Viable models for large Non-Standard neutrino Interactions (NSI)

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Effects of NSI on neutrinos

 Neutral current Non-Standard Interaction (NSI): propagation of neutrinos in matter

Charged current Non-Standard Interaction (NSI): production and detection

Effects of NSI on neutrinos

Neutral current Non-Standard Interaction (NSI): propagation of neutrinos in matter Focus of this talk

Charged current Non-Standard Interaction (NSI): production and detection

Non-standard neutral current interaction

 $\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha} \gamma^{\mu} L \nu_{\beta}) (\bar{f} \gamma_{\mu} P f)$ Chirality Projection Matter field matrix

Neutrino propagation:

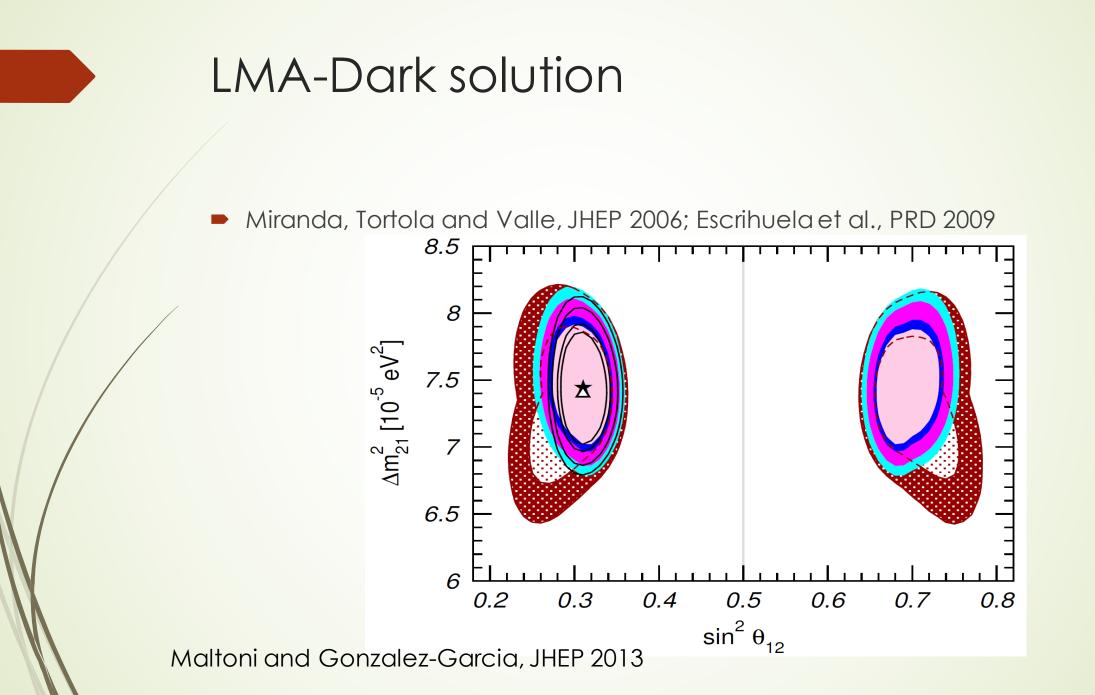
$$\epsilon^{f}_{\alpha\beta} \equiv \epsilon^{fL}_{\alpha\beta} + \epsilon^{fR}_{\alpha\beta}$$

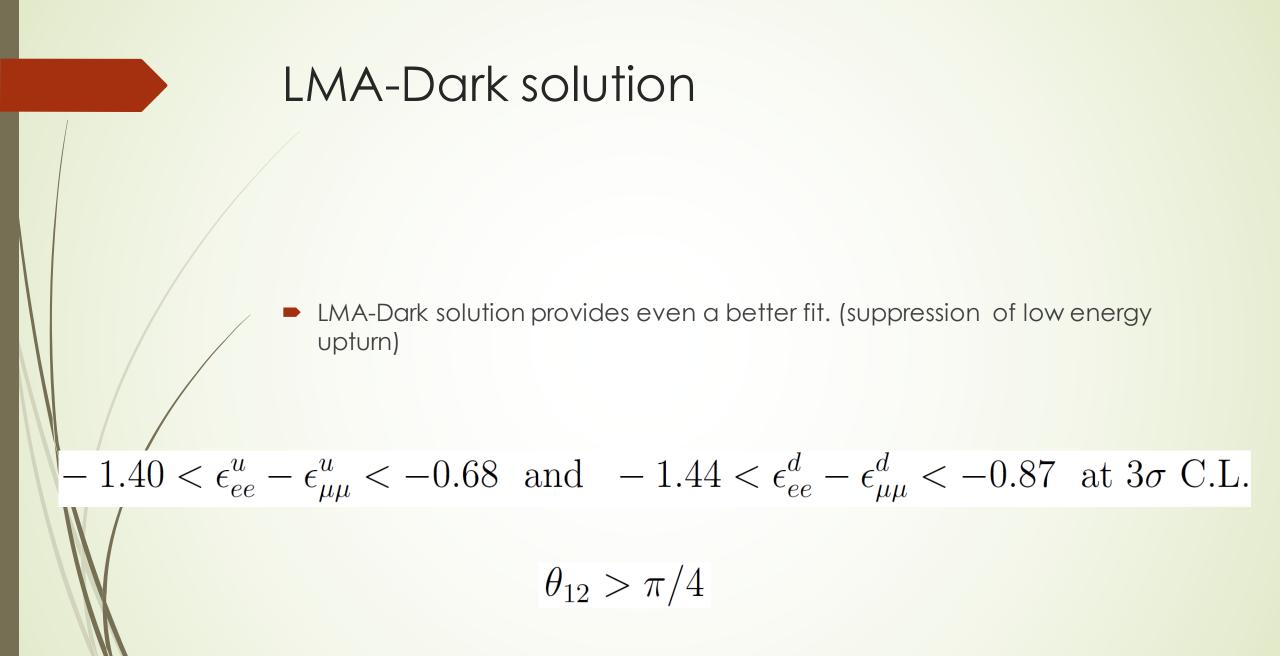
$$i\frac{d}{dx}\begin{pmatrix}\nu_e\\\nu_\mu\\\nu_\tau\end{pmatrix} = H^{\nu}\begin{pmatrix}\nu_e\\\nu_\mu\\\nu_\tau\end{pmatrix}$$

$$H^{\nu} = H_{\text{vac}} + H_{\text{mat}} \quad \text{and} \quad H^{\bar{\nu}} = (H_{\text{vac}} - H_{\text{mat}})^*$$
$$H_{mat} = \sqrt{2}G_F \left[N_e \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} + \sum_{f=e,u,d} N_f \begin{pmatrix} \epsilon_{ee}^f & \epsilon_{e\mu}^f & \epsilon_{e\tau}^f \\ \epsilon_{e\mu}^{f*} & \epsilon_{\mu\mu}^f & \epsilon_{\mu\tau}^f \\ \epsilon_{e\tau}^{f*} & \epsilon_{\mu\tau}^{f*} & \epsilon_{\tau\tau}^f \end{pmatrix} \right]$$

[$90\% \ \mathrm{CL}$	
	Param.	best-fit	LMA	$LMA \oplus LMA-D$
ſ	$\varepsilon^{u}_{ee} - \varepsilon^{u}_{\mu\mu}$	+0.298	[+0.00, +0.51]	$\oplus \left[-1.19, -0.81 ight]$
/	$\varepsilon^u_{\tau\tau} - \varepsilon^u_{\mu\mu}$	+0.001	[-0.01, +0.03]	[-0.03, +0.03]
	$arepsilon^u_{e\mu}$	-0.021	[-0.09, +0.04]	[-0.09, +0.10]
	$arepsilon_{e au}^{u}$	+0.021	[-0.14, +0.14]	[-0.15, +0.14]
	$arepsilon^u_{\mu au}$	-0.001	[-0.01, +0.01]	[-0.01, +0.01]
	ε^u_D	-0.140	[-0.24, -0.01]	$\oplus [+0.40, +0.58]$
/	ε^u_N	-0.030	[-0.14, +0.13]	[-0.15, +0.13]
[$\varepsilon^d_{ee} - \varepsilon^d_{\mu\mu}$	+0.310	[+0.02, +0.51]	$\oplus \left[-1.17, -1.03 ight]$
	$\varepsilon^d_{ au au} - \varepsilon^d_{\mu\mu}$	+0.001	[-0.01, +0.03]	[-0.01, +0.03]
	$arepsilon_{e\mu}^d$	-0.023	[-0.09, +0.04]	[-0.09, +0.08]
	$arepsilon_{e au}^d$	+0.023	[-0.13, +0.14]	[-0.13, +0.14]
	$\varepsilon^{d}_{\mu\tau}$	-0.001	[-0.01, +0.01]	[-0.01, +0.01]
	ε_D^d	-0.145	[-0.25, -0.02]	\oplus [+0.49, +0.57]
	ε^d_N	-0.036	[-0.14, +0.12]	[-0.14, +0.12]

Maltoni and Gonzalez-Garcia, JHEP 2013





Total flux measurement at SNO

Neutral current

Deuteron dissociation

$$\nu + D \rightarrow \nu + p + n$$

- Gamow-Teller transition
- Sensitive only to axial-vector interaction
- No effect from $\epsilon^f_{\alpha\beta} \equiv \epsilon^{fL}_{\alpha\beta} + \epsilon^{fR}_{\alpha\beta}$

Scattering experiments

$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha}\gamma^{\mu}L\nu_{\beta})(\bar{f}\gamma_{\mu}P\ f)$$

NuTeV and CHARM rule out a large part (but not all) of parameter space of LMA-Dark solution.

Davidson, Pena-Garay, Rius, SantaMaria, JHEP 2003

Scattering experiments

$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha} \gamma^{\mu} L \nu_{\beta}) (\bar{f} \gamma_{\mu} P f)$$

NuTeV and CHARM rule out a large part (but not all of) parameter space of LMA-Dark solution.

