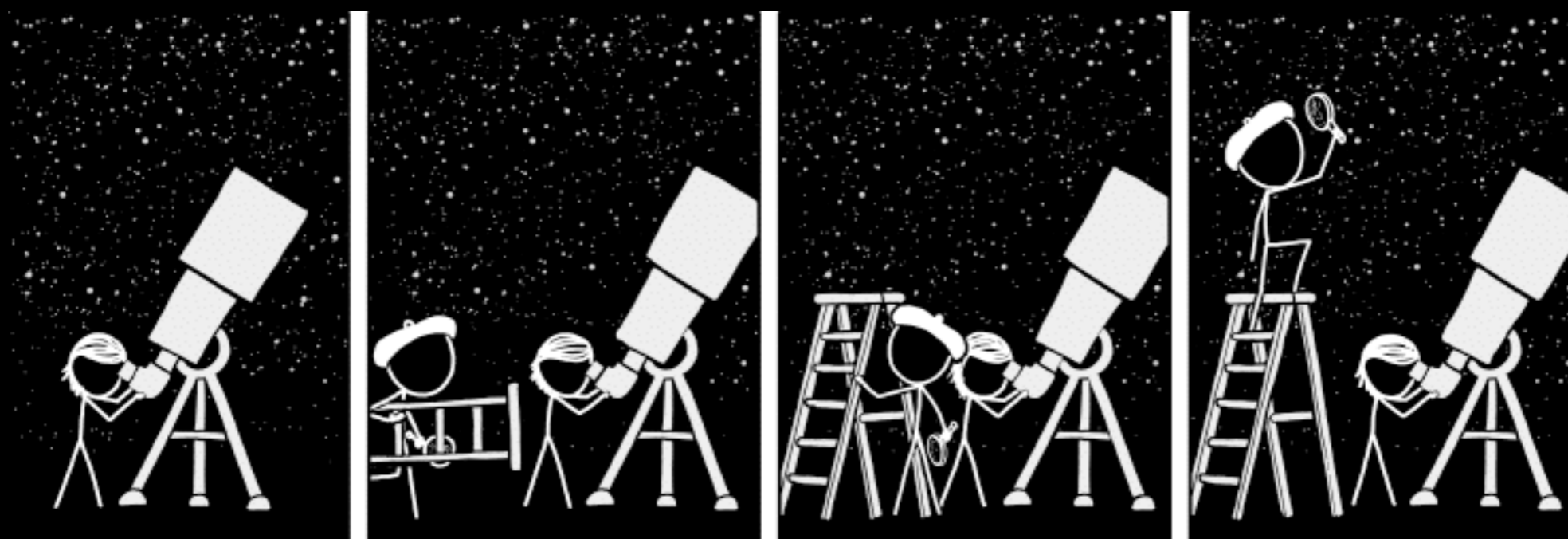


Astrometry of Water Fountains

(and some circumstellar OH masers)

