

GLOBAL CONSTRAINTS ON SEESAW NEUTRINO MIXING

JOSU HERNANDEZ-GARCIA

Based on:

- JHEP 1608 (2016) 033: E. Fernandez-Martinez, JHG and J. Lopez-Pavon
- JHEP 1510 (2015) 130: E. Fernandez-Martinez, JHG, J. Lopez-Pavon and M. Lucente



OUTLINE

- MOTIVATION
- INTRODUCTION
- THE TWO SCENARIOS: G-SS & 3N-SS
- PARAMETRIZATIONS
- OBSERVABLES
- RESULTS OF THE GLOBAL FIT
- 1-LOOP EFFECT
- SUMMARY

MOTIVATION

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Neutrino masses are one of the most promising open windows to physics beyond the Standard Model (SM).

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By adding heavy ν_R to SM particle content, neutrino masses arise in a simple and natural way.



$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{2} \overline{N_R^i} (M_N)_{ij} N_R^{cj} \\ - (Y_N)_{i\alpha} \overline{N_R^i} \phi^\dagger \ell_L^\alpha + \text{h.c.}$$

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By adding heavy ν_R to SM particle content, neutrino masses arise in a simple and natural way.

A set of EW and flavor observables are going to be used to constrain the additional neutrino mixing.

INTRODUCTION

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Once the new heavy states are integrated out, the SM-Seesaw can be considered as a low energy effective theory:

- dim-5 Weinberg op. gives masses to the light ν :

$$\frac{c_{\alpha\beta}^{\text{dim-5}}}{\Lambda} \left(\overline{L^c}_\alpha \tilde{\phi}^* \right) \left(\tilde{\phi}^\dagger L_\beta \right) \xrightarrow{\text{EWSB}} \hat{m} = m_D^t M_N^{-1} m_D$$

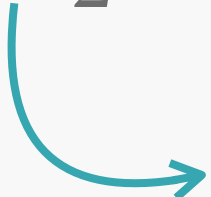
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- dim-6 op. induces non-unitarity in the mixing matrix N of lepton charged current interactions:

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$$N = (I - \eta) U_{\text{PMNS}}$$

since η is Hermitian \Rightarrow the most general parametrization for N .

INTRODUCTION


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violates L

dim-6: $\eta = \frac{1}{2} m_D^\dagger M_N^{-2} m_D$

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If smallness of m_ν comes **only** from the **suppression** with M_N

mixing η much more **suppressed** 

experimental verification extremely **challenging** 

INTRODUCTION

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Meaningful bounds imply $\begin{cases} M_i \sim \mathcal{O}(\Lambda_{\text{EW}}) \\ Y_N \sim \mathcal{O}(1) \end{cases} \Rightarrow \hat{m} \text{ too large}$

Alternatively, **smallness** of m_ν may naturally stem from an **approximate** L instead of a huge hierarchy of masses

 Inverse or Lienar Seesaw

R. Mohapatra and J. Valle, Phys.Rev. **D34**, 1642 (1986)

J. Bernabeu, A. Santamaria, J. Vidal, A. Mendez, and J. Valle,
Phys. Lett. **B187**, 303 (1987)

G.C. Branco, W. Grimus, and L. Lavoura, Nucl. Phys. **B312**, 492 (1989)

INTRODUCTION

In particular:

$$m_D = \frac{v_{\text{EW}}}{\sqrt{2}} \begin{pmatrix} \begin{matrix} \nu_e & \nu_\mu & \nu_\tau \\ 1 & 1 & 1 \\ Y_e & Y_\mu & Y_\tau \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{matrix} & \begin{matrix} 1 & N_1 \\ -1 & N_2 \\ 0 & N_3 \end{matrix} \end{pmatrix} \quad M_N = \begin{pmatrix} \begin{matrix} N_1 & N_2 & N_3 \\ 1 & -1 & 0 \\ 0 & \Lambda & 0 \\ \Lambda & 0 & 0 \\ 0 & 0 & \Lambda' \end{matrix} & \begin{matrix} 1 & N_1 \\ -1 & N_2 \\ 0 & N_3 \end{matrix} \end{pmatrix}$$

where N_i is an arbitrary number of extra heavy fields.

If L is **exact**: $\hat{m} = 0$ while $\eta \neq 0$ and arbitrarily large.

R. Alonso, M. Dhen, M. Gavela, and T. Hambye, JHEP **1301**, 118 (2013)

A. Abada, D. Das, A. Teixeira, A. Vicente, and C. Weiland, JHEP **1302**, 048 (2013)

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where N_i is an arbitrary number of extra heavy fields.

If ϵ_i and μ_j small \nexists terms introduced $\Rightarrow L$ mildly broken $\Rightarrow \begin{cases} \hat{m} \neq 0 \\ m_i \sim \mathcal{O}(\text{eV}) \end{cases}$

while $\eta \neq 0$ and arbitrarily large.

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 - SM is extended with an **arbitrary** number of ν_R
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 - **no** further assumptions
- 3N-SS: a 3 heavy neutrino scenario
 - SM is only extended with **3** ν_R
 - they are **heavier** than Λ_{EW}
 - **large** New Physics effects in spite of the smallness of m_ν
 - m_ν **radiatively stable**

PARAMETRIZATIONS

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- G-SS: a completely general scenario

Where N parametrized by:

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represents the total mixing from all N_{R_i} with the flavor α .

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- 3N-SS: a 3 heavy neutrino scenario

Where the only Seesaw that **saturates** the bounds:

$M_{1,2} \sim \Lambda$ (pseudo Dirac pair), $M_3 \sim \Lambda'$ (decoupled) but

$$\Theta \simeq \frac{1}{\sqrt{2}} \begin{pmatrix} -i\theta_e & \theta_e & 0 \\ -i\theta_\mu & \theta_\mu & 0 \\ -i\theta_\tau & \theta_\tau & 0 \end{pmatrix} \quad \text{with } \theta_\alpha \equiv \frac{vY_\alpha}{\sqrt{2}\Lambda}$$

Since $\eta = \frac{\Theta\Theta^\dagger}{2}$  $\eta = \frac{1}{2} \begin{pmatrix} |\theta_e|^2 & \theta_e\theta_\mu^* & \theta_e\theta_\tau^* \\ \theta_\mu\theta_e^* & |\theta_\mu|^2 & \theta_\mu\theta_\tau^* \\ \theta_\tau\theta_e^* & \theta_\tau\theta_\mu^* & |\theta_\tau|^2 \end{pmatrix}$

$|\eta_{\alpha\beta}| = \sqrt{\eta_{\alpha\alpha}\eta_{\beta\beta}}$  Schwarz inequality is saturated


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Fixing ν osc. data: θ_{ij} & $\Delta m_{ij}^2 \Rightarrow Y_\tau = Y_\tau(m_{1,3}, \delta, \phi_1, \phi_2)$

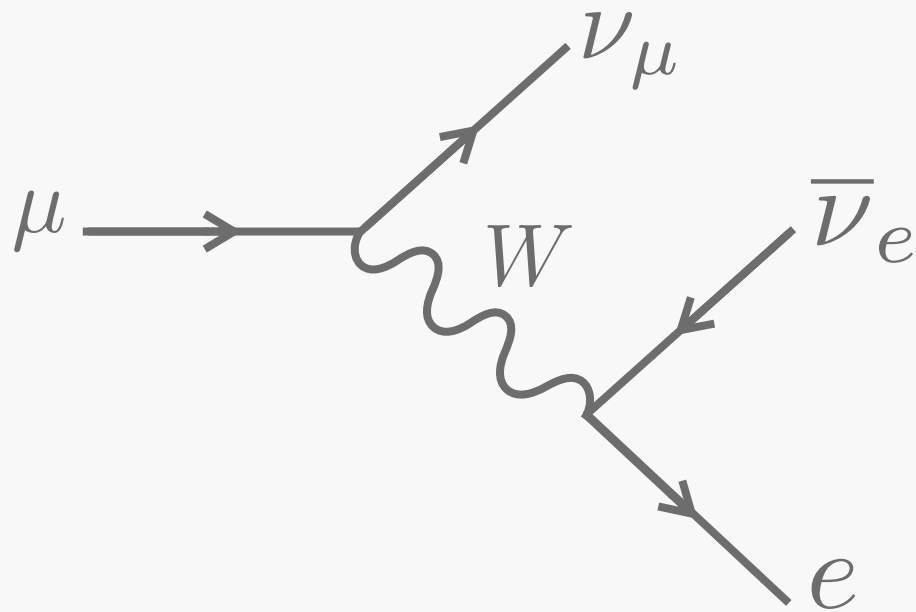
OBSERVABLES

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The 28 observables are computed in terms of α , G_μ and M_Z .