Davidson, Pena-Garay, Rius, SantaMaria, JHEP 2003

But not in the model that we shall present

Underlying theory ? $\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha} \gamma^{\mu} L \nu_{\beta}) (\bar{f} \gamma_{\mu} P f)$ $\epsilon^{f}_{\alpha\beta} \equiv \epsilon^{fL}_{\alpha\beta} + \epsilon^{fR}_{\alpha\beta}$ $-1.40 < \epsilon_{ee}^u - \epsilon_{\mu\mu}^u < -0.68$ and $-1.44 < \epsilon_{ee}^d - \epsilon_{\mu\mu}^d < -0.87$ at 3σ C.L.

$$\epsilon \sim \left(\frac{g_X^2}{m_X^2}\right) G_F^{-1}$$

Underlying theory

$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha}\gamma^{\mu}L\nu_{\beta}) (\bar{f}\gamma_{\mu}P f)$$

$$\epsilon^f_{\alpha\beta} \equiv \epsilon^{fL}_{\alpha\beta} + \epsilon^{fR}_{\alpha\beta}$$

Various model with heavy intermediate particle For a review see:

T. Ohlsson, "Status of non-standard neutrino interactions," Rept. Prog. Phys. 76 (2013) 044201 [arXiv:1209.2710 [hep-ph]].

Too small NSI

$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha}\gamma^{\mu}L\nu_{\beta})(\bar{f}\gamma_{\mu}P \ f)$$

$$\epsilon^{f}_{\alpha\beta} \equiv \epsilon^{fL}_{\alpha\beta} + \epsilon^{fR}_{\alpha\beta}$$

$$-1.40 < \epsilon_{ee}^u - \epsilon_{\mu\mu}^u < -0.68$$
 and $-1.44 < \epsilon_{ee}^d - \epsilon_{\mu\mu}^d < -0.87$ at 3σ C.L.

$$\epsilon \sim \left(\frac{g_X^2}{m_X^2}\right) G_F^{-1}$$

 $\epsilon \ll 1$ $m_X \gg 100 {
m ~GeV}$

ATLAS $\sqrt{s} = 8$ TeV bound

$$\epsilon \sim \left(\frac{g_X^2}{m_X^2}\right) G_F^{-1}$$

Aad et al., PRD90 (2014) 52005

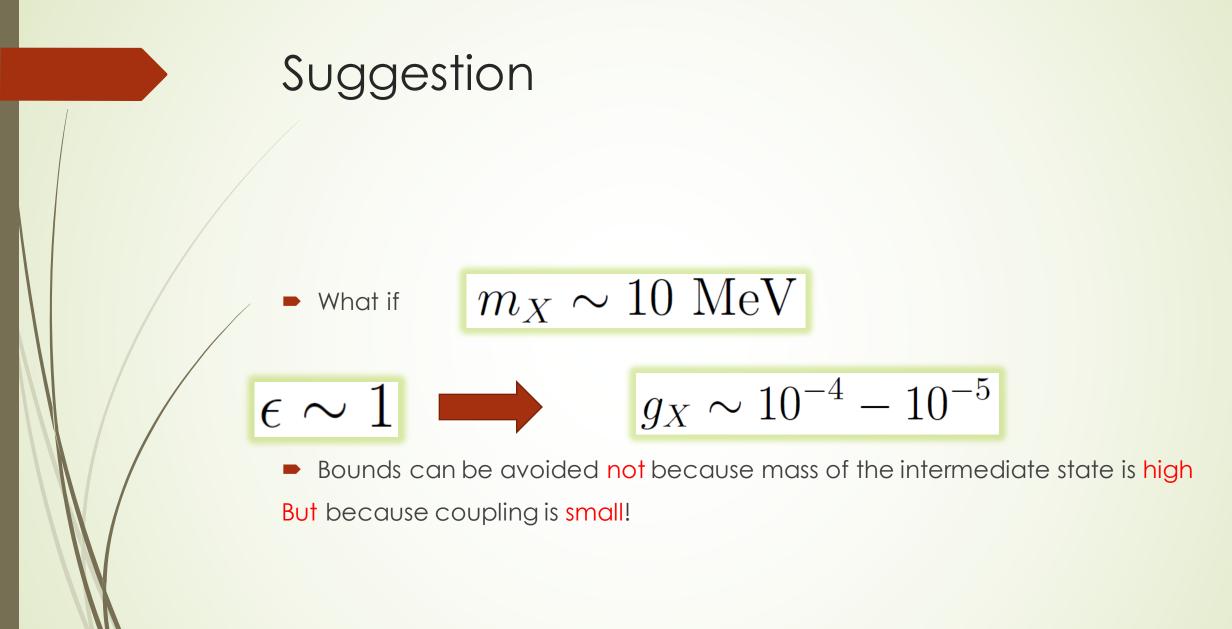
$$m_{Z'_{SM}} > 2900 \text{ GeV}$$

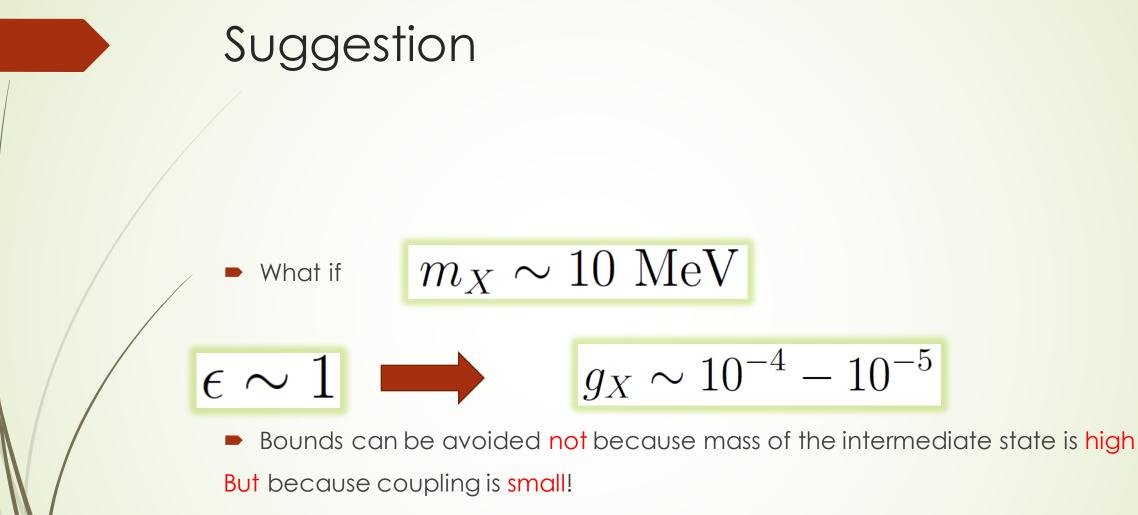
 $\epsilon < 10^{-3}$



- What if $m_X \sim 10 \; {
m MeV}$

YF, A model for large non-standard interactions leading to LMA-Dark solution, Phys. Lett. B748 (2015) 311-315; YF and Shoemaker, Lepton flavor violating non-standard interactions, JHEP 1607 (2016) 033. YF and Heeck, Neutrinophilic non-standard interactions, 1607.07616





$$\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha}\gamma^{\mu}L\nu_{\beta}) (\bar{f}\gamma_{\mu}P \ f)$$

For forward scattering we can still use the effective Lagrangian.

Model for LMA-Dark with $\epsilon_{\mu\mu} = \epsilon_{\tau\tau} \sim 1$

$$-g'\left(Y_L\sum_{\alpha\in\{\mu,\tau\}}\bar{L}_{\alpha}\gamma^{\mu}L_{\alpha}+Y_{Q_1}\bar{Q}_1\gamma^{\mu}Q_1+Y_{u_1}\bar{u}_R\gamma^{\mu}u_R+Y_{d_1}\bar{d}_R\gamma^{\mu}d_R\right)Z'_{\mu}\in\mathcal{L}$$

$$\mathcal{L}_{NSI} = -\frac{g^{\prime 2}Y_L}{m_{Z^{\prime}}^2} \left(\sum_{\alpha \in \{\mu,\tau\}} \bar{L}_{\alpha} \gamma^{\mu} L_{\alpha}\right) \left(Y_{Q_1} \bar{Q}_1 \gamma_{\mu} Q_1 + Y_{u_1} \bar{u}_R \gamma_{\mu} u_R + Y_{d_1} \bar{d}_R \gamma_{\mu} d_R\right)$$

YF PLB 748 (2015) 311

$$-g'\left(Y_L\sum_{\alpha\in\{\mu,\tau\}}\bar{L}_{\alpha}\gamma^{\mu}L_{\alpha}+Y_{Q_1}\bar{Q}_1\gamma^{\mu}Q_1+Y_{u_1}\bar{u}_R\gamma^{\mu}u_R+Y_{d_1}\bar{d}_R\gamma^{\mu}d_R\right)Z'_{\mu}\in\mathcal{L}$$

NO coupling to the electron

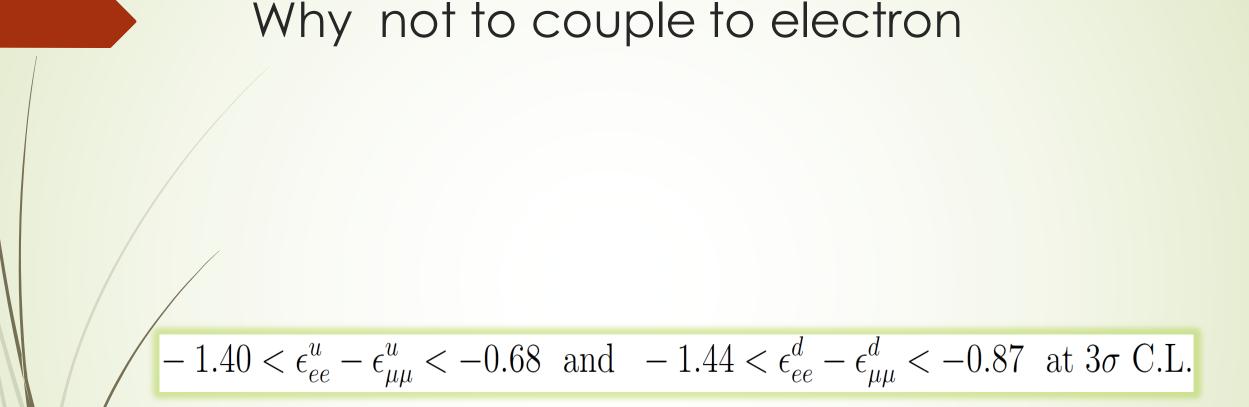
$$\mathcal{L}_{NSI} = -\frac{g^{\prime 2}Y_L}{m_{Z^{\prime}}^2} \left(\sum_{\alpha \in \{\mu,\tau\}} \bar{L}_{\alpha} \gamma^{\mu} L_{\alpha}\right) \left(Y_{Q_1} \bar{Q}_1 \gamma_{\mu} Q_1 + Y_{u_1} \bar{u}_R \gamma_{\mu} u_R + Y_{d_1} \bar{d}_R \gamma_{\mu} d_R\right)$$

$$\epsilon_{\tau\tau}^{u} = \epsilon_{\mu\mu}^{u} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{u_{1}})}{2\sqrt{2}G_{F}}$$

$$\epsilon_{\tau\tau}^{d} = \epsilon_{\mu\mu}^{d} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{d_{1}})}{2\sqrt{2}G_{F}}$$

$$\epsilon_{ee}^{u} = \epsilon_{ee}^{d} = 0$$

$$\epsilon^u_{\alpha\beta} = \epsilon^d_{\alpha\beta} = 0 \quad \alpha \neq \beta$$