Gabor Orosz
Kagoshima, Japan

2016.08.10.
Quy Nhon, Vietnam



Hiroshi Imai

Richard Dodson

Maria Rioja

Akiharu Nakagawa

Ross Burns

Dieter Engels

Sandra Etoaka

Yoshiharu Asaki

Hiroyuki Nakanishi

Sandor Frey

Jose Francisco Gomez

Ricardo Rizzo

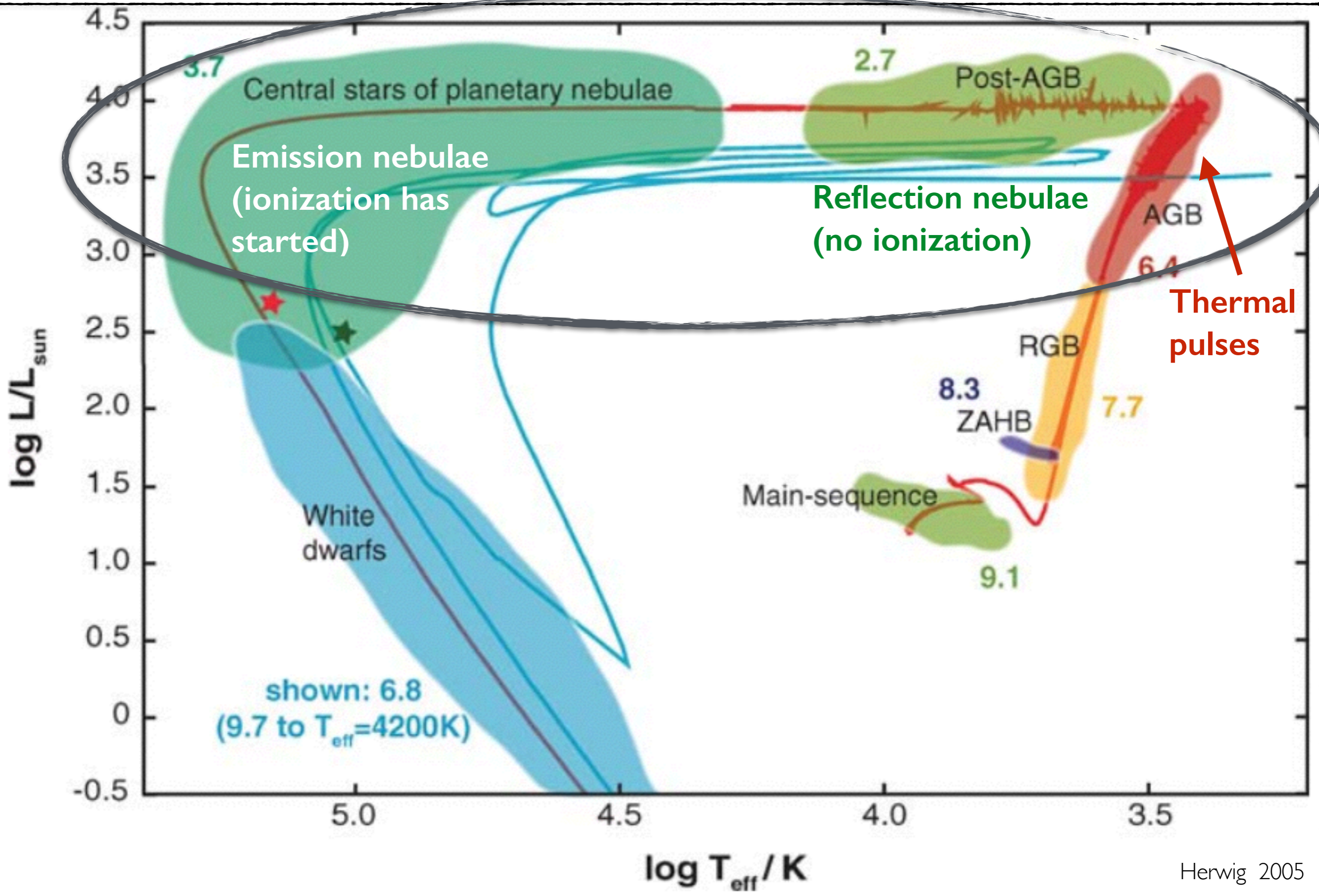
Olga Suarez

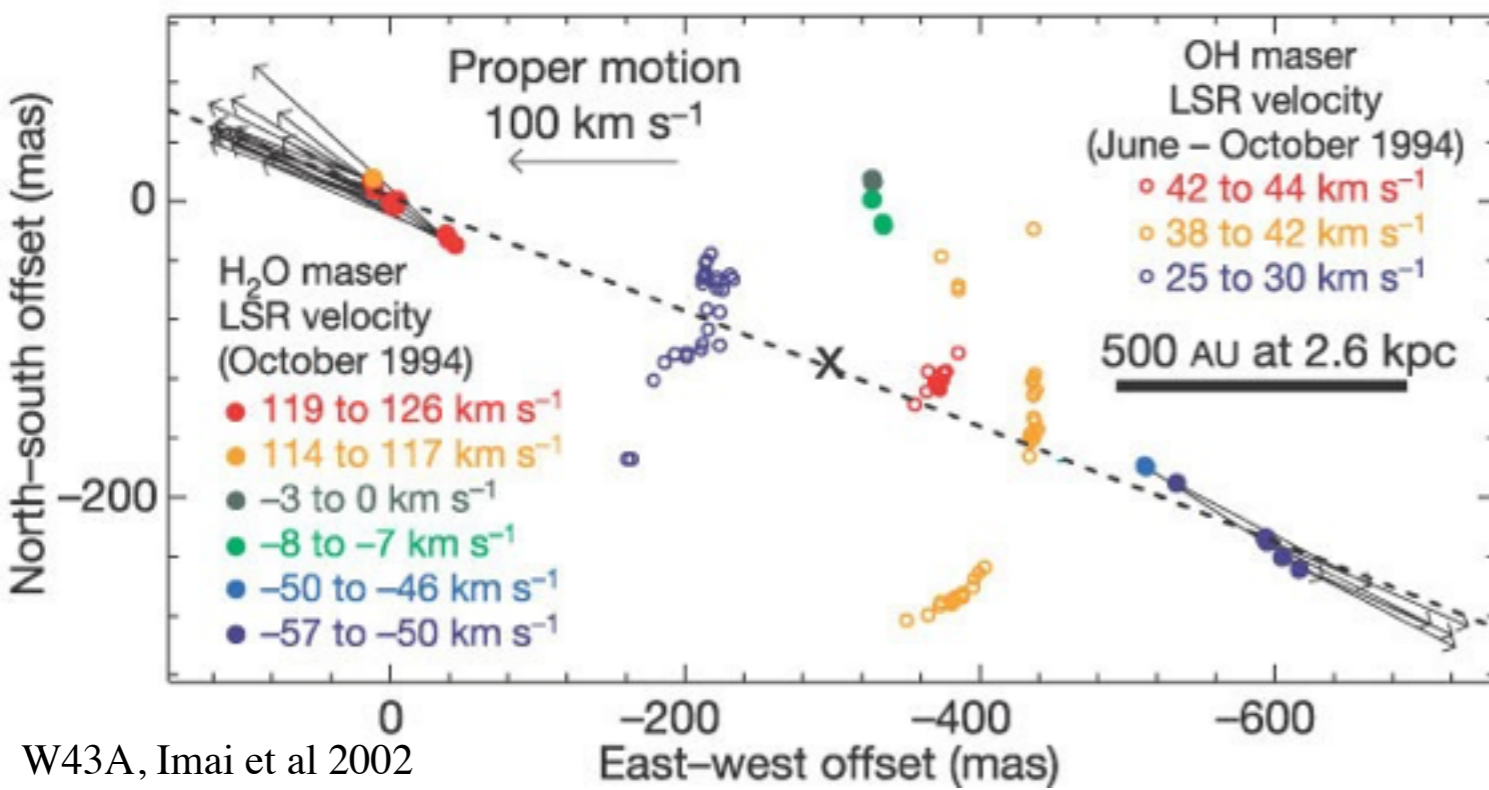
Steve Goldman

Daniel Tafoya

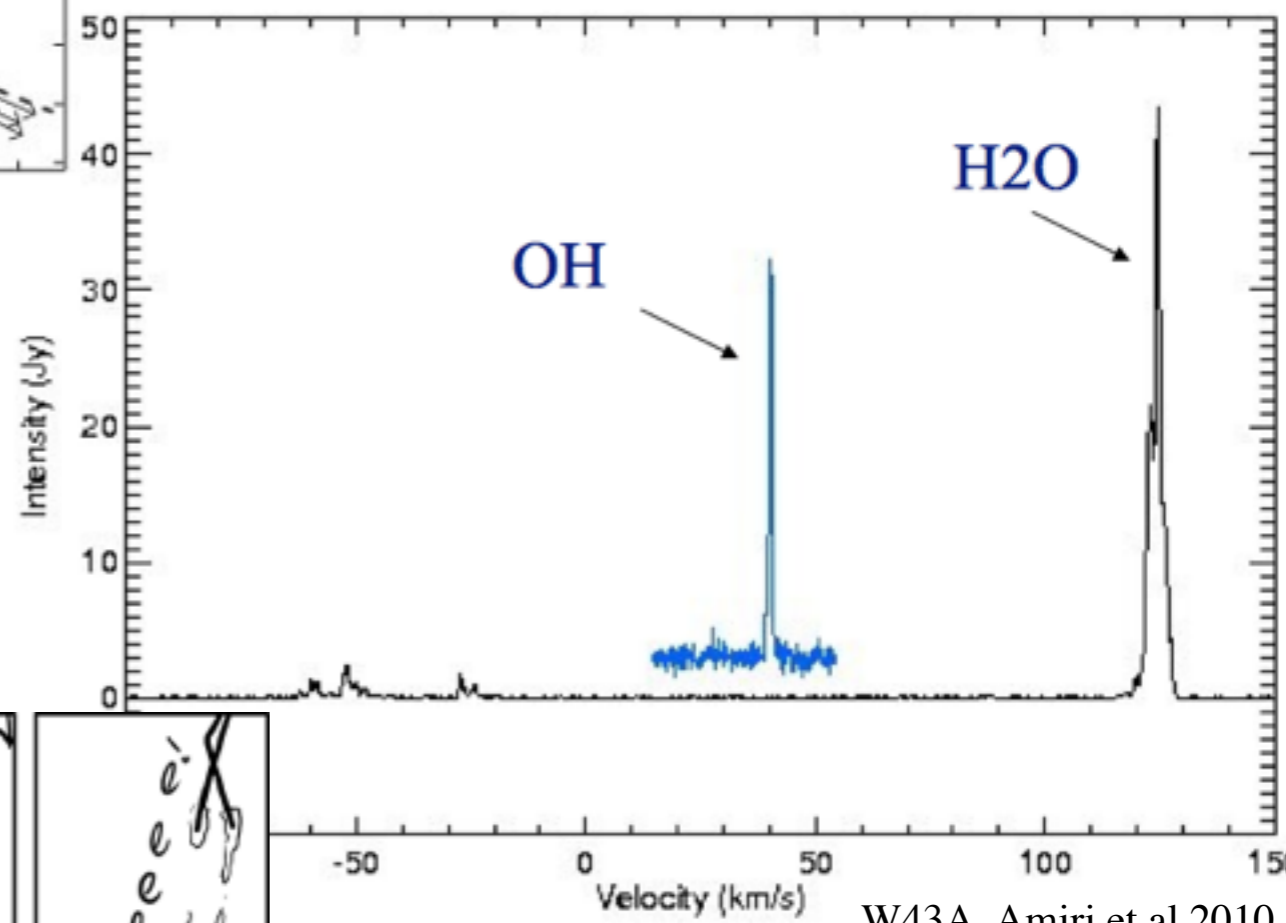
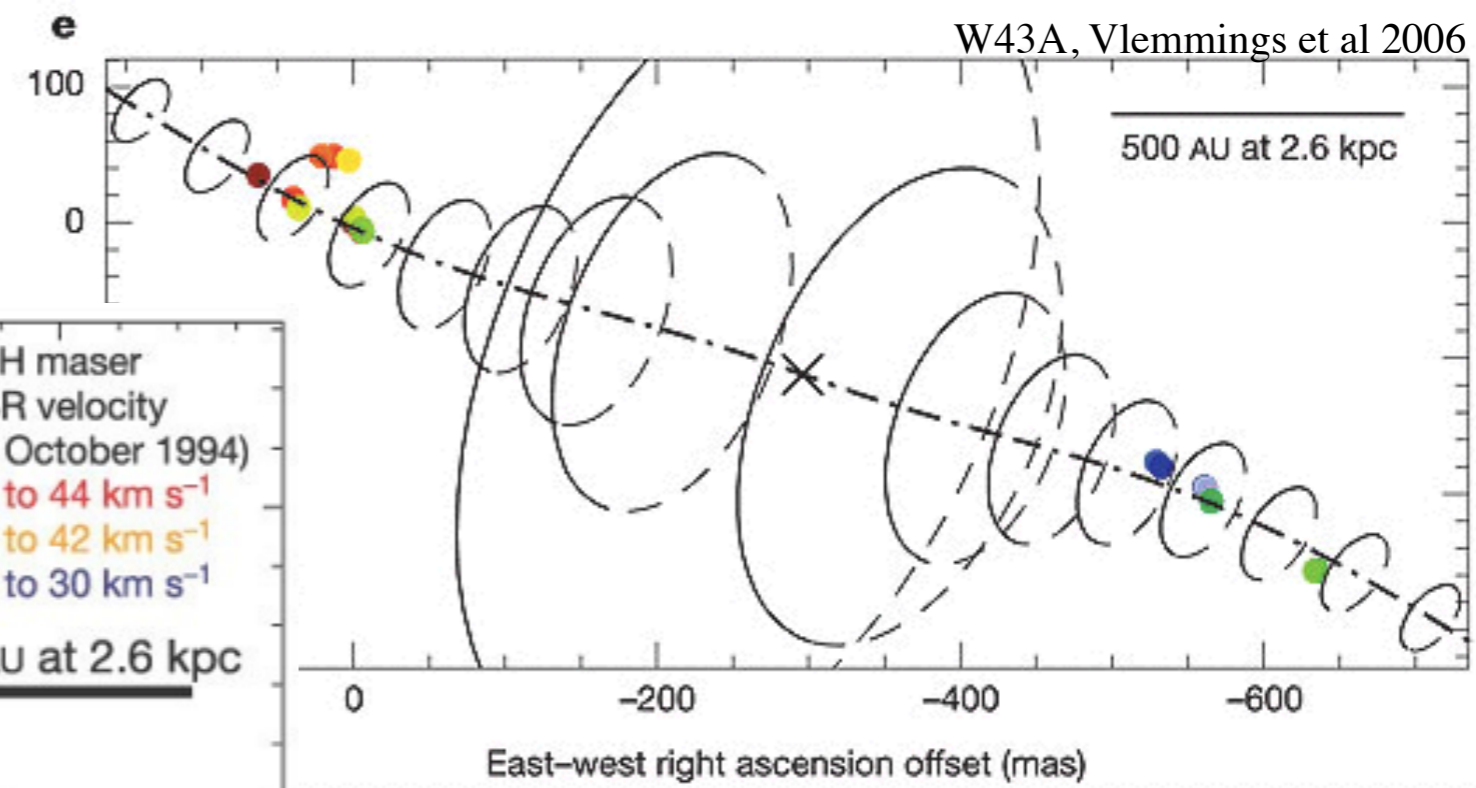
Andres Perez Sanchez

HR diagram of stellar evolution



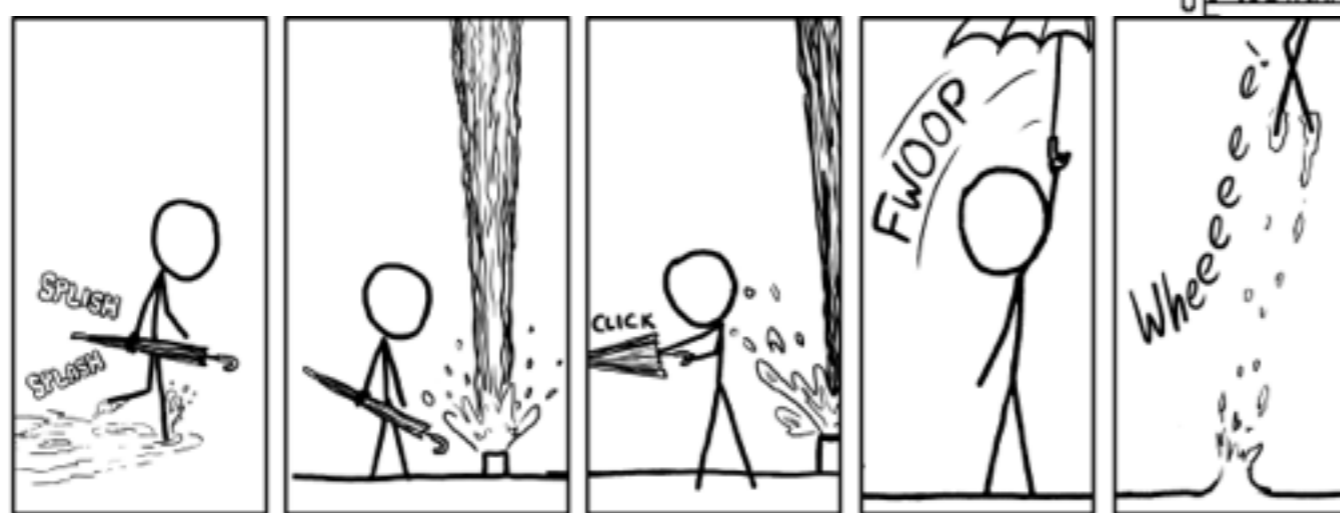
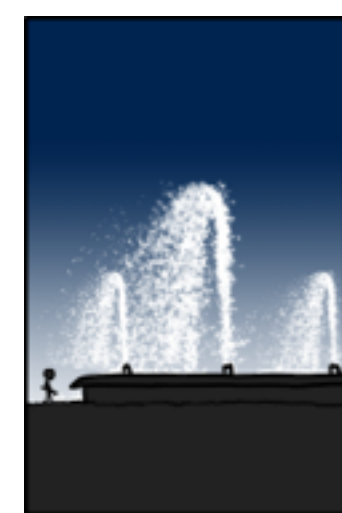


W43A, Imai et al 2002



W43A, Amiri et al 2010

Water fountain sources



XKCD: Fountain
<http://xkcd.com/1038/>

AGB CSE

~ 10 km/s

H₂O masers
22GHz, T \sim 400K
321GHz?

Cavity

Jet

fast molecular outflow

V \sim 100 km/s M \sim 0.1 M $_{\odot}$
very high momentum and energy
episodic event?

T \sim 10-20 K
H H₂ CO etc.

