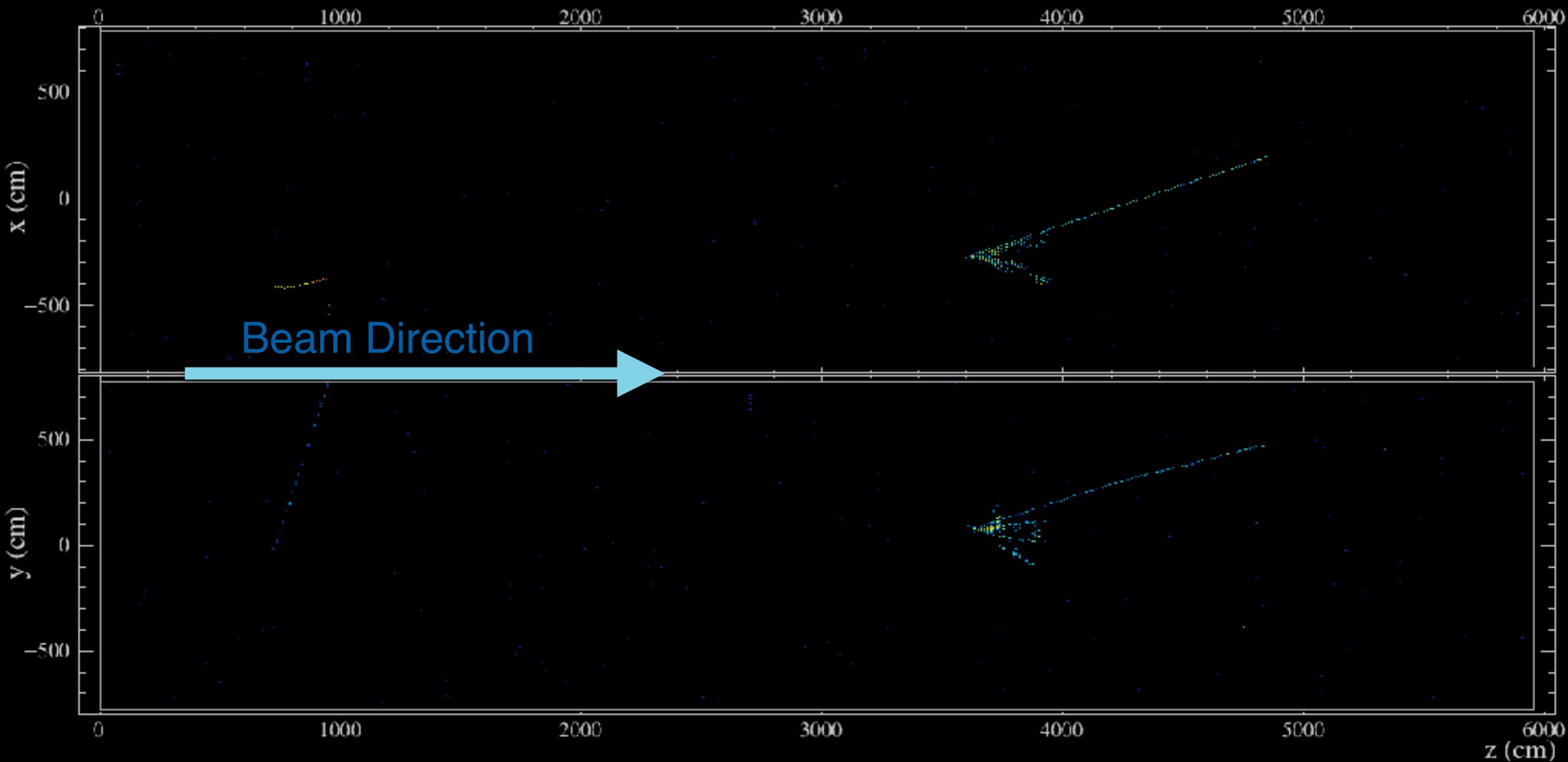


- Two functionally identical detectors
- Extruded plastic cells alternating vertical and horizontal orientation filled with liquid scintillator
- Charged particles passing through cells produce light which is collected by a wavelength shifting fibre and read out with an APD

Far Detector 550 μ s Readout Window



Far Detector 10 μ s NuMI Beam Window



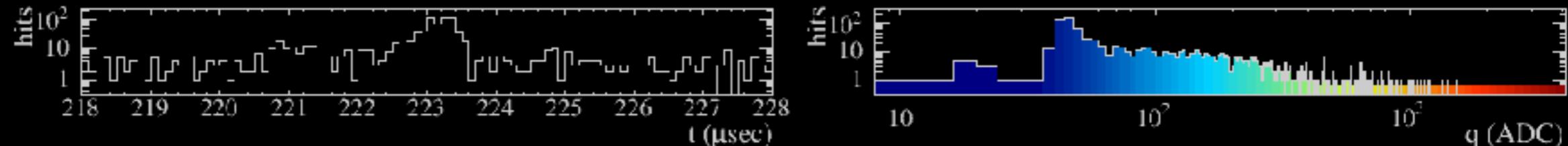
NOvA - FNAL E929

Run: 18620 / 13

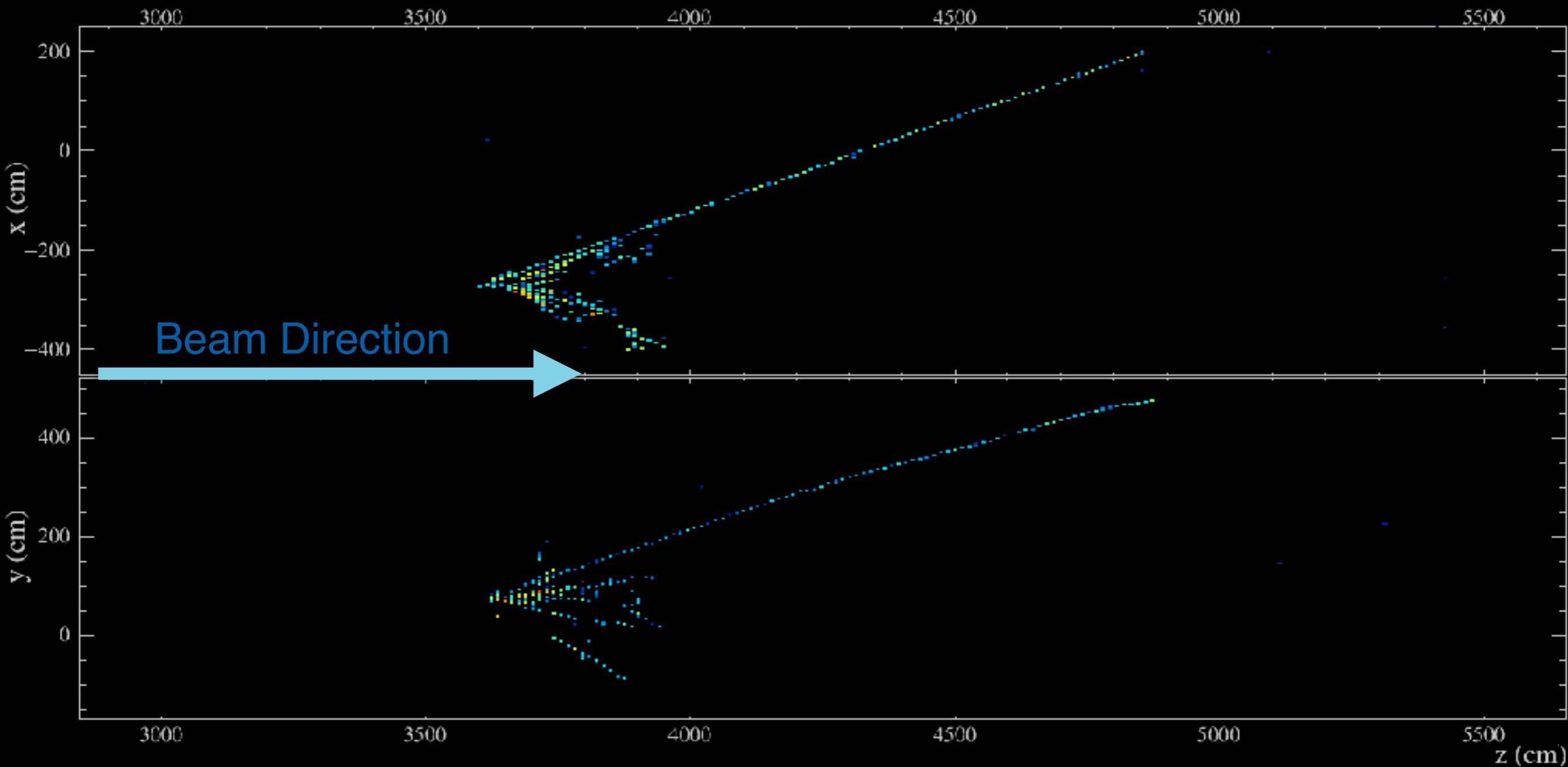
Event: 178402 / --

UTC Fri Jan 9, 2015

00:13:53.087341608



Far Detector Neutrino Interaction



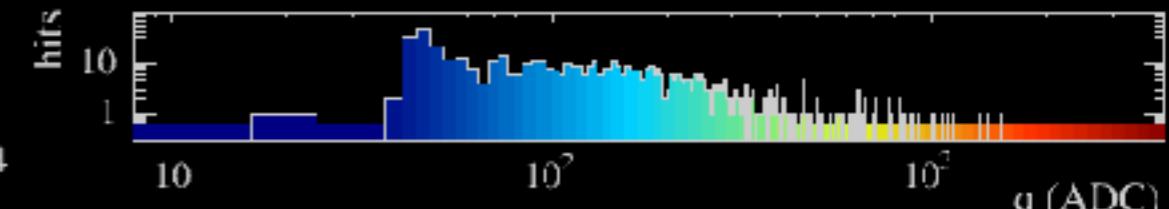
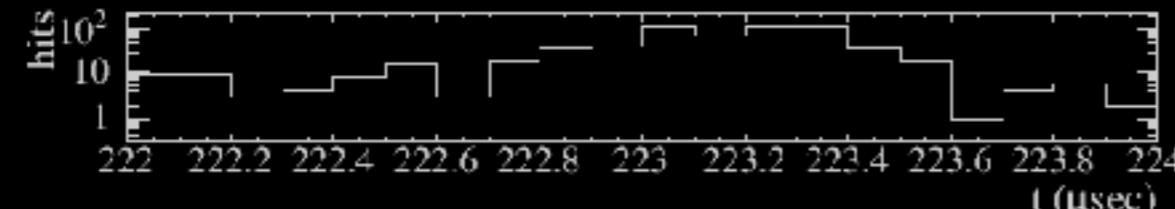
NOvA - FNAL E929