LMA-Dark solution

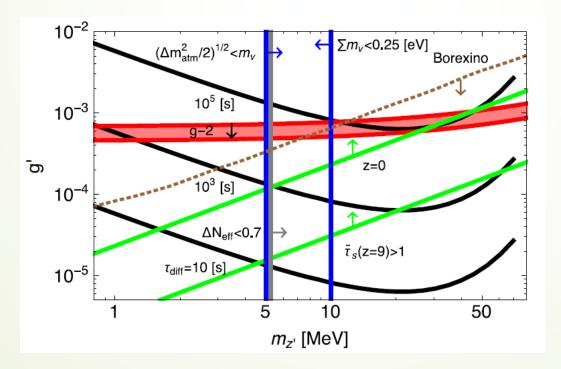
 $\epsilon^{u,d}_{\mu\mu} \sim 1$

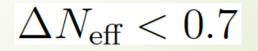
$$\epsilon_{\tau\tau}^{u} = \epsilon_{\mu\mu}^{u} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{u_{1}})}{2\sqrt{2}G_{F}}$$
$$\epsilon_{\tau\tau}^{d} = \epsilon_{\mu\mu}^{d} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{d_{1}})}{2\sqrt{2}G_{F}}$$
$$g^{\prime} \sim 7 \times 10^{-5} \frac{m_{Z^{\prime}}}{10 \text{ MeV}}$$

Big Bang Nucleosynthesis

Kamada and Yu, PRD 92 (2015)

$m_{Z'} > 5 \text{ MeV}$





Neutrino scattering experiments

 $q^2 \gg m_{Z'}^2$ $\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha} \gamma^{\mu} L \nu_{\beta}) (\bar{f} \gamma_{\mu} P f)$

Suppression factor $m_{Z^\prime}^2/(q^2-m_{Z^\prime}^2)$

Neutrino scattering experiments

 $q^2 \gg m_{Z'}^2$

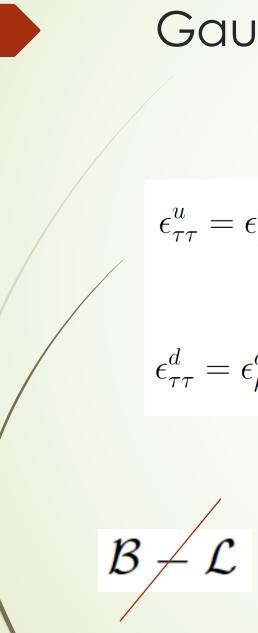
 $\mathcal{L}_{NSI} = -2\sqrt{2}G_F \epsilon^{fP}_{\alpha\beta} (\bar{\nu}_{\alpha}\gamma^{\mu}L\nu_{\beta}) (\bar{f}\gamma_{\mu}P f)$ $10 \text{ MeV} \stackrel{<}{\sim} m_{Z'} \ll 1 \text{ GeV}$

Relaxing bounds from scattering experiments, NuTeV and CHARM



$$\epsilon_{\tau\tau}^{u} = \epsilon_{\mu\mu}^{u} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{u_{1}})}{2\sqrt{2}G_{F}}$$
$$\epsilon_{\tau\tau}^{d} = \epsilon_{\mu\mu}^{d} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{d_{1}})}{2\sqrt{2}G_{F}}$$

$$\epsilon^{u,d}_{\mu\mu} \simeq \epsilon^{u,d}_{\tau\tau} \sim 1$$



$\mathcal{L}_{\mu} - \mathcal{L}_{ au}$

$$\epsilon^{d}_{\tau\tau} = \epsilon^{d}_{\mu\mu} = \frac{g^{\prime 2}}{m_{Z^{\prime}}^{2}} \frac{Y_{L}(Y_{Q_{1}} + Y_{d_{1}})}{2\sqrt{2}G_{F}}$$

$$\epsilon^{u,d}_{\mu\mu} \simeq \epsilon^{u,d}_{\tau\tau} \sim 1$$

$$\epsilon^{u}_{\tau\tau} = \epsilon^{u}_{\mu\mu} = \frac{g^{\prime 2}}{m^{2}_{Z^{\prime}}} \frac{Y_{L}(Y_{Q_{1}} + Y_{u_{1}})}{2\sqrt{2}G_{F}}$$

$$\epsilon^{u}_{\tau\tau} = \epsilon^{u}_{\mu\mu} = \frac{g^{\prime 2}}{m^{2}_{Z^{\prime}}} \frac{Y_{L}(Y_{Q_{1}} + Y_{u_{1}})}{2\sqrt{2}G_{F}}$$

$$\mathcal{L}_{\mu} + \mathcal{L}_{\tau} + \mathcal{B}_{1} - a\mathcal{B}_{2} - (3 - a)\mathcal{B}_{3}$$
$$Y_{Q_{1}} = Y_{u_{1}} = Y_{d_{1}} = 1/3 , \quad Y_{Q_{2}} = Y_{u_{2}} = Y_{d_{2}} = -a/3$$
$$Y_{Q_{3}} = Y_{u_{3}} = Y_{d_{3}} = -1 + a/3$$

U(1)'

U(1)'

$$\mathcal{L}_{\mu} + \mathcal{L}_{\tau} + \mathcal{B}_{1} - a\mathcal{B}_{2} - (3 - a)\mathcal{B}_{3}$$
$$Y_{e} = Y_{L_{e}} = 0 \text{ and } Y_{\mu} = Y_{\tau} = Y_{L_{\mu}} = Y_{L_{\tau}} = 1$$

MODEL FOR LMA-DARK

Desired sign and magnitude

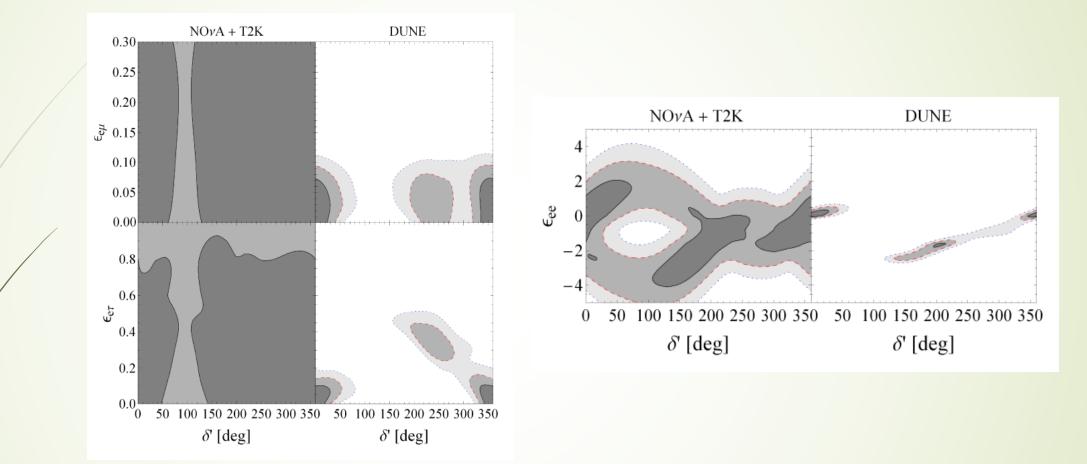
$$-1.40 < \epsilon_{ee}^u - \epsilon_{\mu\mu}^u < -0.68$$
 and $-1.44 < \epsilon_{ee}^d - \epsilon_{\mu\mu}^d < -0.87$ at 3σ C.L.

YF, A model for large non-standard interactions leading to LMA-Dark solution, Phys. Lett. B748 (2015) 311-315;

Effects of NSI in long baseline experiments

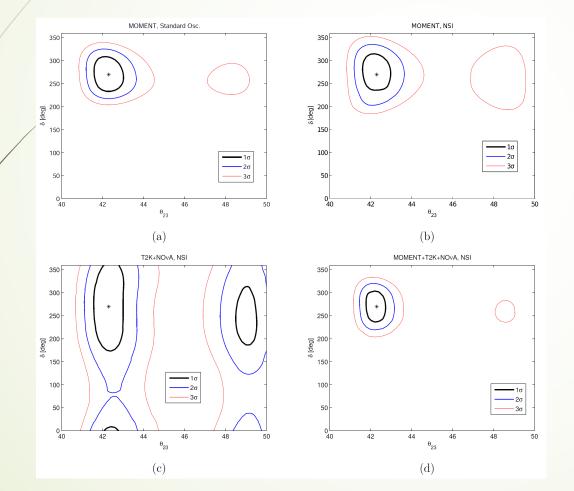
Renewed interest in NSI

NSI can fake CP-violation and lead to wrong determination of θ_{23} octant Masud and Mehta, PRD 94(2016); Forero and Huber, PLB 117 (2016); Liao, Marfatia and Whistnant PRD 93 (2016); Agarwalla, Chatterjee and Palazzo, 1607.01745,



Liao Marfatia Whisnant, PRD93 (2016)

MuOn decay MEdium baseline NeuTrino beam (MOMENT)



Bakhti and YF, JHEP 1607 (2016) 109

Lepton flavor violating NSI

$$\tilde{L} \equiv \begin{pmatrix} \tilde{L}_{\alpha} \\ \tilde{L}_{\beta} \end{pmatrix} \xrightarrow{U(1)'} e^{i\zeta g'\sigma_1 \alpha} \tilde{L},$$

, Charged lepton mass basis

$$L \equiv \begin{pmatrix} L_{\alpha} \\ L_{\beta} \end{pmatrix} = \begin{pmatrix} \cos \theta_L & -\sin \theta_L \\ \sin \theta_L & \cos \theta_L \end{pmatrix} \hat{I}$$

