

Charged Lepton Flavor Violation - Introduction -

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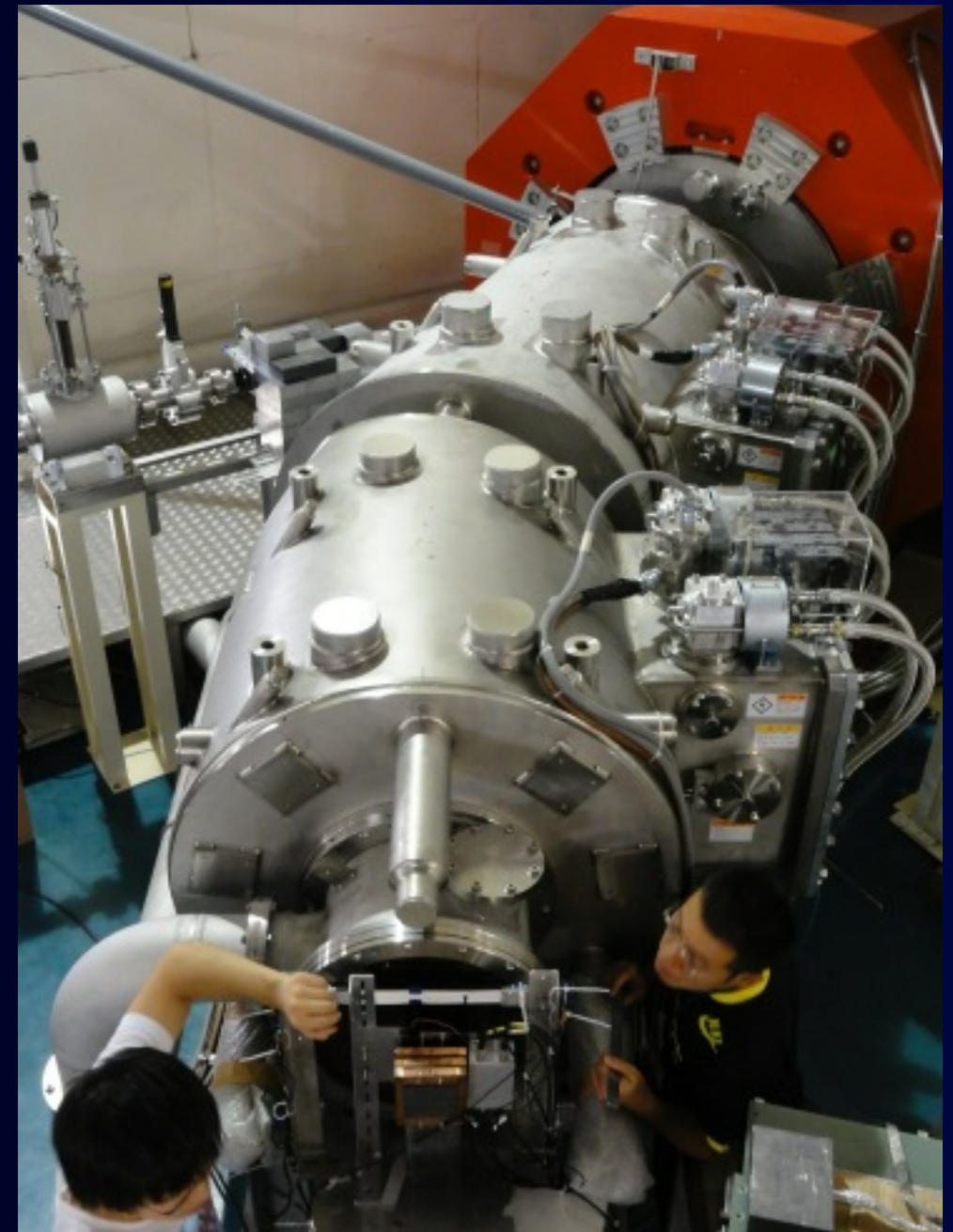
Xth Rencontres du
Vietnam
Flavor Physics

July 31st, 2014



Outline

- Why Charged Lepton Flavor Violation (CLFV) ?
- New Physics in CLFV
- CLFV Experiments
- CLFV in Tau Decays
- CLFV in Muon Decays
 - $\mu \rightarrow e \gamma$
 - $\mu N \rightarrow e N$
- Highly Intense Muon Sources
 - MuSIC at Osaka University
- Summary



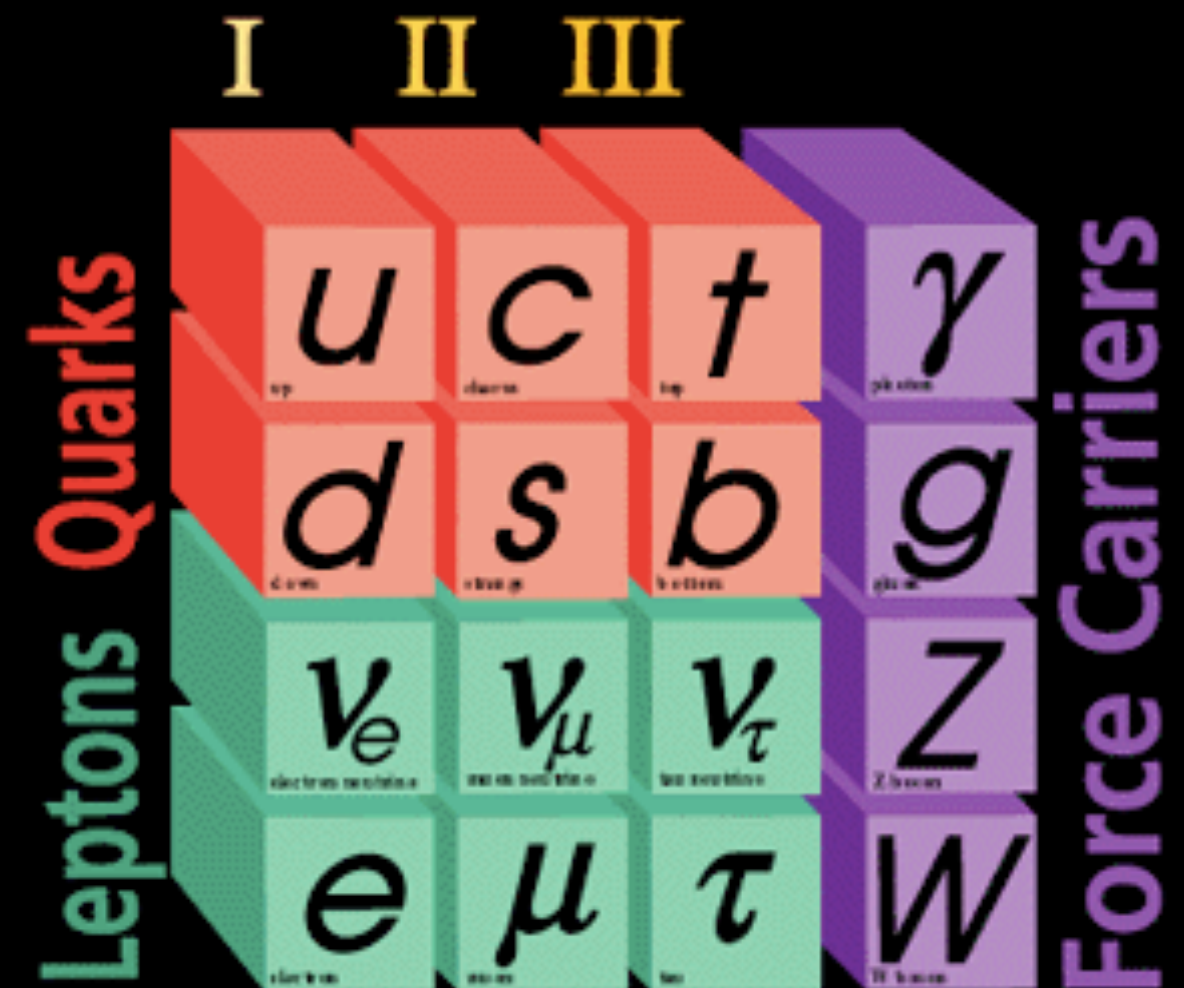
Why CLFV ?



The Higgs boson has been found,
but no new particles are found yet...

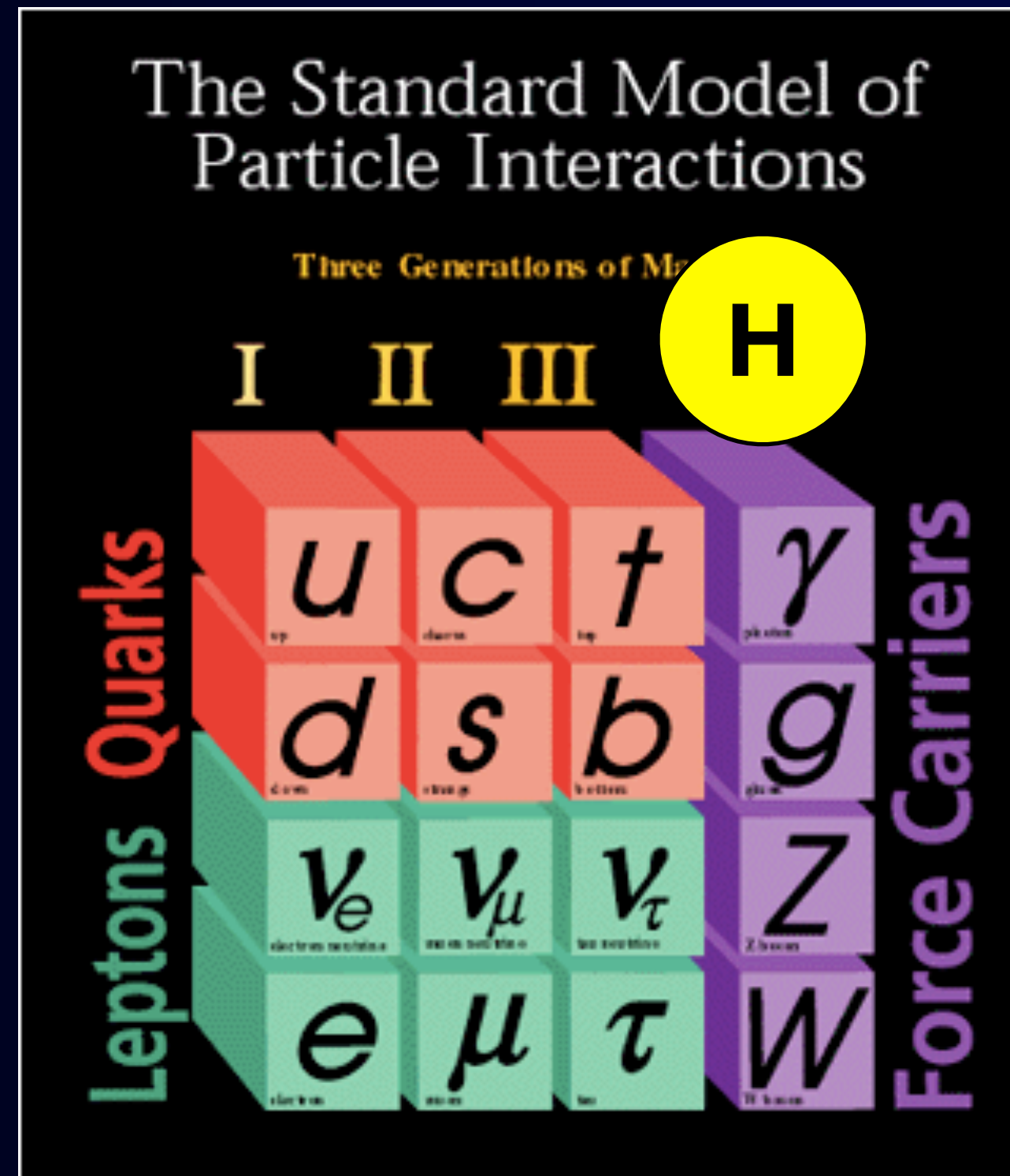
The Standard Model of Particle Interactions

Three Generations of Matter



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The discovery of the Higgs
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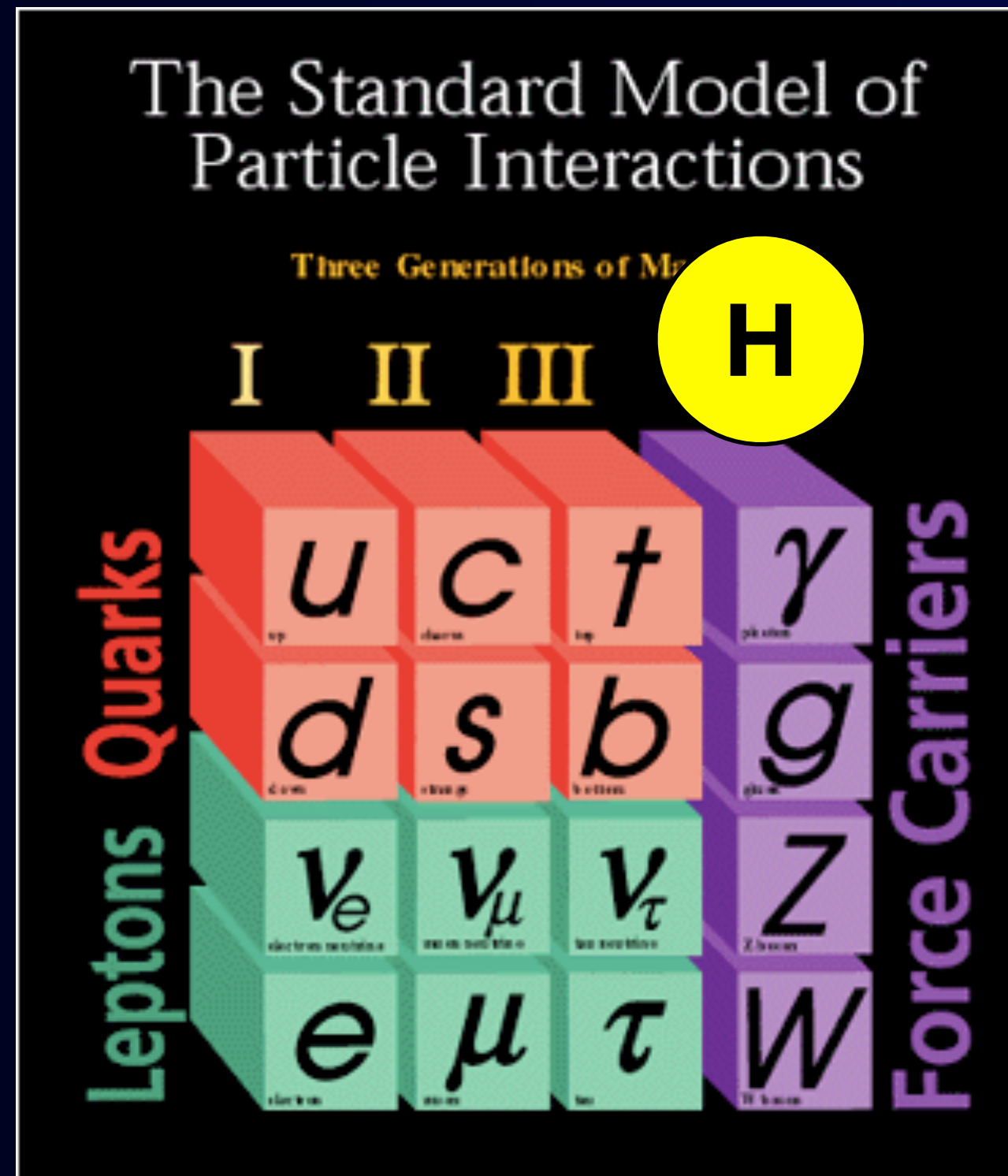


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The Standard Model can explain most of the experimental results. However, there are many undetermined parameters and issues.



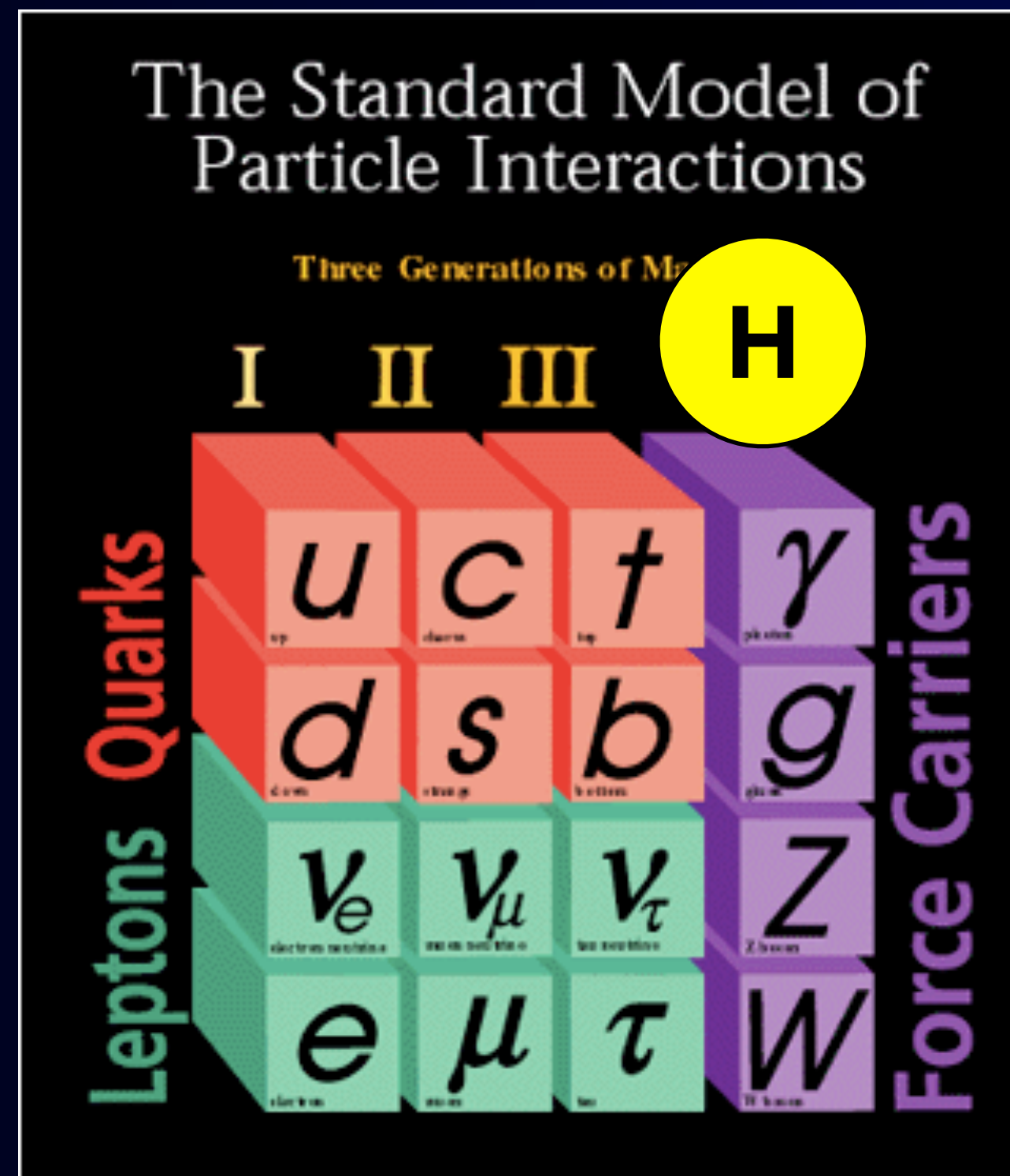
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The Standard Model is considered to be incomplete.
New Physics is needed.



New Physics Search in Charged Lepton Flavor



New Physics Search in Charged Lepton Flavor



with new physics contributions

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)},$$

Dimension 6 Operation for
New Physics
 Λ is the energy scale of new
physics

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Charged Lepton Flavor

For instance, $\mu \rightarrow e \gamma$ ($B < 5.7 \times 10^{-13}$),

$$\frac{C_{\text{NP}}}{\Lambda^2} O_{ij}^{(6)} \rightarrow \frac{C_{\mu e}}{\Lambda^2} \bar{e}_L \sigma^{\rho\nu} \mu_R \Phi F_{\rho\nu}$$
$$\Lambda > 2 \times 10^5 \text{ TeV} \times (C_{\mu e})^{\frac{1}{2}}.$$

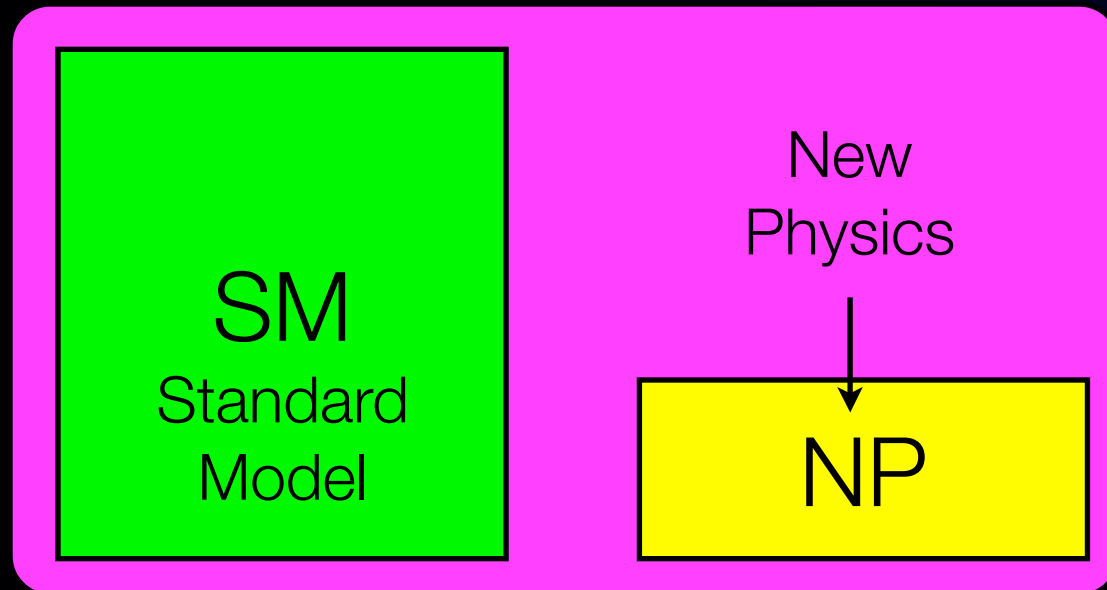
$$\Lambda > O(10^5) \text{ TeV}$$

Guideline for Rare Decay Searches

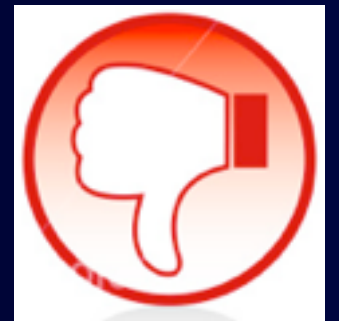
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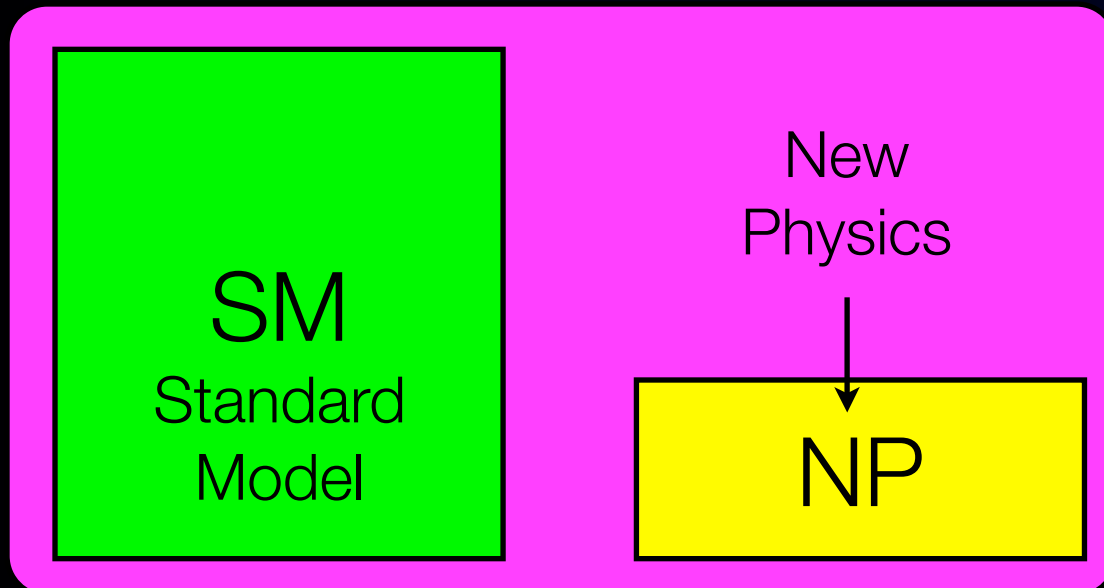


SM contribution is dominant.