Run: 18620 / 13

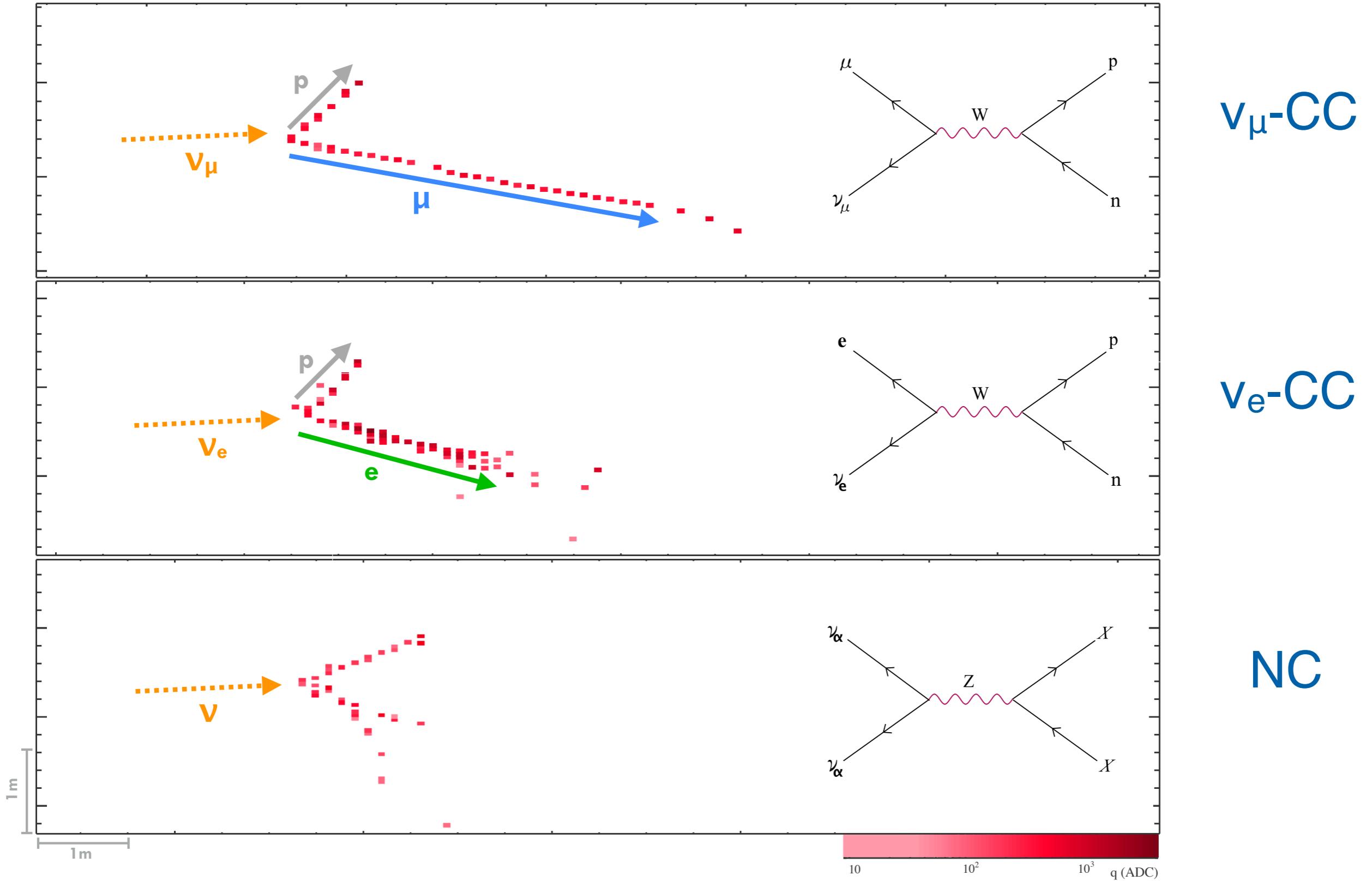
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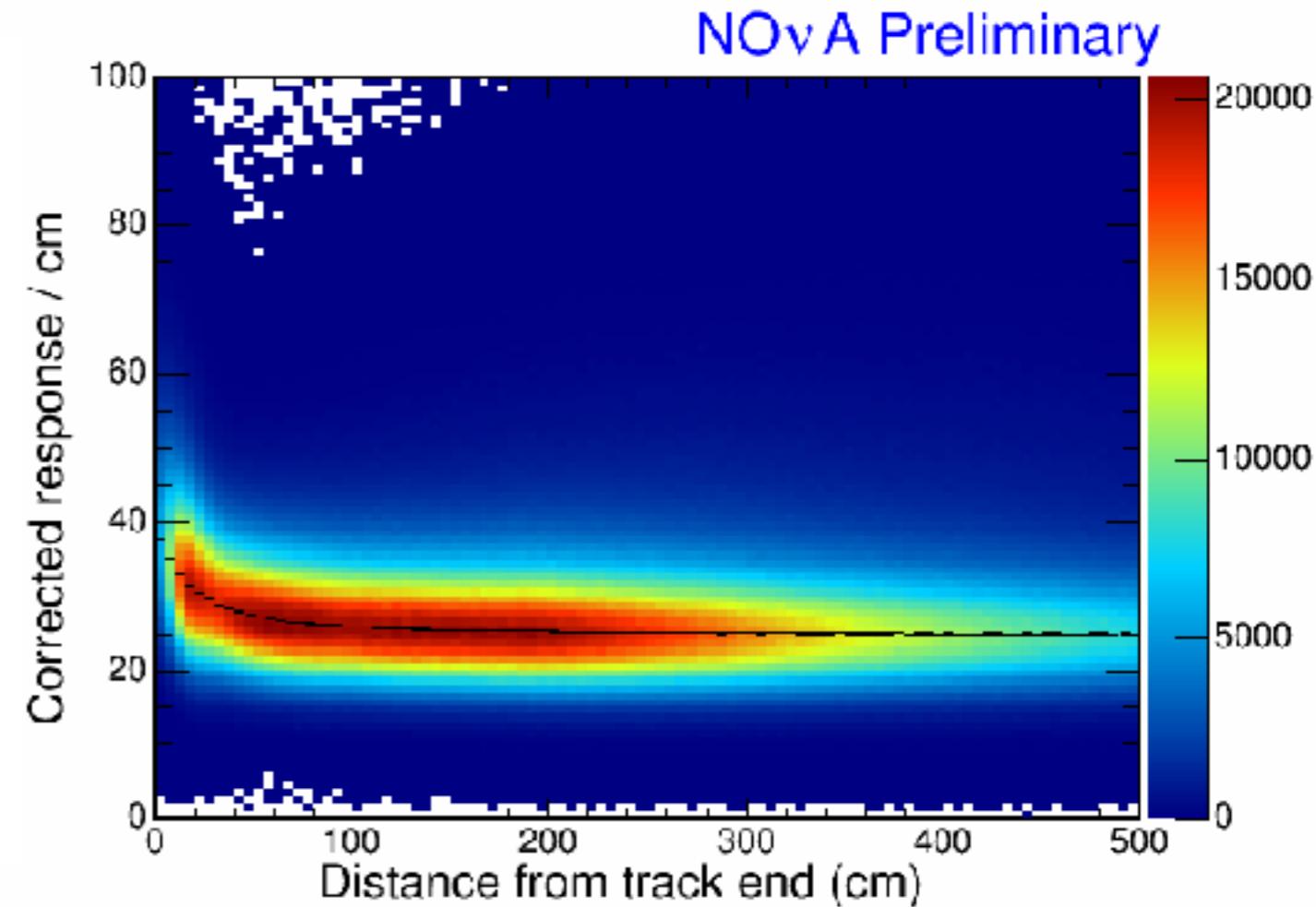
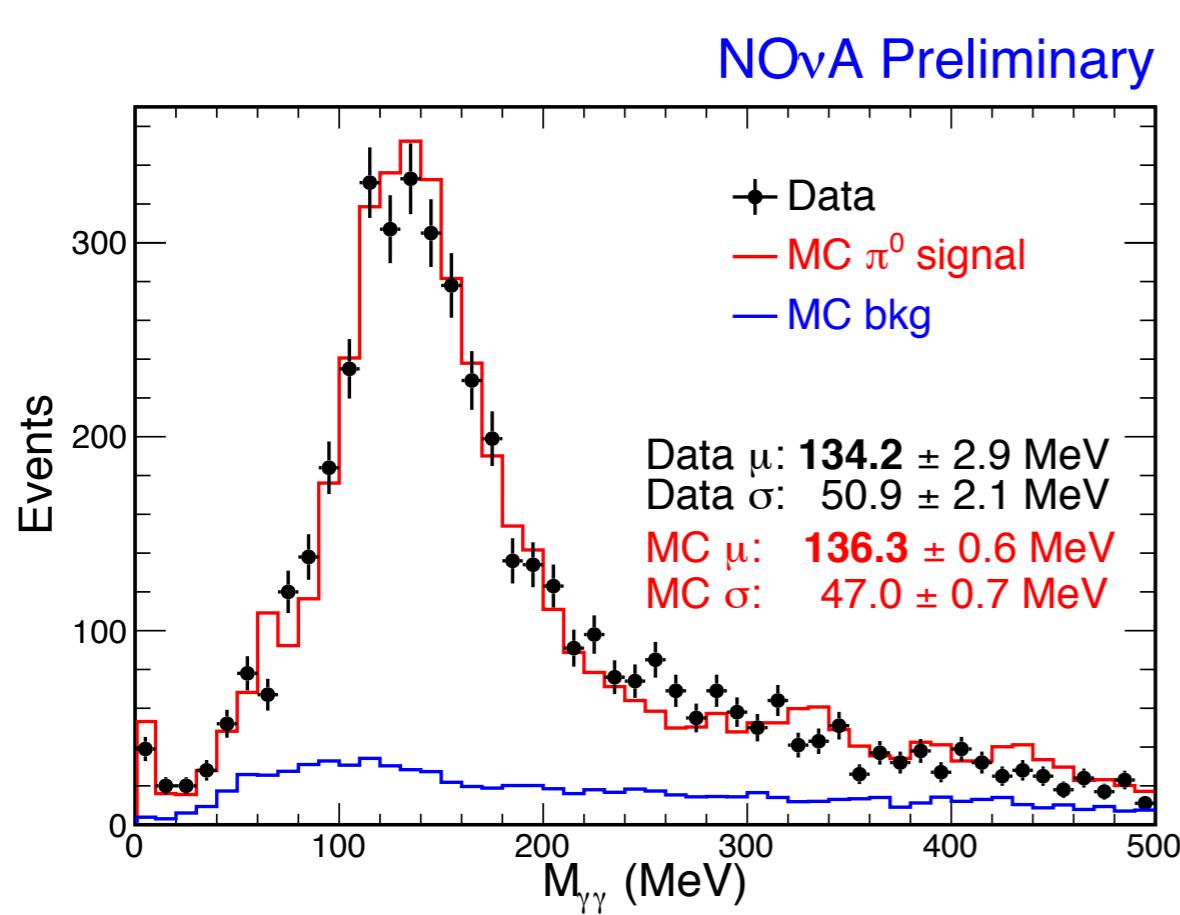
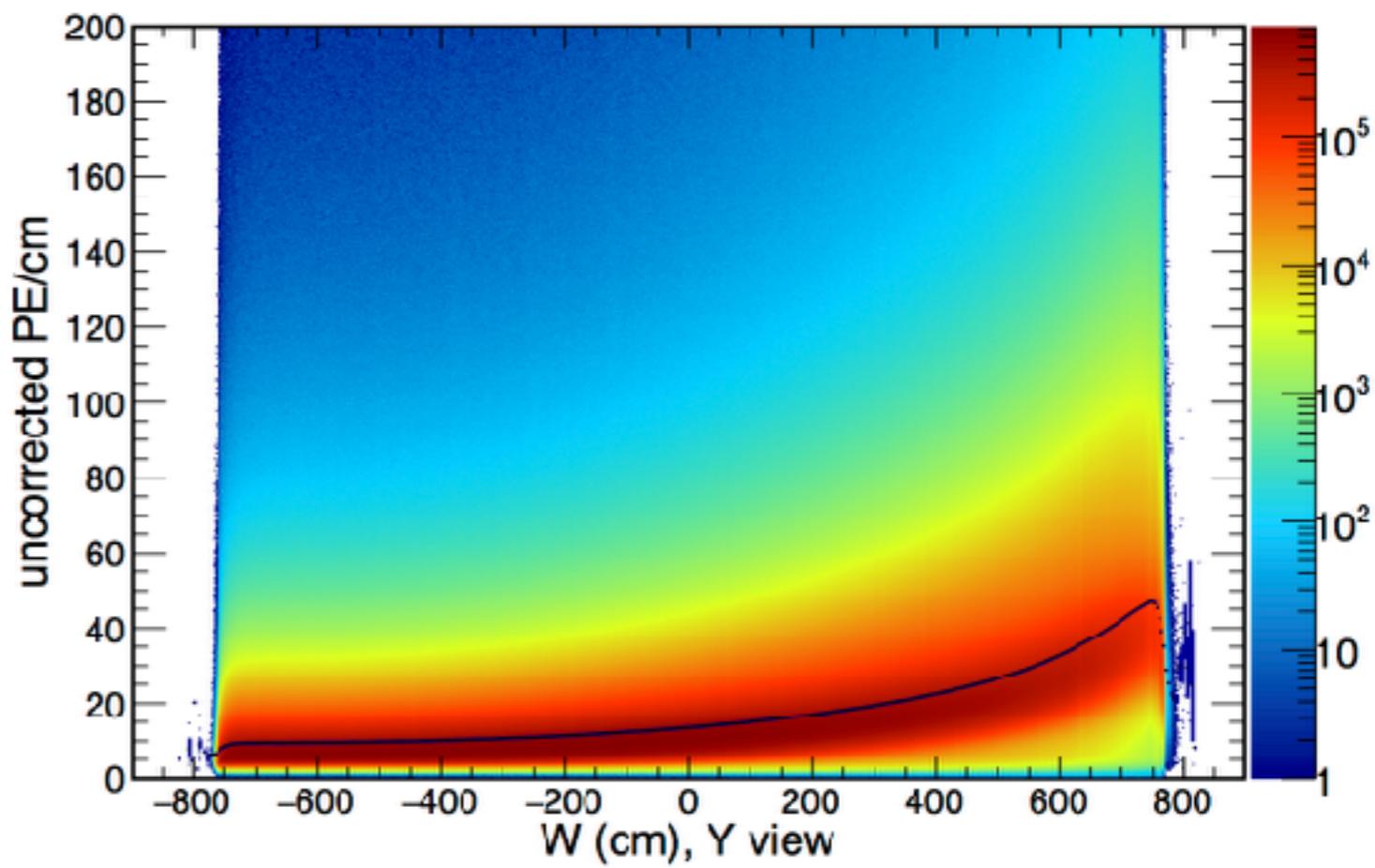


Event Topologies



Detector Calibration

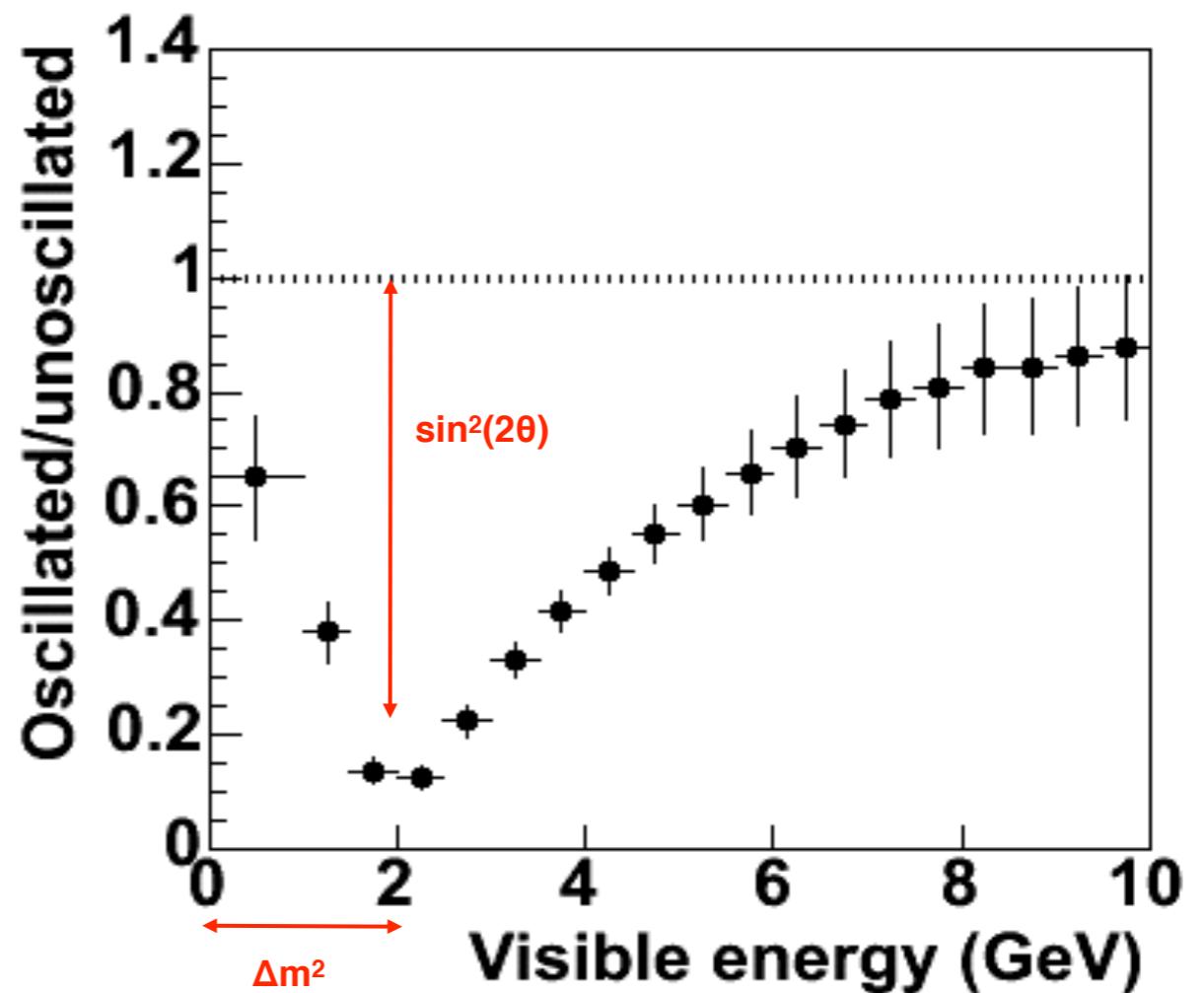
- Cosmic ray muons used to correct attenuation
- Stopping muons used as a standard candle