Lepton flavor violating NSI

$$\tilde{L} \equiv \begin{pmatrix} \tilde{L}_{\alpha} \\ \tilde{L}_{\beta} \end{pmatrix} \xrightarrow{U(1)'} e^{i\zeta g' \sigma_1 \alpha} \tilde{L},$$

$$\frac{\tilde{L}_{\alpha} + \tilde{L}_{\beta}}{\sqrt{2}} \to e^{i\alpha\zeta g'} \frac{\tilde{L}_{\alpha} + \tilde{L}_{\beta}}{\sqrt{2}} \quad \text{and} \quad \frac{\tilde{L}_{\alpha} - \tilde{L}_{\beta}}{\sqrt{2}} \to e^{-i\alpha\zeta g'} \frac{\tilde{L}_{\alpha} - \tilde{L}_{\beta}}{\sqrt{2}}$$

Lepton flavor violating NSI

$$\tilde{L} \equiv \begin{pmatrix} \tilde{L}_{\alpha} \\ \tilde{L}_{\beta} \end{pmatrix} \xrightarrow{U(1)'} e^{i\zeta g'\sigma_1 \alpha} \tilde{L},$$

$$\tilde{R} \equiv \begin{pmatrix} \tilde{l}_{R\alpha}^{-} \\ \tilde{l}_{R\beta}^{-} \end{pmatrix} \xrightarrow{U(1)'} e^{i\zeta g'\sigma_1 \alpha} \tilde{R}.$$

Yukawa coupling with SM Higgs

 $b_0 \tilde{R}^{\dagger} H^{\dagger} \tilde{L} + b_1 \tilde{R}^{\dagger} \sigma_1 H^{\dagger} \tilde{L},$

Charges of baryons under new U(1)'

$$\eta_1 B_1 + \eta_2 B_2 + \eta_3 B_3$$

Flavor structure of NSI

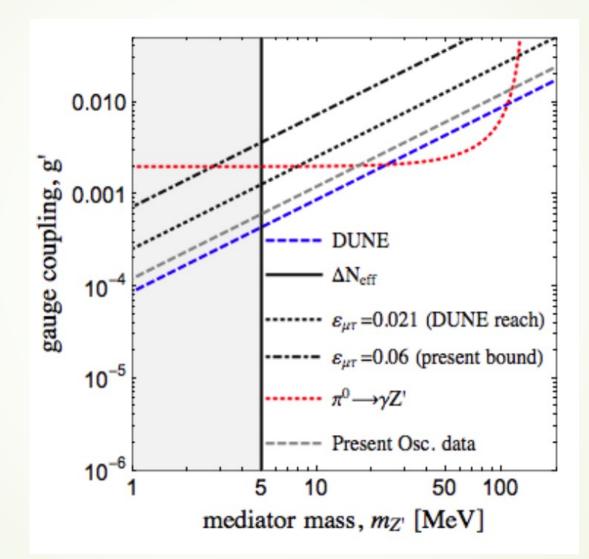
$$\epsilon^{uL} = \epsilon^{uR} = \epsilon^{dL} = \epsilon^{dR} = \frac{\zeta \eta_1(g')^2}{m_{Z'}^2} \frac{1}{2\sqrt{2}G_F} \begin{pmatrix} 0 & 0 & 0\\ 0 & -\sin 2\theta_L & \cos 2\theta_L\\ 0 & \cos 2\theta_L & \sin 2\theta_L \end{pmatrix}$$



$$\Gamma(l_{\beta}^{-} \to Z' l_{\alpha}^{-}) = \frac{(g')^2 \zeta^2}{32\pi} \frac{m_{l_{\beta}}^3}{m_{Z'}^2} (\cos^2 2\theta_L + \cos^2 2\theta_R)$$

$$Br(\tau \to Z'\mu) < 5 \times 10^{-3}$$

$$\zeta (\cos^2 2\theta_L + \cos^2 2\theta_R)^{1/2} < 3 \times 10^{-9} (\frac{1}{g'}) (\frac{m_{Z'}}{10 \text{ MeV}})^{1/2}$$



$$b_0 \tilde{R}^{\dagger} H^{\dagger} \tilde{L} + b_1 \tilde{R}^{\dagger} \sigma_1 H^{\dagger} \tilde{L},$$

$$\operatorname{Br}(H \to \tau \mu) > \frac{\operatorname{Br}(H \to \tau \tau)}{2} \left(\frac{\cos 2\theta_L}{1 + \sin 2\theta_L} \right)^2$$

$$\tan 2\theta_L = (-\epsilon_{\mu\mu} + \epsilon_{\tau\tau})/(2\epsilon_{\mu\tau}).$$

YF and J Heeck, 1607.07616

Benefits: We do not need to worry about the interactions of charged lepton. Especially $l^-_{\alpha} \to l^-_{\beta} + Z'$

Q: How? A: Mixing with a sterile Dirac fermion charged under new U(1)

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New U(1)

New Dirac fermion:

 $\Psi \to e^{ig_{\Psi}\alpha}\Psi$ New scalar electroweak doublet: $H' \to e^{ig_{\Psi}\alpha}H'$

$$\mathcal{L} = -\sum_{\alpha} y_{\alpha} \overline{L}_{\alpha} \tilde{H}' P_R \Psi + \text{h.c.},$$

$$\kappa_{\alpha} = \frac{y_{\alpha} \langle H' \rangle}{M_{\Psi}} = \frac{y_{\alpha} v \cos \beta}{\sqrt{2} M_{\Psi}} \qquad \text{tan}$$

$$\mathcal{L} = \frac{1}{2} (\bar{\nu}^c, \bar{\Psi}_L^c, \bar{\Psi}_R) \begin{pmatrix} \mathcal{M}_\nu & 0 & y \langle H' \rangle \\ 0 & 0 & M_\Psi \\ y \langle H' \rangle & M_\Psi & 0 \end{pmatrix} \begin{pmatrix} \nu \\ \Psi_L \\ \Psi_R^c \end{pmatrix} + \text{h.c.},$$

 $\beta \equiv \langle H \rangle / \langle H' \rangle$, and $v \simeq 246 \,\text{GeV}$.

$$g_{\Psi}\kappa^*_{lpha}\kappa_{eta}Z'_{\mu}ar{
u}_{lpha}\gamma^{\mu}
u_{eta}$$
= Standard Quarks: $q
ightarrow e^{ig_Blpha/3}q$

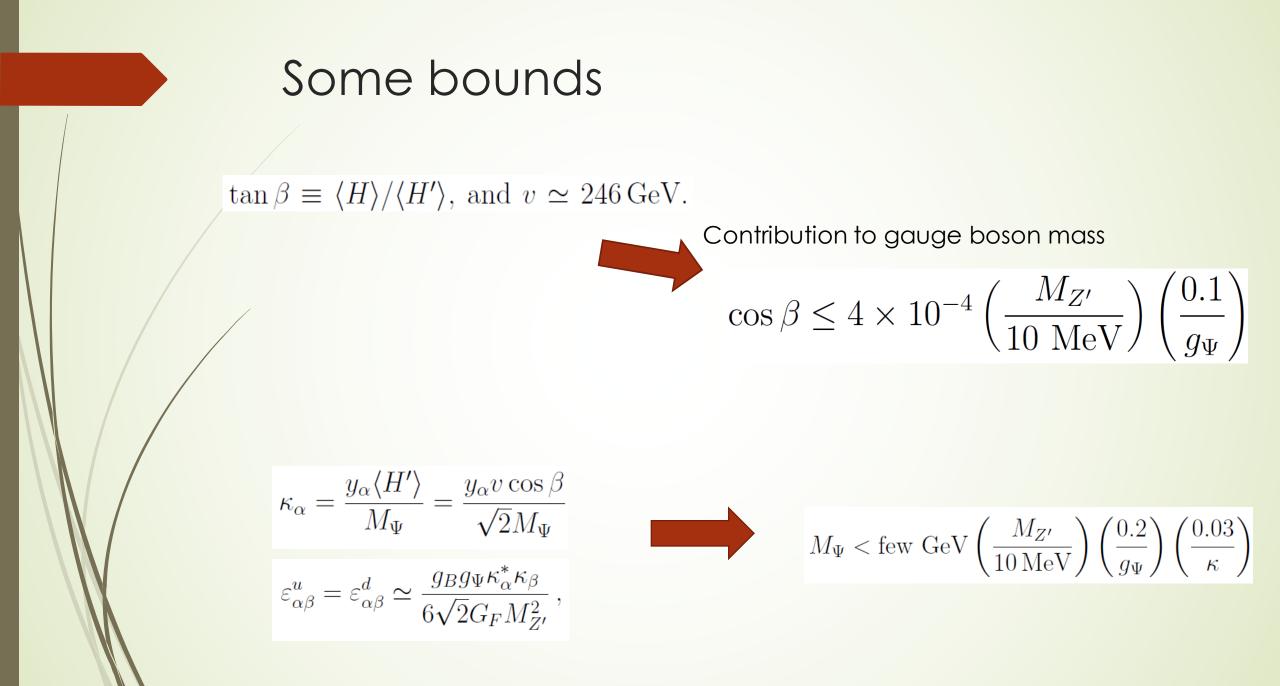
Non-standard Interaction

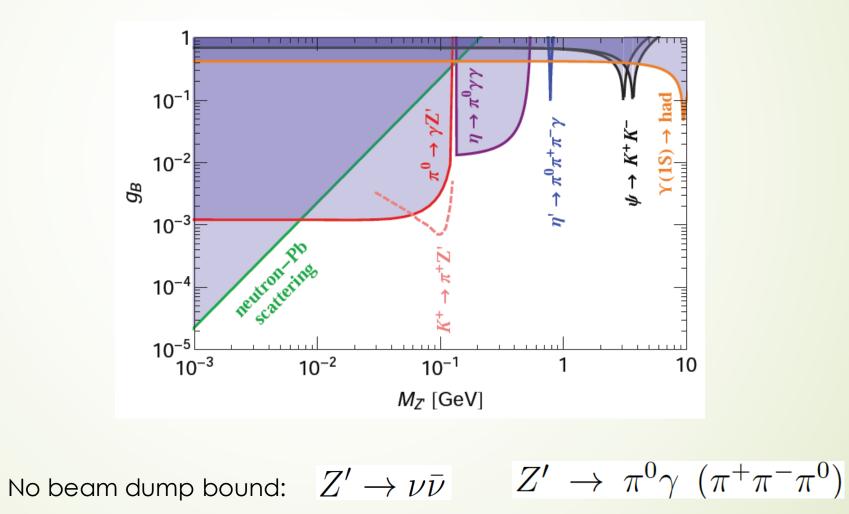
$$\varepsilon_{\alpha\beta}^{u} = \varepsilon_{\alpha\beta}^{d} \simeq \frac{g_{B}g_{\Psi}\kappa_{\alpha}^{*}\kappa_{\beta}}{6\sqrt{2}G_{F}M_{Z'}^{2}}, \qquad |\varepsilon_{\alpha\beta}| = \sqrt{\varepsilon_{\alpha\alpha}\varepsilon_{\beta\beta}}.$$

