



Physics prospects and status of JUNO

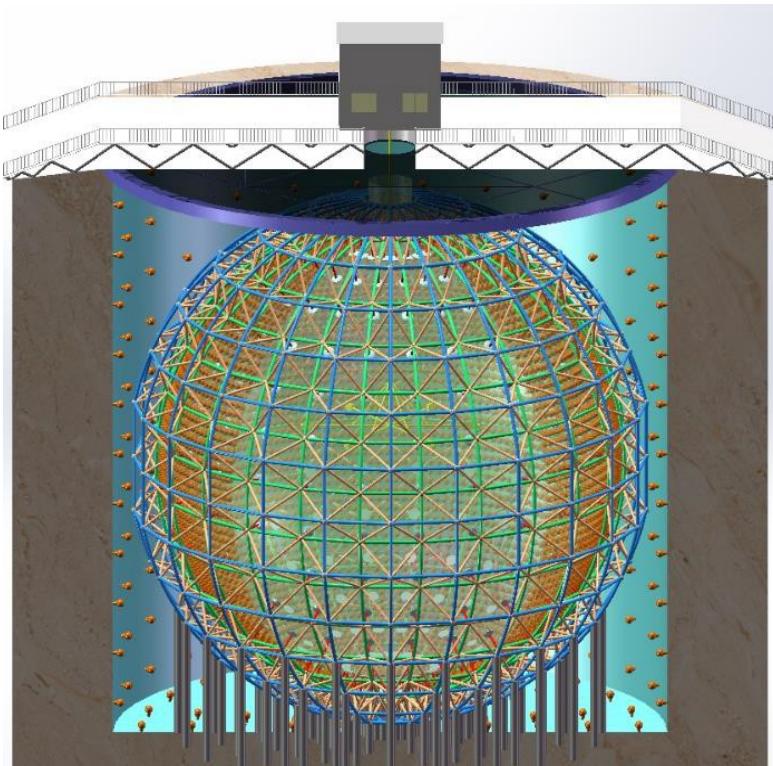


Jian Tang

(On behalf of the JUNO collaboration)

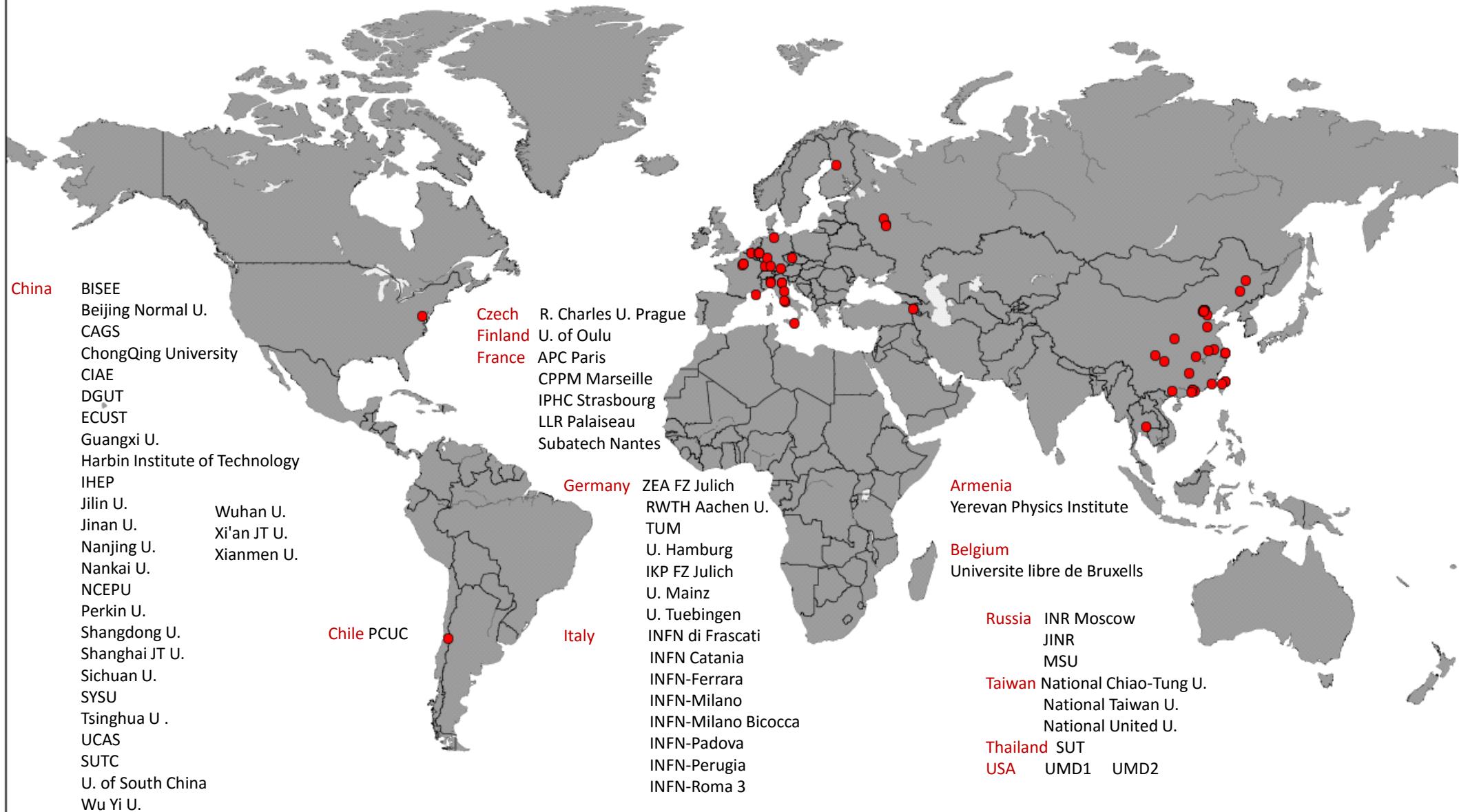
NuFact2016, Quy Nhon, Vietnam

2016.08.23





JUNO Collaboration



International Collaborations with 64 institutions and ~450 members until June in 2016



JUNO Collaboration



8th JUNO Collaboration Meeting
July 25–29, 2016, IHEP





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- **Status of JUNO**
 - **Central Detector**
 - **Liquid Scintillator**
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 - **Calibration system**
- **Summary**

