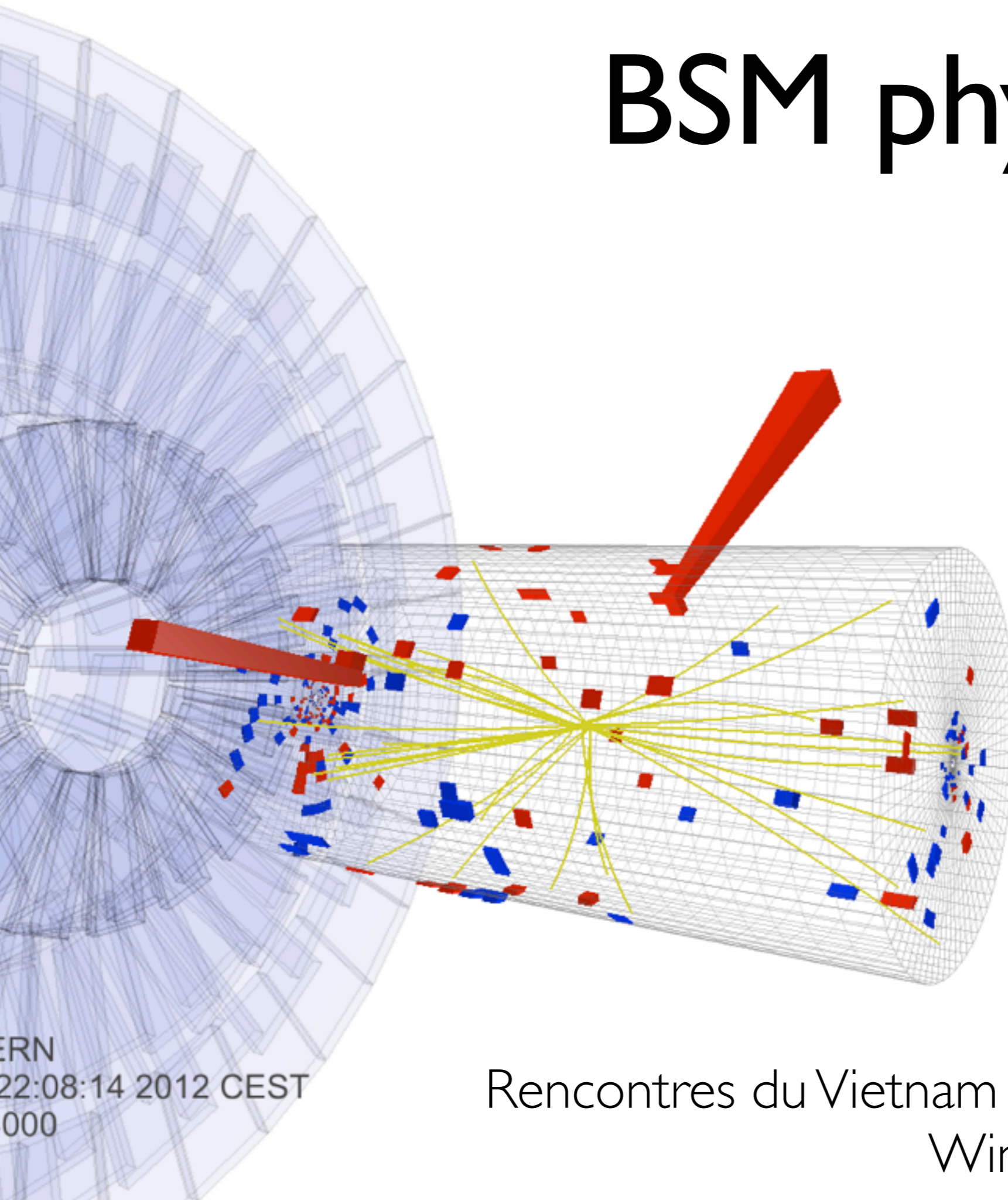


BSM physics after LHC8

Andreas Weiler
(DESY)



ERN
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Rencontres du Vietnam | Inaugural Conference
Windows on the Universe

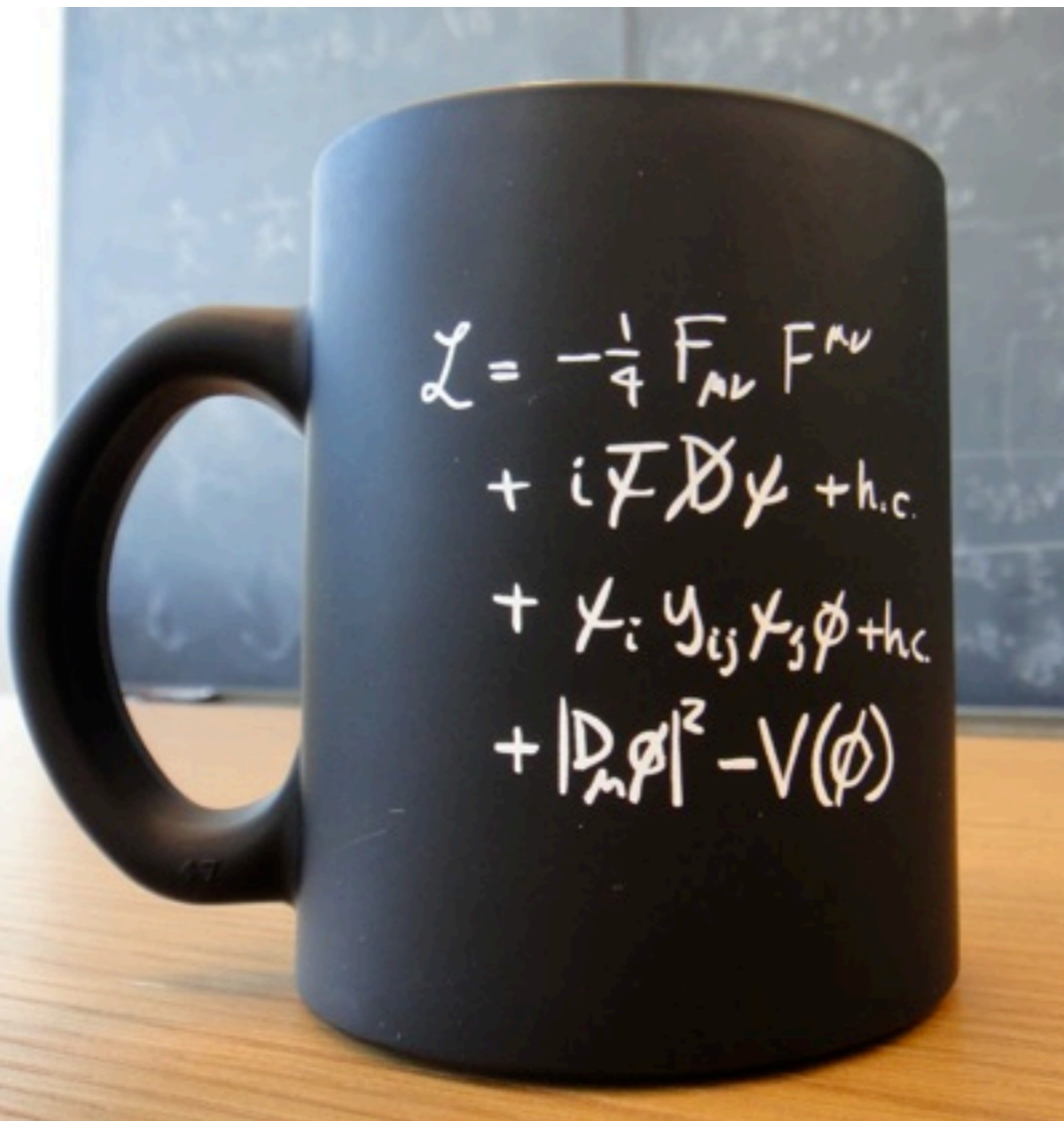
De-motivation



DESPAIR

It's always darkest just before it goes pitch black.


The SM rules



Agrees with all collider-based tests

Accidental symmetries of SM are fully observed (B,L)

Small breaking of global symmetries (flavor, $SU(2)_V$) measured as predicted

broken symmetry	operators	scale Λ
B, L	$(QQQL)/\Lambda^2$	10^{13} TeV
flavor (1,2 nd family), CP	$(\bar{d}s\bar{d}s)/\Lambda^2$	1000 TeV 
flavor (2,3 rd family)	$m_b(\bar{s}\sigma_{\mu\nu}F^{\mu\nu}b)/\Lambda^2$	50 TeV

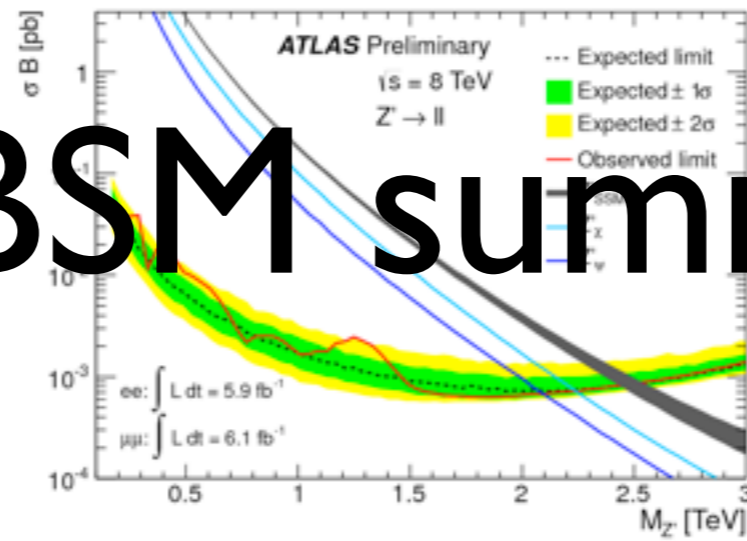
10^5 TeV

Generic new physics must be at $\Lambda \gg \text{TeV}$

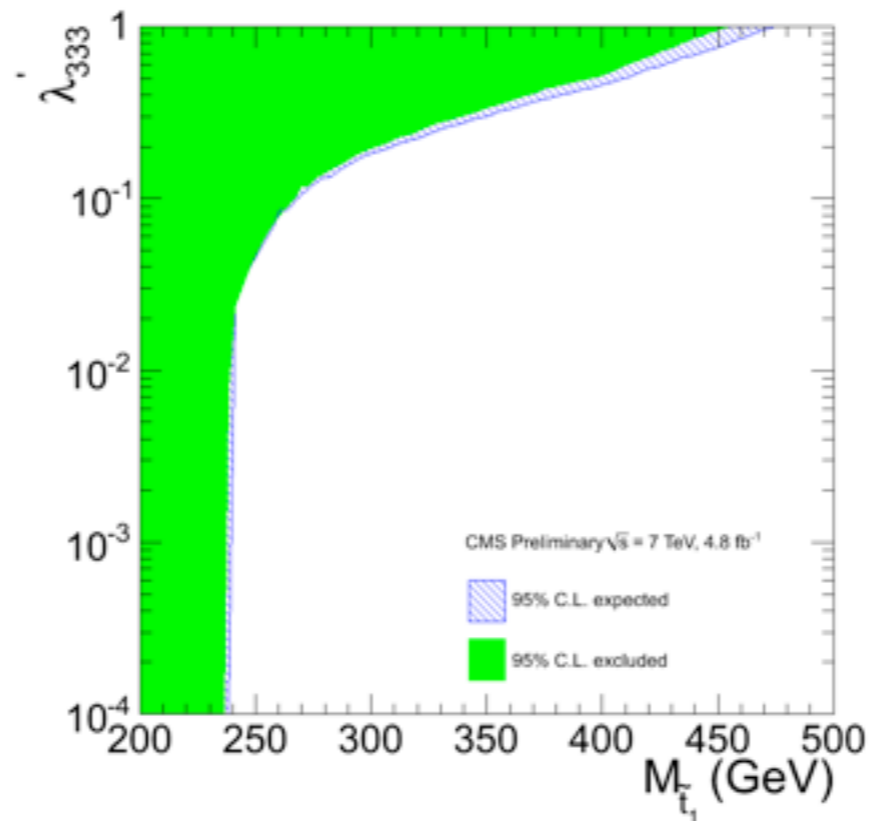
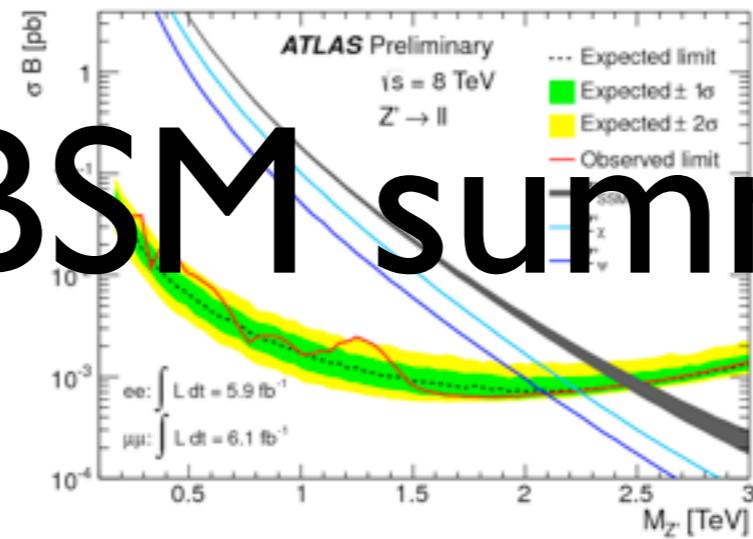
Flavor must be aligned with the SM,
or SM-like

LHC BSM summary:

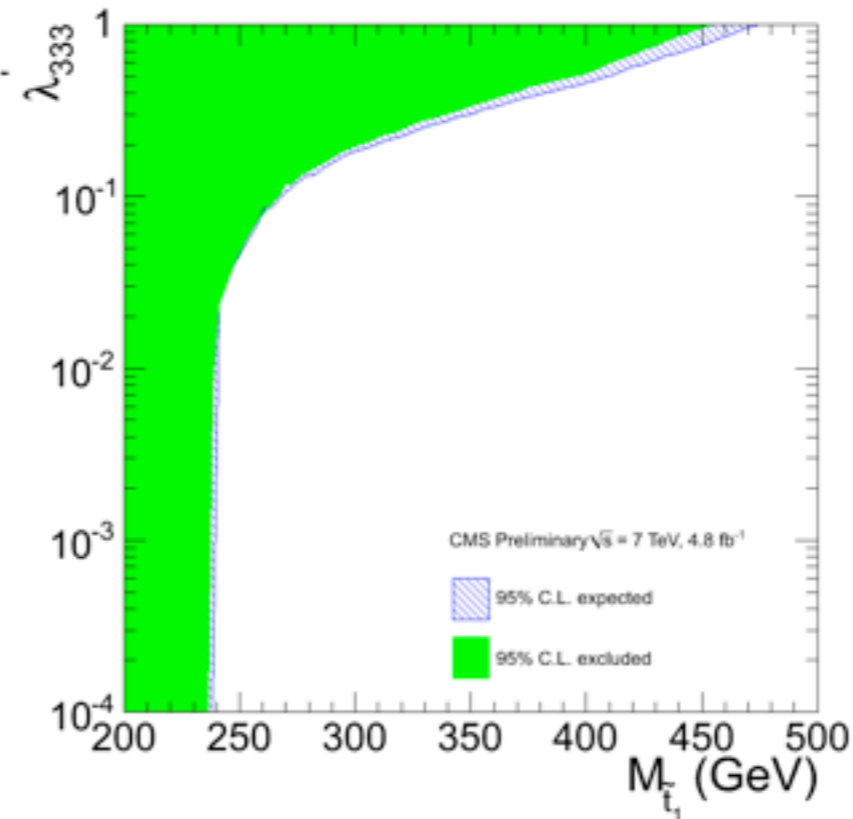
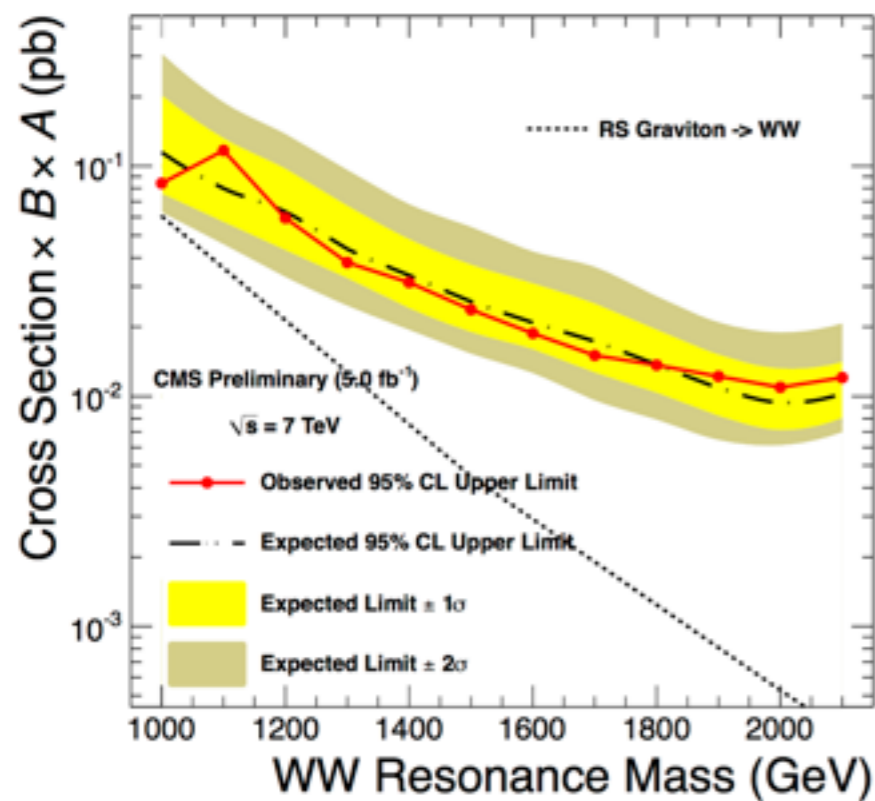
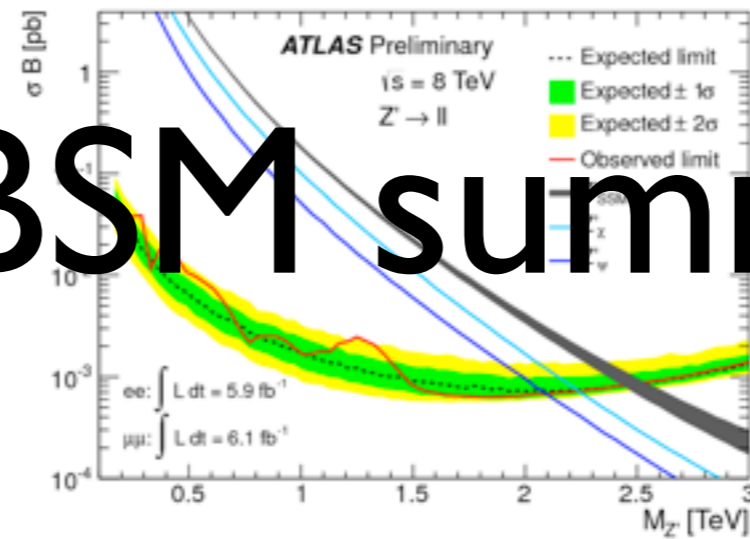
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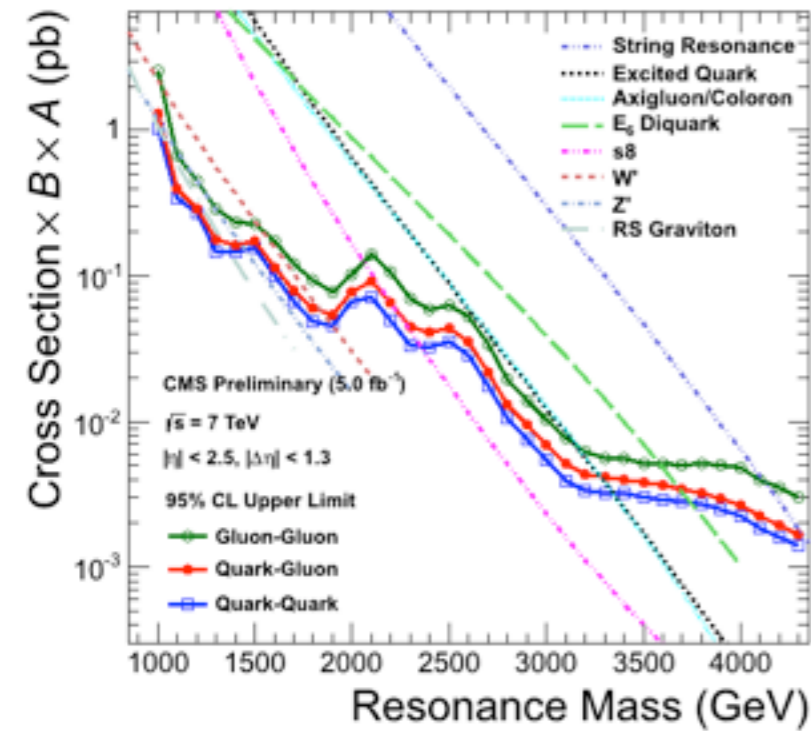


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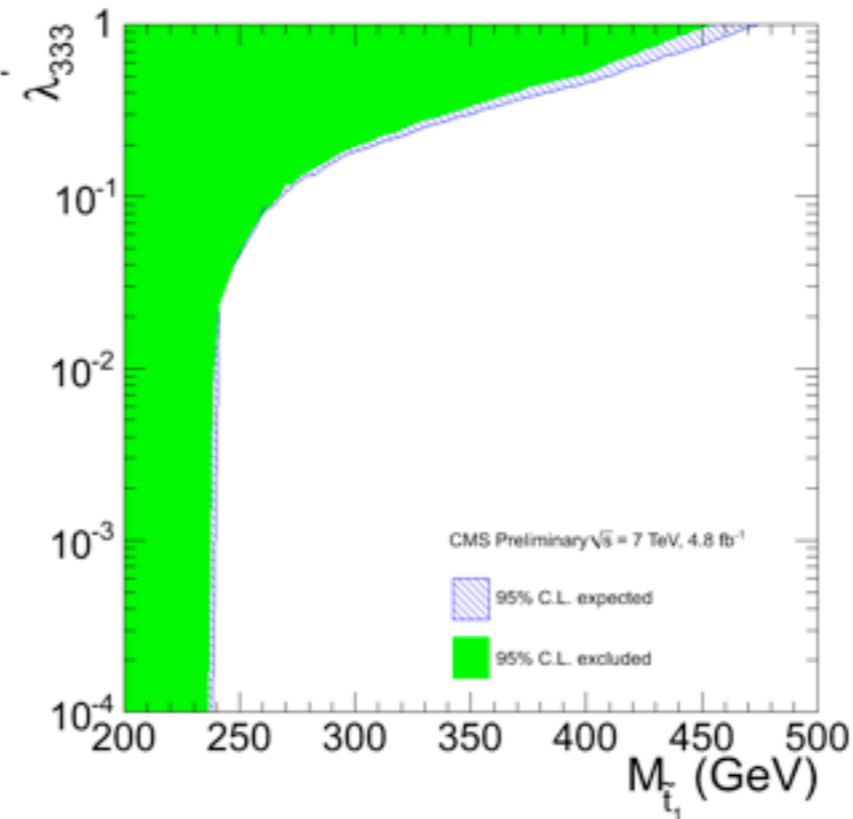
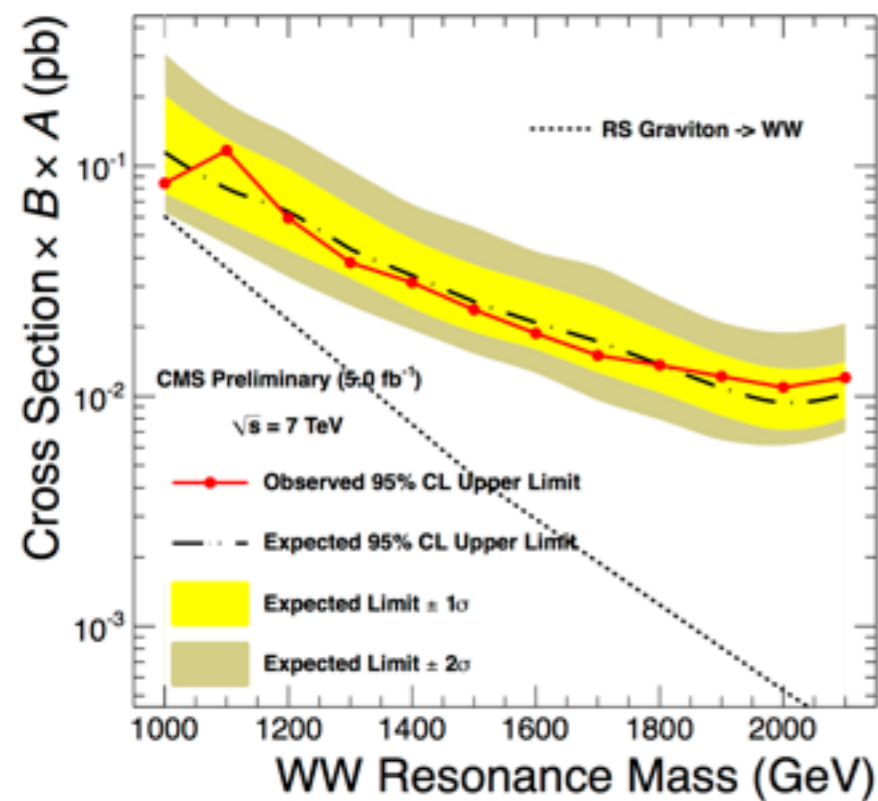
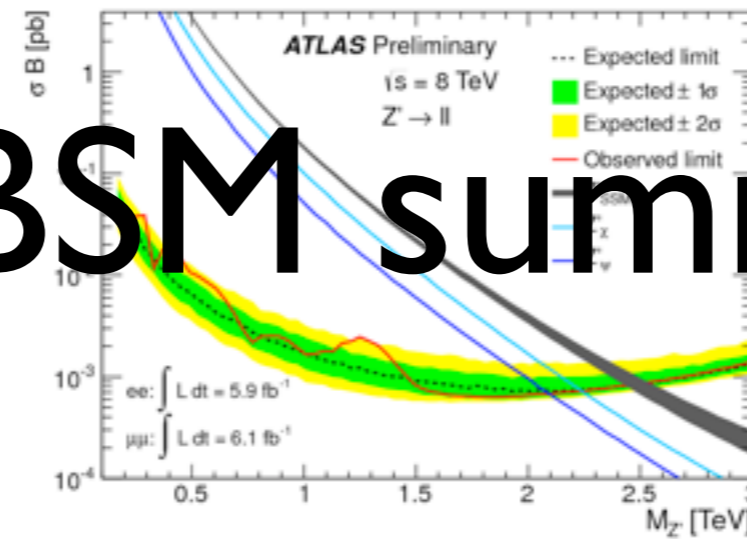