Muon-neutrino disappearance

Muon-Neutrino Disappearance

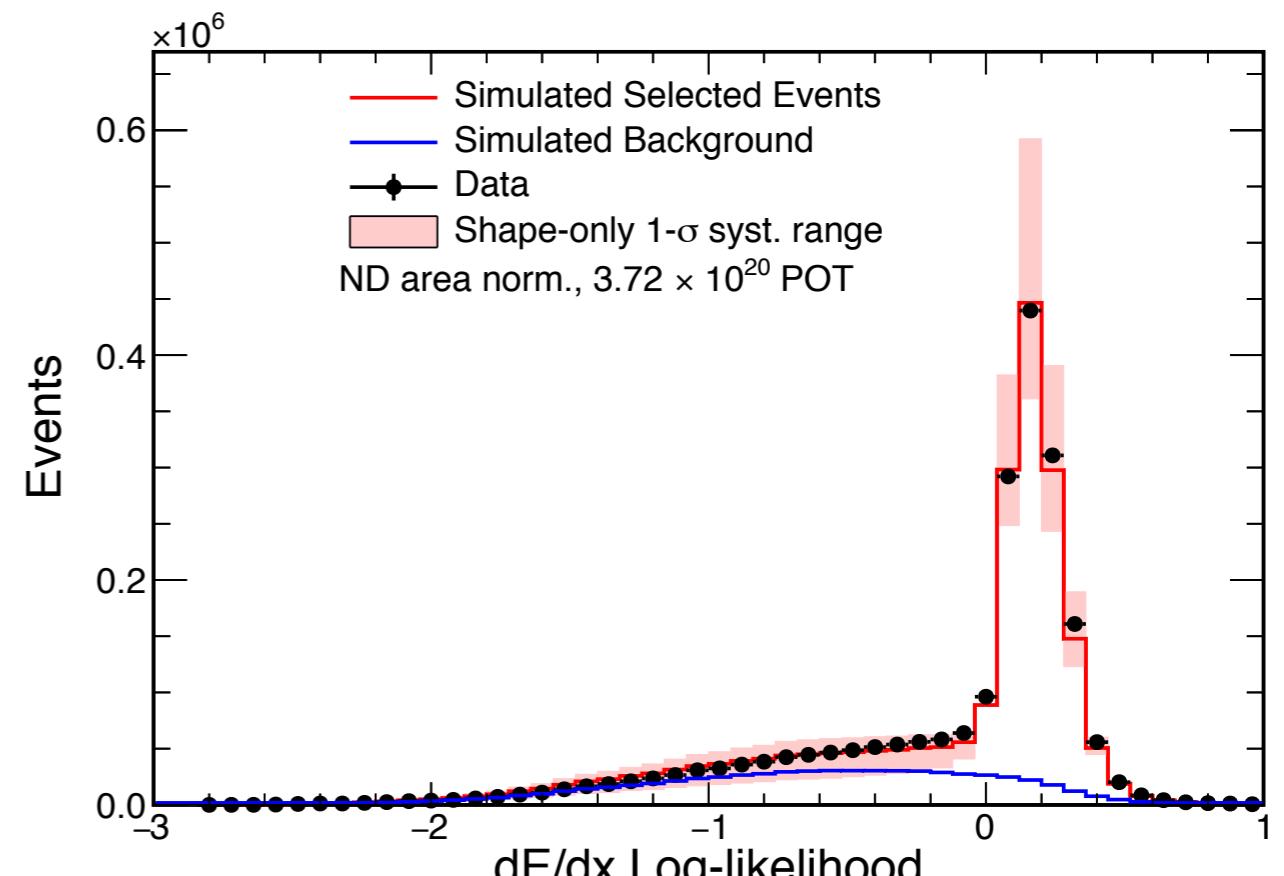
- Two-flavour approximation still basically valid (although analysis uses full three-flavour formalism)
- Measure neutrinos in the ND
- ‘Extrapolate’ measurements to form FD prediction
 - **Taking into account decay kinematics, geometry, efficiencies, purities, energy resolutions, etc.**
- Compare FD data to predictions to find the best fit oscillation parameters



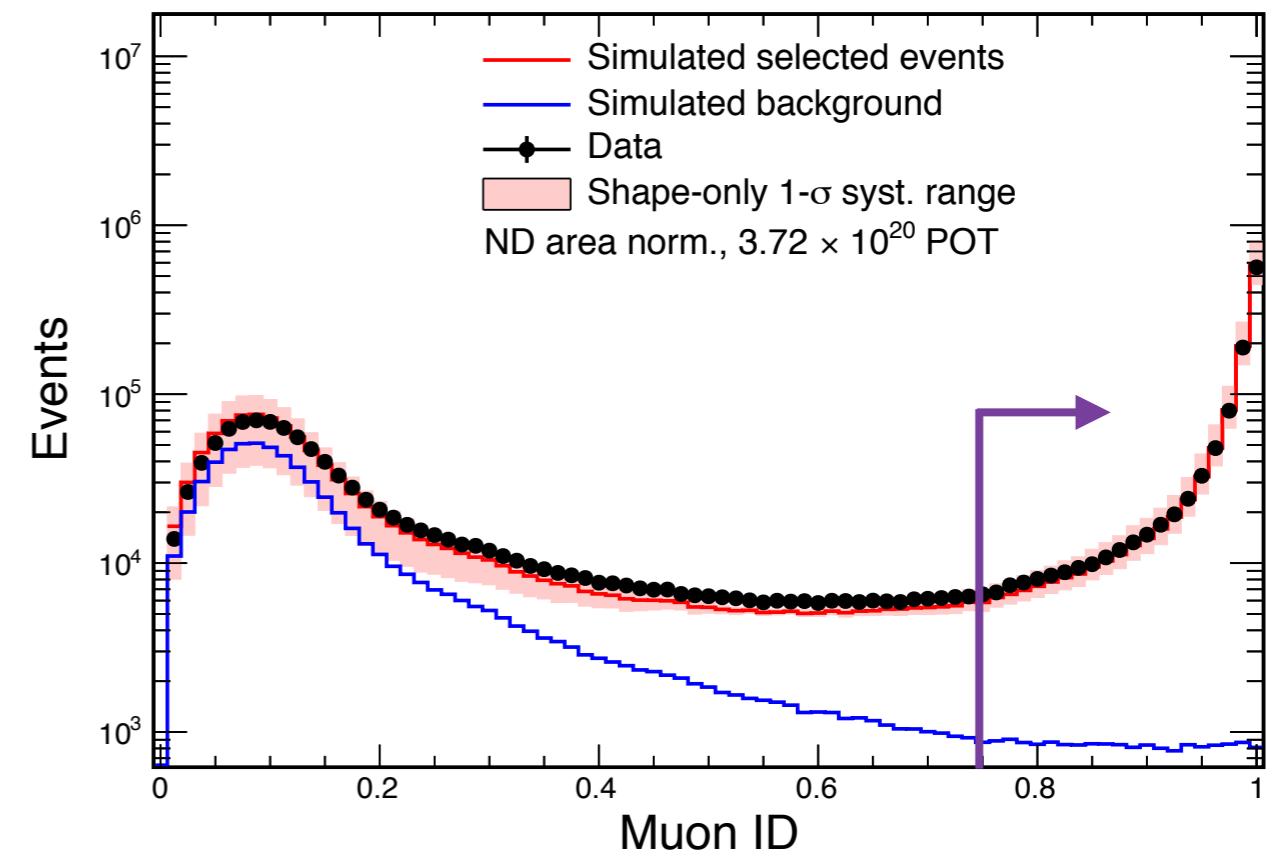
$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta) \sin^2(1.27 \Delta m^2 L / E)$$

Muon-Neutrino Selection

- Separate ν_μ -CC interactions from NC and cosmic-ray backgrounds
- Containment cuts remove activity near walls
- Four variable k-Nearest Neighbour to select muons
 - **Track length**
 - **dE/dx along track**
 - **Scattering along track**
 - **Track-only plane fraction**
- Selection is 81% efficient and 91% pure

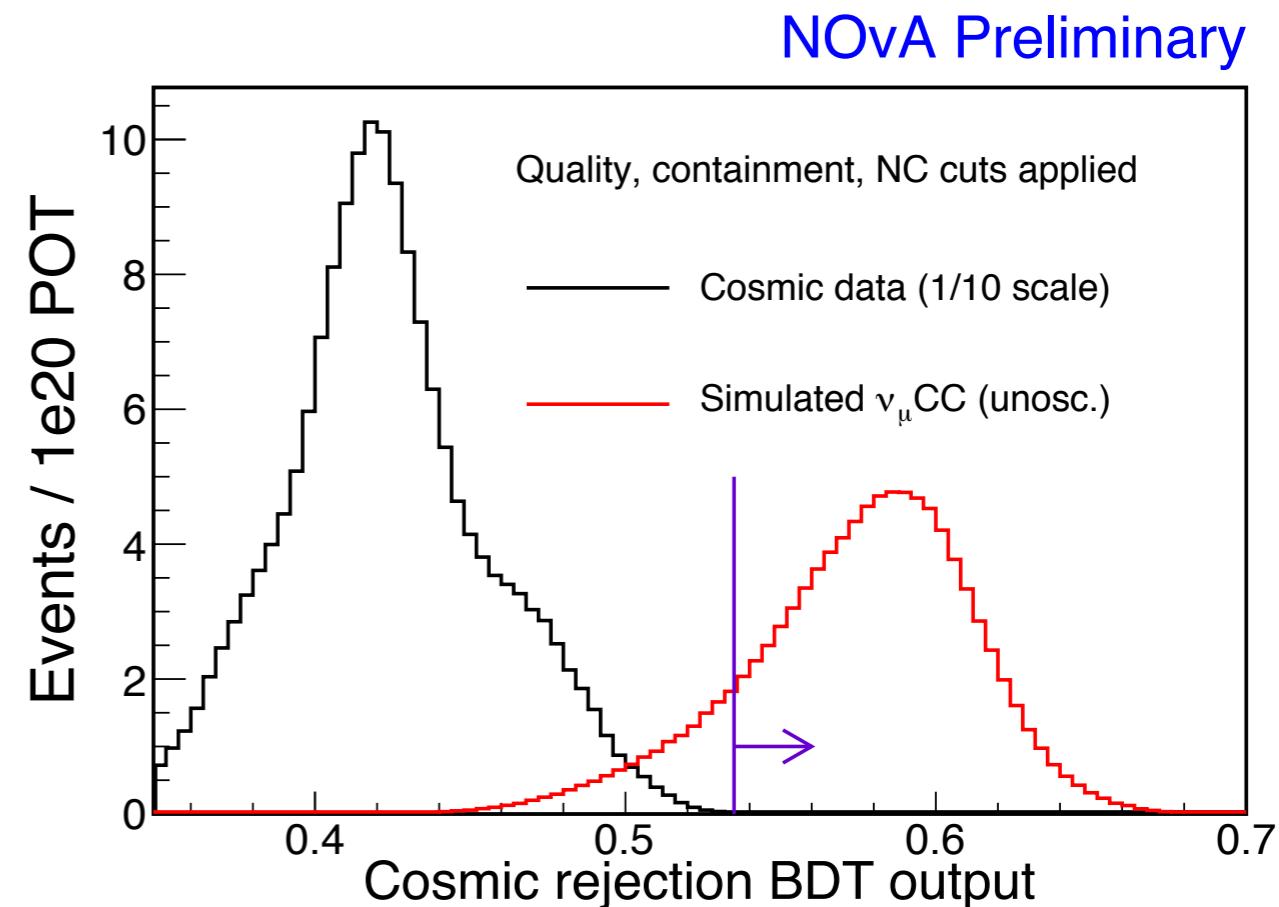
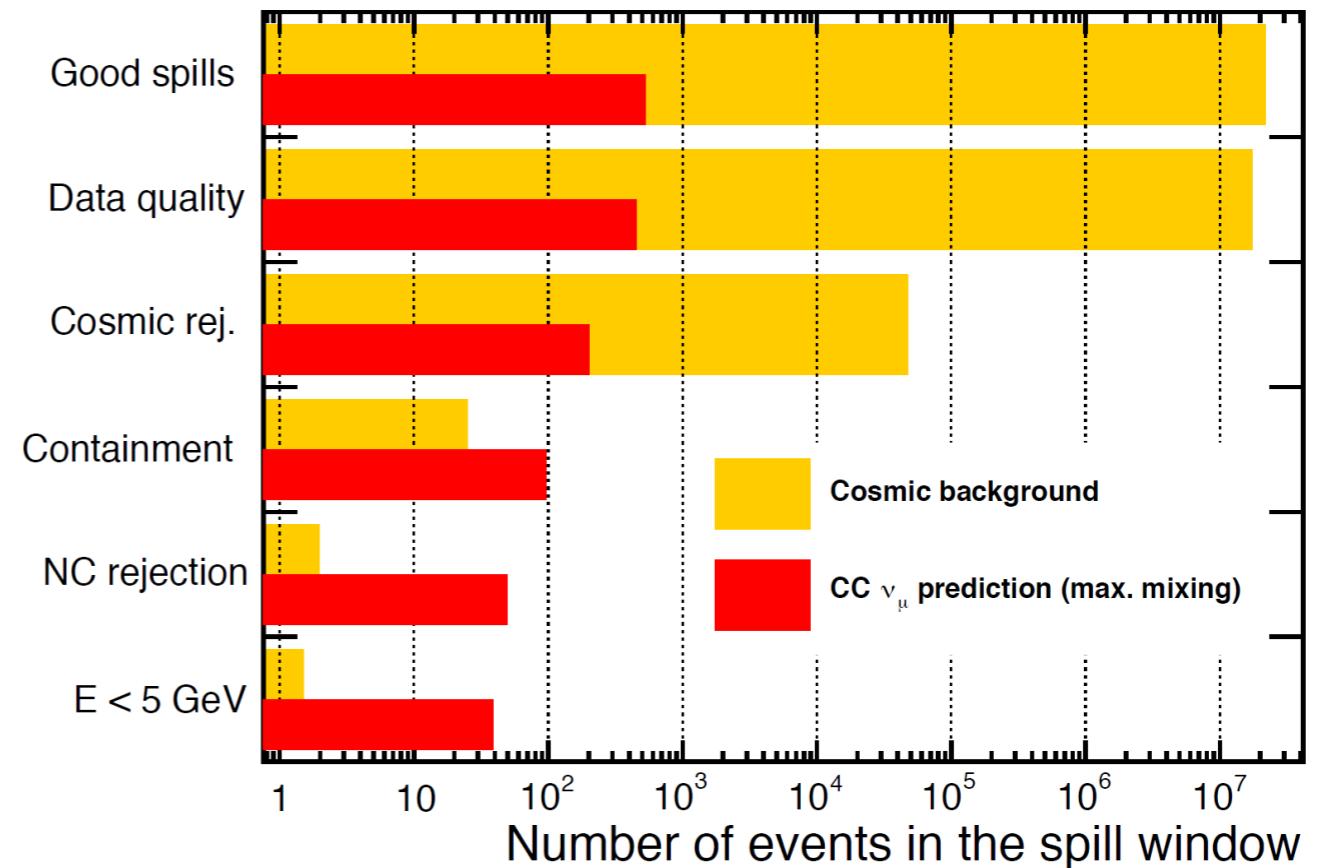


NOvA Preliminary



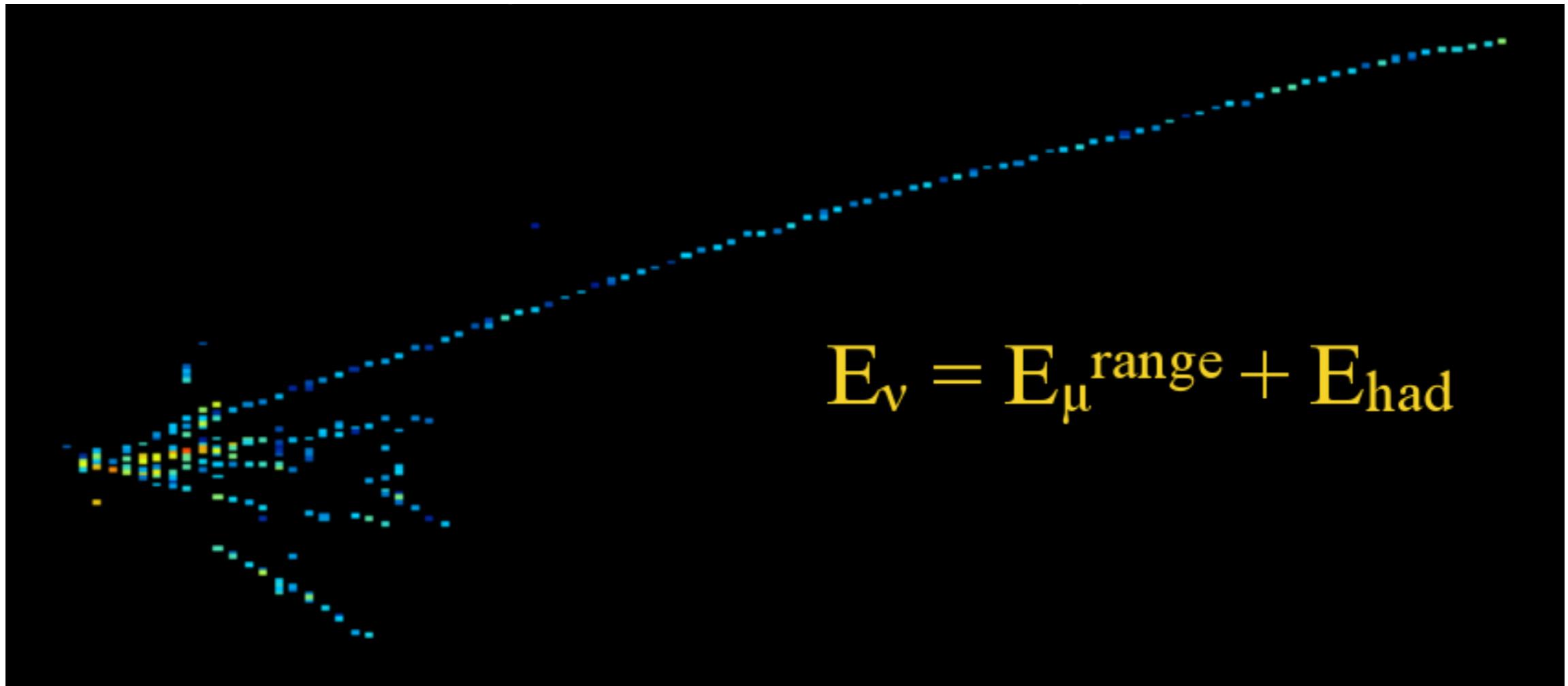
Cosmic Rejection

- Far Detector is on the surface and sees 150 kHz of cosmic induced events
- 10 μ s beam window every 1.3s reduces background by 10^5
- Additional factor of 10^7 rejection achieved from event topology and a boosted decision tree (BDT) based on:
 - **track direction**
 - **start/end points of track**
 - **track length**
 - **energy**
 - **number of hits**
- Predict 2.7 cosmic background events