Probably a binary system

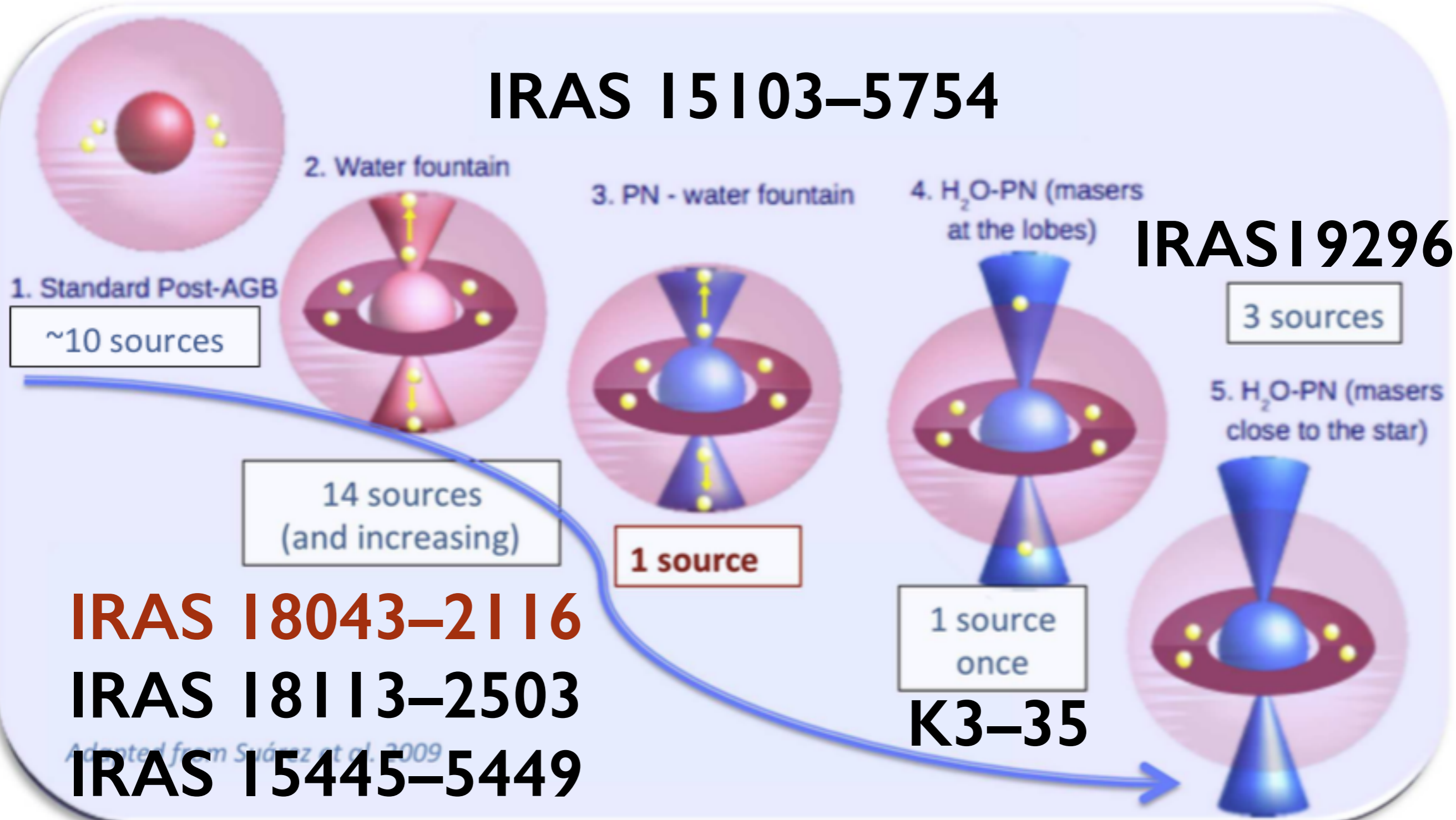
stellar or planetary (Jupiter-class)
common envelope binaries?

Magnetic field

Accretion disk

Sketch of the possible water maser evolution

IRAS 15103-5754



IRAS 18043-2116

IRAS 18113-2503

IRAS 15445-5449

Adapted from Suárez et al. 2009

All known and confirmed water fountains

as of March, 2016

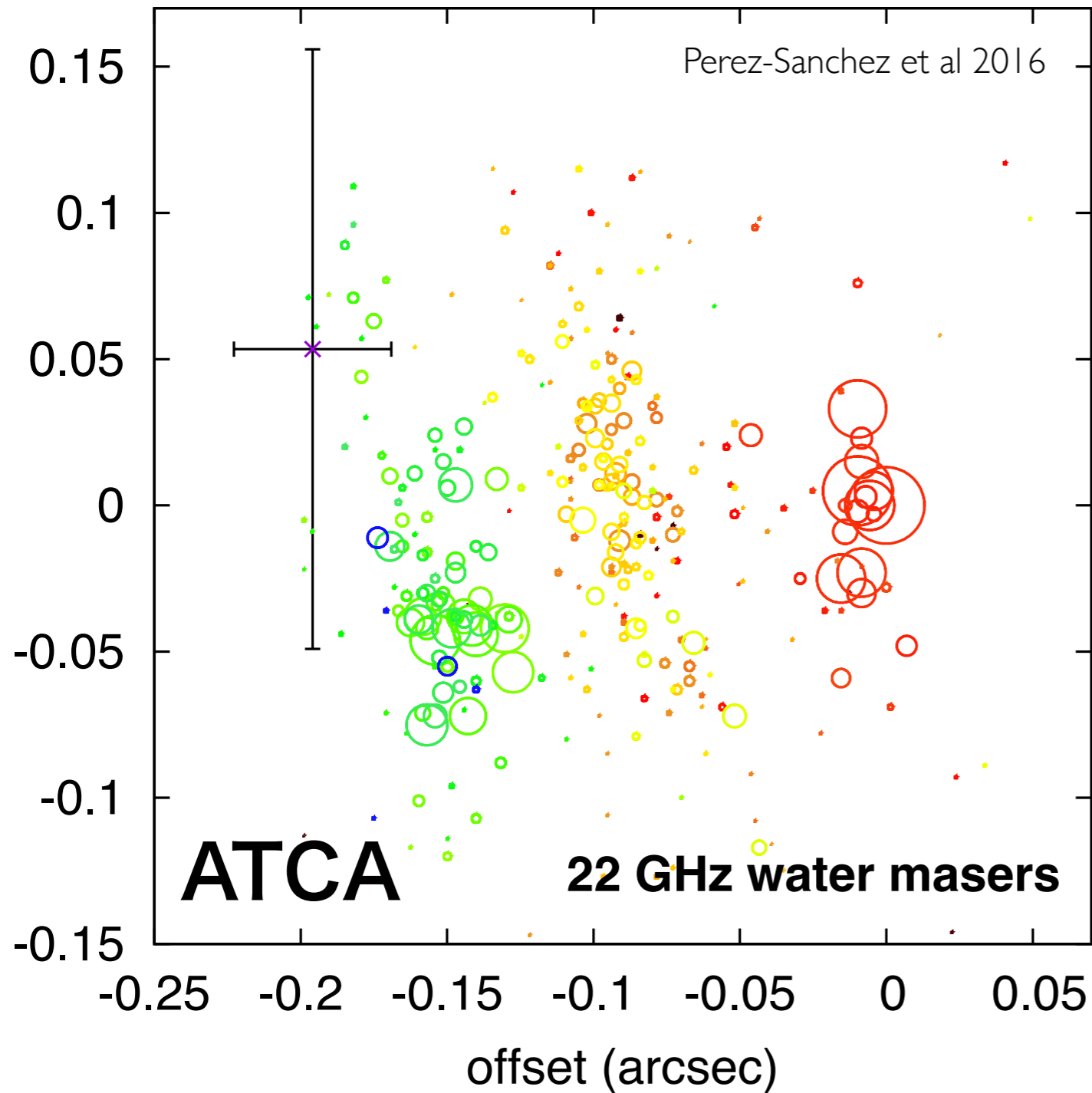
IRAS name	Alt. name	V_{range}	Morphology	References
15103–5754	GLMP 405	70	BJA	Suárez et al. 2009
15445–5449	OH 326.5–0.4	90	TJA	Pérez-Sánchez et al. 2011
16342–3814	OH 344.1+5.8	260	TJ	Claussen et al. 2009
16552–3050	GLMP 498	170	TJ	Suárez et al. 2008
18043–2116	OH 9.1–0.4	400	BE	Walsh et al. 2009
18113–2503	PM 1–221	500	TJ	Gómez et al. 2011
18139–1816	OH 12.8–0.9	50	TJ	Boboltz & Marvel 2005
18286–0959	OH 21.80–0.13	200	BJ	Yung et al. 2011
18450–0148	W43A	180	TE	Imai et al. 2002
18455+0448		40	BJ	Vlemmings et al. 2014
18460–0151	OH 31.0–0.2	310	TEA	Imai et al. 2013
19134+2131		100	TJ	Imai et al. 2007b
19190+1102	PM 1–298	100	TE	Day et al. 2010

Tidbinbilla 70m single-dish monitoring project (T215)

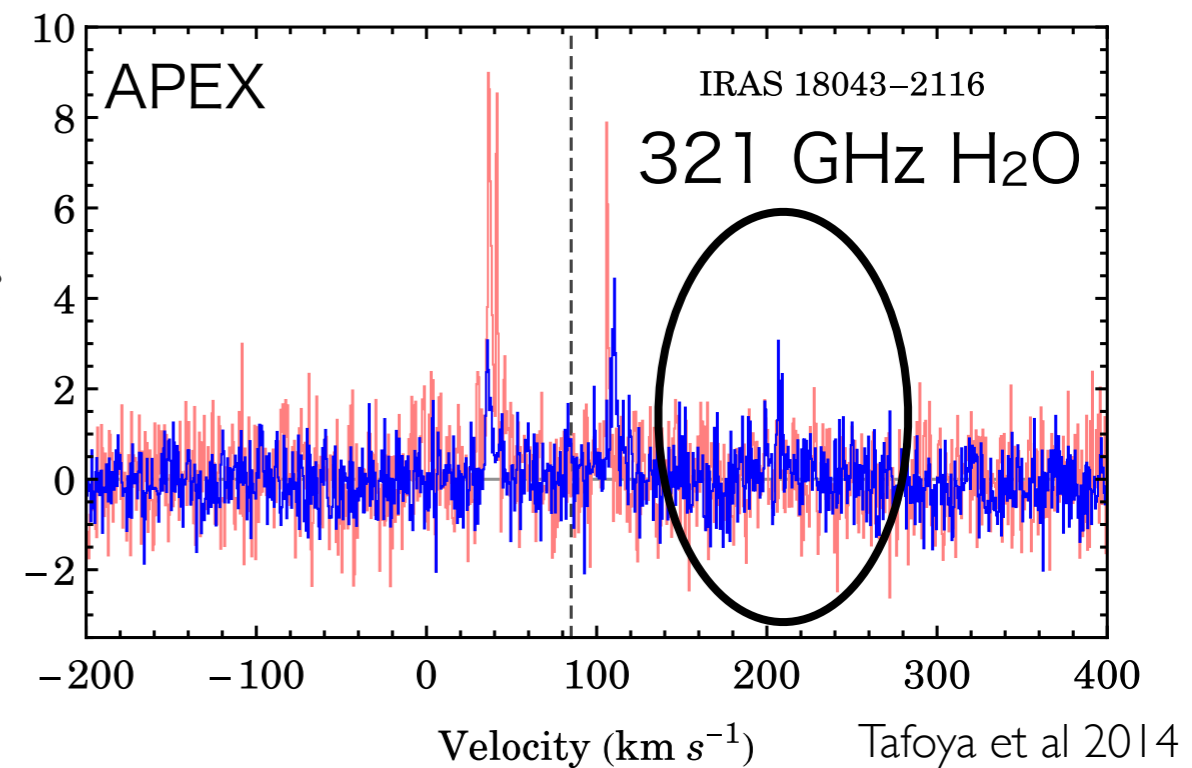
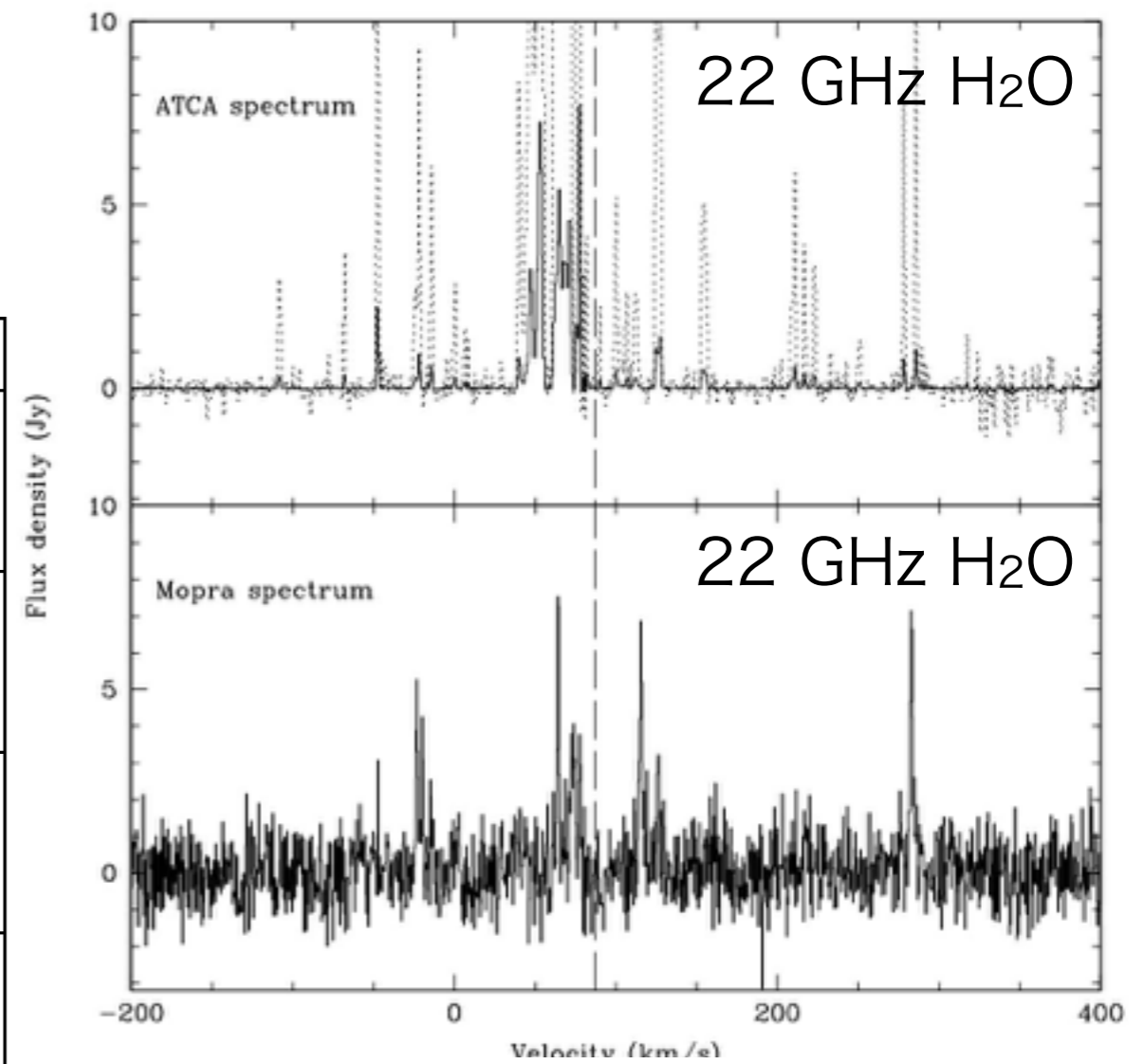
Today's detective story: the mystery of IRAS 18043...