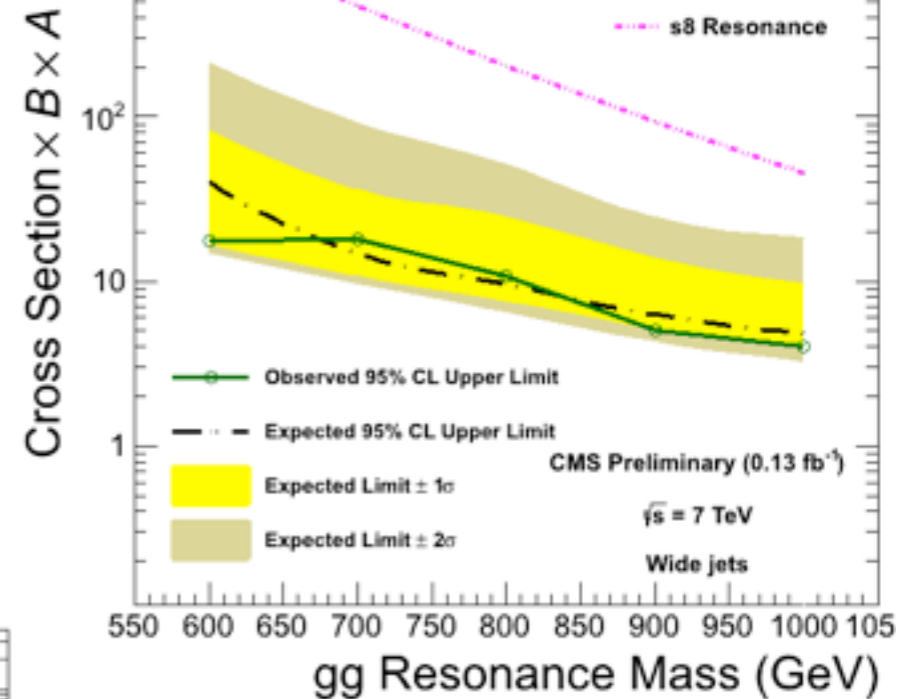
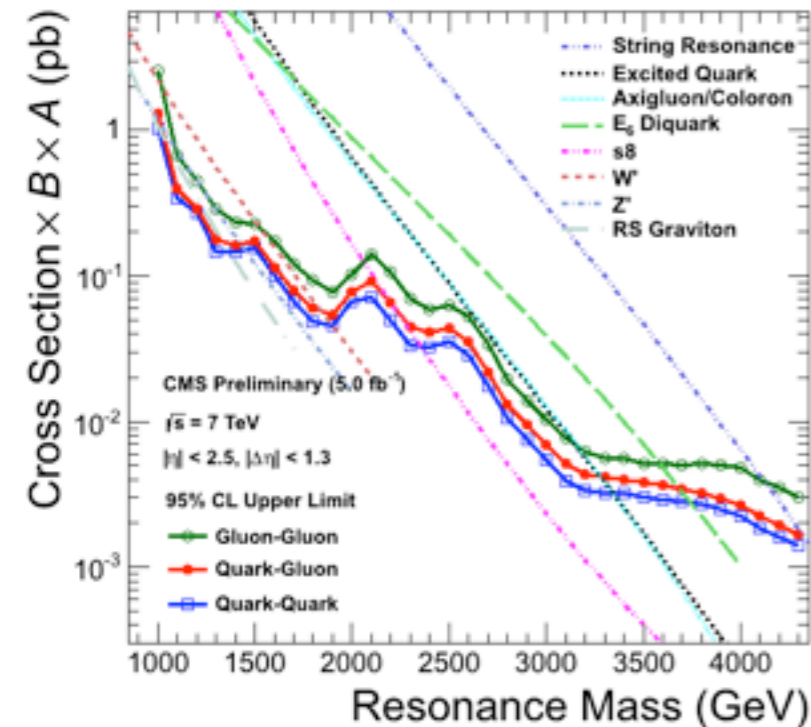
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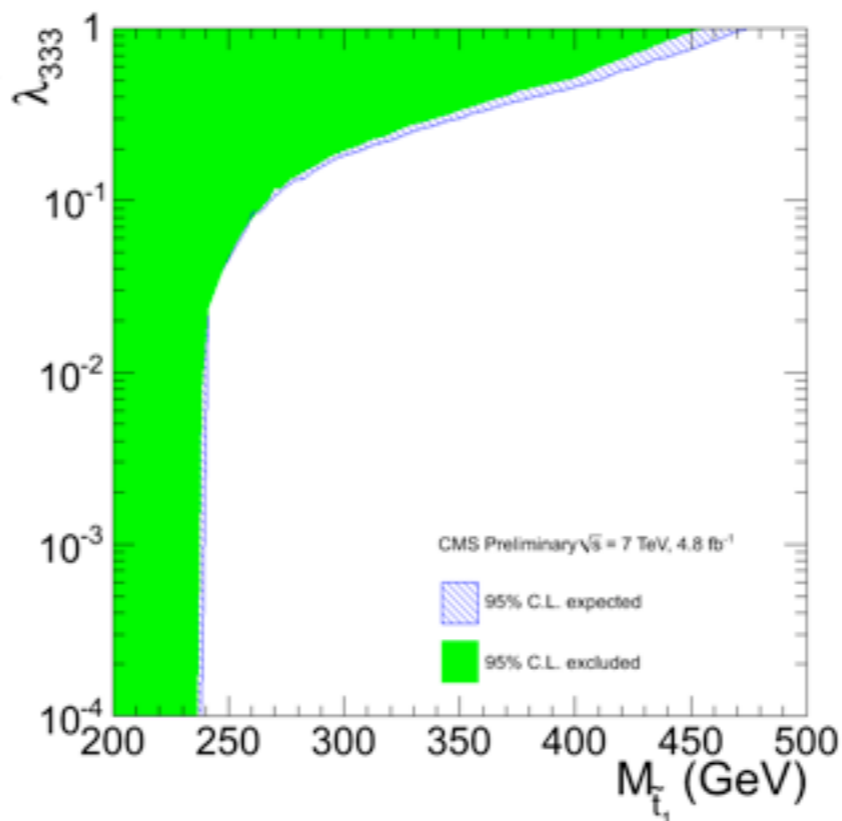
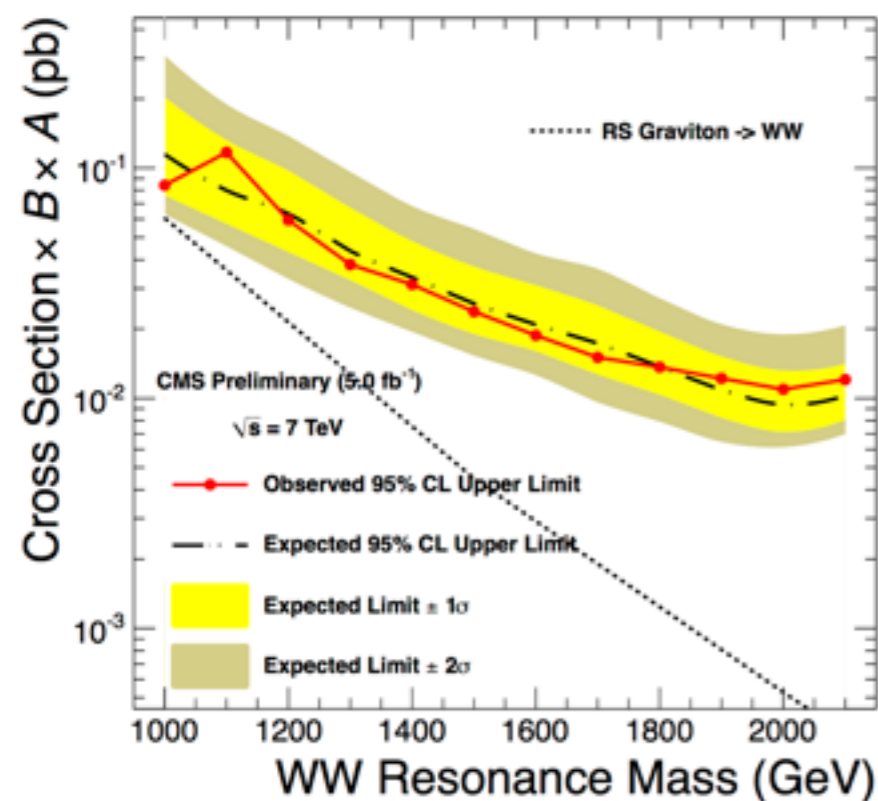
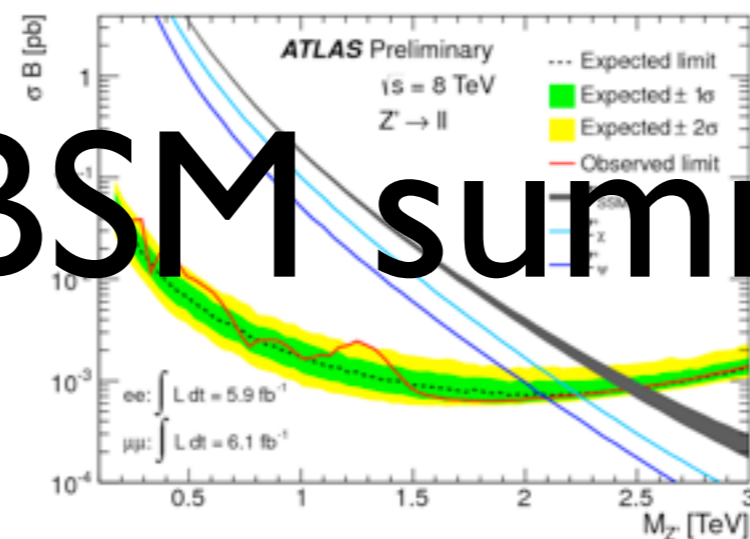


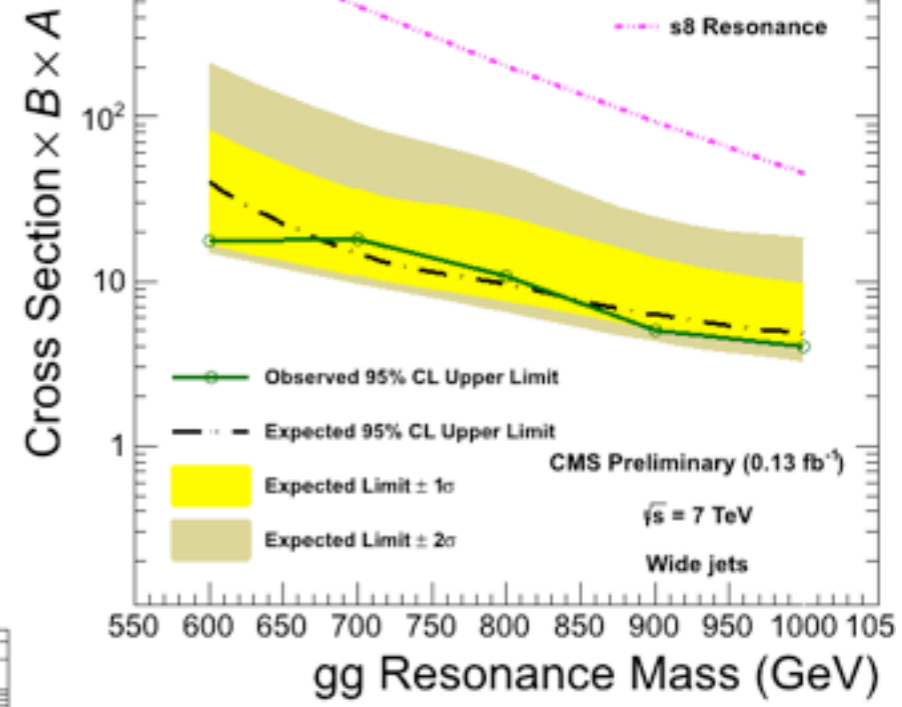
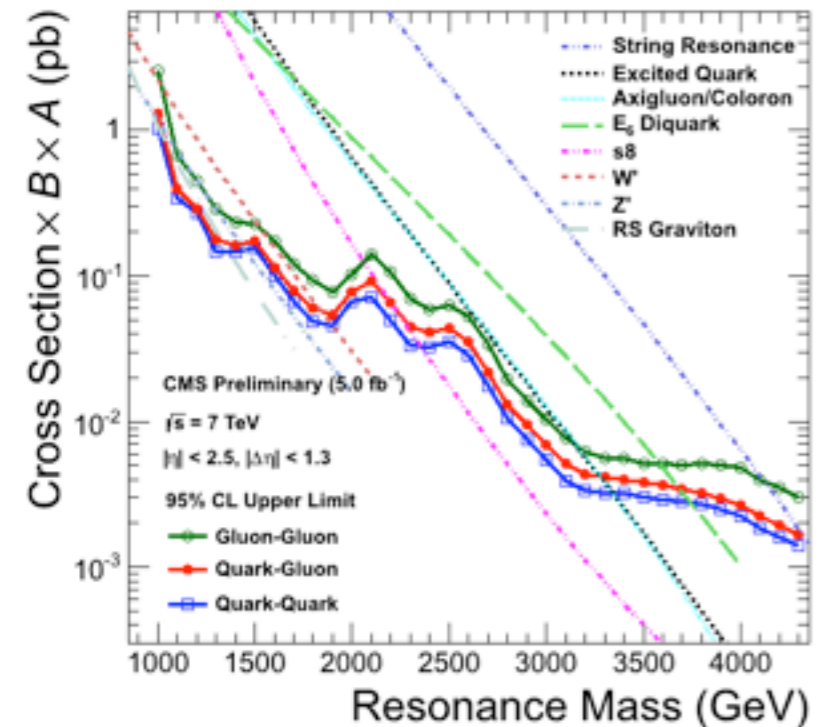
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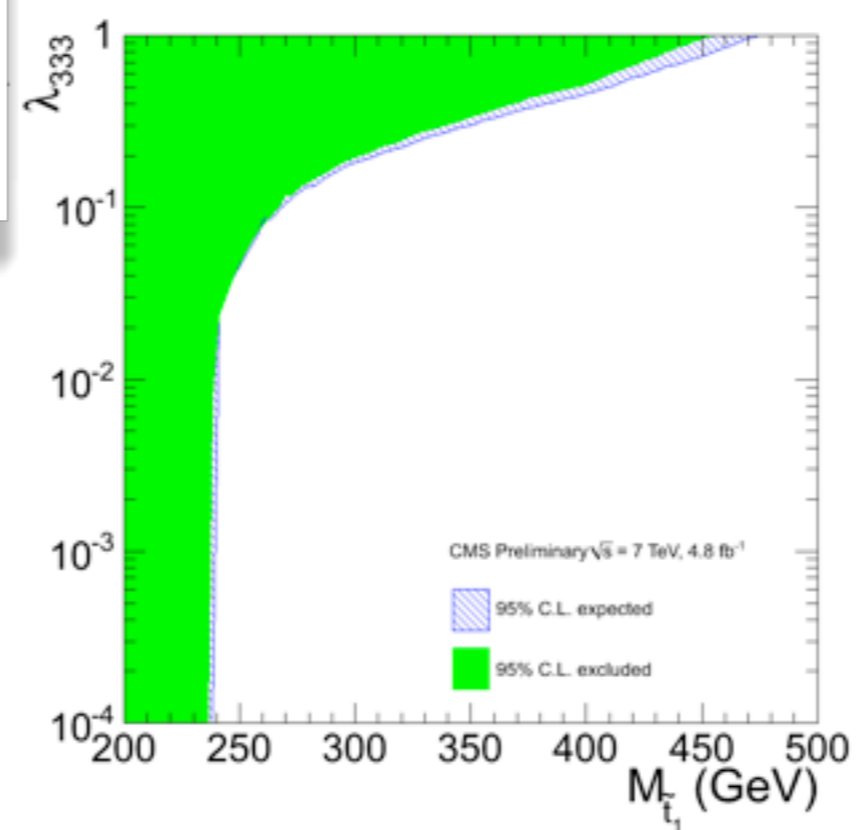
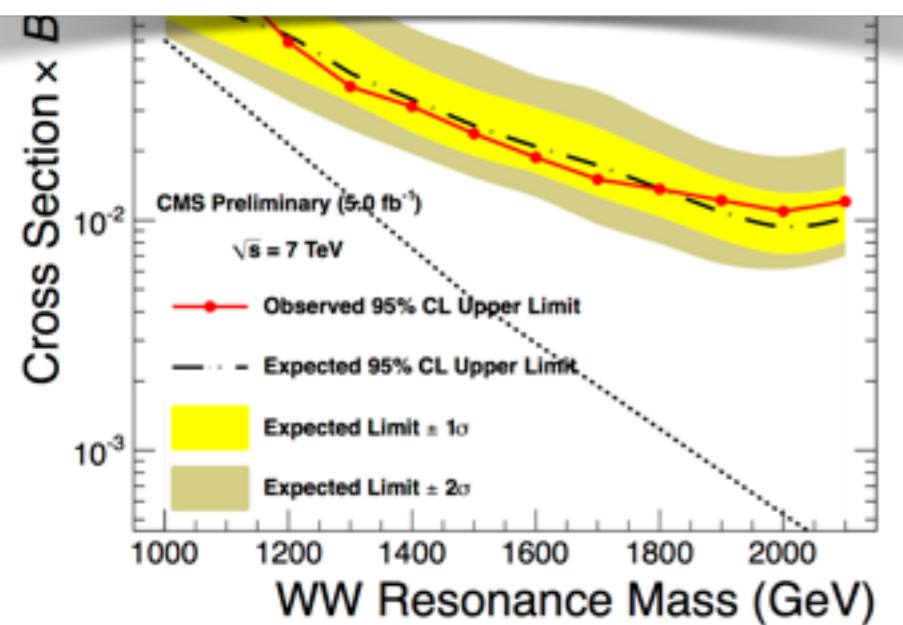
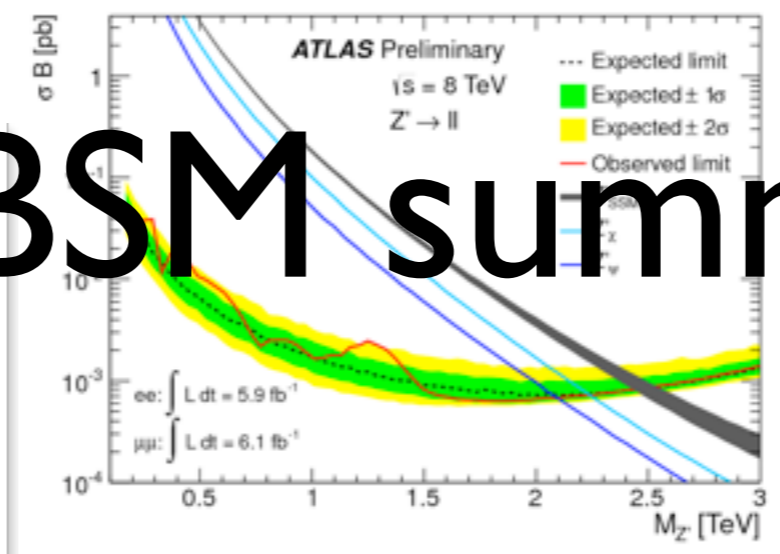
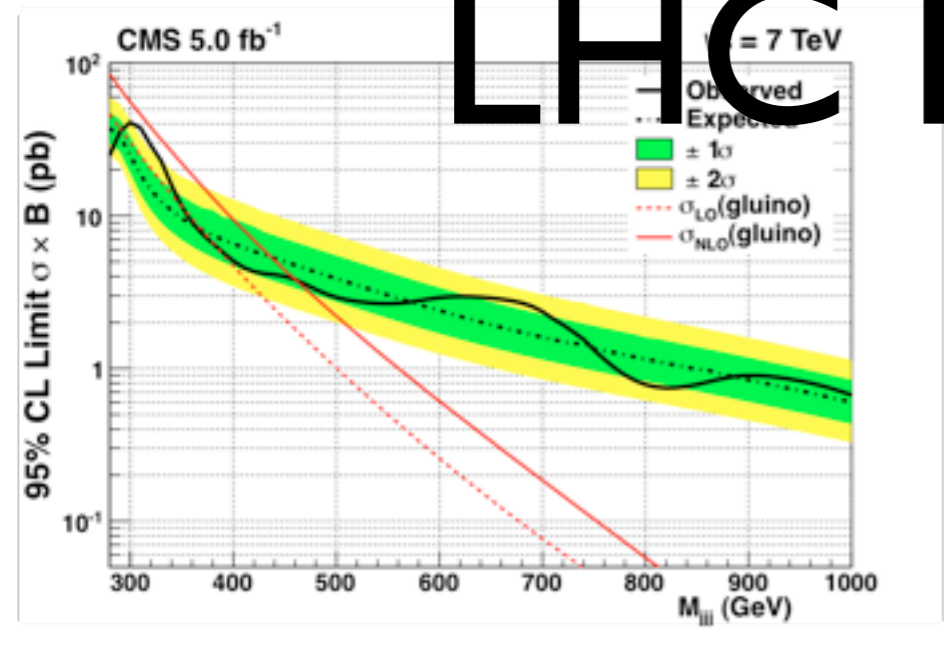