- The W boson mass M_W

G_F measured in the μ decay

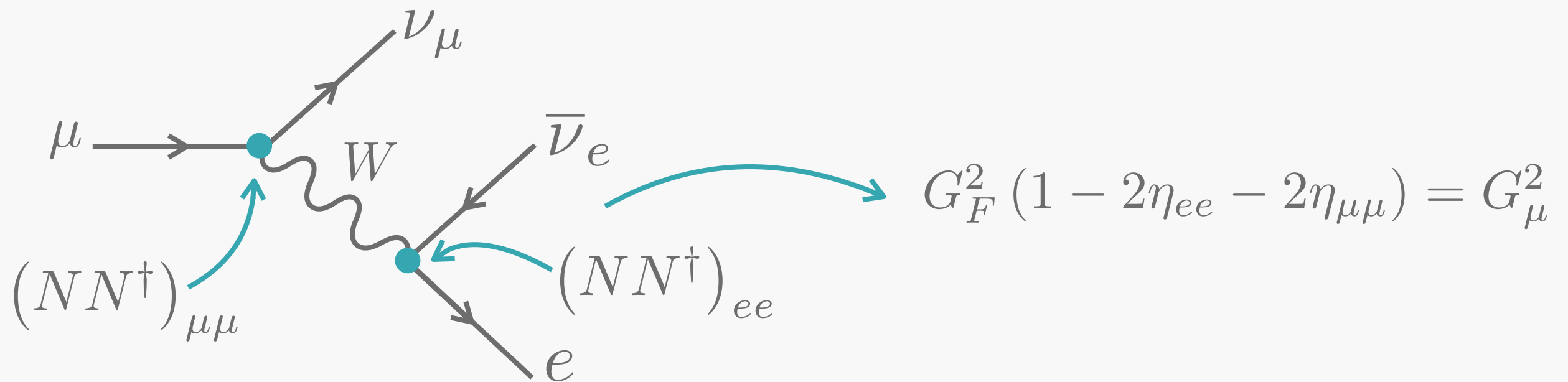


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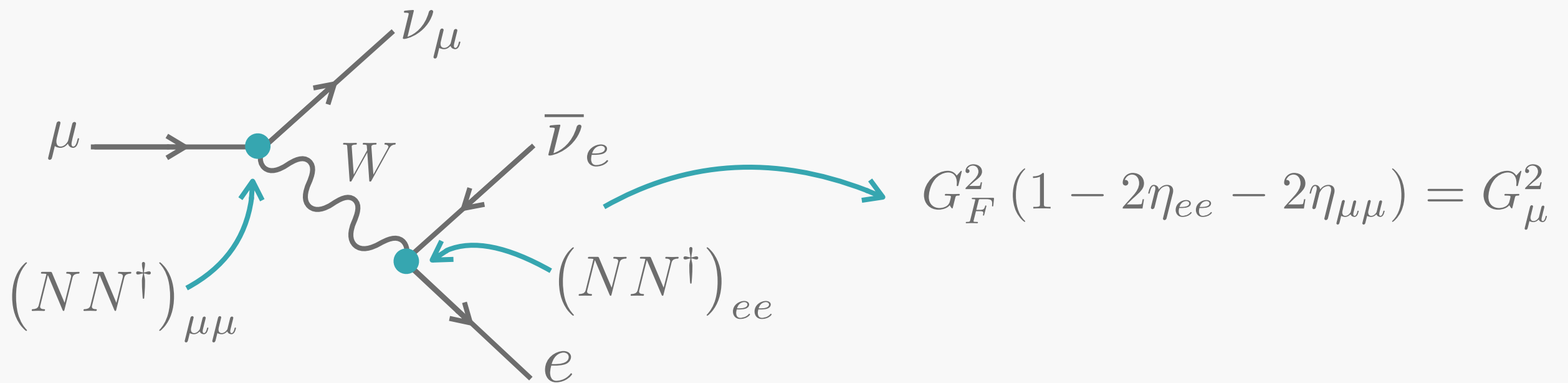


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$$M_W^2 = \frac{\pi\alpha}{\sqrt{2}G_F s_W^2 (1 - \Delta r)} = \frac{\pi\alpha(1 - \eta_{ee} - \eta_{\mu\mu})}{\sqrt{2}G_\mu s_W^2 (1 - \Delta r)}$$

kinematical measurements of M_W constrain η_{ee} and $\eta_{\mu\mu}$

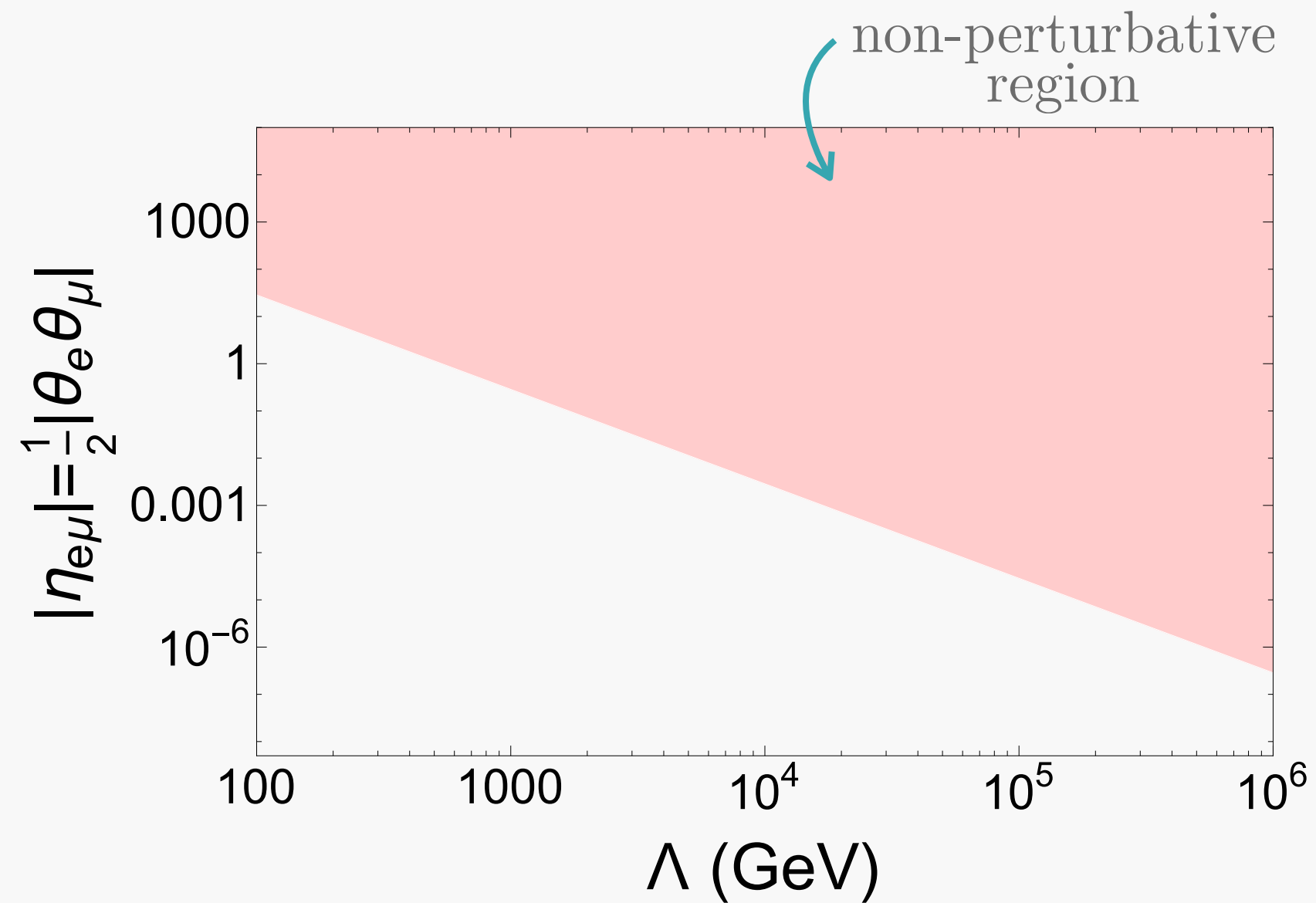
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- The W boson mass M_W
- The effective weak mixing angle θ_W : $s_{W\text{ eff}}^{2\text{ lep}}$ & $s_{W\text{ eff}}^{2\text{ had}}$
- 4 ratios of Z fermionic decays: R_l , R_c , R_b & σ_{had}^0
- The invisible Z width Γ_{inv}
- Universality ratios: $R_{\mu e}^\pi$, $R_{\tau\mu}^\pi$, $R_{\mu e}^W$, $R_{\tau\mu}^W$, $R_{\mu e}^K$, $R_{\tau\mu}^K$, $R_{\mu e}^l$ & $R_{\tau\mu}^l$
- 9 decays constraining the CKM unitarity