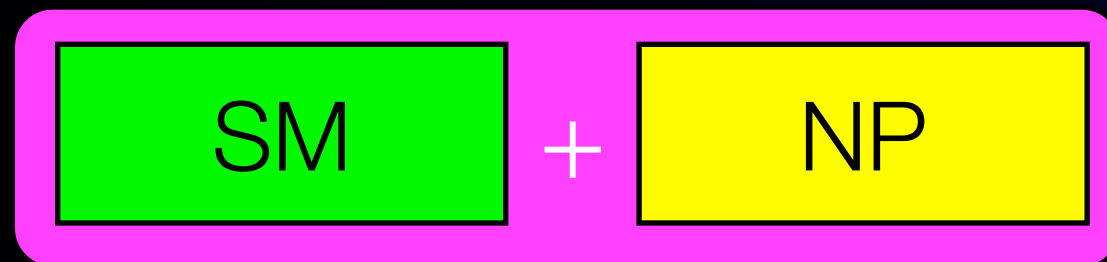


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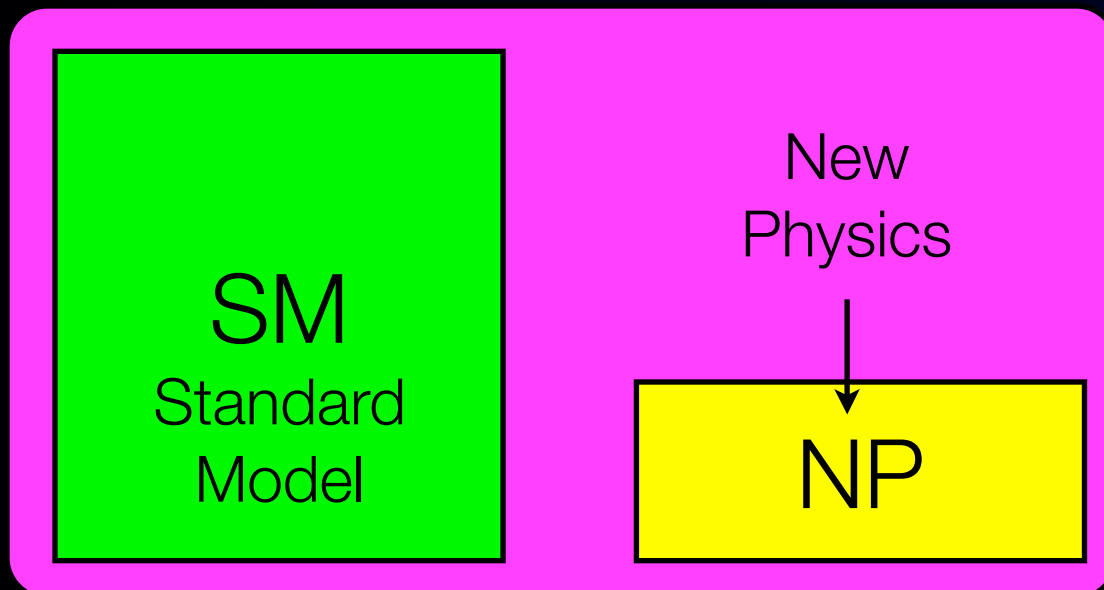


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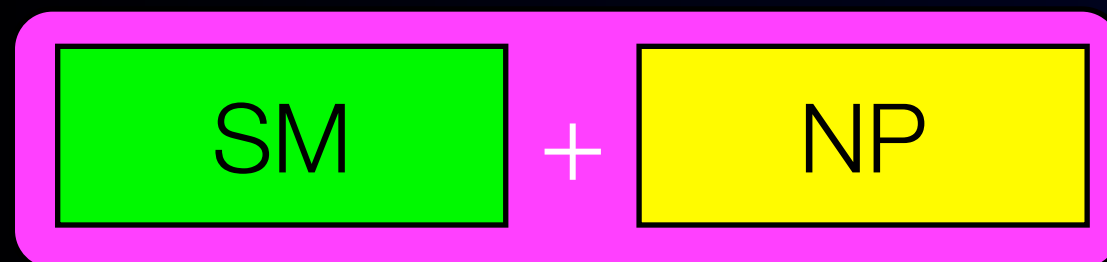
$$B \sim \frac{1}{\sqrt{N}}$$

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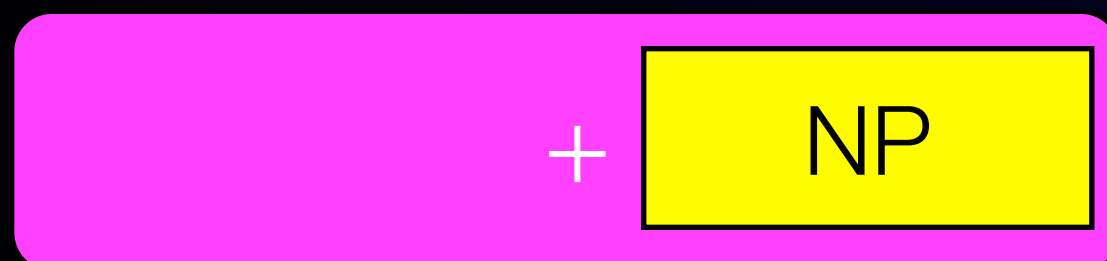


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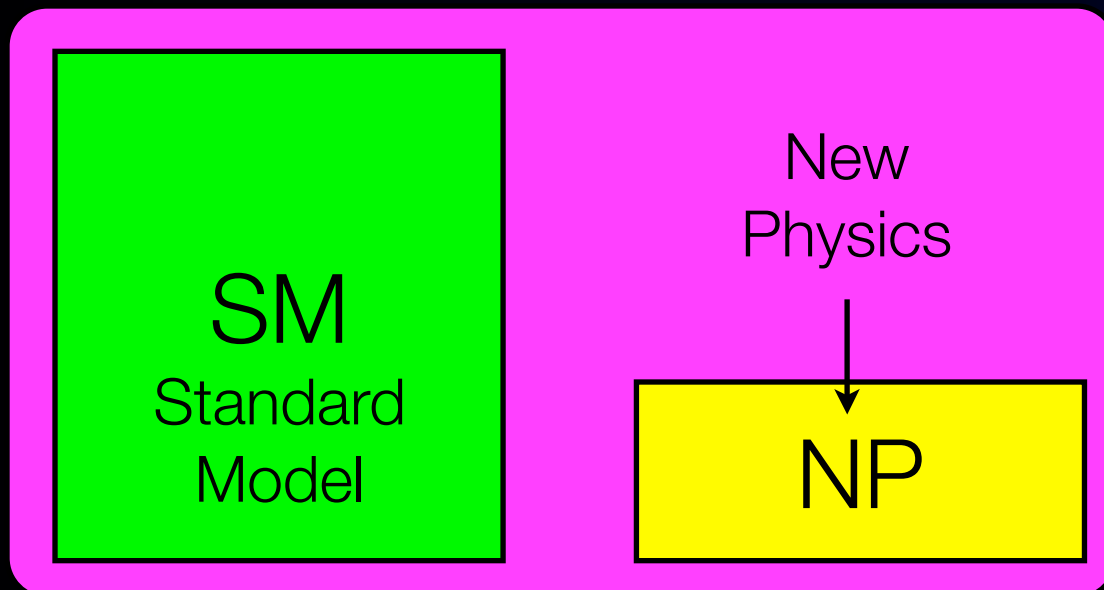


SM contribution is forbidden.

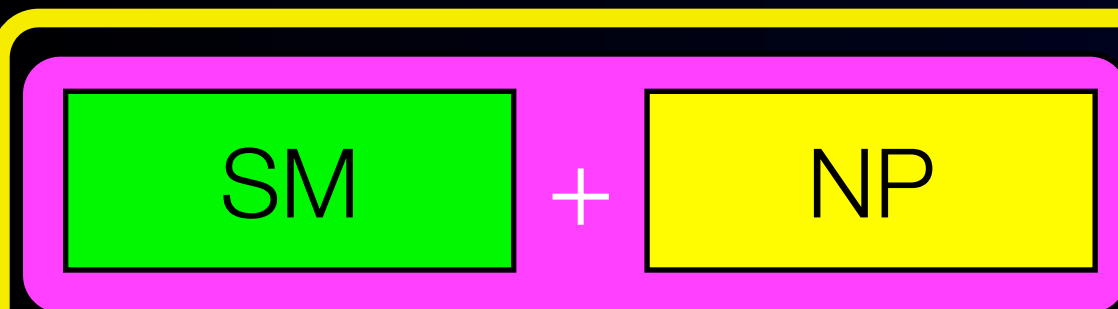
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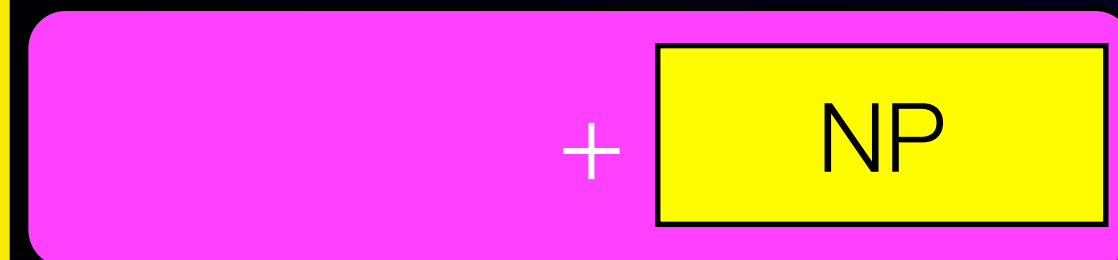


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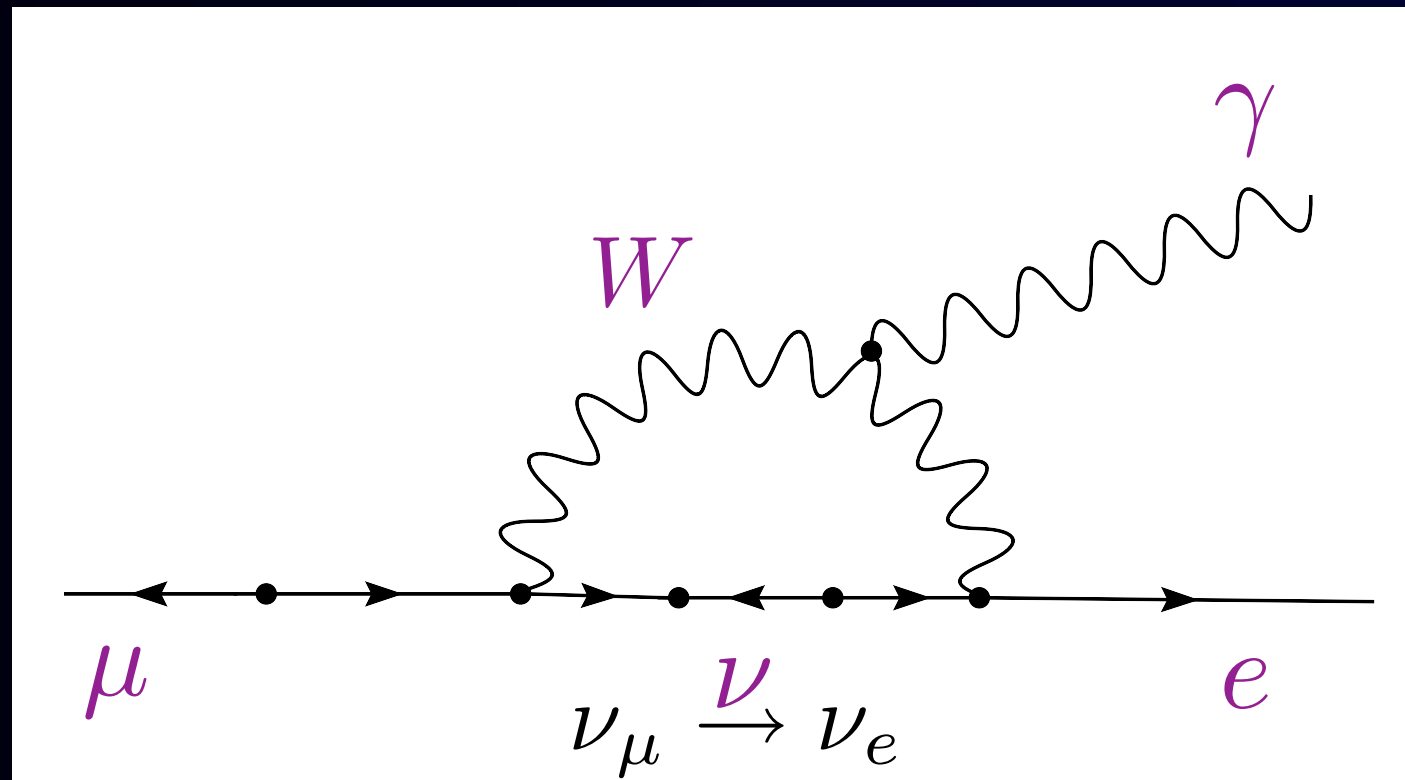
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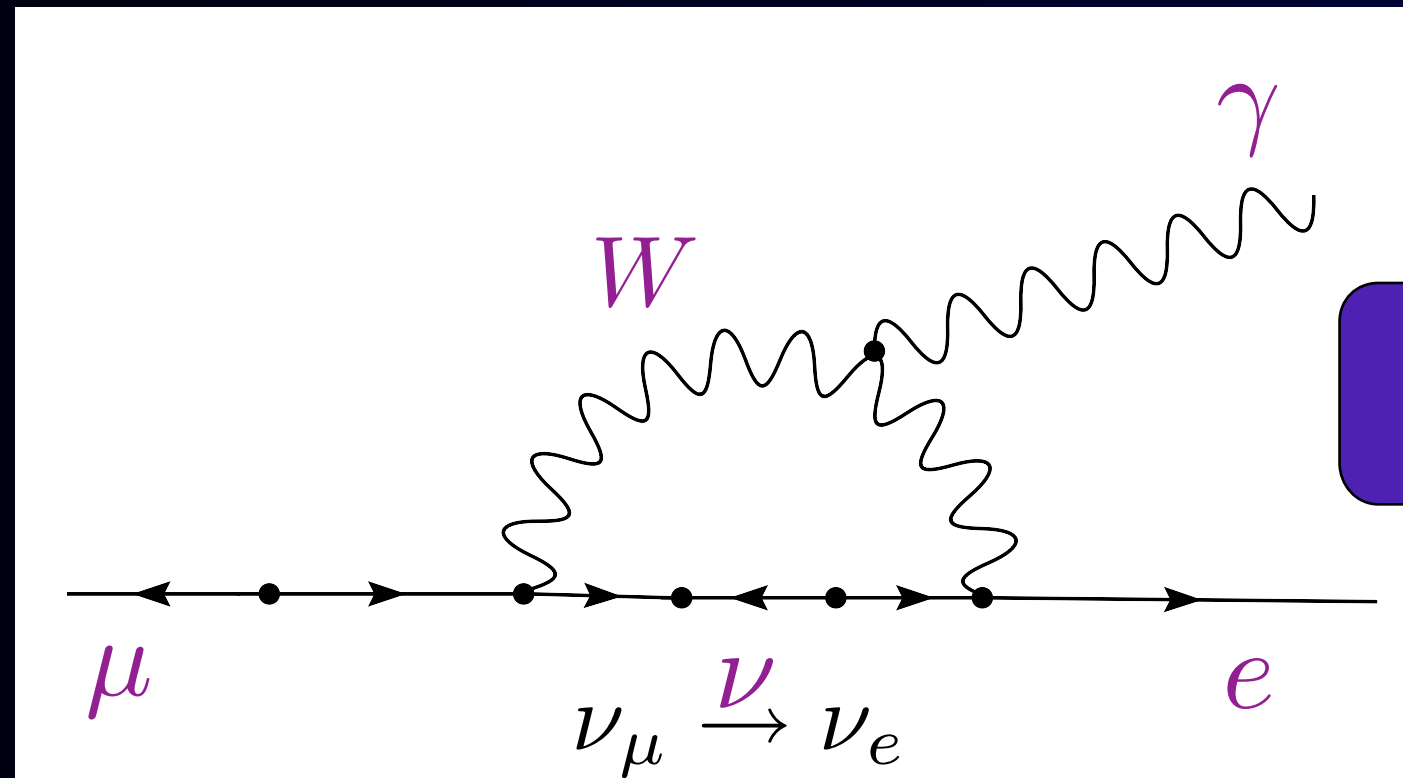
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$$B(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_l (V_{MNS})_{\mu l}^* (V_{MNS})_{el} \frac{m_{\nu_l}^2}{M_W^2} \right|^2$$



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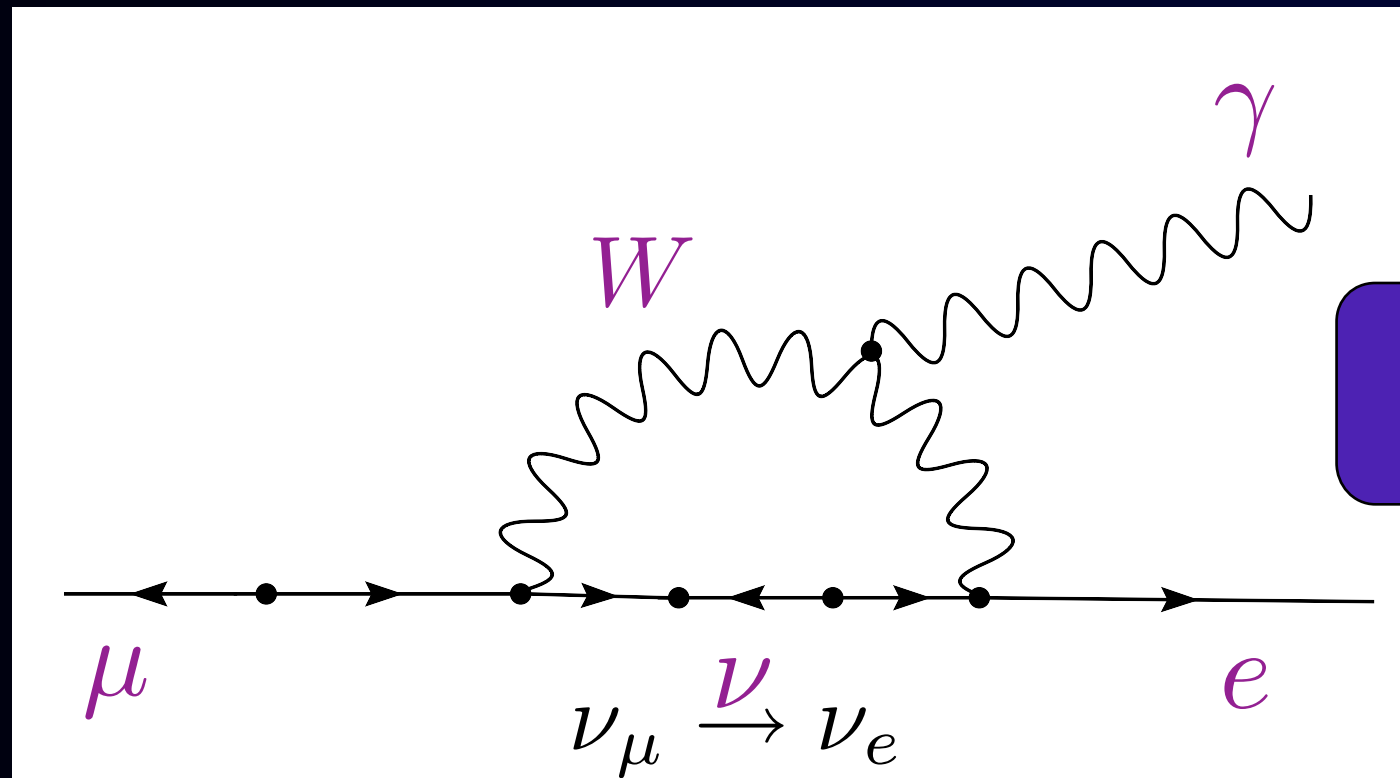
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BR $\sim O(10^{-54})$

Observation of CLFV would indicate a clear signal of physics beyond the SM with massive neutrinos.

Quark FCNC vs. Lepton FCNC (CLFV)



FCNC: The Standard Model contributions are either highly suppressed or forbidden.

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Quark (suppressed)

amplitude

$$|A_{\text{SM}} + \varepsilon_{\text{NP}}|^2 \sim |A_{\text{SM}}|^2 + \underline{2\text{Re}(A_{\text{SM}}\varepsilon_{\text{NP}})} + |\varepsilon_{\text{N}}|^2$$

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no limitation from uncertainty of SM prediction (can go to higher energy scale)

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CLFV Drawback : Rate $\sim 1/\Lambda^4$, high sensitivity is required.

New Physics in CLFV



Various Models Predict CLFV

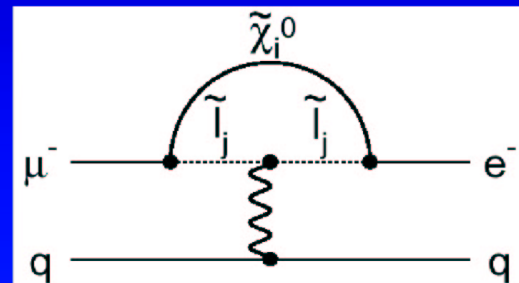


Various Models Predict CLFV

Sensitivity to Different Muon Conversion Mechanisms

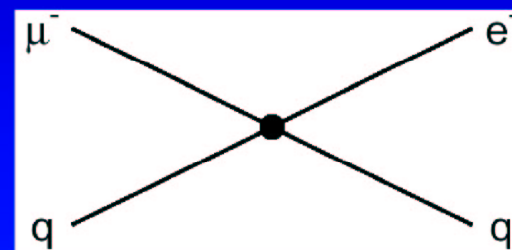


Supersymmetry
Predictions at 10^{-15}



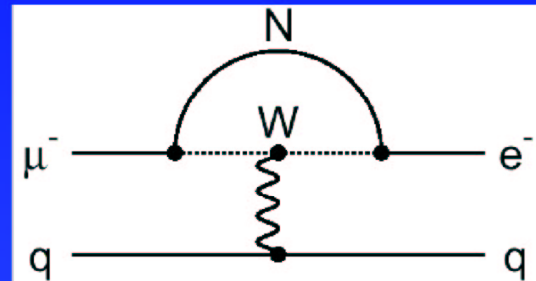
Compositeness

$$\Lambda_c = 3000 \text{ TeV}$$



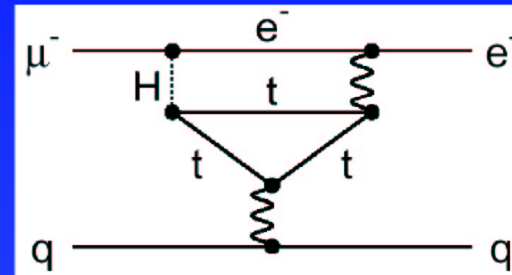
Heavy Neutrinos

$$|U_{\mu N}^* U_{eN}|^2 = 8 \times 10^{-13}$$



Second Higgs doublet

$$g_{H\mu e} = 10^{-4} \times g_{H\mu\mu}$$

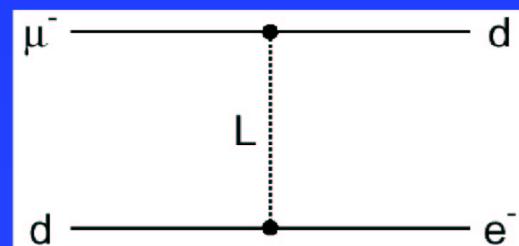


Leptoquarks

$$M_L =$$

$$3000 (\lambda_{\mu d} \lambda_{e d})^{1/2} \text{ TeV}/c^2$$

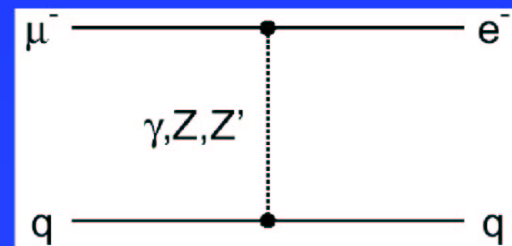
After W. Marciano



Heavy Z' ,
Anomalous Z
coupling

$$M_{Z'} = 3000 \text{ TeV}/c^2$$

$$B(Z \rightarrow \mu e) < 10^{-17}$$



Example of Sensitivity to NP in High Energy Scale : SUSY models



■ For loop diagrams,

$$\text{BR}(\mu \rightarrow e\gamma) = 1 \times 10^{-11} \times \left(\frac{2\text{TeV}}{\Lambda} \right)^4 \left(\frac{\theta_{\mu e}}{10^{-2}} \right)^2 \quad y = \frac{g^2}{16\pi^2} \theta_{\mu e}$$

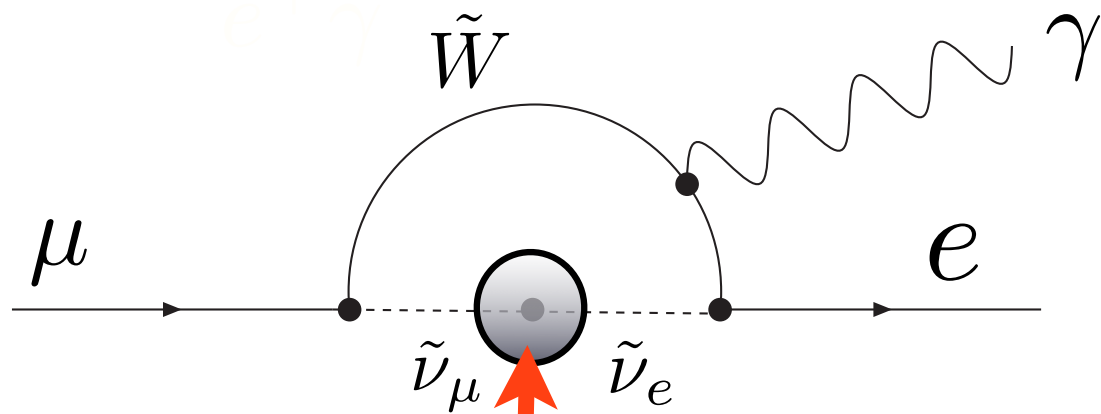
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example diagram for SUSY (~TeV)

Physics at about 10^{16} GeV

slepton mixing
(from RGE)

$$(m_{\tilde{L}}^2)_{21} \sim \frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 V_{td} V_{ts} \ln \frac{M_{GUT}}{M_{R_s}}$$

$$(m_L^2)_{21} \sim \frac{3m_0^2 + A_0^2}{8\pi^2} h_\tau^2 U_{31} U_{32} \ln \frac{M_{GUT}}{M_R}$$

SUSY-GUT model

SUSY neutrino
seesaw model

CLFV Predictions

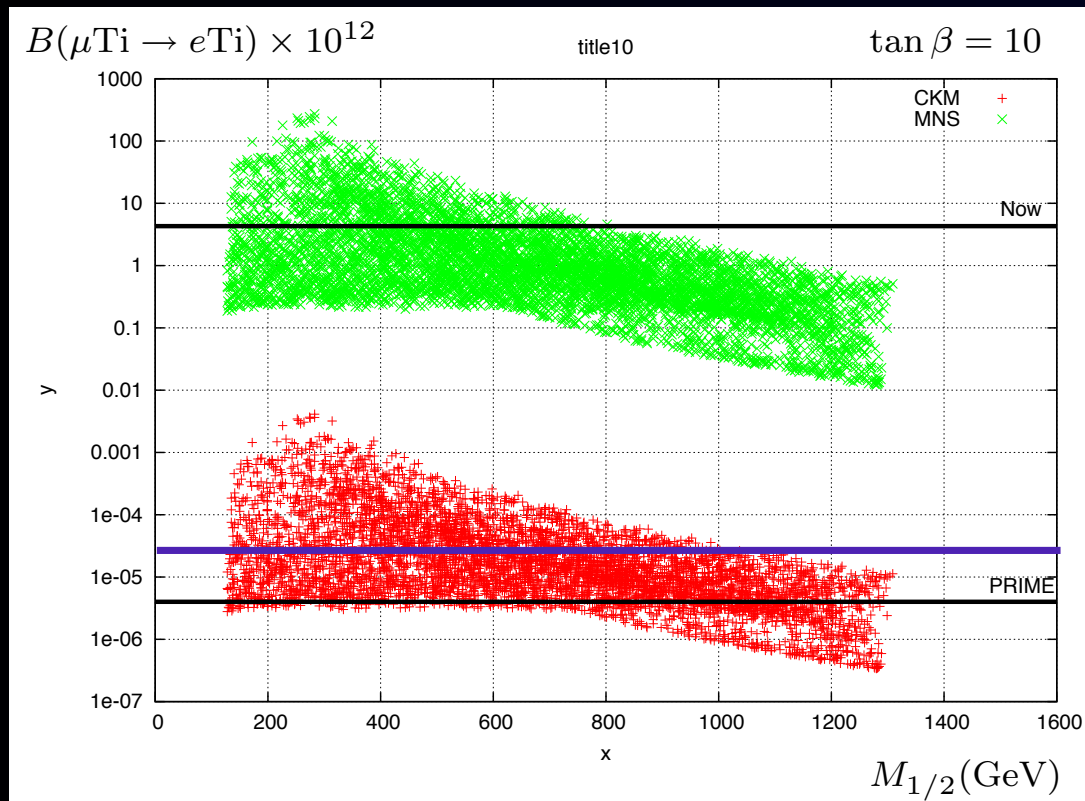


Various BSM models
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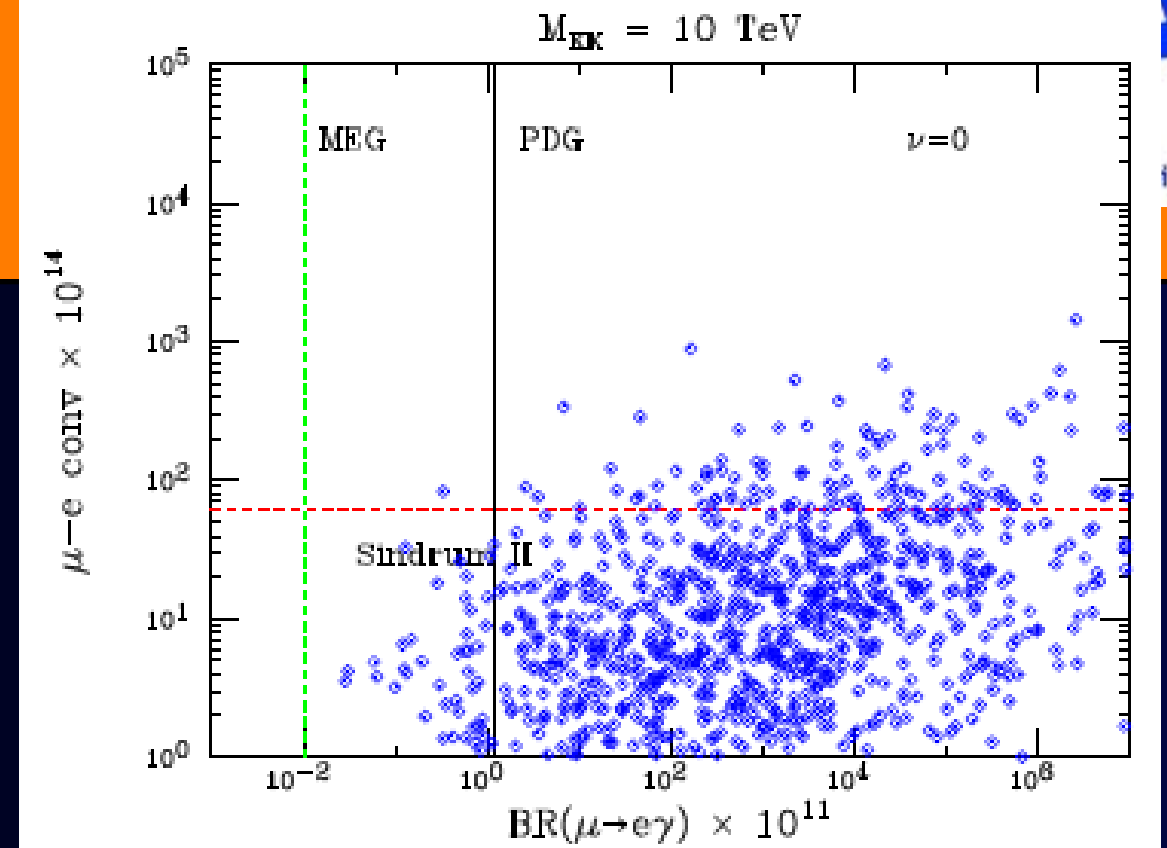
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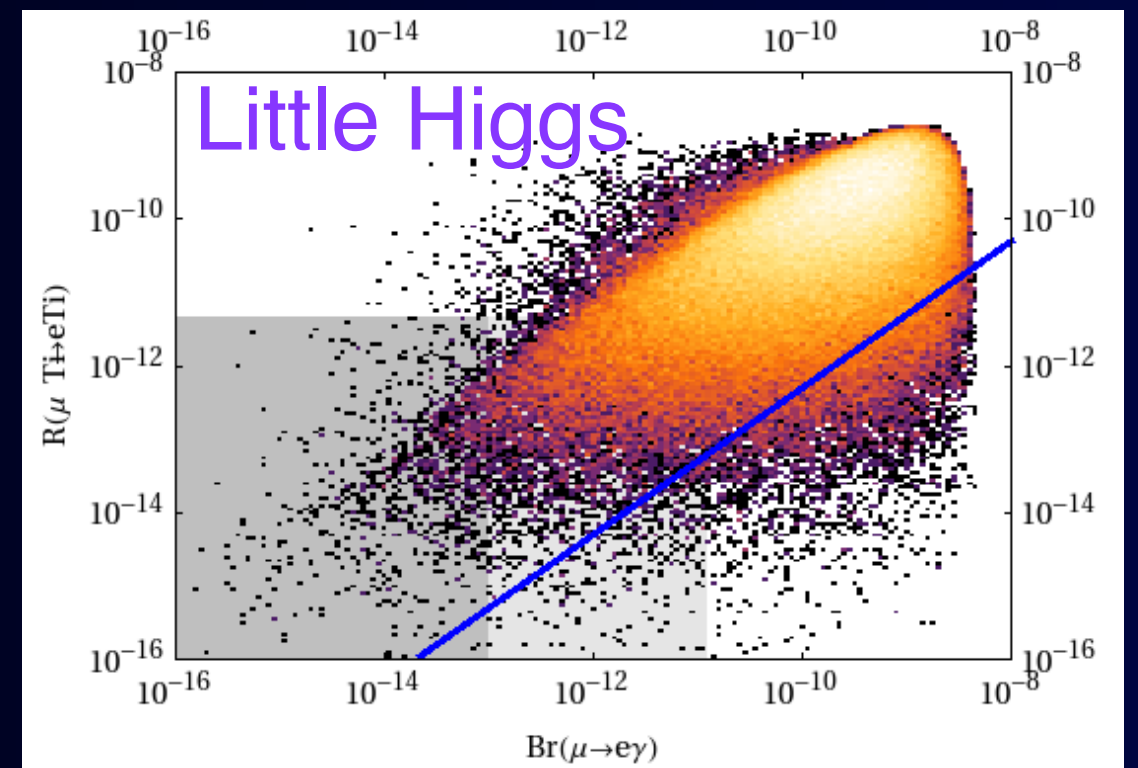
SUSY model



extra dimension model



little Higgs model



“DNA of New Physics” (a la Prof. Dr. A.J. Buras)