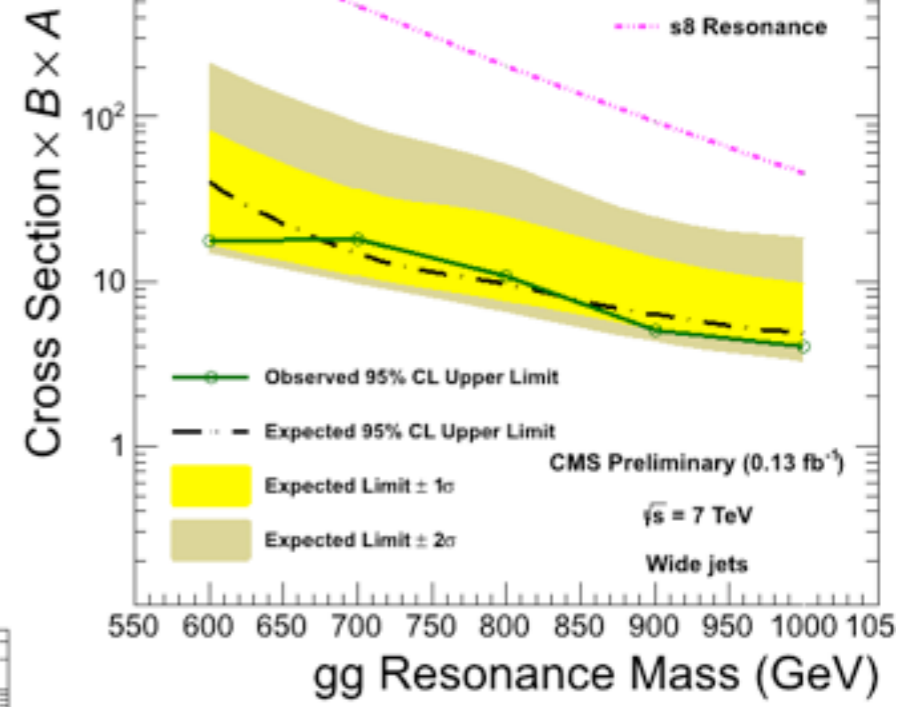
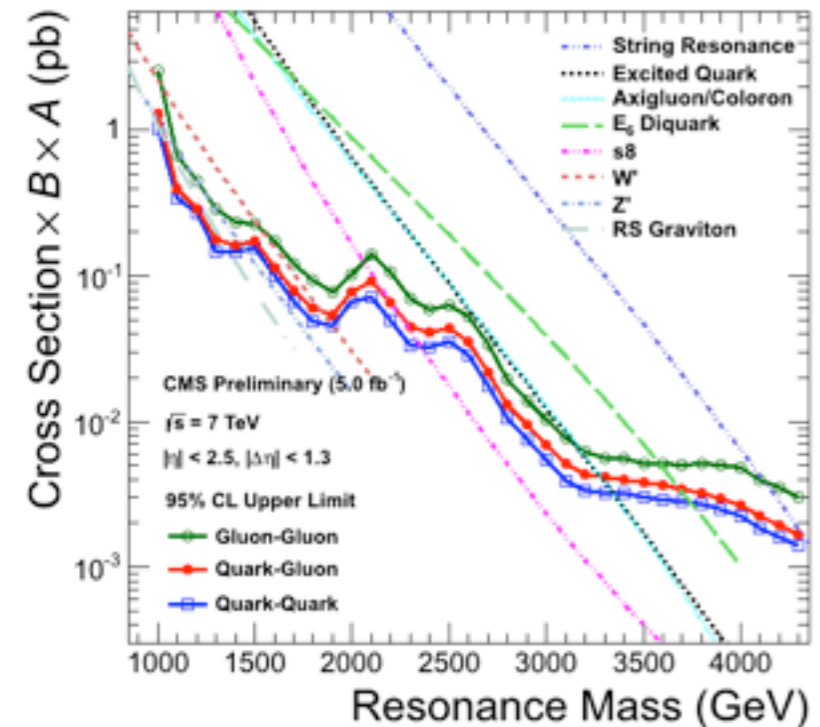
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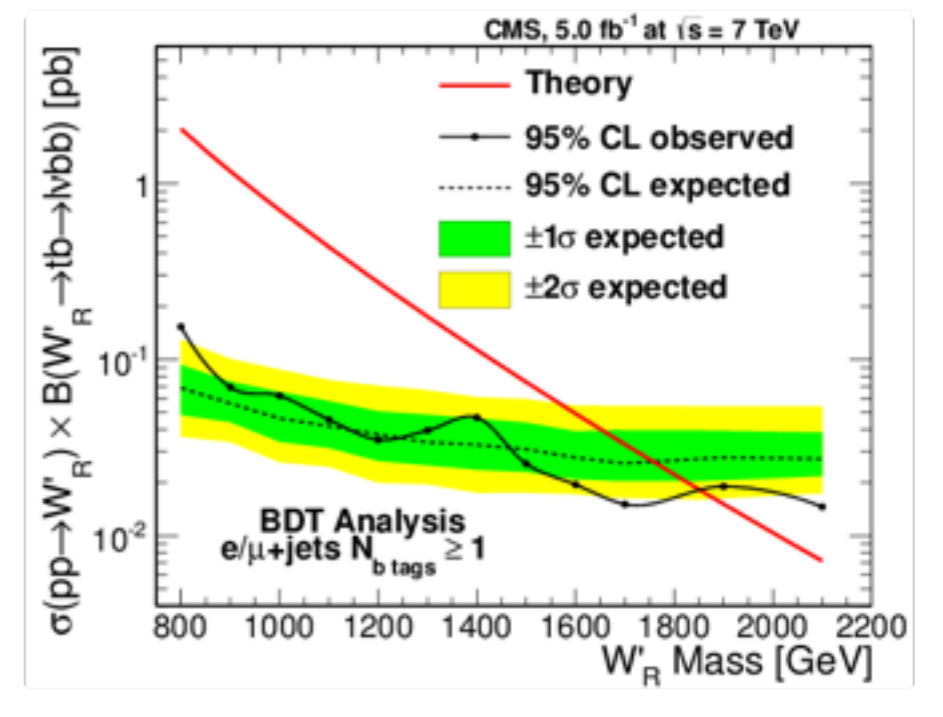
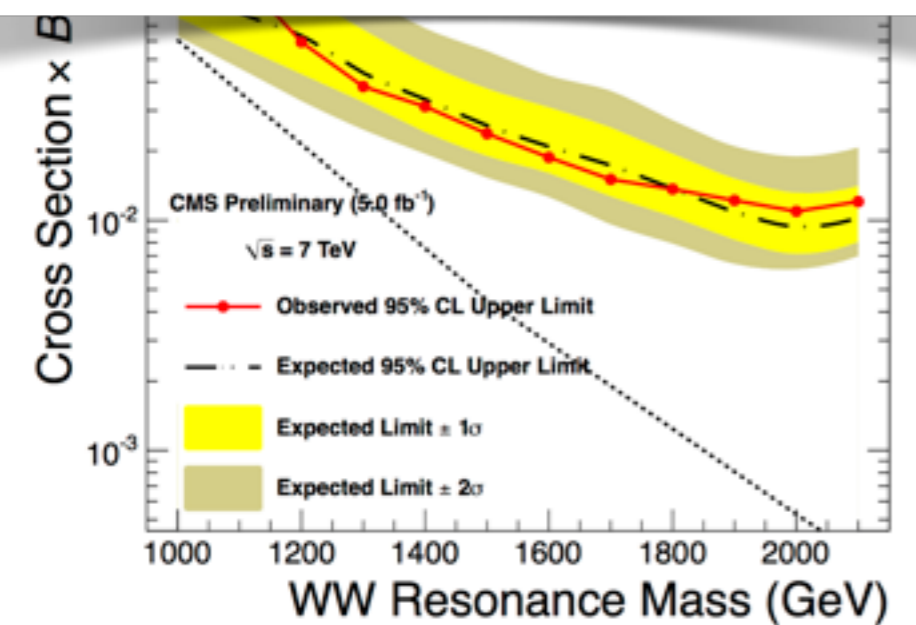
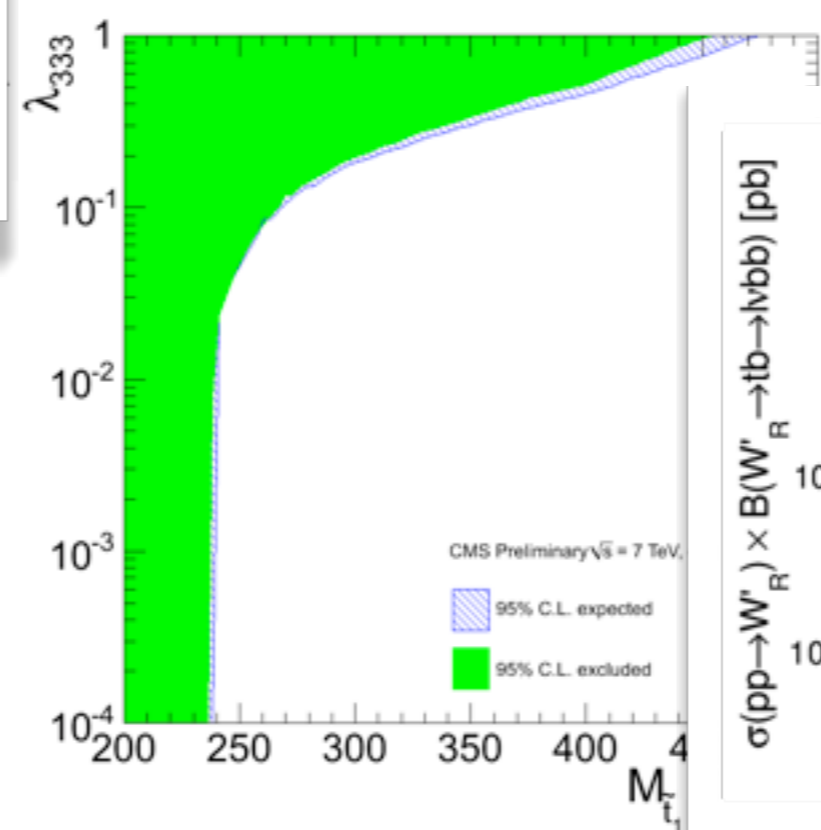
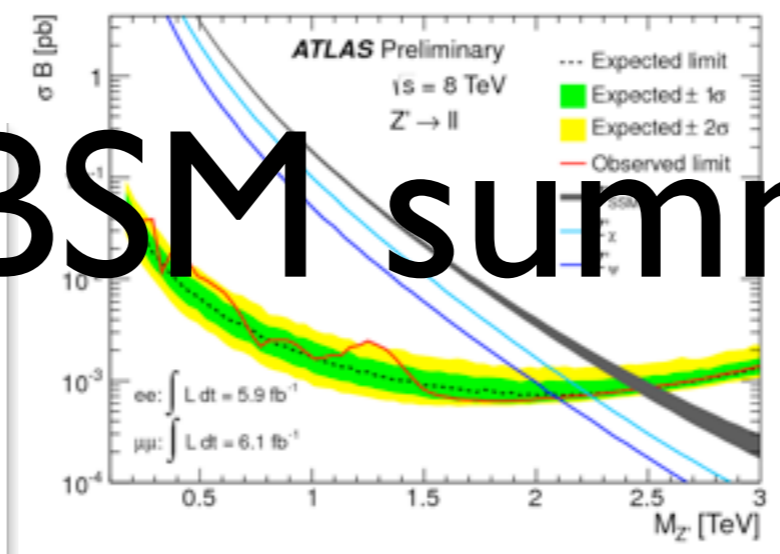
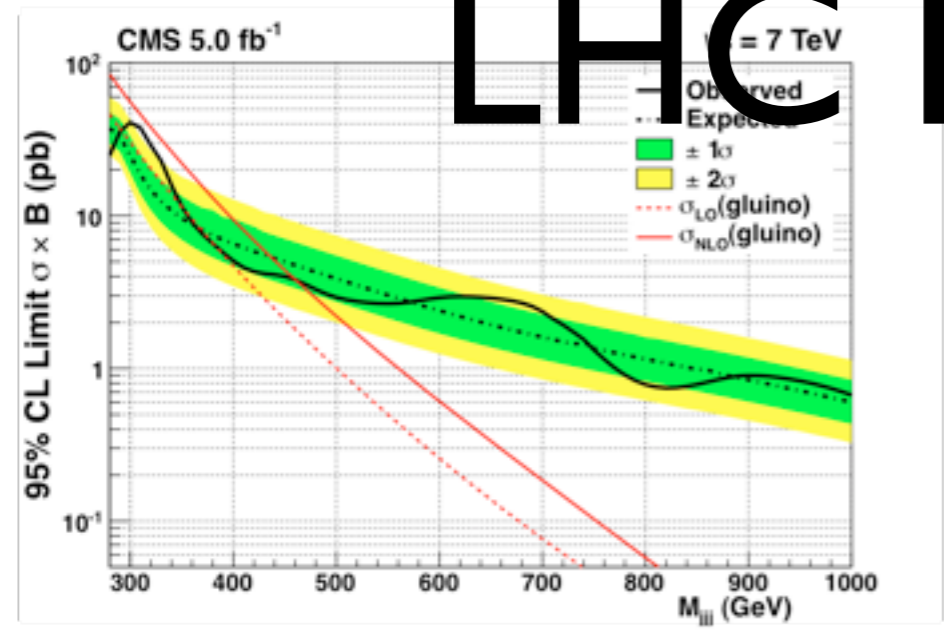


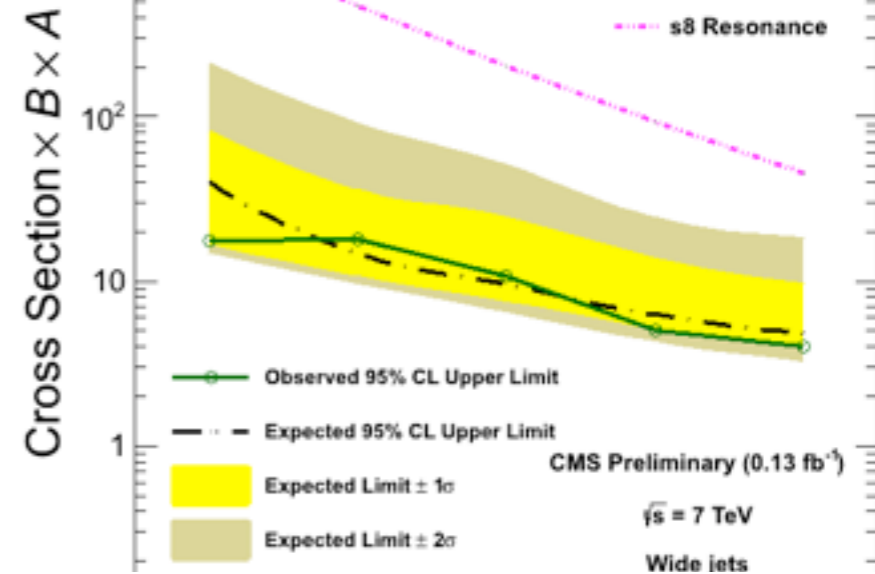
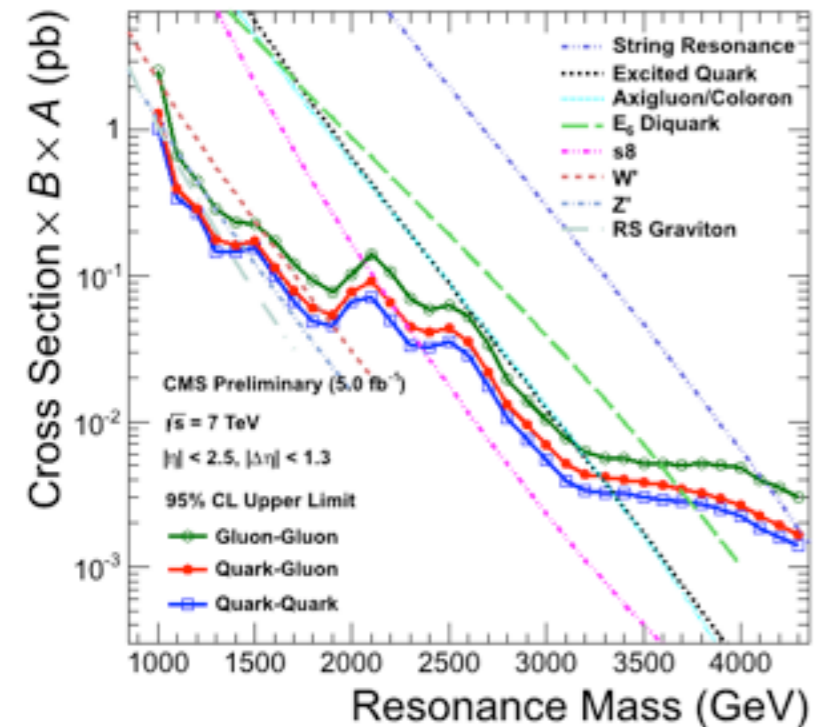
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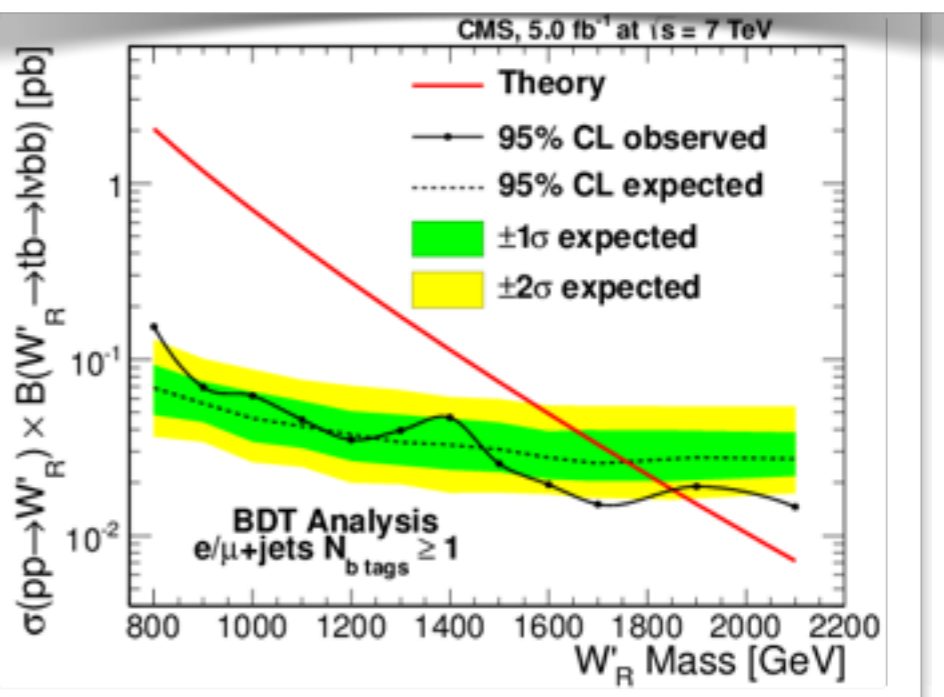
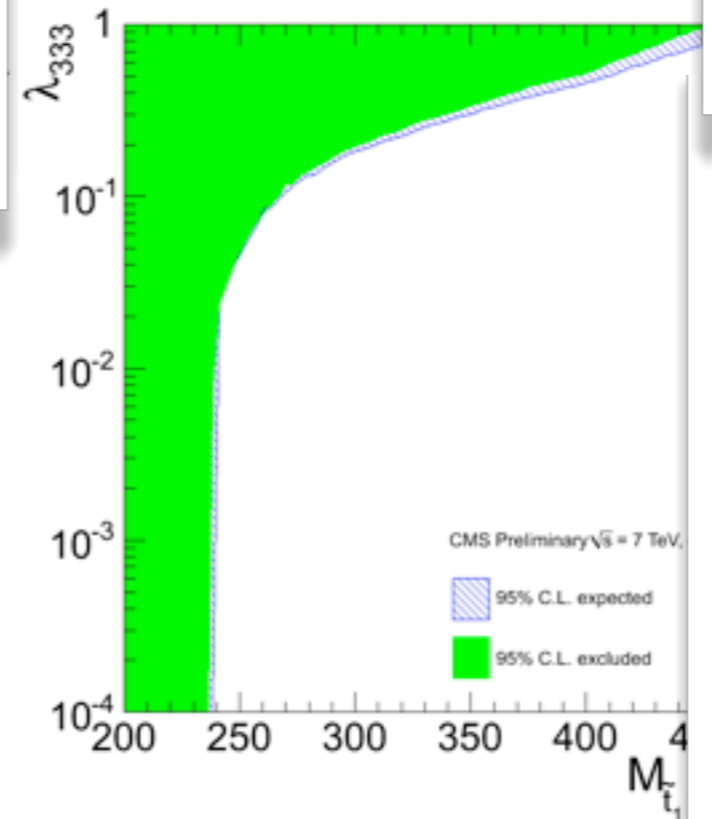
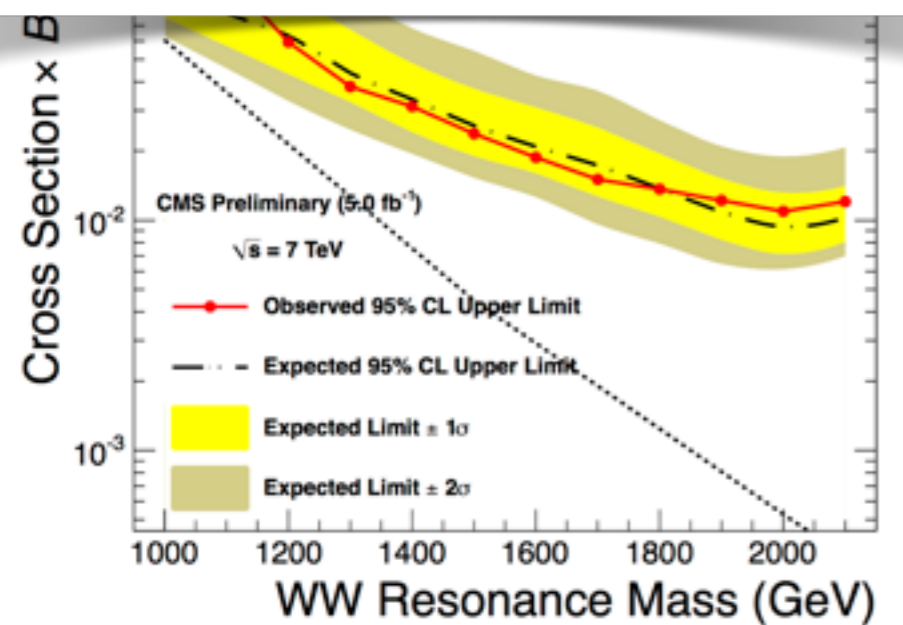
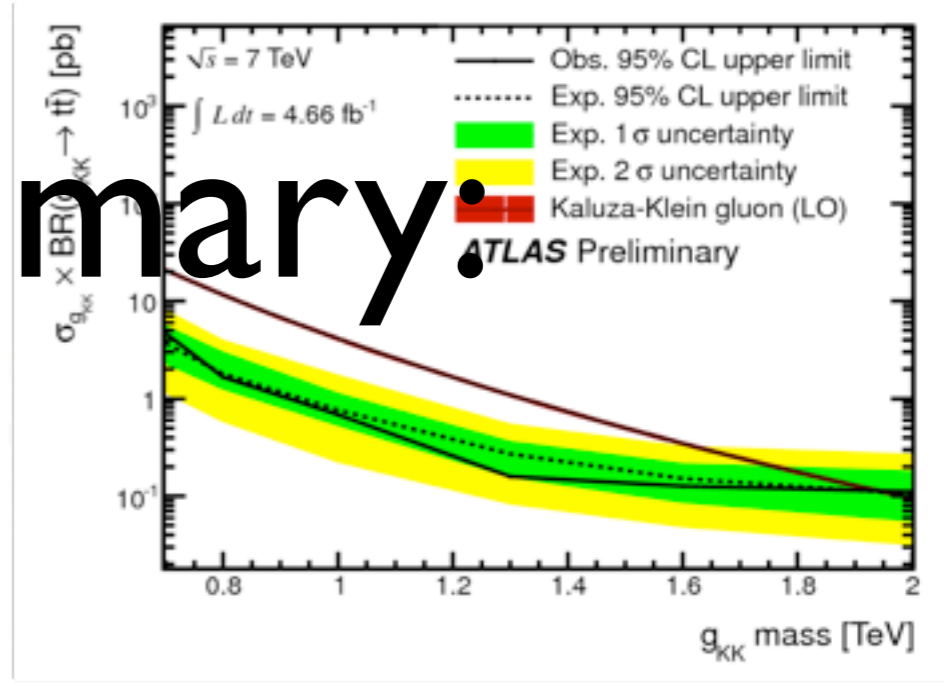
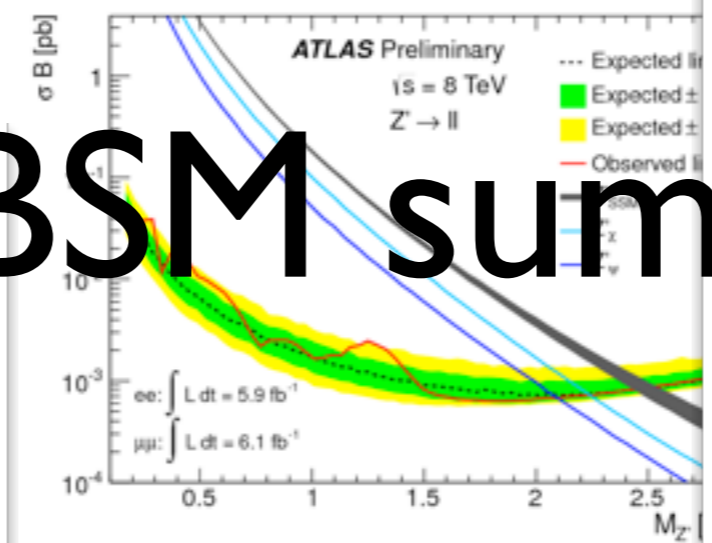
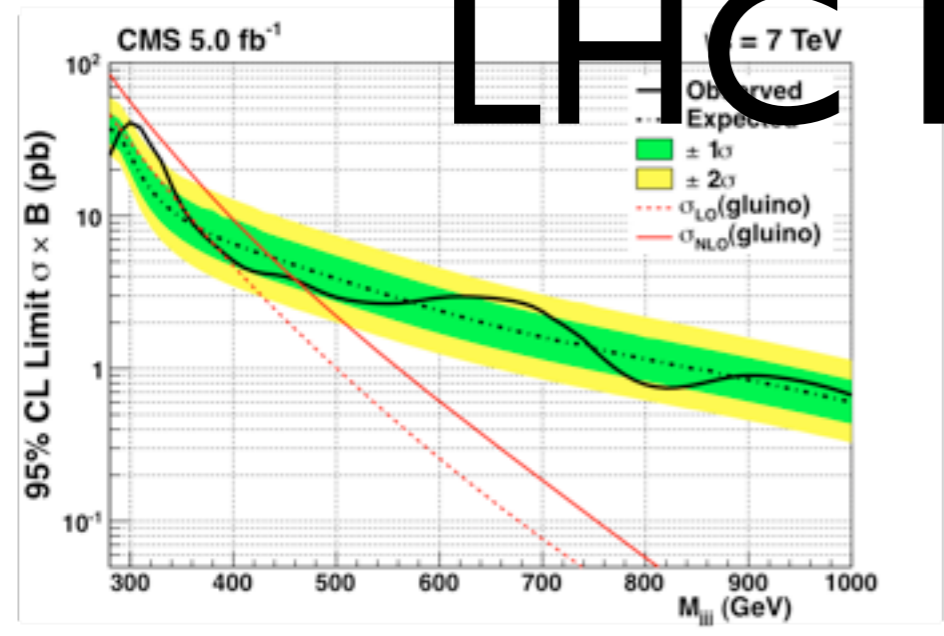