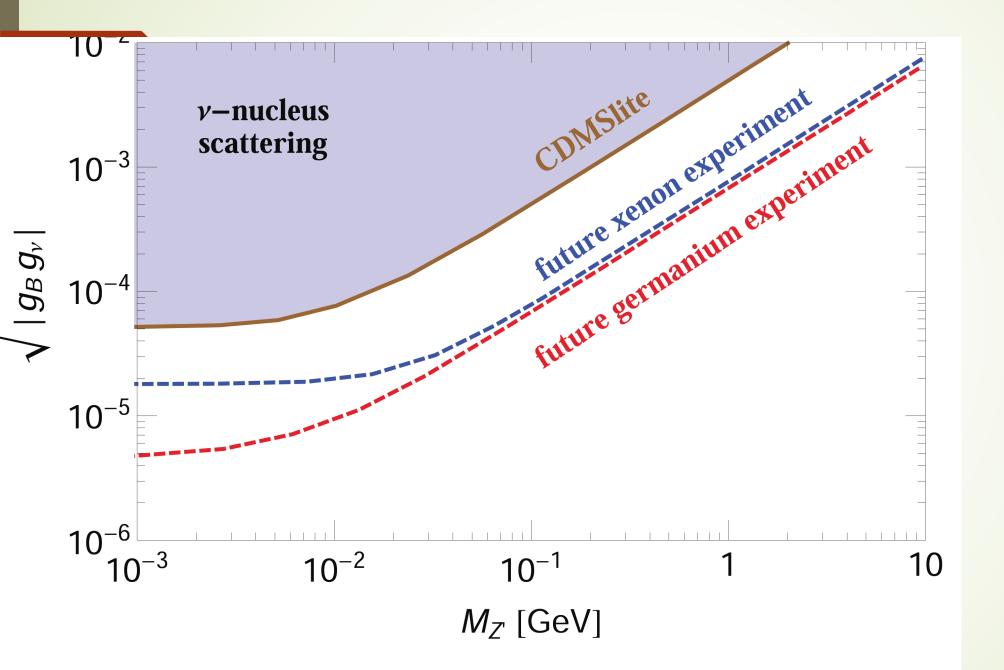
Schwartz inequality:

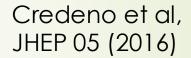
$$g_{\Psi}\kappa^*\kappa^T \to \sum_j g_{\Psi_j}\kappa^*_{\Psi_j}\kappa^T_{\Psi_j}$$

$$|\varepsilon_{\alpha\beta}| \leq \sqrt{\varepsilon_{\alpha\alpha}\varepsilon_{\beta\beta}}$$









LUX-ZEPLIN SuperCDMS

Violation of the unitarity of the PMNS matrix

Muon decay and tests of lepton flavor universality

 $|\kappa_e|^2 < 2.5 \times 10^{-3}, \ |\kappa_\mu|^2 < 4.4 \times 10^{-4}, \text{ and } |\kappa_\tau|^2 < 5.6 \times 10^{-3} \text{ at } 2\sigma.$

Fernandez-Martinez, Hernandez-Garcia and Lopex-Pavon, JHEP 1608 (2016)

 $|\kappa_{\mu}\kappa_{e}| < 10^{-3}$, $|\kappa_{\mu}\kappa_{\tau}| < 1.6 \times 10^{-3}$, and $|\kappa_{e}\kappa_{\tau}| < 3.7 \times 10^{-3}$

Bounds from $l_{\alpha} \rightarrow l_{\beta}\gamma$ and $l_{\alpha} \rightarrow Z'l_{\beta}$ are weaker.

UV completion

Anomaly cancelation

• One example: Two generation of colorless fermions with opposite chiralities and U(1)' charges B_1 and B_2 ,

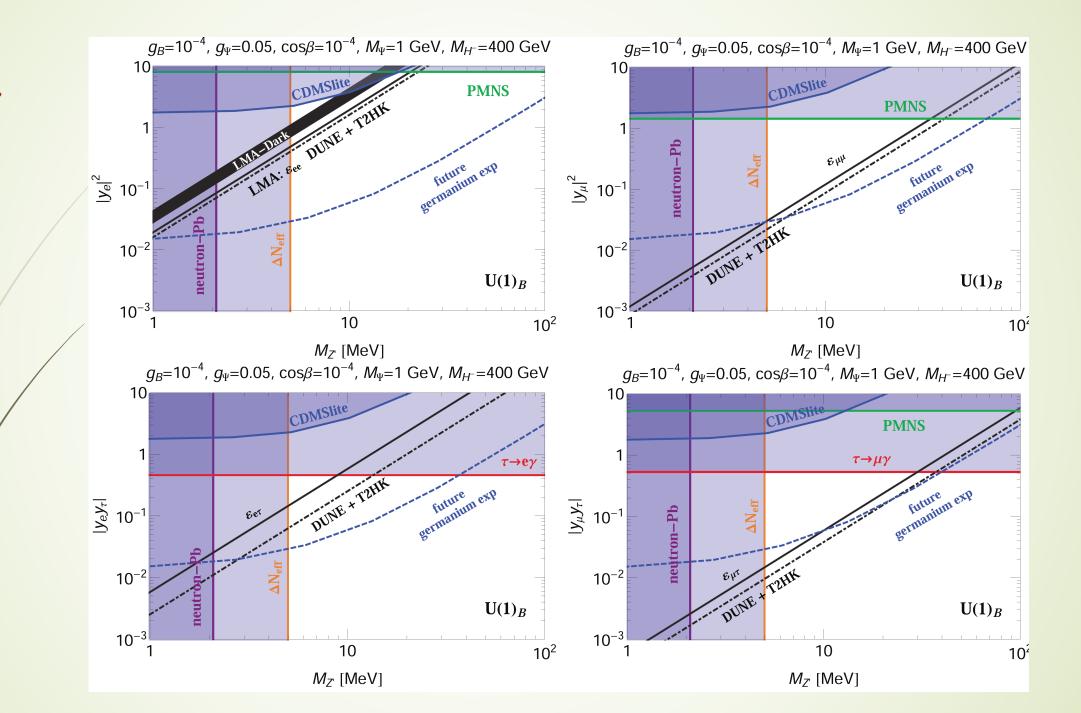
$$B_1 - B_2 = -3$$
 Duerr et al., PRL 110 (2013); PRD 91 (2015)

Electroweak singlet giving mass to new fermions

Taking $\langle S_B \rangle \gtrsim 1 \,\mathrm{TeV}$

 $M_{Z'}$ given by $g_B \langle S_B \rangle$

$$g_B \lesssim 10^{-4} \left(\frac{M_{Z'}}{100 \,\mathrm{MeV}}\right)$$



Collider phenomenology

LEP+LHC
$$L = 19.5 \, \text{fb}^{-1}$$
 and $\sqrt{s} = 8 \, \text{TeV}$
 $M_{H^-} > 275 \, \text{GeV}$

CMS Khachatryan et al., Eur Phys J C74 (2014) no 9, 3036

$$\frac{\operatorname{Br}(H^{-} \to \ell_{\alpha} \Psi)}{\operatorname{Br}(H^{-} \to \ell_{\beta} \Psi)} \simeq \frac{|y_{\alpha}|^{2}}{|y_{\beta}|^{2}} \simeq \frac{\varepsilon_{\alpha\alpha}}{\varepsilon_{\beta\beta}}$$
$$\Psi \to \nu Z' \qquad Z' \to \nu \bar{\nu}$$

Observational consequences

Emission in Supernova

• Similar to $\mathcal{L}_{\mu} - \mathcal{L}_{ au}$

Kamada and Yu, PRD 92 (2015)

 $c\tau_{Z'} \sim 10^{-9} \mathrm{km} (g'/7 \times 10^{-5})^{-2} (T/10 \text{ MeV}) (10 \text{ MeV}/m_{Z'})^2$

Reduced mean free path for neutrinos

prolong the diffusion time

High energy cosmic neutrino

Kamada and Yu, PRD 92 (2015)

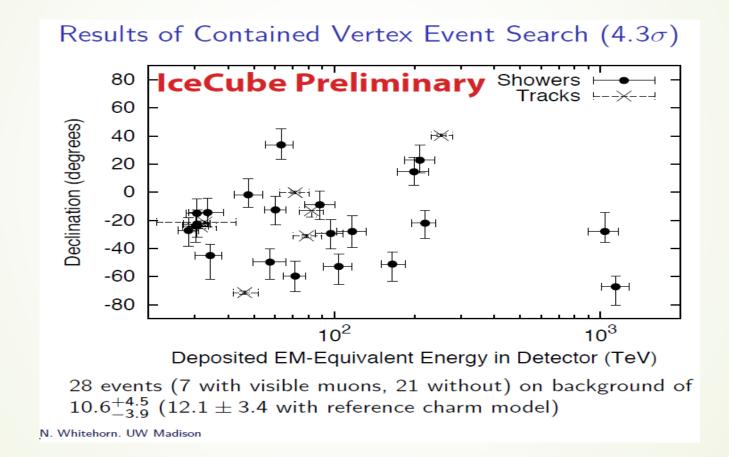
$$\mathcal{L}_{\mu} - \mathcal{L}_{ au}$$