W. Altmannshofer, A.J. Buras, S. Gori, P. Paradisi and D.M. Straub

	AC	RVV2	AKM	δ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★★	★	★	★	★	★★★★	?
ϵ_K	★	★★★★	★★★★	★	★	★★	★★★★
$S_{\psi\phi}$	★★★★	★★★★	★★★★	★	★	★★★★	★★★★
$S_{\phi K_S}$	★★★★	★★	★	★★★★	★★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★★	★★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★★	★★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★★	★★★★	★★★★	★★★★	★★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$\mu \rightarrow e \gamma$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
$\tau \rightarrow \mu \gamma$	★★★★	★★★★	★	★★★★	★★★★	★★★★	★★★★
$\mu + N \rightarrow e + N$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
d_n	★★★★	★★★★	★★★★	★★	★★★★	★	★★★★
d_e	★★★★	★★★★	★★	★	★★★★	★	★★★★
$(g-2)_\mu$	★★★★	★★★★	★★	★★★★	★★★★	★	?

These are a subset of a subset listed by Buras and Girschbach
MFV, CMFV, 2HDM_{MFV}, LHT, SM4, SUSY flavor, SO(10) – GUT,
SSU(5)_{HN}, FBMSSM, RHMfV, L-R, RS₀, gauge flavor,

The pattern of measurement:
 ★ ★ ★ large effects
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 is characteristic,
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GLOSSARY

AC [10]

RH currents & U(1) flavor symmetry

RVV2 [11]

SU(3)-flavored MSSM

AKM [12]

RH currents & SU(3) family symmetry

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CKM-like currents

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Little Higgs with T Parity

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in More Details.



in More Details.



see Frank Deppisch' talk

ex. CLFV and neutrino mass generation
and more

CLFV Experiments



Quarks, Neutrinos, and Charged Leptons



Quarks, Neutrinos, and Charged Leptons

Quarks



Quark mixing
observed

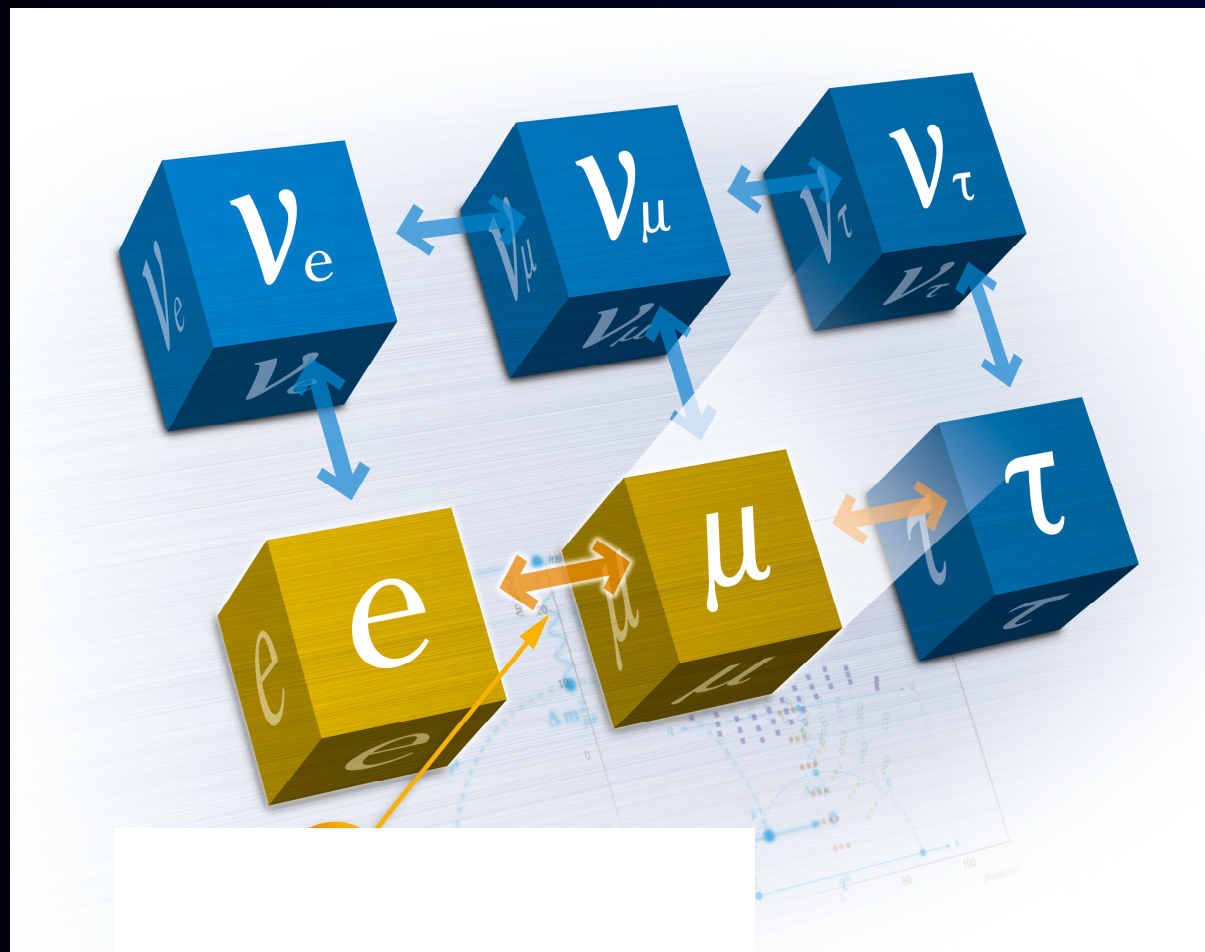
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Leptons



Neutrino
mixing
observed

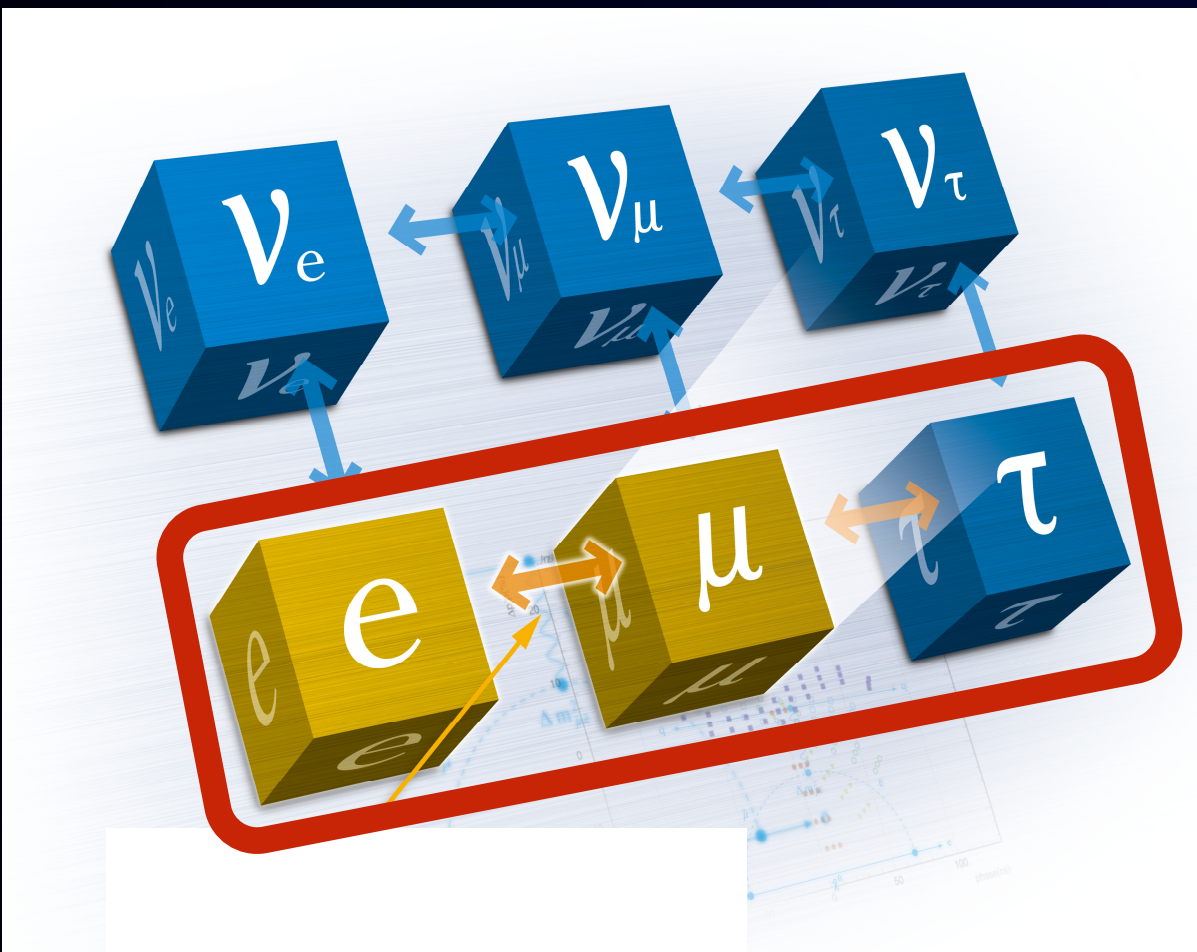
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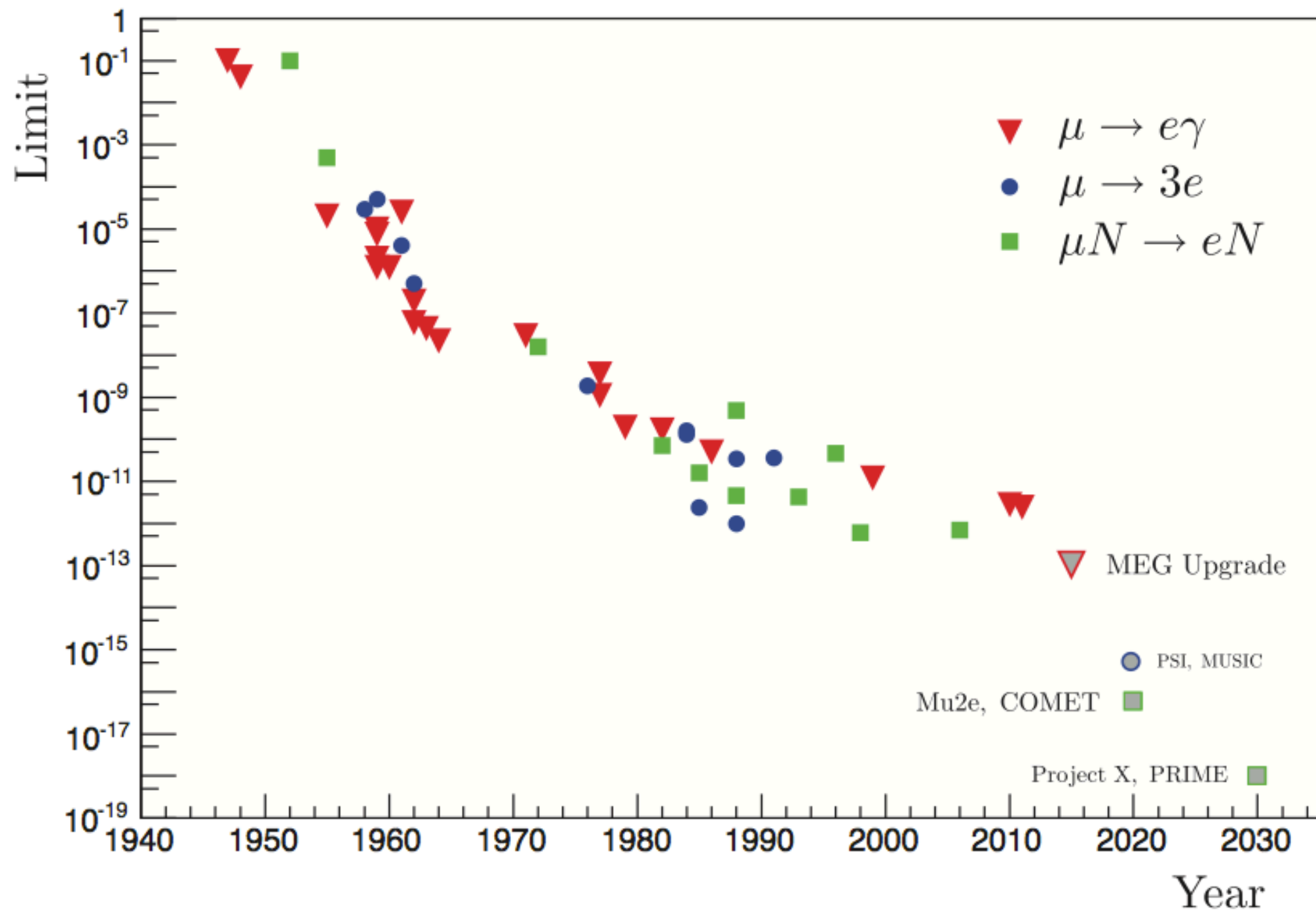
Leptons



Neutrino
mixing
observed

Charged
lepton mixing
not observed.

CLFV History



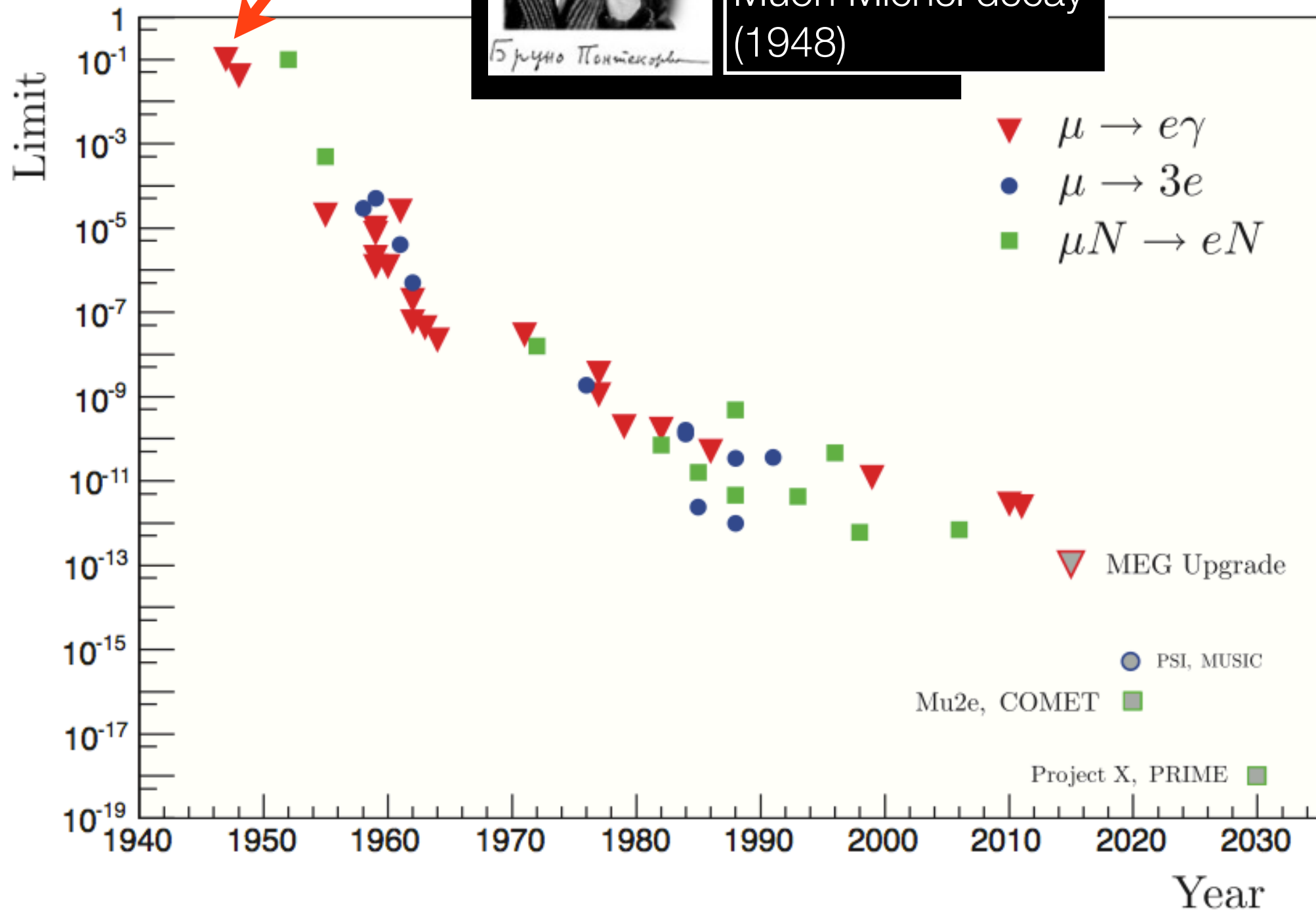
CLFV History

First CLFV search



Pontecorvo
in 1947

Muon Michel decay
(1948)



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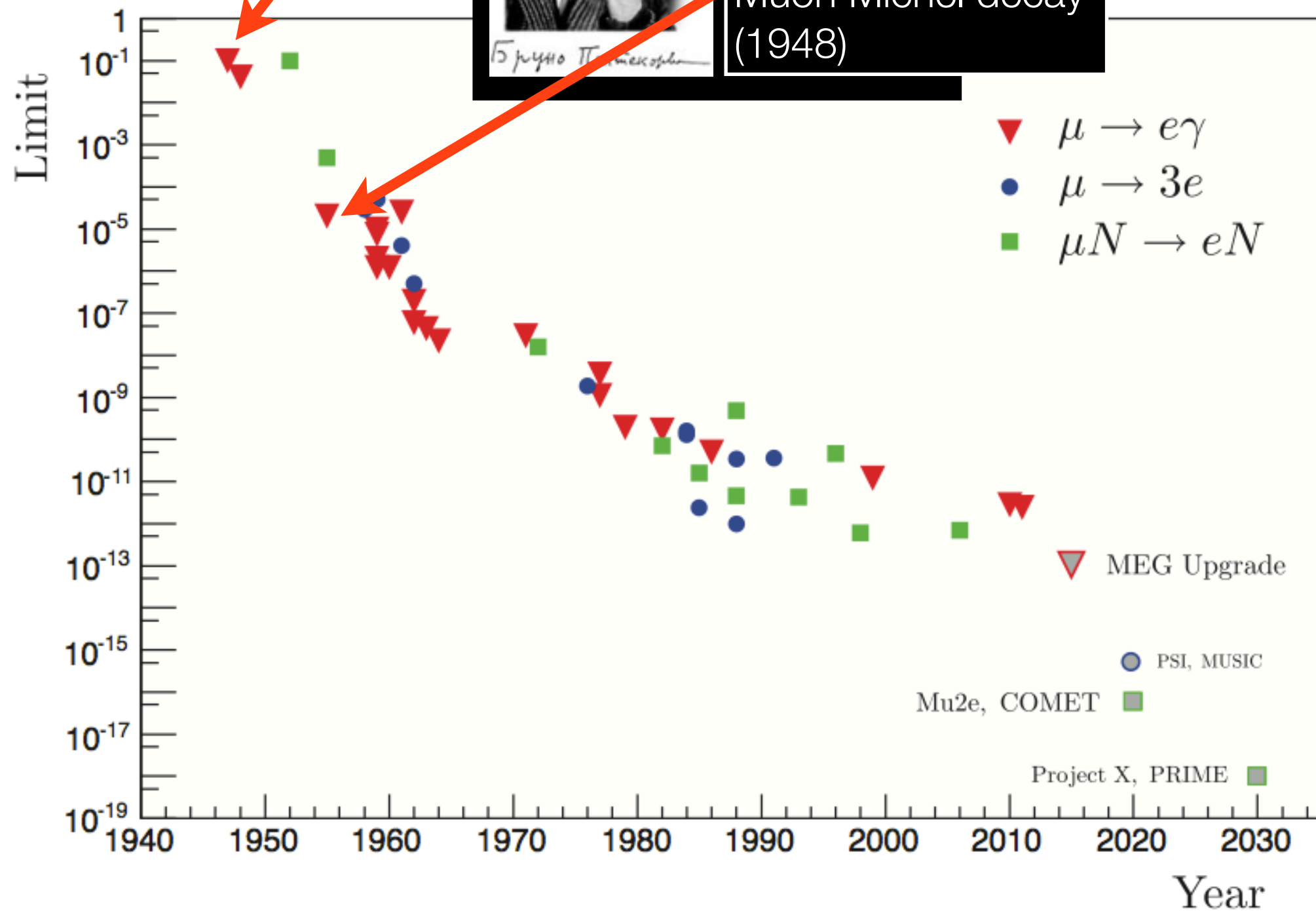


Pontecorvo
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Accelerators
producing muons

Feinberg's $\mu \rightarrow e\gamma$
crisis (1955)



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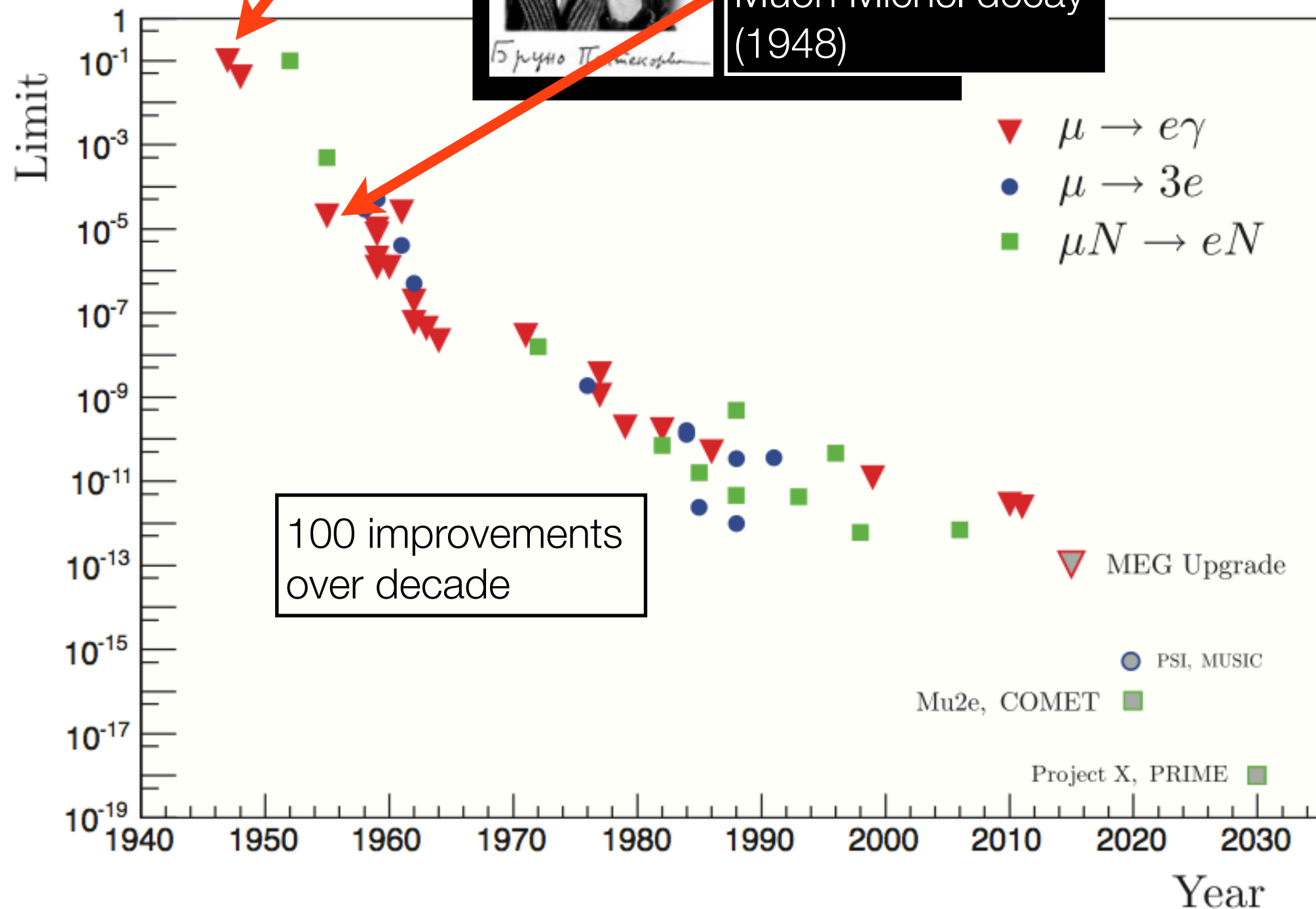