IRAS 18043–2116

a **HOT** water fountain?

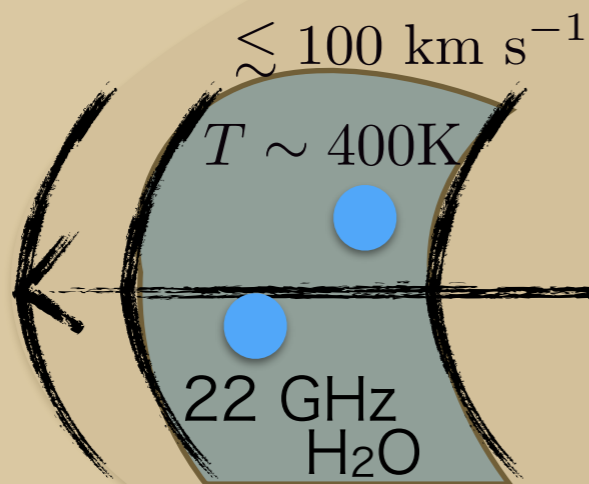


unresolved continuum emission at 22 GHz
(shock-induced synchrotron dampened by free-free)



AGB CSE

Cavity



Jet 1 $\sim 100 \text{ km s}^{-1}$

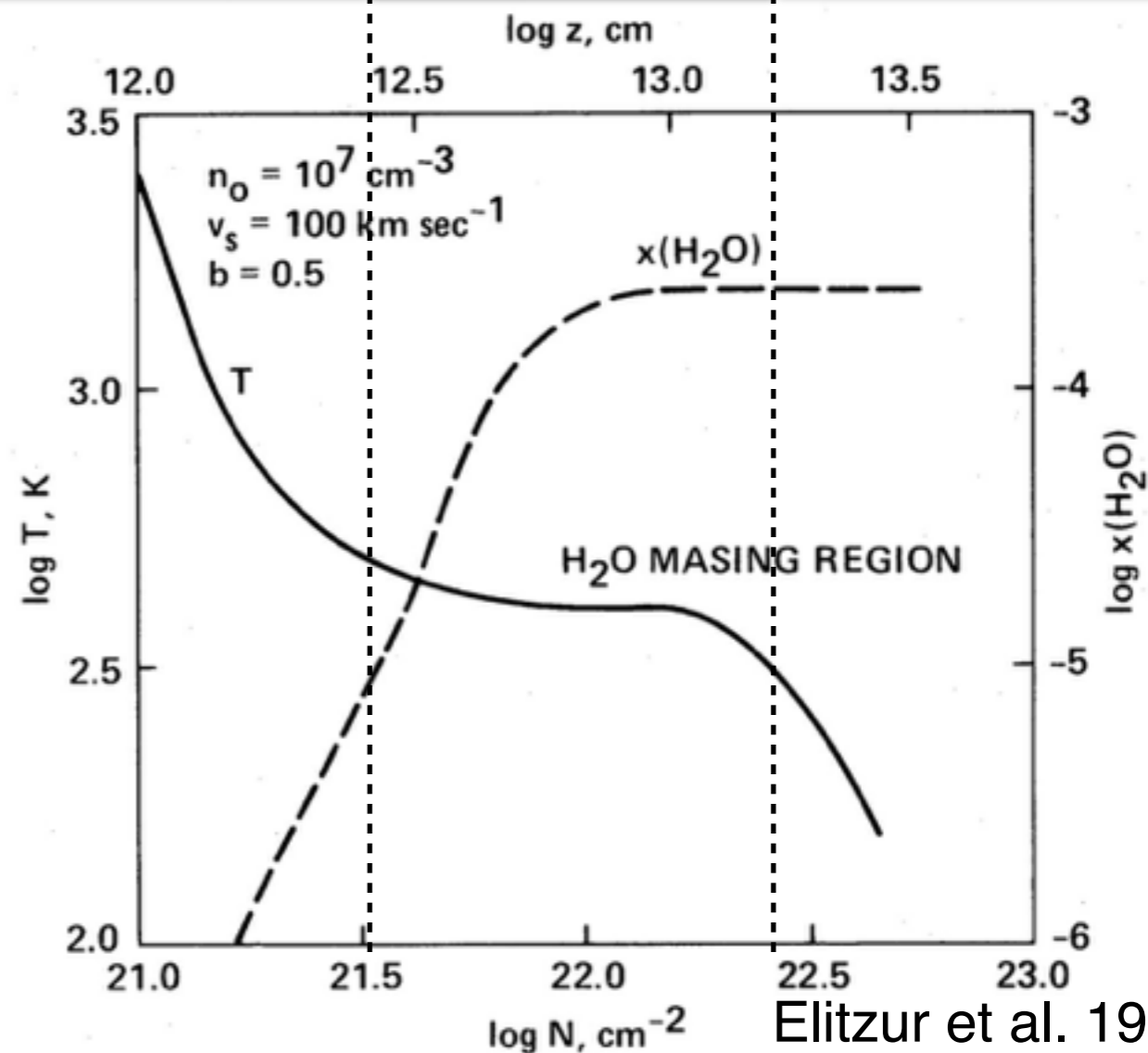
high temperature
but too little water

high-enough
temperature
enough water

enough water
but too cold

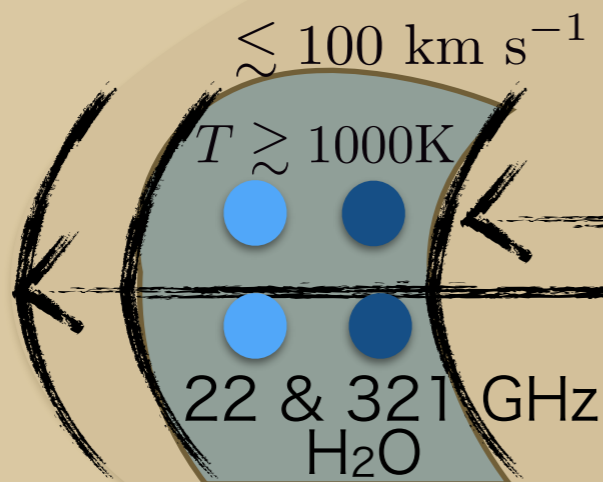