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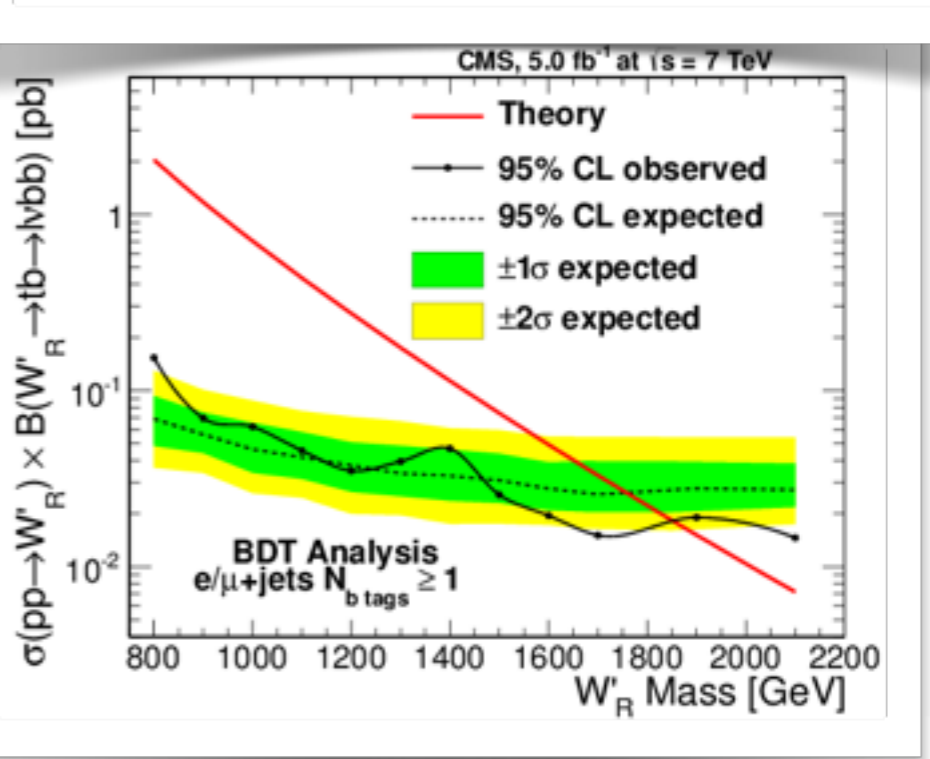
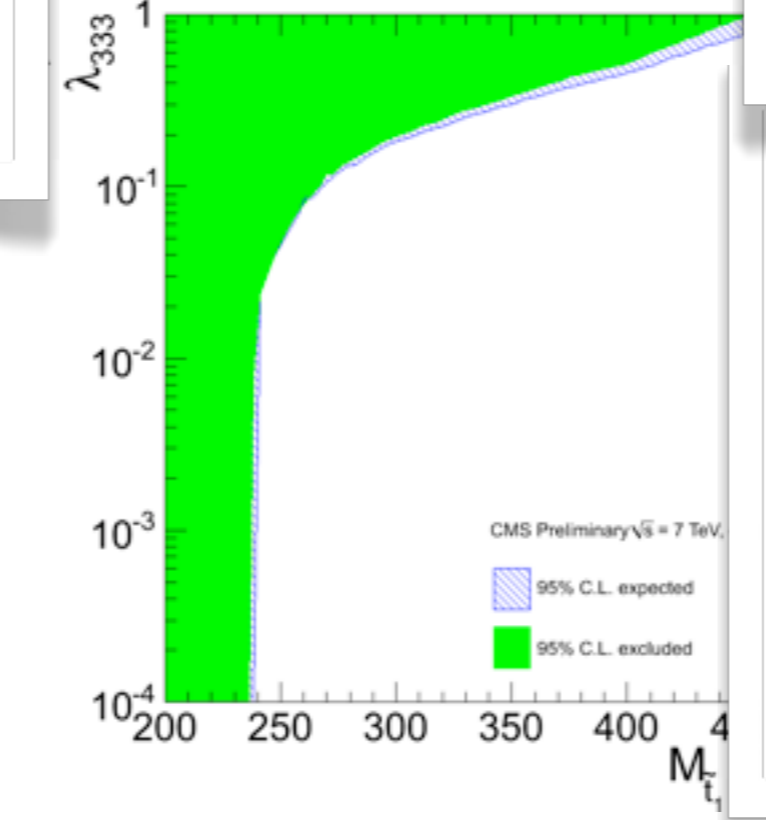
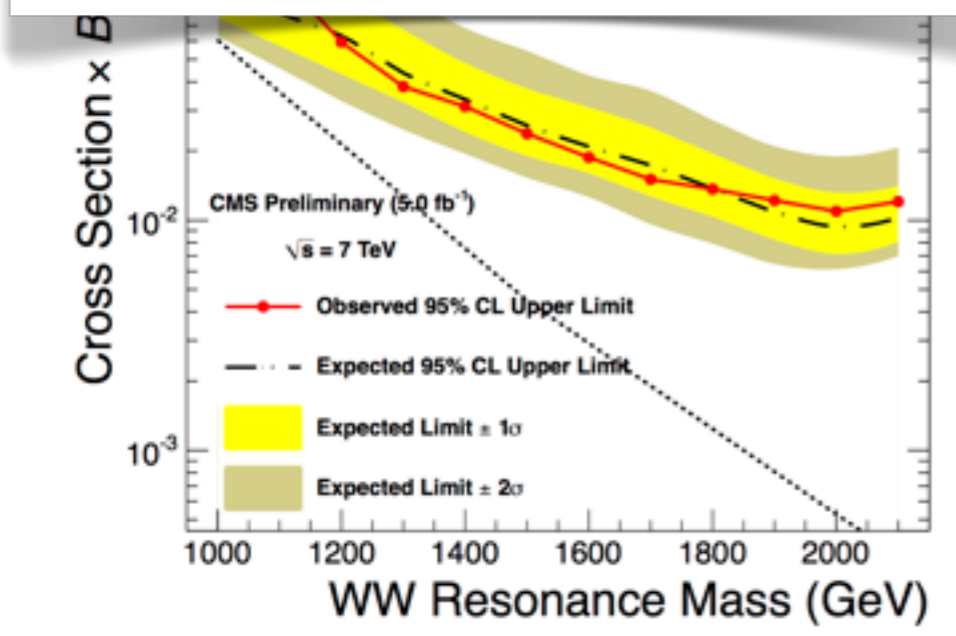
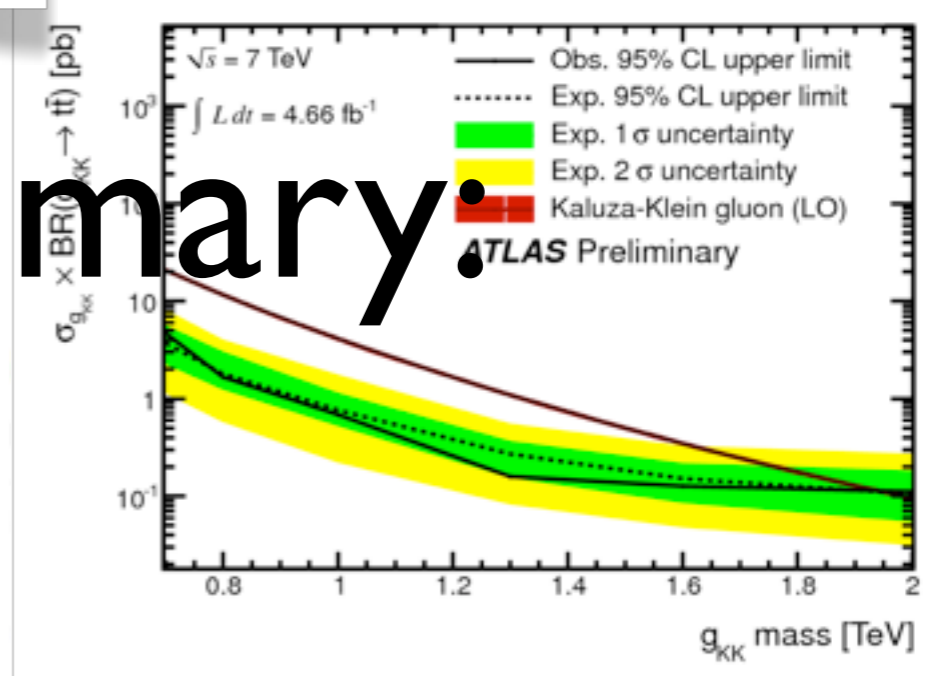
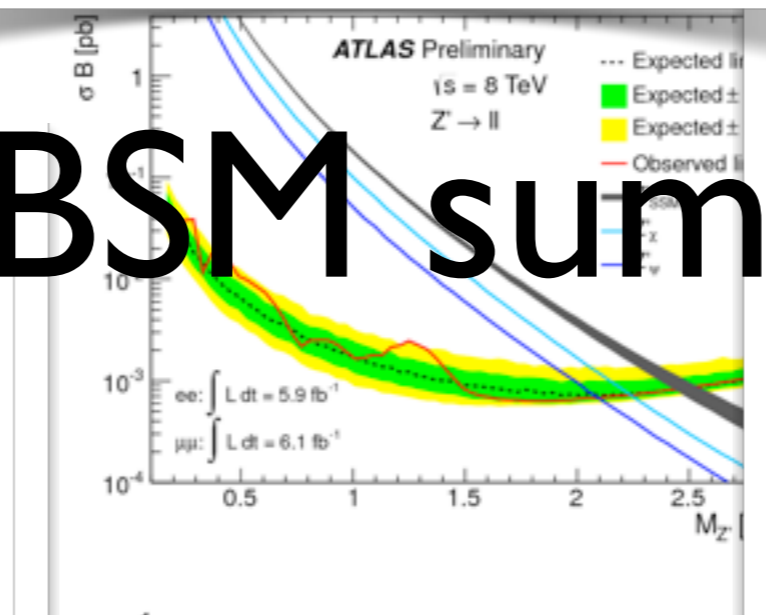
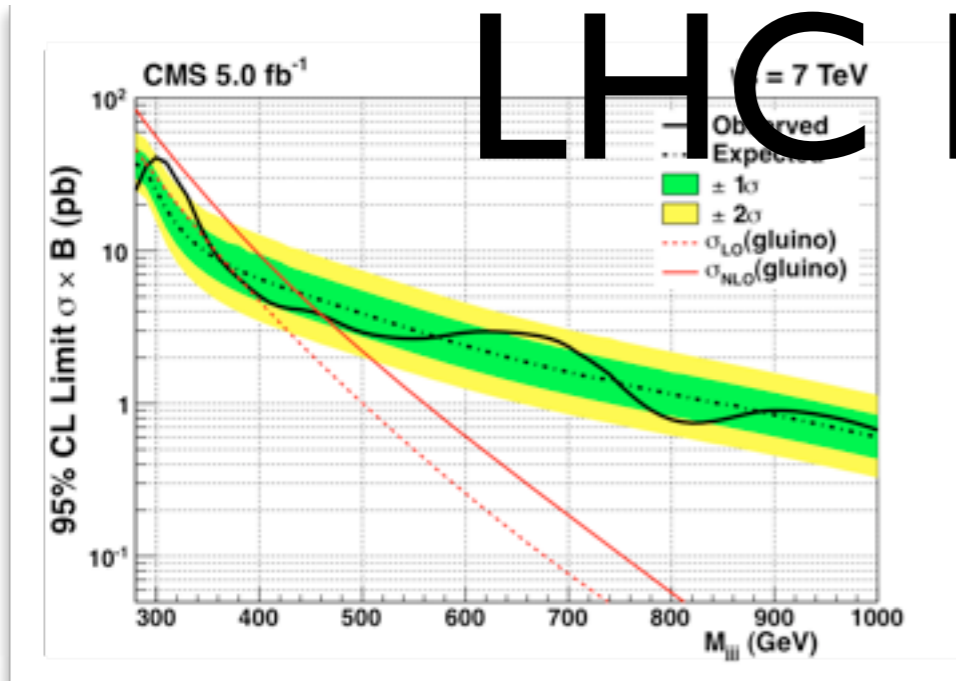
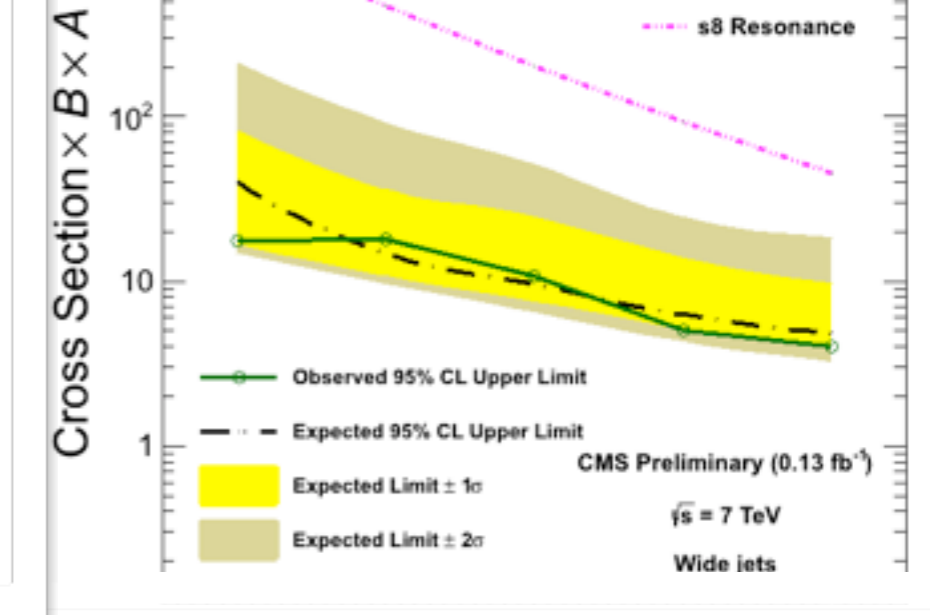
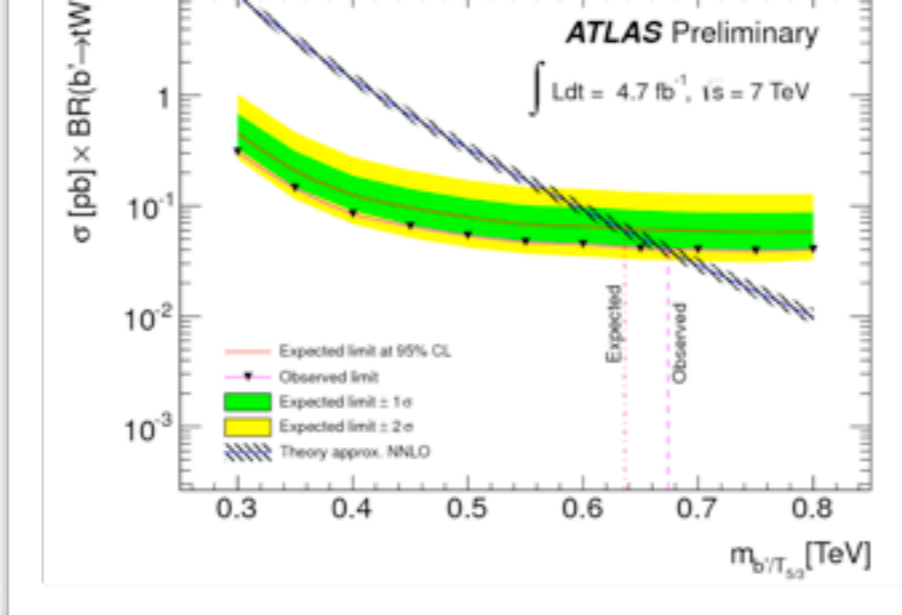
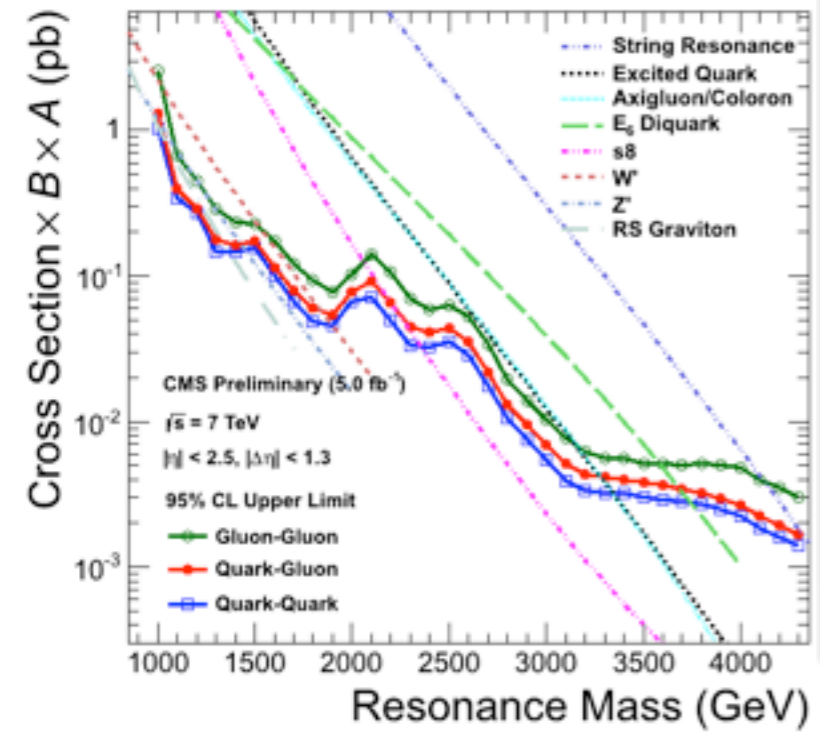