$$\nu\nu \to Z' \to \nu\nu$$

Background neutrino at rest

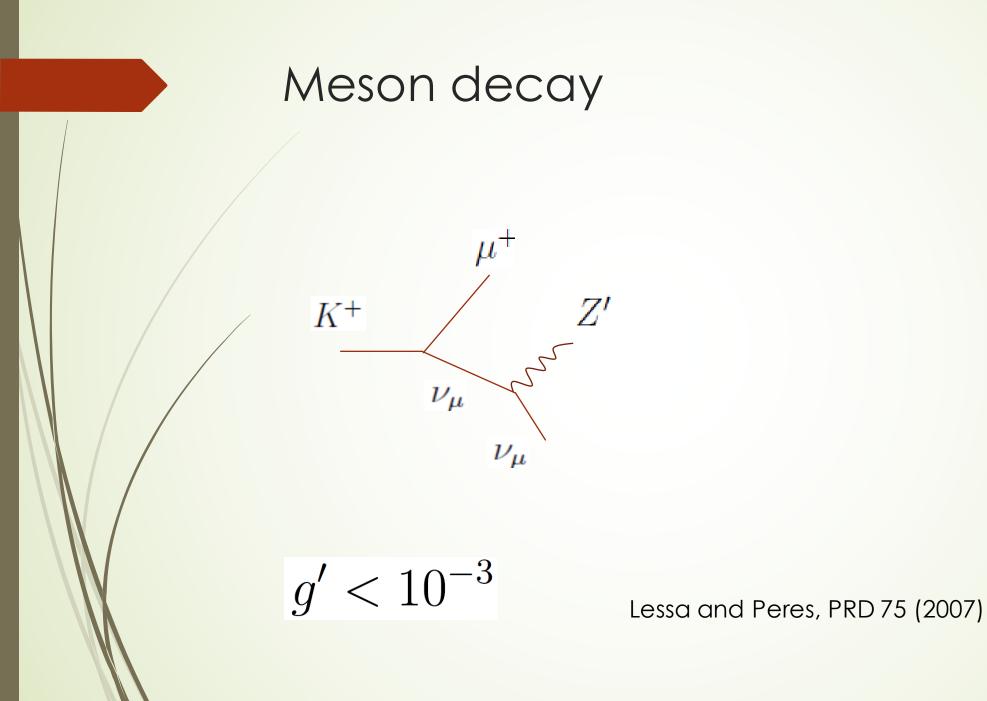
 $400~{\rm TeV}$ to ${\rm PeV}$

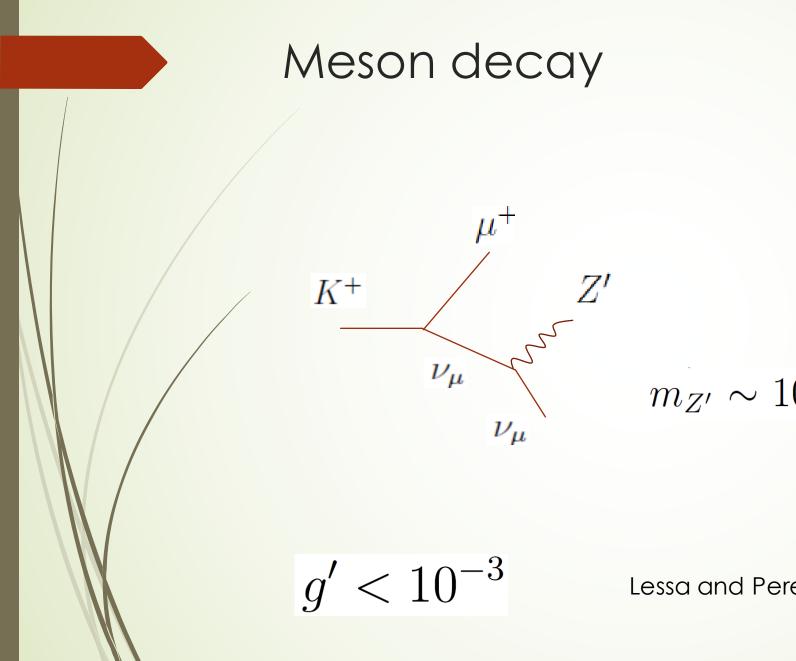
Dip or gap in ICECUBE spectrum











$m_{Z'} \sim 10 \text{ MeV} \text{ and } g' \sim 7 \times 10^{-5}$

Lessa and Peres, PRD 75 (2007)

Summary

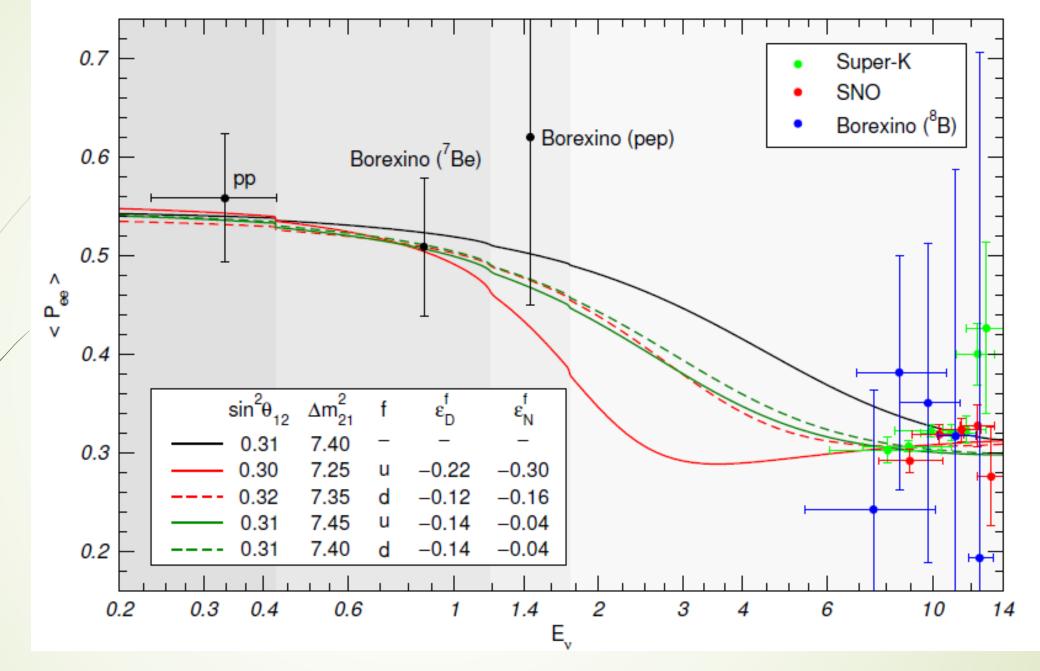
Viable models for sizeable neutral current NSI based on U(1)'

$$\sqrt{g_{\nu}g_B} \sim 7 \times 10^{-5} \frac{m_{Z'}}{10 \text{ MeV}}$$

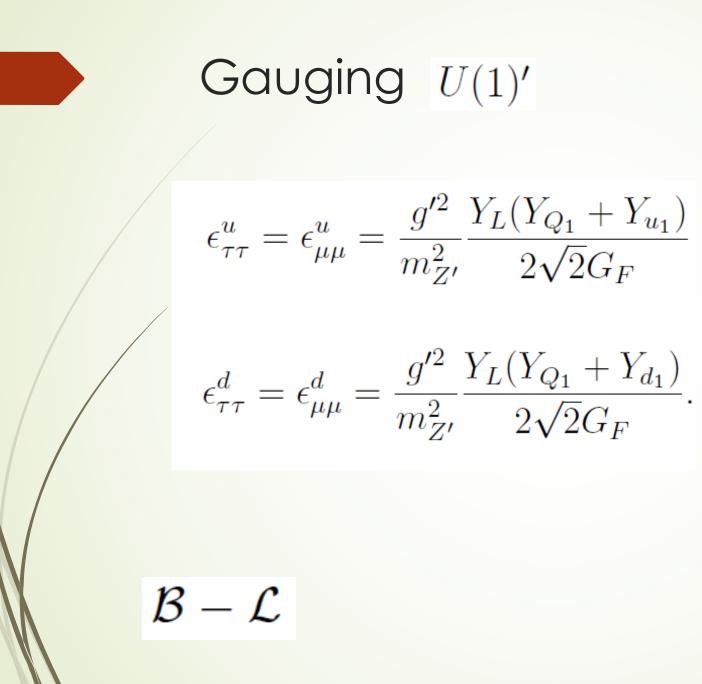
- Rich phenomenology: Prospect of testing via
- 1) SN neutrino
- 2) Dip below PeV in the spectrum of cosmic neutrinos
- 3) Rare meson decay
- 4) Interaction rate of solar neutrinos at dark matter direct detection experiments



Present 90% C.L. COLOMA, JHEP 03 (2016) $|\varepsilon_{e\mu}^{u} + \varepsilon_{e\mu}^{d}| < 0.12, \qquad |\varepsilon_{e\tau}^{u} + \varepsilon_{e\tau}^{d}| < 0.18, \qquad |\varepsilon_{\mu\tau}^{u} + \varepsilon_{\mu\tau}^{d}| < 0.018$ $0.11 < \varepsilon_{\scriptscriptstyle ee}^u + \varepsilon_{\scriptscriptstyle ee}^d - \varepsilon_{\tau\tau}^u - \varepsilon_{\tau\tau}^d < 0.60\,, \quad \text{and} \quad -0.04 < \varepsilon_{\mu\mu}^u + \varepsilon_{\mu\mu}^d - \varepsilon_{\tau\tau}^u - \varepsilon_{\tau\tau}^d < 0.037$ DUNE 90 % C.L. $\left|\varepsilon_{e\mu}^{u} + \varepsilon_{e\mu}^{d}\right| < 0.024, \qquad \left|\varepsilon_{e\tau}^{u} + \varepsilon_{e\tau}^{d}\right| < 0.08, \qquad \left|\varepsilon_{\mu\tau}^{u} + \varepsilon_{\mu\tau}^{d}\right| < 0.012,$ $0.017 < \varepsilon_{ee}^{u} + \varepsilon_{ee}^{d} - \varepsilon_{\tau\tau}^{u} - \varepsilon_{\tau\tau}^{d} < 0.43, \quad \text{and} \quad -0.027 < \varepsilon_{\mu\mu}^{u} + \varepsilon_{\mu\mu}^{d} - \varepsilon_{\tau\tau}^{u} - \varepsilon_{\tau\tau}^{d} < 0.025.$



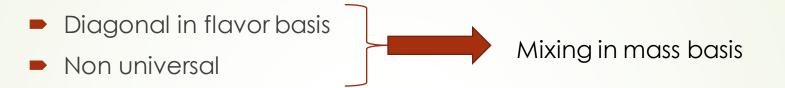
Maltoni and Gonzalez-Garcia, JHEP 2013



 $\epsilon_{\mu\mu}^{u,d} \simeq \epsilon_{\tau\tau}^{u,d} \sim 1$

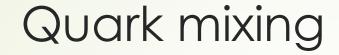
 $\mathcal{L}_{\mu} - \mathcal{L}_{ au}$

Flavor physics in quark sector



Universal coupling of first and second generations

Avoiding large contribution to $K - \overline{K}$ or $D - \overline{D}$ mixing,

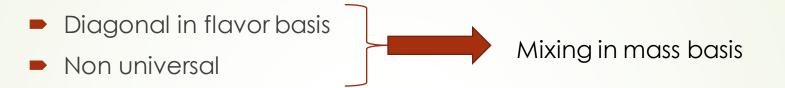


$$\mathcal{L}_{\mu} + \mathcal{L}_{\tau} + \mathcal{B}_{1} - a\mathcal{B}_{2} - (3 - a)\mathcal{B}_{3}$$
$$a = -1$$