$\sim 10 \text{ km/s}$

from Daniel Tafoya



AGB CSE

Cavity

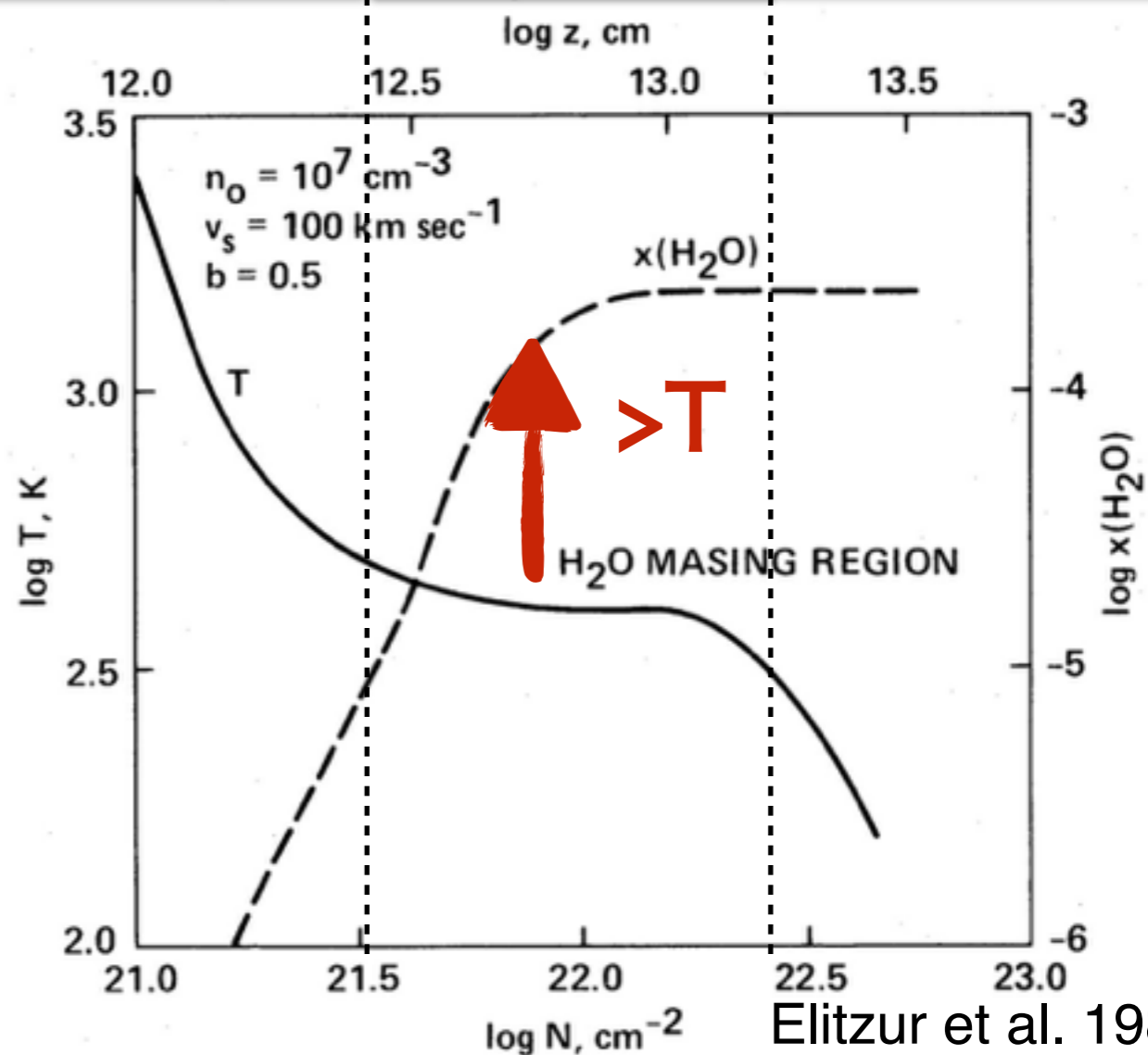


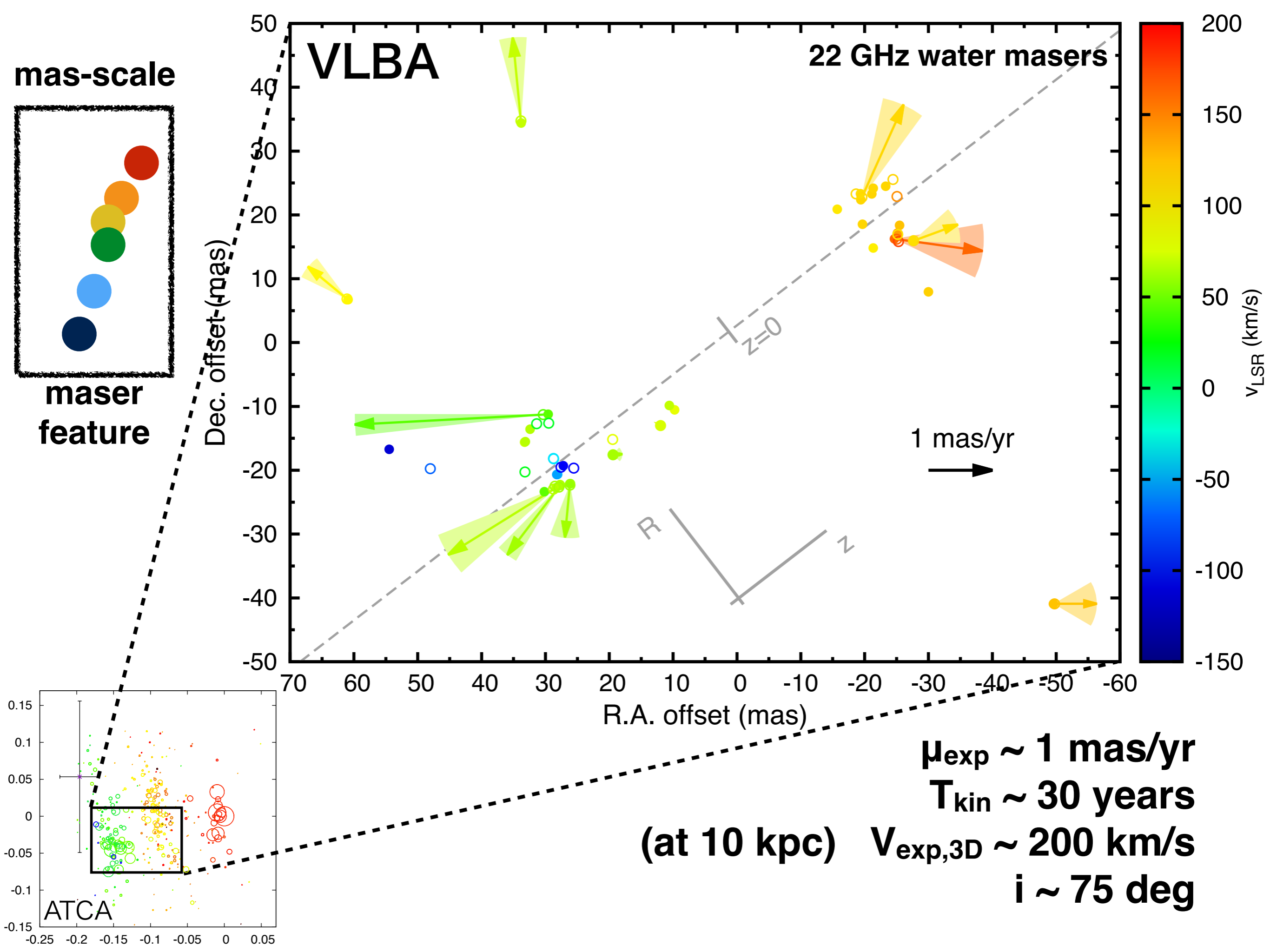
Jet 2 $\sim 100 \text{ km s}^{-1}$

Jet 1 $\sim 100 \text{ km s}^{-1}$

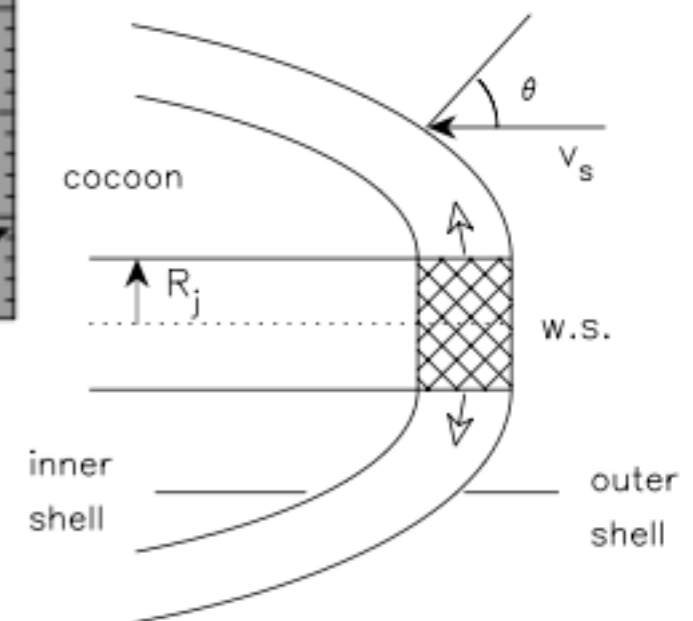
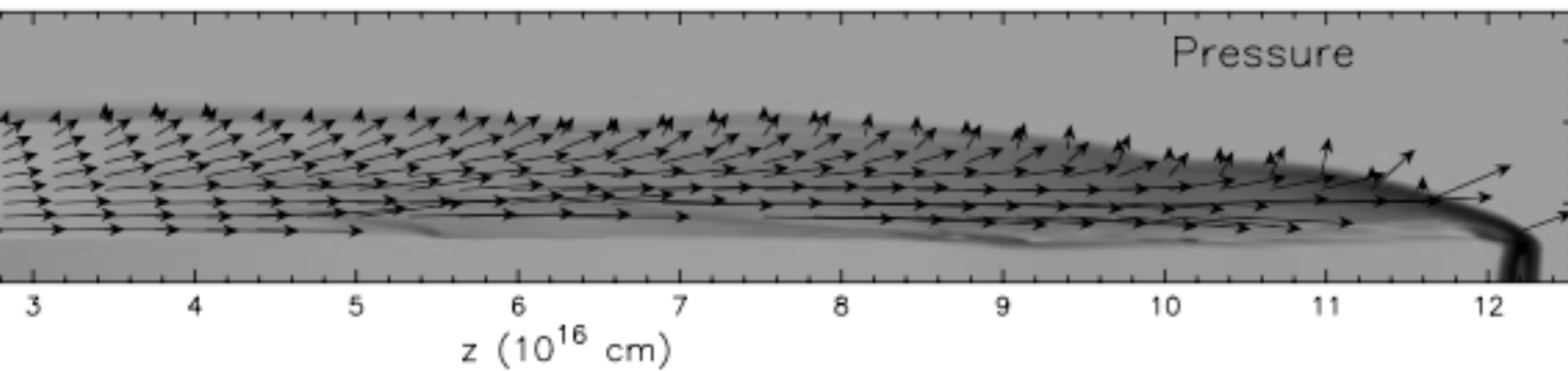
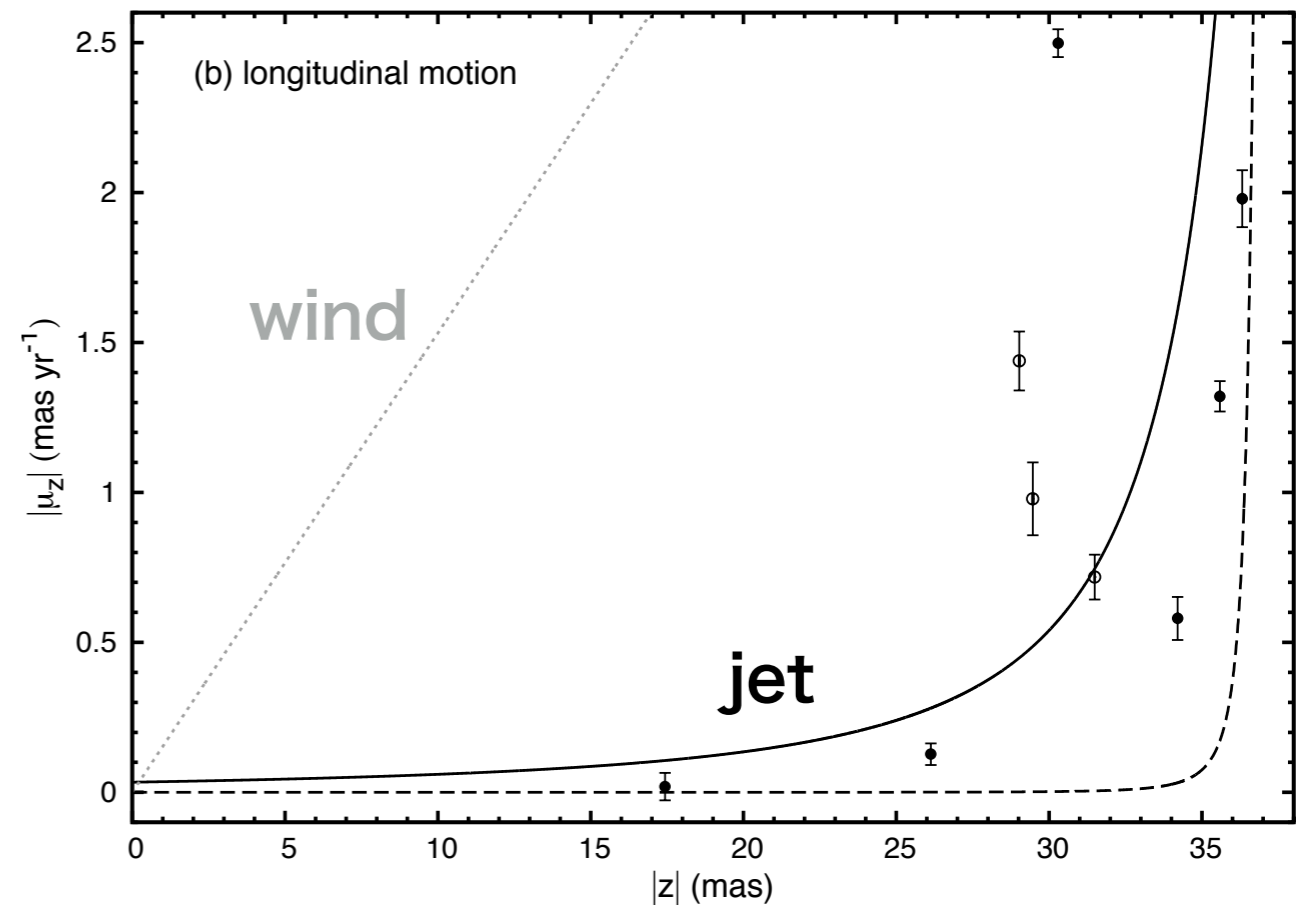
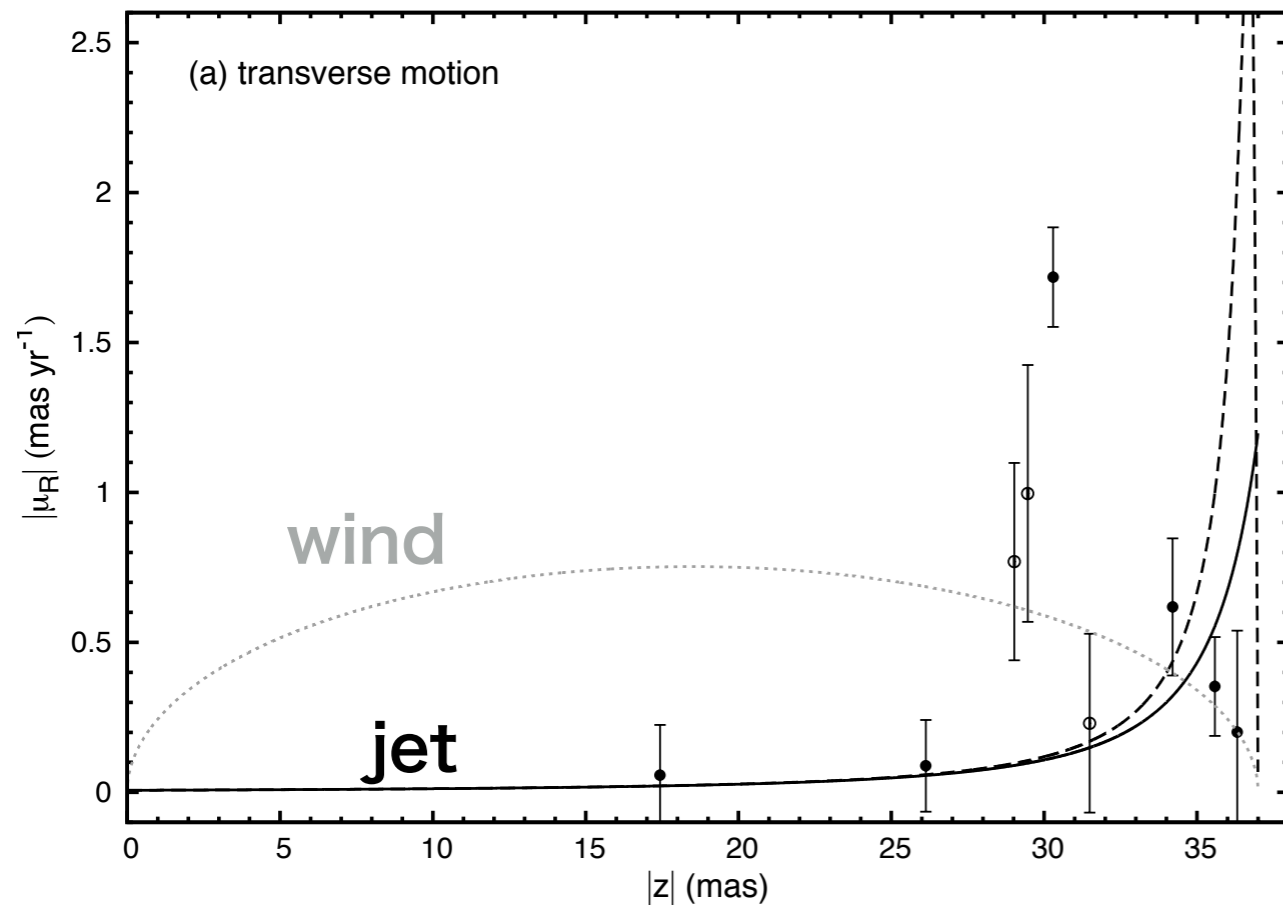
$\sim 10 \text{ km/s}$