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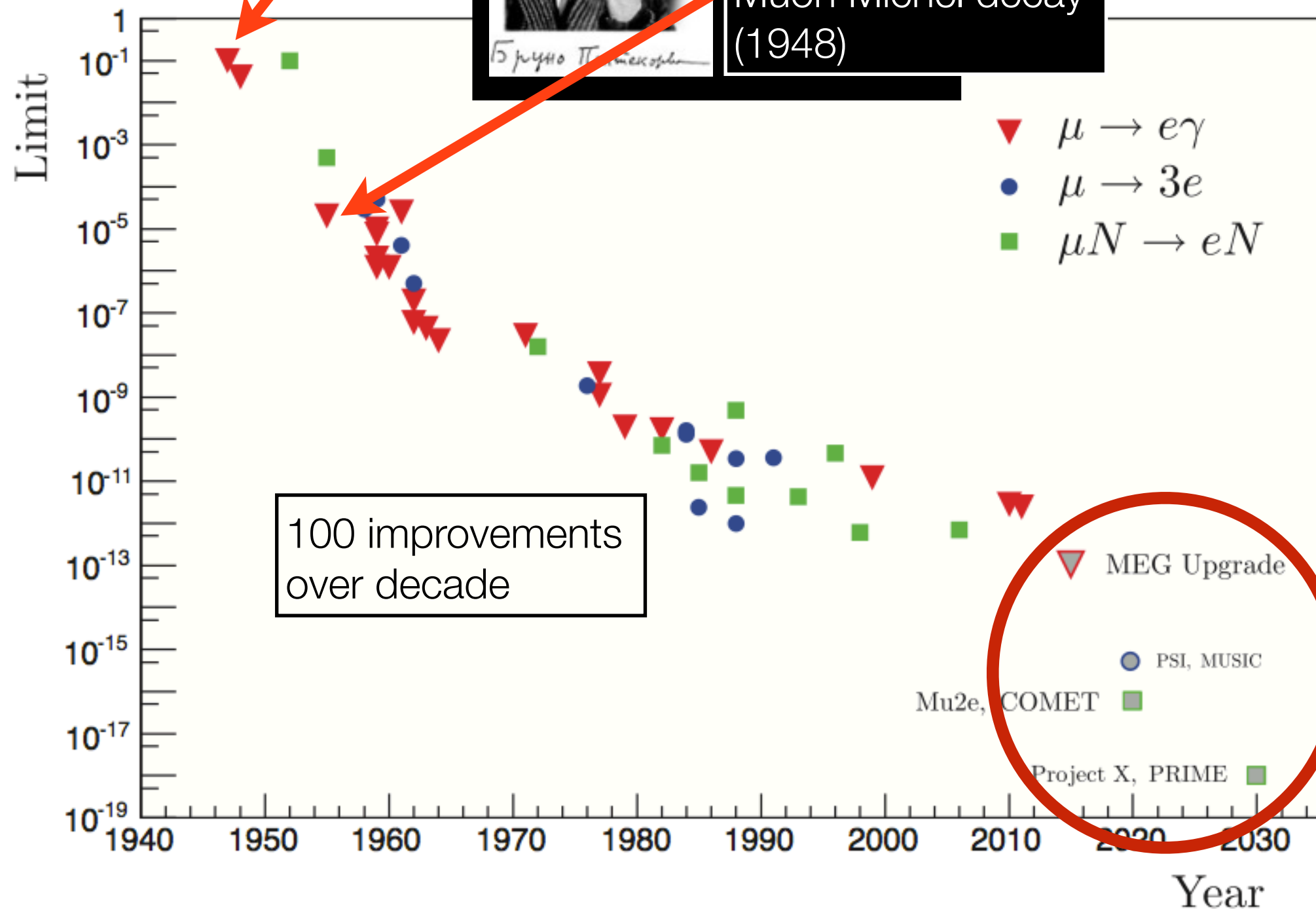


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100 improvements
over decade

Present Limits and Future Expectations

process	present limit	future	
$\mu \rightarrow e \gamma$	$< 5.7 \times 10$	< 10	MEG
$\mu \rightarrow e e e$	$< 1.0 \times 10$	< 10	Mu3e
$\mu N \rightarrow e N$ (in Al)	none	< 10	Mu2e / COMET
$\mu N \rightarrow e N$ (in Ti)	$< 4.3 \times 10$	< 10	Mu2e / COMET
$\tau \rightarrow e \gamma$	$< 3.3 \times 10$	< 10	super KEKB
$\tau \rightarrow e e e$	$< 3.4 \times 10$	< 10	super KEKB
$\tau \rightarrow \mu \gamma$	$< 4.4 \times 10$	< 10	super KEKB
$\tau \rightarrow \mu \mu \mu$	$< 2.1 \times 10$	< 10	super KEKB/LHCb

CLFV in Tau Decays



CLFV in Tau Decays





Future Prospects on CLFV in Tau Decays



Future Prospects on CLFV in Tau Decays



Super KEKB
(50 ab^{-1})



LHCb upgrade

Future Prospects on CLFV in Tau Decays



Super KEKB
(50 ab^{-1})



LHCb upgrade

- $\tau \rightarrow l \nu$ at level $(0.2-1) \times 10^{-9}$ by Belle-II
- $\tau \rightarrow \mu \mu \mu$ at the level of $(0.1-8) \times 10^{-9}$ by LHCb with 50 fb^{-1} .

CLFV in Muon Decays



List of cLFV Processes with Muons

$\Delta L=1$

- $\mu^+ \rightarrow e^+ \gamma$
- $\mu^+ \rightarrow e^+ e^+ e^-$
- $\mu^- + N(A, Z) \rightarrow e^- + N(A, Z)$
- $\mu^- + N(A, Z) \rightarrow e^+ + N(A, Z - 2)$

$\Delta L=2$

- $\mu^+ e^- \rightarrow \mu^- e^+$
- $\mu^- + N(A, Z) \rightarrow \mu^+ + N(A, Z - 2)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ + N(A, Z - 1)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ \mu^+ \mu^- + N(A, Z - 1)$

List of cLFV Processes with Muons

$\Delta L=1$

- $\mu^+ \rightarrow e^+ \gamma$
- $\mu^+ \rightarrow e^+ e^+ e^-$
- $\mu^- + N(A, Z) \rightarrow e^- + N(A, Z)$
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$\Delta L=2$

- $\mu^+ e^- \rightarrow \mu^- e^+$
- $\mu^- + N(A, Z) \rightarrow \mu^+ + N(A, Z - 2)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ + N(A, Z - 1)$
- $\nu_\mu + N(A, Z) \rightarrow \mu^+ \mu^+ \mu^- + N(A, Z - 1)$

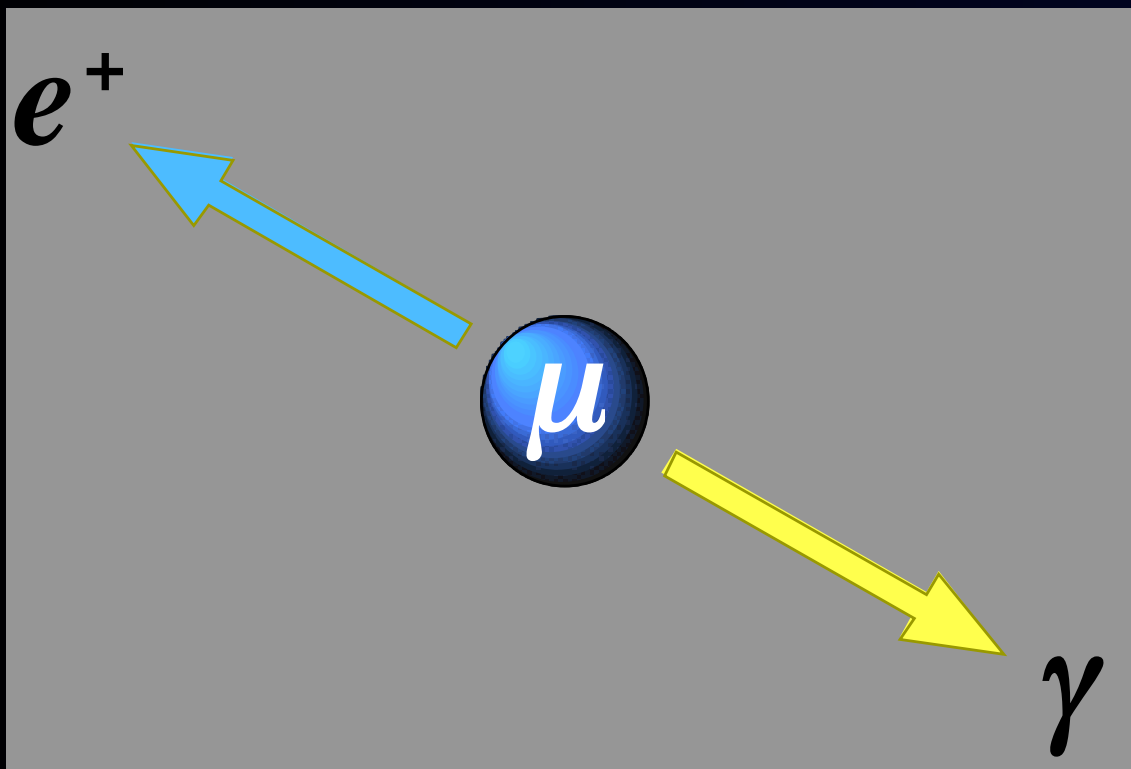
What is $\mu \rightarrow e\gamma$?

- Event Signature

- $E_e = m_\mu/2$, $E_\gamma = m_\mu/2$ (=52.8 MeV)
- angle $\theta_{\mu e}=180$ degrees (back-to-back)
- time coincidence

- Backgrounds

- prompt physics backgrounds
 - radiative muon decay $\mu \rightarrow e\nu\bar{\nu}\gamma$ when two neutrinos carry very small energies.
- accidental backgrounds
 - positron in $\mu \rightarrow e\nu\bar{\nu}$
 - photon in $\mu \rightarrow e\nu\bar{\nu}\gamma$ or photon from e^+e^- annihilation in flight.

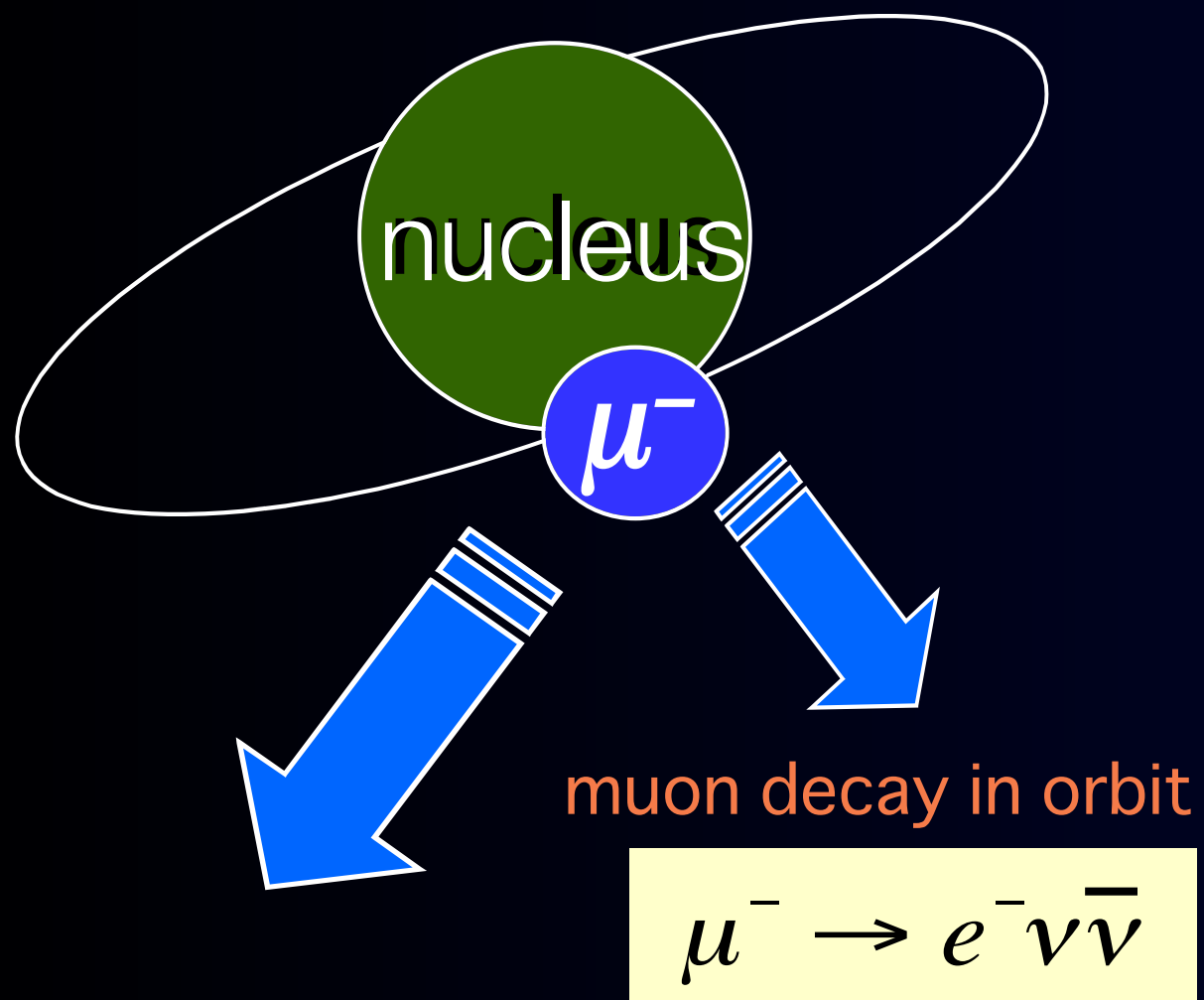


What is Muon to Electron Conversion?



What is Muon to Electron Conversion?

1s state in a muonic atom

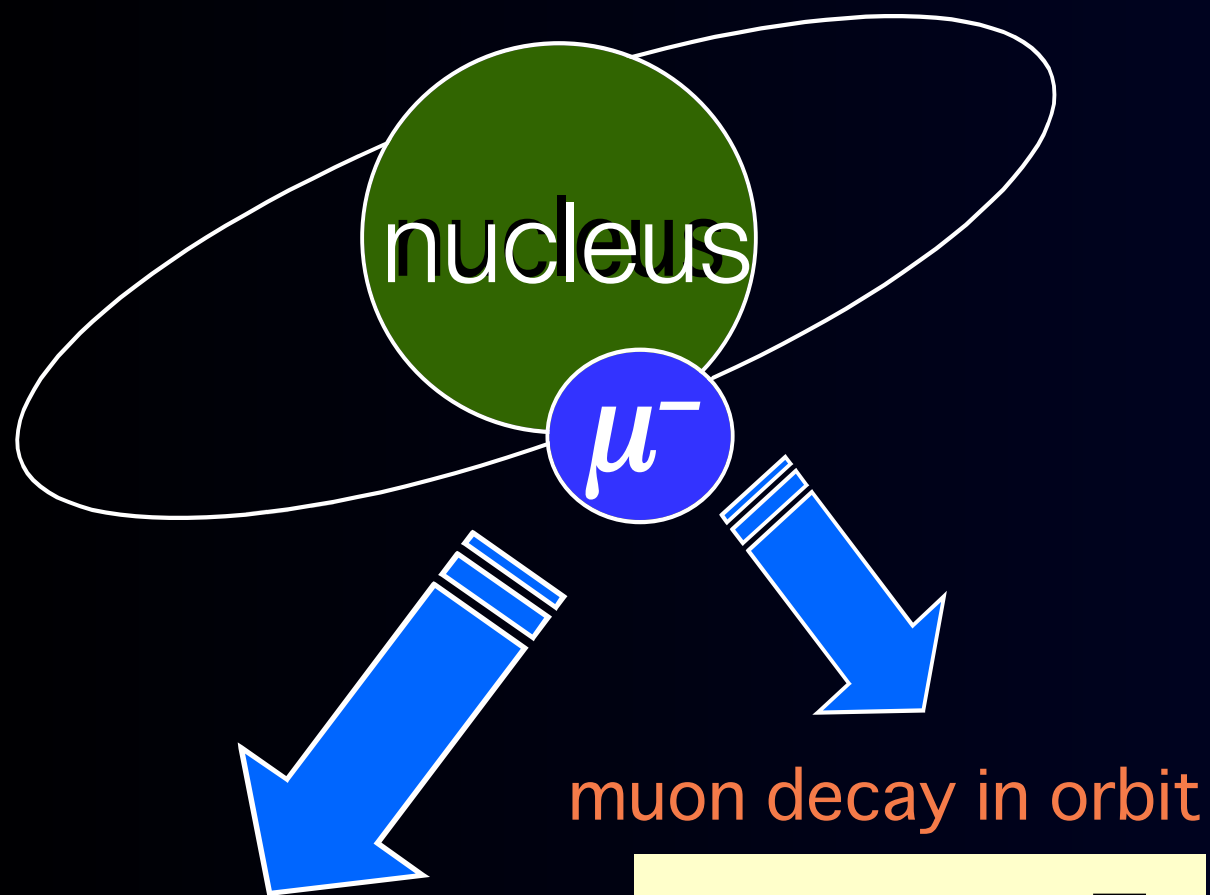


nuclear muon capture

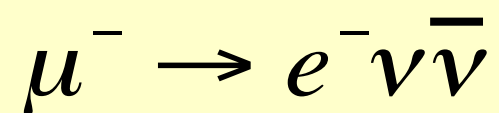
$$\mu^- + (A, Z) \rightarrow \nu_\mu + (A, Z - 1)$$

What is Muon to Electron Conversion?

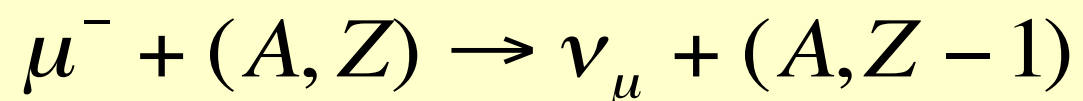
1s state in a muonic atom



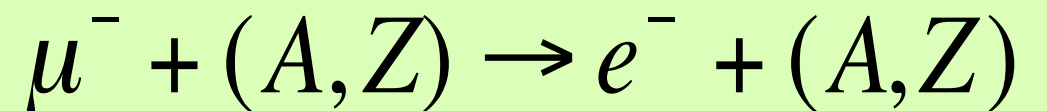
muon decay in orbit



nuclear muon capture



Neutrino-less muon
nuclear capture



Event Signature :

a single mono-energetic
electron of 100 MeV

Backgrounds:

- (1) physics backgrounds
ex. muon decay in orbit (DIO)
- (2) beam-related backgrounds
ex. radiative pion capture,
muon decay in flight,
- (3) cosmic rays, false tracking

Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. μ -e conversion



Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. μ -e conversion



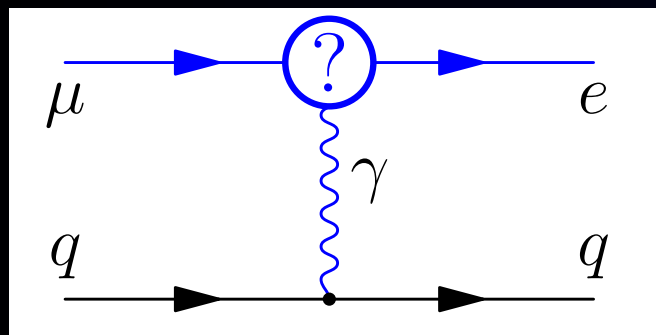
$$L_{\text{CLFV}} = \frac{1}{1 + \kappa} \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma^{\mu\nu} e_L F_{\mu\nu} + \frac{\kappa}{1 + \kappa} \frac{1}{\Lambda^2} (\bar{\mu}_L \gamma^\mu e_L) (\bar{q}_L \gamma_\mu q_L)$$

Physics Sensitivity: $\mu \rightarrow e \gamma$ vs. μ -e conversion

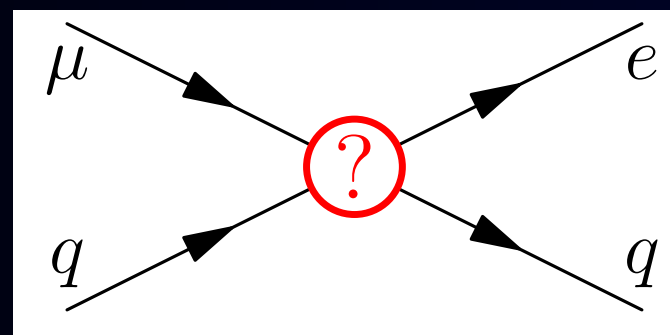


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Photonic (dipole)
interaction



Contact
interaction

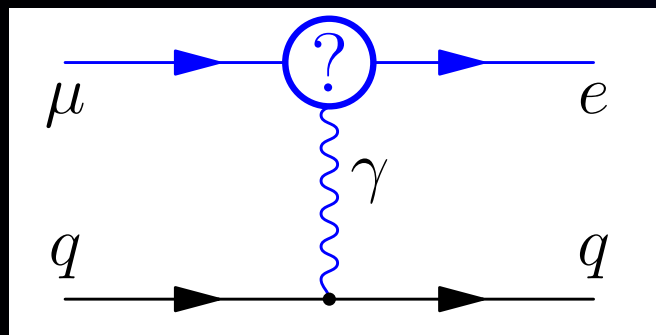


Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. μ -e conversion

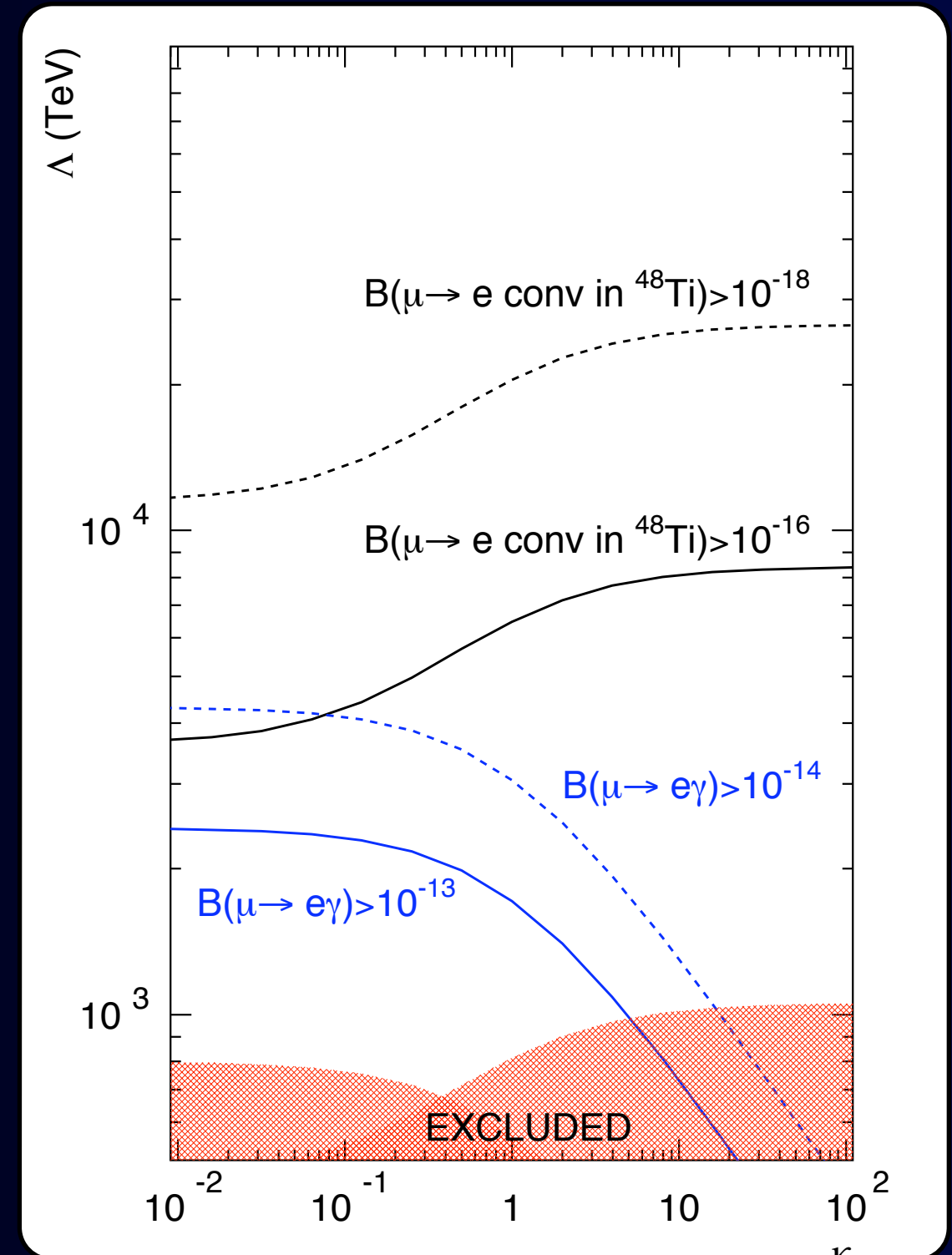
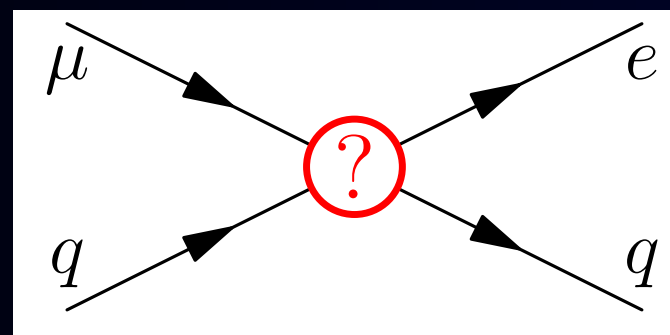


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Photonic (dipole)
interaction



Contact
interaction

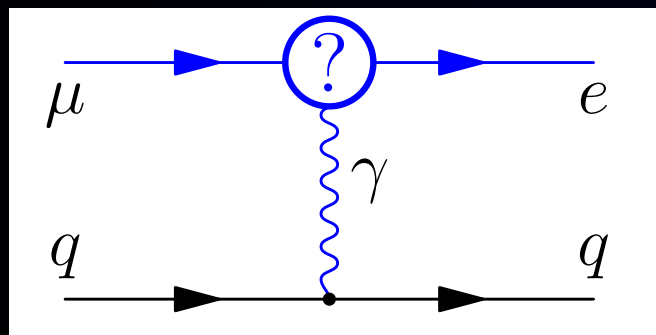


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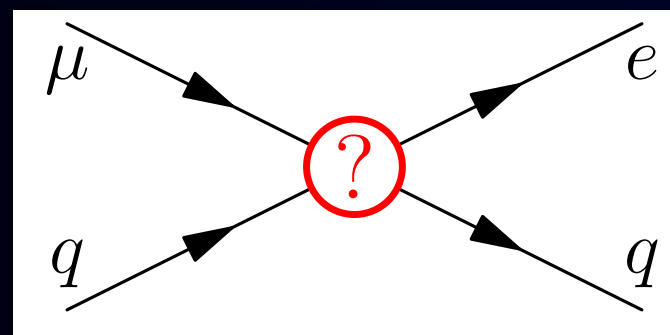


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Photonic (dipole) interaction