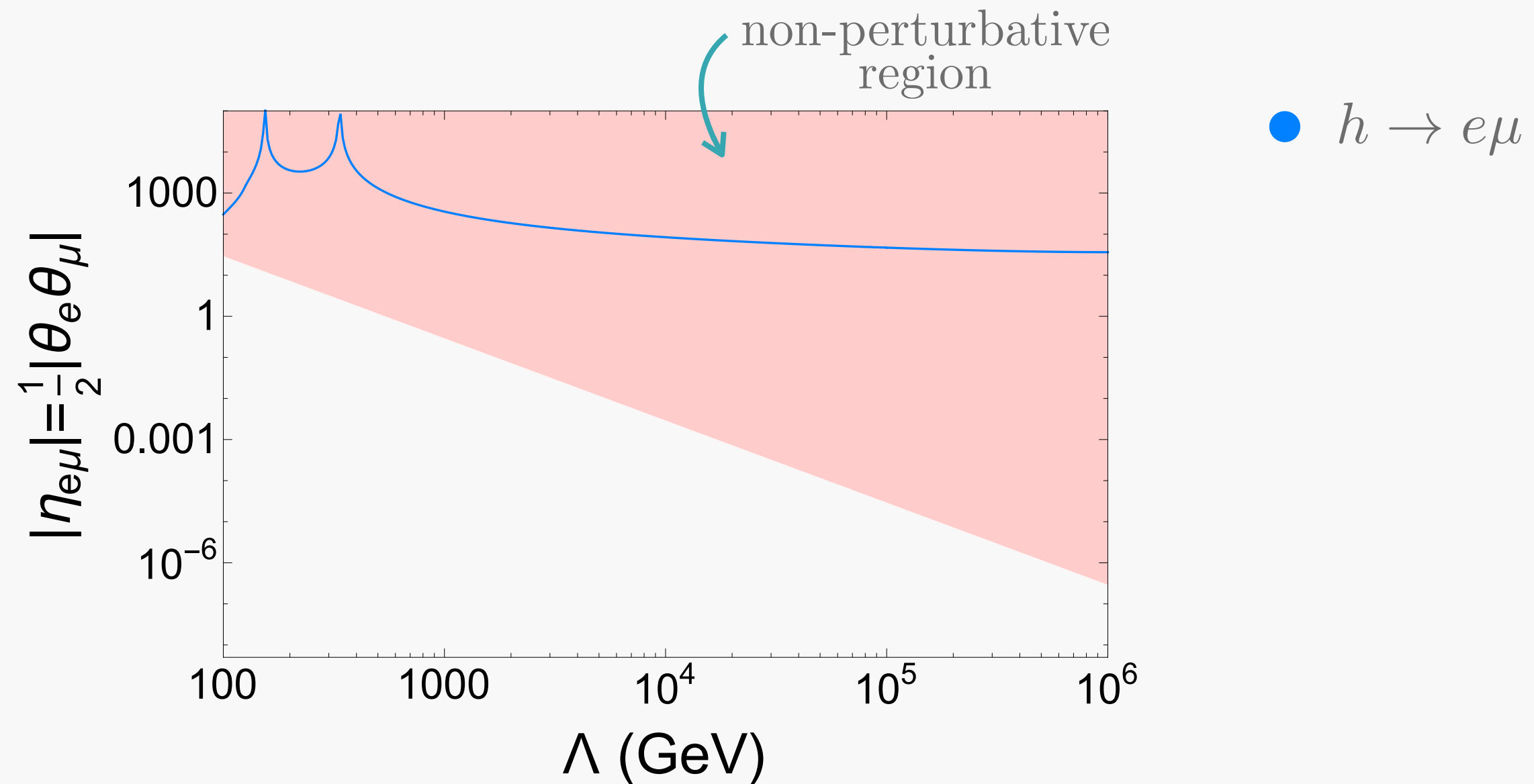
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LFV decays: $\mu - e$ transitions