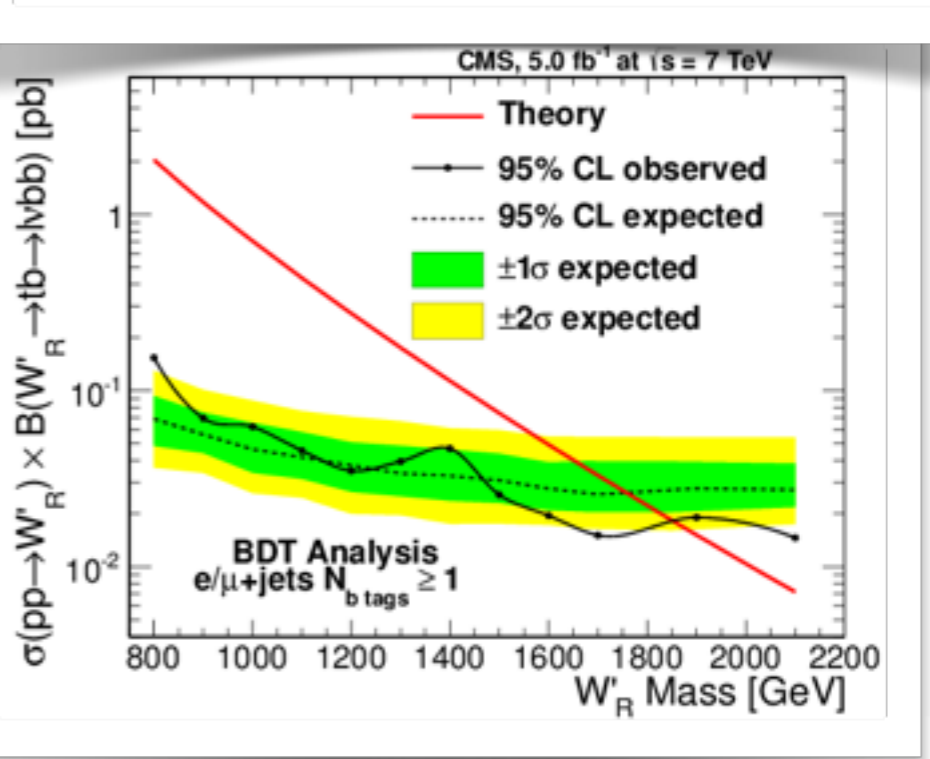
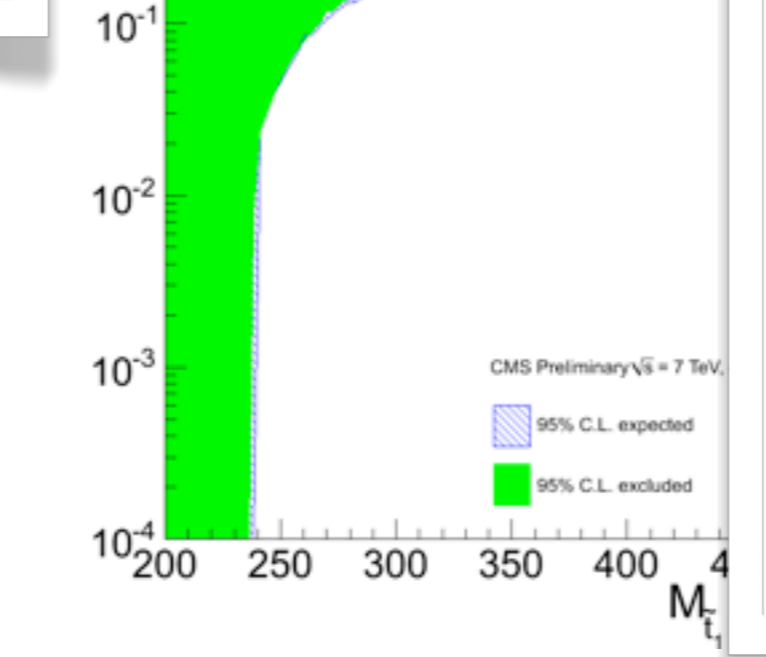
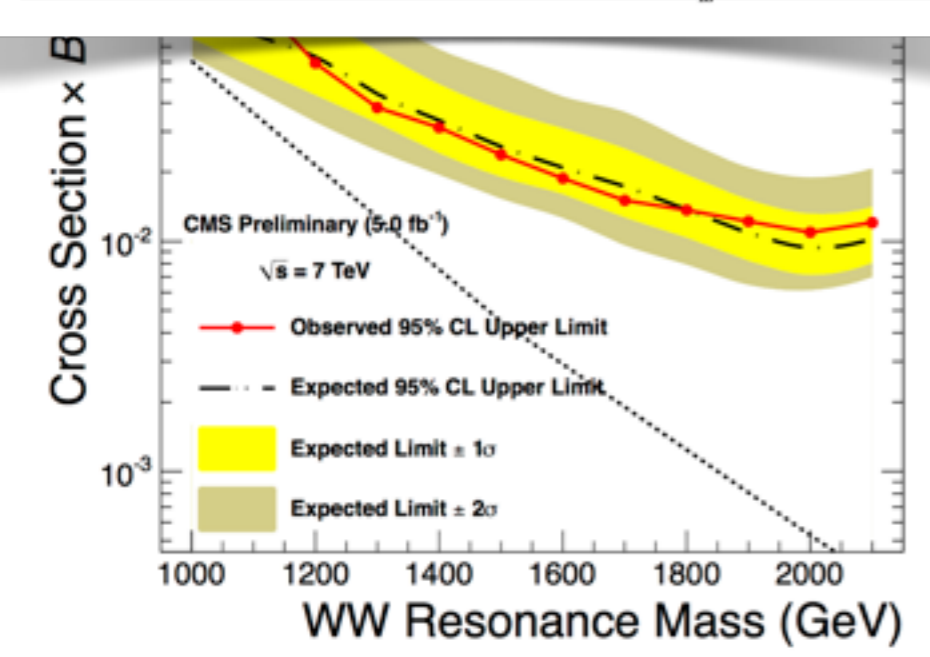
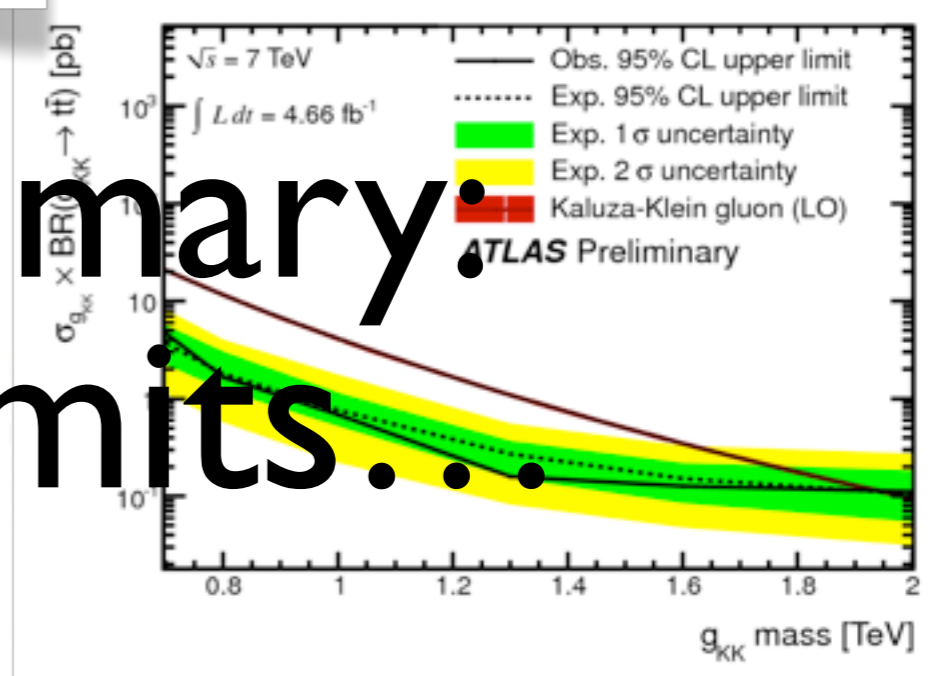
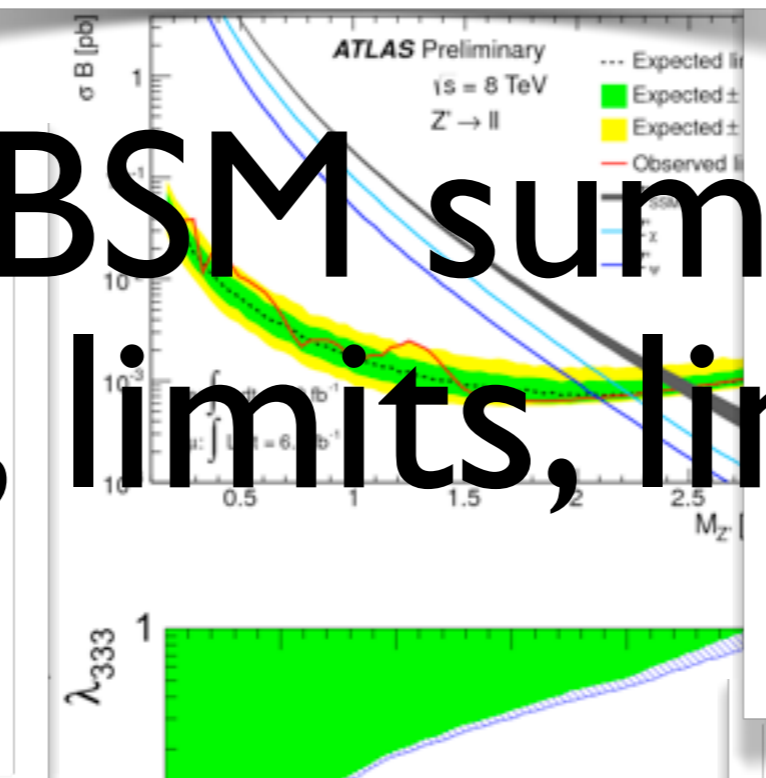
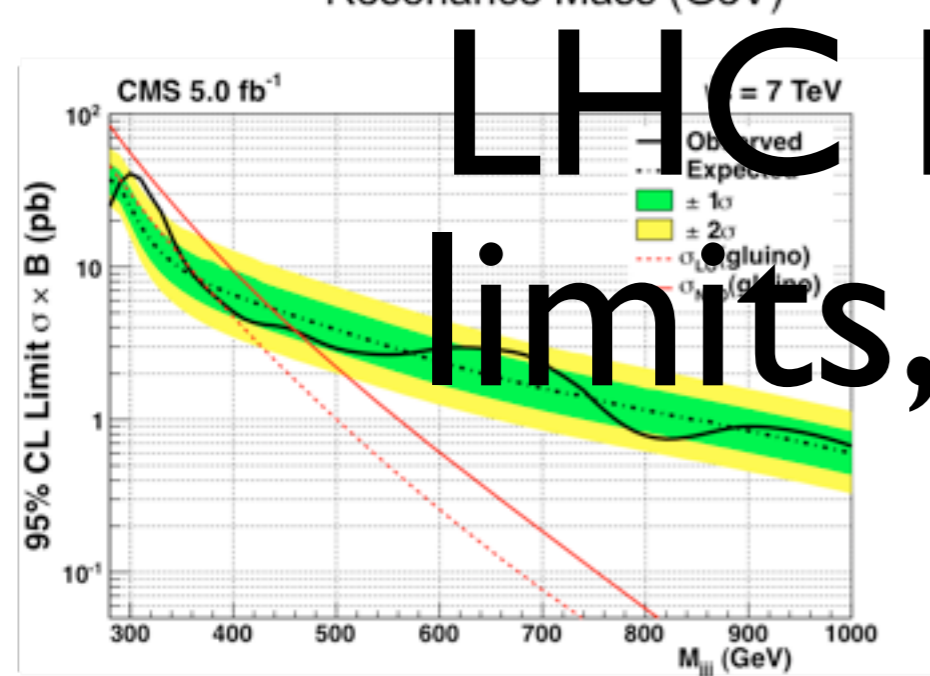
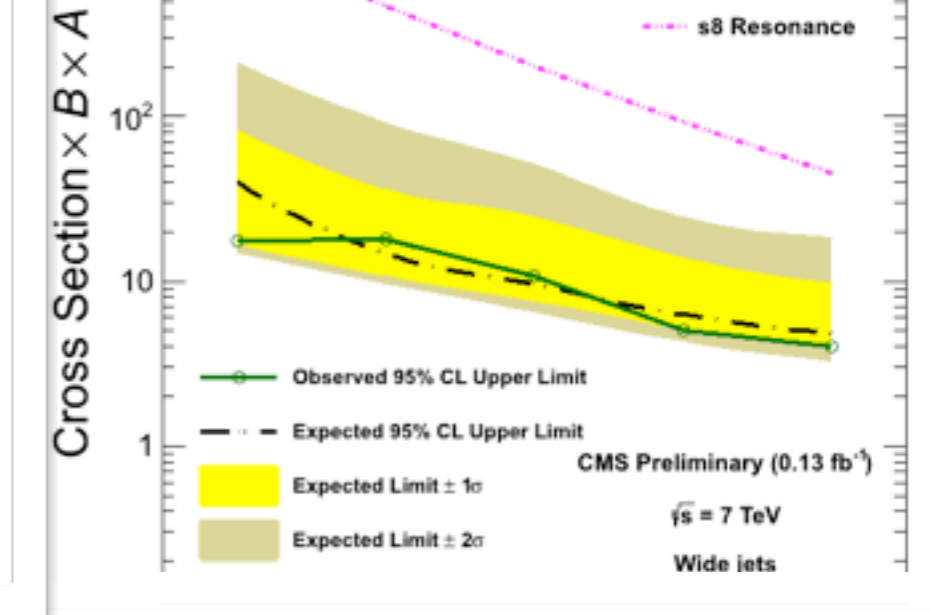
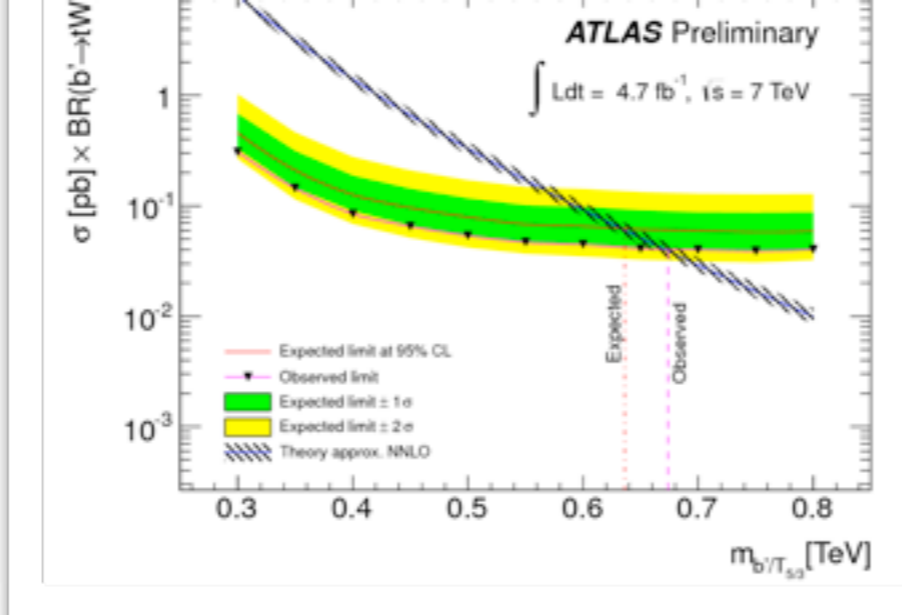
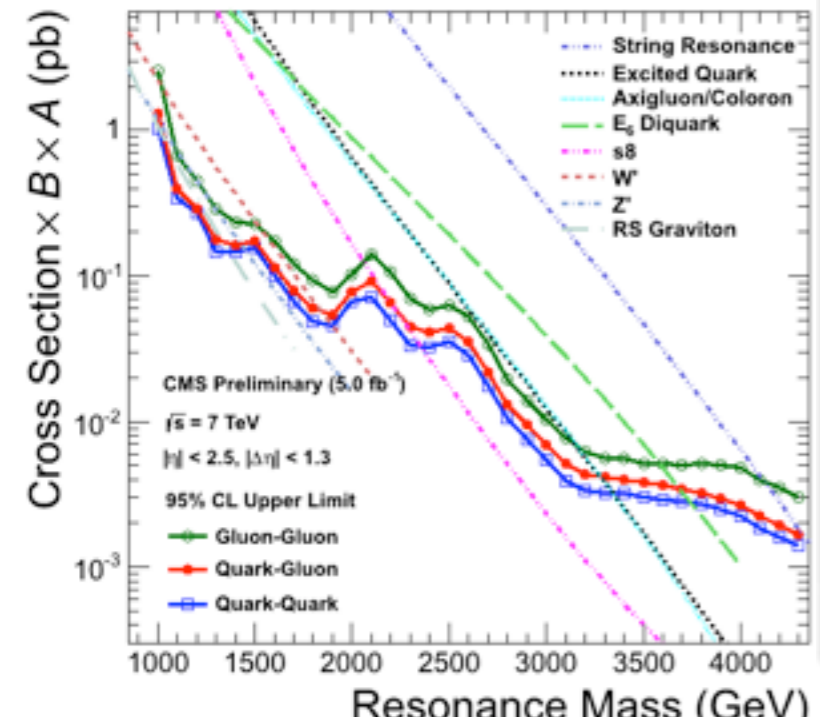


LHC BSM summary:



LHC BSM summary:





**LHC BSM summary:
 limits, limits, limits...**

Why go beyond the
SM?

Dark matter?

Dark Energy?

Origin of quark mass and mixing hierarchies?

Strong CP?

EW strong coupling/unitarity problem?

Matter-Antimatter asymmetry?

Neutrino masses?

Inflation?

Quantum instability of Higgs mass?

Charge quantization (GUT?)?

Quantum Gravity?

...



**Why expect new
physics
at the LHC?**

Dark matter? 10^{-15} GeV ? 10^{12} GeV ?

~~Dark Energy?~~

Origin of quark mass and mixing hierarchies?

~~Strong CP?~~

EW strong coupling/unitarity problem

Matter-Antimatter asymmetry? 100 GeV? 10^{13} GeV ?

Neutrino masses? 10^{13} GeV ? 100 GeV?

~~Inflation?~~

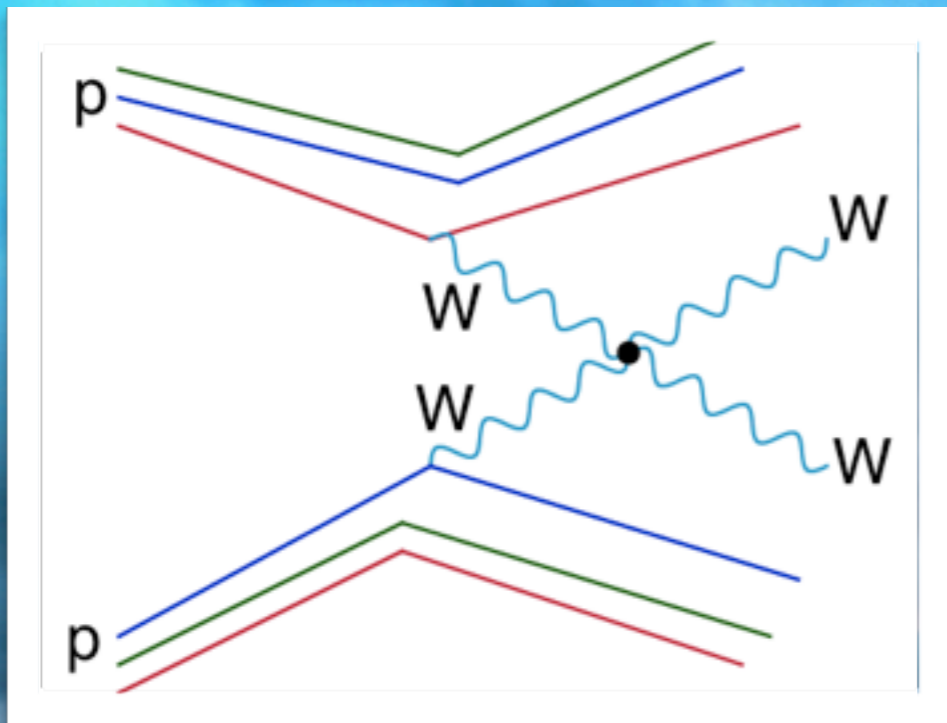
Quantum instability of the Higgs mass

~~Charge quantization (GUT?)?~~

Quantum Gravity? TeV or M_{Planck} ...

...

Before the Higgs discovery

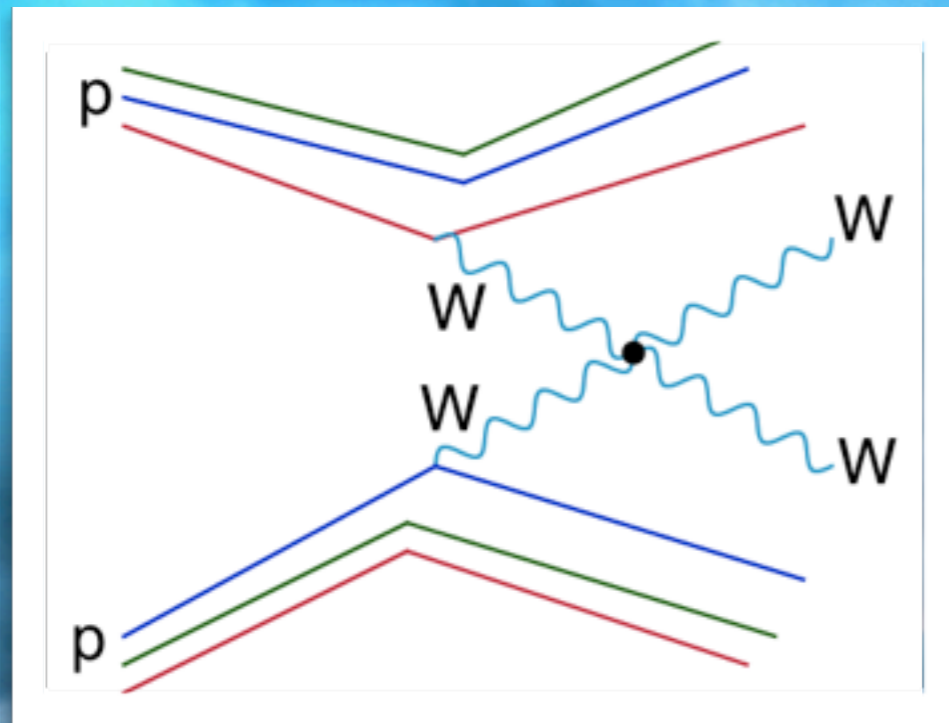


$$\Lambda \approx 4\pi v \approx 3 \text{ TeV}$$

Energy

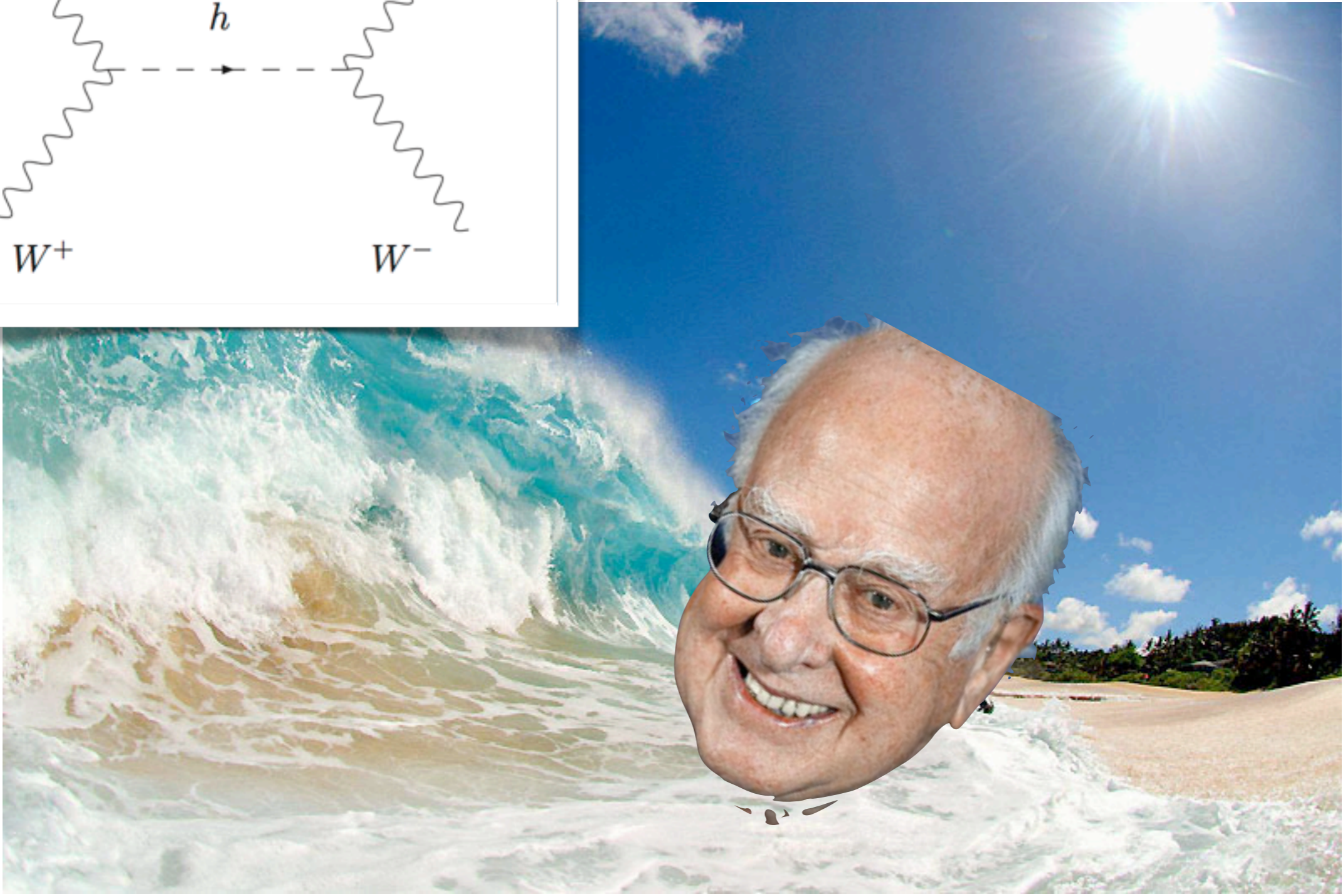
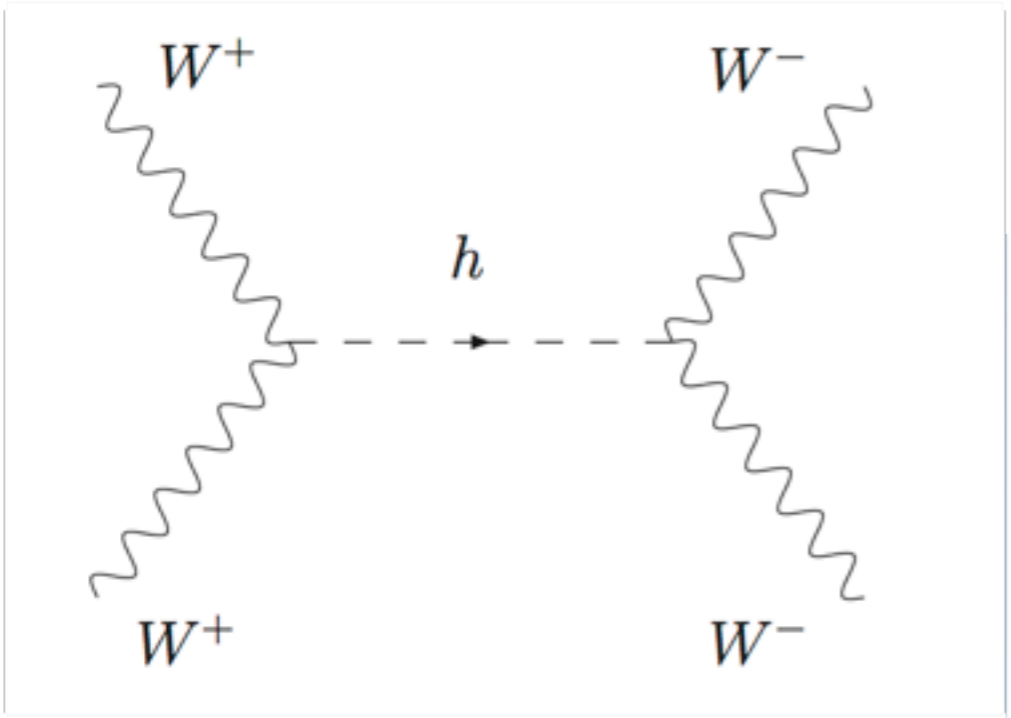


Before the Higgs discovery



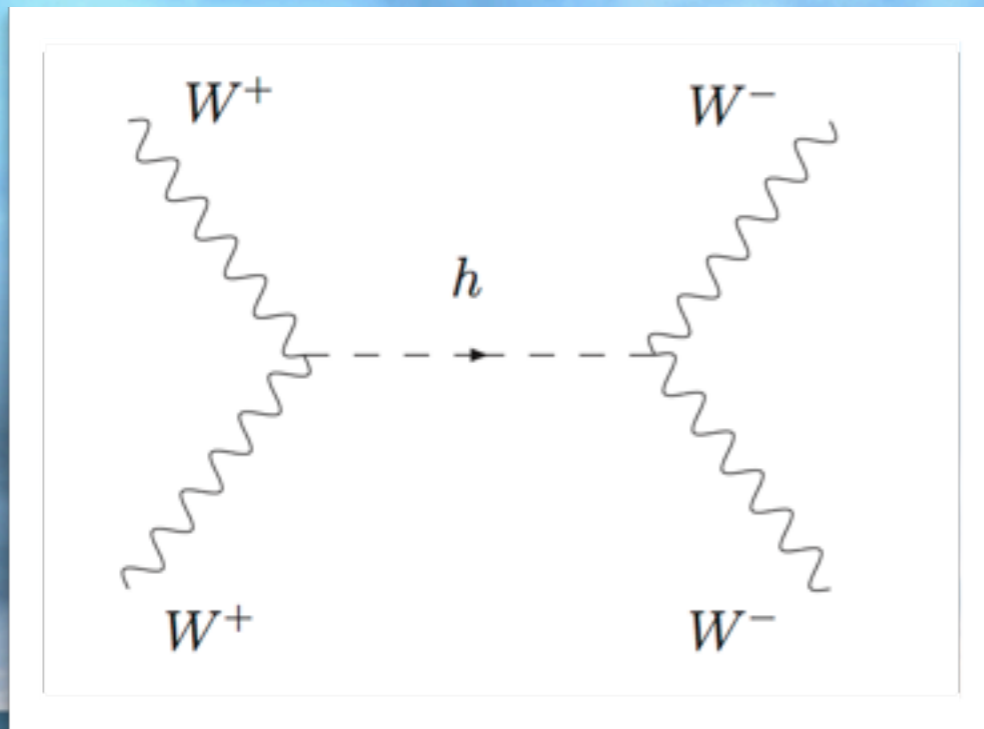
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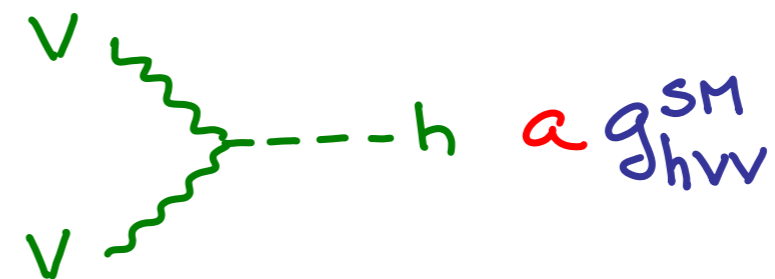
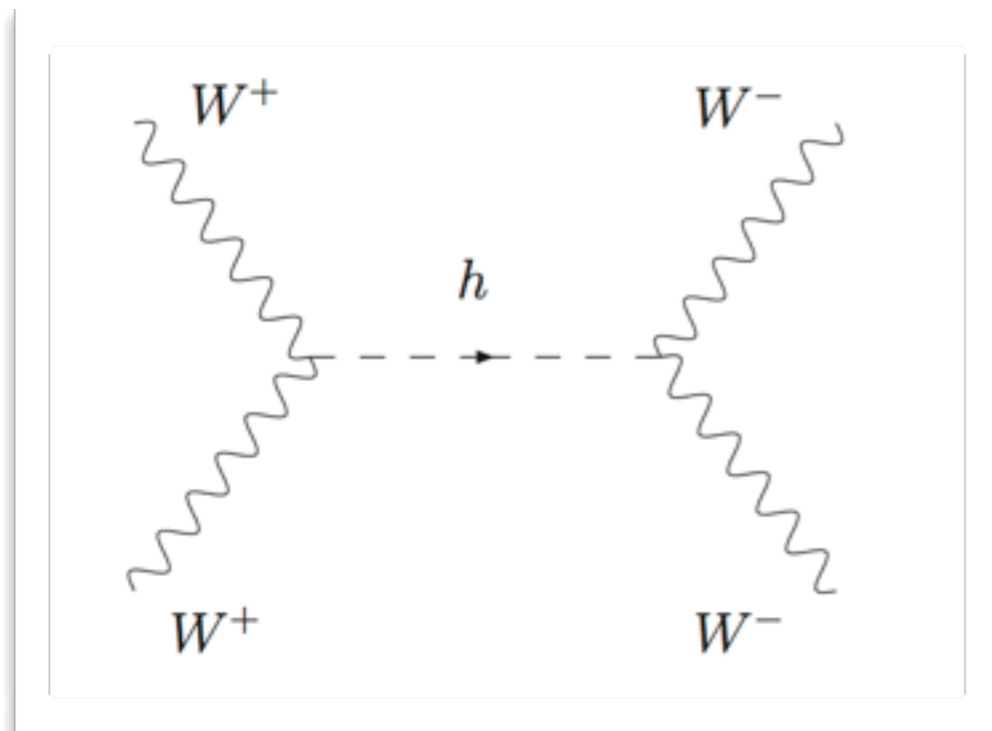
Adding SM-like Higgs

SM works up to $\Lambda \gg \text{LHC}$



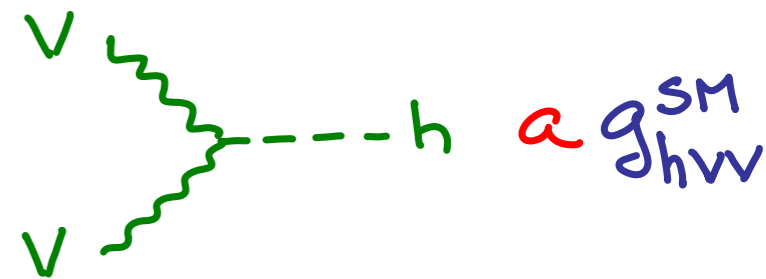
Adding SM-like Higgs

What if the coupling is not exactly like in the SM?



$$\Lambda \approx 4\pi v \longrightarrow \frac{4\pi v}{\sqrt{1 - a^2}}$$

$$\Lambda \approx 4\pi v \longrightarrow \frac{4\pi v}{\sqrt{1 - a^2}}$$



Even if we measure $a < 1$, no guarantee for new physics in reach of LHC.

