

New Physics in Astrophysical Neutrino Flavor

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Talk based on: [Phys.Rev.Lett. 115 \(2015\) 161303 \[arXiv:1506.02043\]](#)

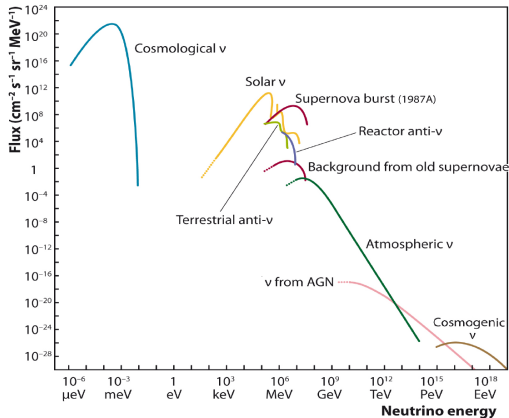
In collaboration with **Carlos Argüelles** and **Teppei Katori**

NuFact 2016



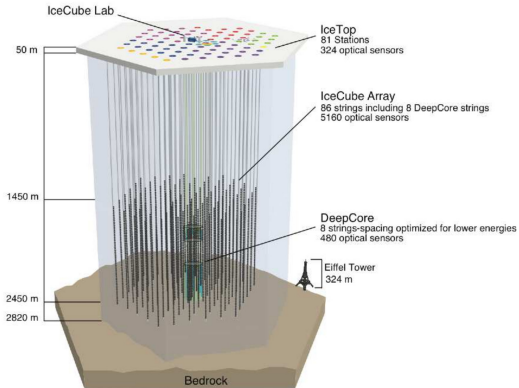
VNIVERSITAT
ID VALÈNCIA

Introduction



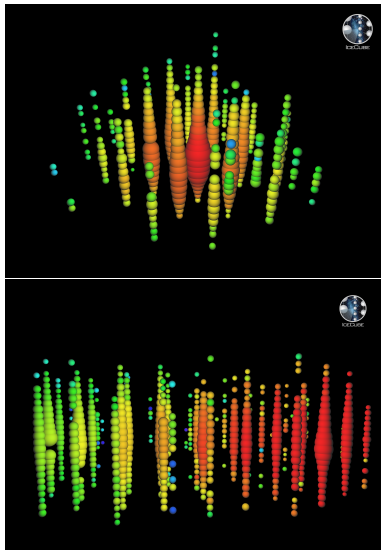
- 1 We measure neutrinos spread more than 21 orders of magnitude in energy.
- 2 The last years IceCube saw the first events in the high energy tail.

IceCube Results



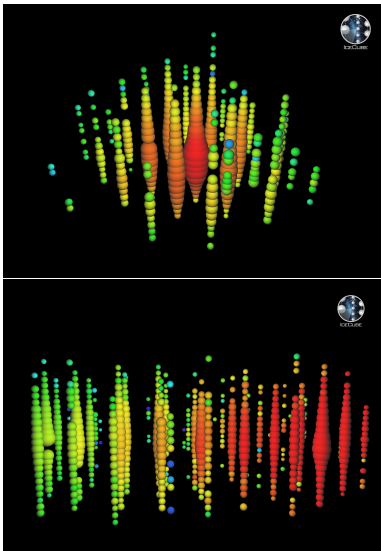
- 1km³ volume
- 86 string
- 5160 PMTs

Introduction



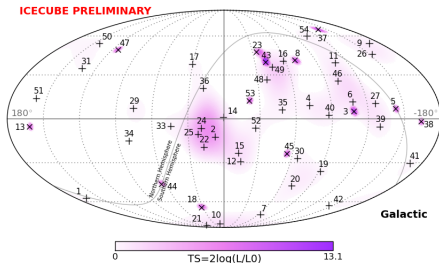
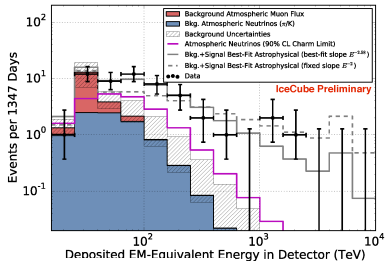
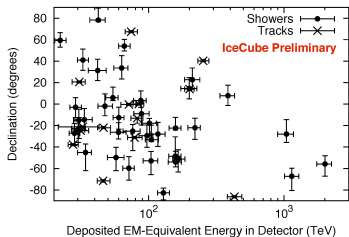
- 1 For the first time we measure **Ultra High Energy Extraterrestrial** neutrinos!
- 2 In 4 Yrs 54 events above 20TeV(Deposited)
- 3 The origin and nature of this neutrinos remains **UNKNOWN!**