Mixing between first and second generations

No mixing between third generation and the rest $H'^{\dagger}car{Q}_{3}u_{1,2}$ $ar{d}_{3}H'^{\dagger}Q_{1,2}$

Flavor physics in quark sector



Universal coupling of first and second generations

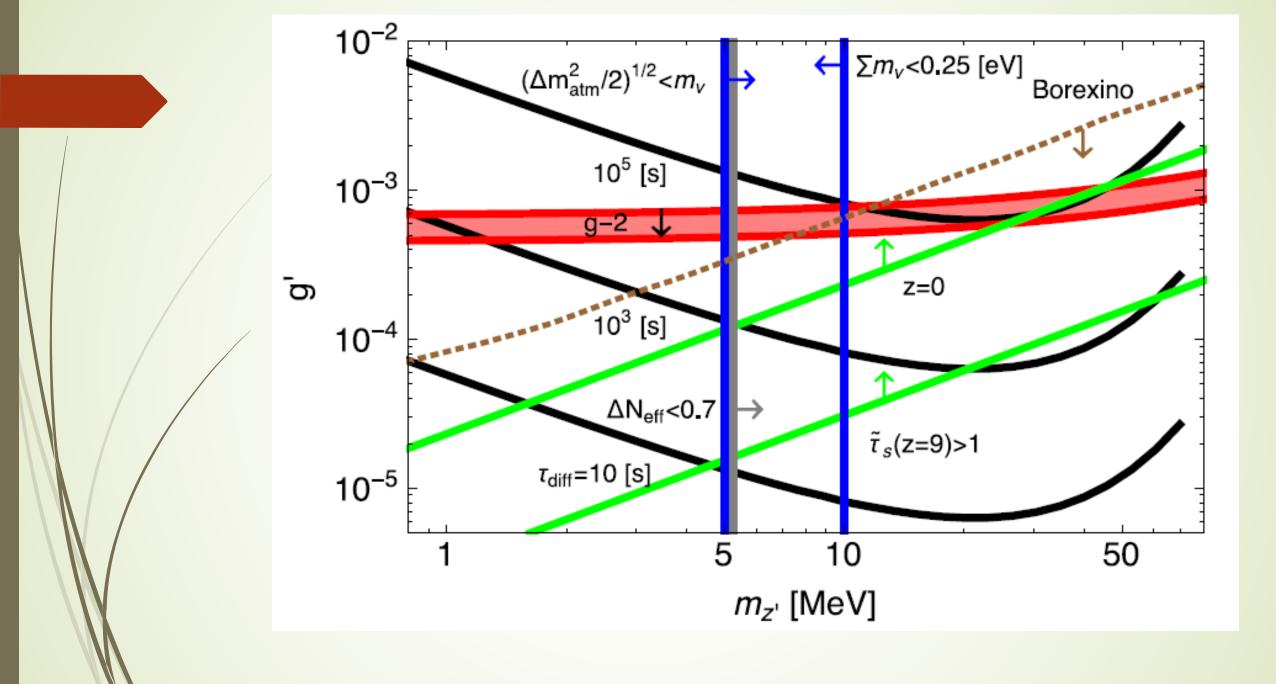
Avoiding large contribution to $K - \overline{K}$ or $D - \overline{D}$ mixing,

Crivellini, D'Amrosio and Heeck, PRD 91 (2015) 075006

$$\begin{split} O_9^{\ell\ell} &= \frac{\alpha_{\rm EM}}{4\pi} \left[\bar{s} \gamma^{\mu} P_L b \right] \left[\bar{\ell} \gamma_{\mu} \ell \right], \\ O_9^{\prime\ell\ell} &= \frac{\alpha_{\rm EM}}{4\pi} \left[\bar{s} \gamma^{\mu} P_R b \right] \left[\bar{\ell} \gamma_{\mu} \ell \right], \\ C_9^{\mu\mu} &\simeq \frac{-g^{\prime 2}}{\sqrt{2}m_{Z^\prime}^2} \frac{\pi}{\alpha_{\rm EM}} \frac{1}{G_F} a \simeq -\left(\frac{a}{1/3} \right) \left(\frac{3 \,{\rm TeV}}{m_{Z^\prime}/g^\prime} \right)^2, \end{split}$$

Global fit R(K) and $B \to K^* \mu^+ \mu^- R(K) = \frac{B \to K \mu^+ \mu^-}{B \to K e^+ e^-} = 0.745^{+0.090}_{-0.074} \pm 0.036$,

 $-0.60(-0.95) \ge C_9^{\mu\mu} \ge (-1.65) - 2.00$



Kamada and Yu, 1504.00711

Determining

Shedding light on LMA-Dark solar neutrino solution by medium baseline reactor experiments: JUNO and RENO-50

YF and Bakhti, JHEP 2014

$$P(\bar{\nu}_e \to \bar{\nu}_e) = \left| |U_{e1}|^2 + |U_{e2}|^2 e^{i\Delta_{21}} + |U_{e3}|^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{12}^2 c_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2 = \left| c_{12}^2 c_{13}^2 + s_{13}^2 e^{i\Delta_{21}} + s_{13}^2 e^{i\Delta_{31}} \right|^2$$

$$c_{13}^4 (1 - \sin^2 2\theta_{12} \sin^2 \frac{\Delta_{21}}{2}) + s_{13}^4 + 2s_{13}^2 c_{13}^2 [\cos \Delta_{31} (c_{12}^2 + s_{12}^2 \cos \Delta_{21}) + s_{12}^2 \sin \Delta_{31} \sin \Delta_{21}]$$

Medium Baseline reactor experiments

DAYA BAY in CHINA
 RENO in South Korea
 Ready for data taking in 2020.

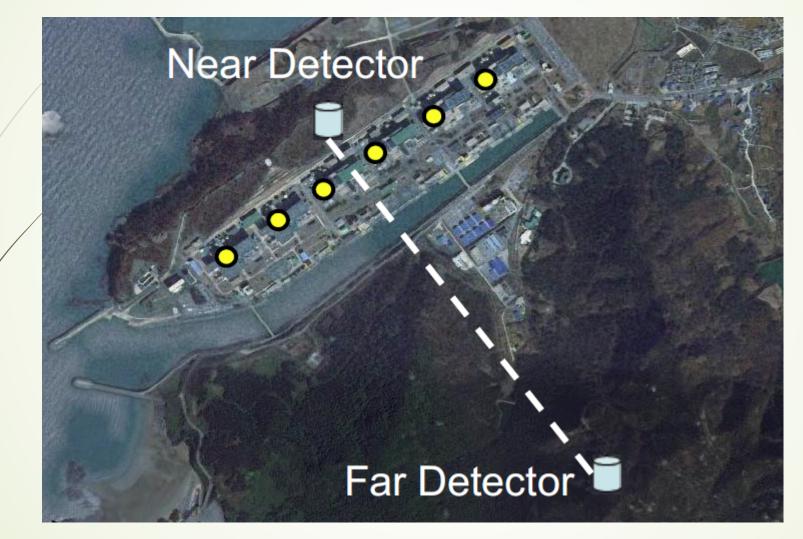
Baseline ~ 50 km

$$\frac{\Delta m_{01}^2 L}{2E_{\nu}} \sim 0.4 \frac{\Delta m_{01}^2}{10^{-5} \text{ eV}^2} \frac{L}{50 \text{ km}} \frac{3 \text{ MeV}}{E_{\nu}}$$

Main goal determination of

$$\operatorname{sgn}(\Delta m_{31}^2)$$

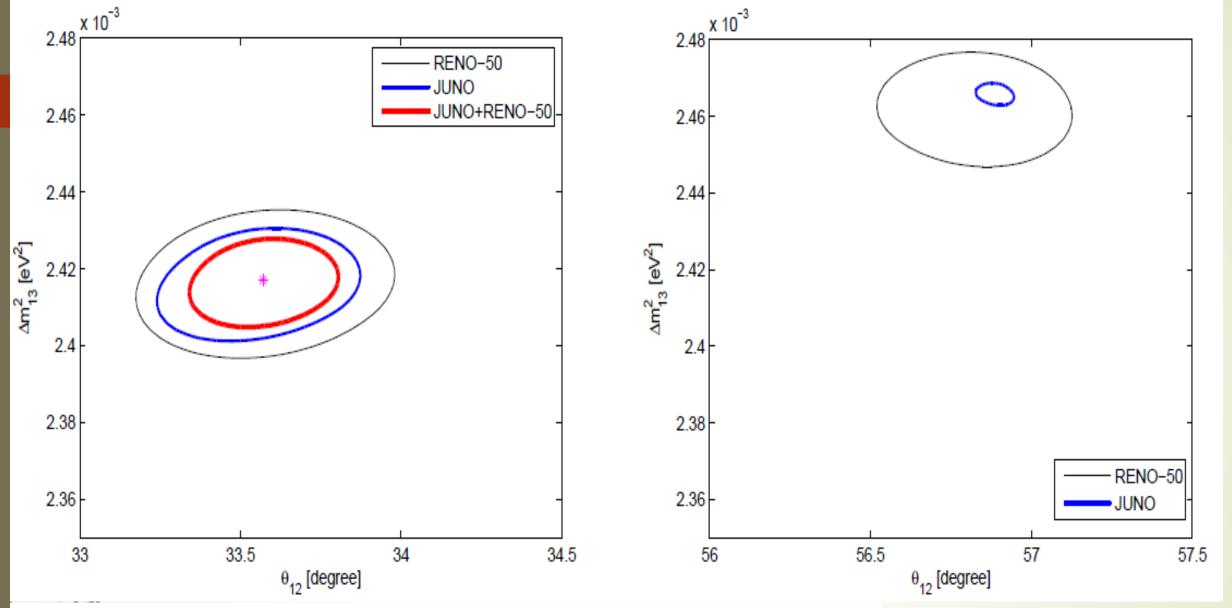
RENO-50 in South Korea





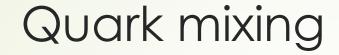
Daya Bay and Juno





Allowed region at 3 σ C.L. after 5 years of data taking by RENO-50 and JUNO.