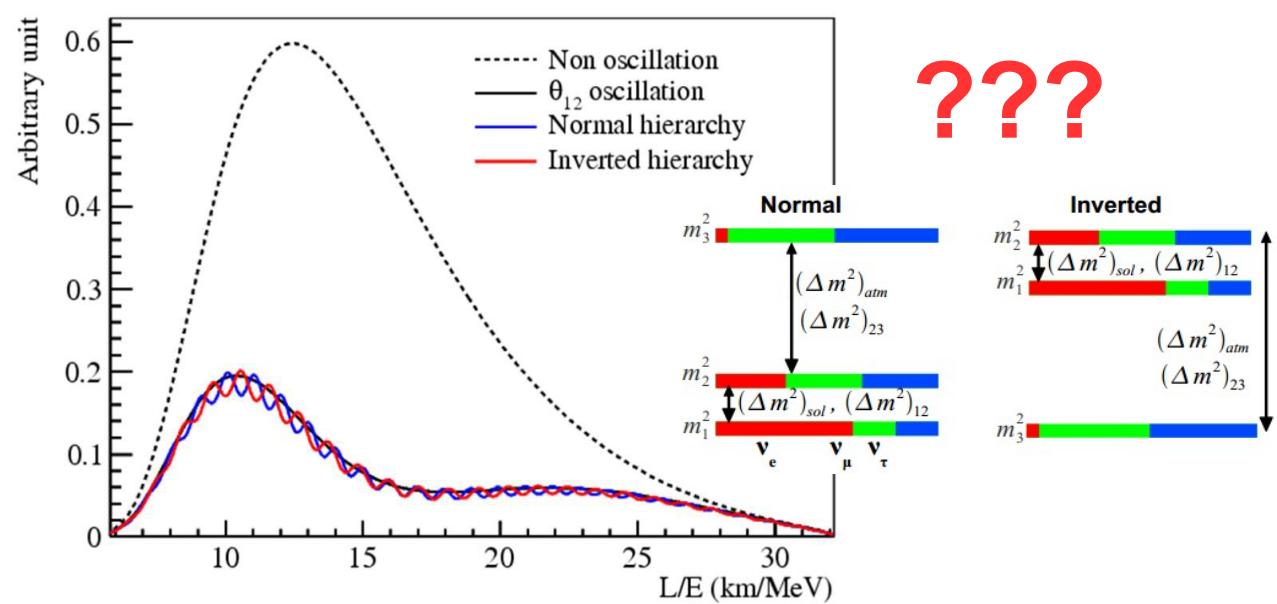
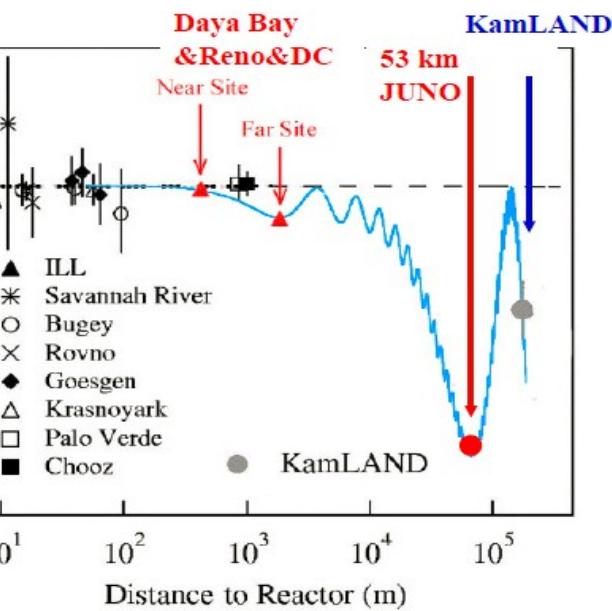
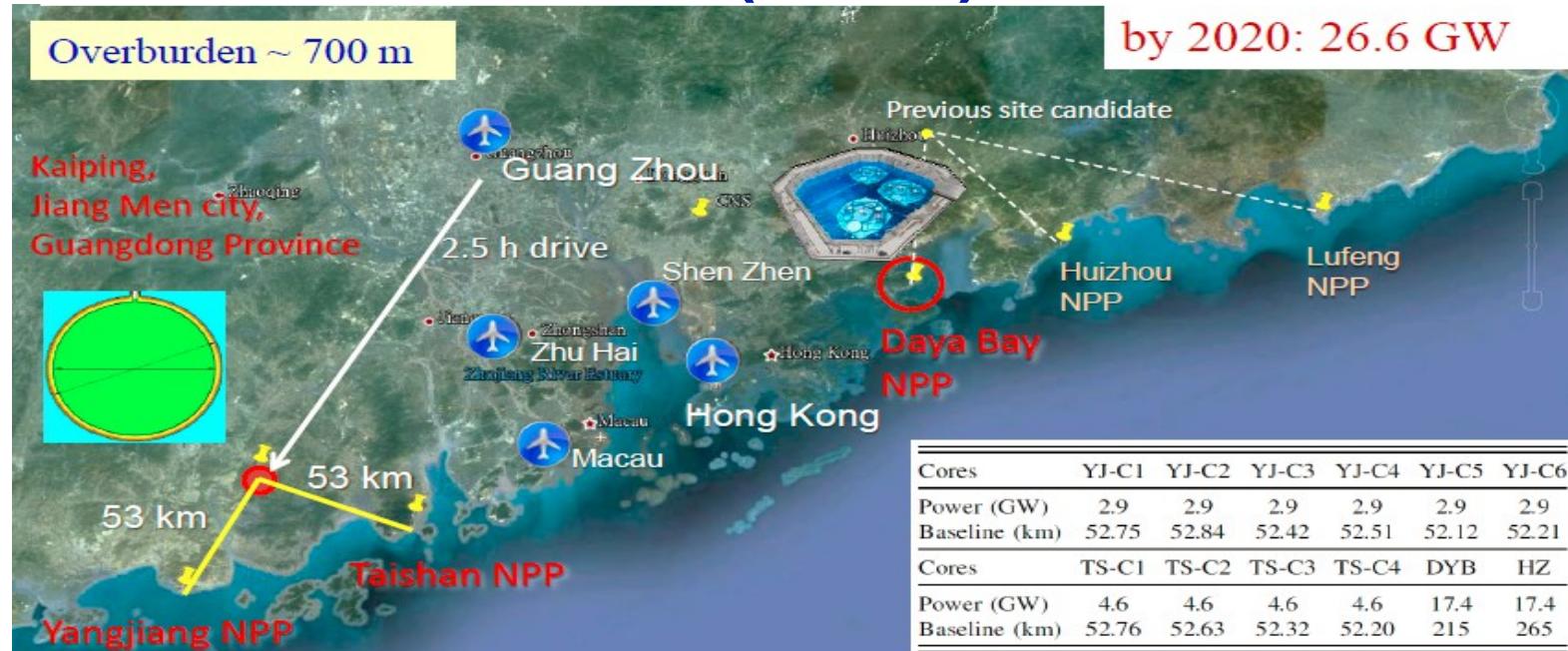


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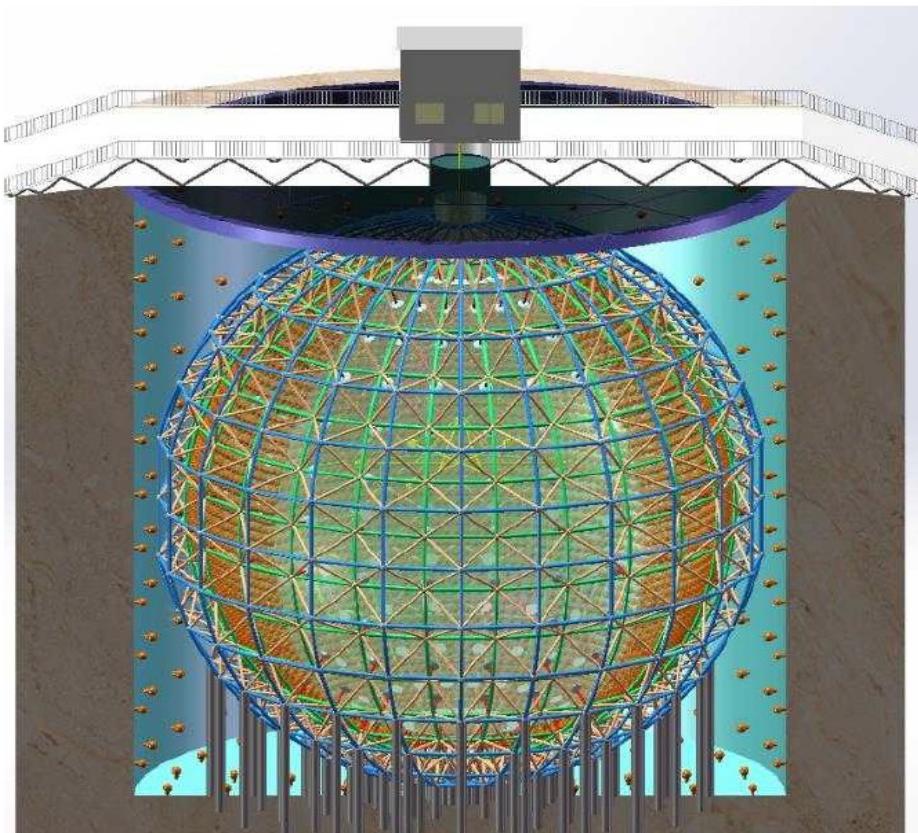