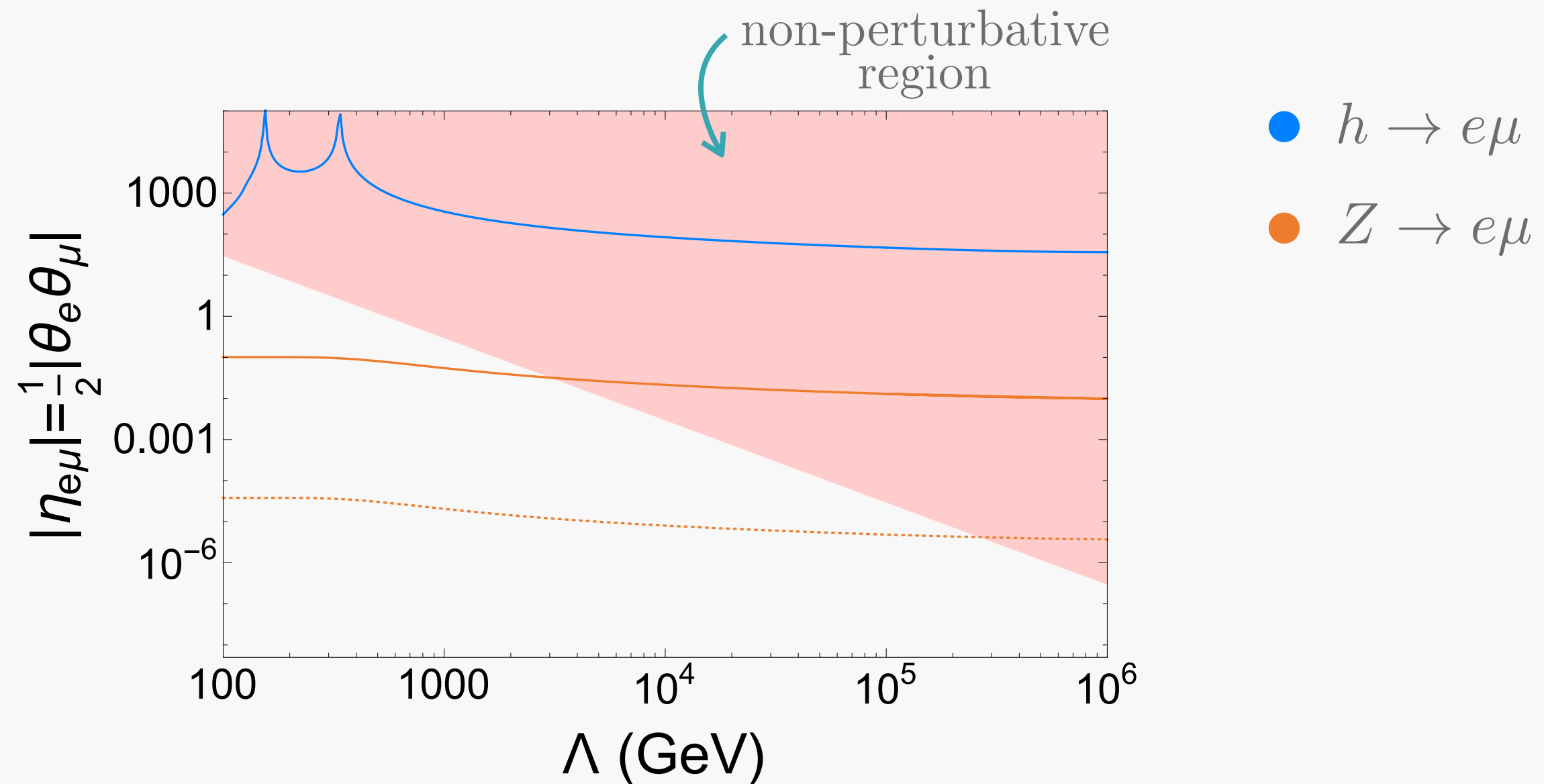
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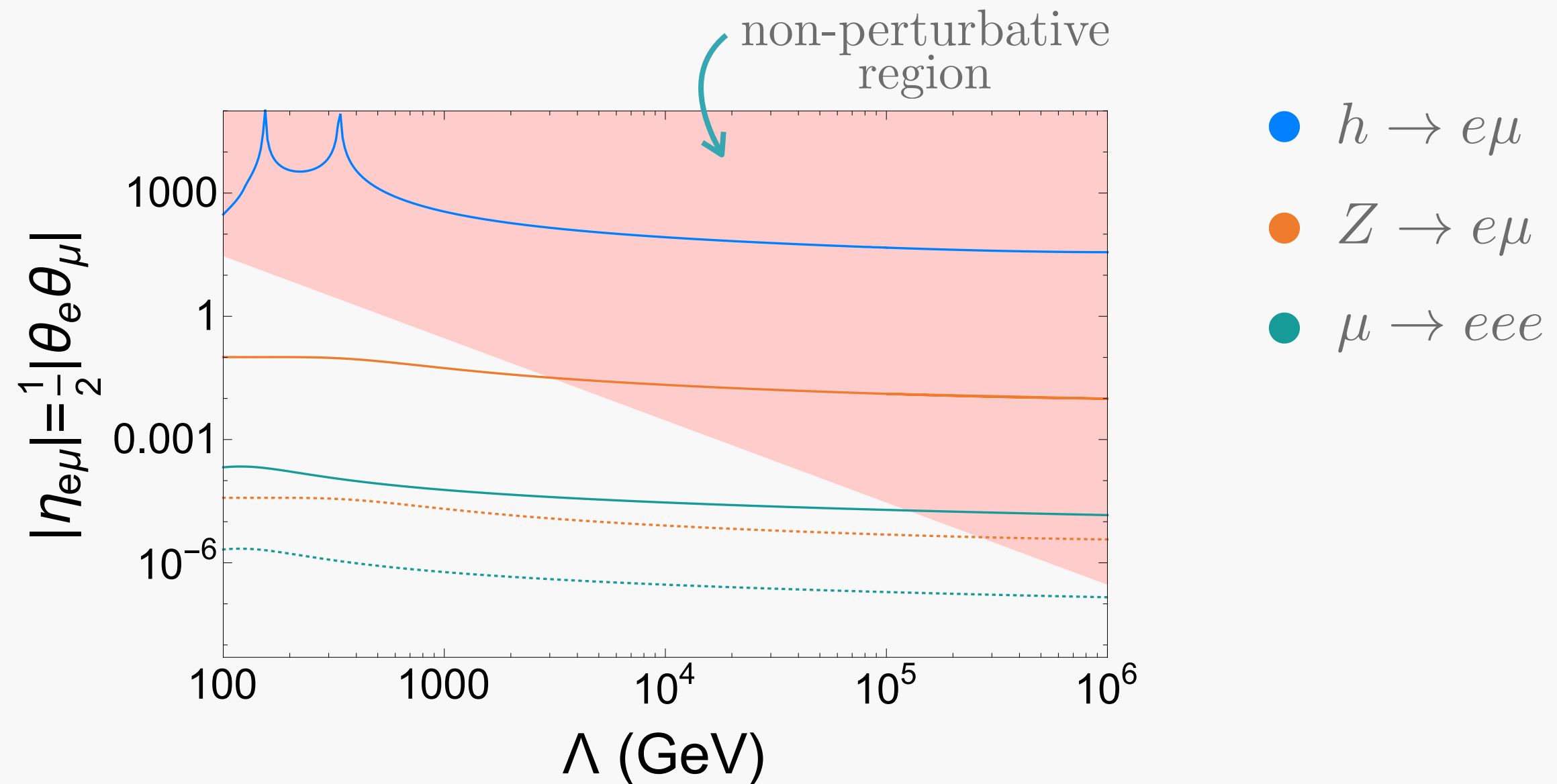
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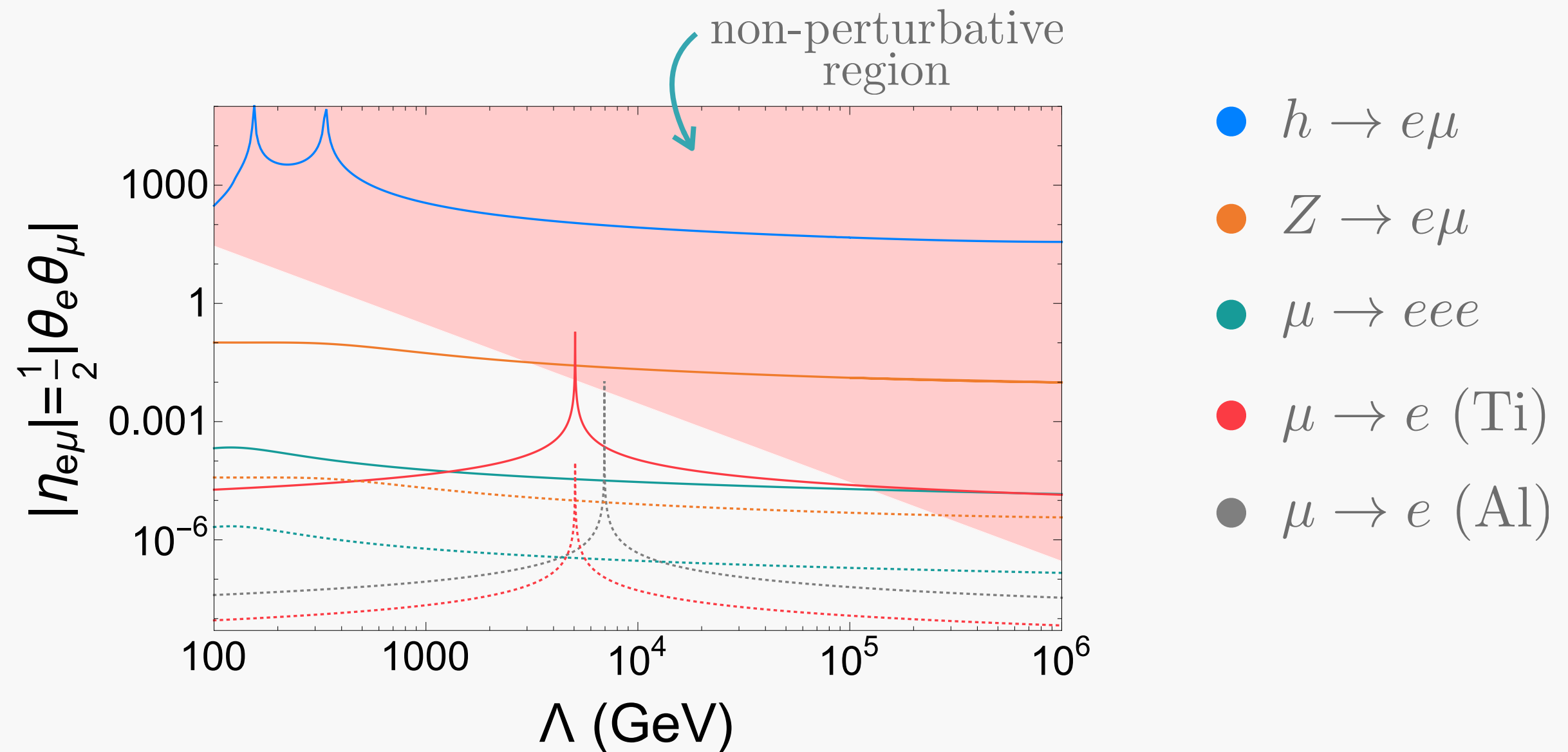
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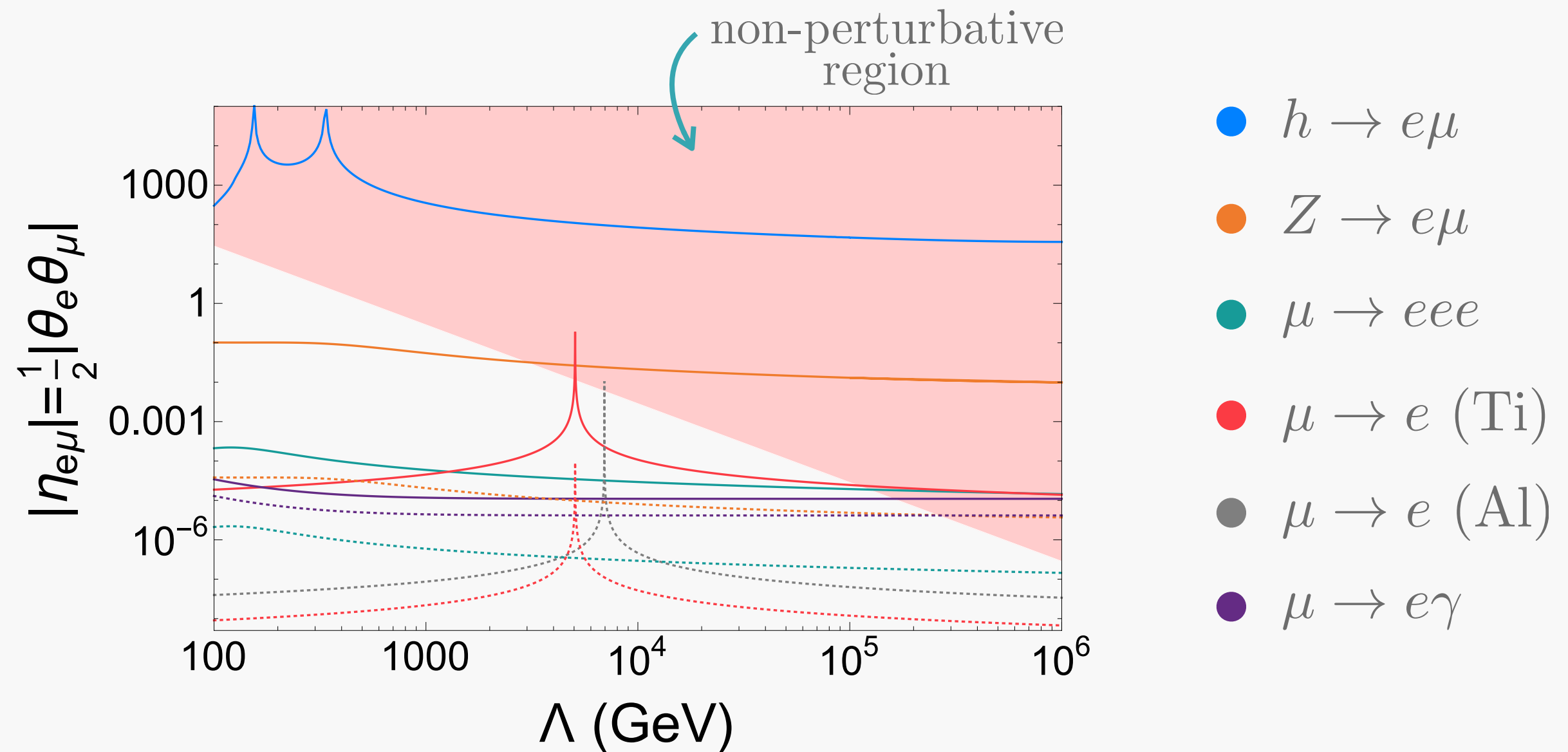
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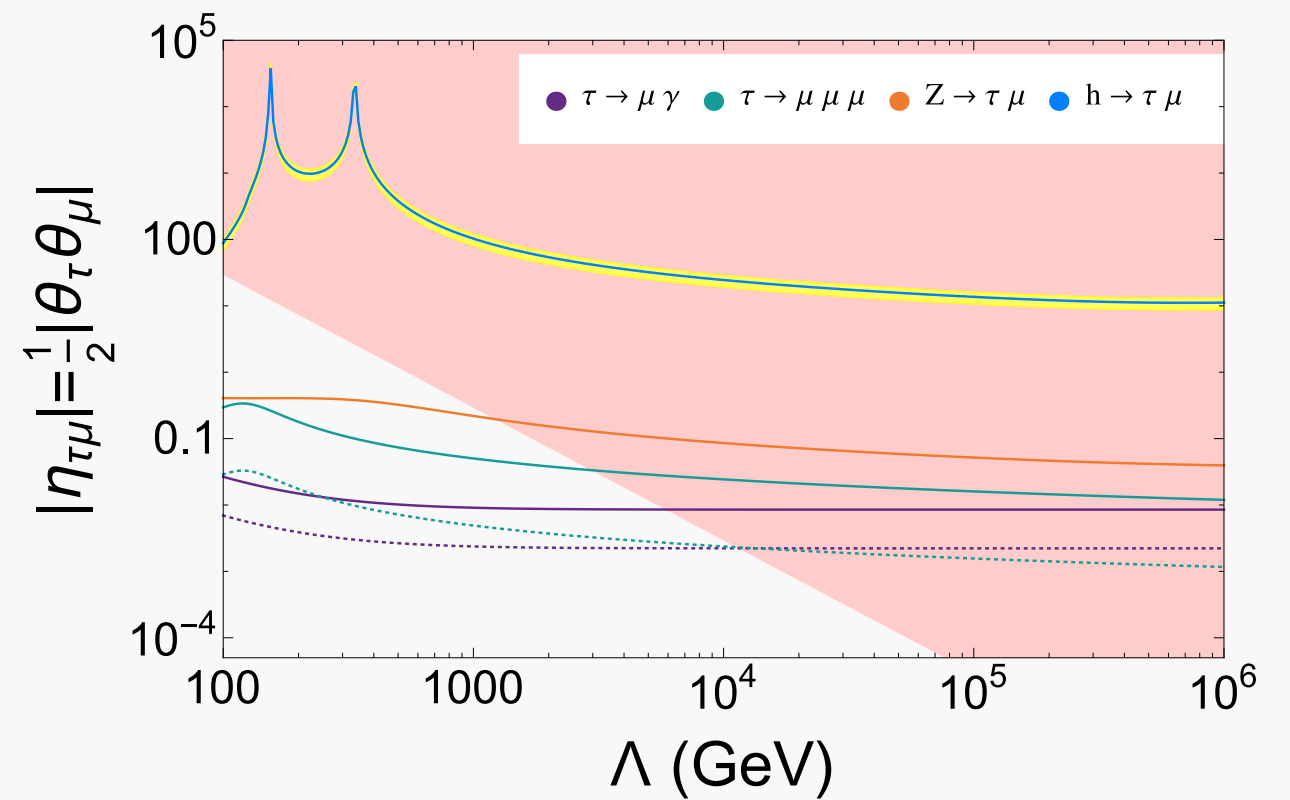
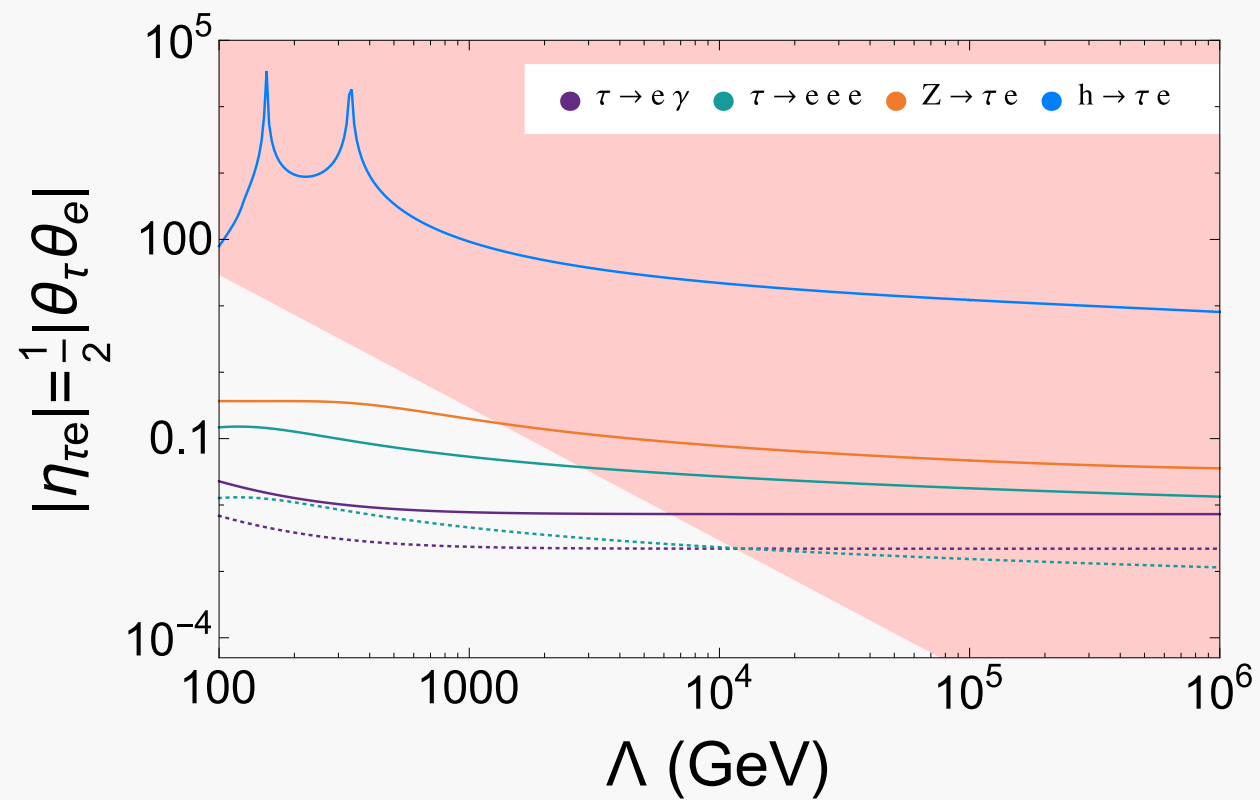
OBSERVABLES