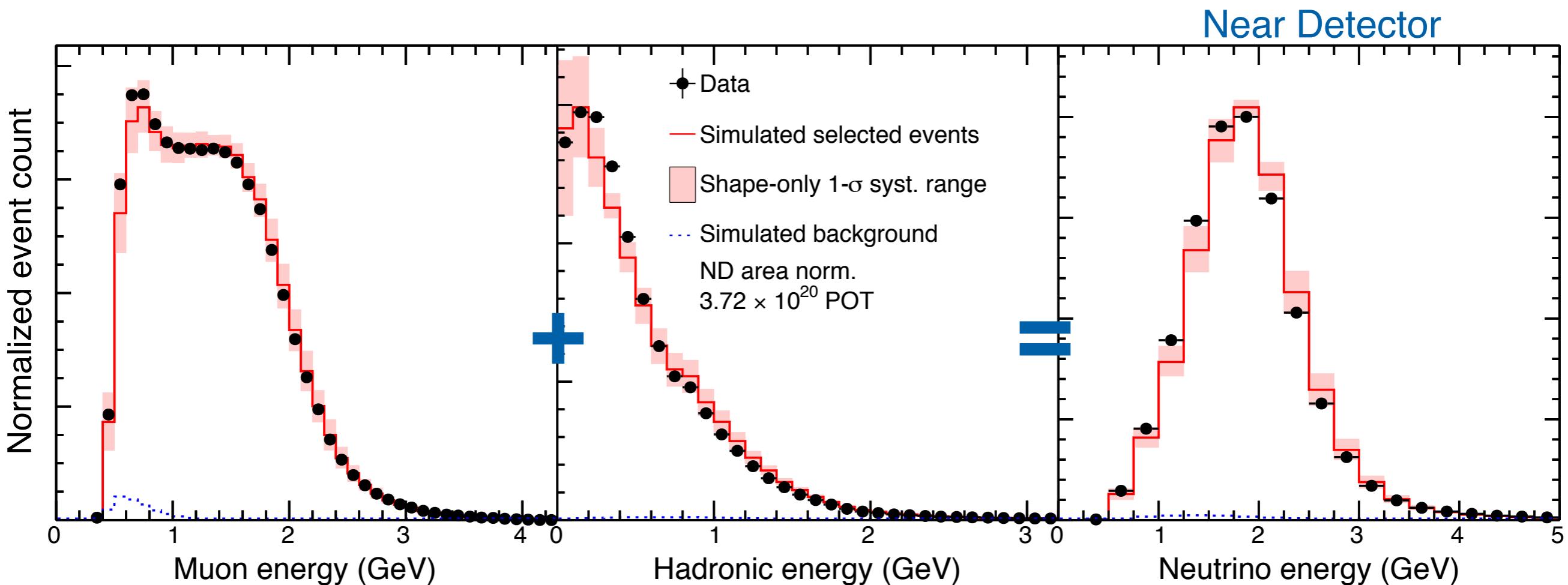
Energy Estimation

- Muon dE/dx used in length-to-energy conversion
- Hadronic energy estimated from calorimetric sum of non-muon hits
- ~7% resolution on neutrino energy



Energy Estimation

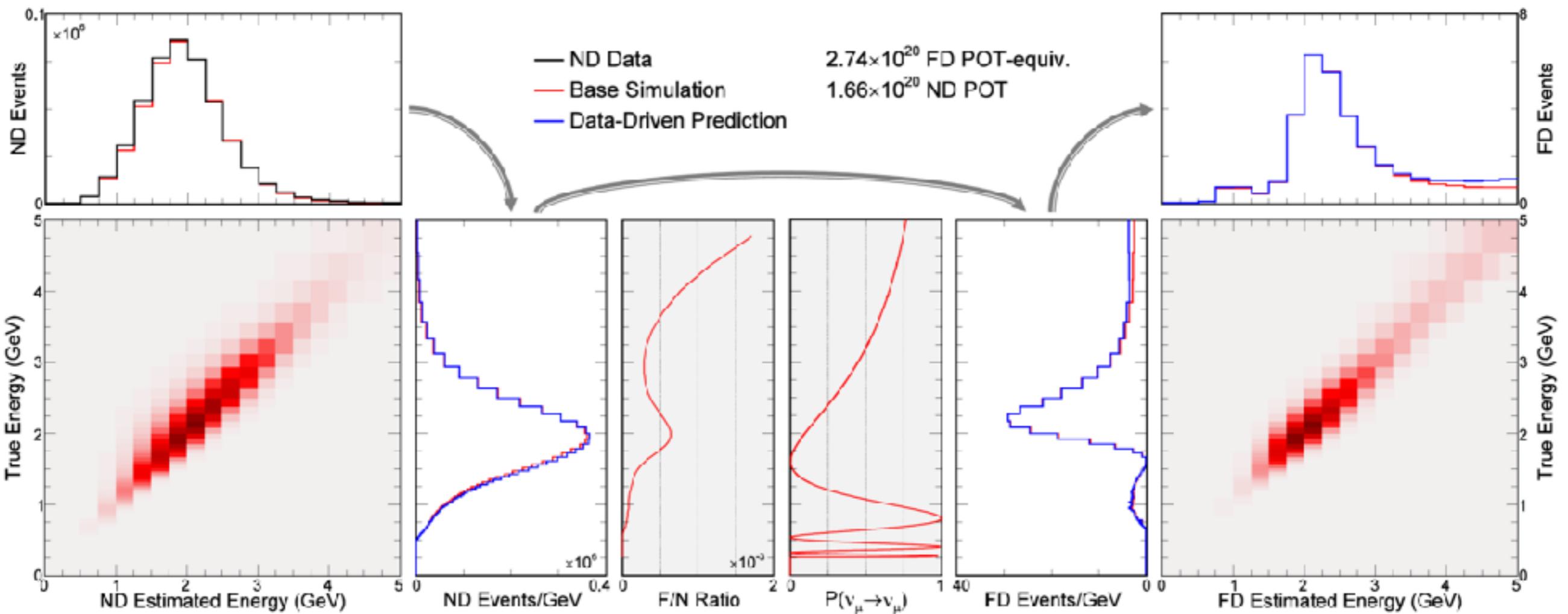
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$$E_\nu = E_\mu(L_\mu) + E_h$$

Extrapolation

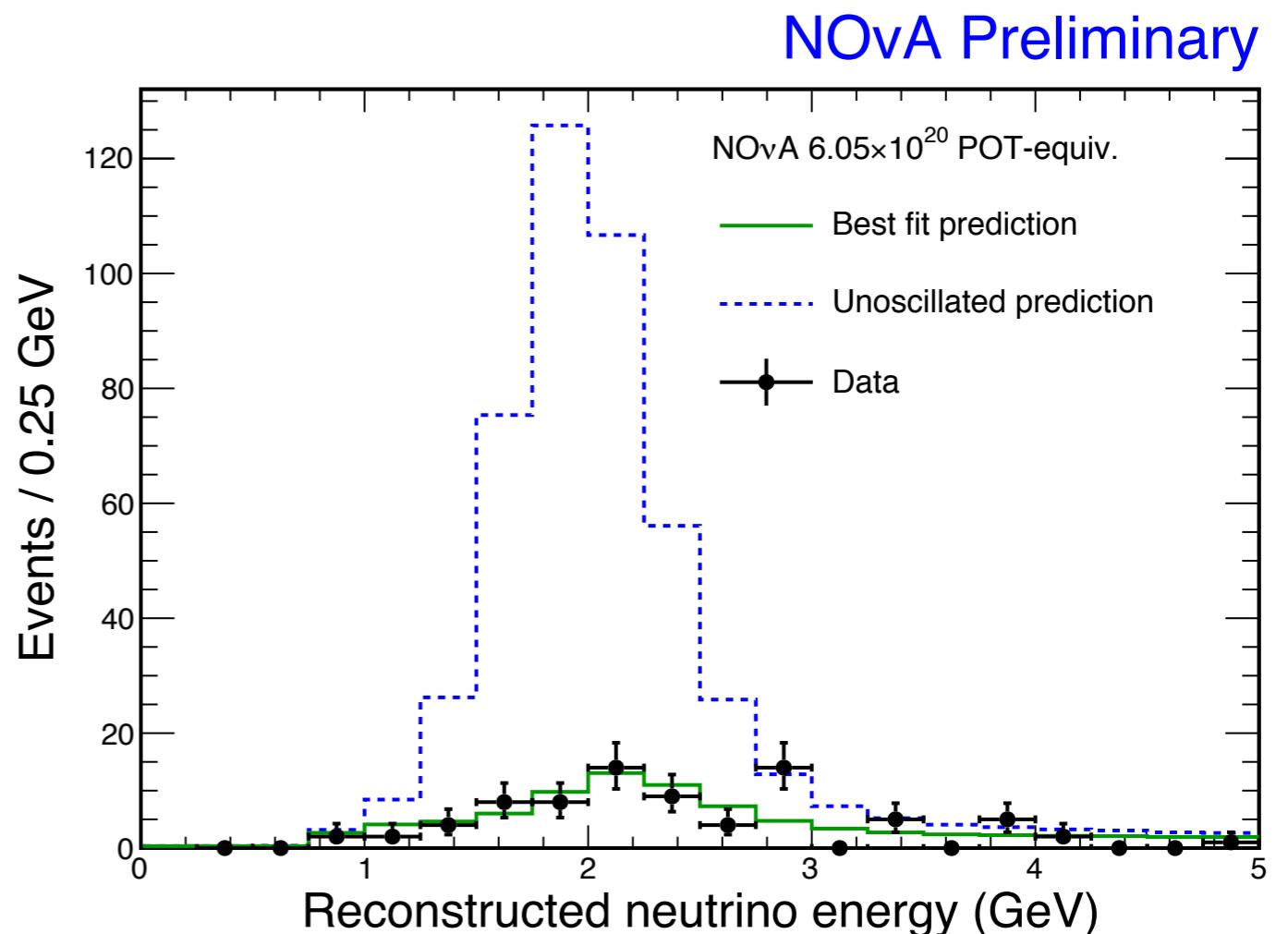
- Use high statistics ND data/MC to adjust prediction at FD
 - Translate ND data/MC observation to true energy
 - Oscillate ratio to the FD
 - Smear back into reconstructed energy



Muon-Neutrino Disappearance

[10.1103/PhysRevLett.118.151802](https://doi.org/10.1103/PhysRevLett.118.151802)

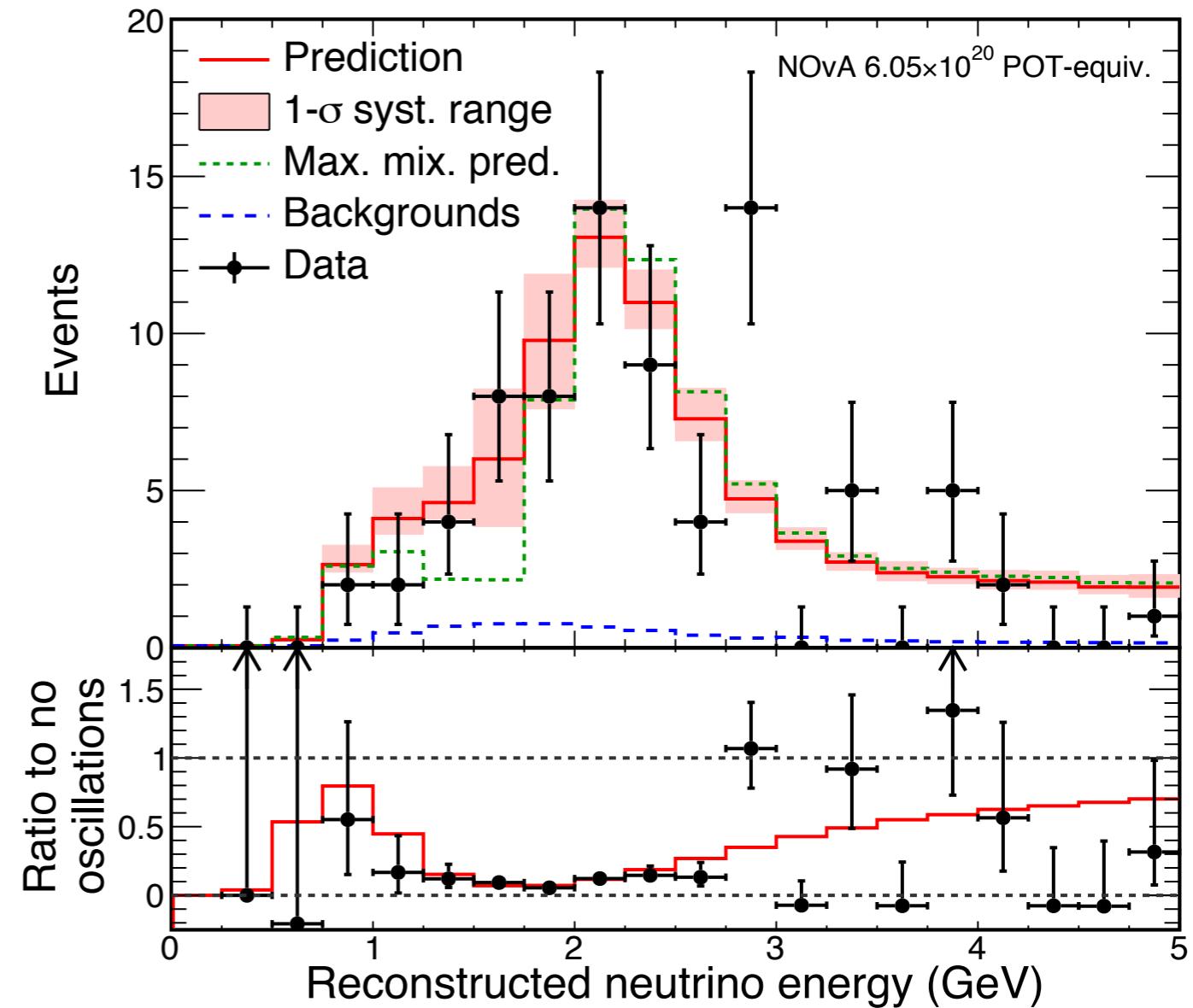
- Using 6.05×10^{20} POT equivalent
- 473 ± 30 events predicted in the absence of oscillations
- Observed 78 events
- 82 events predicted at the best fit point including 3.7 beam background and 2.9 cosmic induced events