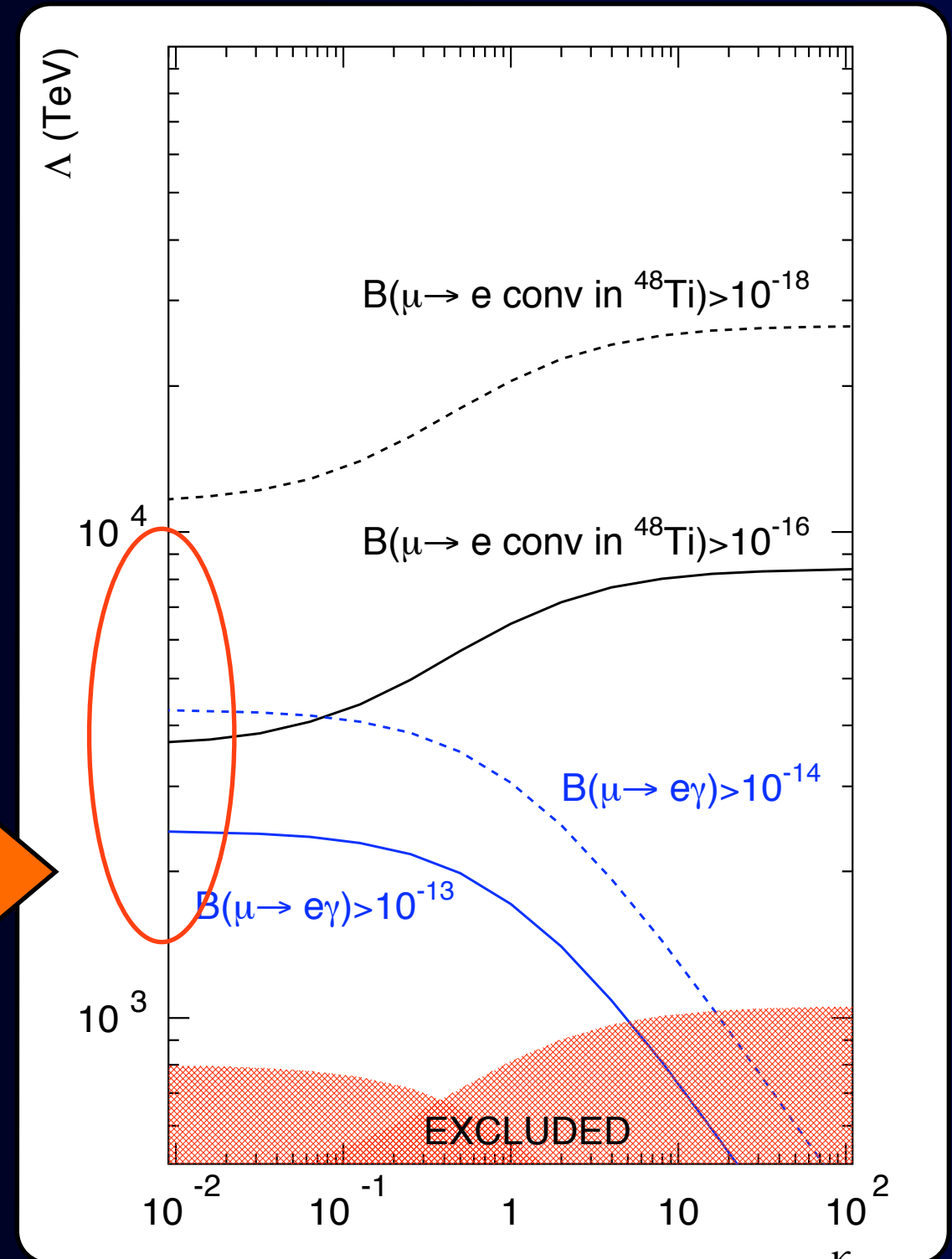
Contact interaction



if photonic contribution dominates,

$$\frac{B(\mu N \rightarrow eN)}{B(\mu \rightarrow e\gamma)} = \frac{G_F^2 m_\mu^4}{96\pi^3 \alpha} \times 3 \times 10^{12} B(A, Z) \sim \frac{B(A, Z)}{428}$$

- for aluminum, about 1/390~0.003
- for titanium, about 1/230



Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. $\mu \rightarrow eee$



Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. $\mu \rightarrow eee$



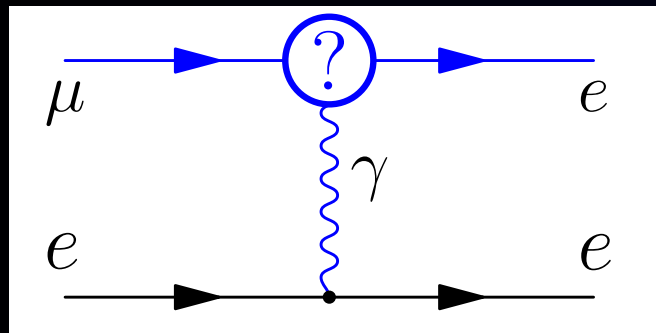
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Physics Sensitivity: $\mu \rightarrow e \gamma$ vs. $\mu \rightarrow e e e$

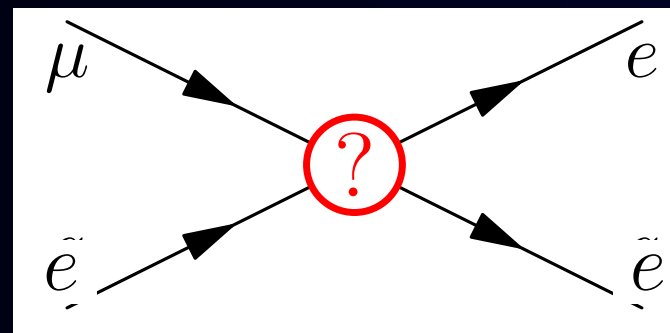


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Photonic (dipole)
interaction



Contact
interaction

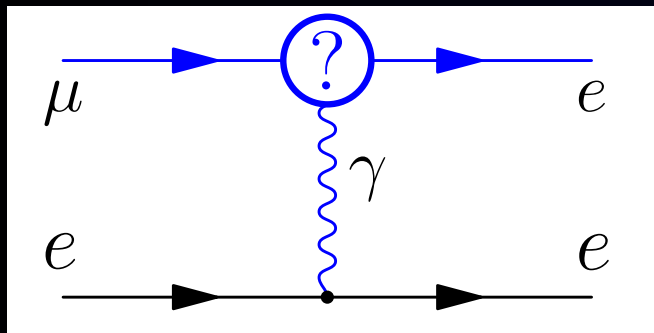


Physics Sensitivity: $\mu \rightarrow e\gamma$ vs. $\mu \rightarrow eee$

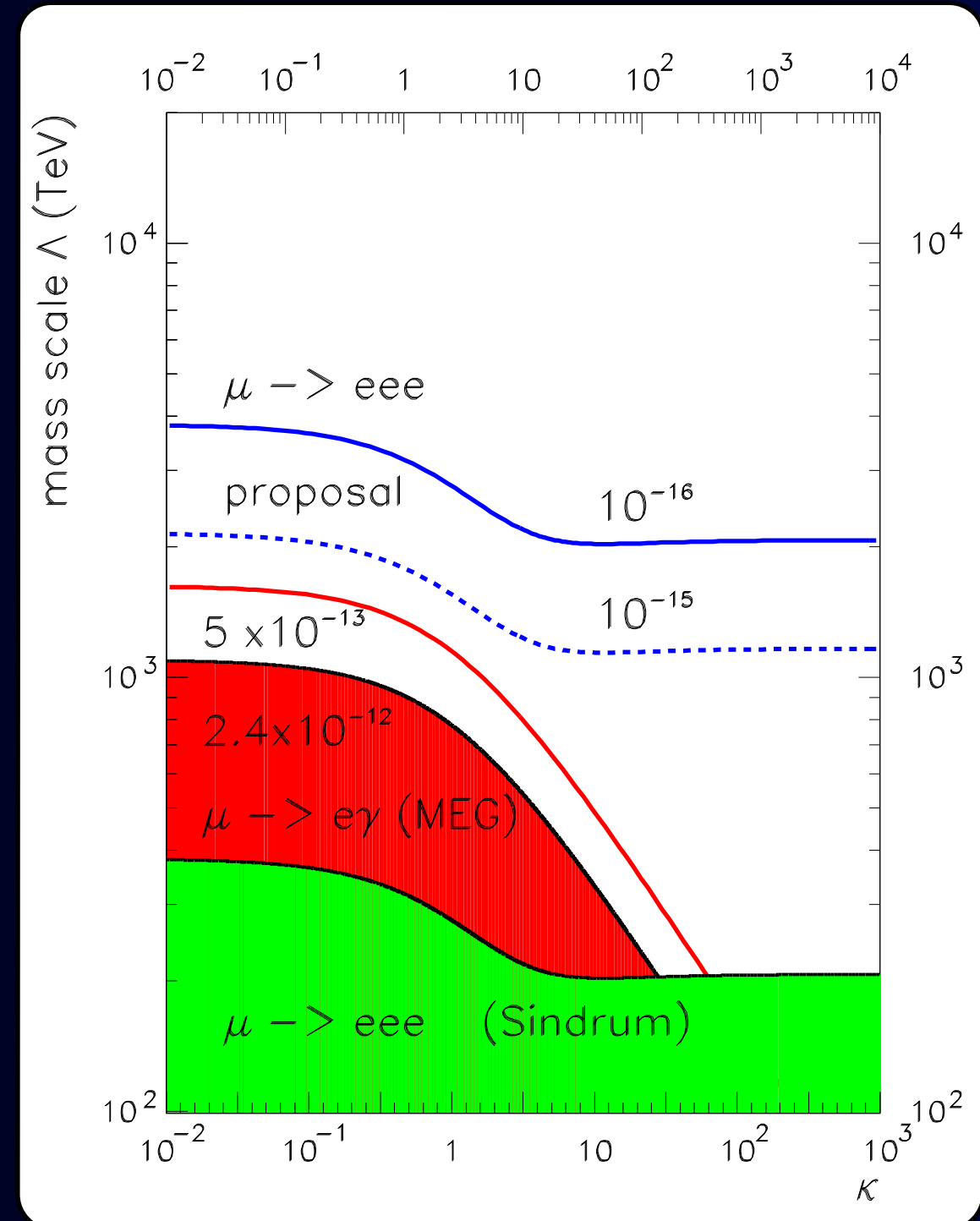
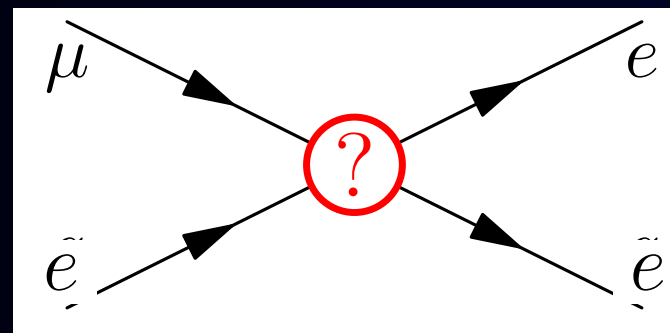


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Photonic (dipole)
interaction



Contact
interaction

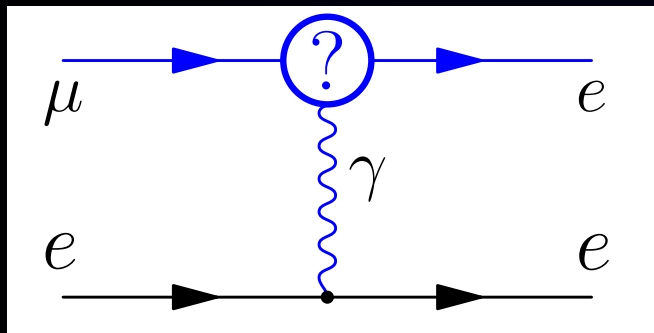


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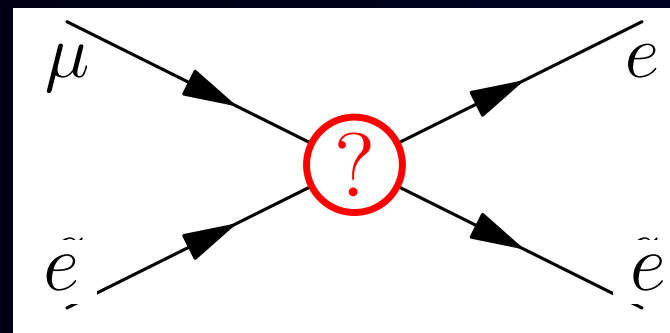


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Photonic (dipole)
interaction

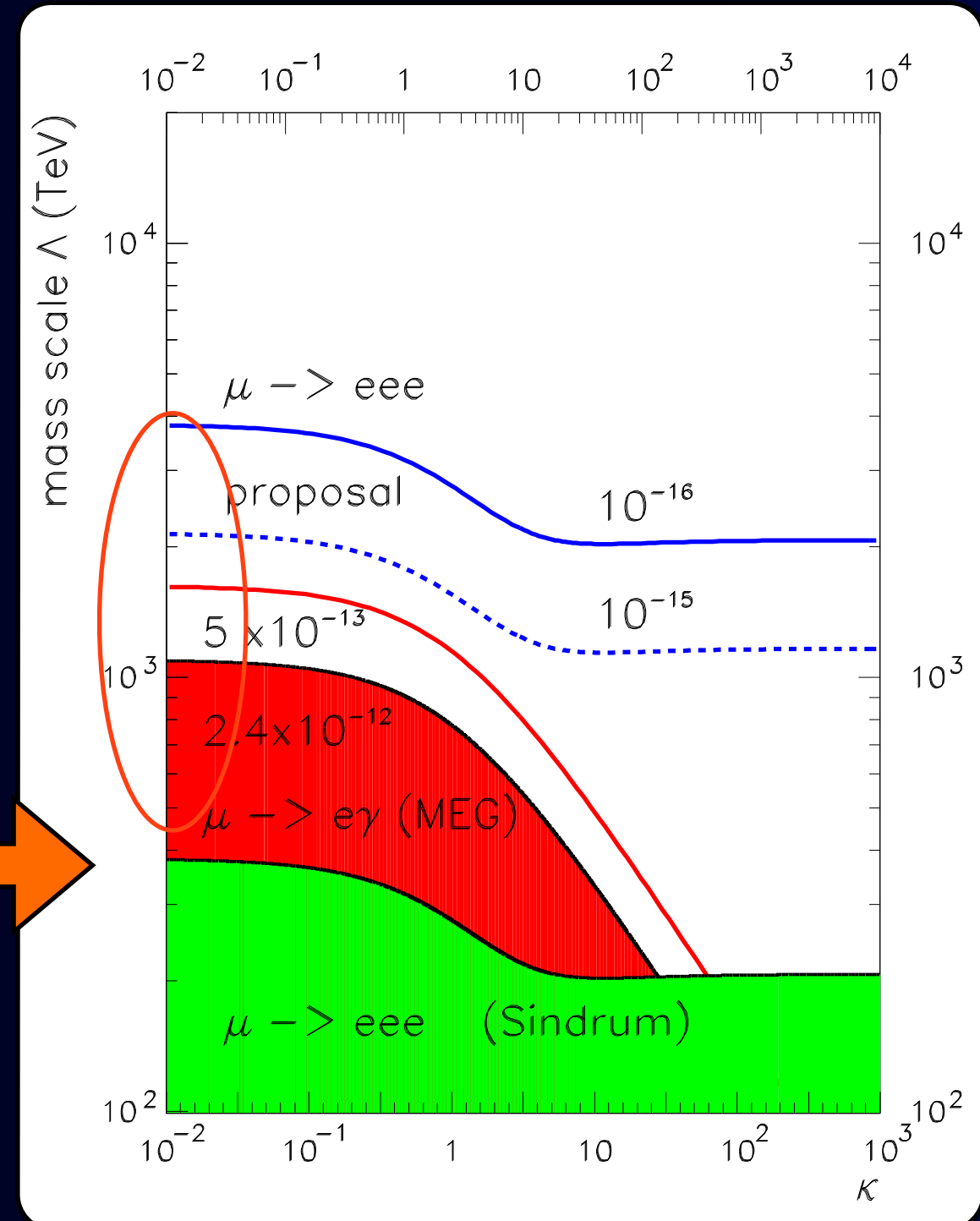


Contact
interaction



if photonic contribution dominates,

$$\frac{B(\mu \rightarrow eee)}{B(\mu \rightarrow e\gamma)} \approx 0.006$$



Experimental Comparison between $\mu \rightarrow e\gamma/\mu \rightarrow eee$ and μ -e Conversion



Experimental Comparison between $\mu \rightarrow e\gamma/\mu \rightarrow eee$ and μ -e Conversion



Process	Major backgrounds	Beam	Issues
$\mu^+ \rightarrow e^+ \gamma$	accidental	DC beam	detector resolution
$\mu^+ \rightarrow e^+ e^+ e^-$	accidental	DC beam	detector resolution
$\mu^- N \rightarrow e^- N$	beam-related	pulsed beam	beam qualities

Experimental Comparison between $\mu \rightarrow e\gamma/\mu \rightarrow eee$ and μ -e Conversion



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$\mu \rightarrow e\gamma$ and $\mu \rightarrow eee$:

Accidental background is given by $(\text{rate})^2$. The detector resolutions have to be improved.

Experimental Comparison between $\mu \rightarrow e\gamma/\mu \rightarrow eee$ and μ -e Conversion



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$\mu \rightarrow e\gamma$ and $\mu \rightarrow eee$:

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μ -e conversion:

A higher beam intensity can be taken because of no coincidence. Beam backgrounds can be under control.

CLFV Experiments in Muon Decays



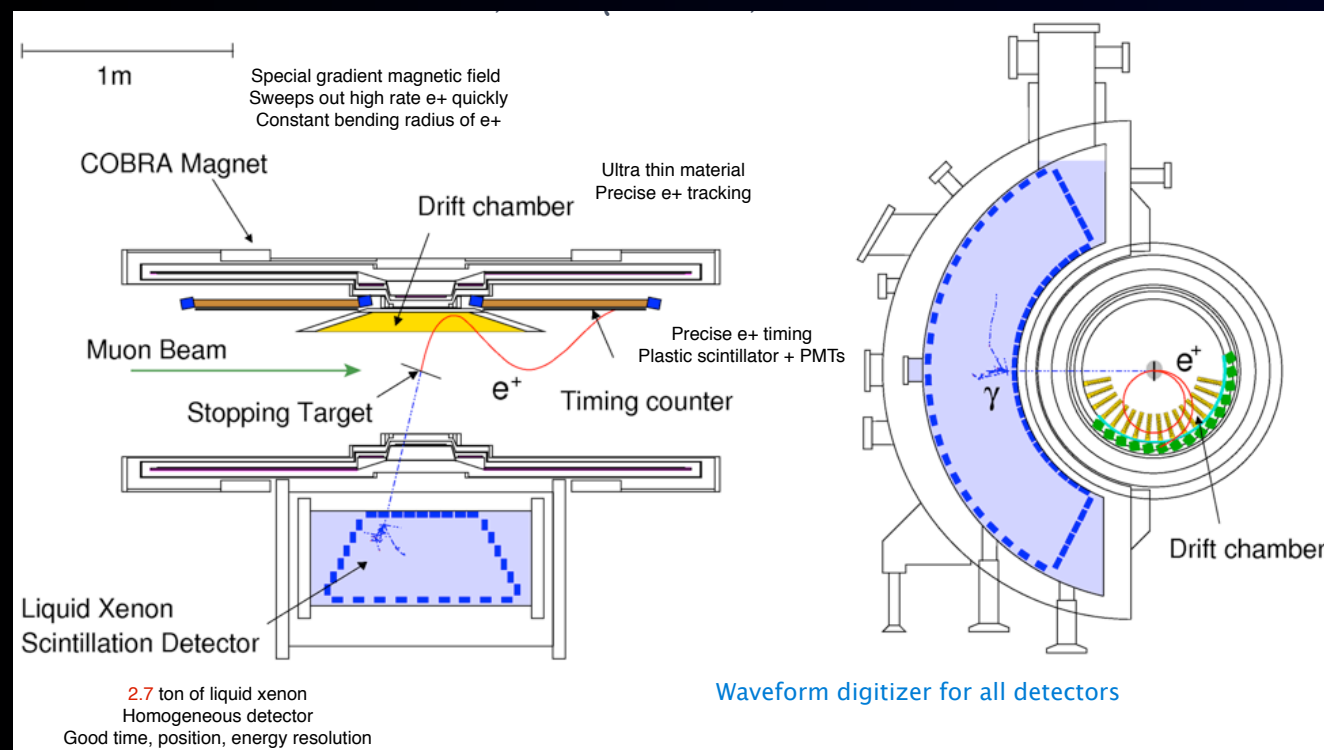
CLFV Experiments in Muon Decays

$\mu \rightarrow e\gamma$

MEG

@PSI

- Detector upgrade would include e^+ tracking and liq.Xe detector.
- The upgrade MEG will start in 2015 or 2016, aiming $O(10^{-14})$



CLFV Experiments in Muon Decays

$\mu \rightarrow e\gamma$

MEG

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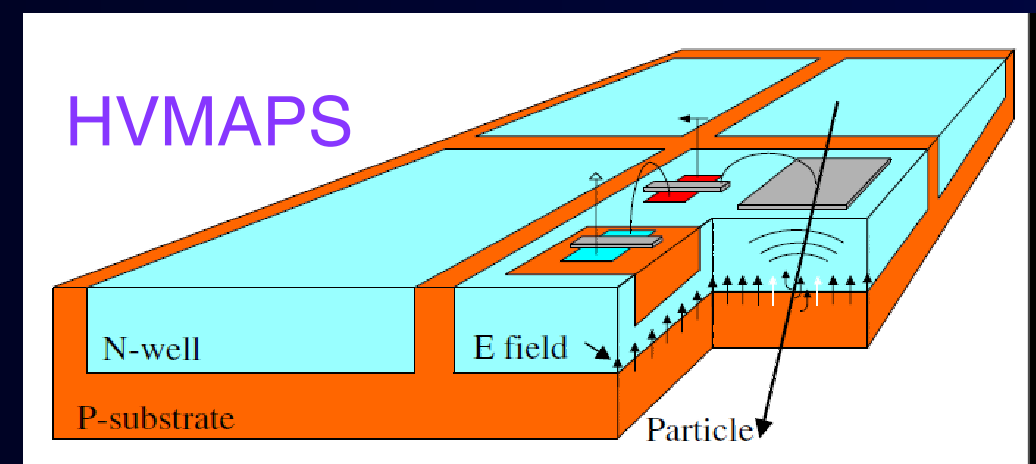
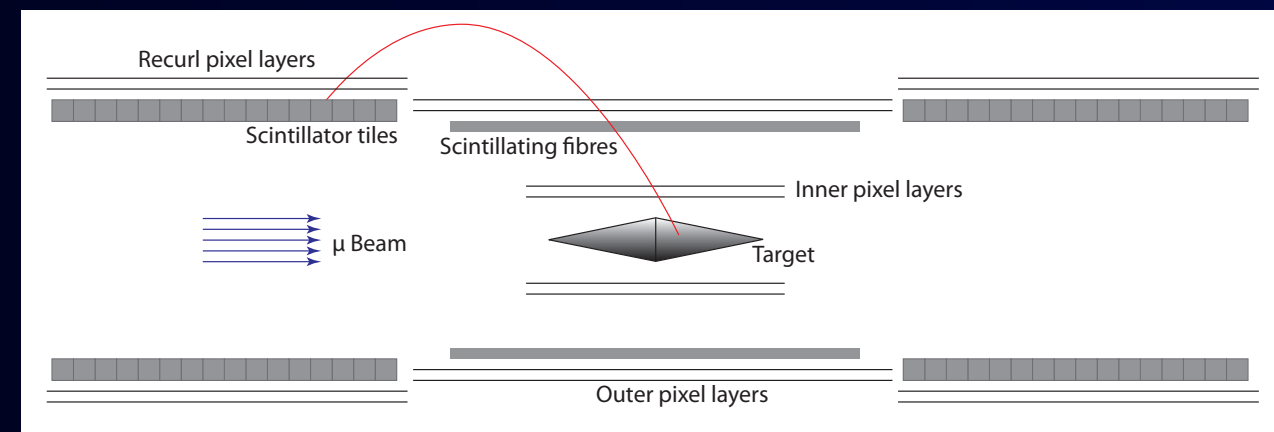
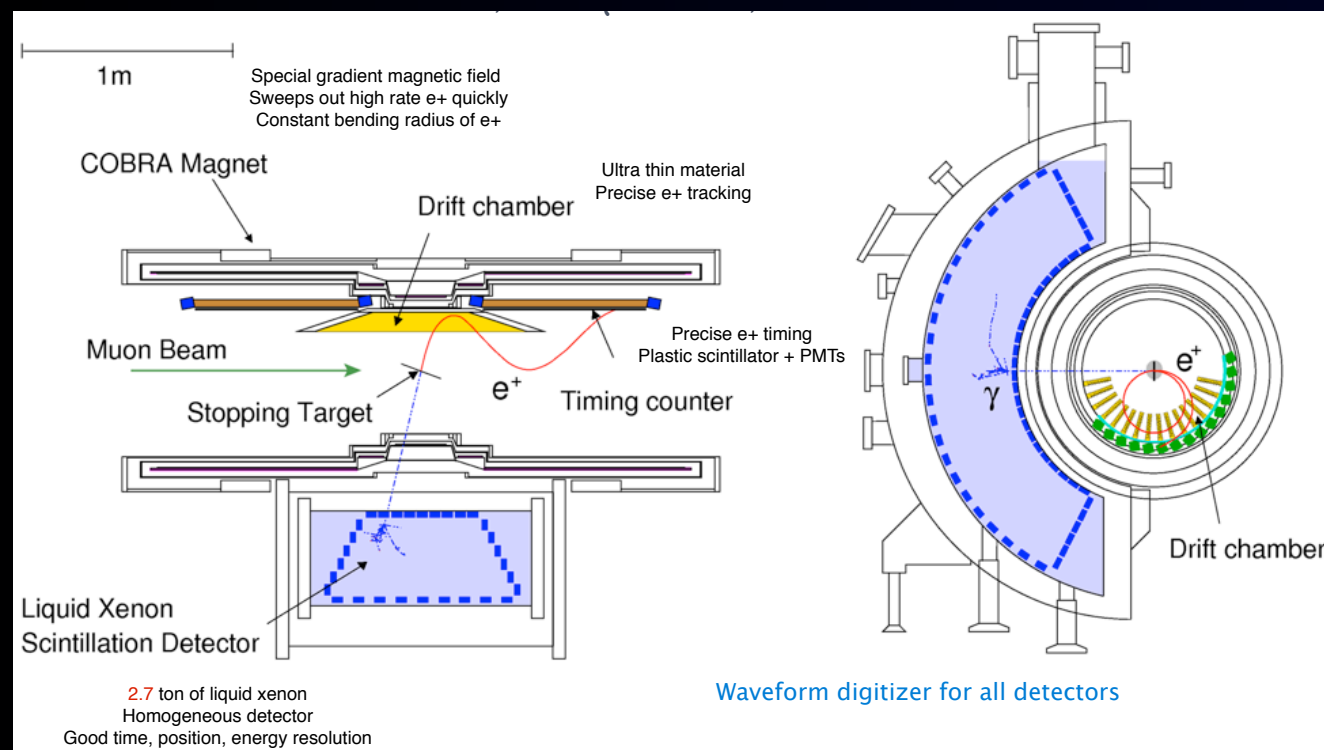
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$\mu \rightarrow eee$

Mu3e

@PSI

- search for $\mu \rightarrow eee$.
- approved at PSI last week
- staged approach, 10^{-14} in 2015, and 10^{-16} in 2017.