LFV decays: $\mu - e$ transitions



OBSERVABLES

LFV decays: $\tau - e$ & $\tau - \mu$ transitions



OBSERVABLES

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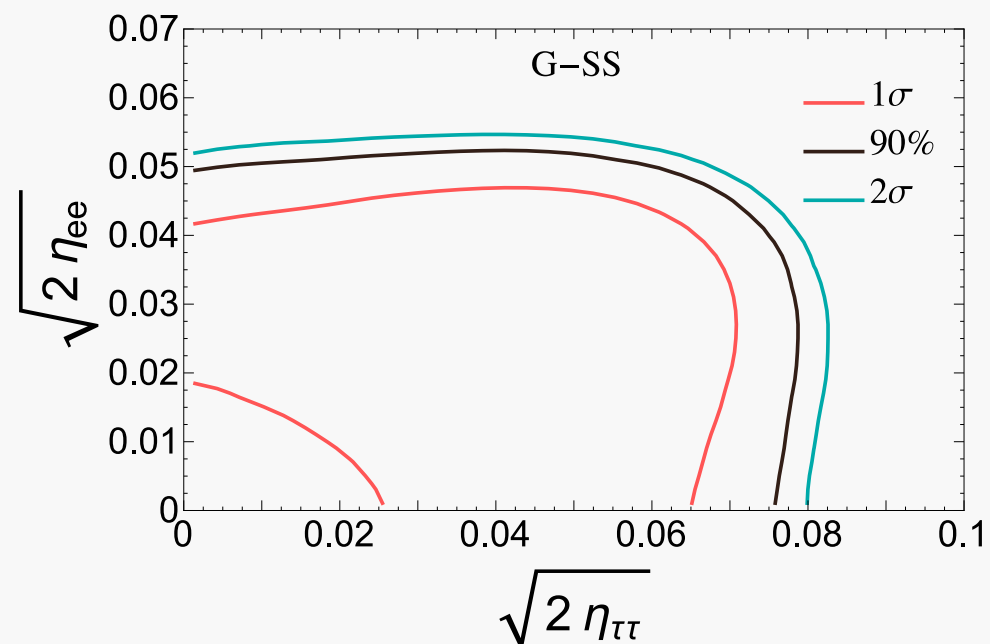
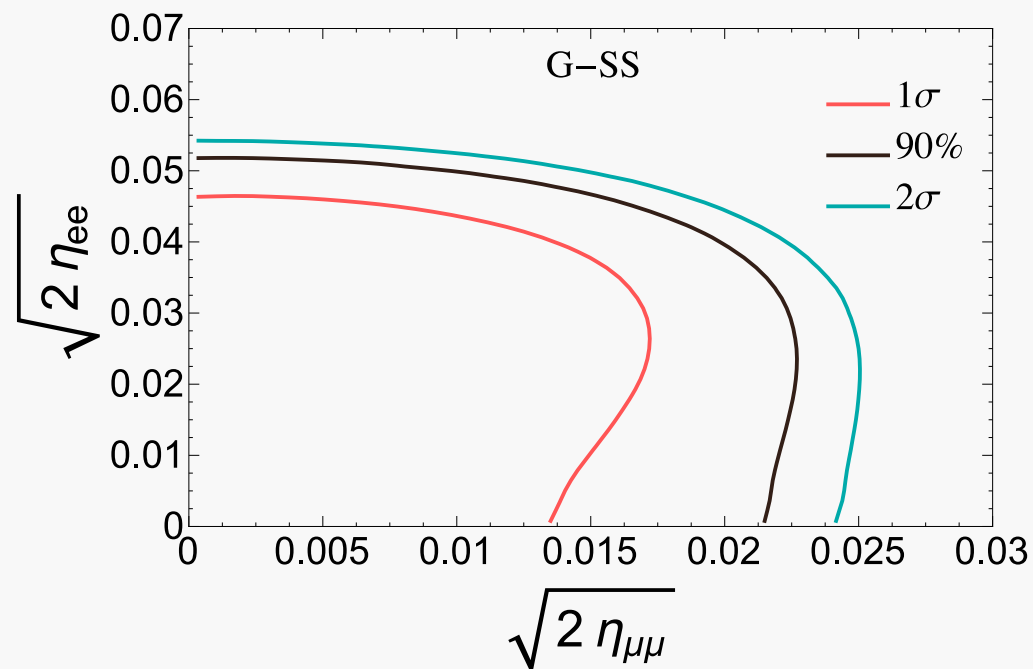
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- 9 decays constraining the CKM unitarity
- 3 rare LFV decays: $\mu \rightarrow e\gamma$, $\tau \rightarrow \mu\gamma$ & $\tau \rightarrow e\gamma$