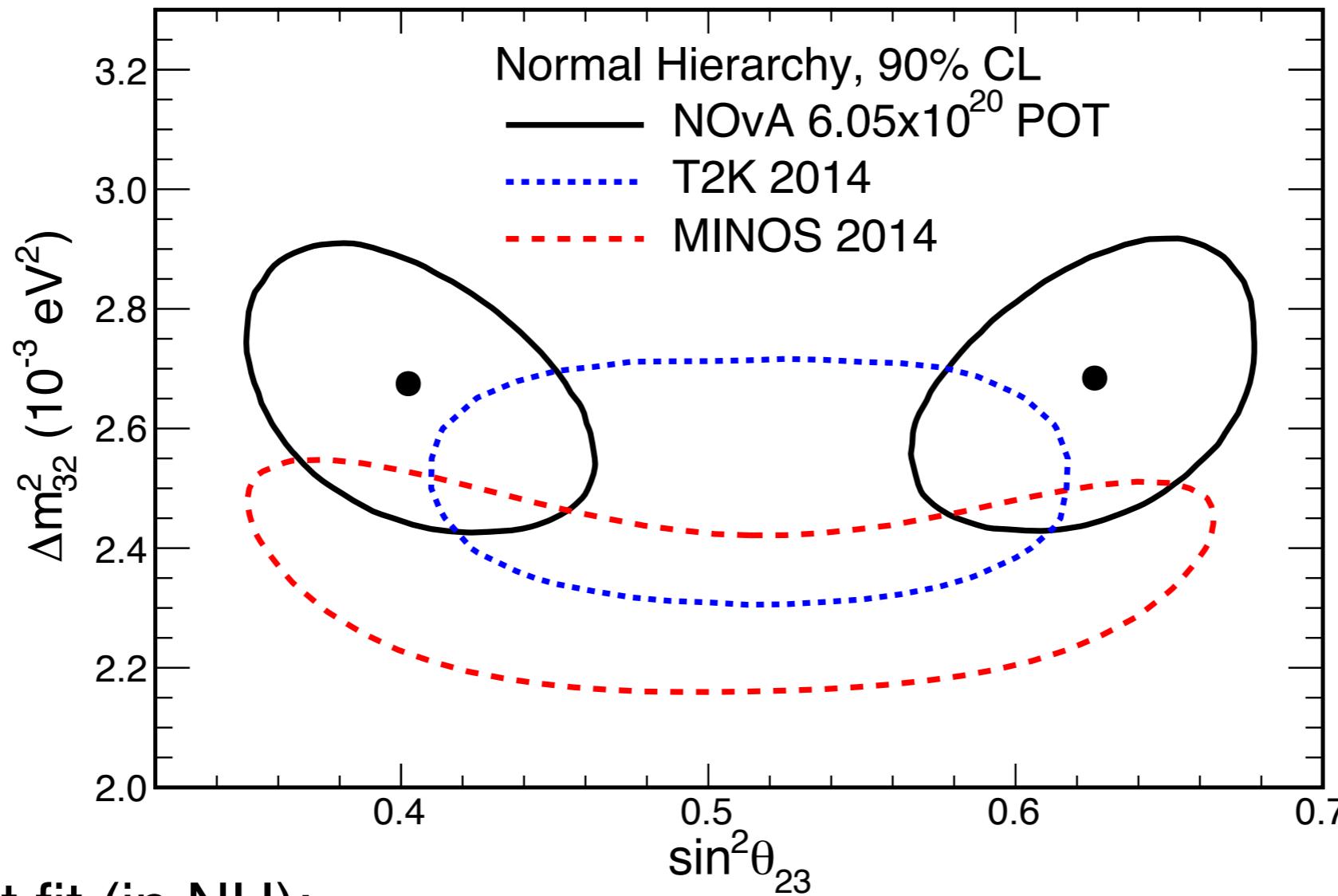


Muon-Neutrino Disappearance

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Best fit (in NH):

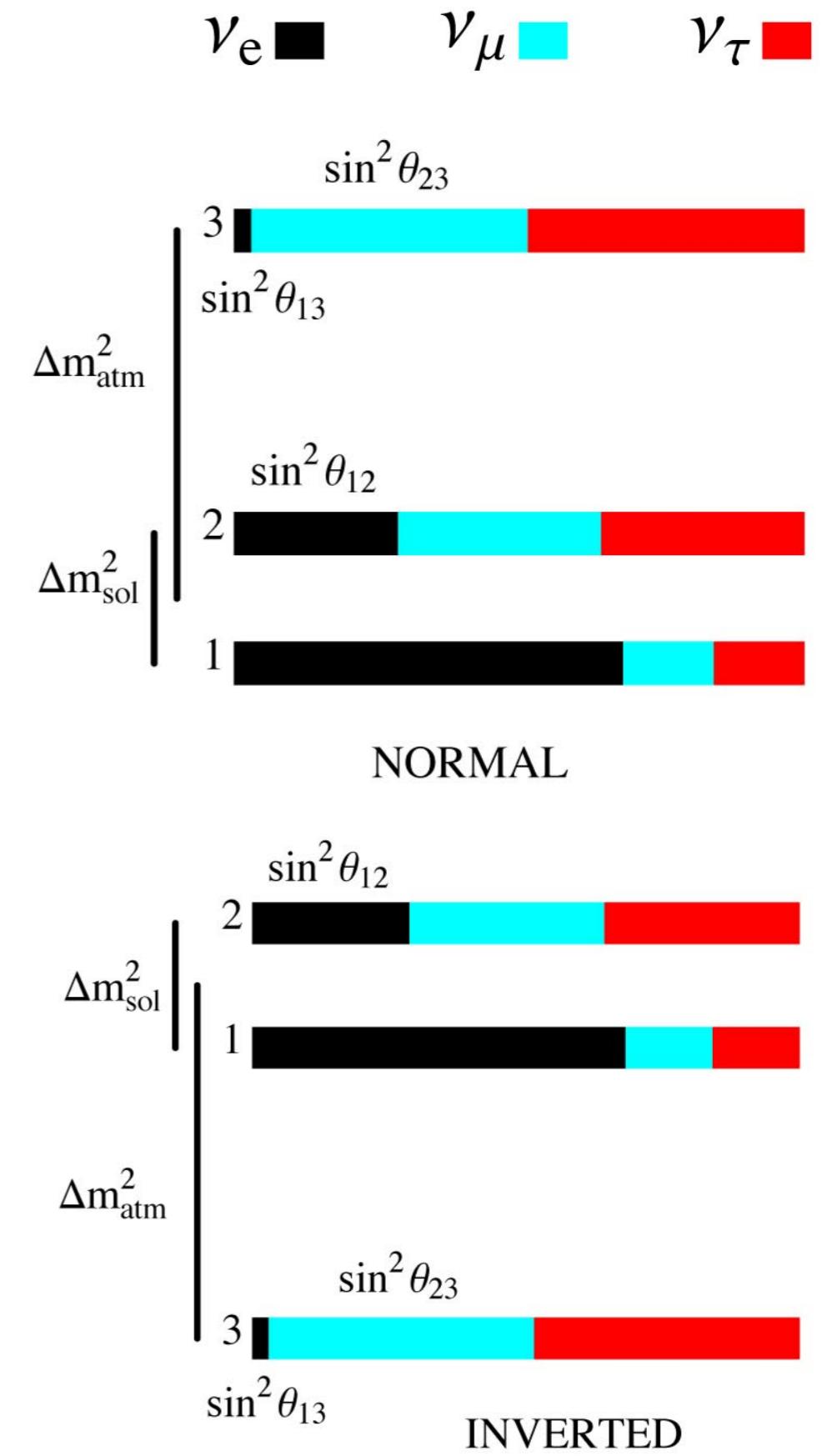
$$|\Delta m_{32}^2| = 2.67 \pm 0.11 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} = 0.404^{+0.030}_{-0.022} (0.624^{+0.022}_{-0.030})$$

- Maximal-mixing disfavoured at 2.6 sigma
- Interesting tension between NOvA and T2K, new results eagerly anticipated

Electron-Neutrino Appearance

- Electron-neutrino appearance is a sub-dominant oscillation mode at the NOvA L/E
- Matter effects matter (almost 3 times longer baseline than T2K)
- Sensitive to
 - **Mass hierarchy**
 - **CP violating phase**
 - **Octant of θ_{23}**



New Classification Algorithm

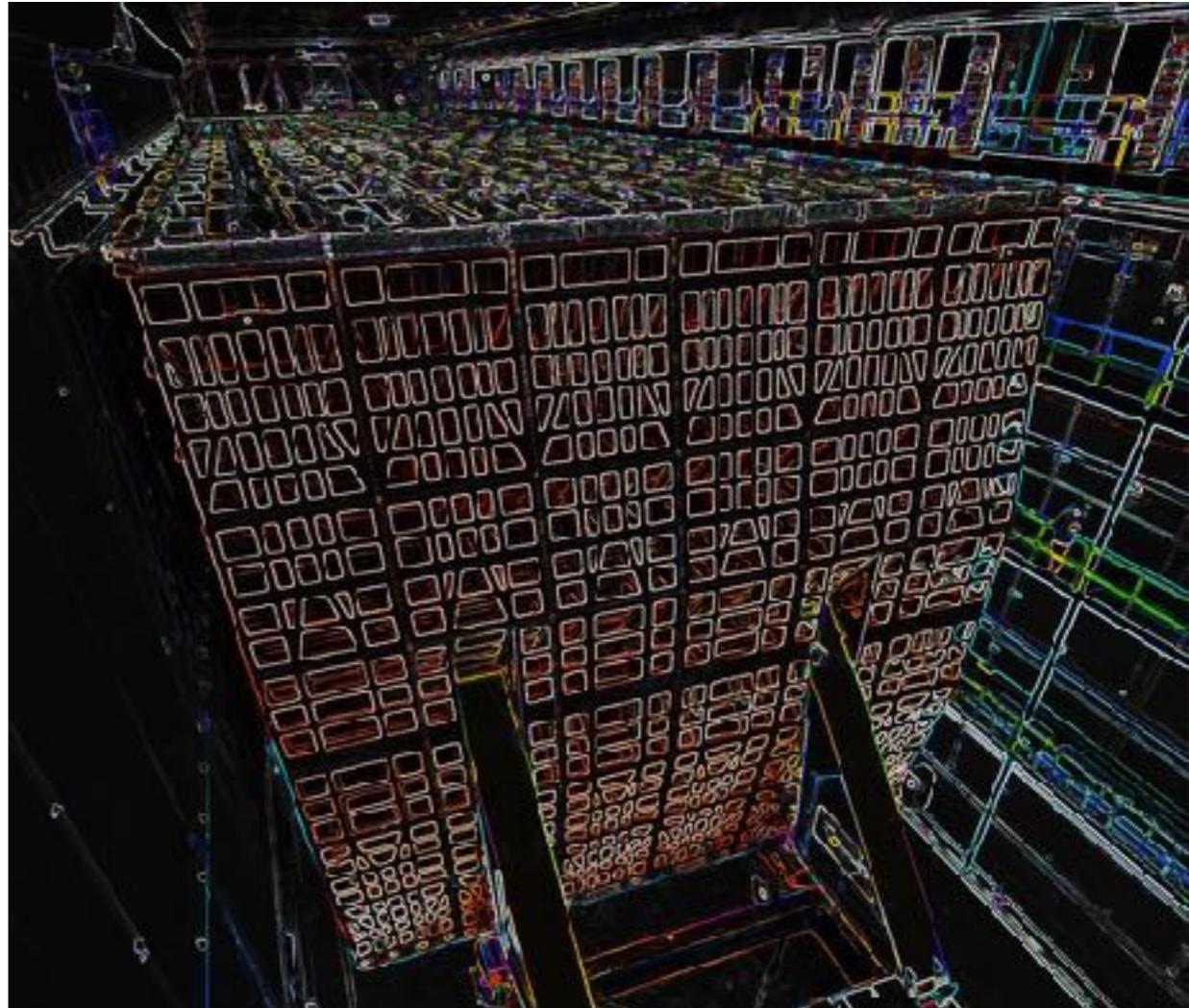
35	40	41	45	50
40	40	42	46	52
42	46	50	55	55
48	52	56	58	60
56	60	65	70	75



0	1	0
0	0	0
0	0	0



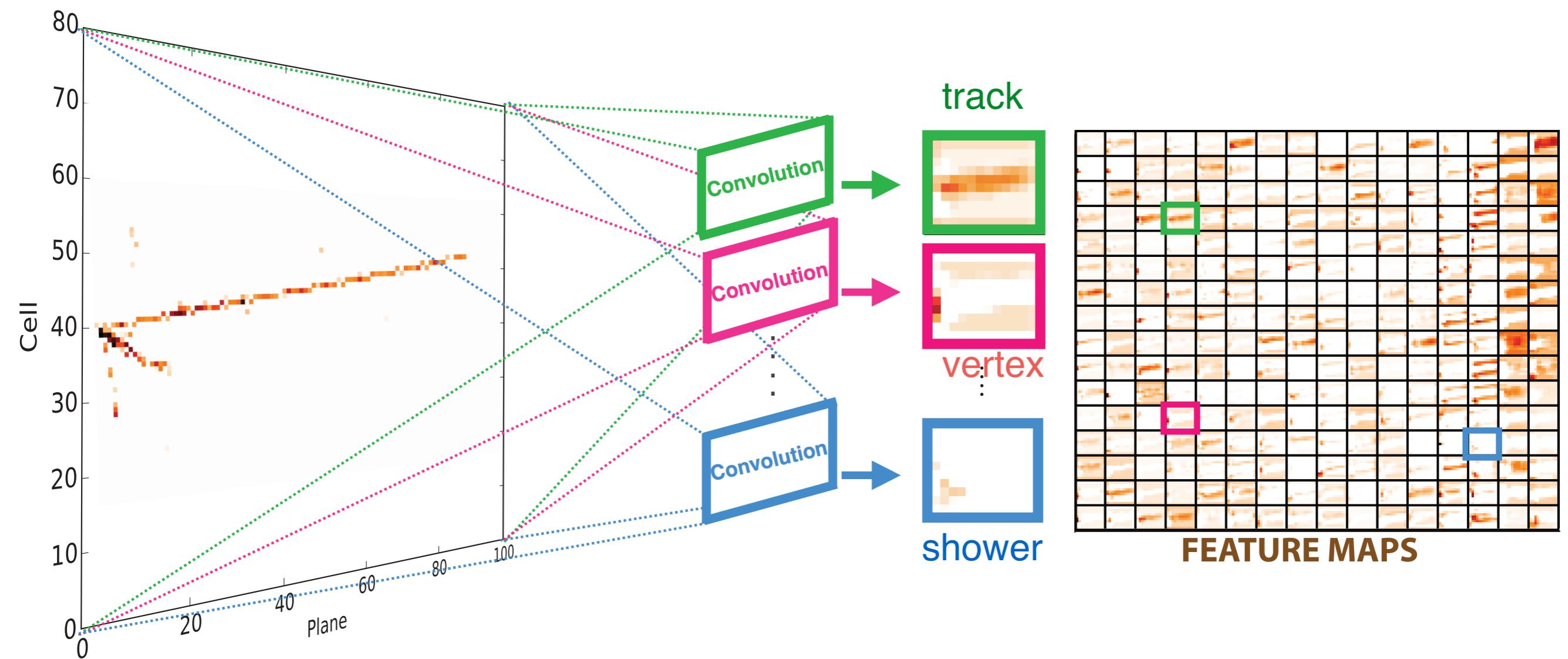
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- Take advantage of recent advances in machine learning/computer vision
- Deep networks extract complex features from input data, GPUs greatly improve training time
- Inputs to the network are pixels in image
- Apply convolutional kernels to pull out event features
 - **Image on right has gone through a convolutional kernel looking for edges**