CLFV Experiments in Muon Decays

$\mu \rightarrow e\gamma$

MEG

@PSI

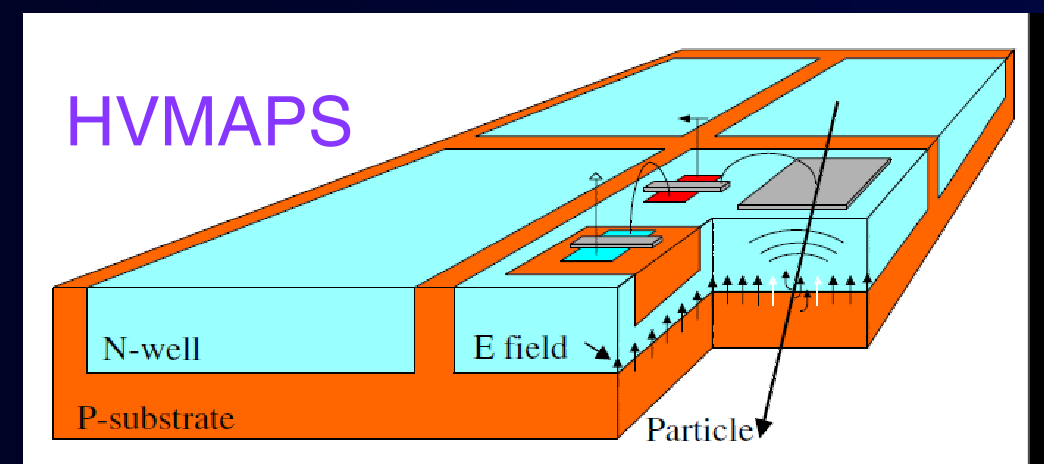
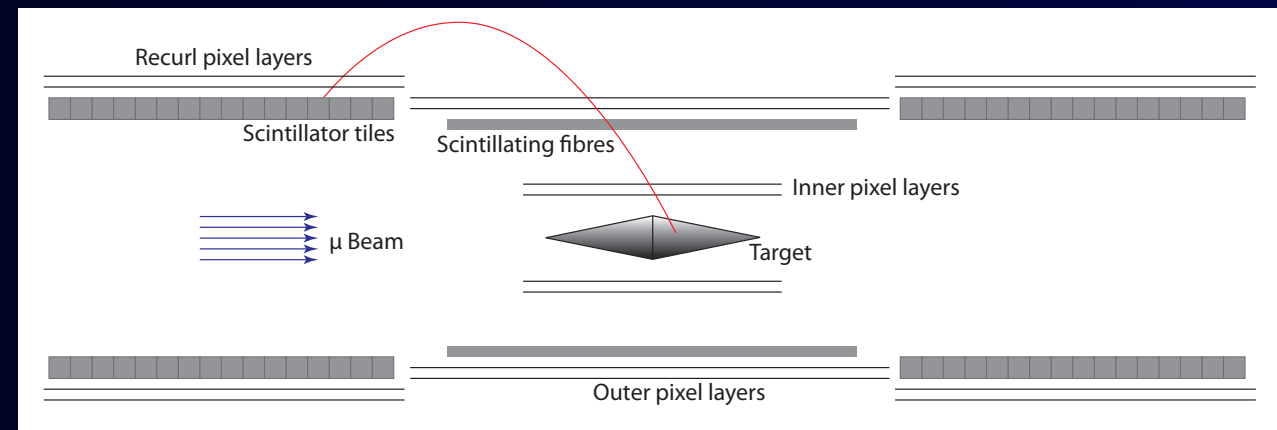
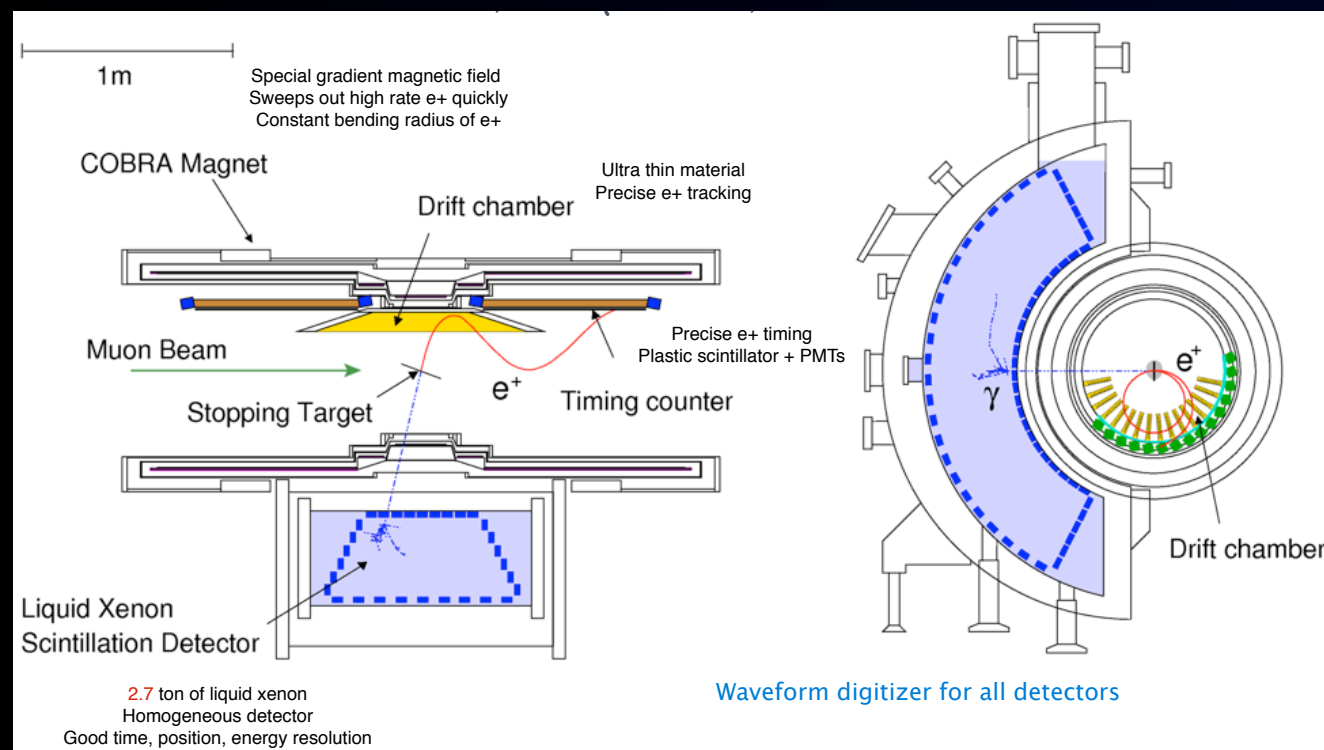
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see Dmitry Grigoriev's talk

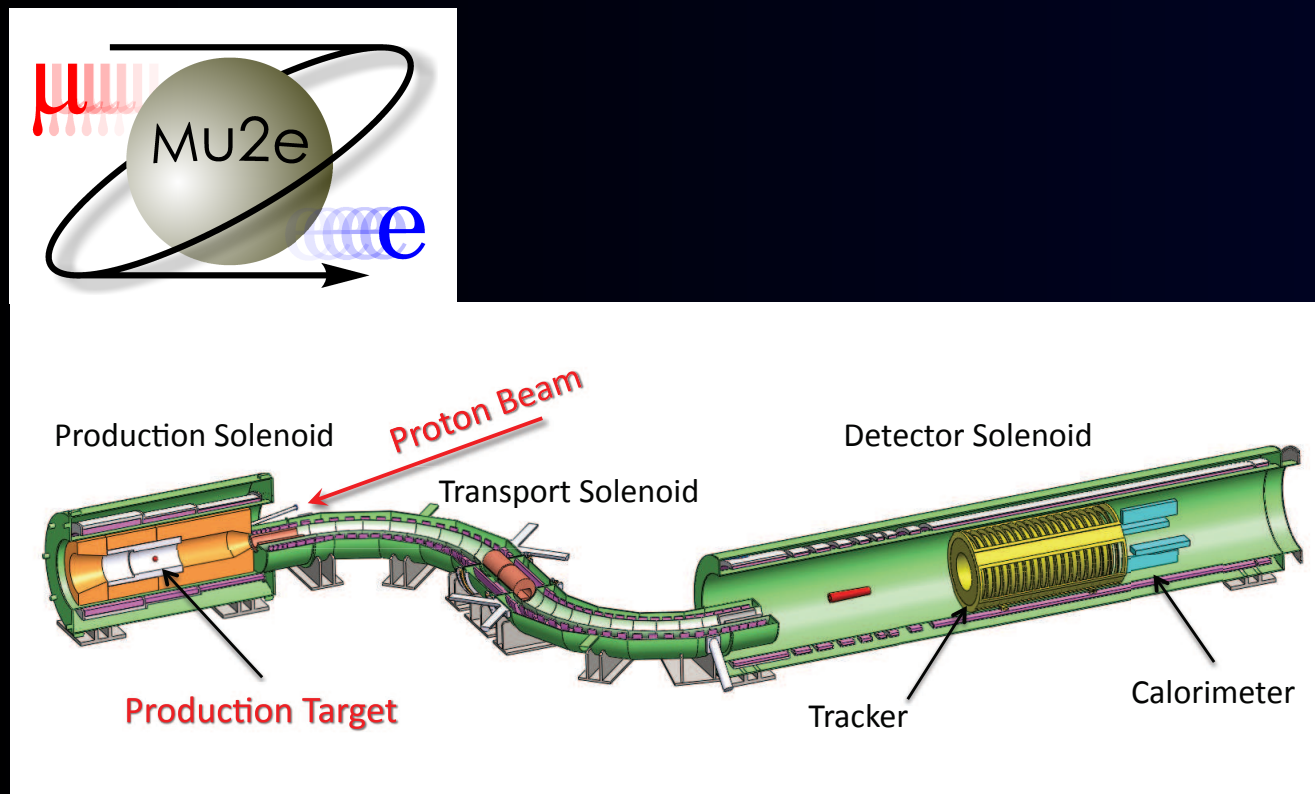
CLFV Experiments in Muon Decays



CLFV Experiments in Muon Decays

$\mu N \rightarrow e N$ Mu2e @FNAL

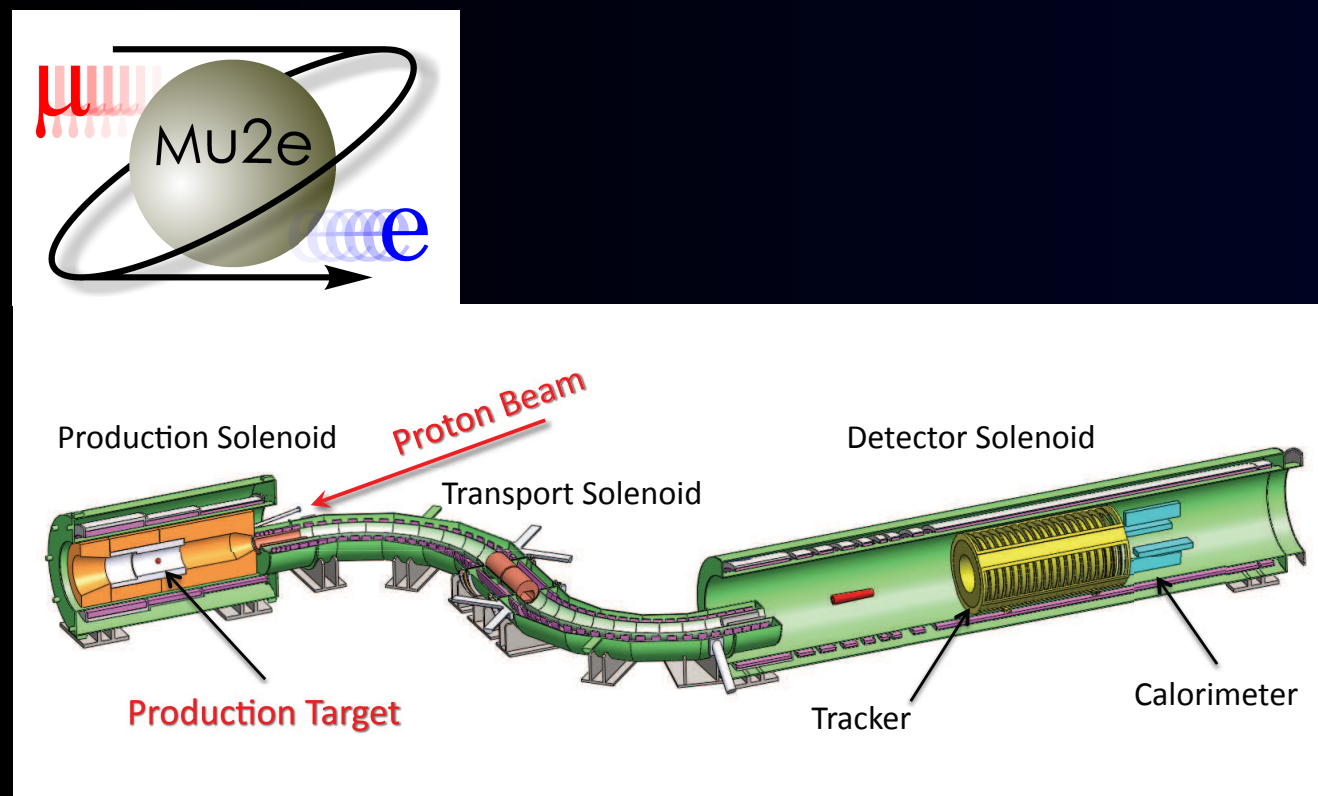
- $\text{SES} < 3 \times 10^{-17}$
- DOE CD2 review in 2014
- start in 2019



CLFV Experiments in Muon Decays

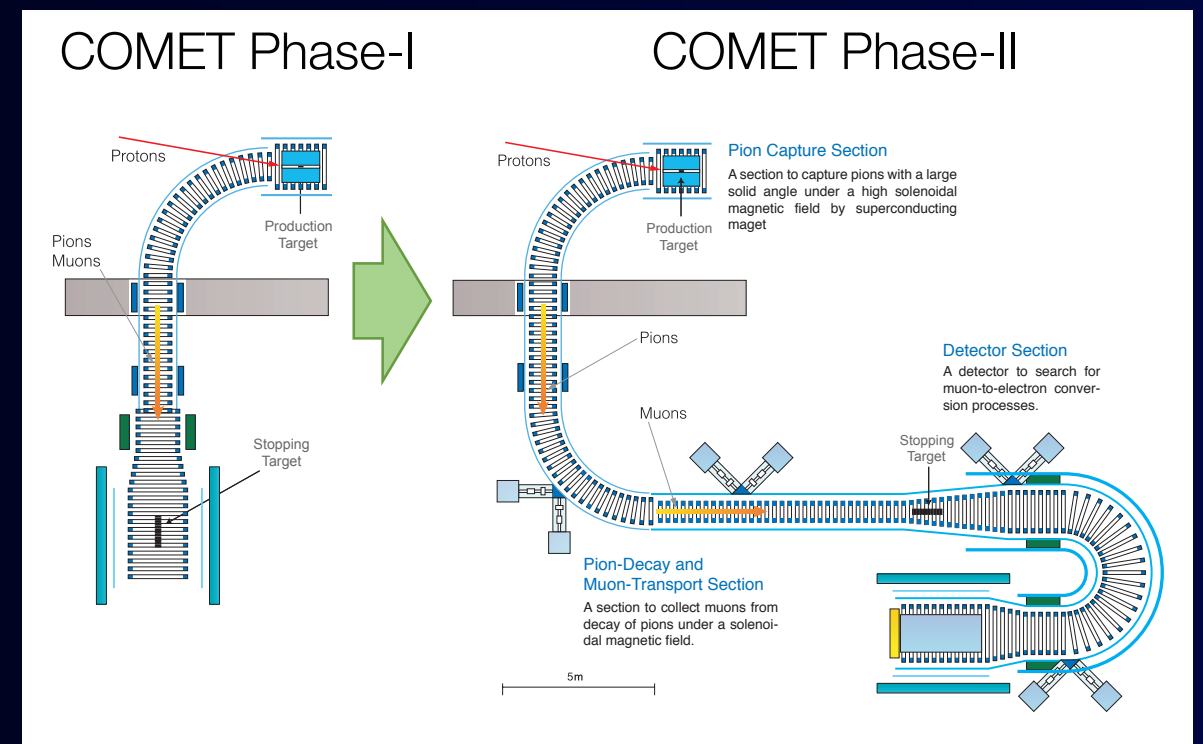
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$\mu N \rightarrow e N$ COMET @J-PARC

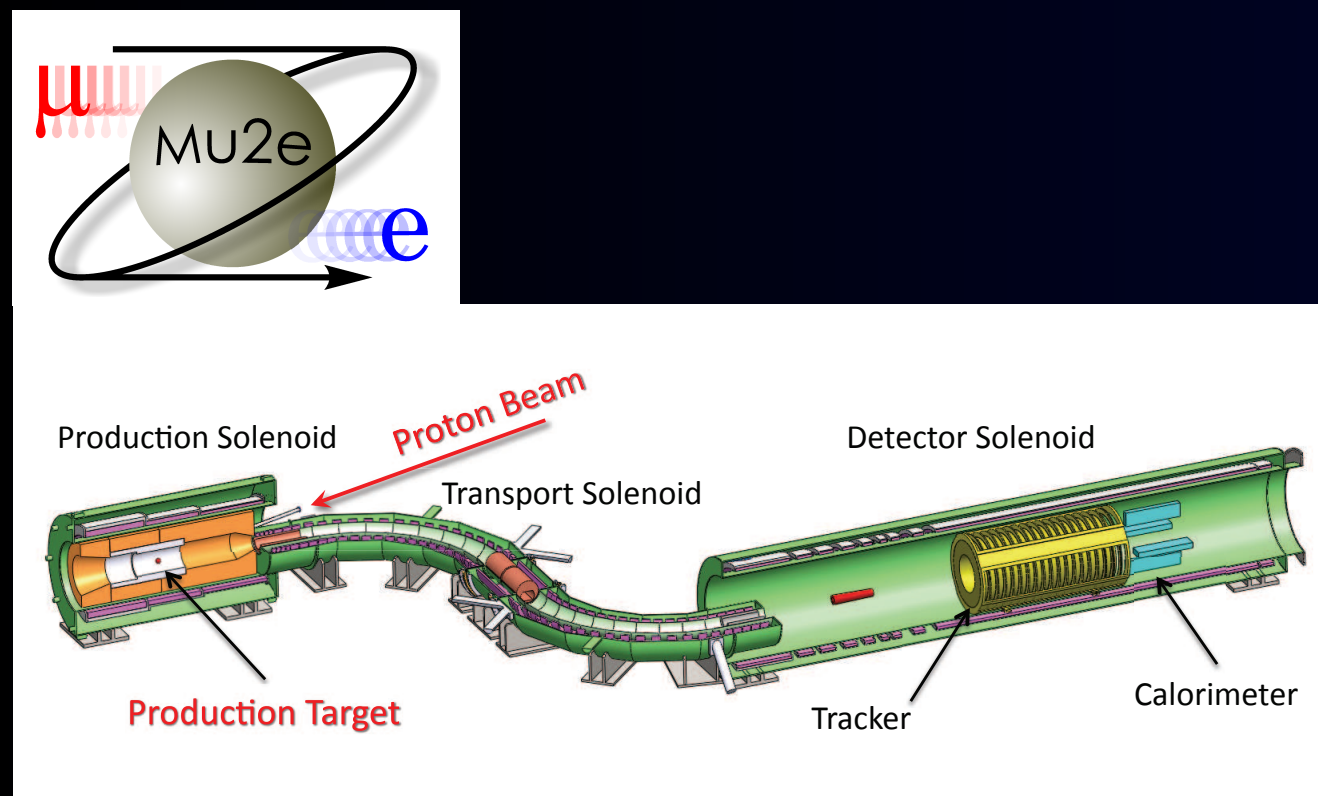
- staged approach
- COMET Phase-I $< 3 \times 10^{-15}$, 2016
- COMET Phase-II $< 3 \times 10^{-17}$, 2019



CLFV Experiments in Muon Decays

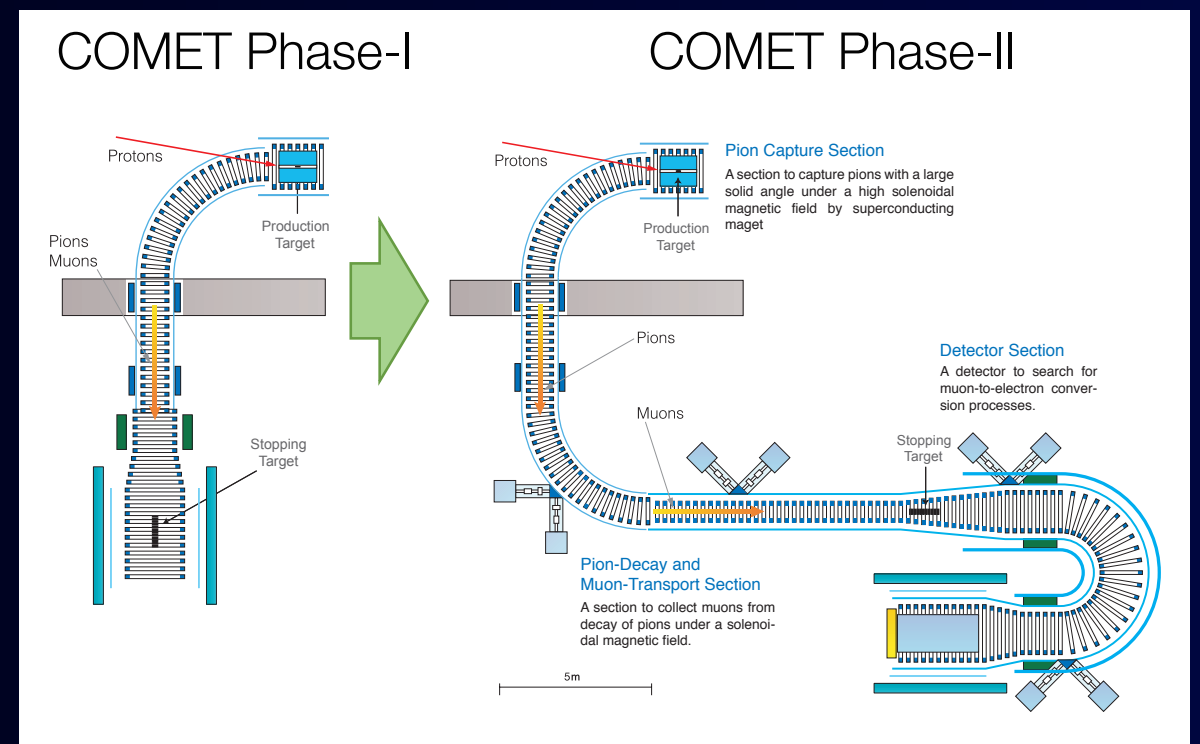
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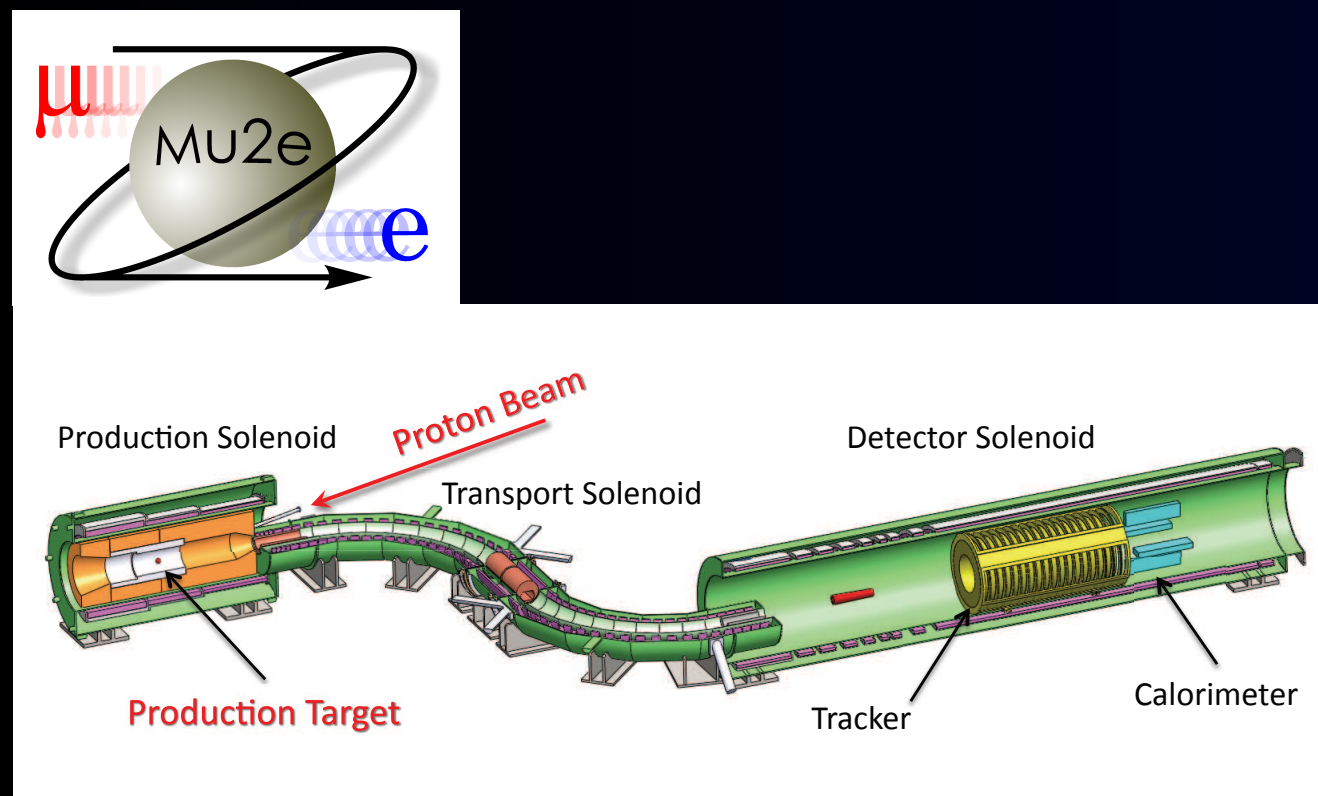


see Giovanni Tassielli's talk

CLFV Experiments in Muon Decays

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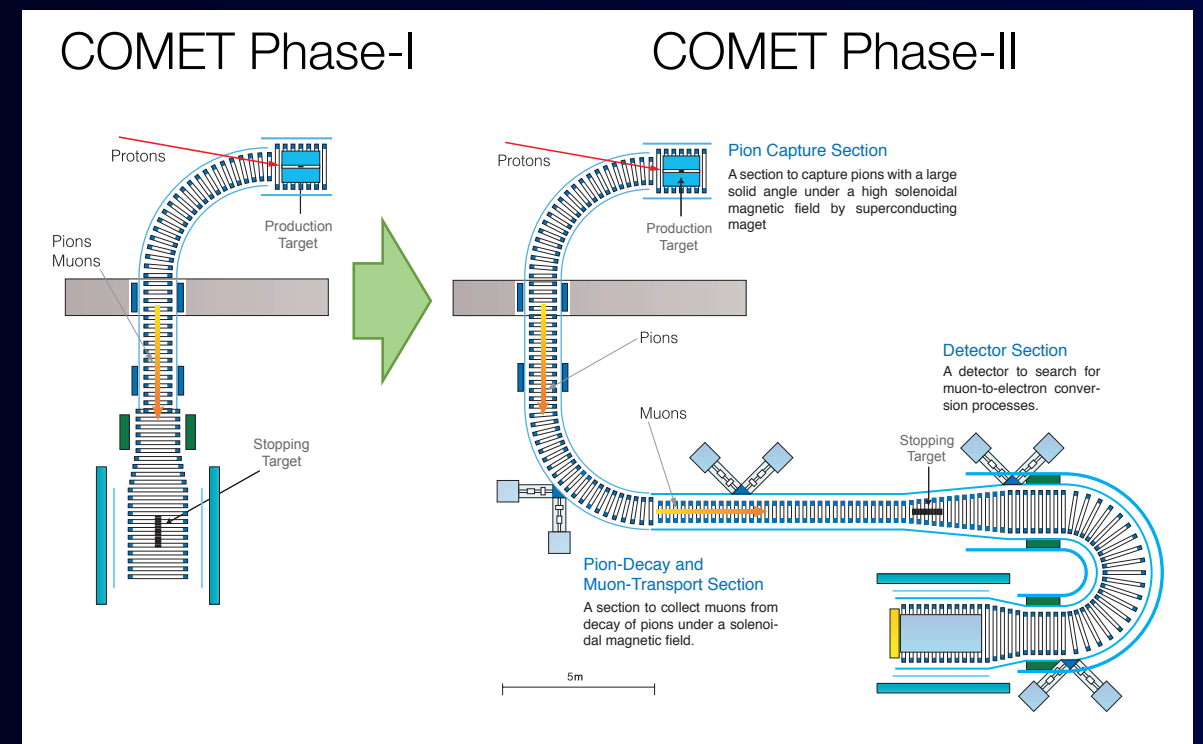
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see Giovanni Tassielli's talk

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- COMET Phase-II $< 3 \times 10^{-17}$, 2019



see Tran Hoai Nam's talk

Breakthrough in Muon Sources

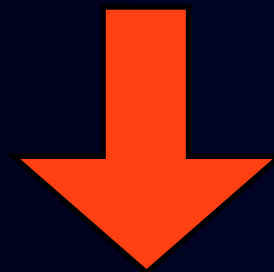


High Energy Scale Reach in CLFV



High Energy Scale Reach in CLFV

$$R \sim \frac{1}{\Lambda^4}$$



Can we improve the Λ reach by an order of magnitude ?
We must have at least 10^4 times the number of parent particles in rare decays.

An aerial photograph of the CERN Large Hadron Collider (LHC) in Switzerland. The LHC's circular tunnel is visible as a thin, light-colored line snaking through a dense, green forested landscape. In the bottom left corner, a portion of an airport with runways and taxiways is visible. The sky above is blue with scattered white clouds. A bright yellow rectangular box is positioned in the top left corner, containing the word "CERN" in a large, black, sans-serif font.