RESULTS OF THE GLOBAL FIT

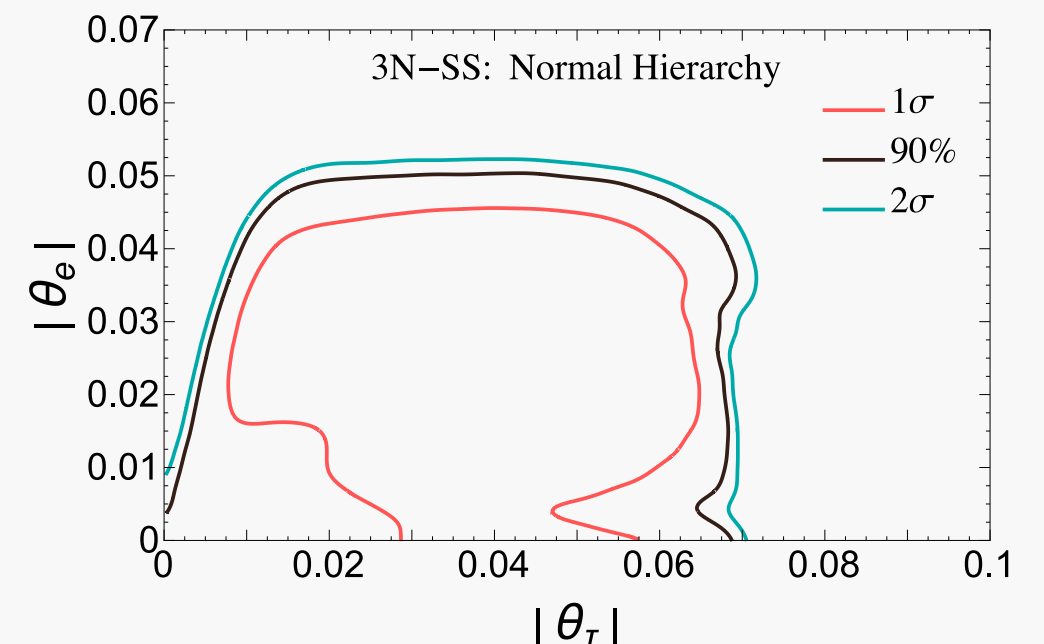
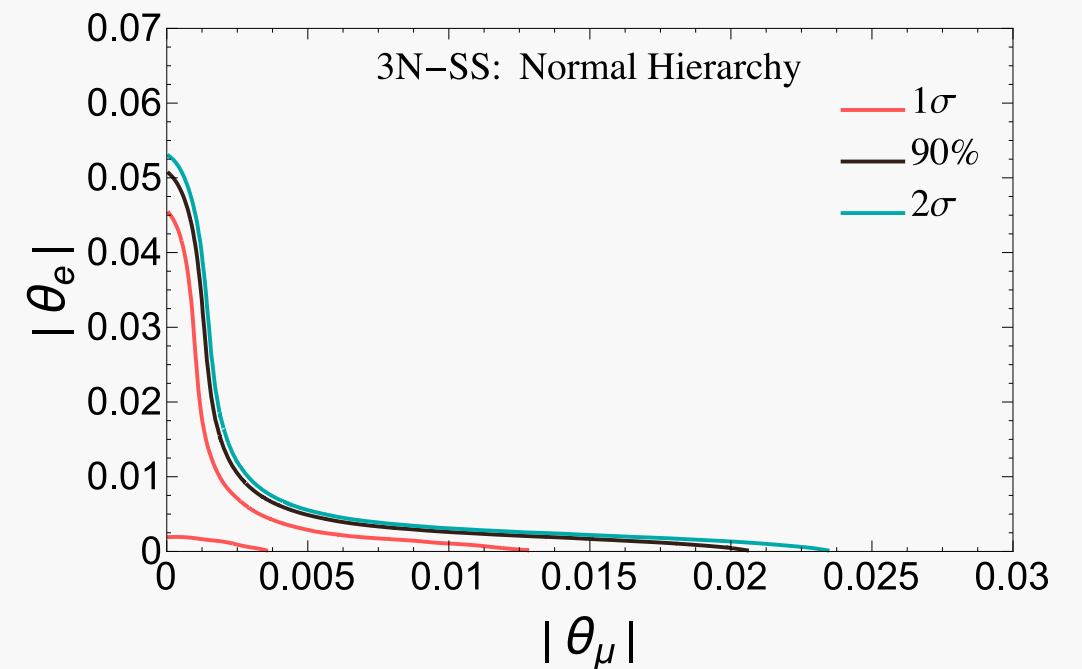
RESULTS OF THE GLOBAL FIT

MCMC with the 28 observables scanning over the free parameters

G-SS:



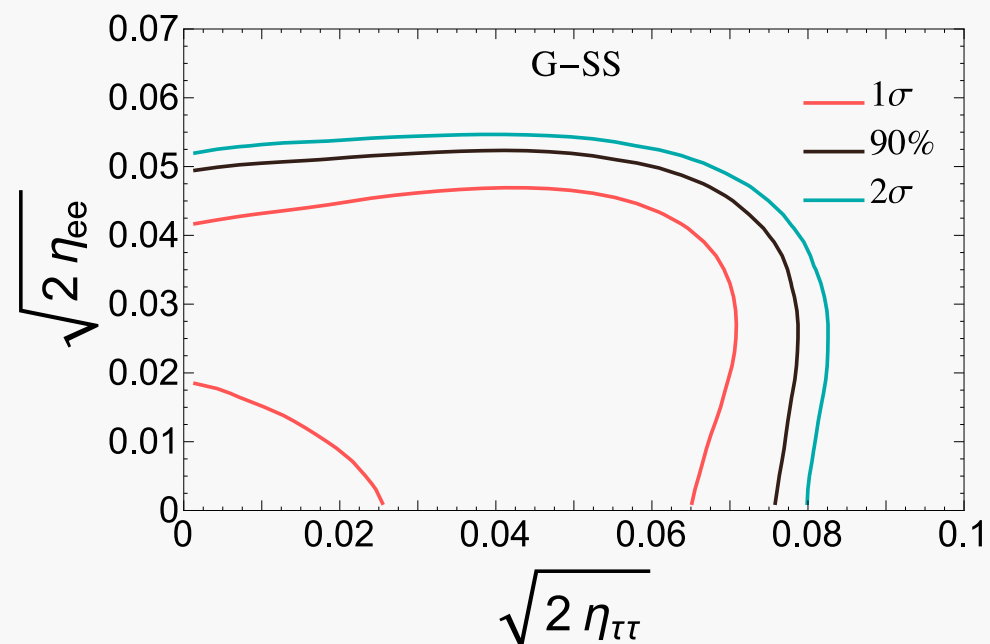
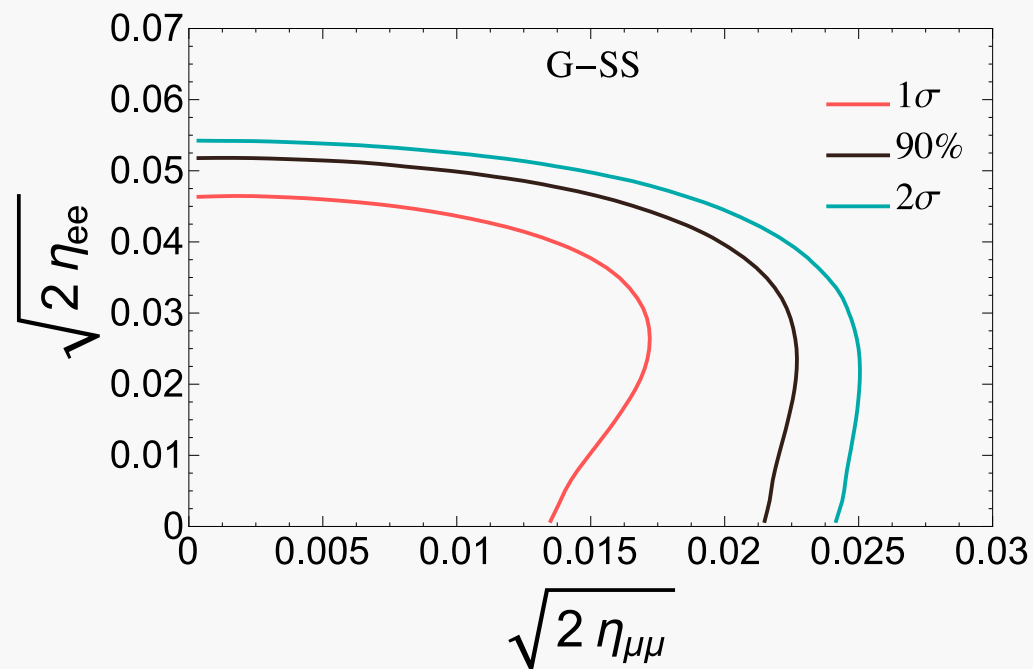
3N-SS: Normal Hierarchy