Jiangmen Underground Neutrino Observatory (JUNO)



Rich Physics Programs

- ◆ JUNO has been approved in Feb. 2013. ~ 300 M\$.
- ◆ A multiple-purpose neutrino experiment



- ◆ 20 kton LS detector
- ◆ 3% energy resolution
- ◆ 700 m underground
- ◆ Rich physics possibilities
- ◆ Reactor neutrino
 - ⇒ for Mass hierarchy and precision measurement of oscillation parameters
 - ⇒ Supernovae neutrino
 - ⇒ Geoneutrino
 - ⇒ Solar neutrino
 - ⇒ Atmospheric neutrino
 - ⇒ Exotic searches including proton decay, dark matter

Neutrino Physics with JUNO, J. Phys. G 43, 030401 (2016)

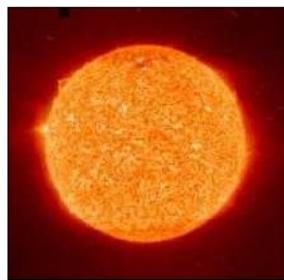


JUNO Event Rates

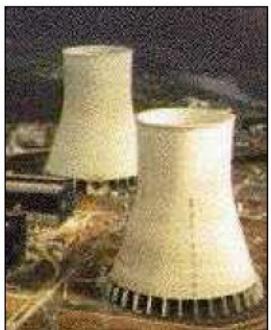
after selection



Supernova ν
5-7k in 10s for 10kpc

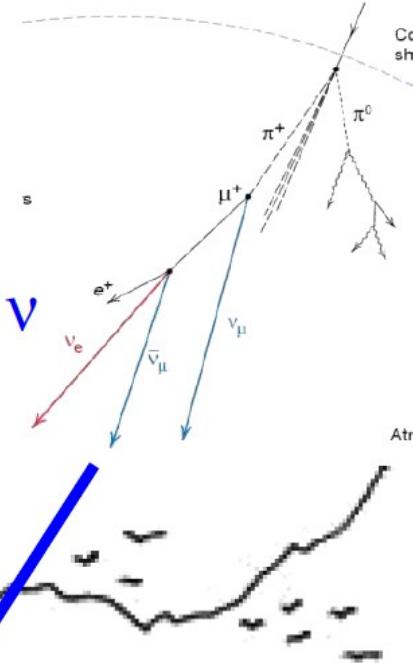


Solar ν
(10s-1000s)/day



36 GW, 53 km
reactor ν , 60/day
Bkg: 3.8/day

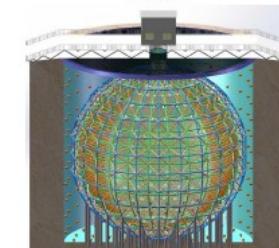
Atmospheric ν
several/day



Cosmic muons
 $\sim 250k/\text{day}$

0.003 Hz/m²
215 GeV
10% multiple-muon

Geo-neutrinos
1.1/day





JUNO Sensitivity on MH



PRD 88, 013008
(2013)

Statistics only

Relative Meas.

w/ absolute Δm^2

Realistic case

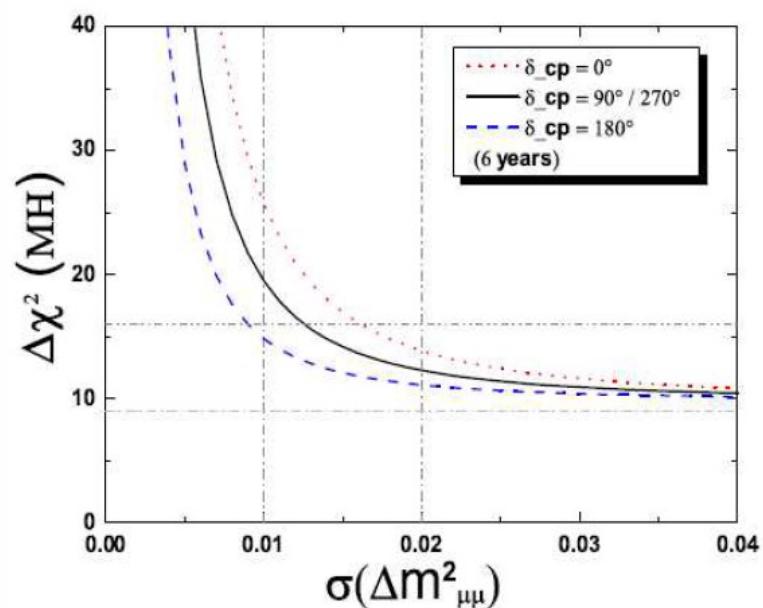
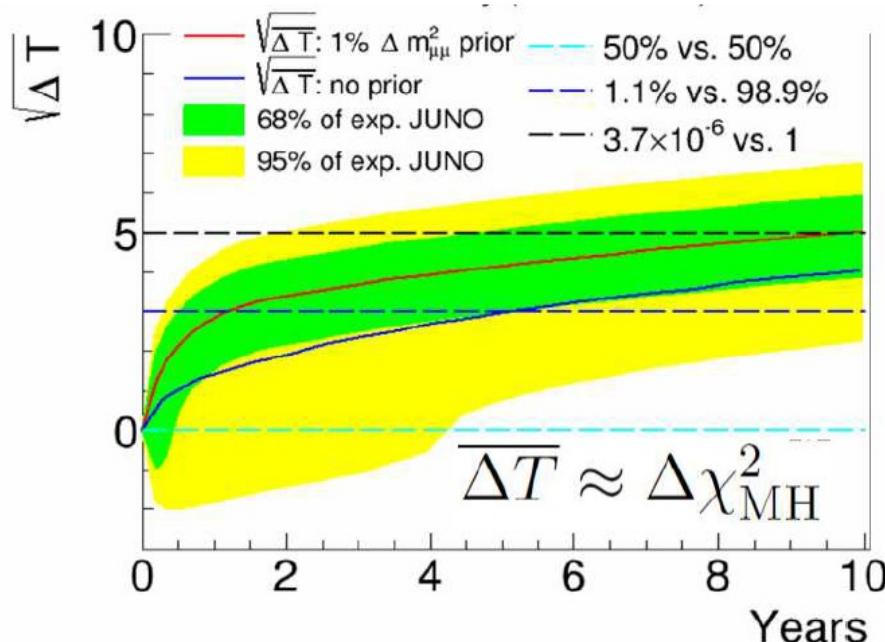
4σ

5σ

3σ

4σ

**JUNO MH
sensitivity
with 6
years' data:**



	Ideal	Core distr.	DYB & HZ	Shape	B/S (stat.)	B/S (shape)	$ \Delta m_{\mu\mu}^2 $
Size	52.5 km	Real	Real	1%	6.3%	0.4%	1%
$\Delta \chi^2_{\text{MH}}$	16	- 3	-1.7	- 1	- 0.6	- 0.1	+ (4-12)