FNAL

Fermilab Accelerator Complex 2012

Advanced Superconducting Test Accelerator (under construction)

CDF

Linac and Booster

Tevatron (decommissioned)

DZero

Test-Beam Fixed-Target Beamlines

MTA Muon Test Area

Muon Campus (under construction)

MINOS - NOvA To Minnesota

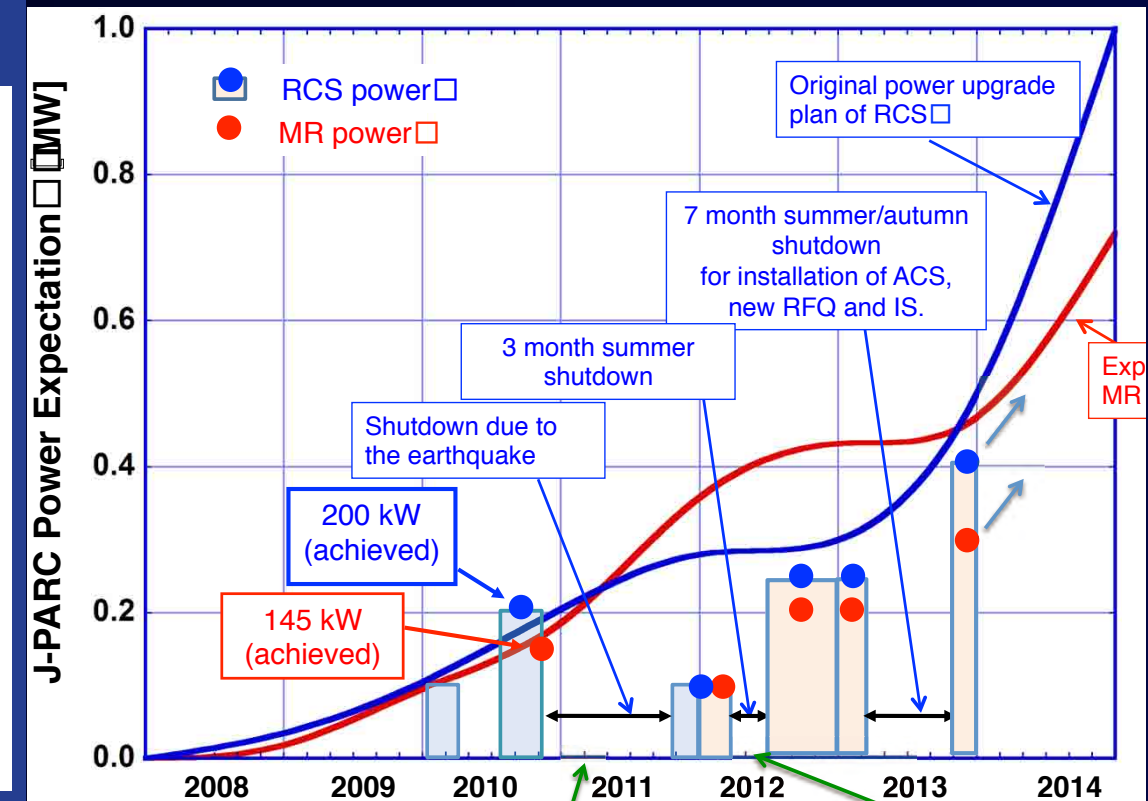
Booster Neutrino Beam

under development

Main Injector and Recycler

An aerial photograph of the J-PARC (Japan Proton Accelerator Research Complex) facility. The image shows a large, circular accelerator ring (AR) in the center, surrounded by various buildings and infrastructure. A yellow rectangular box is overlaid in the top left corner, containing the text "J-PARC" in black, sans-serif font. The facility is situated in a wooded area, with a road and parking lots visible. The overall scene is captured from a high angle, providing a clear view of the complex's layout.

The graph displays the projected proton flux for various neutrino experiments from 2011 to 2021. The y-axis represents Protons/Hour, ranging from 0.00E+00 to 2.50E+17. The x-axis represents years from 2011 to 2021. The graph includes data for Main Injector (blue diamonds), Booster Neutrinos (red squares), g-2 (green triangles), mu2e (purple crosses), and Total (cyan asterisks). A 'Shutdown' arrow points to the period between 2012 and 2013. Labels for experiments include MINOS, MINERvA, MiniBooNE, MicroBooNE, Muon g-2, and LBNE. Energy thresholds are marked: 120 GeV ν , 8 GeV ν , 8 GeV μ , and 120 GeV ν .



MuSIC Facility at Osaka University

- Frontend of Mu2e / COMET -

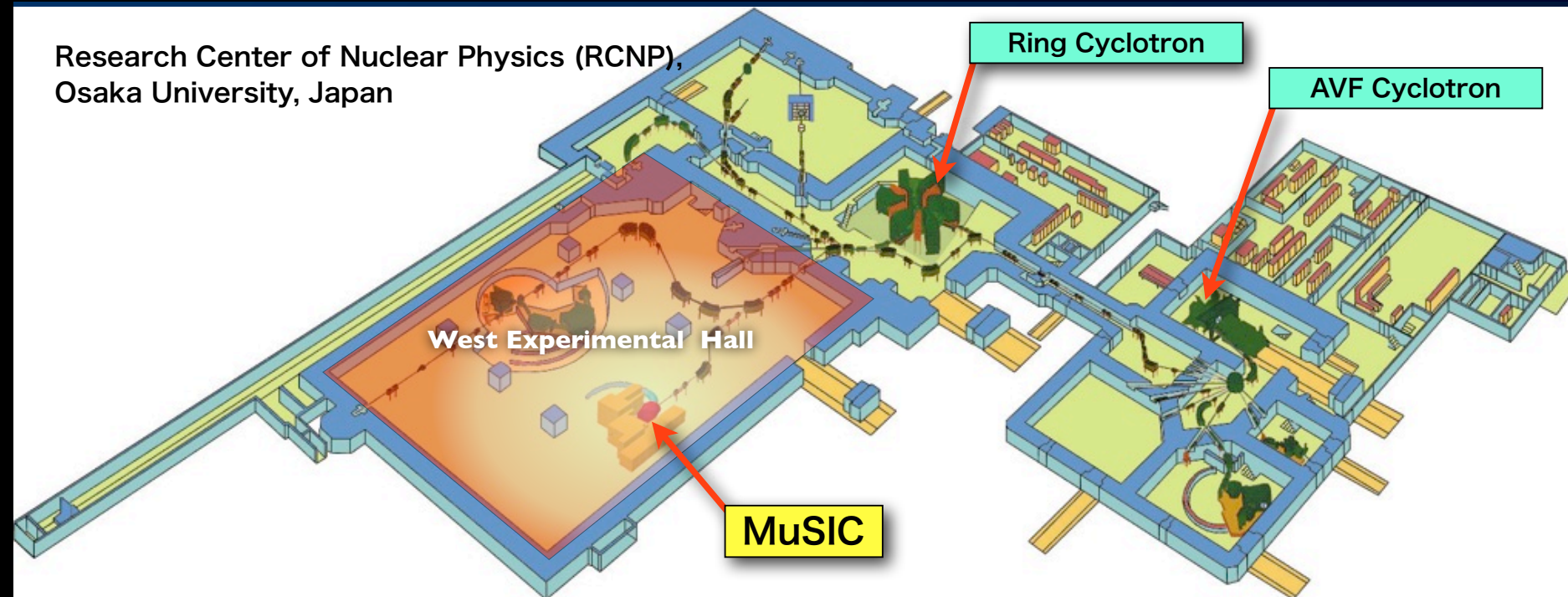


MuSIC Facility at Osaka University

- Frontend of Mu2e / COMET -



Research Center of Nuclear Physics (RCNP),
Osaka University, Japan



RCNP cyclotron
400 MeV, 1 μ A

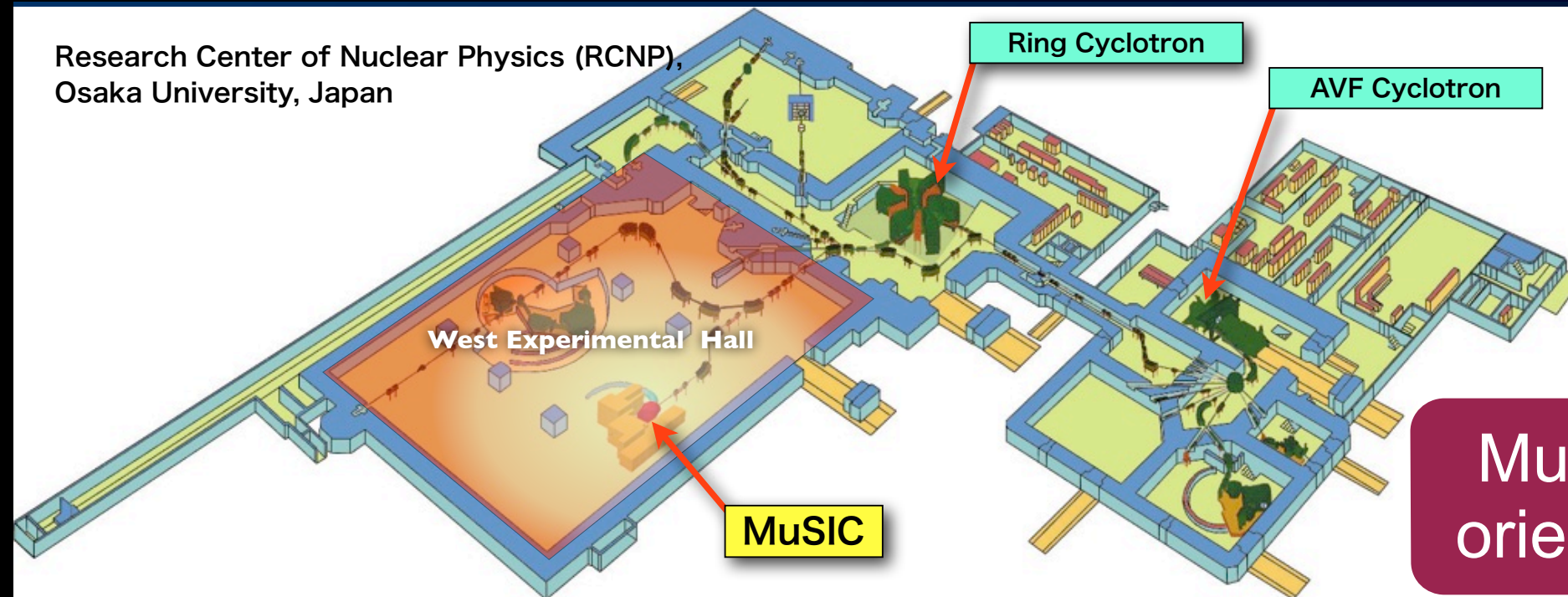
- RCNP has two cyclotrons. A proton beam with 392 MeV, 1 μ A is provided from the Ring Cyclotron (up to 5 μ A in near future).
- The MuSIC is in the largest experimental hall, the west experimental hall.

MuSIC Facility at Osaka University

- Frontend of Mu2e / COMET -



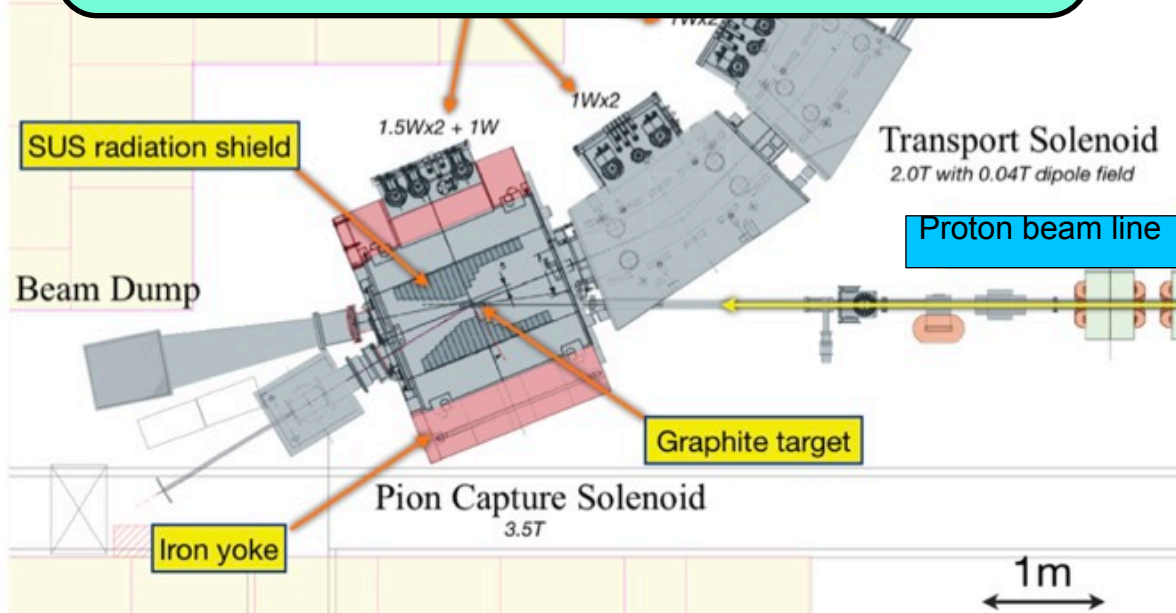
Research Center of Nuclear Physics (RCNP),
Osaka University, Japan



RCNP cyclotron
400 MeV, $1\mu\text{A}$

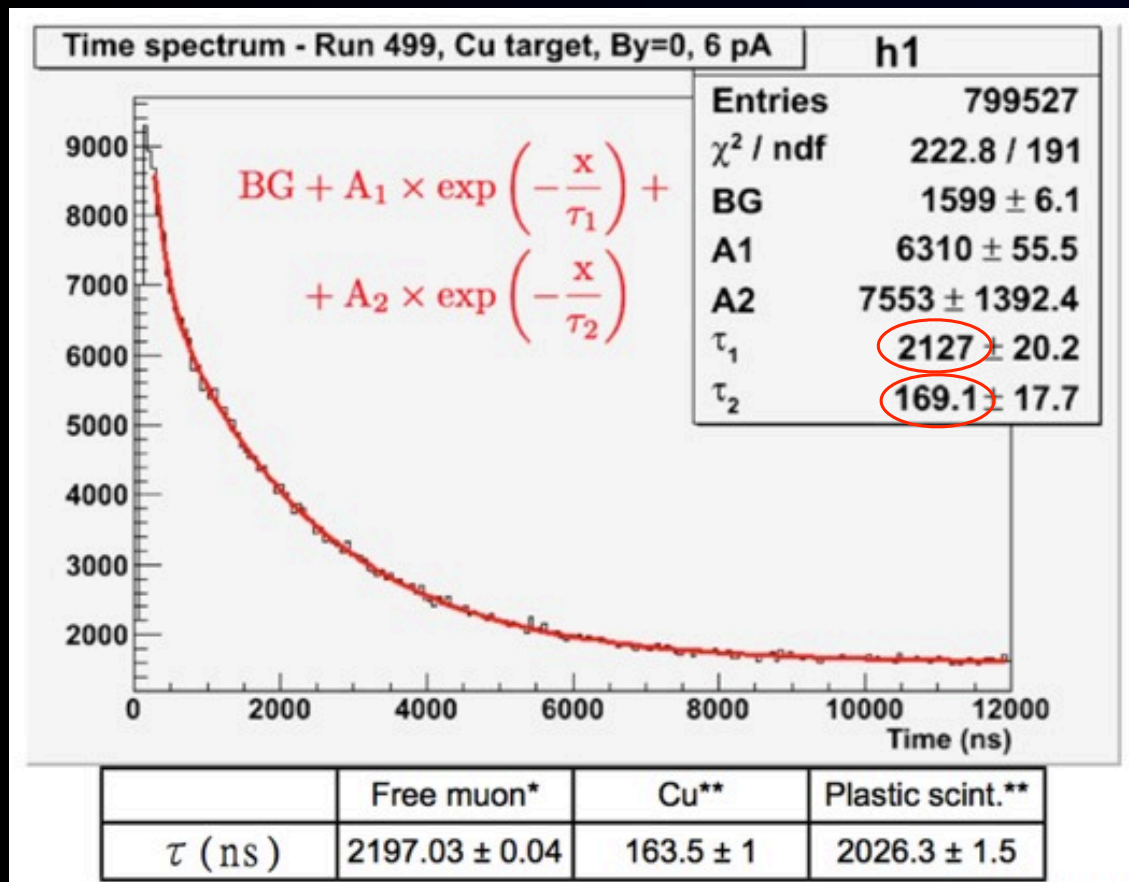
MuSIC = Muon Science
oriented Intense Channel

Pion Capture System

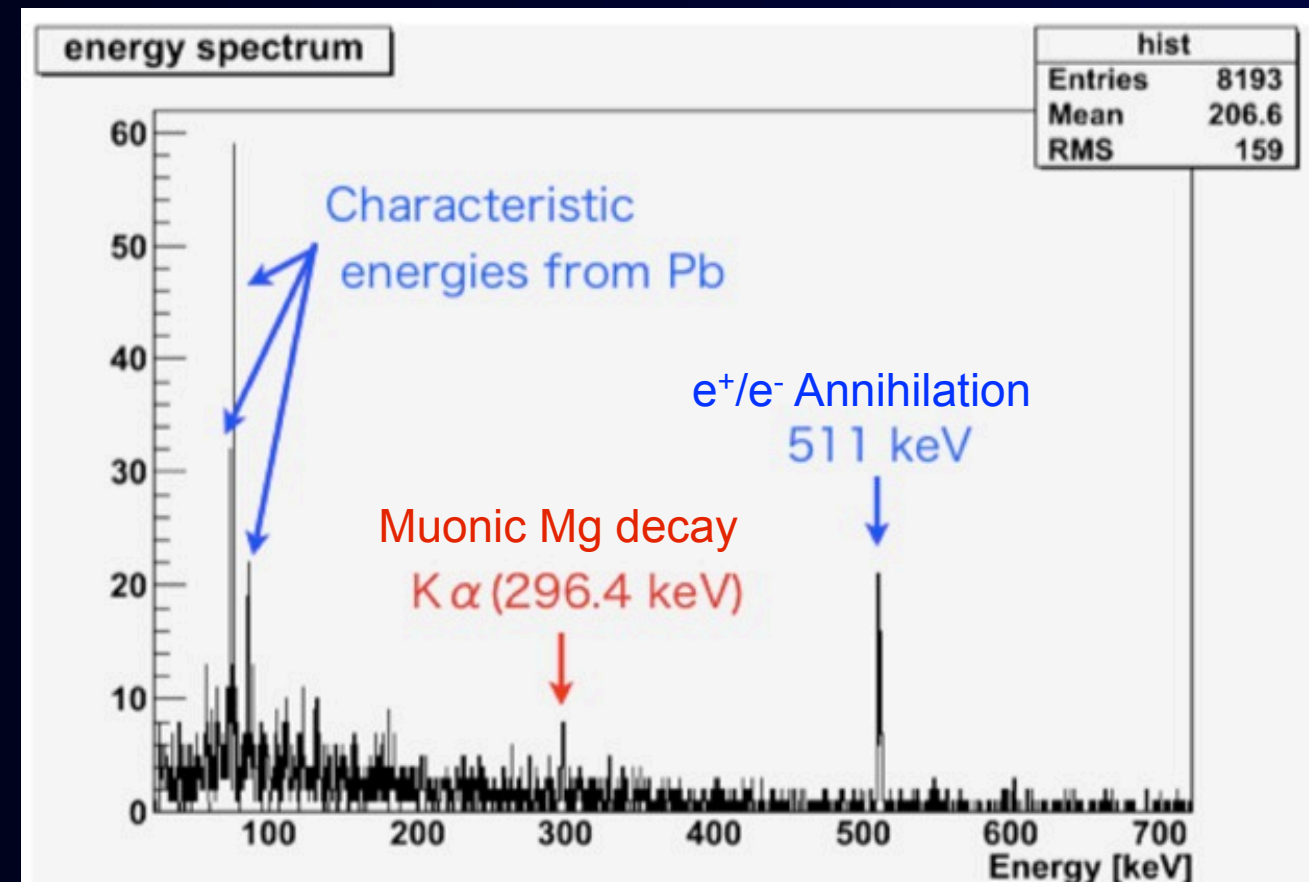


MuSIC Facility at Osaka University

Muon Production Efficiency (x1000)



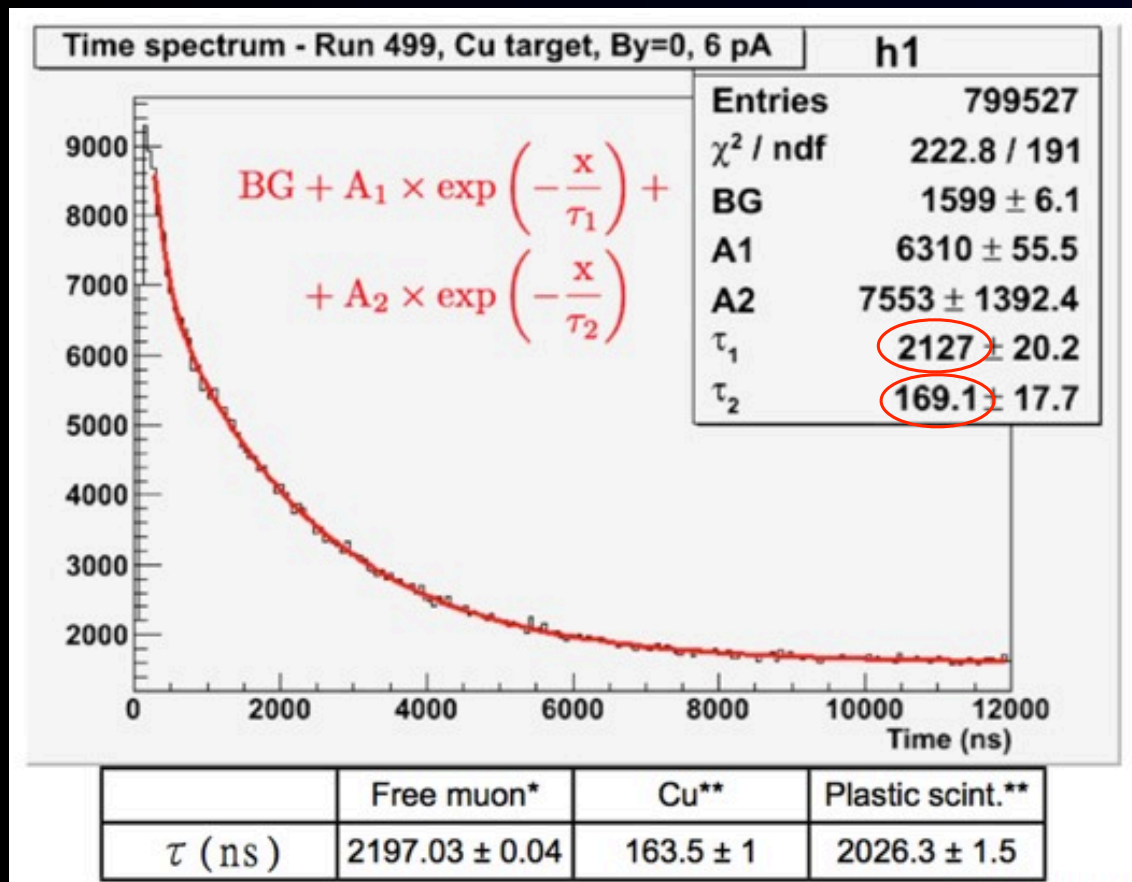
positive muons



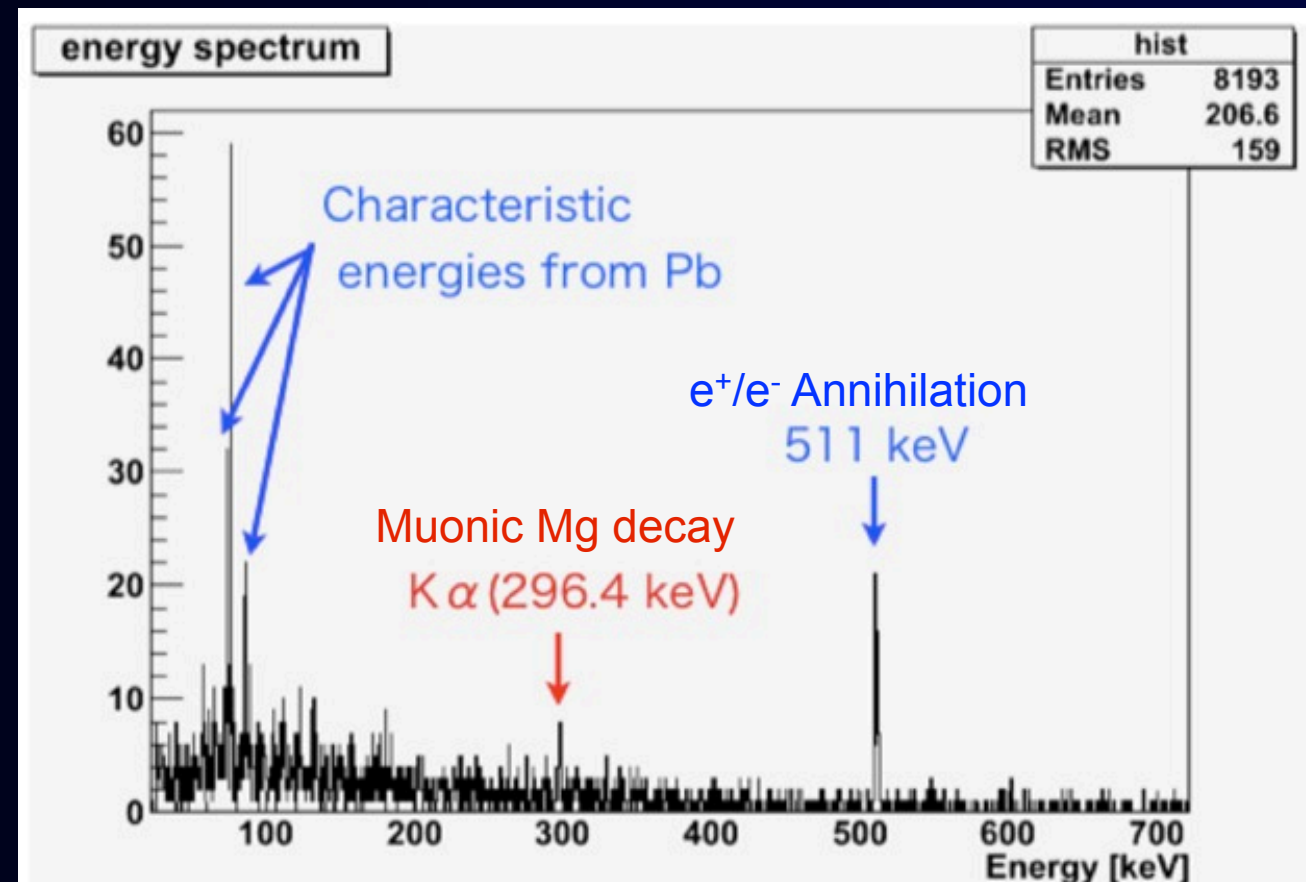
negative muons

MuSIC Facility at Osaka University

Muon Production Efficiency (x1000)



positive muons



negative muons

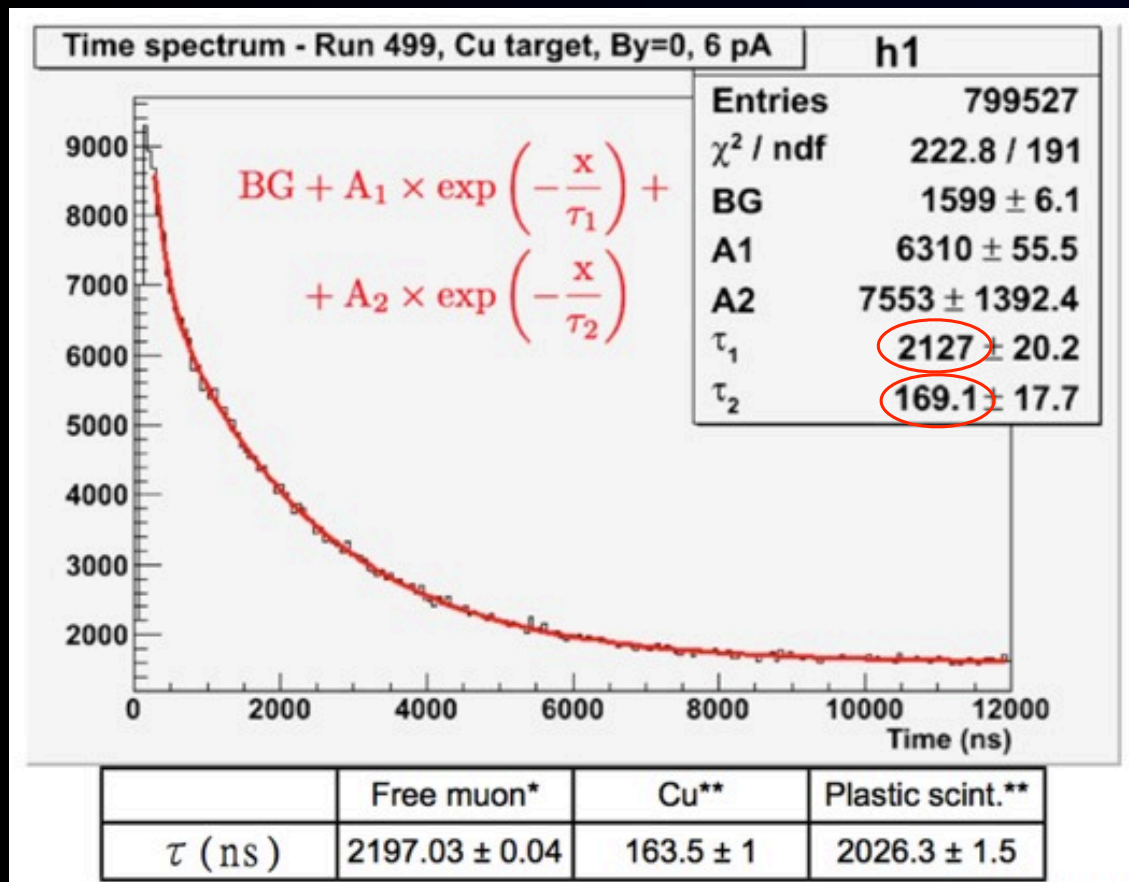
MuSIC muon yields

μ^+ : $3 \times 10^8/\text{s}$ for 400W

μ^- : $1 \times 10^8/\text{s}$ for 400W

MuSIC Facility at Osaka University

Muon Production Efficiency (x1000)