from Daniel Tafoya



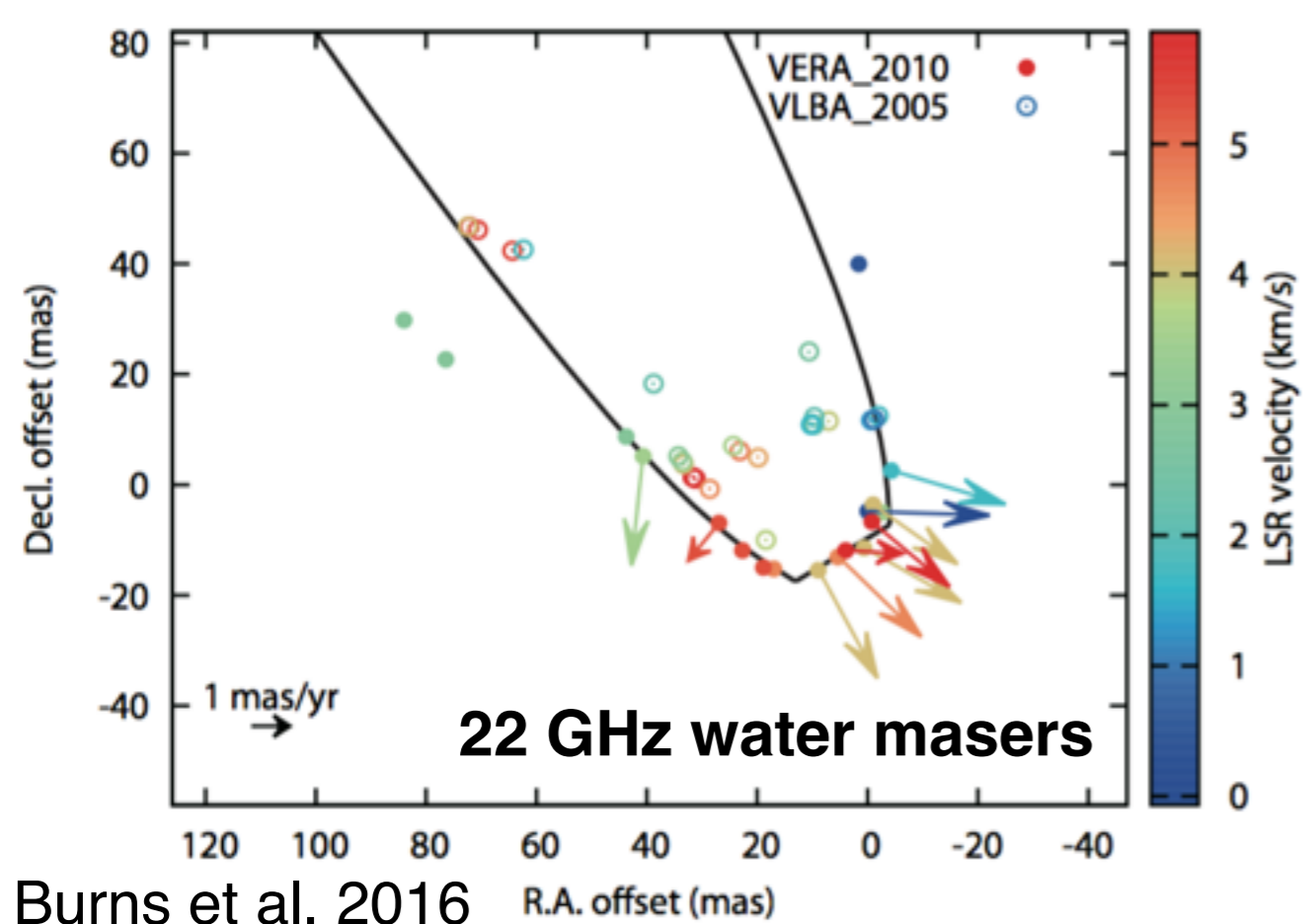


Ballistic bow shock model

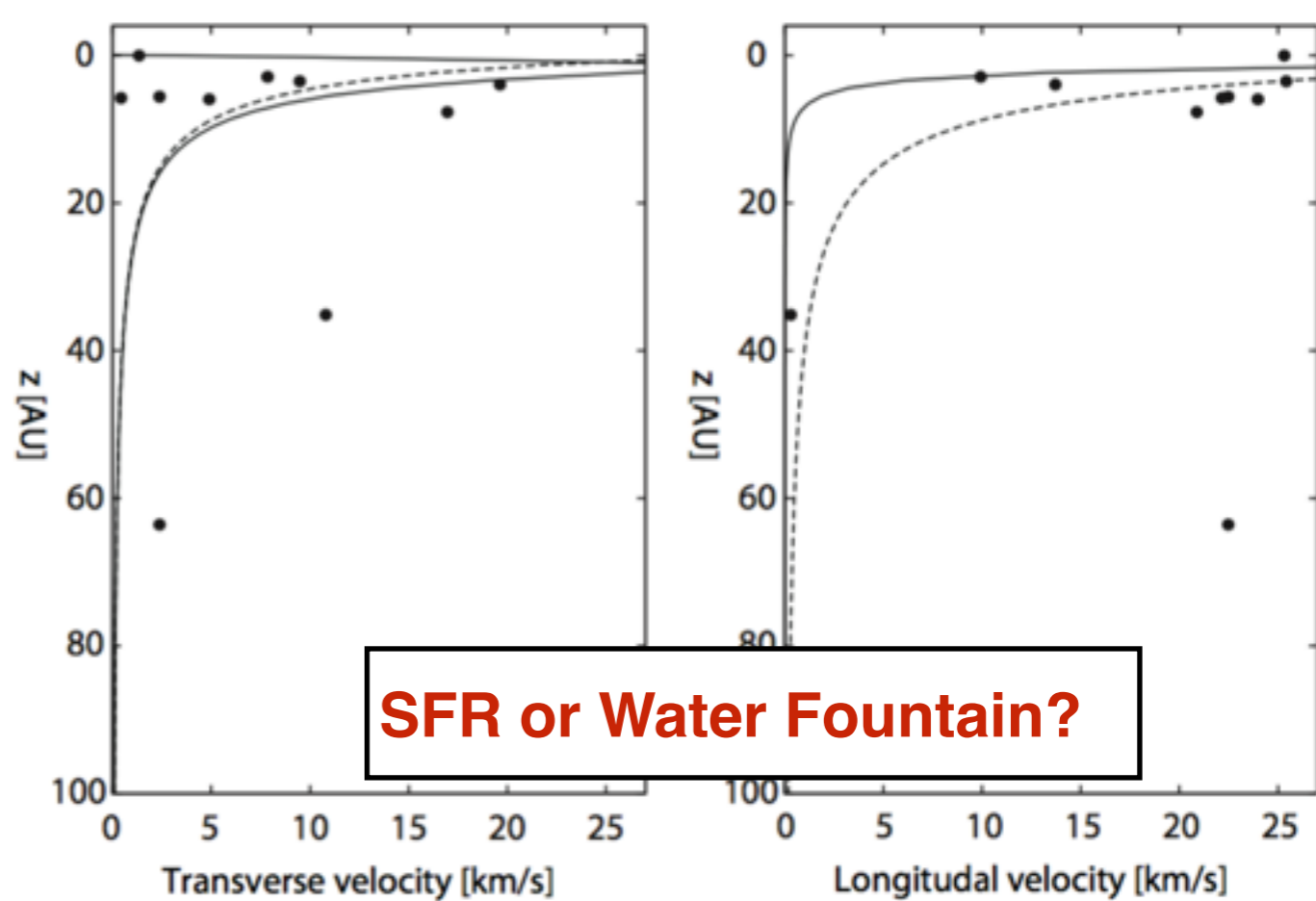


Jet-driven protostellar outflow shells:

Ostriker et al 2001 (model), Lee et al 2001 (simulation)

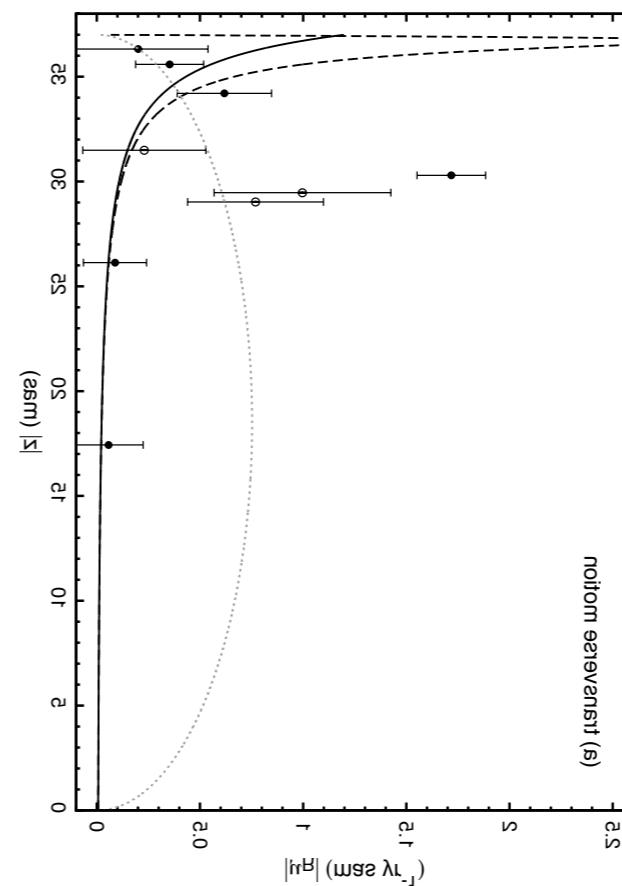
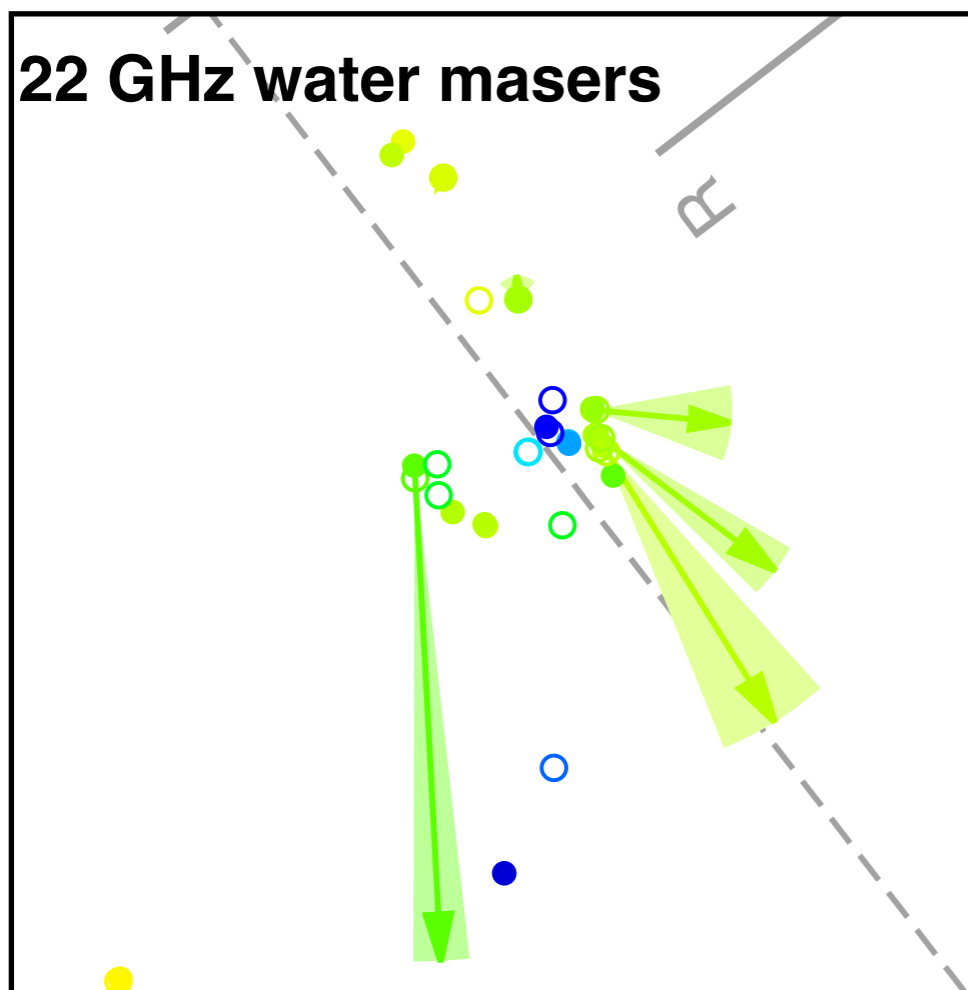