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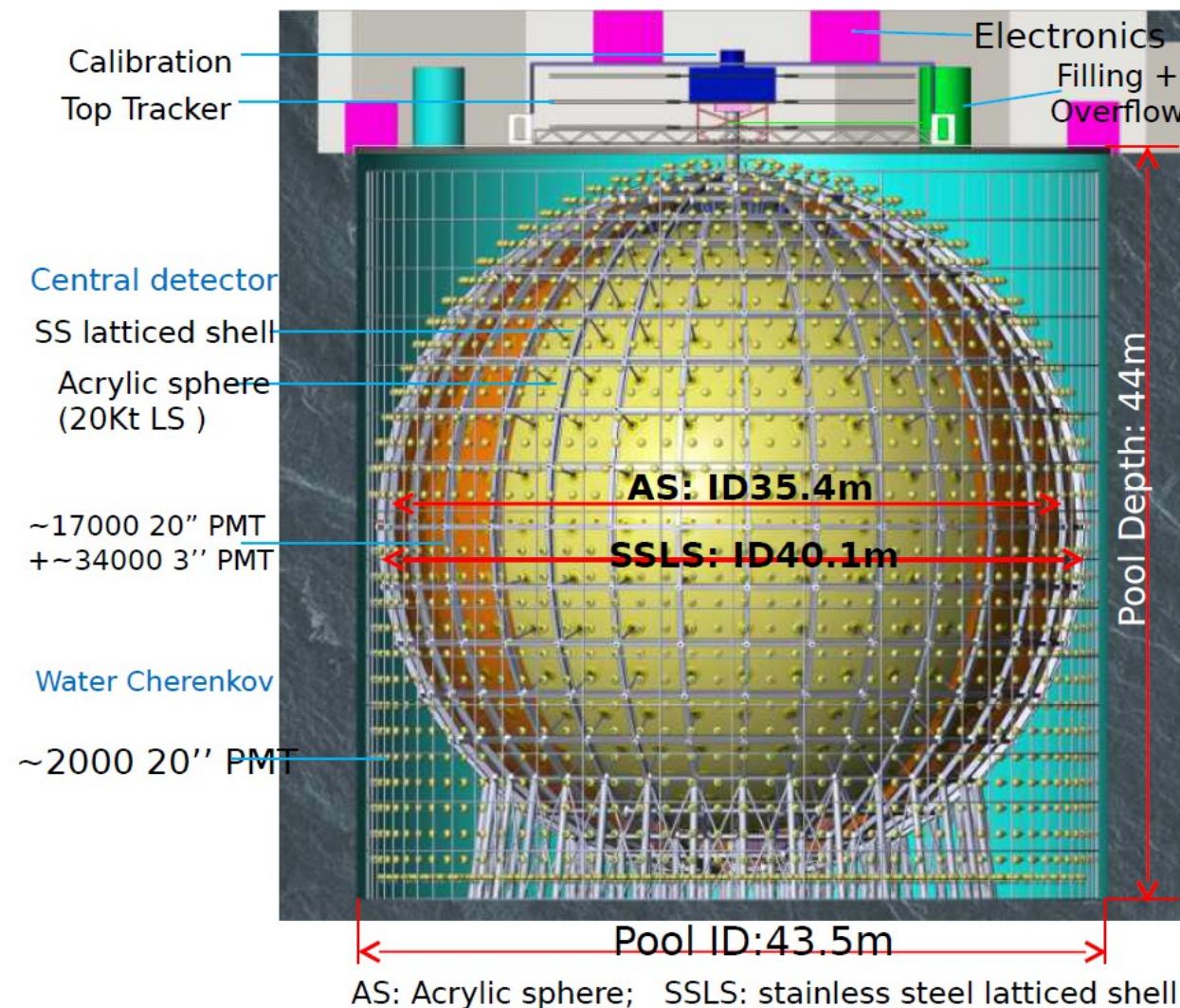


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JUNO detector



**Acrylic Option:
sphere + stainless
steel latticed shell**

A multi-purpose neutrino observatory

- **Largest LS detector:** ~20kt@ Φ 35.4m
- **Energy resolution:** <3%@1MeV under \geq 1200p.e./MeV with high QE 20" PMTs coverage
- **1GHz sampling waveform readout electronics for better energy understanding and more possibilities**
- **Double calorimetry both for low/high energy and cross checking**

Most of the key designs determined



CD Structure design

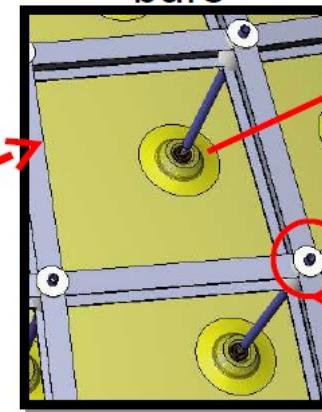
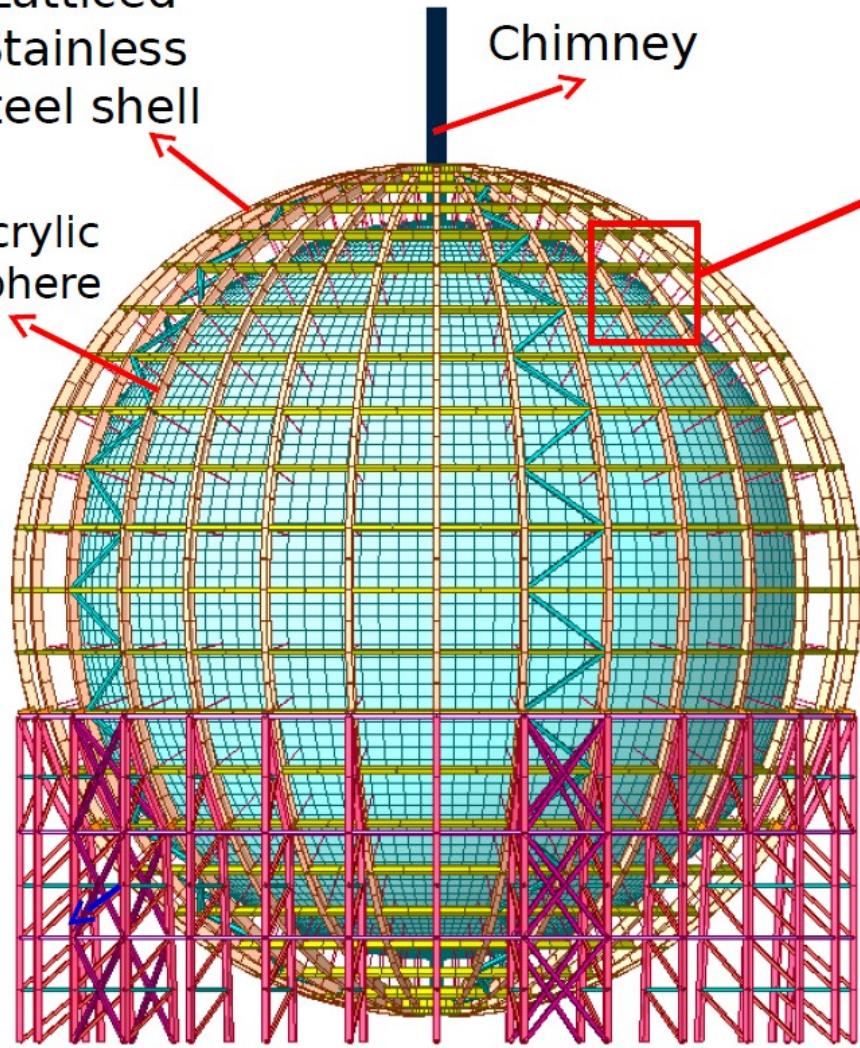