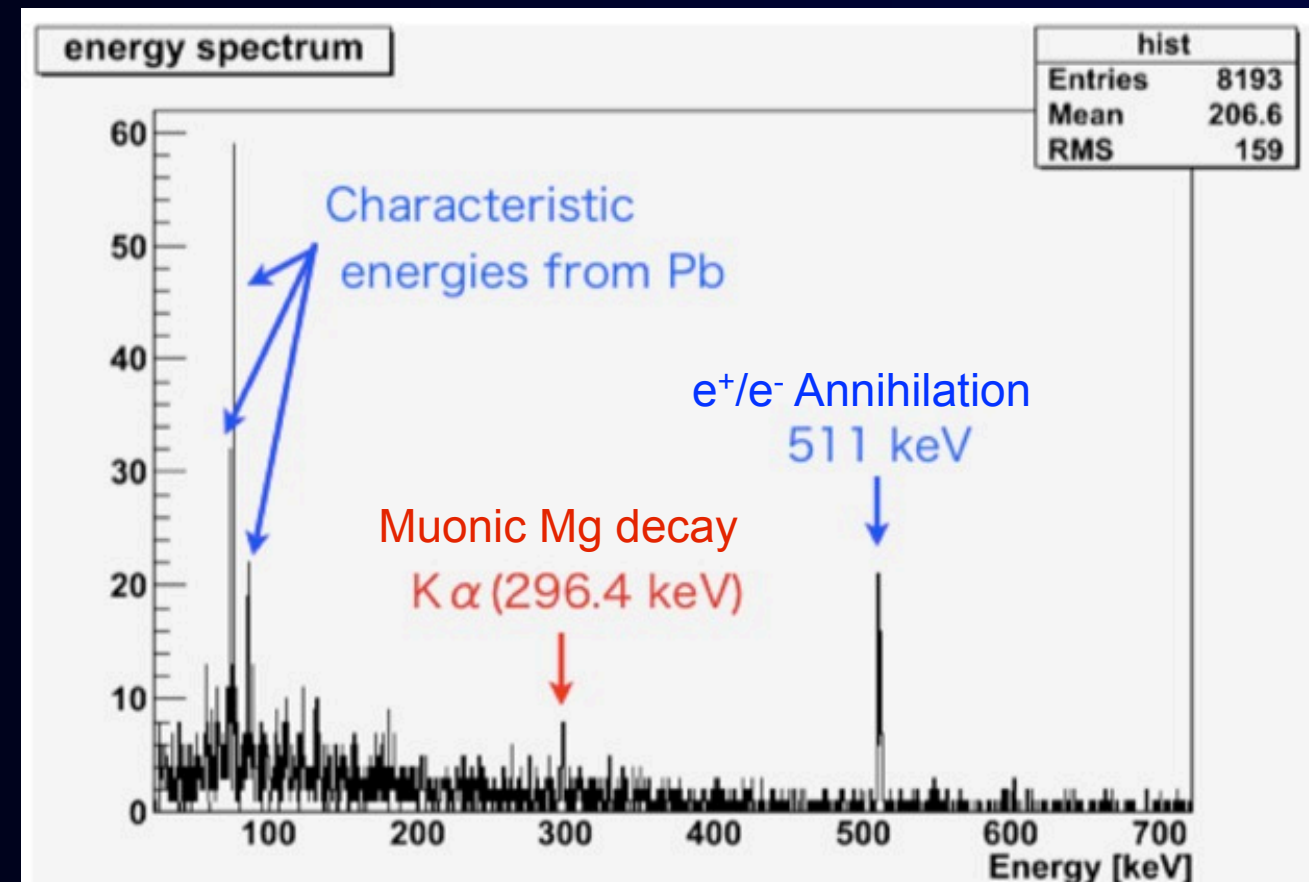


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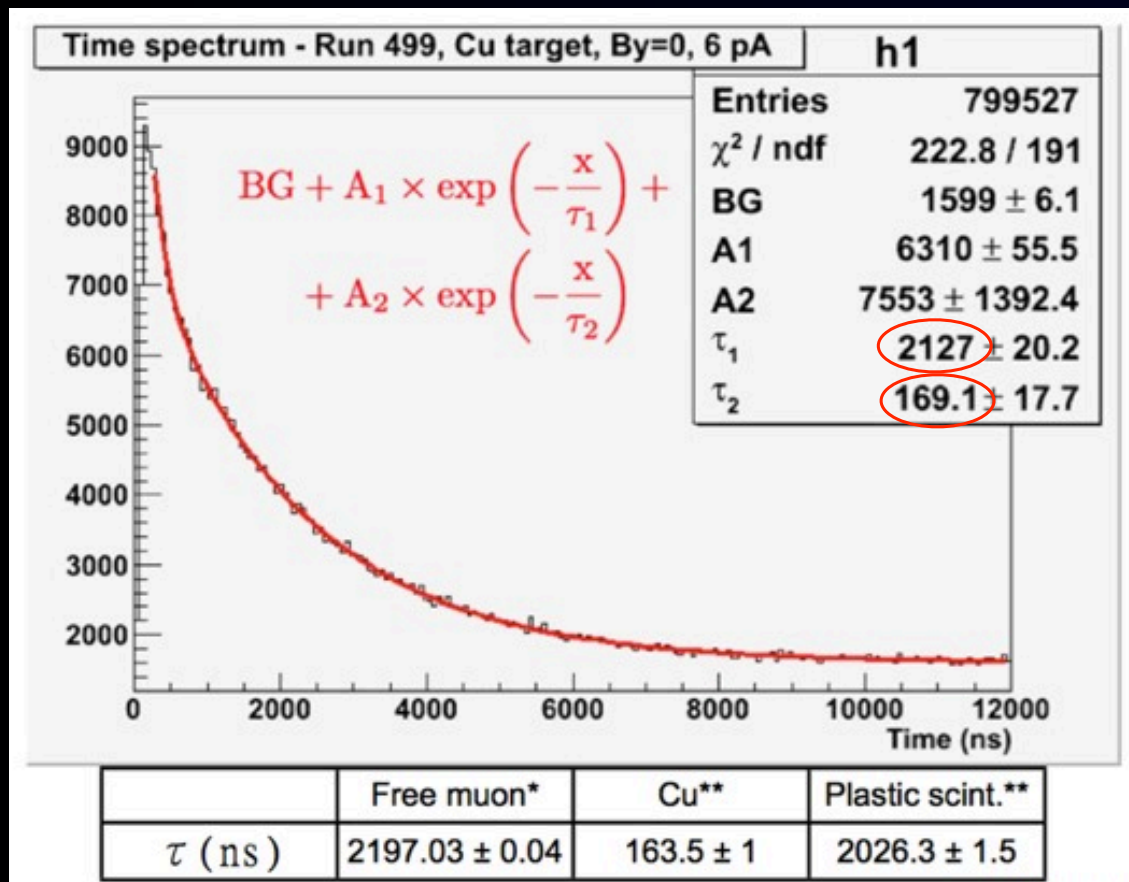


negative muons

cf. $10^8/\text{s}$ for 1.3MW @PSI
Requirements of $\times 10^3$ achieved...

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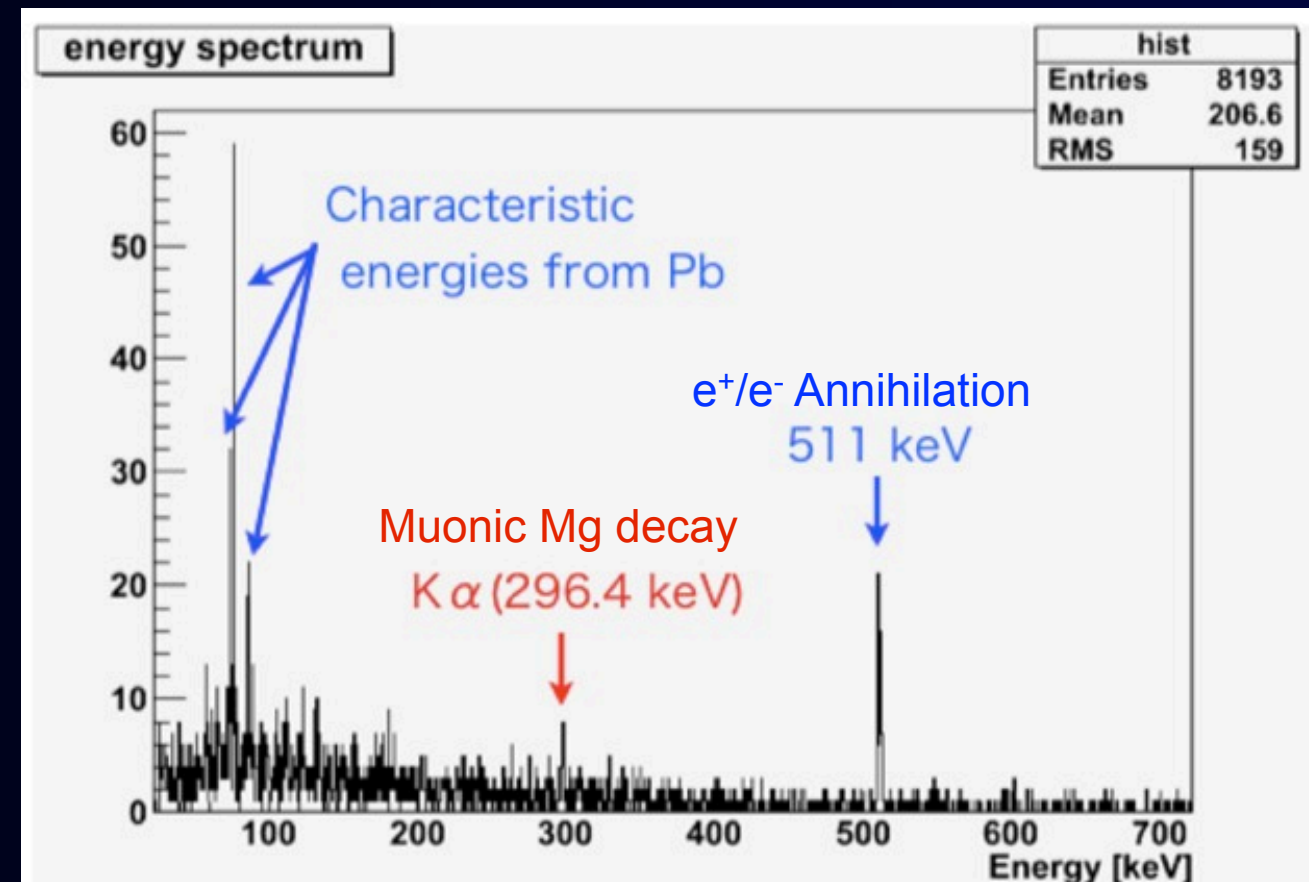


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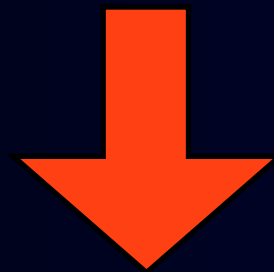
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**Demonstration of
Pion Capture System**

High Energy Scale Reach in CLFV

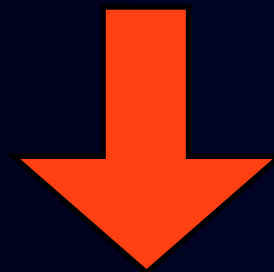
$$R \sim \frac{1}{\Lambda^4}$$



Can we improve the Λ reach by an order of magnitude ?
We must have at least 10^4 times the number of parent particles in rare decays.

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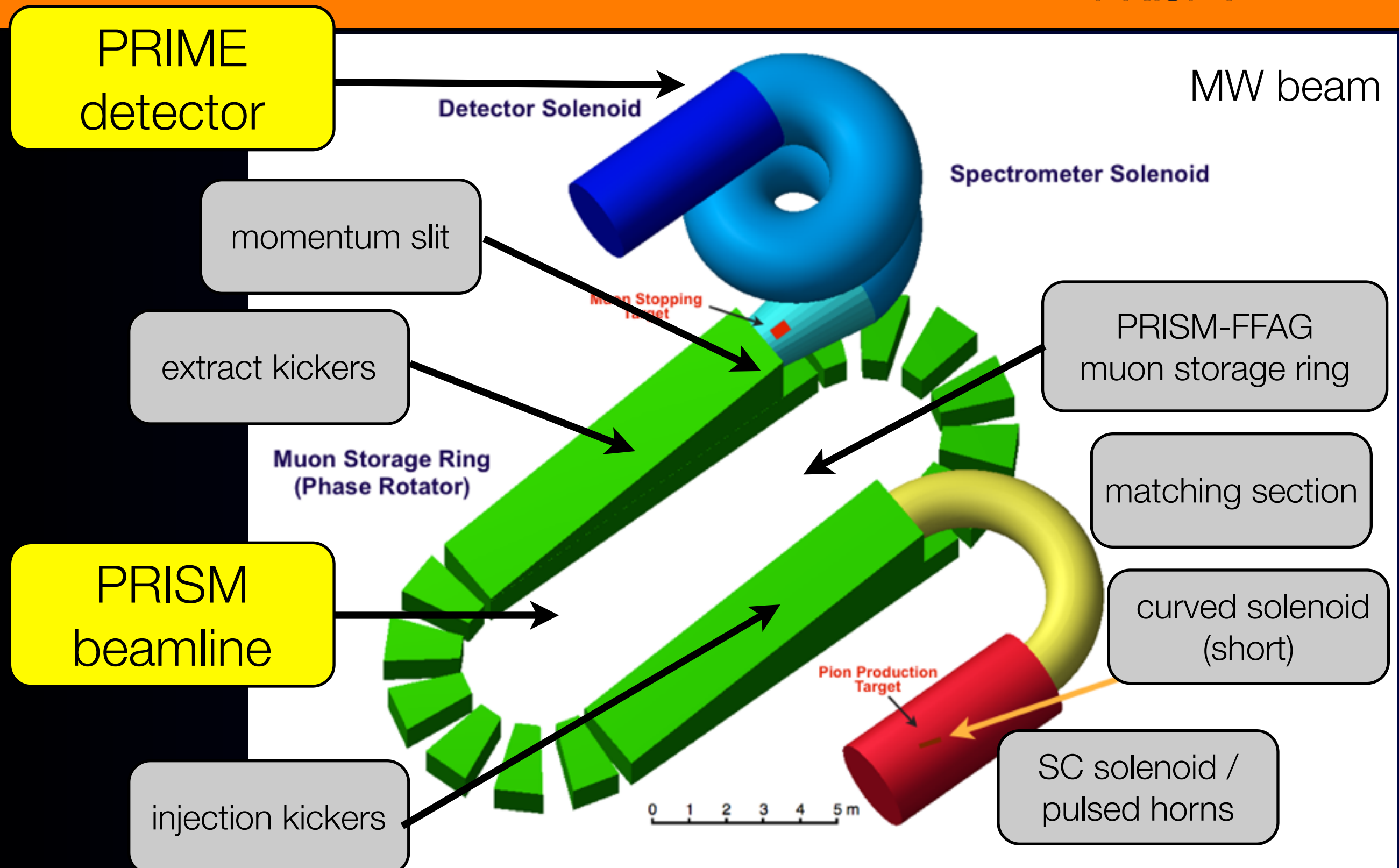
$$R \sim \frac{1}{\Lambda^4}$$



Can we improve the Λ reach by an order of magnitude ?
We must have at least 10^4 times the number of parent particles in rare decays.

Yes, now it is possible for muons with the novel pion capture system.

PRISM/PRIME : Future Search for μ -e Conversion at 3×10^{-19}



Summary



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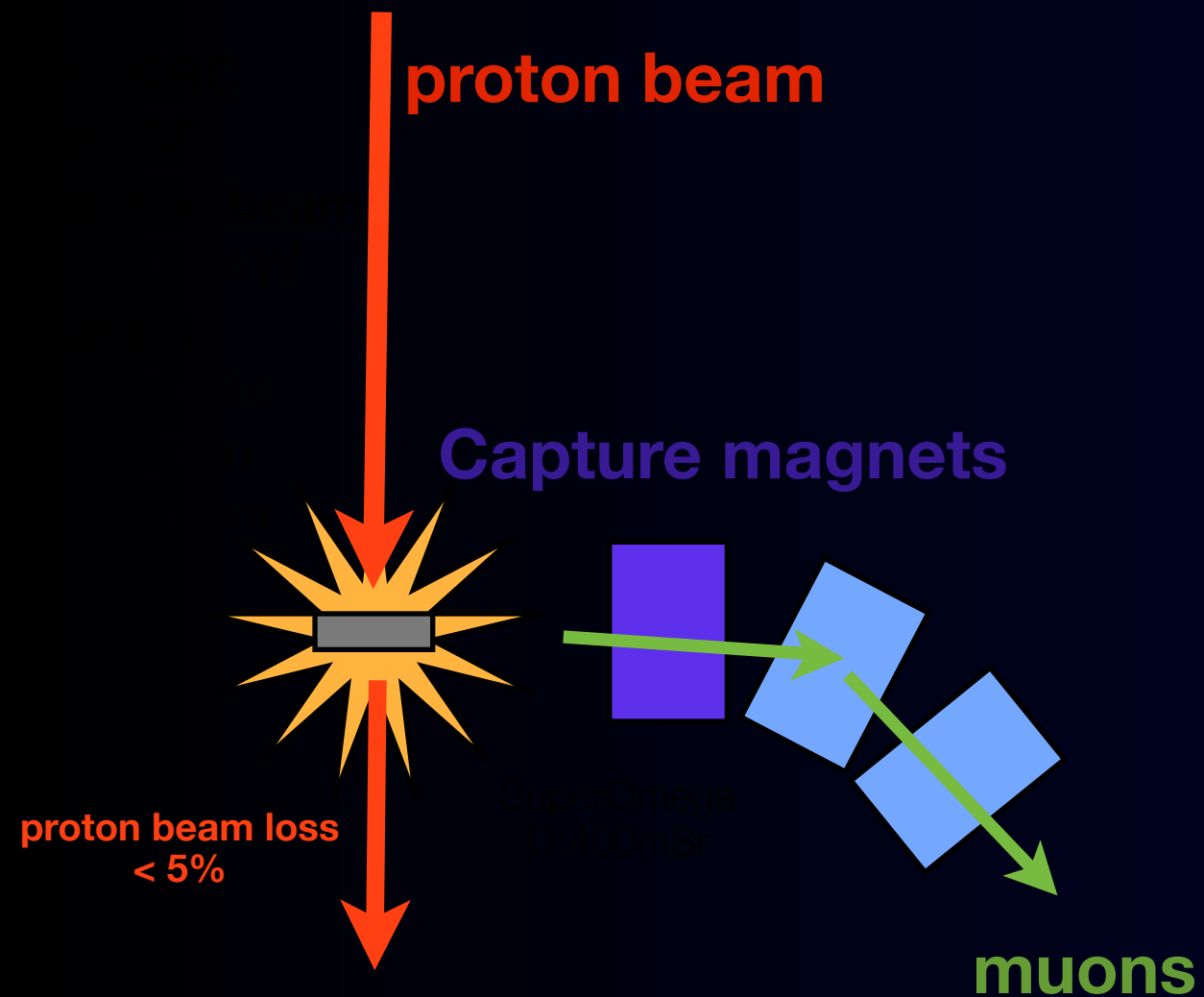
Backup

Production and Collection of Pions and Muons



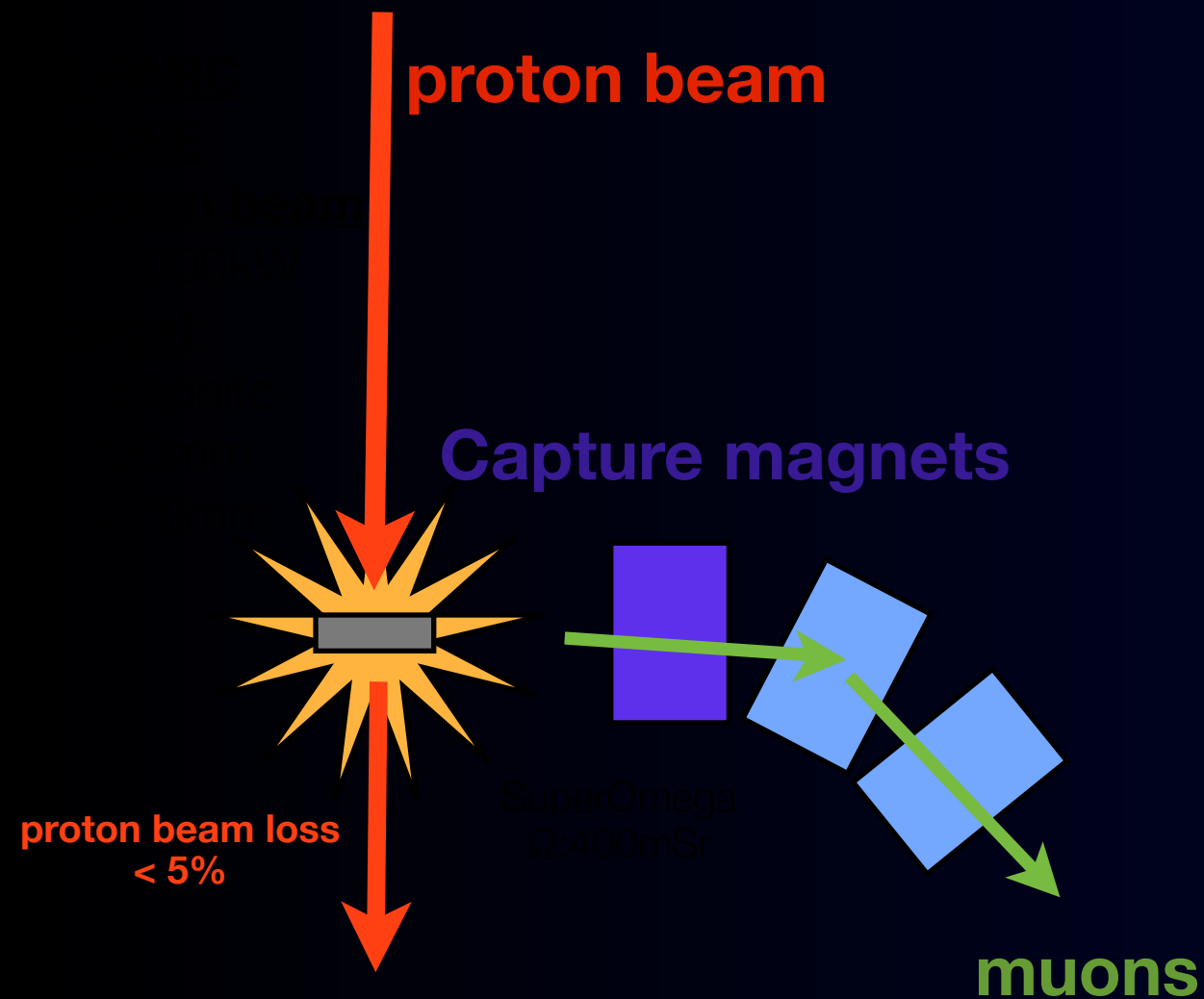
Production and Collection of Pions and Muons

Conventional muon beam line



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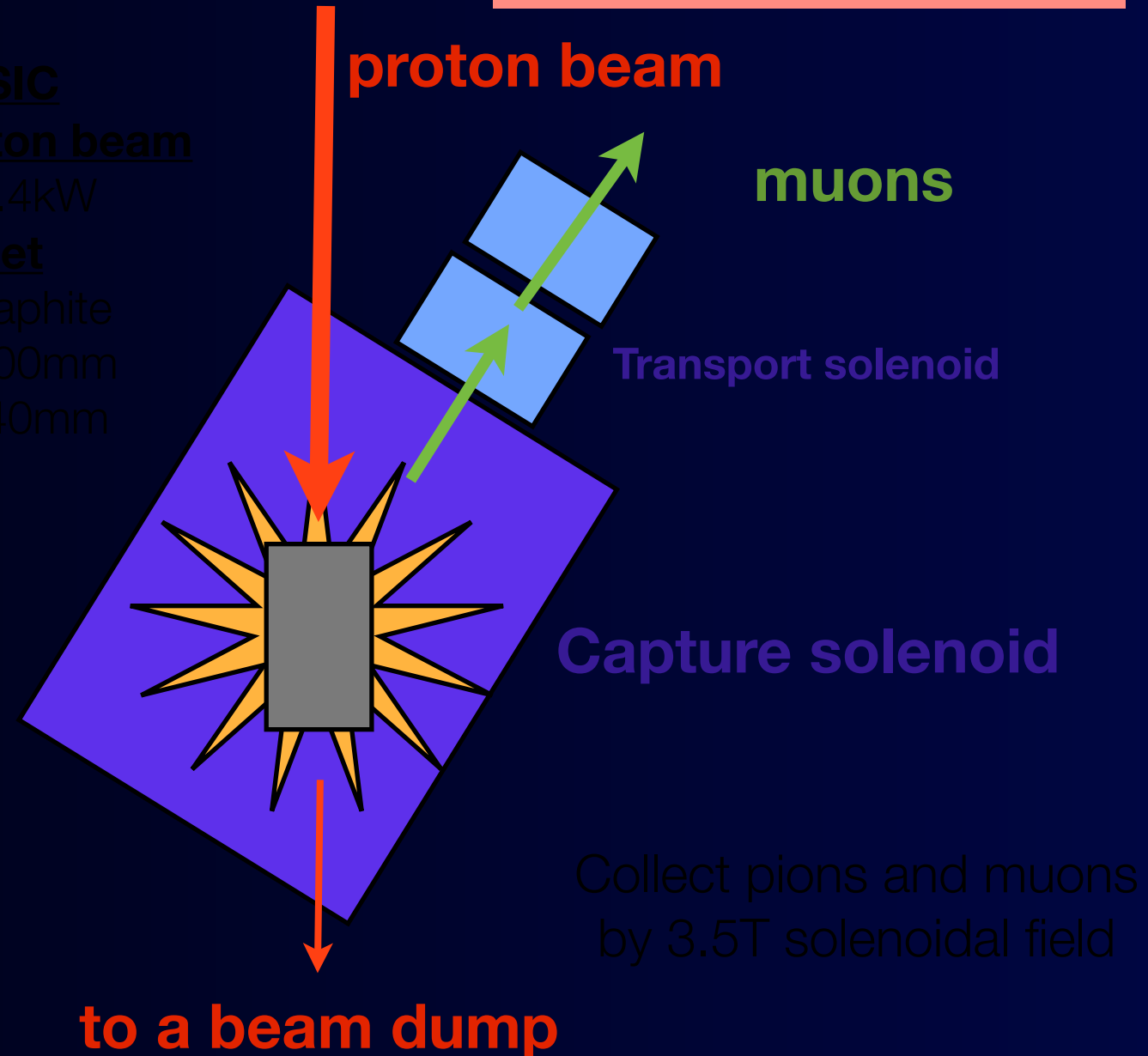
Much efficient

MuSIC
proton beam

-0.4kW

target

graphite
t200mm
φ40mm



MuSIC, COMET, PRISM,
Neutrino factory,
Muon collider

Large solid angle & thick target