 $\frac{\delta E_{\nu}}{E_{\nu}} \simeq 3\% \times (\frac{E_{\nu}}{\text{MeV}})^{1/2}$



$$\mathcal{L}_{\mu} + \mathcal{L}_{\tau} + \mathcal{B}_{1} - a\mathcal{B}_{2} - (3 - a)\mathcal{B}_{3}$$
$$a = -1$$

Mixing between first and second generations

No mixing between third generation and the rest $H'^{\dagger}car{Q}_{3}u_{1,2}$ $ar{d}_{3}H'^{\dagger}Q_{1,2}$



Flavor violation

 $H'^{\dagger} c \bar{Q}_{3} u_{1,2} = \bar{d}_{3} H'^{\dagger} Q_{1,2}$

 $m_{H'} \gg 100 \,\,\mathrm{GeV}$

 $\langle H' \rangle \stackrel{<}{\sim} m_{EW}$

Small VEV

$$-m_{S}^{2}|S|^{2} + \lambda_{S}|S|^{4} + m_{SHH'}SH^{\dagger}H' + m_{H'}^{2}H'^{\dagger}H'$$

$$0 < m_S^2 \ll m_{H'}^2$$
 and $m_{EW}^2 < m_{H'}^2$

$$\langle S \rangle = \left(\frac{m_S^2}{2\lambda_S}\right)^{1/2}, \quad \langle H' \rangle = -m_{SHH'} \frac{\langle H \rangle \langle S \rangle}{2m_{H'}^2}.$$

Small VEV $\left|-m_{S}^{2}|S|^{2}+\lambda_{S}|S|^{4}+m_{SHH'}SH^{\dagger}H'+m_{H'}^{2}H'^{\dagger}H'\right|$ $0 < m_S^2 \ll m_{H'}^2$ and $m_{EW}^2 < m_{H'}^2$ $\langle S \rangle = \left(\frac{m_S^2}{2\lambda_S}\right)^{1/2}, \quad \langle H' \rangle = -m_{SHH'} \frac{\langle H \rangle \langle S \rangle}{2m_{H'}^2}.$ $m_{SHH'}, \langle H \rangle \ll m_{H'} \text{ and } \langle S \rangle \stackrel{<}{\sim} \langle H \rangle$

Small VEV $\left|-m_{S}^{2}|S|^{2}+\lambda_{S}|S|^{4}+m_{SHH'}SH^{\dagger}H'+m_{H'}^{2}H'^{\dagger}H'\right|$ $0 < m_S^2 \ll m_{H'}^2$ and $m_{EW}^2 < m_{H'}^2$ $\langle S \rangle = \left(\frac{m_S^2}{2\lambda_S}\right)^{1/2}, \quad \langle H' \rangle = -m_{SHH'} \frac{\langle H \rangle \langle S \rangle}{2m_{H'}^2}.$ $m_{SHH'}, \langle H \rangle \ll m_{H'} \text{ and } \langle S \rangle \stackrel{<}{\sim} \langle H \rangle$ $\ll m_{H'}, m_{EW}, \langle S \rangle$

Mass of new gauge boson

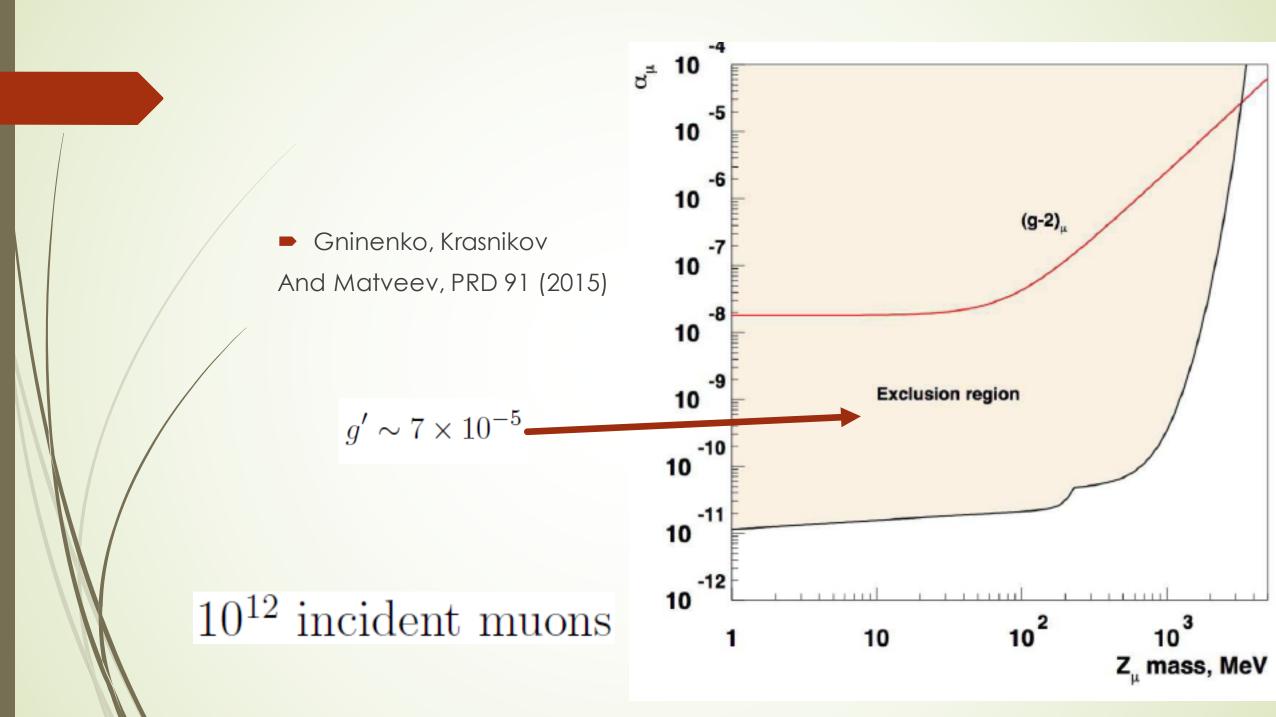
$$m_{Z'} = g' \left(Y_{S_1}^2 \langle S_1 \rangle^2 + Y_{S_2}^2 \langle S_2 \rangle^2 + Y_S^2 \langle S \rangle^2 + Y_{H'}^2 \langle H' \rangle^2 \right)^{1/2}$$
$$m_{Z'} \sim 10 \text{ MeV and } g' \sim 7 \times 10^{-5}$$
$$\langle S_1 \rangle \sim \langle S_2 \rangle \sim 100 \text{ GeV}$$



 $\mu + A \rightarrow \mu + A + Z', Z' \rightarrow \nu \bar{\nu}$

muon beam with energy of 150 GeV CERN SPS

Gninenko, Krasnikov and Matveev PRD 91 (2015)



Muon magnetic dipole moment Z' $\Delta (g-2)_{\mu}/2 = g'^2/8\pi^2$ $(g-2)_{\mu}/2 \sim 5 \times 10^{-11}$

Neutrino trident scattering

 $\nu + A \rightarrow \nu + A + \mu^+ + \mu^-$

CCFR collaboration:

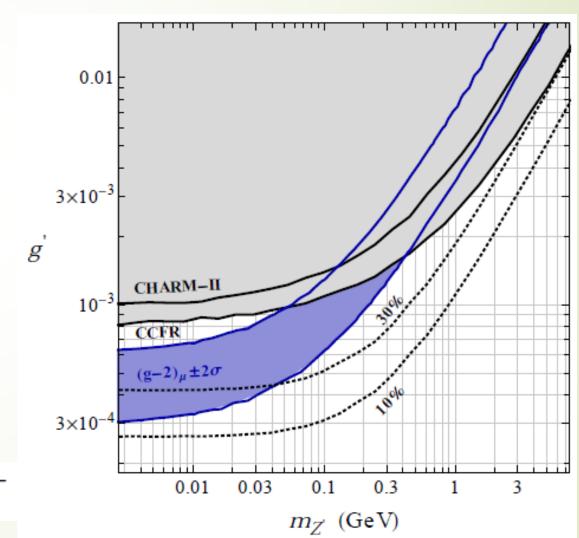
scattering of $\sim 160 \text{ GeV}$ neutrino beam off an iron target

PRL66 (1991)

CHARM II collaboration

scattering of $\sim 20 \text{ GeV}$ neutrino beam off a glass target

Neutrino trident scattering



Altmannshofer et al., PRL113 (2014)

 $\nu + A \rightarrow \nu + A + \mu^+ + \mu^-$

Yukawa coupling of neutrinos

 $\lambda_1 \bar{N}_1 H^T c L_e + \lambda_2 \bar{N}_2 H^T c L_\mu + \lambda_3 \bar{N}_3 H^T c L_\tau + \lambda_4 \bar{N}_2 H^T c L_\tau + \lambda_5 \bar{N}_3 H^T c L_\mu$

$$Y_{N_2} = Y_{N_3} = Y_{L_{\mu}} = Y_{L_{\tau}} = 1$$
$$Y_{N_1} = 0$$

Yukawa coupling of neutrinos

 $\lambda_1 \bar{N}_1 H^T cL_e + \lambda_2 \bar{N}_2 H^T cL_\mu + \lambda_3 \bar{N}_3 H^T cL_\tau + \lambda_4 \bar{N}_2 H^T cL_\tau + \lambda_5 \bar{N}_3 H^T cL_\mu + \text{H.c.}$

Basis change: $\lambda_4 = 0 \text{ or } \lambda_5 = 0$

No mixing:

Mix:

 $\nu_e \text{ and } \nu_\mu \qquad \nu_e \text{ and } \nu_\tau$

 ν_{μ} and ν_{τ}

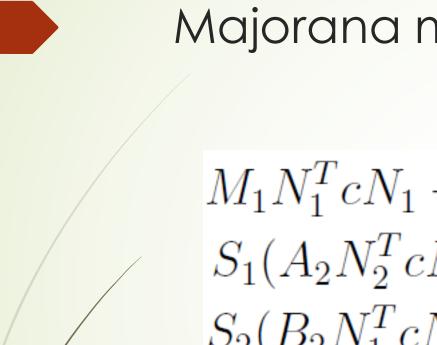


If there is no Majorana mass for right-handed neutrinos:

1) $m_{N_i} \sim m_{\nu}$

 (ΔN_{eff})

2) Smallness of neutrino mass



Majorana masses

$M_1 N_1^T c N_1 +$ $S_1(A_2N_2^TcN_2 + A_3N_3^TcN_3 + A_{23}N_2^TcN_3) +$ $S_2(B_2N_1^T cN_2 + B_3N_1^T cN_3) + \text{H.c.}$