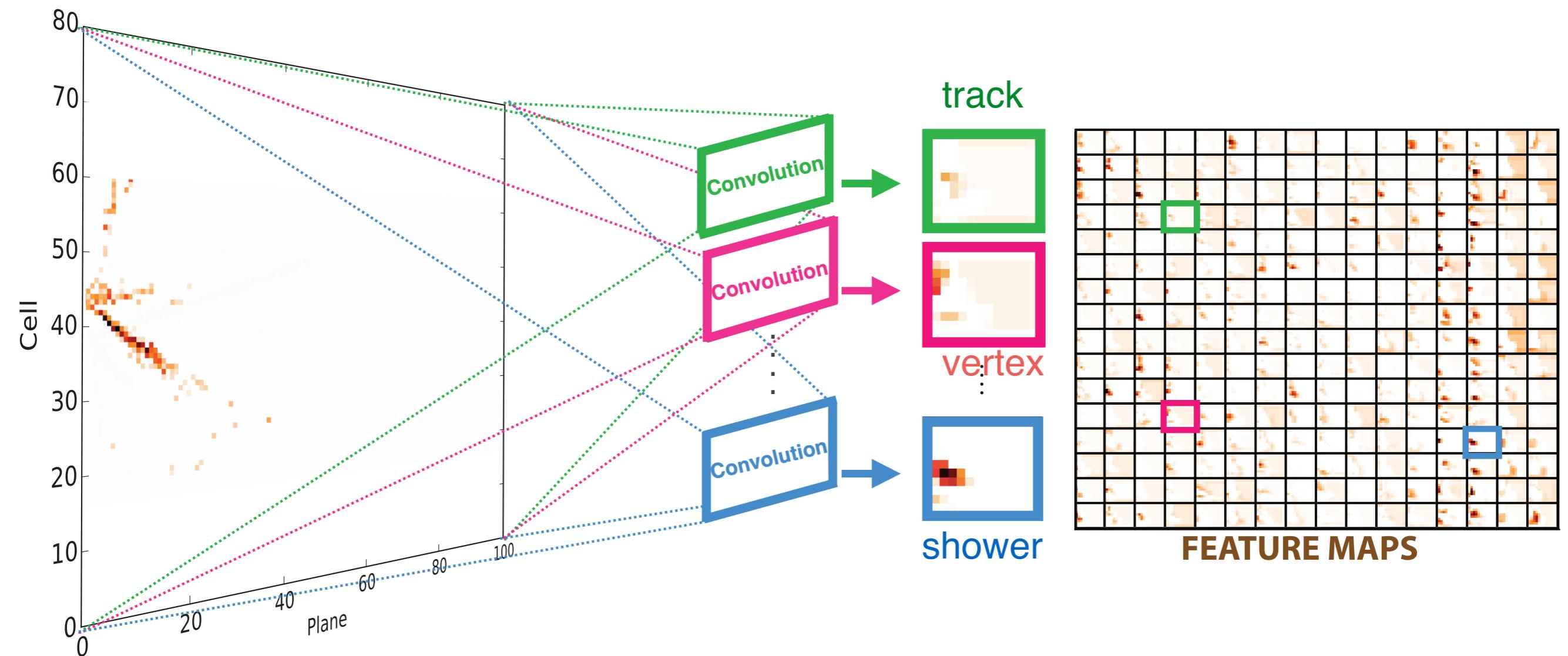
Convolutional Visual Network (CVN) Selection

- Showing a muon neutrino interaction and the first layer of feature maps extracted from the convolutional kernels
- Strong feature here is the track

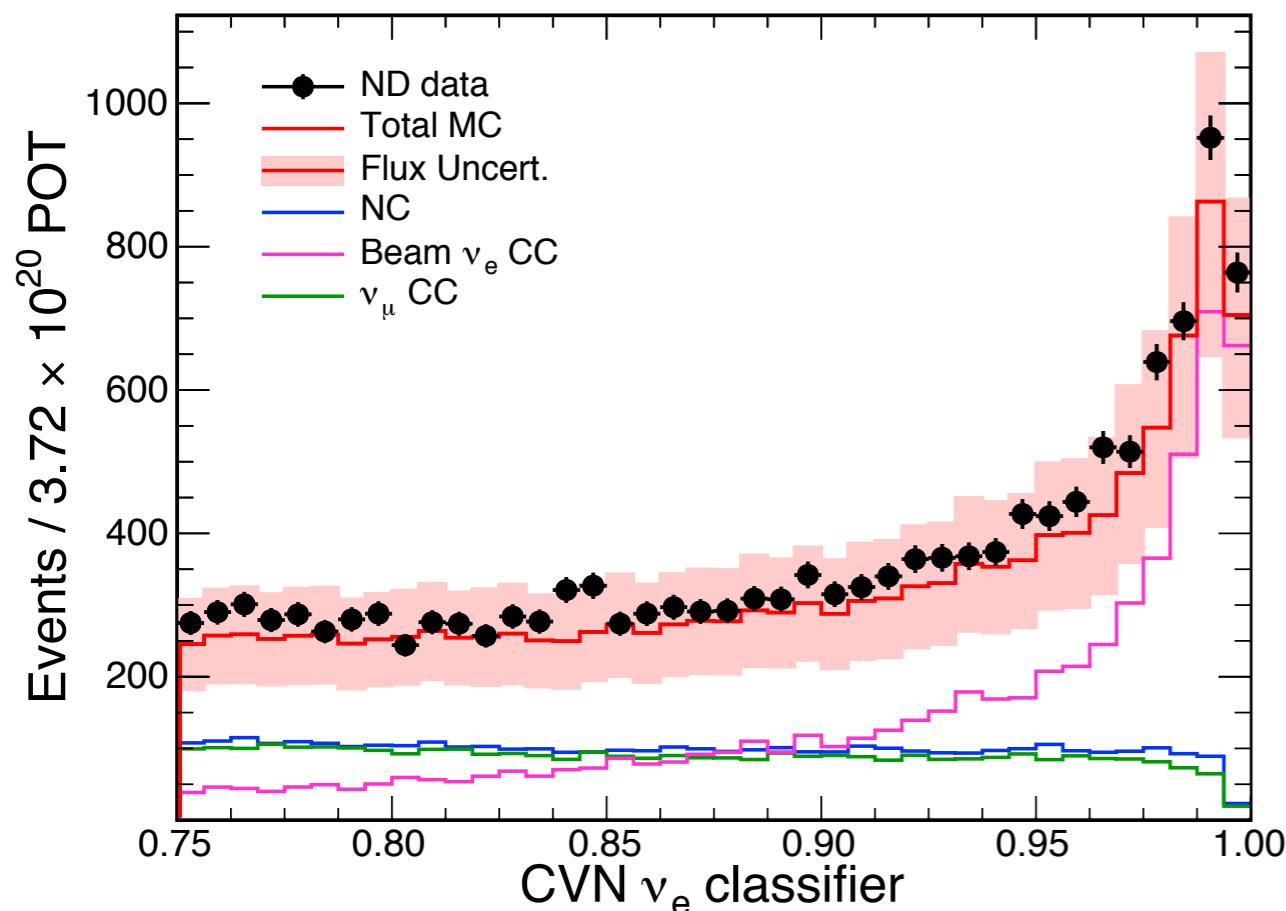


Convolutional Neural Networks

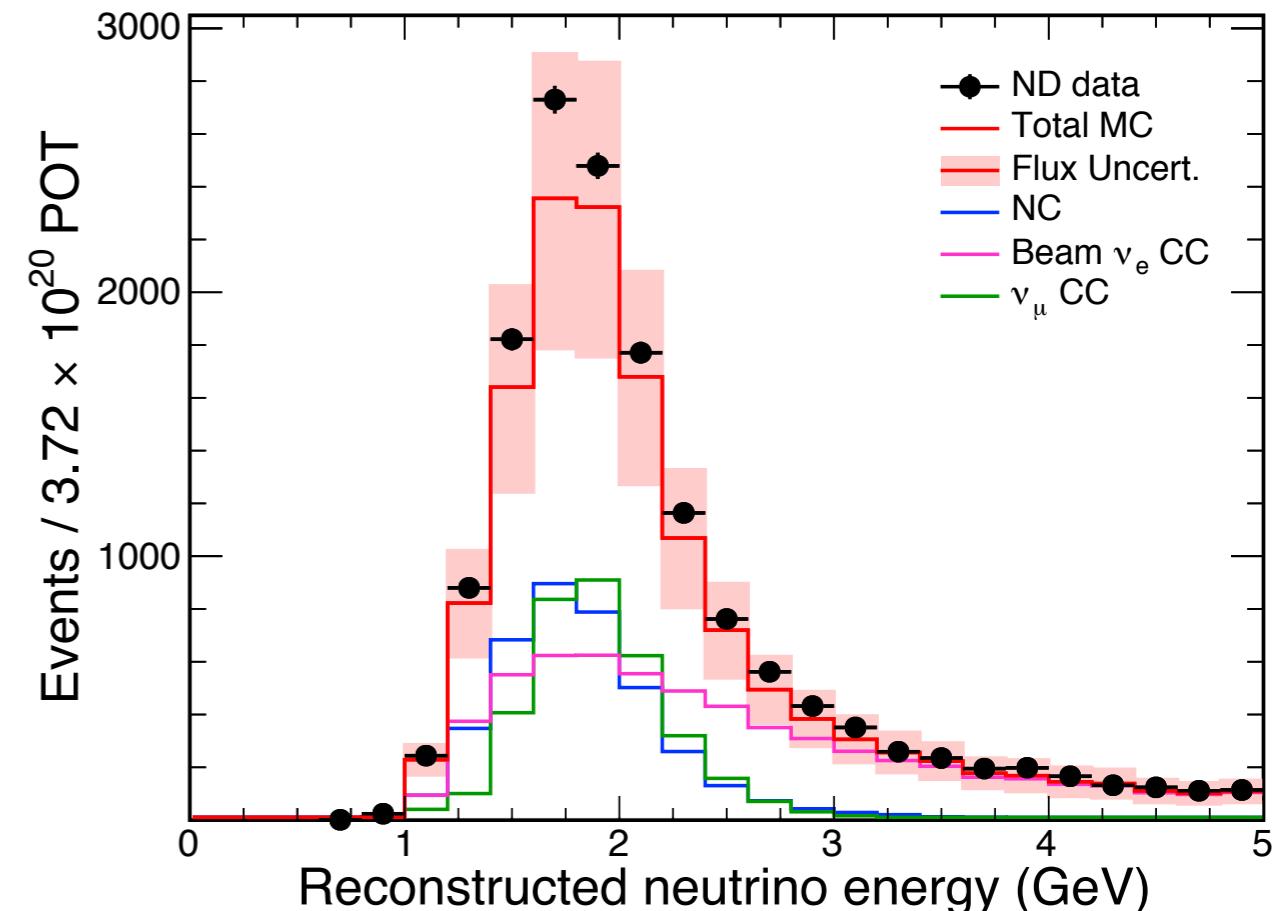
- Showing an electron neutrino interaction and the first layer of feature maps extracted from the convolutional kernels
- The strong features extracted are the shower as opposed to the track