Latticed
Stainless
steel shell

Acrylic
sphere

Chimney

Connection
bars

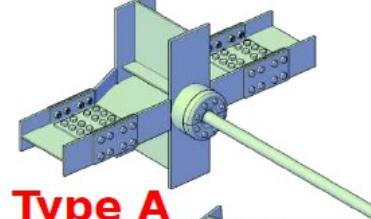


pillars

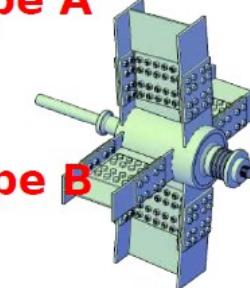
Acrylic
nodes

Type A

Type B



Type A



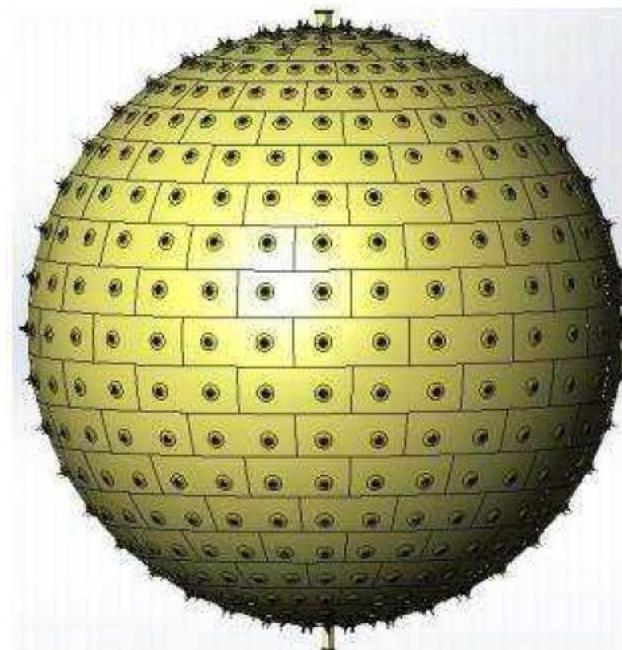
Type B

Shell node

- Shaping 120mm acrylic sphere plate
- Build the biggest acrylic sphere
- Tolerance both for weight and buoyancy under different conditions
- Low background control
- Transparency and less shadowing
- Installation and filling

Acrylic sphere supported
by stainless steel shell

Acrylic sphere R&D



Acrylic divided into 200+panels



The problems of shrinkage and shape variation were resolved.
Three companies had good practices.

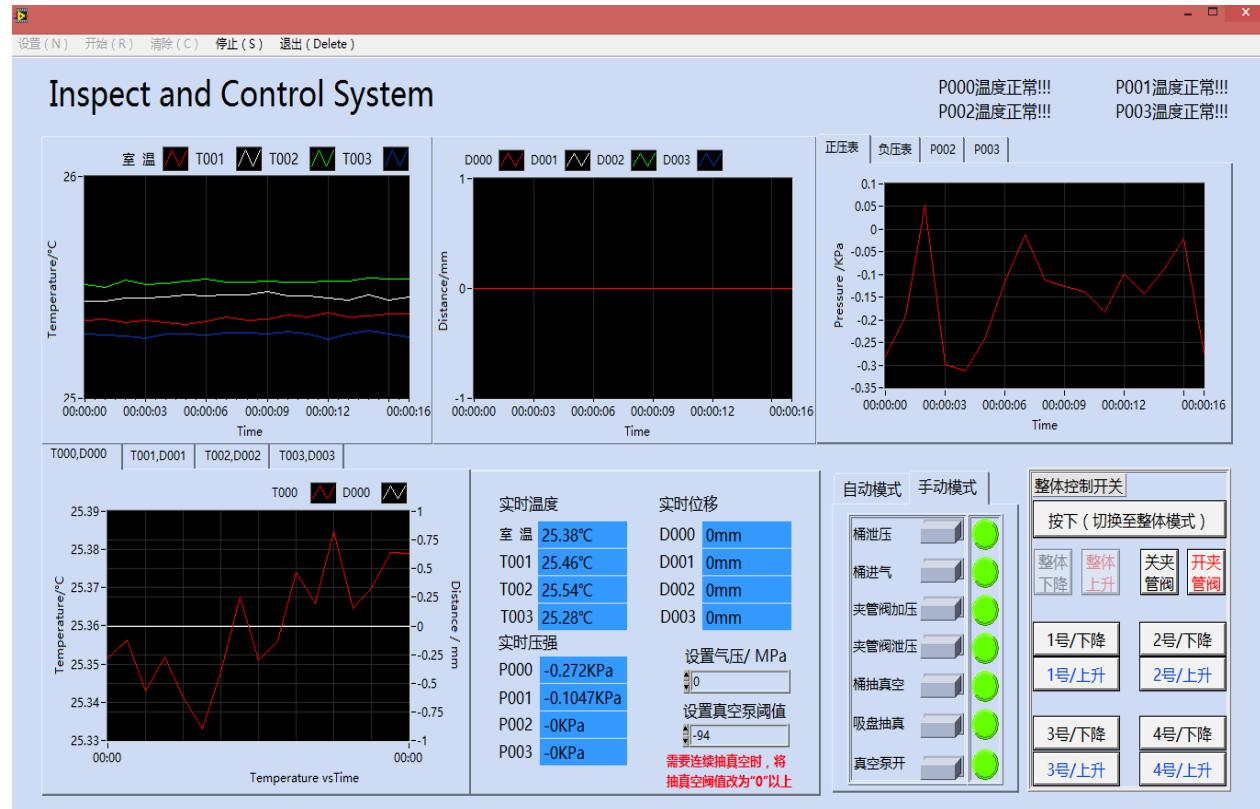
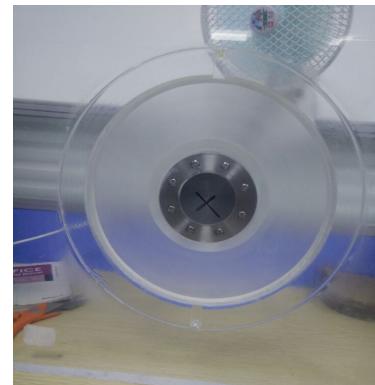
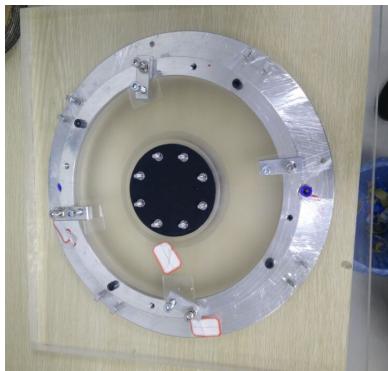


Forming panel size: 3m x 8m x120mm





A bonding machine for CD construction by SYSU



- Air-tight tests to control radio-purity
- Automatic glue injection with ultra-pure Nitrogen
- On-line monitor to keep the consistency
- Multiple bonding in parallel to increase efficiency
- Warning system to avoid dangerous polymerization



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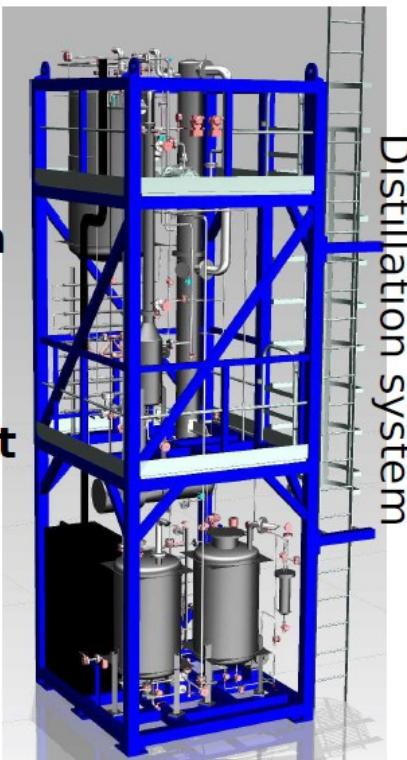
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LS Pilot plant