Example: composite pseudo-Goldstone Higgs:

$$a = \sqrt{1 - (v/f)^2} \approx 0.8 \dots 0.9$$

$$\Lambda > 6 \dots 8 \text{ TeV}$$

So what should be our
guiding principle?



“What’s the problem?”

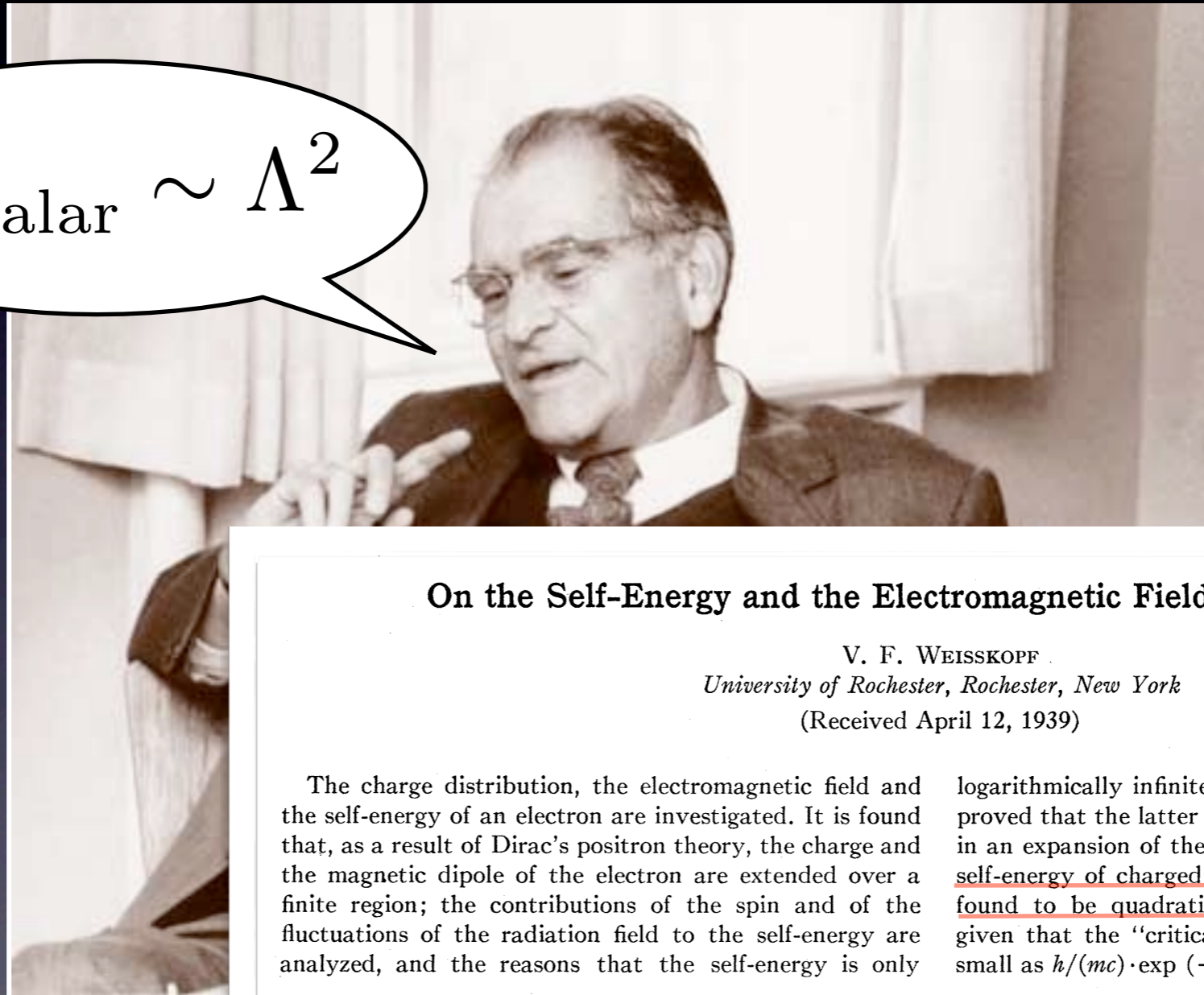
$$m_{\text{scalar}}^2 \sim \Lambda^2$$



Weisskopf, Phys. Rev.56 (**1939**) 72

“What’s the problem?”

$$m_{\text{scalar}}^2 \sim \Lambda^2$$



On the Self-Energy and the Electromagnetic Field of the Electron

V. F. WEISSKOPF

University of Rochester, Rochester, New York

(Received April 12, 1939)

The charge distribution, the electromagnetic field and the self-energy of an electron are investigated. It is found that, as a result of Dirac's positron theory, the charge and the magnetic dipole of the electron are extended over a finite region; the contributions of the spin and of the fluctuations of the radiation field to the self-energy are analyzed, and the reasons that the self-energy is only

logarithmically infinite in positron theory are given. It is proved that the latter result holds to every approximation in an expansion of the self-energy in powers of e^2/hc . The self-energy of charged particles obeying Bose statistics is found to be quadratically divergent. Some evidence is given that the “critical length” of positron theory is as small as $h/(mc) \cdot \exp(-hc/e^2)$.

Weisskopf, Phys. Rev.56 (1939) 72

A light Higgs is unnatural

$$V(h) = \epsilon \Lambda^2 h^2 + \lambda h^4$$

For $\epsilon = \pm \mathcal{O}(1)$

$$\langle h \rangle = 0$$
$$\langle h \rangle = \Lambda$$

Need: $\sqrt{\epsilon} \sim m_W / \Lambda$

Naturalness*

- Higgs mass is sensitive to high scale threshold (GUT, gravity,...)
- Enormous quantum corrections $\mathcal{O}(\text{highest scale})$ exceed Higgs mass physical value, need to **fine-tune** parameters

Naturalness : absence of special conspiracies between **phenomena occurring at very different length scales**

* Caveat emptor: What about the other naturalness disaster (CC)?
Physics at M_{Planck} might be very different. Multiverse alternative?

Example: add a very heavy scalar to the SM

$$M_\Phi \gg m_H \quad \text{with} \quad \lambda_\Phi |H|^2 |\Phi|^2$$

Example: add a very heavy scalar to the SM

$$M_\Phi \gg m_H \quad \text{with} \quad \lambda_\Phi |H|^2 |\Phi|^2$$

- Quadratic corrections: $\delta m_H^2 \approx \frac{\lambda_\Phi}{16\pi^2} M^2 \ln \frac{M^2}{\Lambda^2} + \dots$
- Need new physics to soften UV dependence, e.g. supersymmetry

Example: add a very heavy scalar to the SM

$$M_\Phi \gg m_H \quad \text{with} \quad \lambda_\Phi |H|^2 |\Phi|^2$$

- Quadratic corrections: $\delta m_H^2 \approx \frac{\lambda_\Phi}{16\pi^2} M^2 \ln \frac{M^2}{\Lambda^2} + \dots$

- Need new physics to soften UV dependence, e.g. supersymmetry

$$\delta m_H^2 = \frac{3y_t^2}{8\pi^2} \tilde{m}_t^2 \ln \frac{\tilde{m}_t^2}{\Lambda^2} + \dots$$

... but new degrees better not be too far above m_H

Electro-weak symmetry breaking & new physics in times of austerity

Light Higgs

```
graph TD; LH[Light Higgs] --> S["light stops_{1,2}, sbottom_L, higgsinos, gluinos, ..."]; LH --> T["light top partners (Q=5/3, 2/3, 1/3), anything else?"]; LH --> Q["?"]; S --- SUP[supersymmetry]; T --- CH[composite Higgs]
```

light stops_{1,2}, sbottom_L,
higgsinos, gluinos, ...

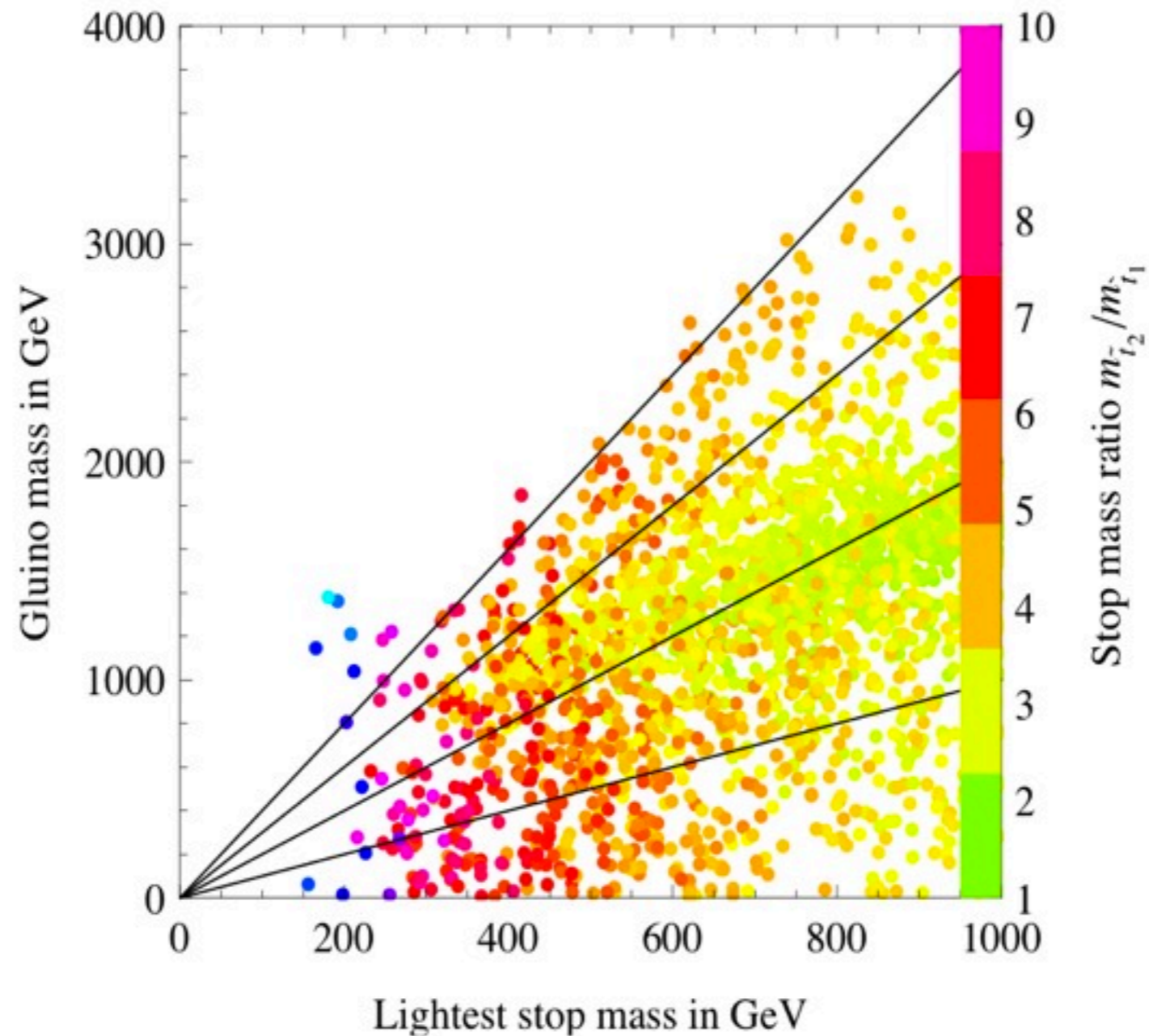
supersymmetry

light top partners
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anything else ?

composite Higgs

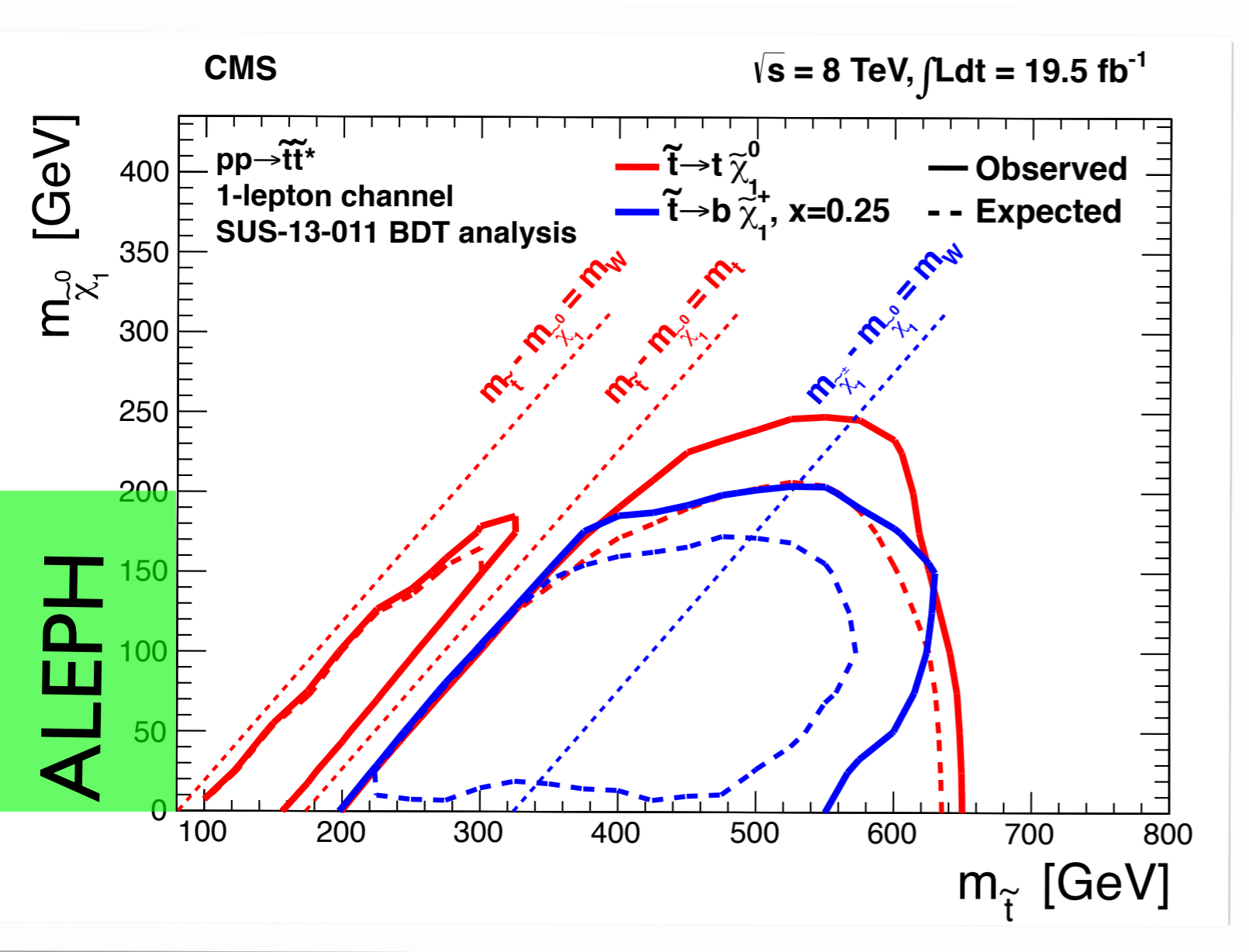
MSSM stops vs. m_H

Delgado et al.



Even easier in NMSSM, ...

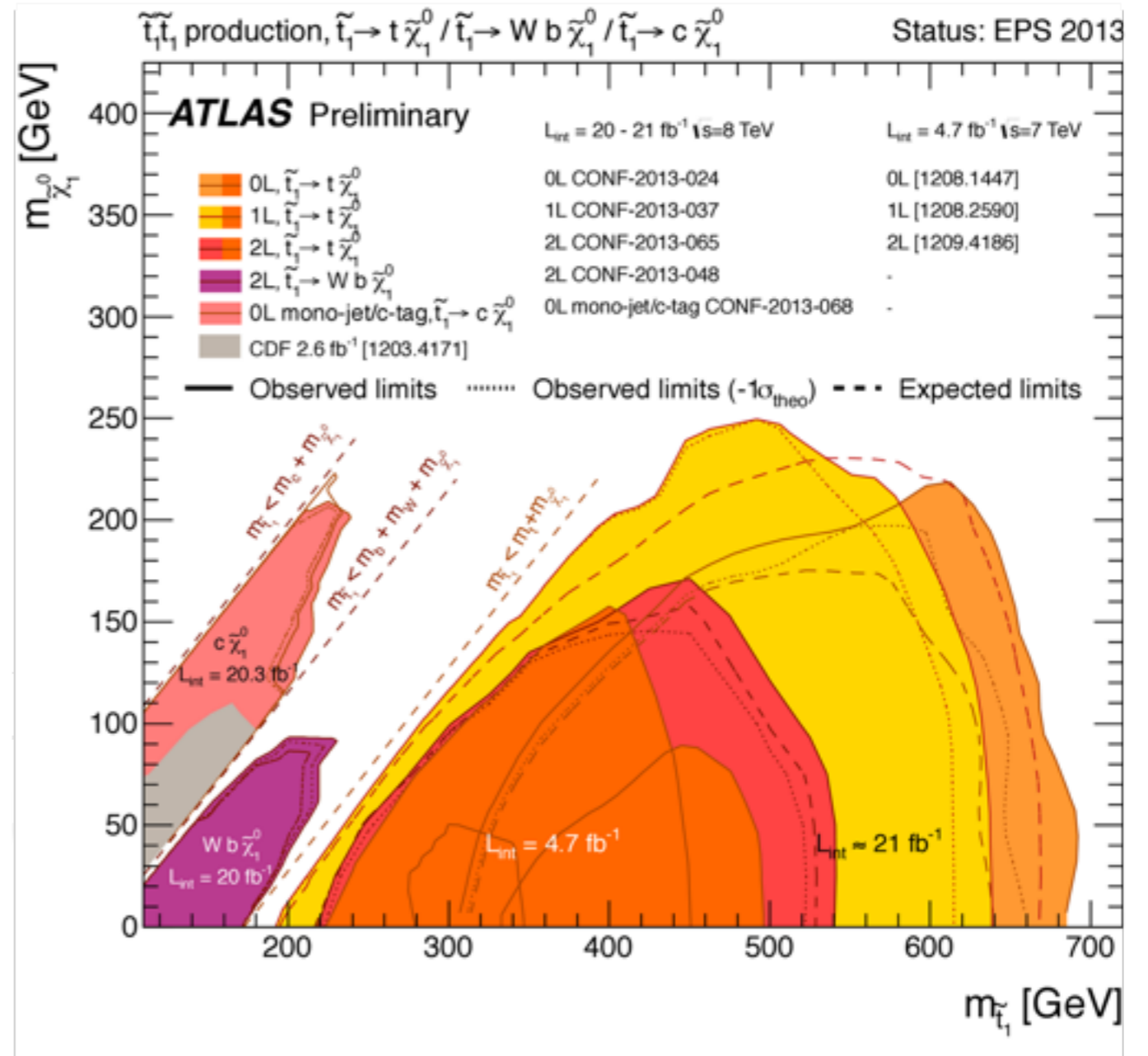
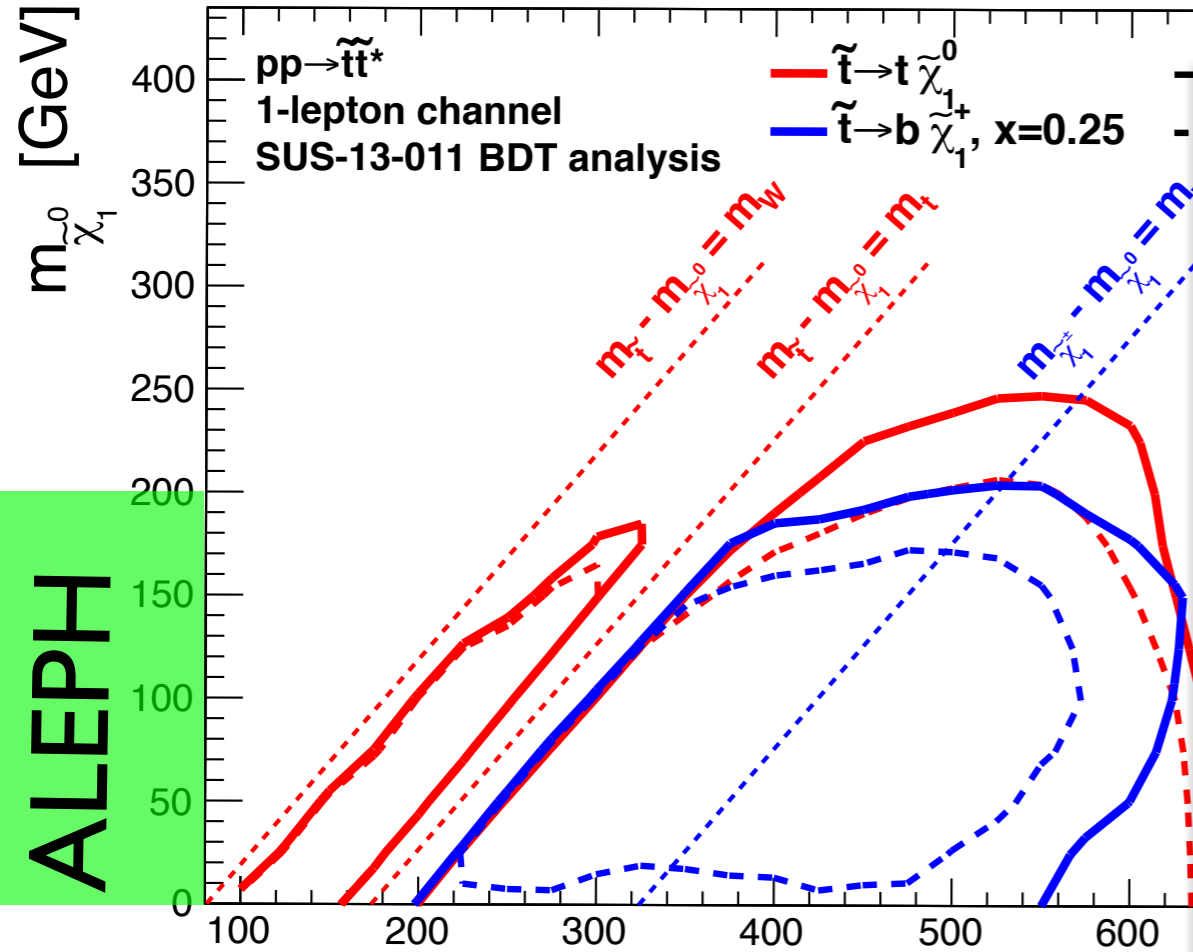
Direct stop searches



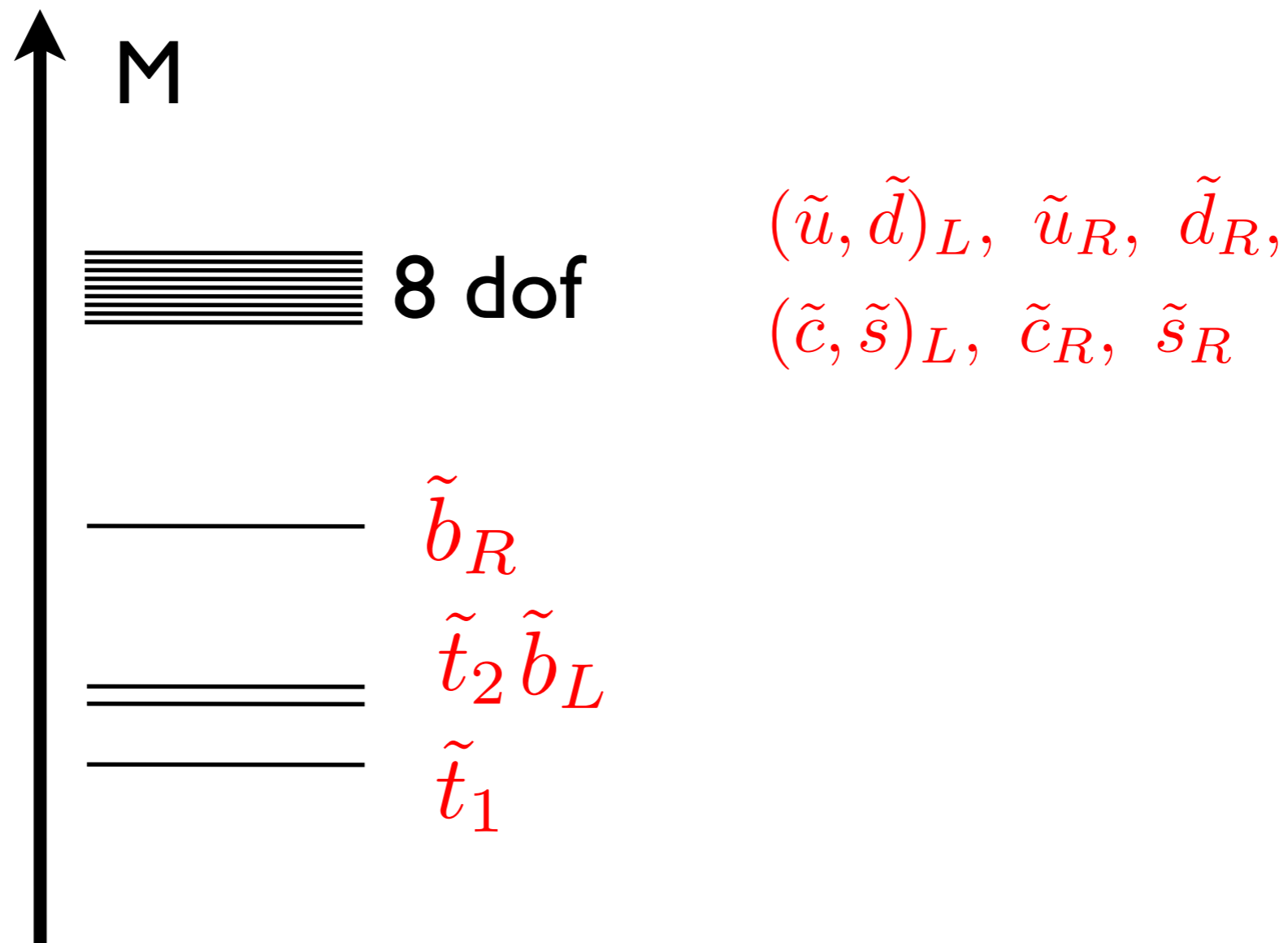
Direct stop searches

CMS

$\sqrt{s} = 8 \text{ TeV}, \int \mathcal{L} dt = 19.5 \text{ fb}^{-1}$



Naturalness prefers split squarks



Splitting via RGE?