Detection Physics

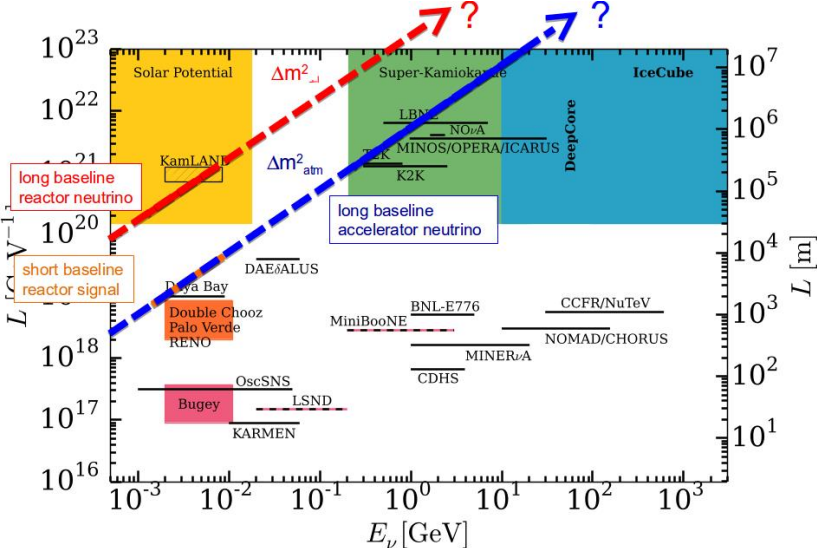


- 1 In the data we can distinguish two event topologies.
 - Tracks, produced by a ν_μ via CC interaction.
 - Cascade, produced by all the ν in NC or by ν_e and ν_τ CC.
- 2 The $\bar{\nu}_e$ may produce cascade events via the Glashow resonance $\approx 6PeV$

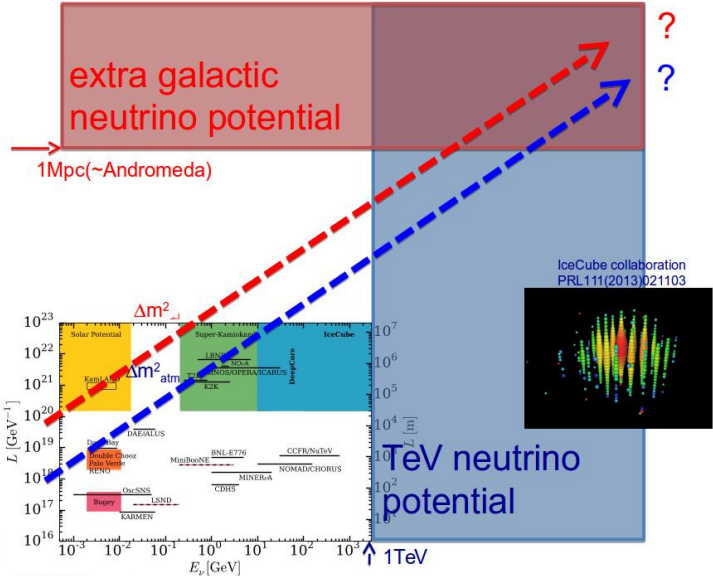
Data



Exploring new physics



Exploring new physics

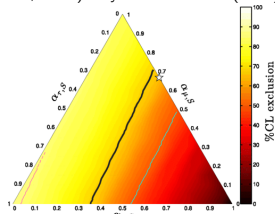


Flavor content of the Extraterrestrial Neutrinos

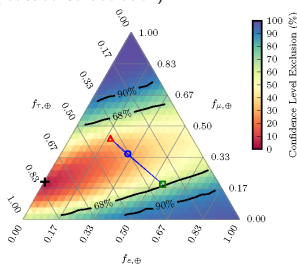
- **Flavor content**, is the ratio of the events that are electron, muon and tau neutrinos in the arrival flux.