NOvA Preliminary

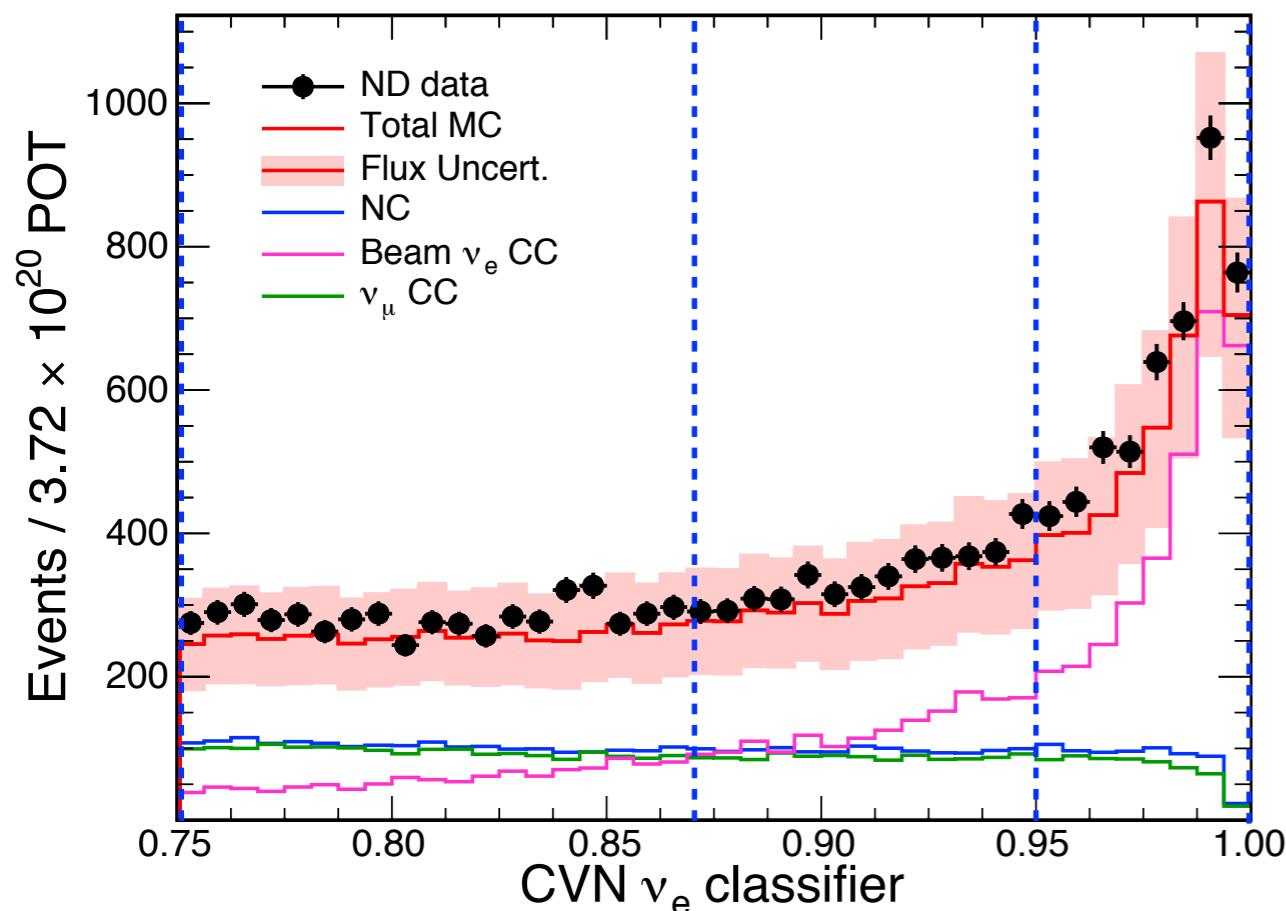


NOvA Preliminary

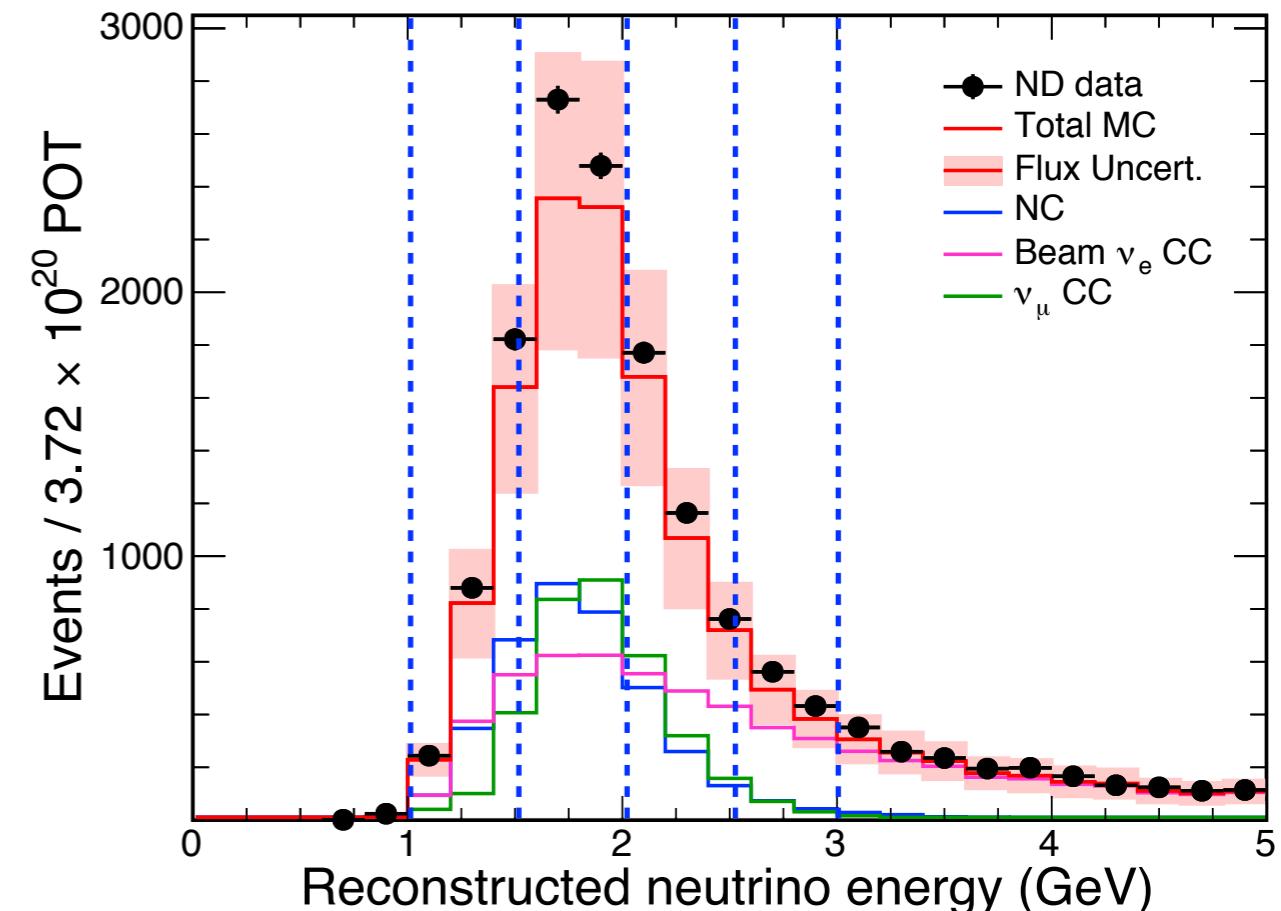


- 73% ν_e CC selection efficiency, 76% purity with CVN classifier
- Good ND Data/MC agreement
- CVN provides better cosmic rejection and similar systematics to 2015 classifiers

NOvA Preliminary



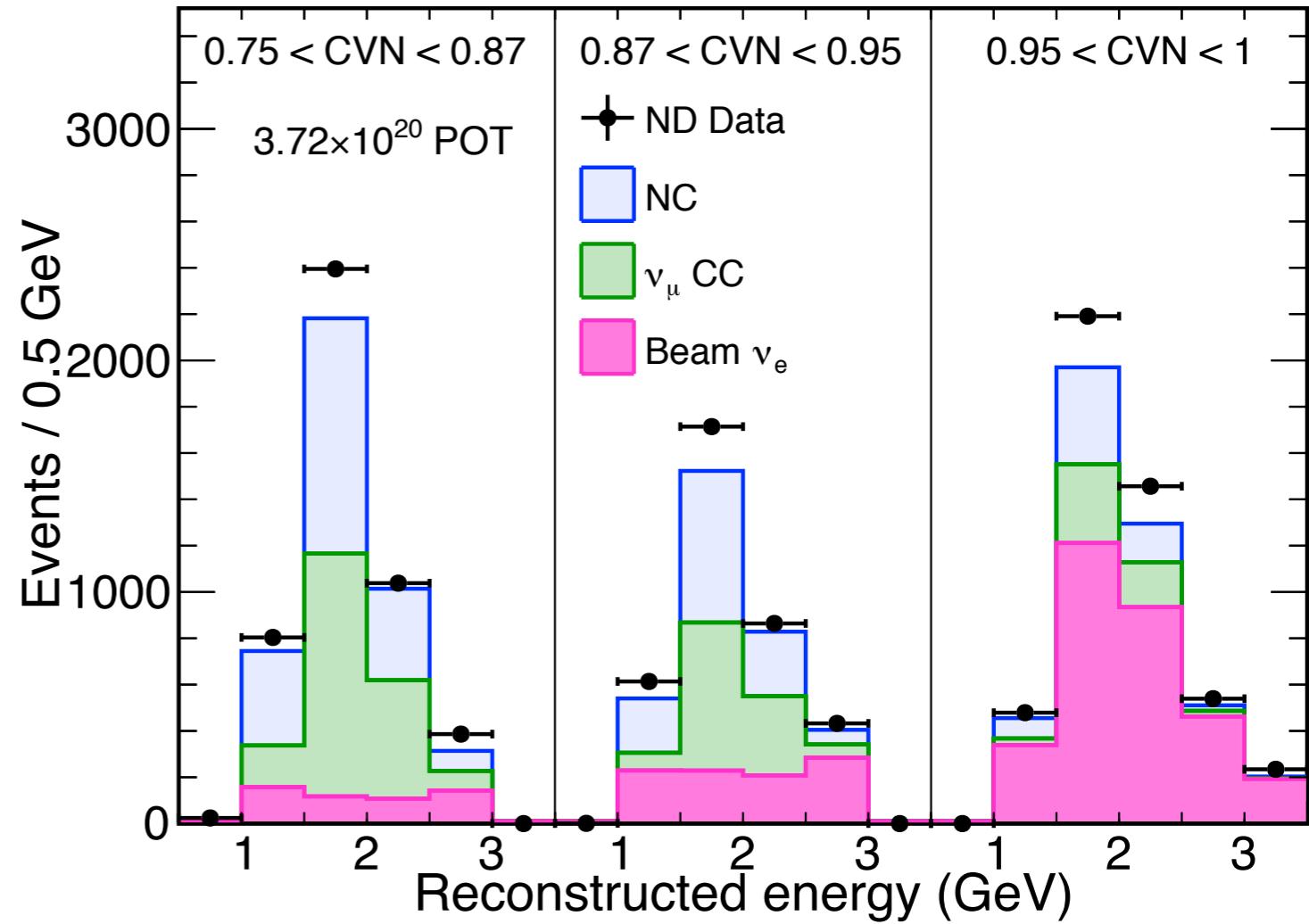
NOvA Preliminary



- Bin analysis in four bins of energy and three of CVN

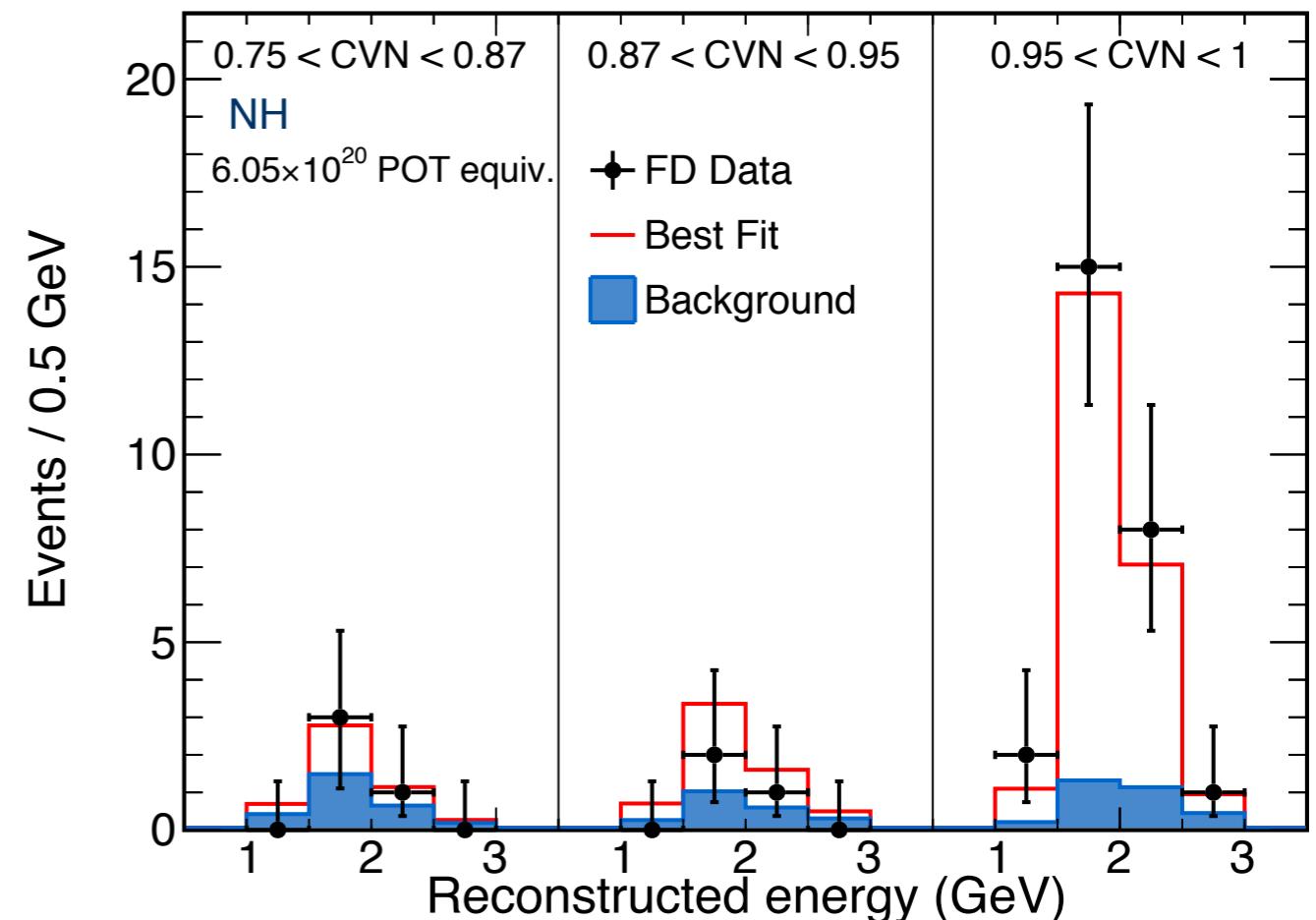
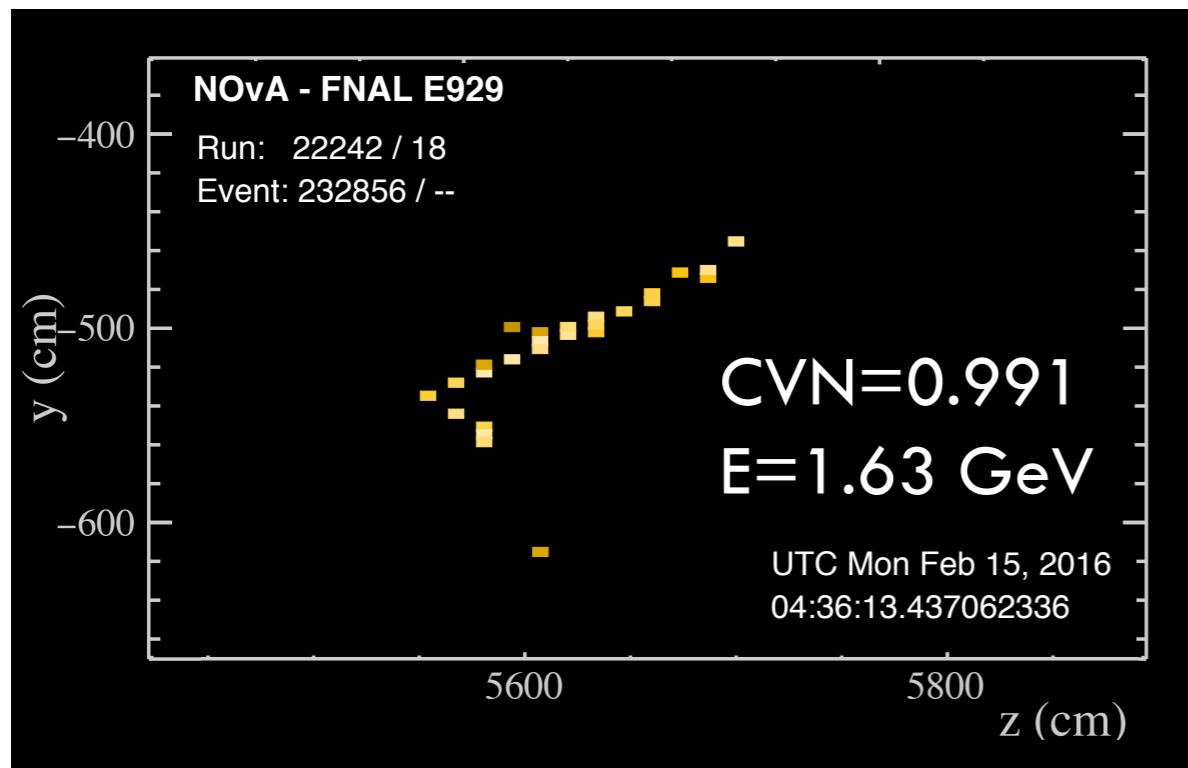
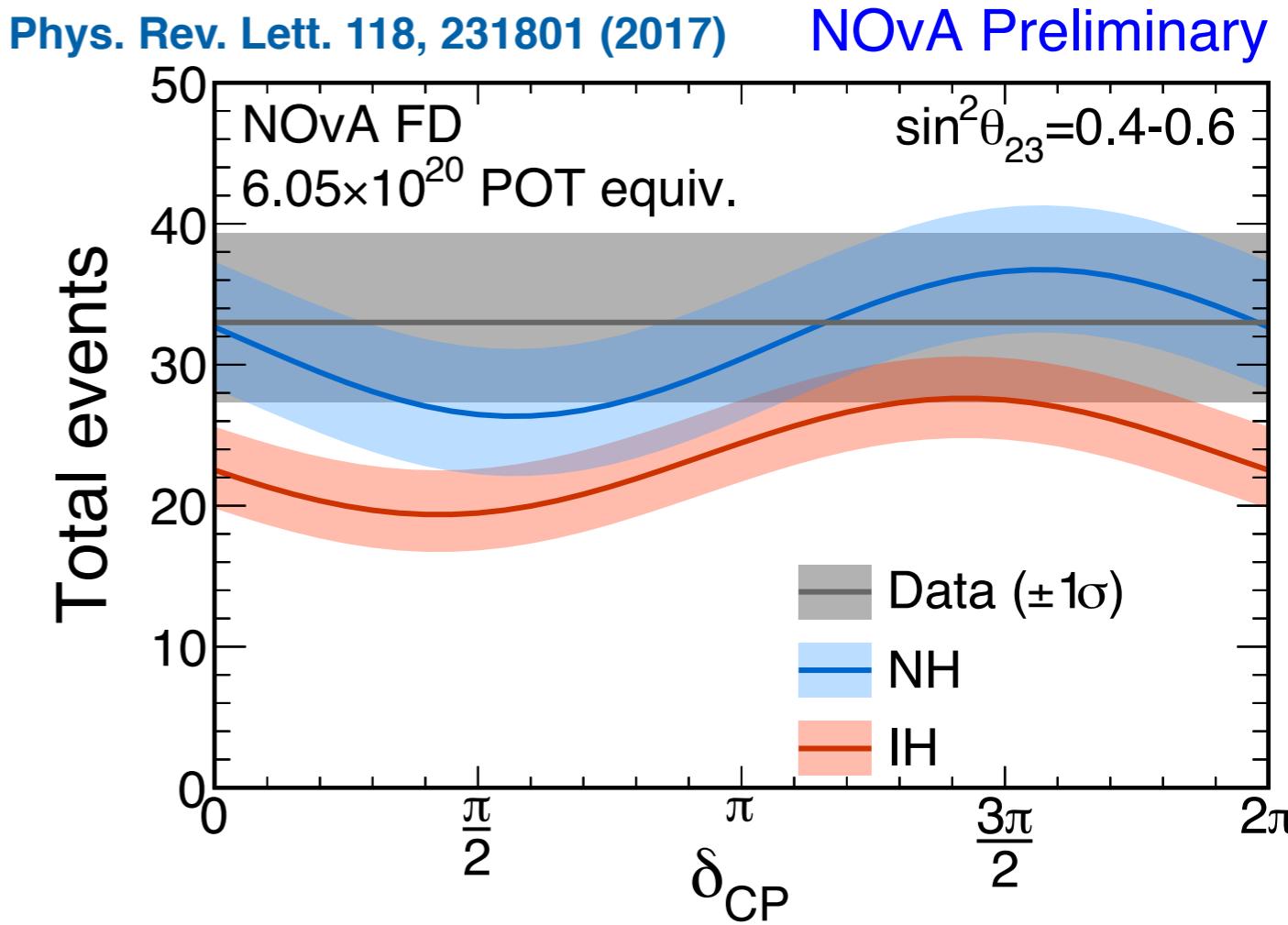
Data Driven Background Corrections