Burns et al. 2016

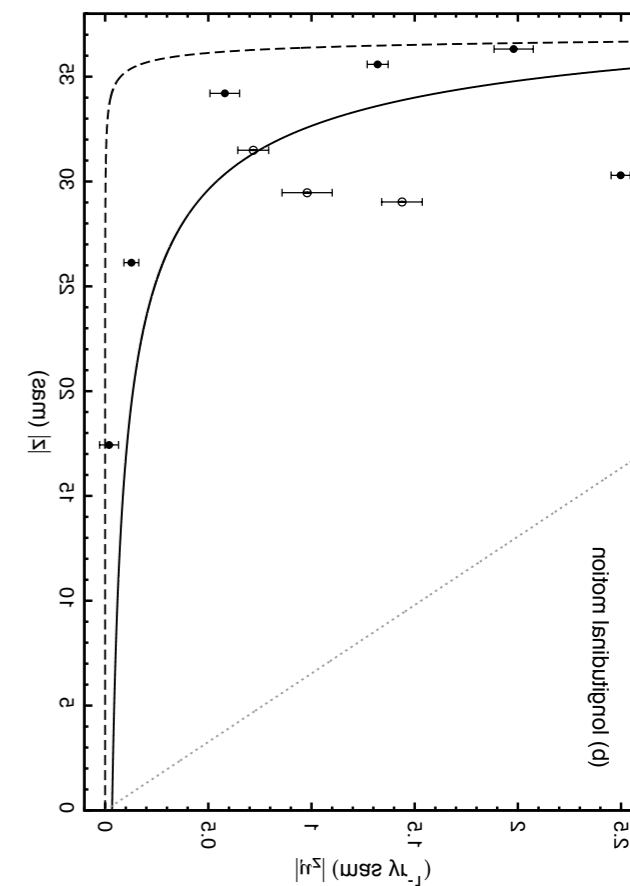


$T_{\text{kin}} \sim 130$ years

$V_{\text{exp,3D}} \sim 20$ km/s



$T_{\text{kin}} \sim 30$ years



$V_{\text{exp,3D}} \sim 200$ km/s

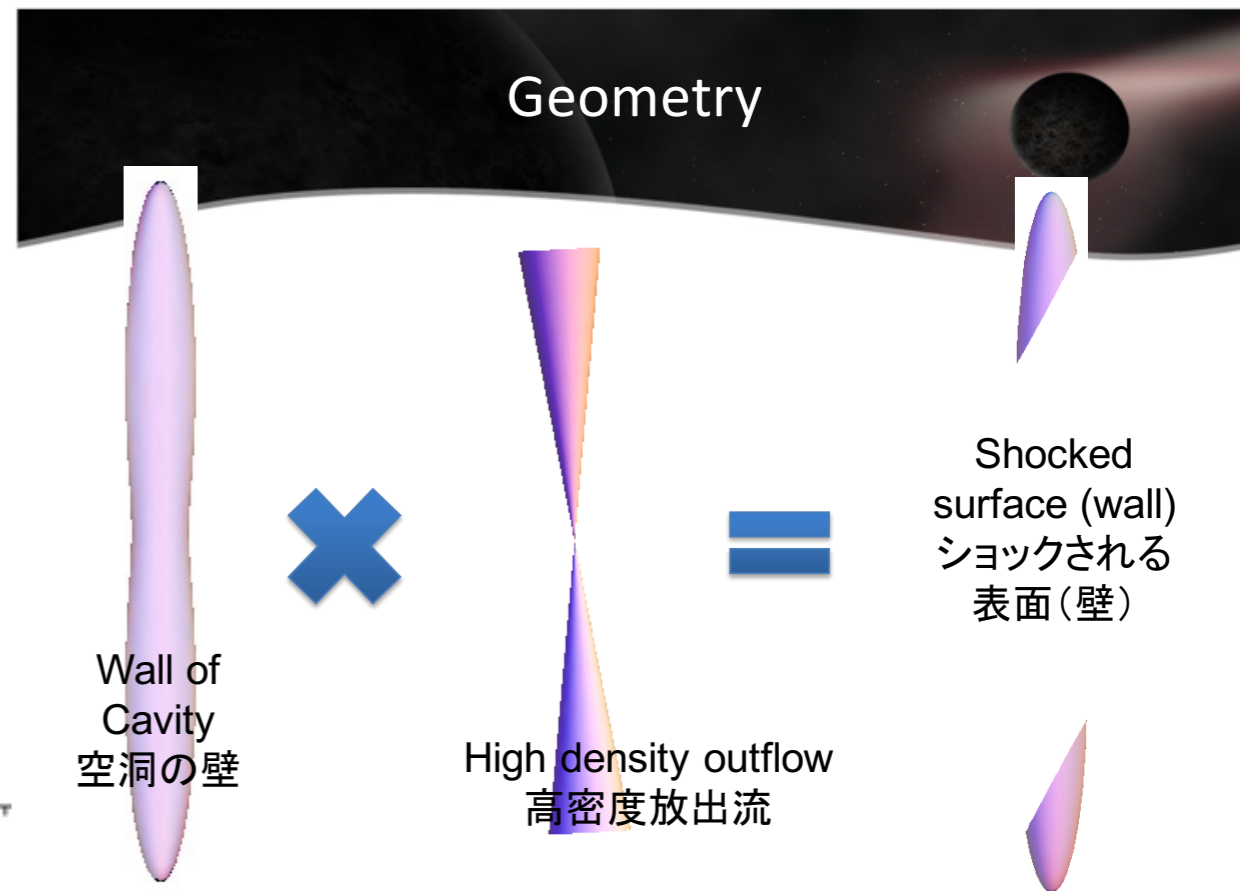
Water fountains outflows

Two possible scenarios

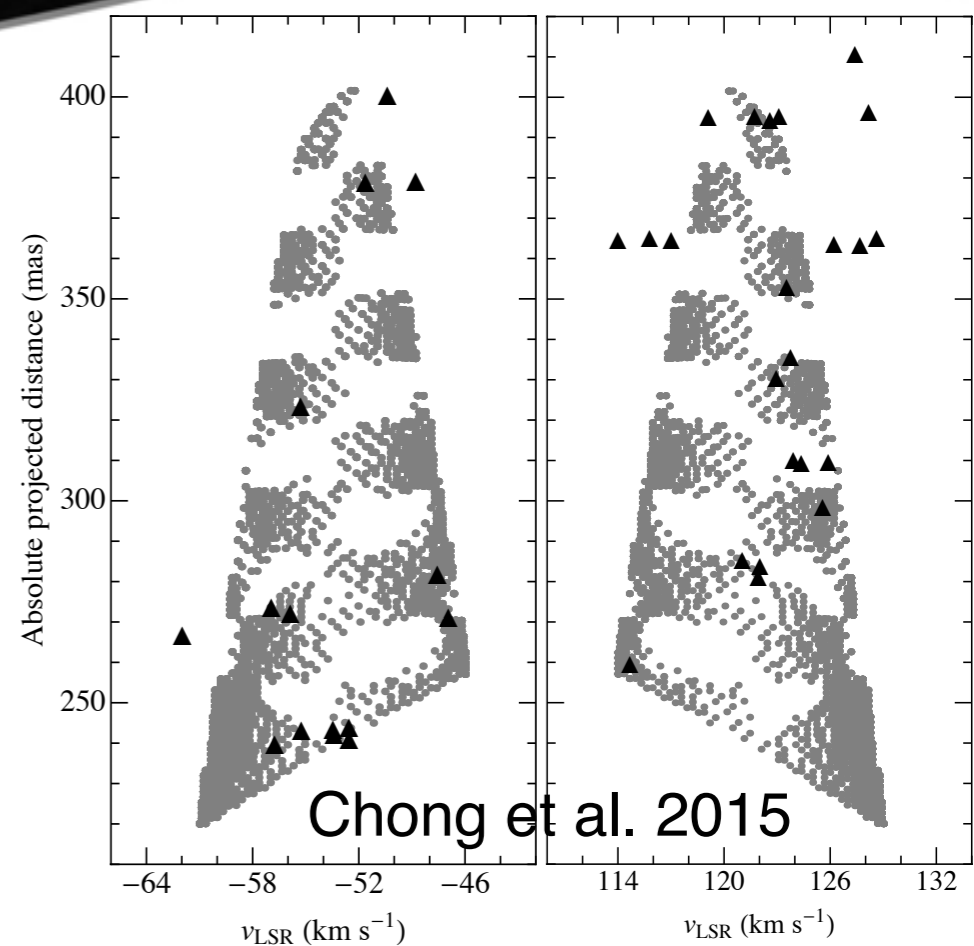
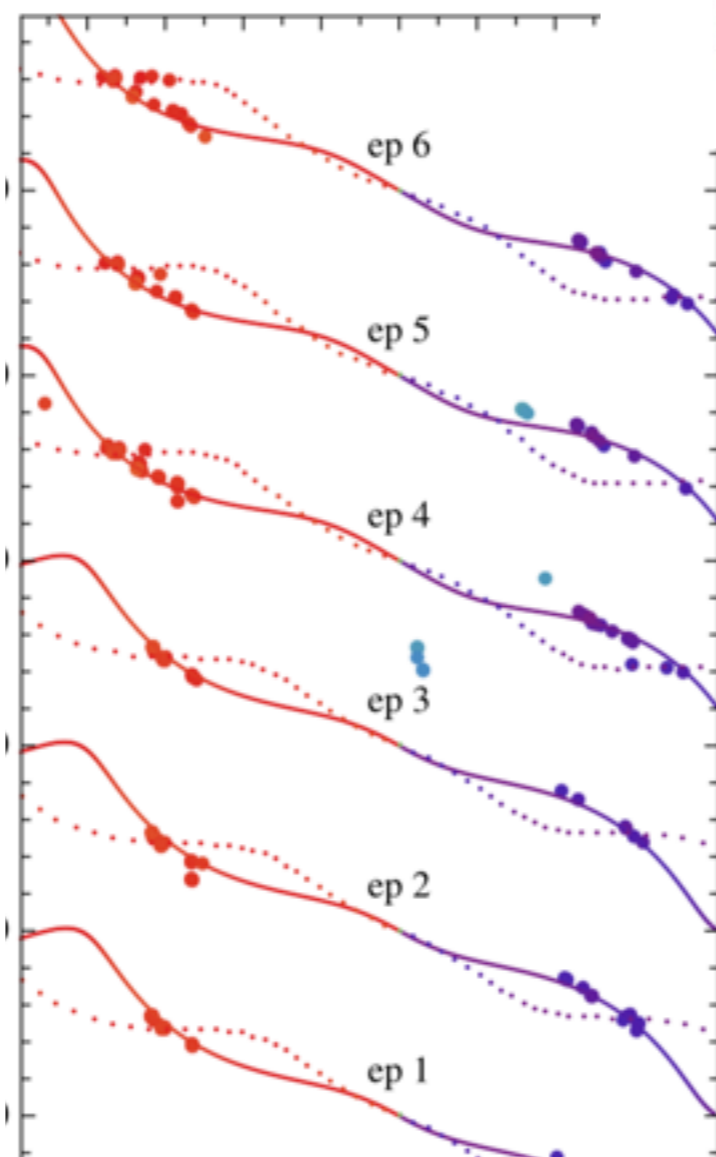
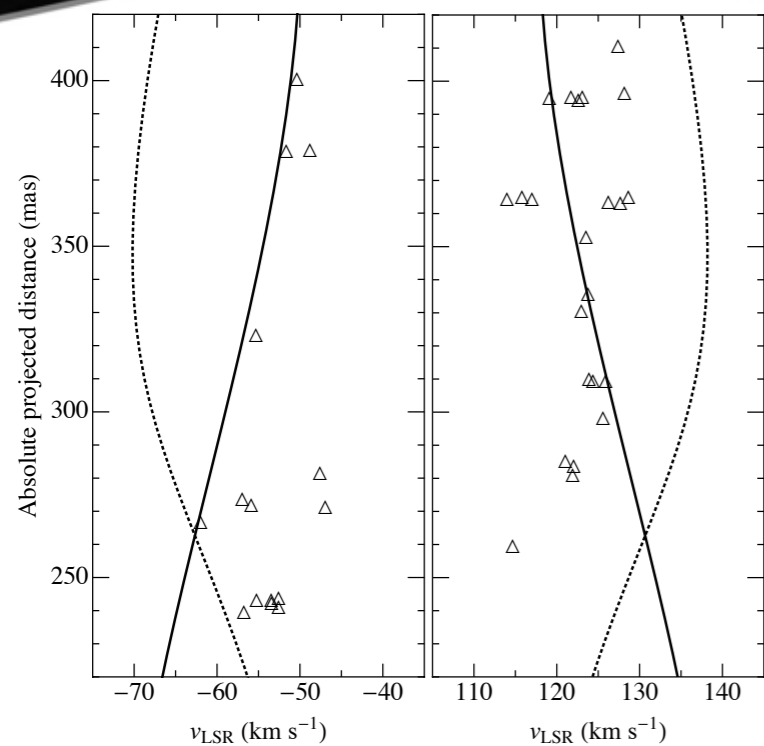
Cavity model

vs

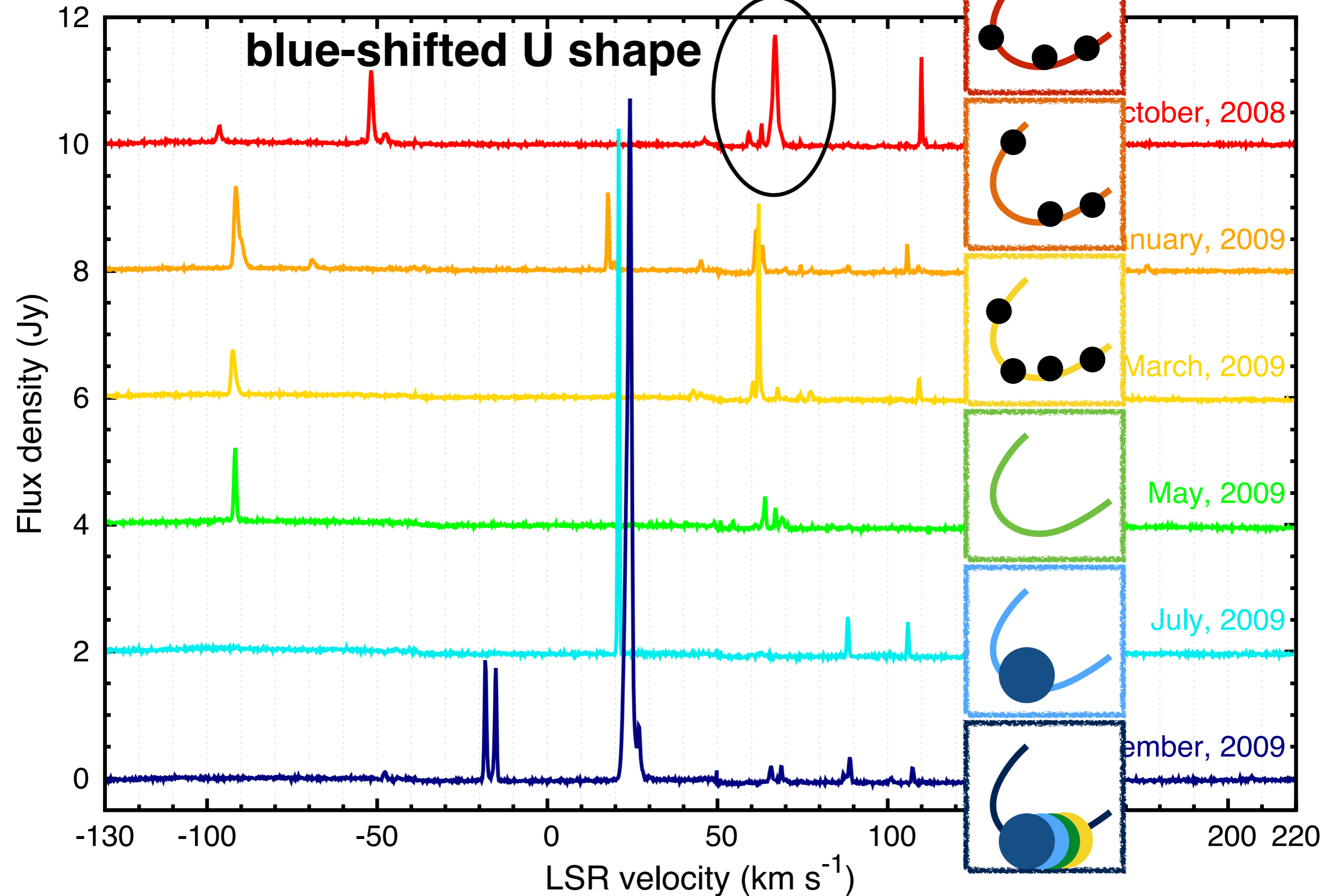
precessing jet model



Precessing Jet Model 歳差ジェットモデル



Episodic event?



So is IRAS 18113 a hot water fountain?

collimated outflow? definitely

shocked surface? definitely

episodic outflows? maybe

ALMA will help: mapping the 321 GHz H₂O masers
(Grade “C” proposal, PI: A. Perez-Sanchez)

Now some marketing and community notice...