- ◆ Purify 20 ton LAB to test the overall design of purification system at Daya Bay. Plan to replace the target LS in one detector.
- ◆ Quantify the effectivities of subsystems
 - ⇒ Optical : >20m A.L @430nm
 - ⇒ Radio-purity: $< 10^{-15}$ g/g (U, Th)
- ◆ Determine the choice of sub-systems
 - ⇒ Al_2O_3 , distillation, gas striping, water extraction

**Distillation
and steam
stripping**

**Installed at
Daya Bay**



**Al_2O_3 column pilot plant
installed in Daya Bay LS hall**





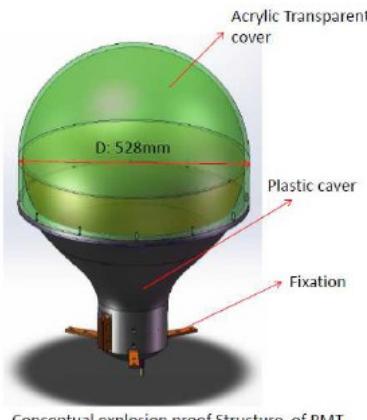
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PMT coverage

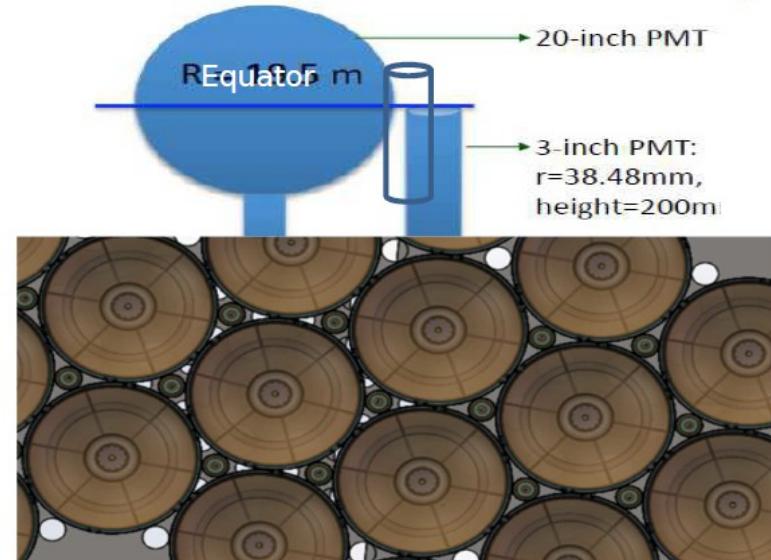
Single PMT in layers



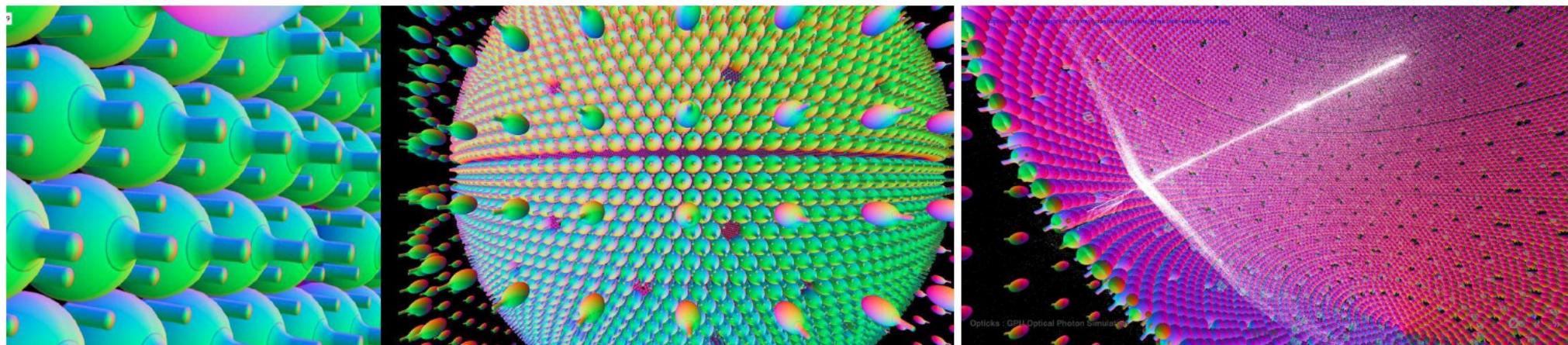
Conceptual explosion proof Structure of PMT
20" PMTs with structure
Installation with few mm gap

PMT No. : ~17000
PMT coverage: ~75%

Double calorimetry

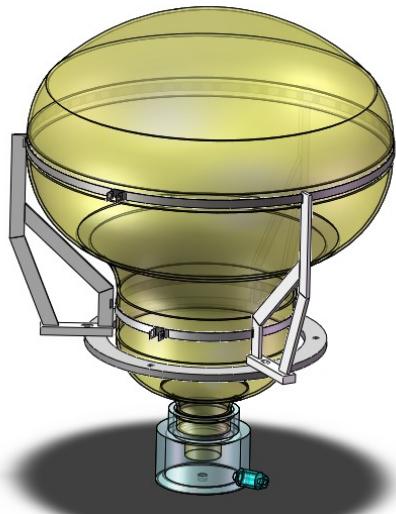


Targeted the largest light level ever detected in LSD ~1200 pe/MeV
(Daya Bay 160 pe/MeV -Borexino 500 pe/MeV -KamLAND 250 pe/MeV)