- ν_e -CC selection in the ND picks out FD backgrounds
 - Beam ν_e -CC
 - ν_μ -CC
 - Neutral current
- ~10% excess of data over MC in the ND
- Extrapolate data/MC differences to adjust FD prediction
- Each component oscillates differently
- Must decompose the data into constituent components



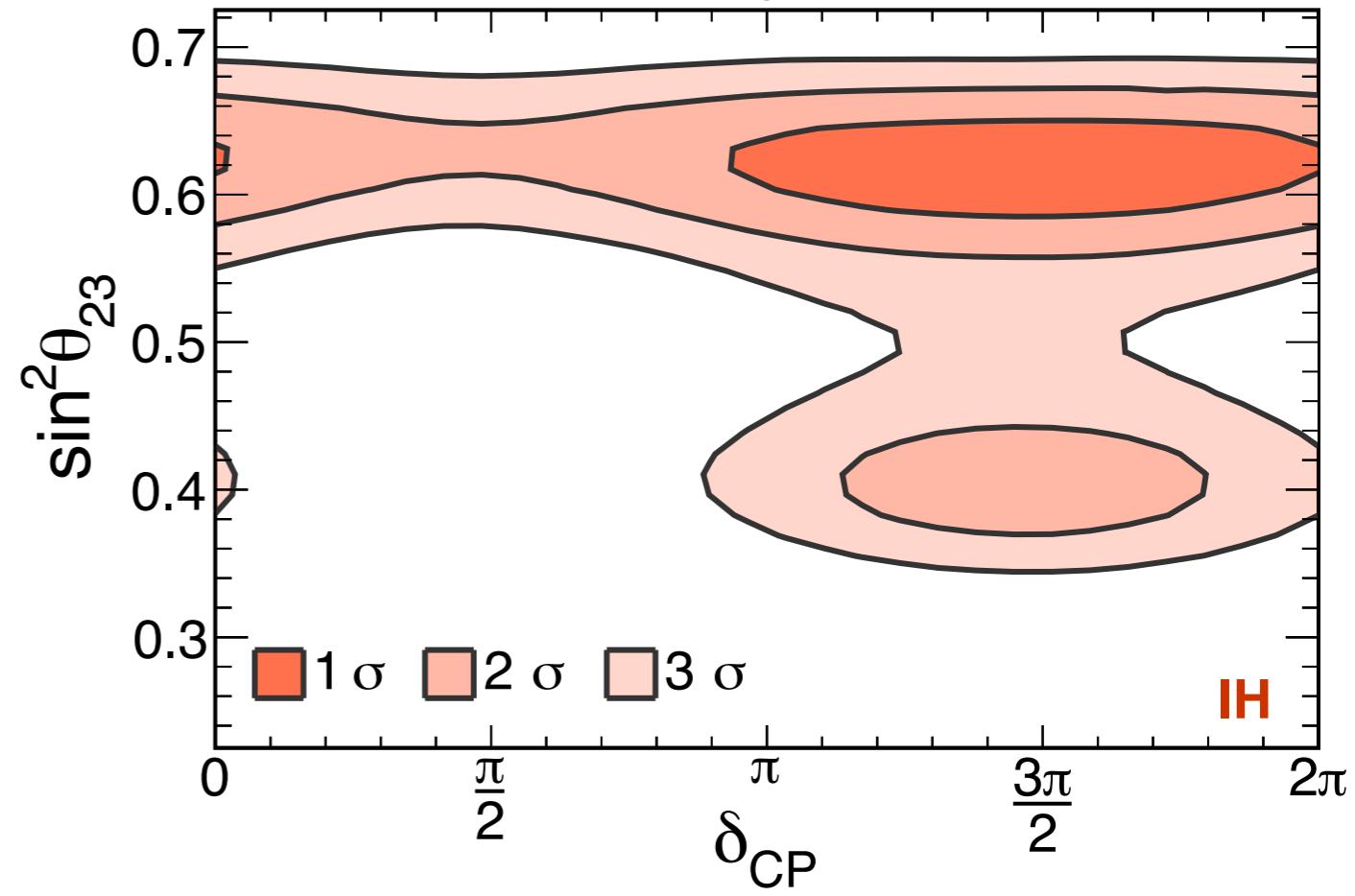
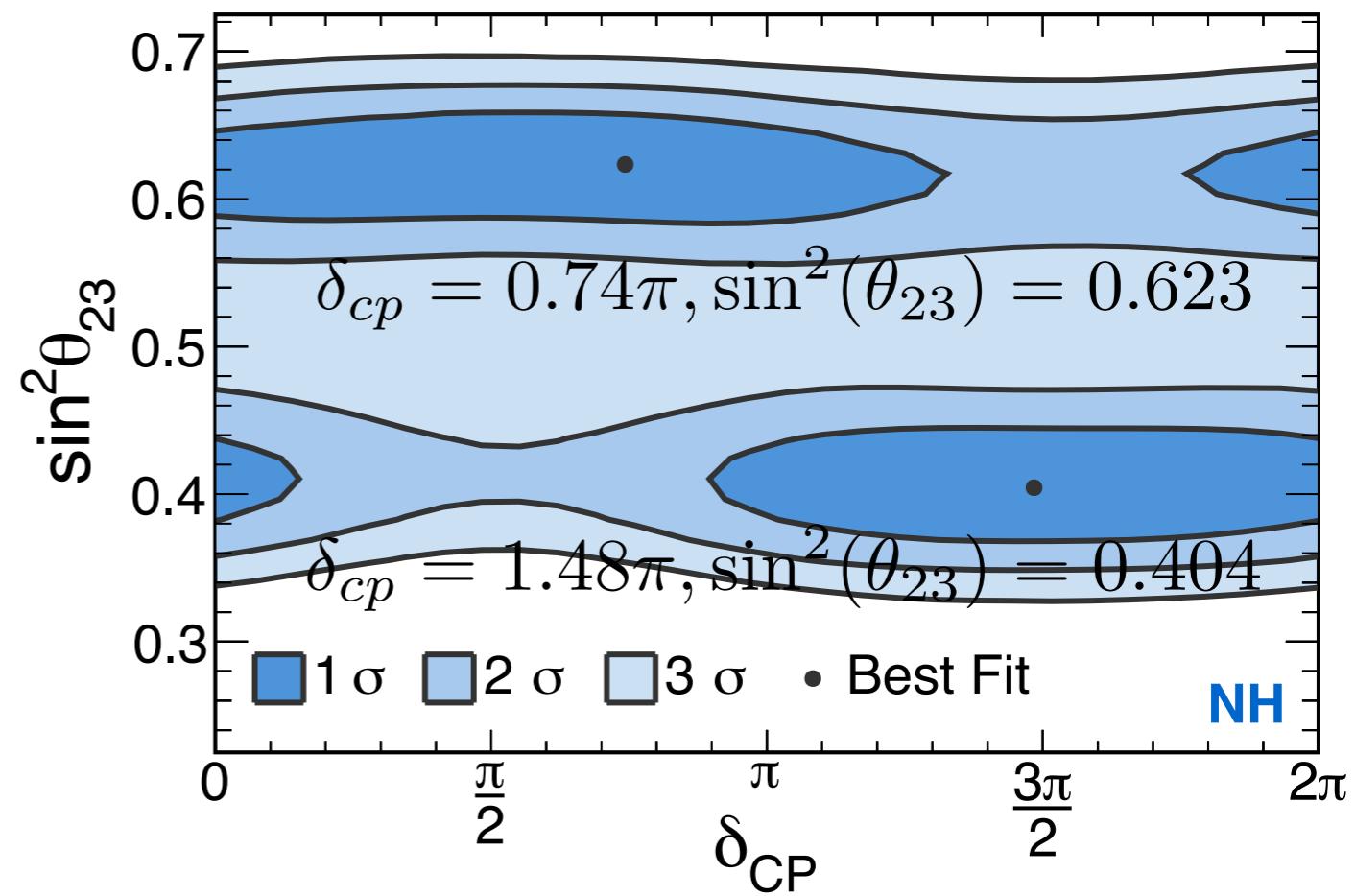
Electron-neutrino appearance

- Observe 33 events on background of 8.2 ± 0.8 events
- Confirmation that electron neutrinos are appearing (8σ)



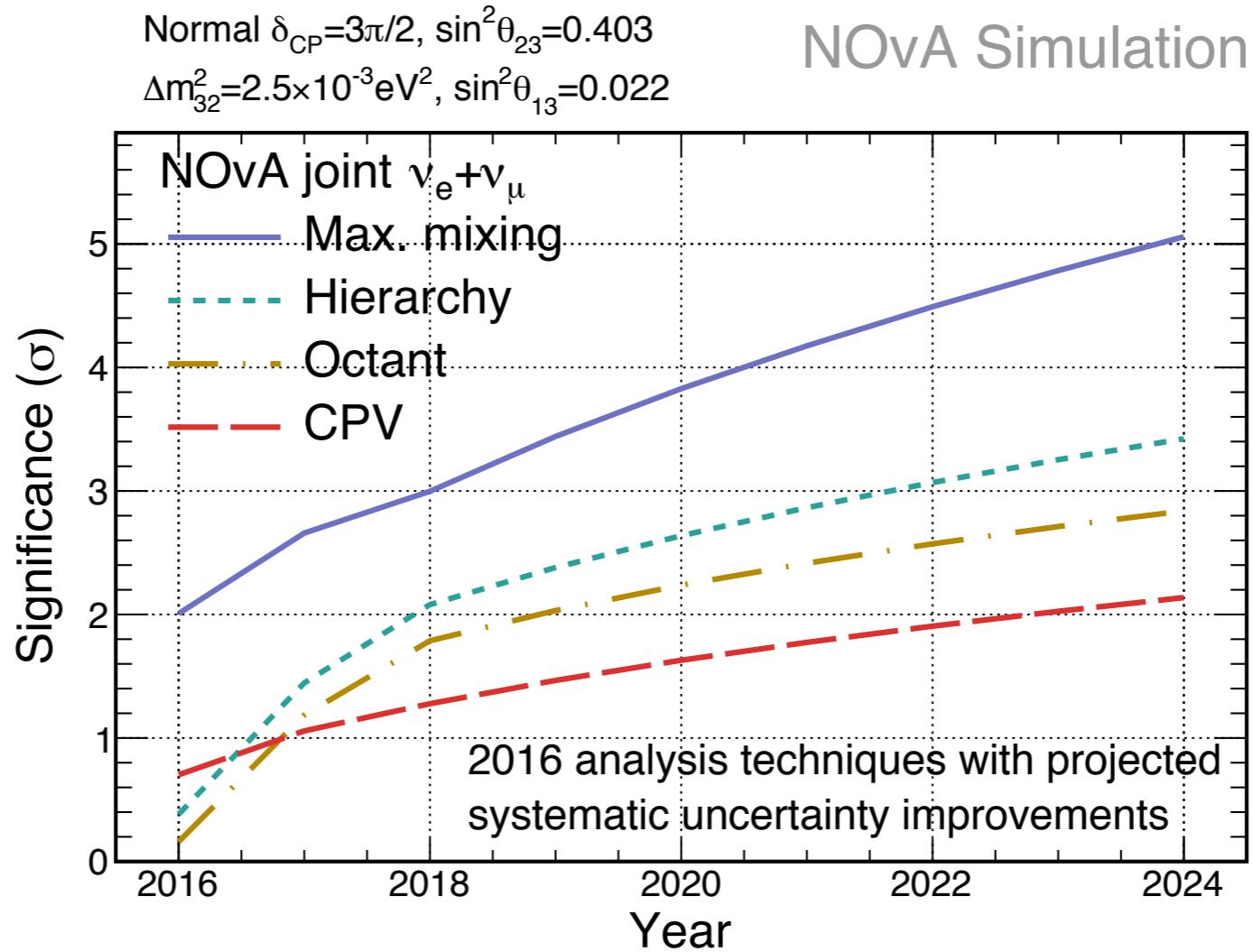
$\nu_\mu \rightarrow \nu_e$ Oscillation Results

- Fit for hierarchy, δ_{CP} , $\sin^2\theta_{23}$
 - Constrain $\sin^2 2\theta_{13} = 0.085 \pm 0.005$ from reactor experiments
 - Simultaneously fit NOvA disappearance and electron appearance data
- Global best fit, two degenerate points in Normal Hierarchy
 - best fit IH-NH, $\Delta\chi^2=0.47$
 - Lower octant, IH is disfavoured at greater than 93% C.L for all values of δ_{CP}



Phys. Rev. Lett. 118, 231801 (2017)

Looking Forward



- NOVA Switched to anti-neutrino running in February 2017
- Plan to run 50% neutrino, 50% anti-neutrino after 2018
 - **3 σ sensitivity to maximal mixing of θ_{23} in 2018**
 - **2 σ sensitivity to mass hierarchy and θ_{23} octant in 2018-2019**

Conclusions

- Analysis of 6.05×10^{20} POT of NOvA data (1 nominal year)
- Muon-neutrino disappearance ([arXiv:1701.05891](#))
 - **Best fit is non-maximal value of θ_{23} , maximal mixing disfavoured at 2.6σ**
- Electron-neutrino appearance ([arXiv:1703.03328](#))
 - **First joint fit of NOvA appearance and disappearance data**
 - **Weak preference for normal hierarchy**
 - **Inverted hierarchy, lower octant is disfavoured at $> 93\%$ C.L.**
- Anti-neutrino run 2017 completed last week
- Watch out for new results at next years Neutrino Conference