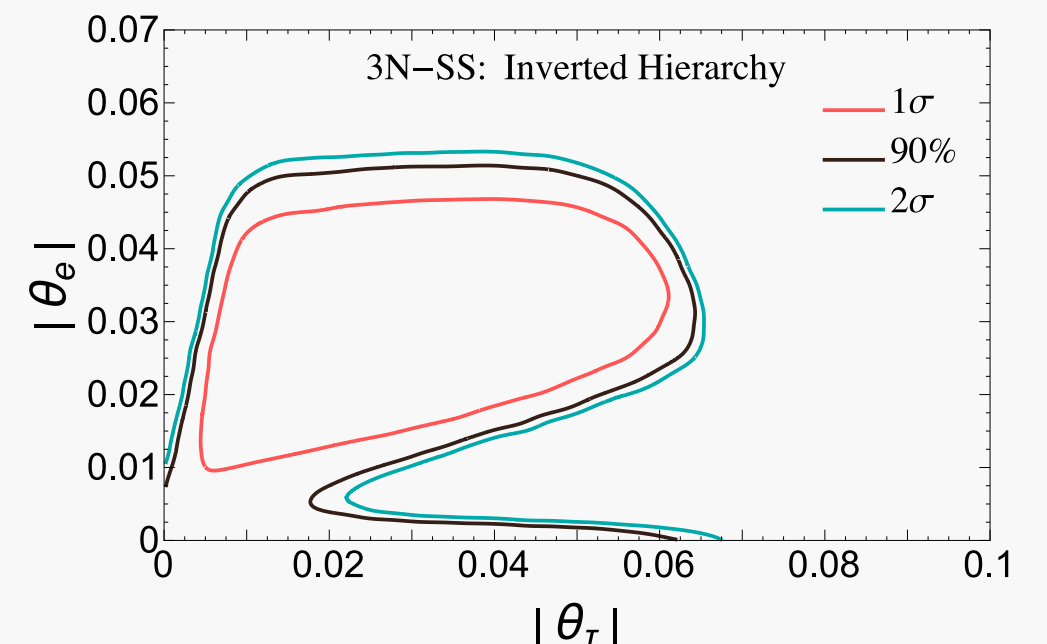
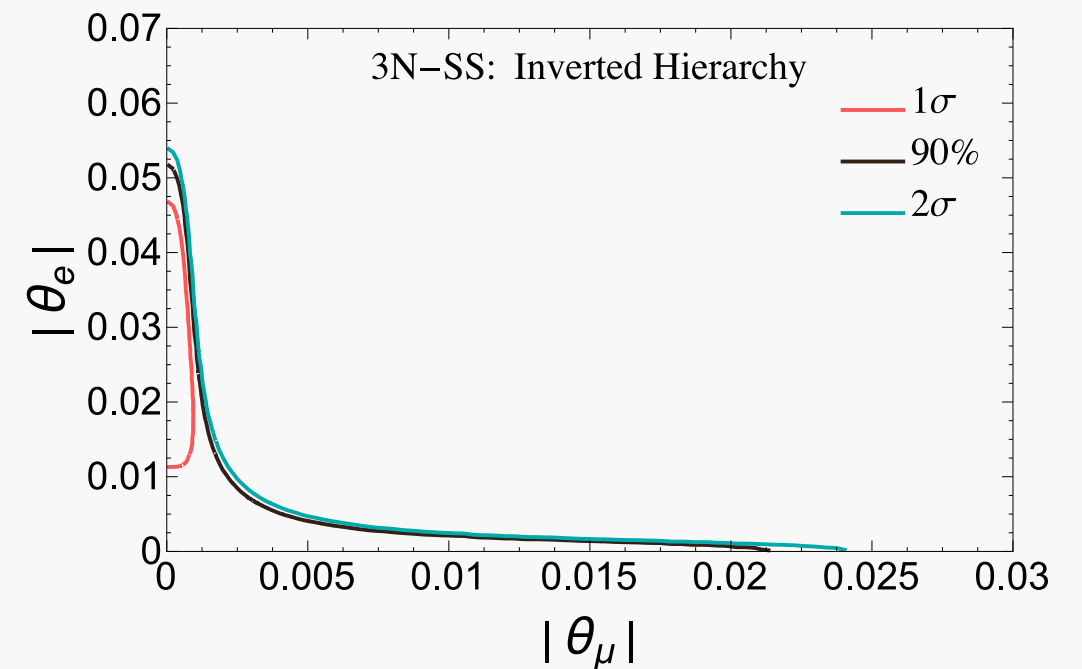
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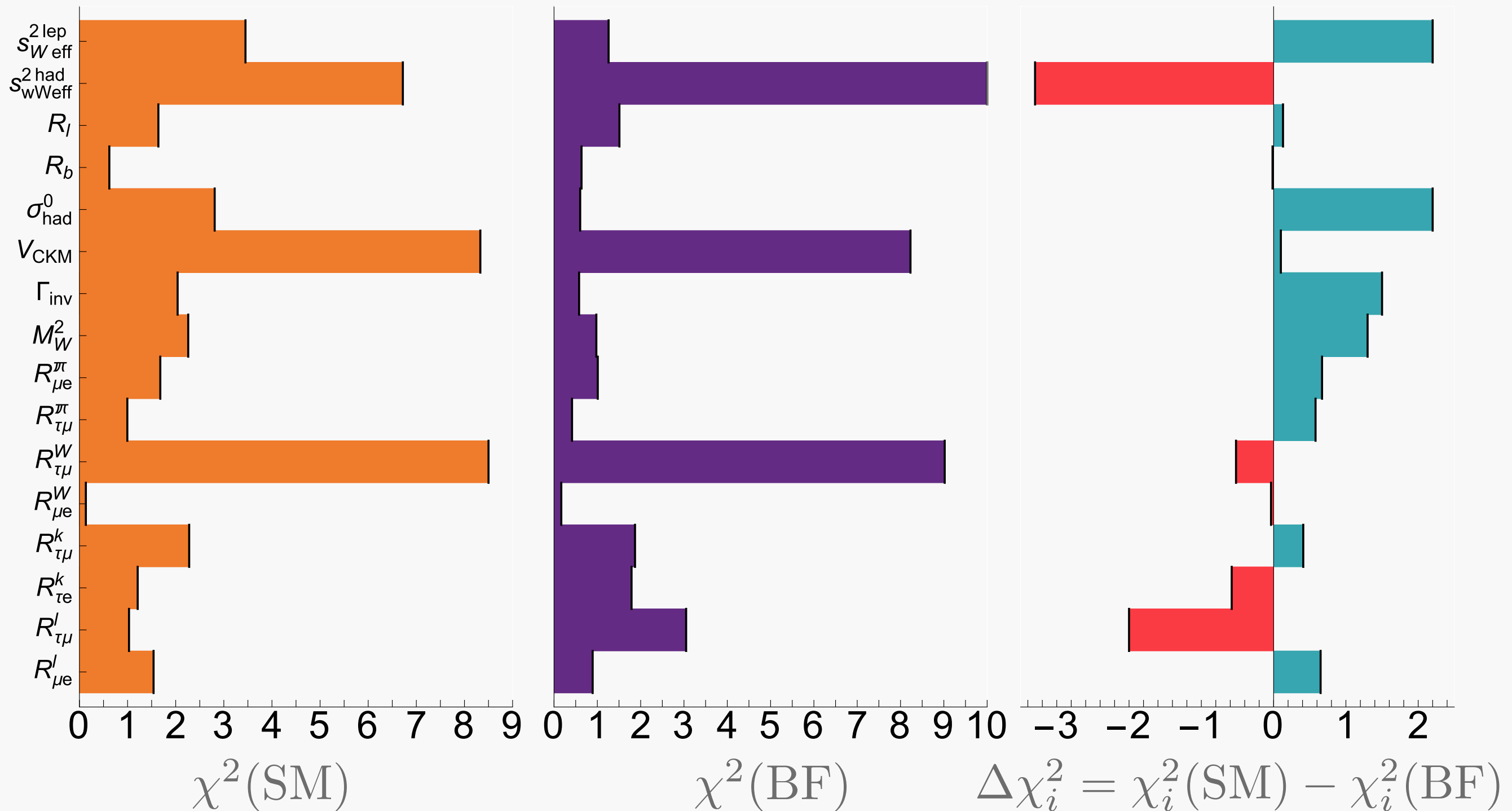


3N-SS: Inverted Hierarchy



RESULTS OF THE GLOBAL FIT

Contributions from the different observables to the χ^2 :