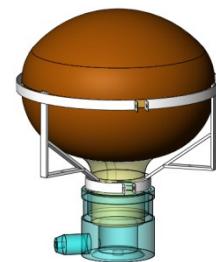
Development of PMTs for JUNO



H



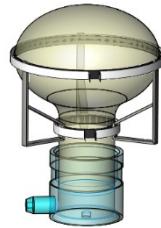
M



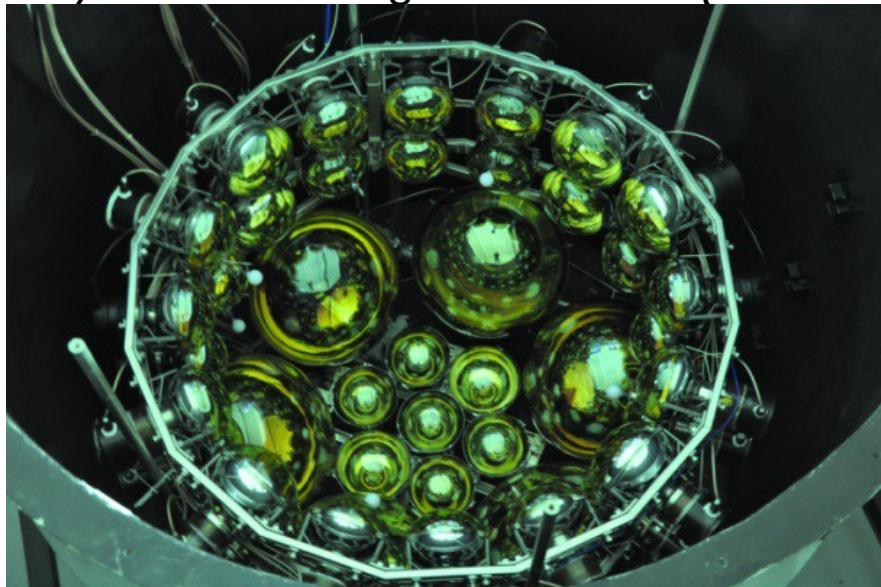
H
7



H
8



M
8

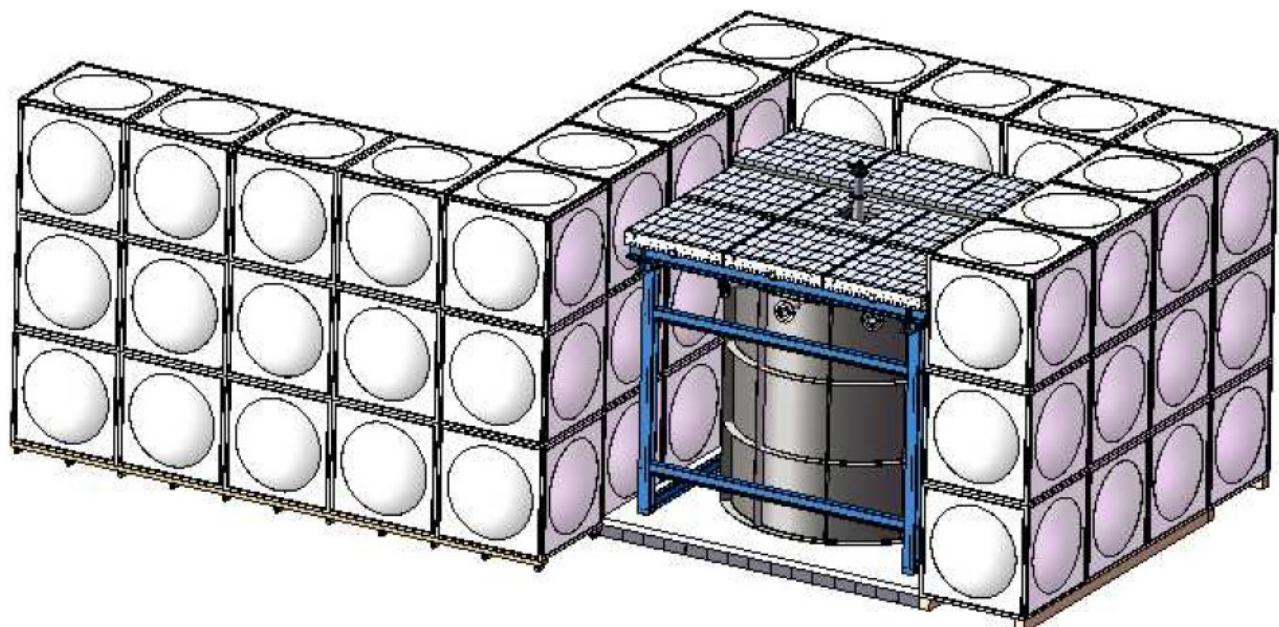
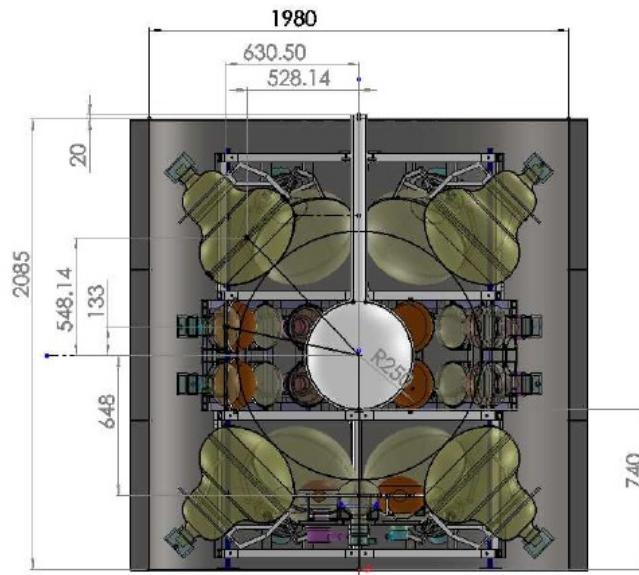




JUNO Detector Prototype



- Goal: Study/Comparison of PMTs' performances in a real LS detector
 - MCP-PMT 8"+20", Hainan Zhanchuang (HZC) 9", Hamamatsu 8"+20"
 - Bench test to cross check company parameters, Bench test to prepare for PMT mass testing, JUNO Prototype detector for more experience on:
 - New PMT testing
 - Large PMT mounting
 - Large PMT installation
 - Water proof PMT potting
 - PMT performance in LS detector
 - Calibration testing