Papucci, Ruderman, AW '11

Splitting via renormalization group does not help

$$\delta m_H^2 \simeq 3 \left(m_{Q_3}^2 - m_{Q_{1,2}}^2 \right) \simeq \frac{3}{2} \left(m_{U_3}^2 - m_{U_{1,2}}^2 \right)$$

1-loop, LLog,
tan β moderate

Higgs fine-tuning = **RGE mass splitting**

Splitting via RGE?

Papucci, Ruderman, AW '11

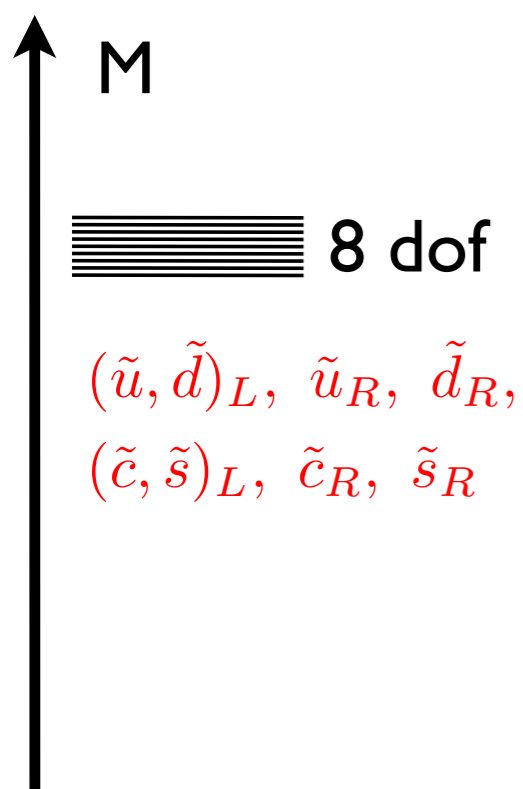
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1-loop, LLog,
tan β moderate

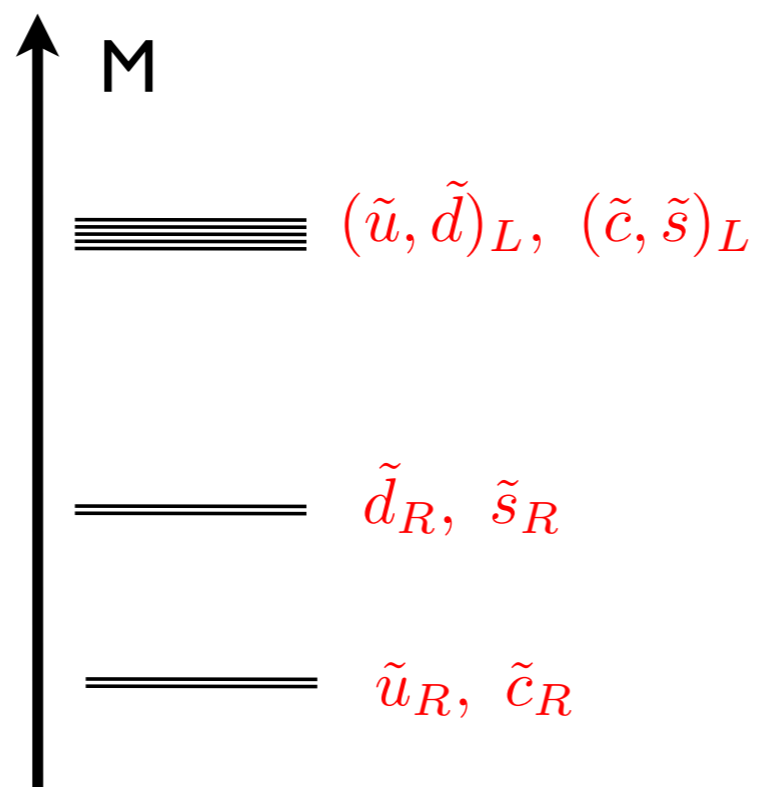
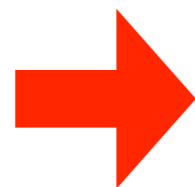
Higgs fine-tuning = RGE mass splitting

→ Flavor non-trivial susy
breaking!

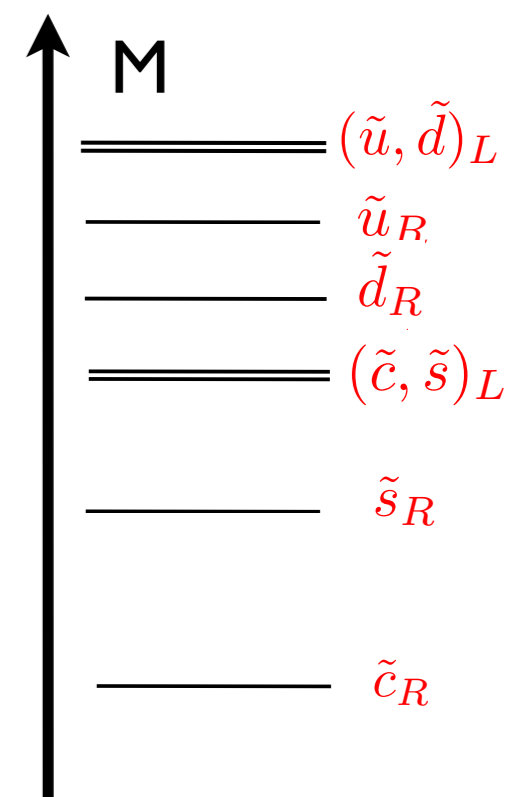
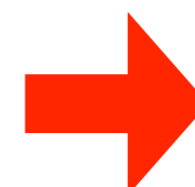


Degenerate

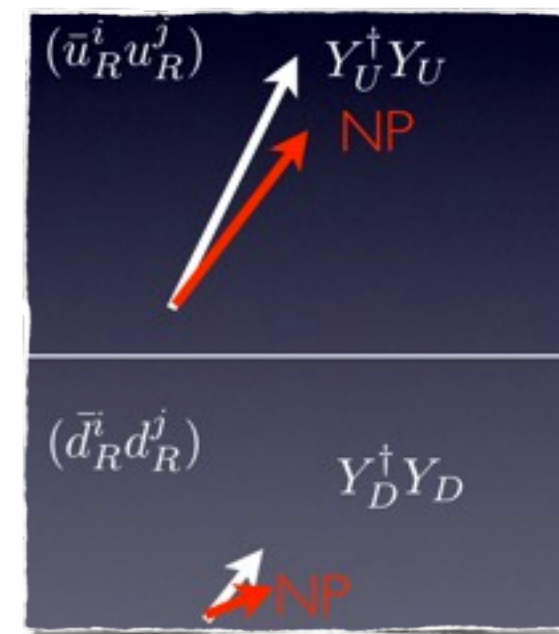
mSugra, CMSSM,
 pMSSM, ...



Minimal Flavor

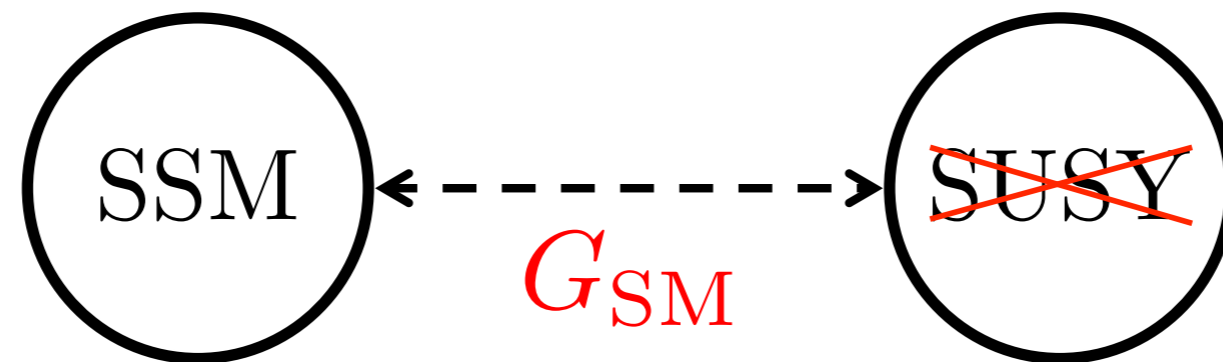


Anarchy!



Gauge Mediation

see e.g. Giudice/Rattazzi review

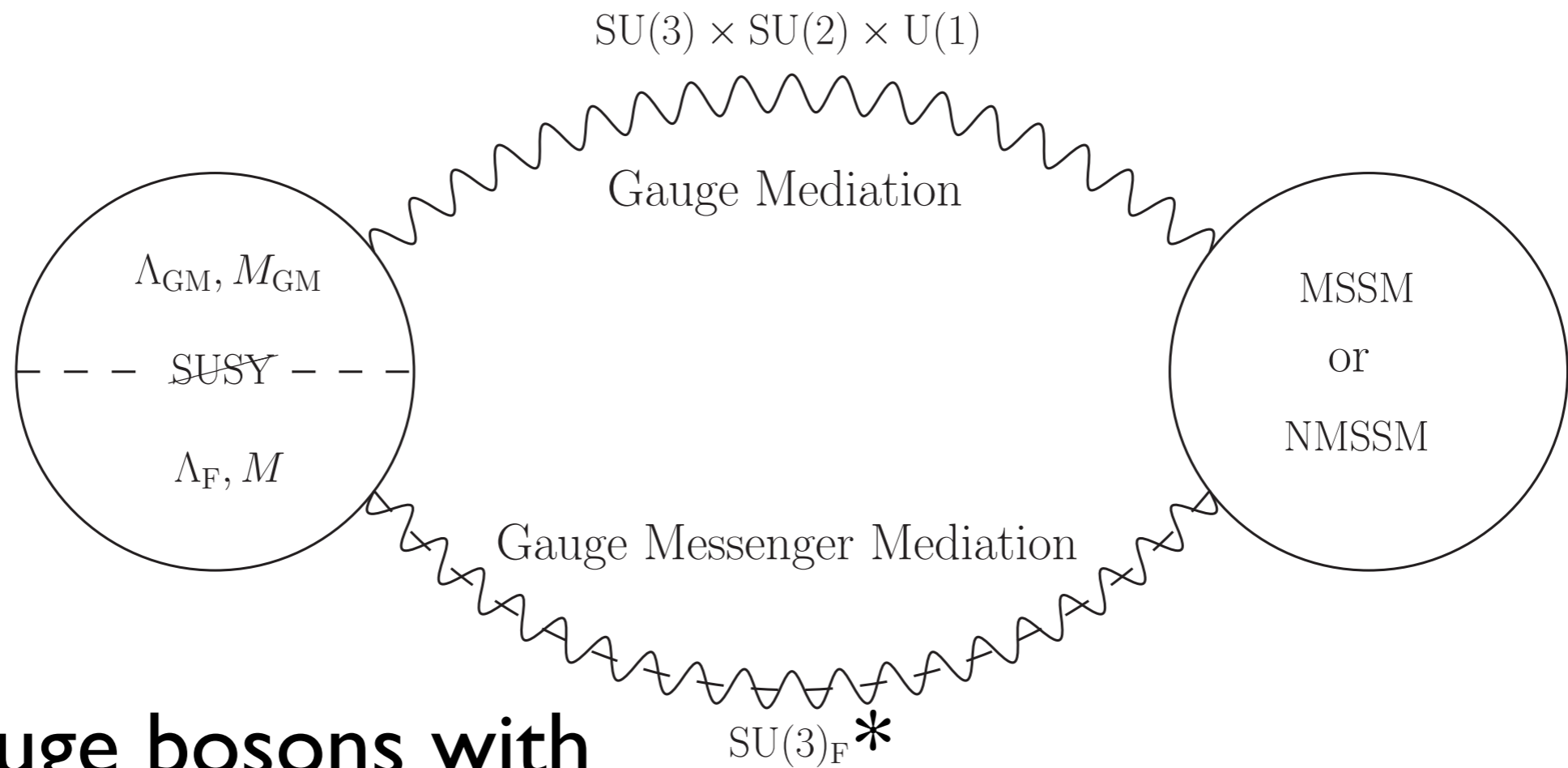


$$G_{SM} = SU(3) \times SU(2) \times U(1)$$

Degenerate quarks!

Split spectrum from flavor gauge mediation

U(1): Kaplan, Kribs '99; Craig, McCullough, Thaler '12
Brümmer, McGarrie, AW (to appear)



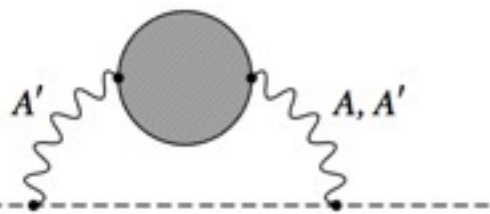
Flavor gauge bosons with
flavor and susy breaking masses

* Diagonal, anomaly-free subgroup of SM w/o Yukawas
 $SU(3)_{Q_L} \times SU(3)_{u_R} \times SU(3)_{d_R}$

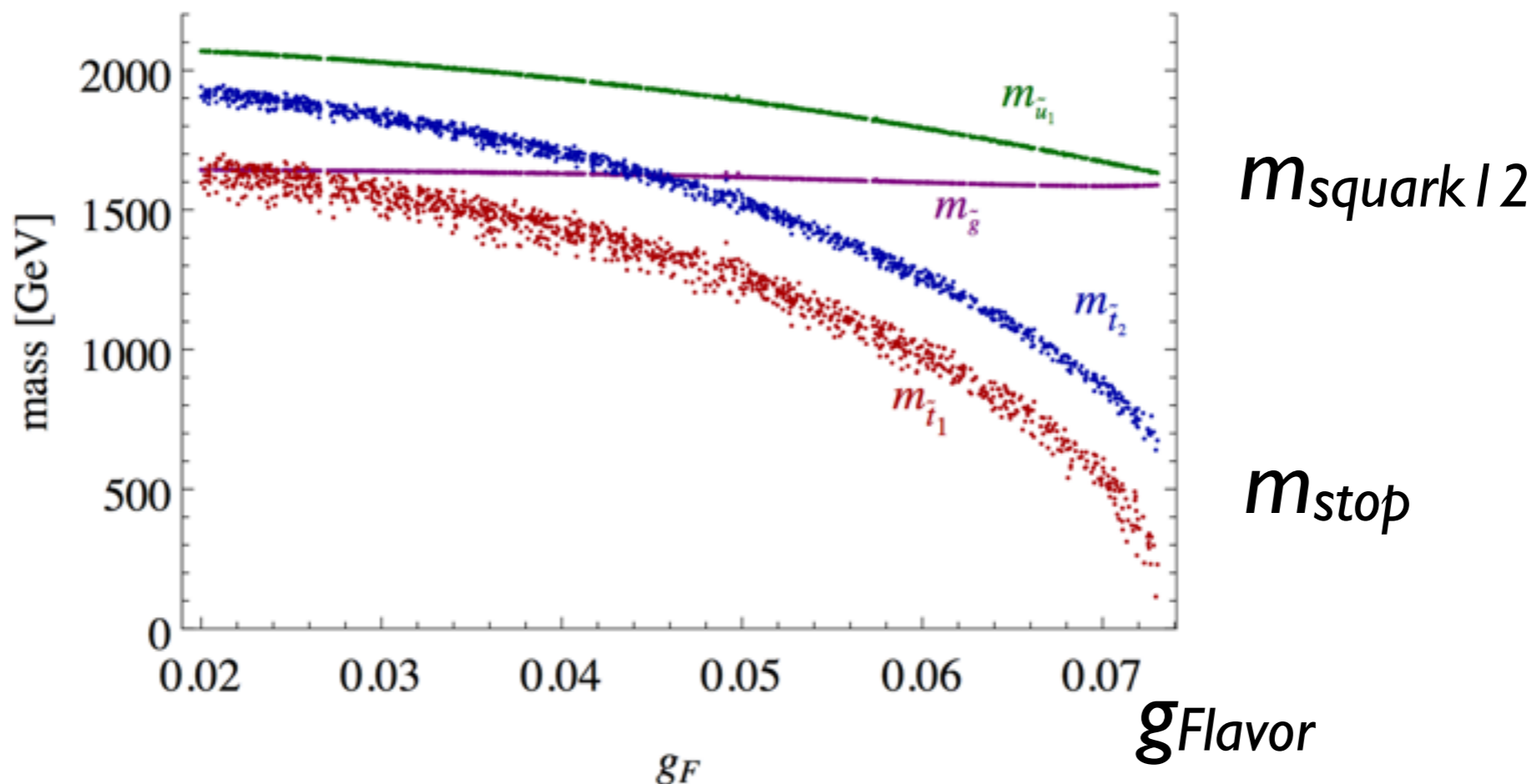
Natural Split spectrum

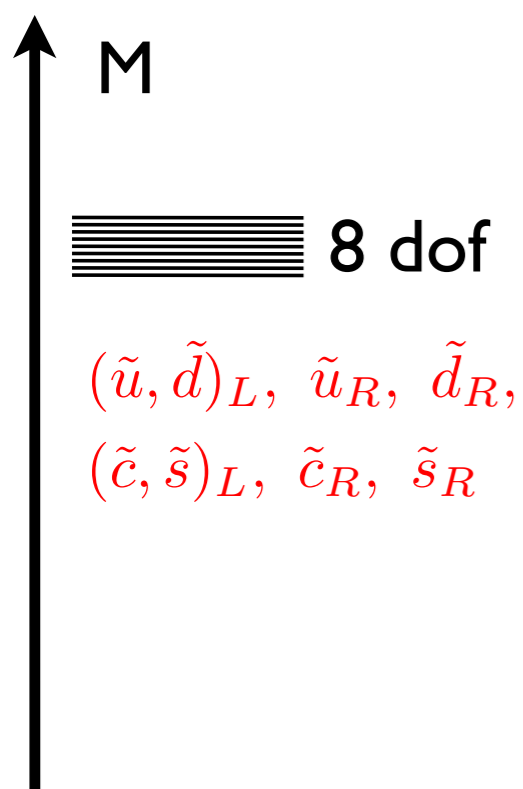
Brümmer, McGarrie, Weiler (to appear)

Negative contribution from gauge messengers



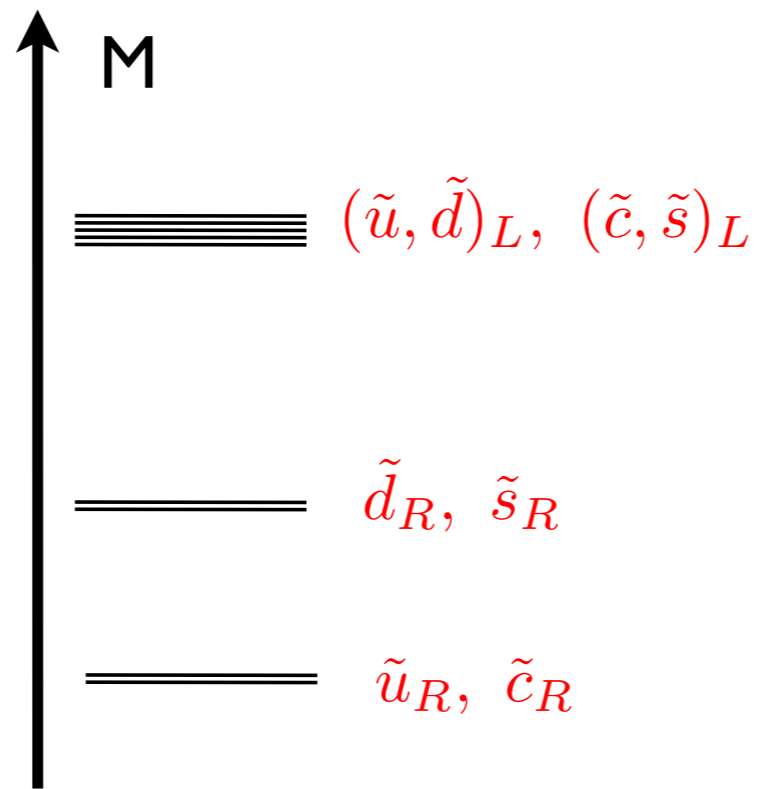
$$\delta m_{Q,U,D}^2 = -\frac{g_F^2}{16\pi^2} \frac{F^2}{M^2} \left[\begin{pmatrix} \frac{7}{6} & 0 & 0 \\ 0 & \frac{7}{6} & 0 \\ 0 & 0 & \frac{8}{3} \end{pmatrix} + \mathcal{O}(\epsilon^2) \right]$$



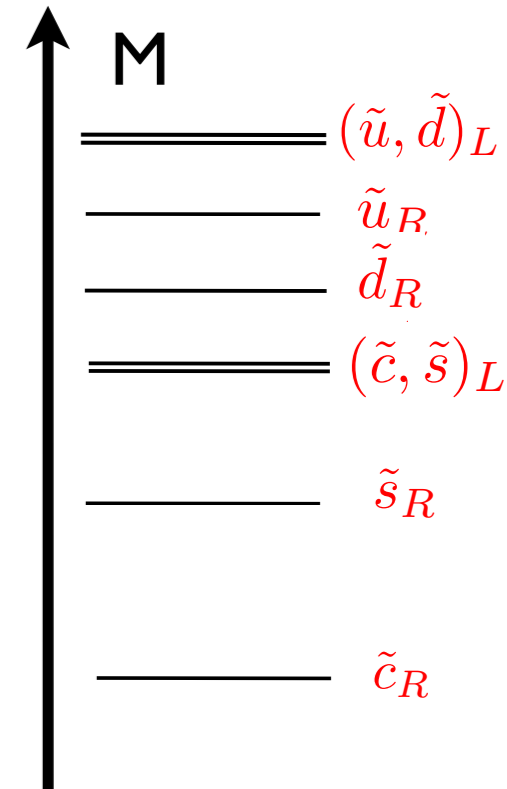
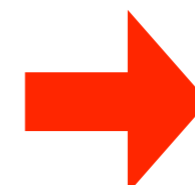


Degenerate

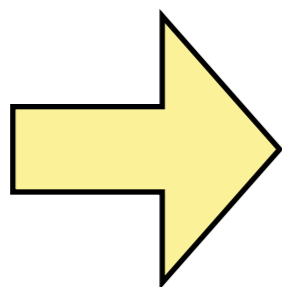
mSugra, CMSSM,
 pMSSM, ...



Minimal Flavor



Anarchy!