RESULTS OF THE GLOBAL FIT

Global fit: diagonal entries of the mixing matrix

G-SS:	3N-SS:	
	NH	IH
$\sqrt{2\eta_{ee}} = 0.031^{+0.010}_{-0.020}$	$ \theta_e = 0.029^{+0.012}_{-0.020}$	$ \theta_e = 0.031^{+0.010}_{-0.012}$
$\sqrt{2\eta_{\mu\mu}} < 0.011$	$ \theta_\mu < 7.6 \cdot 10^{-4}$	$ \theta_\mu < 6.9 \cdot 10^{-4}$
$\sqrt{2\eta_{\tau\tau}} = 0.044^{+0.019}_{-0.027}$	$ \theta_\tau = 0.043^{+0.018}_{-0.027}$	$ \theta_\tau = 0.037^{+0.021}_{-0.032}$

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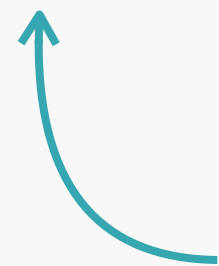
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RESULTS OF THE GLOBAL FIT

Global fit: off-diagonal entries of the mixing matrix

G-SS:		3N-SS:	
LFC	LFV	NH	IH
$\sqrt{2 \eta_{e\mu} } < 0.018$	$\sqrt{2 \eta_{e\mu} } < 4.1 \cdot 10^{-3}$	$\sqrt{ \theta_e\theta_\mu } < 4.1 \cdot 10^{-3}$	$\sqrt{ \theta_e\theta_\mu } < 4.1 \cdot 10^{-3}$
$\sqrt{2 \eta_{e\tau} } < 0.045$	$\sqrt{2 \eta_{e\tau} } < 0.107$	$\sqrt{ \theta_e\theta_\tau } = 0.036^{+0.010}_{-0.016}$	$\sqrt{ \theta_e\theta_\tau } = 0.036^{+0.010}_{-0.023}$
$\sqrt{2 \eta_{\mu\tau} } < 0.024$	$\sqrt{2 \eta_{\mu\tau} } < 0.115$	$\sqrt{ \theta_\mu\theta_\tau } < 0.007$	$\sqrt{ \theta_\mu\theta_\tau } < 0.005$

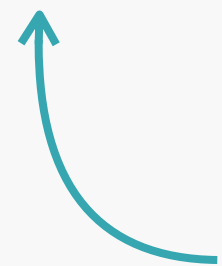


Schwarz inequality

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$\sqrt{2 \eta_{e\tau} } < 0.045$	$\sqrt{2 \eta_{e\tau} } < 0.107$	$\sqrt{ \theta_e\theta_\tau } = 0.036^{+0.010}_{-0.016}$	$\sqrt{ \theta_e\theta_\tau } = 0.036^{+0.010}_{-0.023}$
$\sqrt{2 \eta_{\mu\tau} } < 0.024$	$\sqrt{2 \eta_{\mu\tau} } < 0.115$	$\sqrt{ \theta_\mu\theta_\tau } < 0.007$	$\sqrt{ \theta_\mu\theta_\tau } < 0.005$



Schwarz inequality

1-LOOP EFFECT

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Several observables go with:

$$\frac{|\theta_e|^2}{2} + \frac{|\theta_\mu|^2}{2} + 2\alpha T \quad \curvearrowright \quad T = \frac{\Sigma_{WW}(0)}{M_W^2} - \frac{\Sigma_{ZZ}(0)}{M_Z^2}$$

W and Z boson propagators corrected by the heavy ν_R :

$$W = W + W \text{ (loop with } \ell \text{ and } N_R \text{)} \quad \Sigma_{WW}$$

The diagram illustrates the renormalization of the Z boson propagator. It shows an orange wavy line labeled Z on the left, followed by an equals sign. To the right of the equals sign is a grey wavy line labeled Z , followed by a plus sign. To the right of the plus sign is a grey wavy line labeled Z connected to a loop with two N_R labels, which is then connected to another grey wavy line labeled Z . An orange arrow points to the loop with the label Σ_{ZZ} .

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$$\frac{|\theta_e|^2}{2} + \frac{|\theta_\mu|^2}{2} + 2\alpha T \quad \xrightarrow{\quad} \quad T = \frac{\Sigma_{WW}(0)}{M_W^2} - \frac{\Sigma_{ZZ}(0)}{M_Z^2}$$

W and Z boson propagators corrected by the heavy ν_R :

$$\begin{aligned} \text{Wavy } W &= \text{Wavy } W + \text{Wavy } W \text{---} \text{Circle} \text{---} W \\ &\quad \text{with } \ell \text{ on top and } N_R \text{ on bottom, labeled } \Sigma_{WW} \\ \text{Wavy } Z &= \text{Wavy } Z + \text{Wavy } Z \text{---} \text{Circle} \text{---} Z \\ &\quad \text{with } N_R \text{ on top and } N_R \text{ on bottom, labeled } \Sigma_{ZZ} \end{aligned}$$

A **cancellation** between tree and loop level could be possible.

This **relaxes** some bounds **allowing** to fit some anomalies.

1-LOOP EFFECT

If L is mildly broken $\Rightarrow T \geq 0 \Rightarrow$ No cancellation allowed

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$$m_i^{\text{tree}} \sim v_{\text{EW}}^2 Y^2 \left(\frac{1}{\Lambda} \mathcal{O} \left(\epsilon_1, \frac{\mu_2}{2\Lambda} \right) + \frac{1}{\Lambda'} \mathcal{O} \left(\epsilon_2^2, \frac{\mu_4}{4\Lambda^2} \right) \right)$$

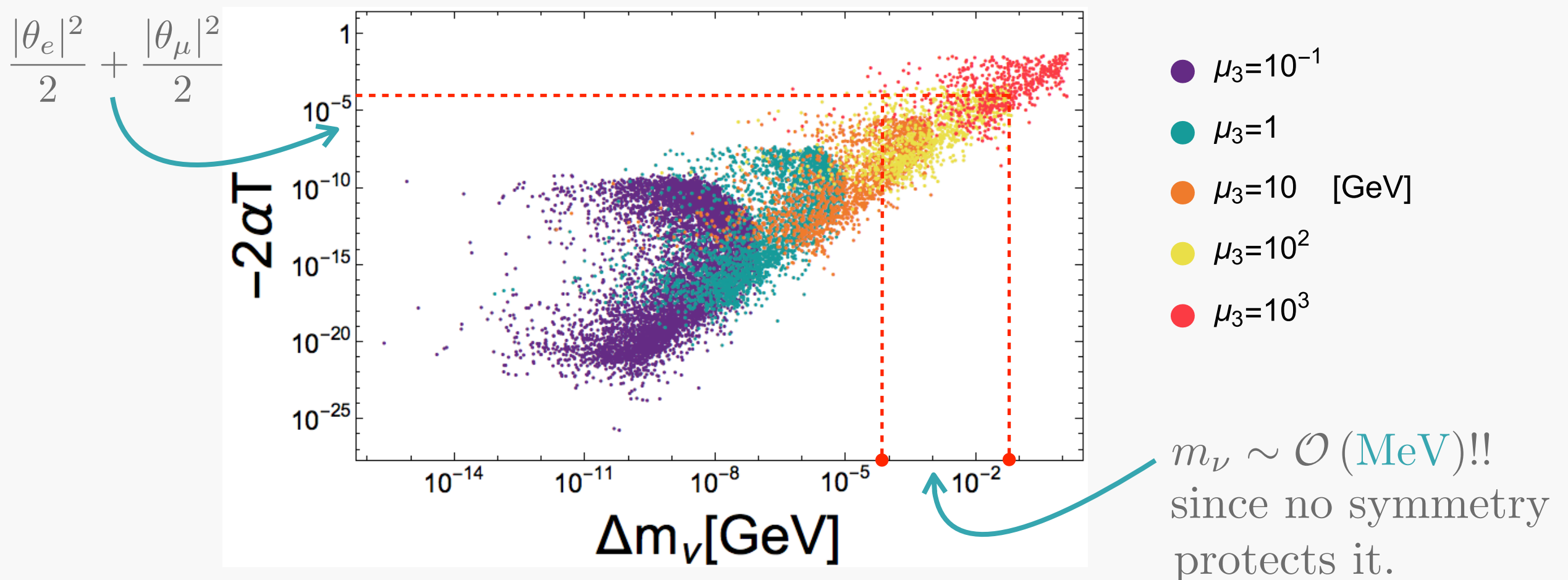
large \mathbb{L} driven by μ_1 and μ_3

$$T \simeq \frac{v_{\text{EW}}^4}{64\pi s_W^2 M_W^2} \left(\sum_{\alpha} |Y_{\alpha}|^2 \right)^2 f(\mu_1, \mu_3)$$

1-LOOP EFFECT

Loop corrections from μ_1 and μ_3 to m_i should be taken into account:

$$\Delta m_{\nu_{\alpha\beta}} = \frac{Y_\alpha Y_\beta}{32\pi^2 \mu} \left(3M_Z^2 f(\mu_1, \mu_3, M_Z) + M_h^2 f(\mu_1, \mu_3, M_h) \right)$$



SUMMARY

A set of EW and flavor observables have been used to constrain the additional mixing in two different scenarios.

A non-zero value for e and τ mixings with a significance of 2σ and an upper bound for μ mixing found in both scenarios.

In the G-SS scenario, $\eta_{e\mu}$ is constrained by $\mu \rightarrow e\gamma$ while $\eta_{\tau e}$ and $\eta_{\tau\mu}$ are constrained by indirect bounds through Schwarz inequality.

In a L -conserving Seesaw model loop effects are negligible.

THANKS