Flavor content of the Extraterrestrial Neutrinos

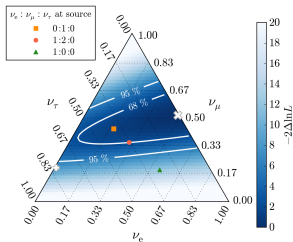
(Olga Mena, et al.) Phys.Rev.Lett. 113 (2014) 091103



(IceCube Collaboration)



(IceCube Collaboration) Phys. Rev. Lett. 114, 171102



- Measuring the flavor content is hard!
- It may depend in details about the energy spectrum. (Olga et al. arXiv:1502.02649)

Flavor content of the Extraterrestrial Neutrinos

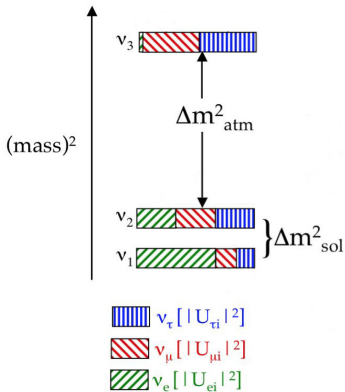
Details of the analysis that may effect the flavor ratio, and I'm not going to talk about

- Miss identification(Track-Cascade)
- Energy spectral index
- Cut-off in energy(especially before the Glashow resonance $\approx O(6PeV)$)
- ν anti $\bar{\nu}$ ratio
- We need a lot more statistics

IceCube extension needed, In the following we assume we have a measure, not necessarily very strong

Neutrino oscillations

Neutrino oscillations : mass eigenstates (ν_i ; $i = 1, 2, 3$) and flavor eigenstates (ν_α ; $\alpha = e, \mu, \tau$) are not the same.



$$\Delta m_{\text{sol}}^2 = 7.5 \times 10^{-5} \text{eV}^2$$

$$|\Delta m_{\text{atm}}^2| = 2.4 \times 10^{-3} \text{eV}^2$$

$$\nu_i = \sum_{\beta} U_{\beta i} \nu_{\beta}$$

$$U = U(\theta_{12}, \theta_{23}, \theta_{13}, \delta^{CP})$$

$$|U| \simeq \begin{pmatrix} 0.8 & 0.5 & 0.1 \\ 0.3 & 0.7 & 0.6 \\ 0.4 & 0.5 & 0.8 \end{pmatrix}$$

[B. Kayser, hep-ph/0506165 (2004)]

[M.C. Gonzalez-Garcia et al., JHEP 12 (2012)]

Flavor content after propagating for long

- The propagation distance is much longer than the oscillation length.

$$\bar{P}_{\nu_\alpha \rightarrow \nu_\beta}(E) = \sum_i |U_{\alpha i}(E)|^2 |U_{\beta i}(E)|^2 ,$$

- U is the unitary matrix that relates the propagation and flavor eigenstates.

New Physics in ν Oscillations

$$H = \frac{1}{2E}UM^2U^\dagger + \sum_n \left(\frac{E}{\Lambda_n}\right)^n \tilde{U}_n O_n \tilde{U}_n^\dagger = V^\dagger(E)\Delta V(E) \quad (1)$$

- $O_n = \text{diag}(O_{n,1}, O_{n,2}, O_{n,3})$
- $\Delta = \text{diag}(\Delta_1, \Delta_2, \Delta_3)$

O_n and Λ_n set the scale of the new physics.

New Physics in ν Oscillations

$$H = \frac{1}{2E} U M^2 U^\dagger + \sum_n \left(\frac{E}{\Lambda_n} \right)^n \tilde{U}_n O_n \tilde{U}_n^\dagger = V^\dagger(E) \Delta V(E) \quad (2)$$

This is very generic, some examples are:

n	New Physics	Current Bound From SK and IC-atm
0	CPT-odd Lorentz Violation Coupling space time torsion Non Standard Neutrino Interactions	$O_0 < 10^{-23} \text{GeV}$
1	CPT-even Lorentz Violation Violation of the equivalence principle	$O_1/\Lambda_1 < 10^{-27}$

Theory:
V. De Sabbata and M. Gasperini, Nuovo Cim. A65, 479 (1981)
V. A. Kostelecky and M. Mewes, Phys.Rev. D69, 016005 (2004), hep-ph/0309025
S. Glashow, A. Halprin, P. Krastev, C. N. Leung, and J. T. Pantaleone, Phys.Rev. D56, 2433 (1997), hep-ph/9703454
J. S. Diaz, A. Kostelecky, and M. Mewes, Phys.Rev. D89, 043005 (2014), 1308.6344

Best bounds:

K. Abe et al. (Super-Kamiokande), Phys.Rev. D91, 052003 (2015), 1410.4267
R. Abbasi et al. (IceCube), Phys.Rev. D82, 112003 (2010), 1010.4096