PuC neutron source loaded in JUNO prototype

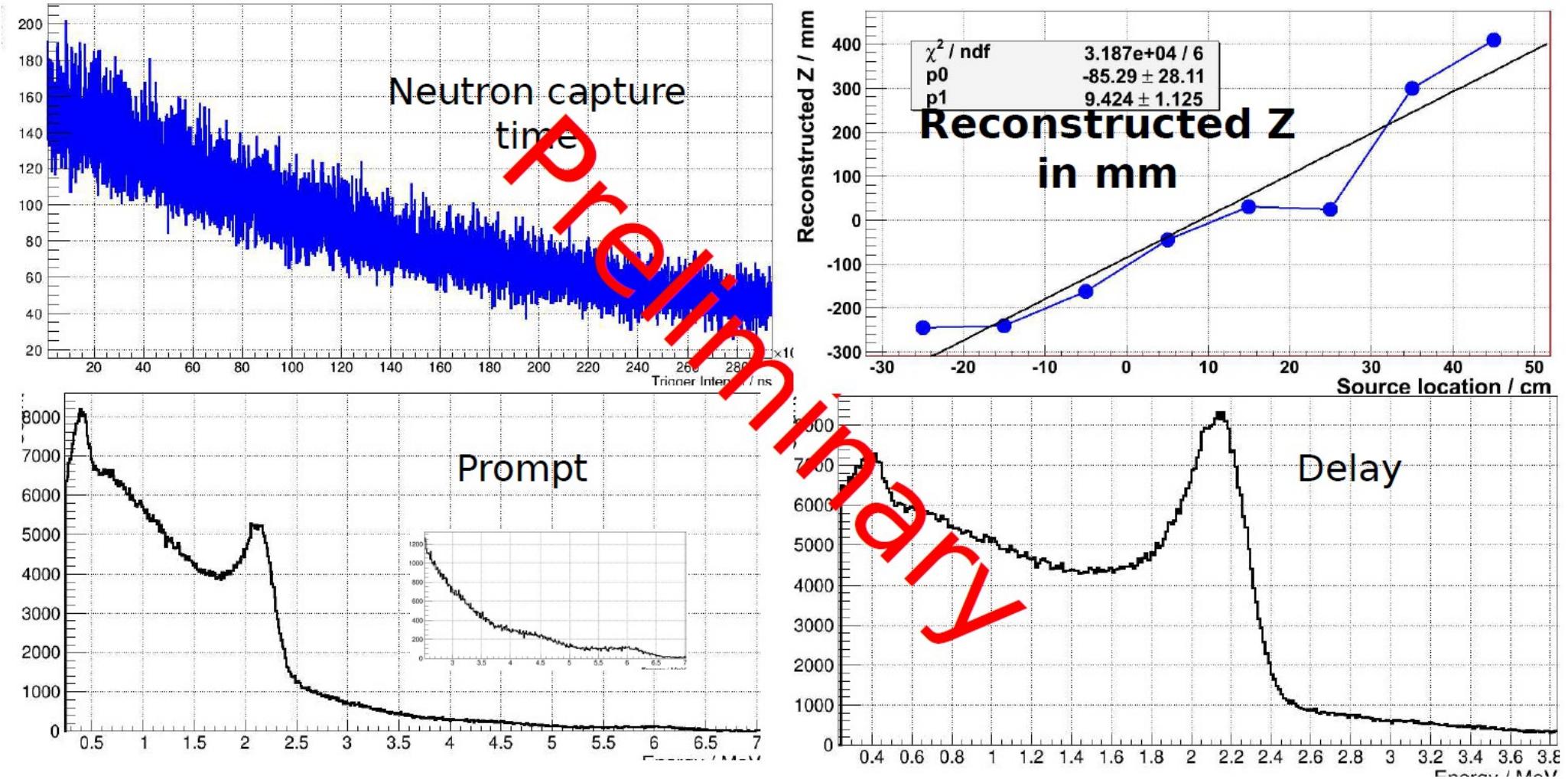




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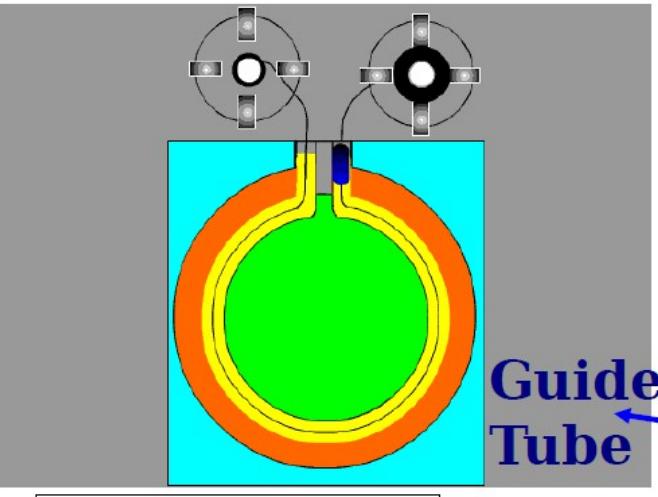
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Calibration System

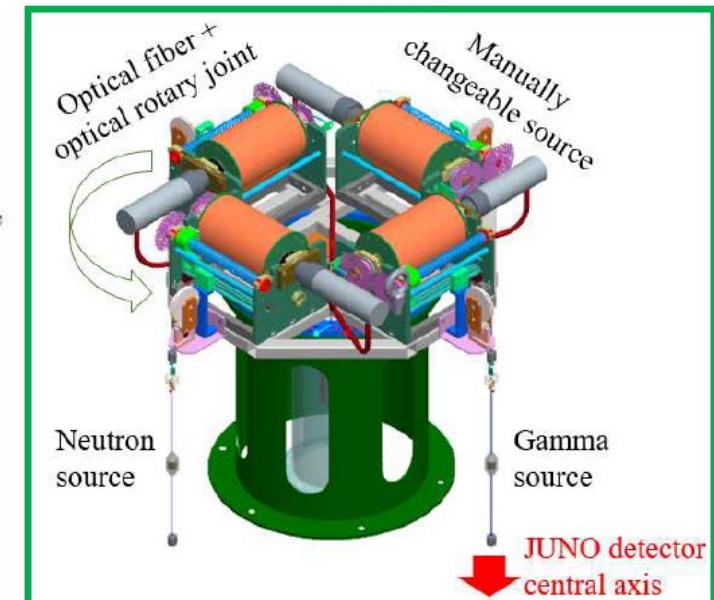
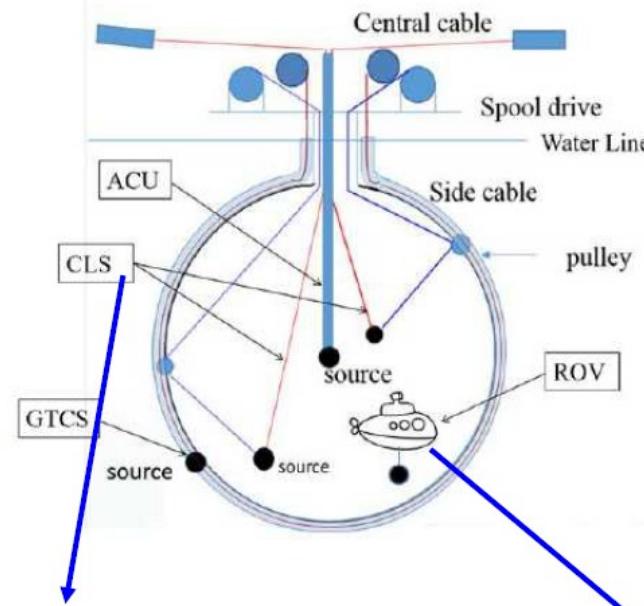


ACU



Guide
Tube

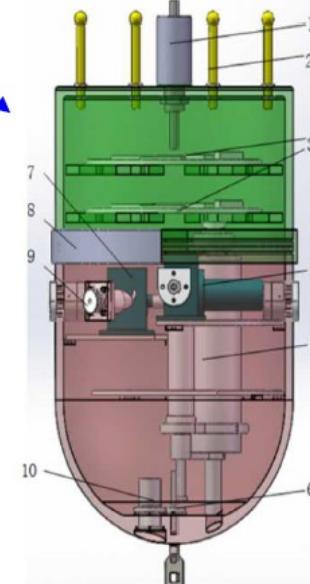
Boundary scan



Scan of central axis (1D)



Scan of vertical plane (2D)



ROV:3D scan



Schedule

ITEMs	Start
CD & water poll equipment installation	2018
PMT potting and testing	2018
CD & VETO PMT installation	2019
AD & VETO water filling	2019
LS filling	2019
Run	2020



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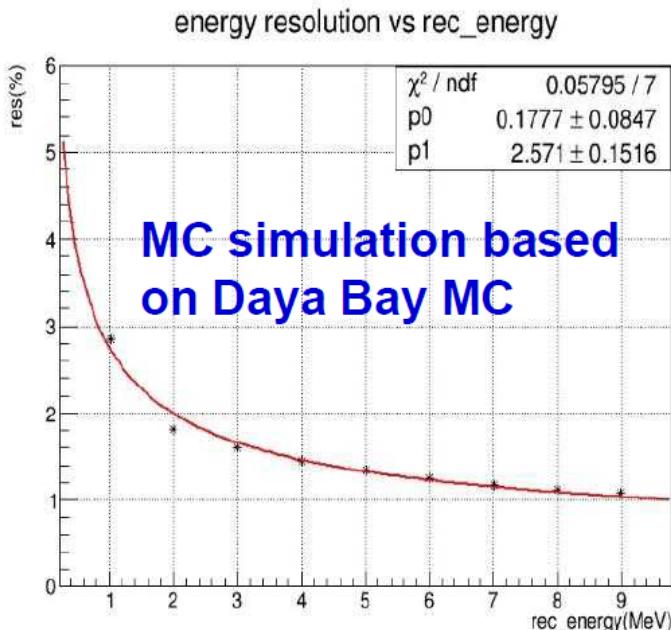
Summary

- **JUNO is a fully funded project and progresses in the fast track.**
- **JUNO central detector designed, matched with simulation for response, background, reconstruction, etc.**
- **PMTs selected and R&D of PMT instrumentation going on**
- **LS Pilot plant is ready to be scaled up for JUNO.**
- **Calibration system just passed the preliminary design review.**
- **Prototype testing for PMT, potting, LS, electronics, analysis etc. going well.**
- **Welcome to joining our efforts in JUNO.**

Thank you for your attention.



Challenges



- 77% photocathode coverage
- PMT peak QE: 35%
- Attenuation length of 20 m
- abs. 60 m + Rayl. scatt. 30m

- ◆ Two Challenges:
 - ⇒ How good is the energy resolution
 - ⇒ How well we know the reactor spectrum
- ◆ Model prediction (2-10%) + energy nonlinearity (1-3%) from LS and electronics/readout
- ◆ Two approaches to mitigate the spectrum uncertainties
 - ⇒ Direct measurement of the spectrum to 1% by SBL reactor exp.
 - ⇒ Constraint from Daya Bay measurements, independent of models, similar LS and similar electronics $\sim 1\%$

	KamLAND	BOREXINO	JUNO
LS mass	1 kt	0.5 kt	20 kt
Energy Resolution	6%/ \sqrt{E}	5%/ \sqrt{E}	3%/ \sqrt{E}
Light yield	250 p.e./MeV	511 p.e./MeV	1200 p.e./MeV