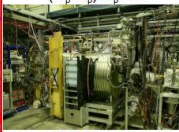
Is a Big Community

Atomic Interferometer
(a,c)^{n,p,e} < 10⁻⁶



PRL106(2011)151102

CERN Antiproton Decelerator
(M_p-M_{p̄})/M_p < 10⁻⁸



Nature419(2002)456

Spin torsion pendulum
b_e < 10⁻³⁰ GeV



PRL9

Neutrino TOF
(v-c)/c < 10⁻⁵



PRD76(2007)072005

Tevatron and LEP
-5.8x10⁻¹² < κ_ν - 4/3c_e⁰⁰ < 1.2x10⁻¹¹



PRL102(2009)170402

KTeV/KLOE (strange)
Δa_K < 10⁻²² GeV
FOCUS (charm)
Δa_c < 10⁻¹⁶ GeV
BaBar/Belle (bottom)
Δa_b/m_B < 10⁻¹⁴



Limits from all these experiments >60 page tables!
Rev.Mod.Phys.83(2011)11
ArXiv:0801.0287v9

laboratory and gravimetric tests of gravity
matter interferometry
neutrino oscillations and propagation, neutrino-antineutrino mixing
oscillations and decays of K, B, D mesons
particle-antiparticle comparisons
post-newtonian gravity in the second- and third-generation gravitational wave detectors
sidereal and annual time variations in space-based missions

JHEP11(2012)049

Cryogenic optical resonator
Δc/c < 10⁻¹⁶



GRB vacuum birefringence



PRL97(2006)140401

Double gas maser
b_n(rotation) < 10⁻³³ GeV
b_n(boost) < 10⁻²⁷ GeV

LSND



MINOS ND



MINOS FD



IceCube



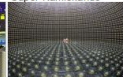
MiniBooNE



Double Chooz



Super-Kamiokande



PRD72(2005)076004 PRL101(2008)151601 PRL105(2010)151601 PRD82(2010)112003 PLB718(2013)1303 PRD86(2013)112009 PRD91(2015)052003

PRL100(2008)131802

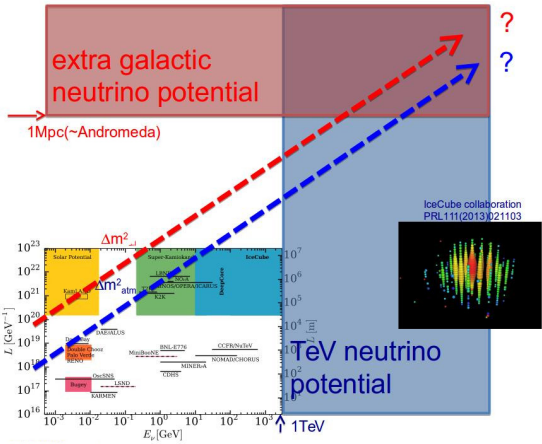
classical and quantum field theory
mathematical foundations, Fine
neutrino oscillations and pro

PRL99(2007)050401

PRL105(2010)151604
PRL107(2010)171604

Exploring new physics

$$H = \frac{1}{2E} U M^2 U^\dagger + \sum_n \left(\frac{E}{\Lambda_n} \right)^n \tilde{U}_n O_n \tilde{U}_n^\dagger = V^\dagger(E) \Delta V(E) \quad (3)$$



Showing the effect in the Flavor Ratio

Showing the effect is not trivial:

- 1 Not only the strength of the new physics is not known but also the flavor structure (new Mixing Matrices \tilde{U})
- 2 PMNS mixings have errors

Prescription

- 1 We use what we call anarchic sampling. The idea is use a flat prior in the Haar measure of $SU(3)$ for the mixing matrices to show the posterior probabilities in the flavor triangle(think Bayesian).¹

$$d\tilde{U}_n = d\tilde{s}_{12}^2 \wedge d\tilde{c}_{13}^4 \wedge d\tilde{s}_{23}^2 \wedge d\tilde{\delta} , \quad (4)$$

¹This measure is the one used to show the possible anarchic origin of the standard mixing matrix. (L. J. Hall, H. Murayama, and N. Weiner, Phys.Rev.Lett. 84, 2572 (2000), hep-ph/9911341)

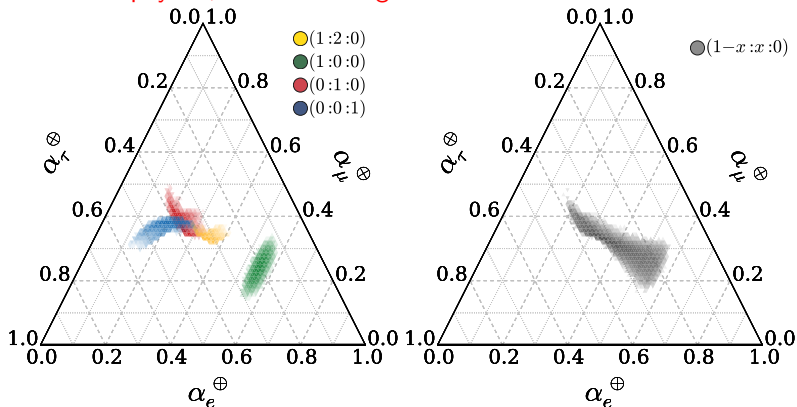
Showing the effect in the Flavor Ratio

Initial Fluxes:

- Pion Decay and Muon Energy Lost (1 : 0 : 0)
- Neutron Decay (0 : 1 : 0)
- Charged Pion Decay (1 : 2 : 0)
- Production of ν_τ (0 : 0 : 1), non know mechanism

Result for the standard oscillation parameters

Result for different initial flavor content, and for a linear combinations. Only standard ν physics, www.nu-fit.org

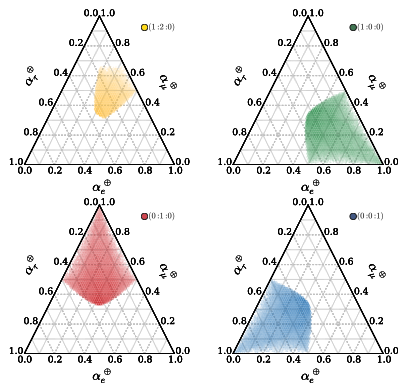


- Even not knowing the initial flux, the final ratios are in a relatively small region!

New Physics Dominated Effect

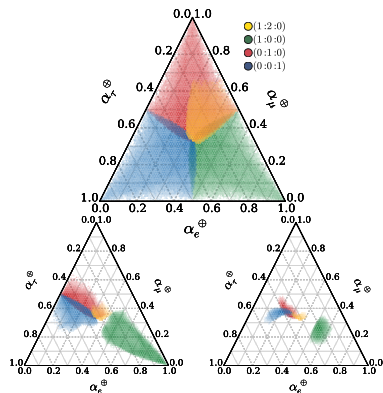
Result for different initial flavor content in the case **dominated by new physics**

$$H = \frac{1}{2E}UM^2U^\dagger + \sum_n \left(\frac{E}{\Lambda_n}\right)^n \tilde{U}_n O_n \tilde{U}_n^\dagger = V^\dagger(E)\Delta V(E) \quad (5)$$



- Even in the New Physics dominated case **not all the flavor ratio space is reachable for each initial flux.**
- Therefore, even having new physics the measure of the flavor content is giving important **information about the flavor ratio at the source.**

Intermediate Case for $n = 0$



- 1 Top $O_0 = 10^{-23}$ Current Bound

Already maximal effect!

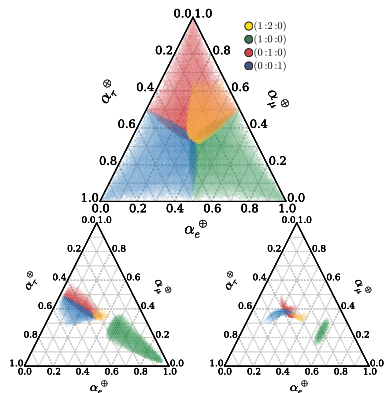
- 2 Bottom-Left
 $O_0 = 3.6 \times 10^{-26}$

Potentially measurable

- 3 Bottom-Right
 $O_0 = 6.3 \times 10^{-28}$

I don't think so

Intermediate Case for $n = 1$

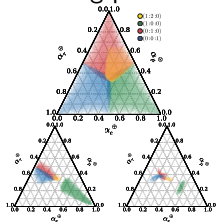


The power is not so important,
same conclusions

- 1 Top $O_1/\Lambda_1 = 10^{-27}$
Current Bound
- 2 Bottom-Left
 $O_1/\Lambda_1 = 1.0 \times 10^{-30}$
- 3 Bottom-Right
 $O_1/\Lambda_1 = 3.2 \times 10^{-34}$

Conclusions

- 1 The flavor content of the astrophysical neutrinos may give very important information about a **wide range of new physics scenarios**.
- 2 Some information about the **production flavor content is always preserved**.
- 3 The **current bounds** on this operators produce the **maximum effect** in the flavor content observable i.e. **any** trustable **measure** of the flavor content **will put the strongest bounds** on the new physics operators.
- 4 2 and 3 imply that a **precise measurement** of the flavor content will answer in a orthogonal way questions about **production and propagation**. The following positive scenarios are possible and discernible:



- Pion Decay and Muon Energy Lost + new physics (upper corner)
- Neutron Decay + new physics (lower right corner)
- Charged Pion Decay + new physics (middle right)
- Production of ν_τ + new physics. (lower left corner), it may imply new physics also in the production.