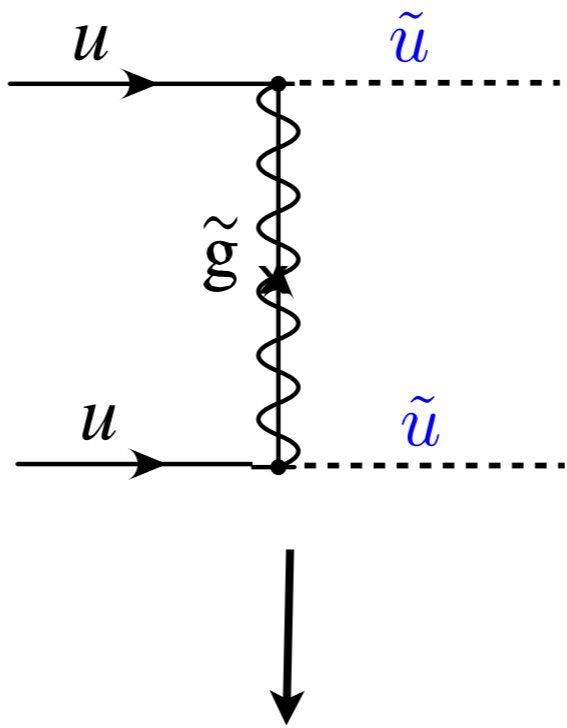
**Consider beyond MFV susy searches.
 Sensitivity can change dramatically..**

$$N_{\text{signal}} = [\text{multiplicity}] \times [\text{pdfs}] \times [\text{signal efficiency}]$$

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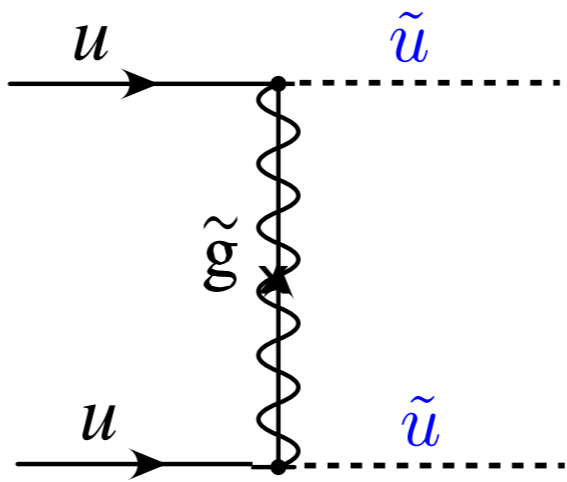


8 degenerate squarks
→ 1 light squark flavor

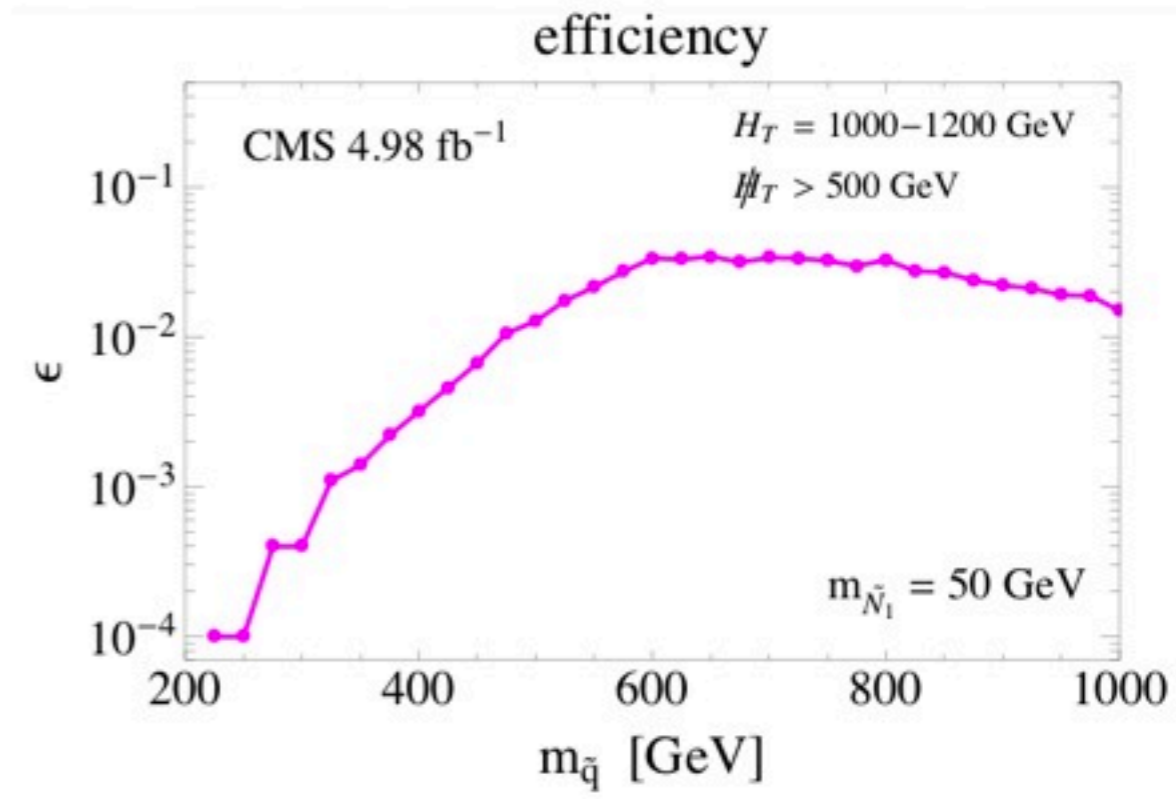


$$N_{\text{signal}} = [\text{multiplicity}] \times [\text{pdfs}] \times [\text{signal efficiency}]$$

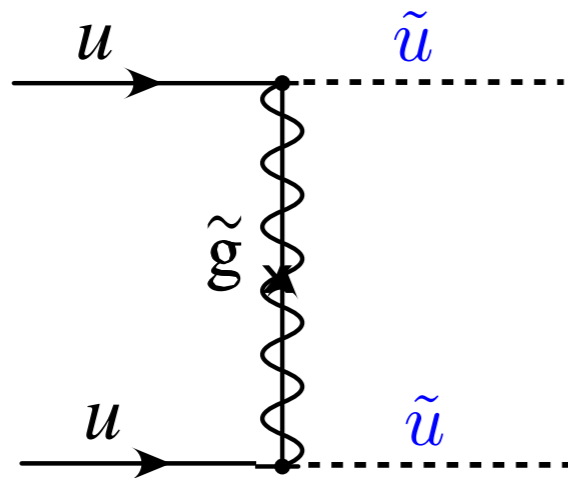
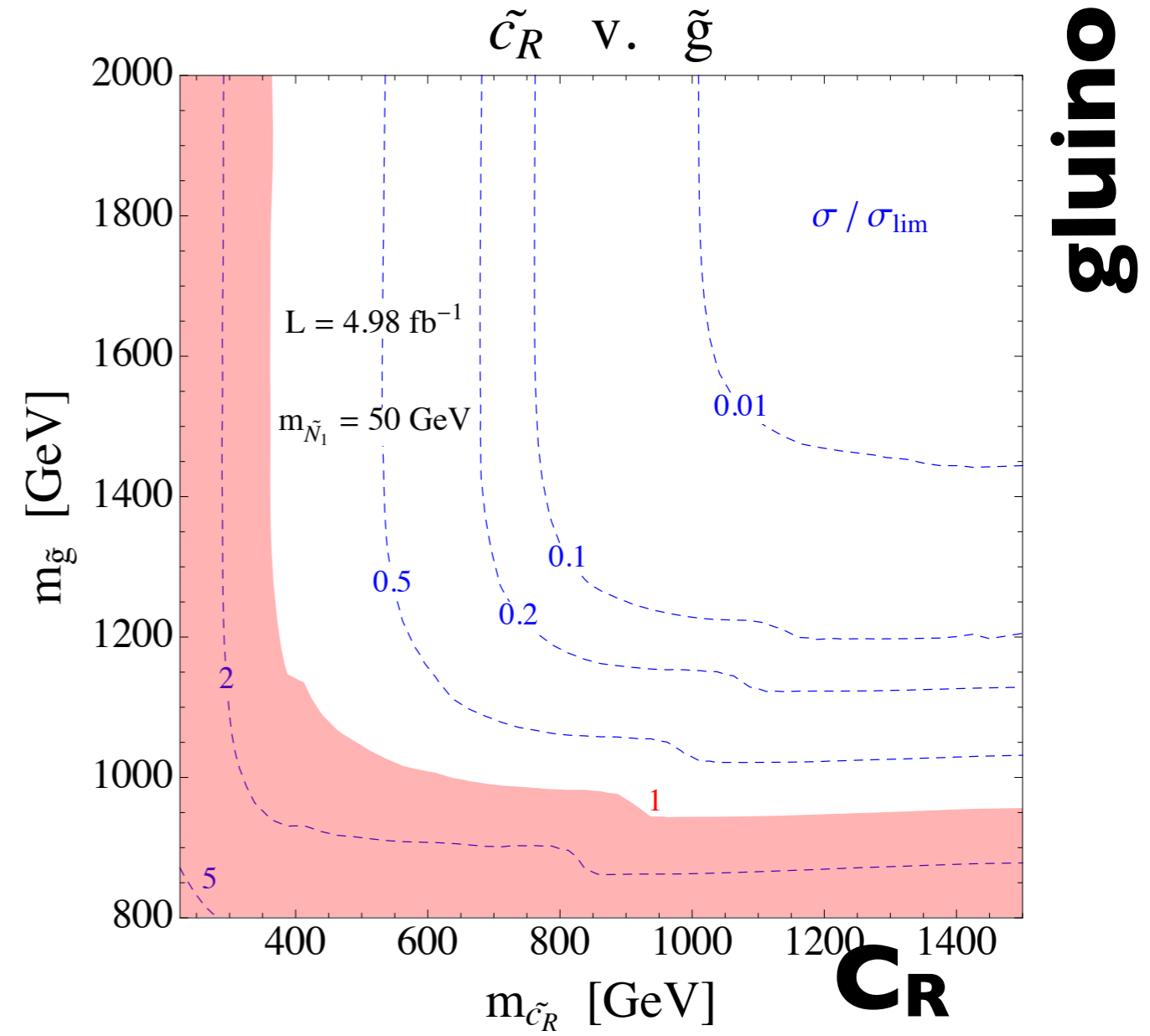
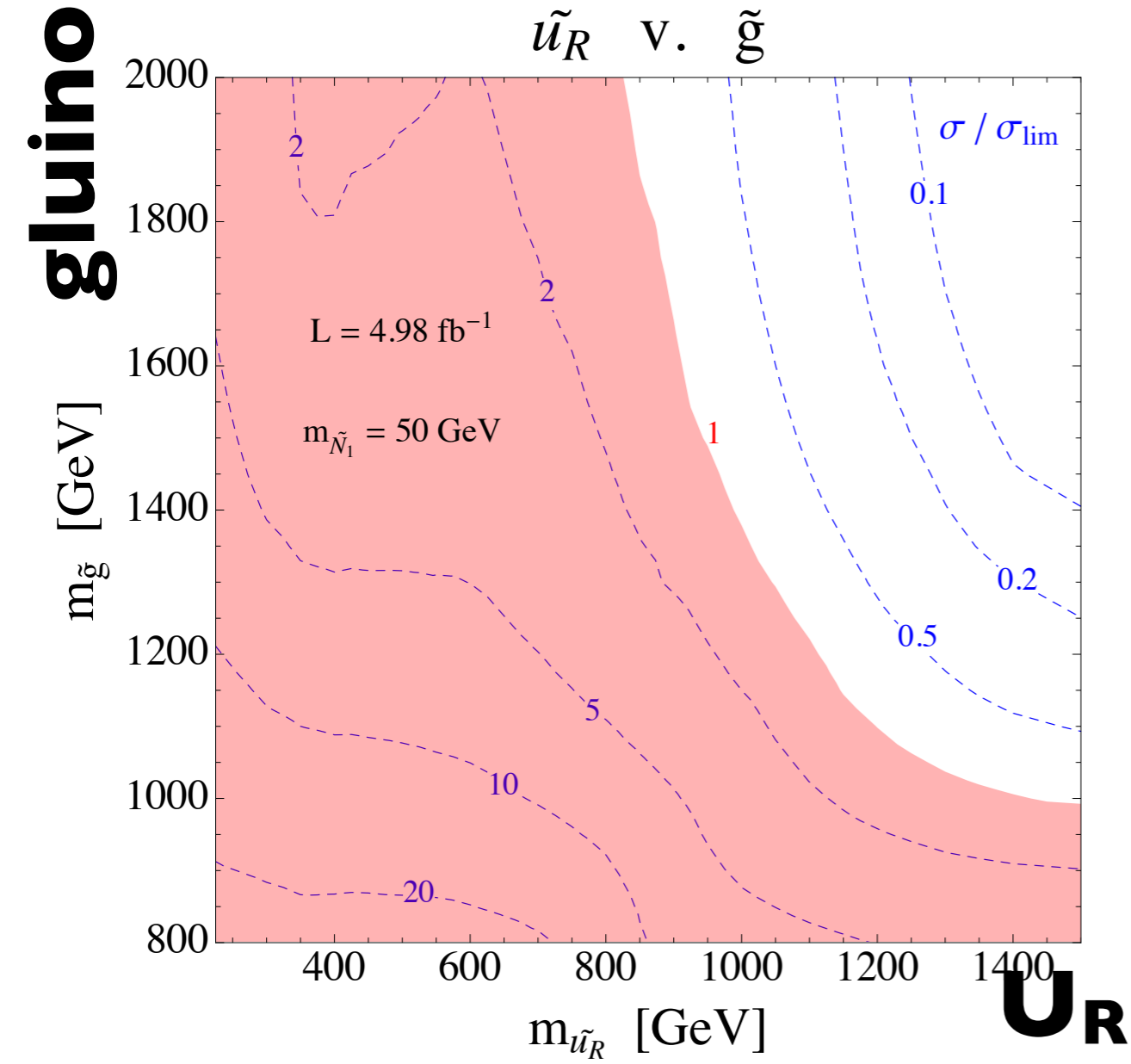
8 degenerate squarks
 → 1 light squark flavor



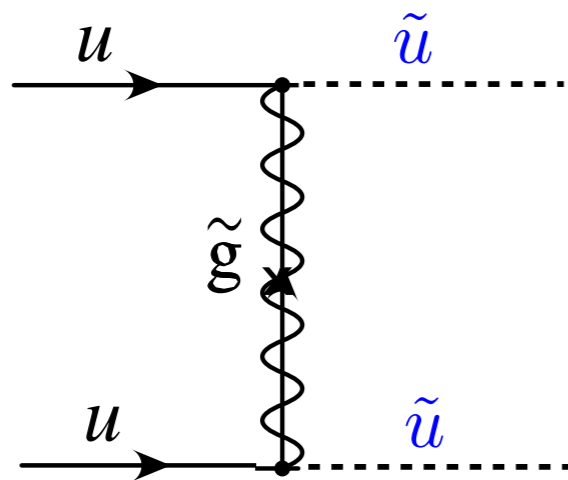
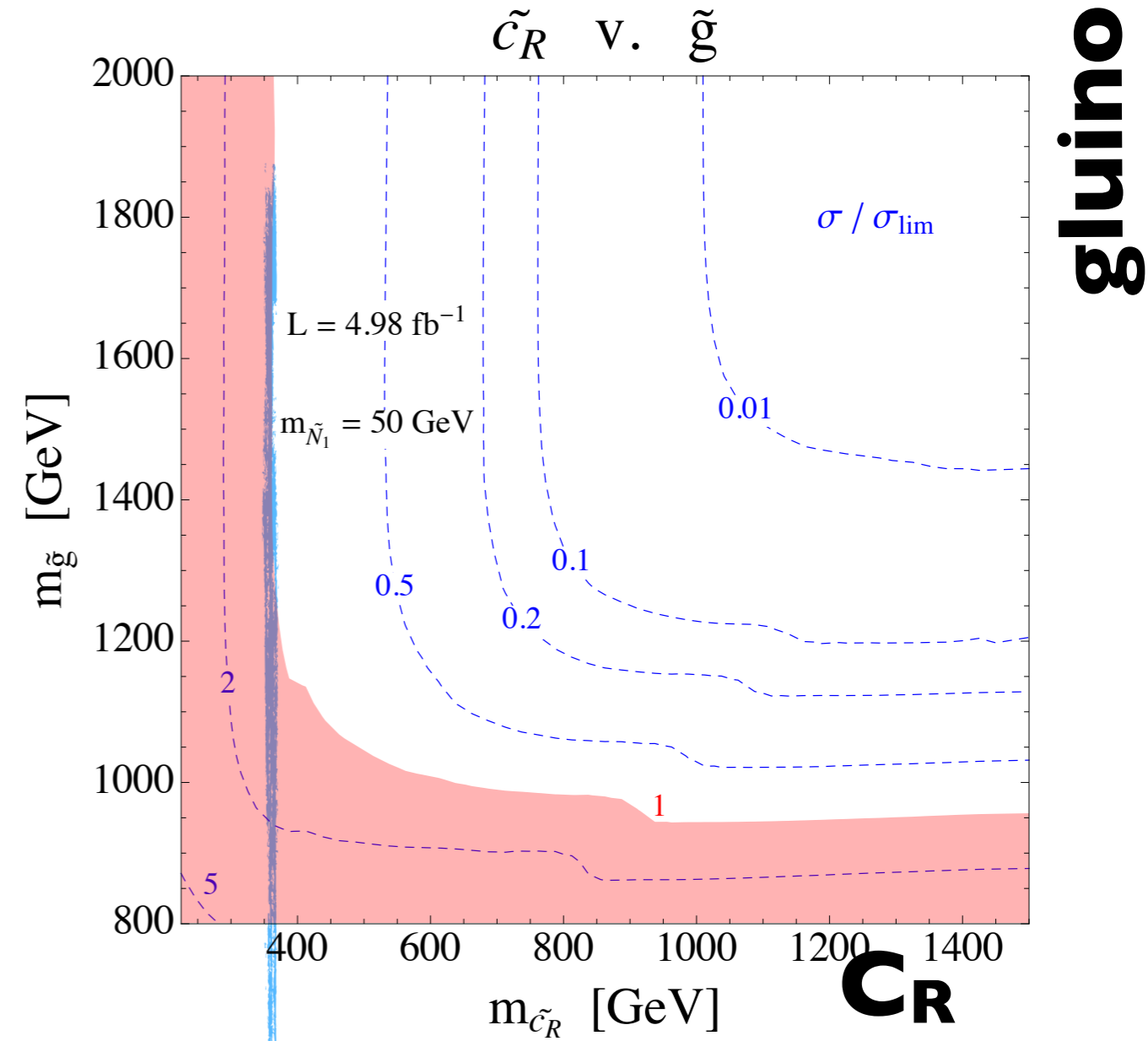
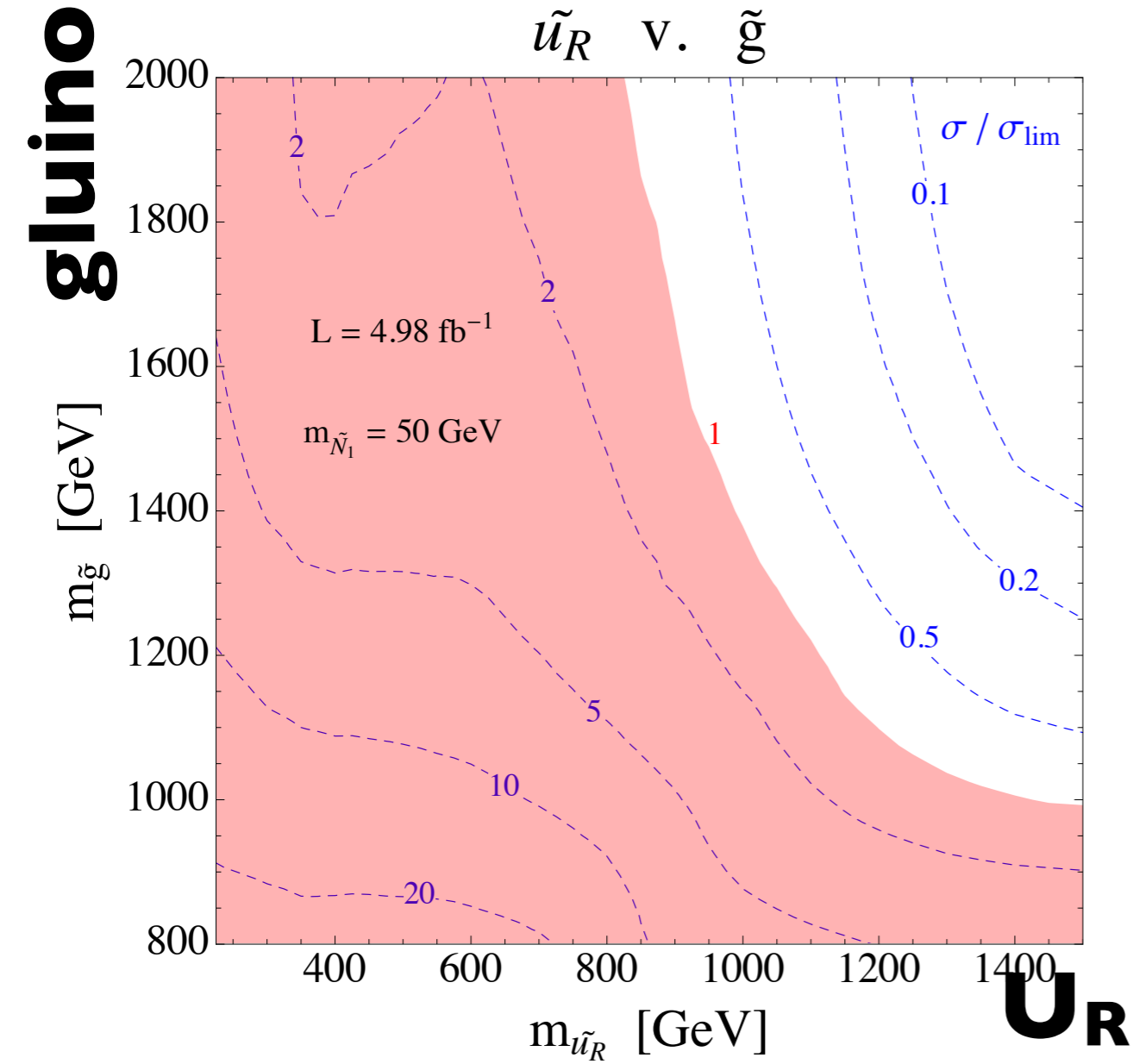
$$N_{\text{signal}} = [\text{multiplicity}] \times [\text{pdfs}] \times [\text{signal efficiency}]$$



8 degenerate squarks
 → 1 light squark flavor



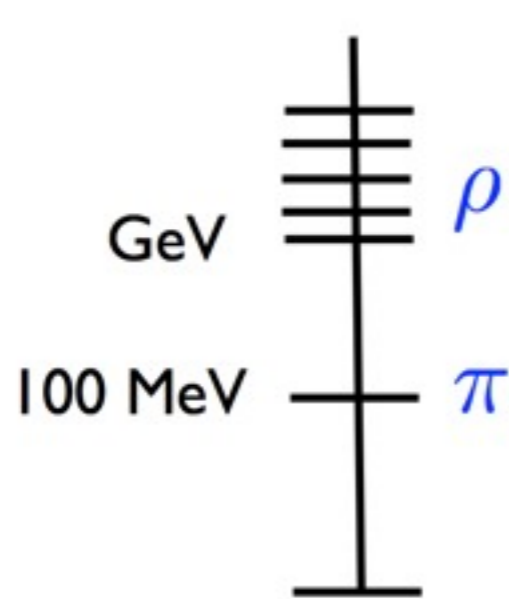
Recent update
(e.g. CMS PAS SUS-13-012)



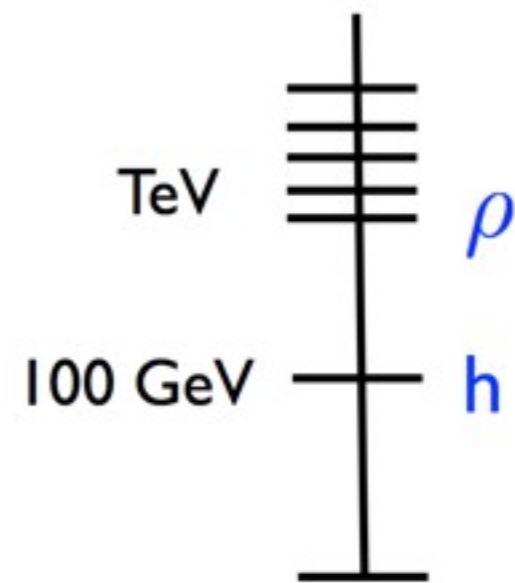
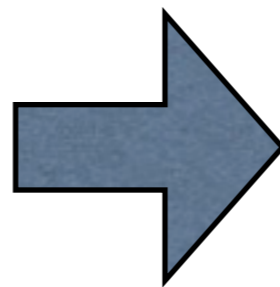
**Sea squarks can
still be < 400 GeV**

Recent update
(e.g. CMS PAS SUS-13-012)

Strong EWSB (Composite Higgs)



QCD



Higgs as a pGB

Why is the Higgs light?

Kaplan; Agashe et. al

Inspired by QCD: (pseudo) scalar pion is the lightest state


Shift symmetry...

$$\pi \rightarrow \pi + c$$

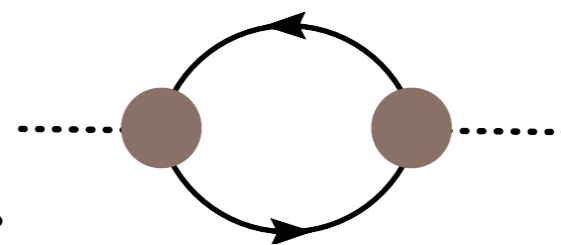
... protects its mass.

Interactions are perturbative for $E \ll 4\pi f$

No pure composite effects due to Goldstone symmetry


$$= 0$$

Shift symmetry broken by elementary-composite couplings:



$$m_h^2 \sim \frac{\lambda^2}{16\pi^2} \Lambda_{comp}^2$$

$$\lambda \ll 4\pi$$

Light Higgs implies light fermionic top partners

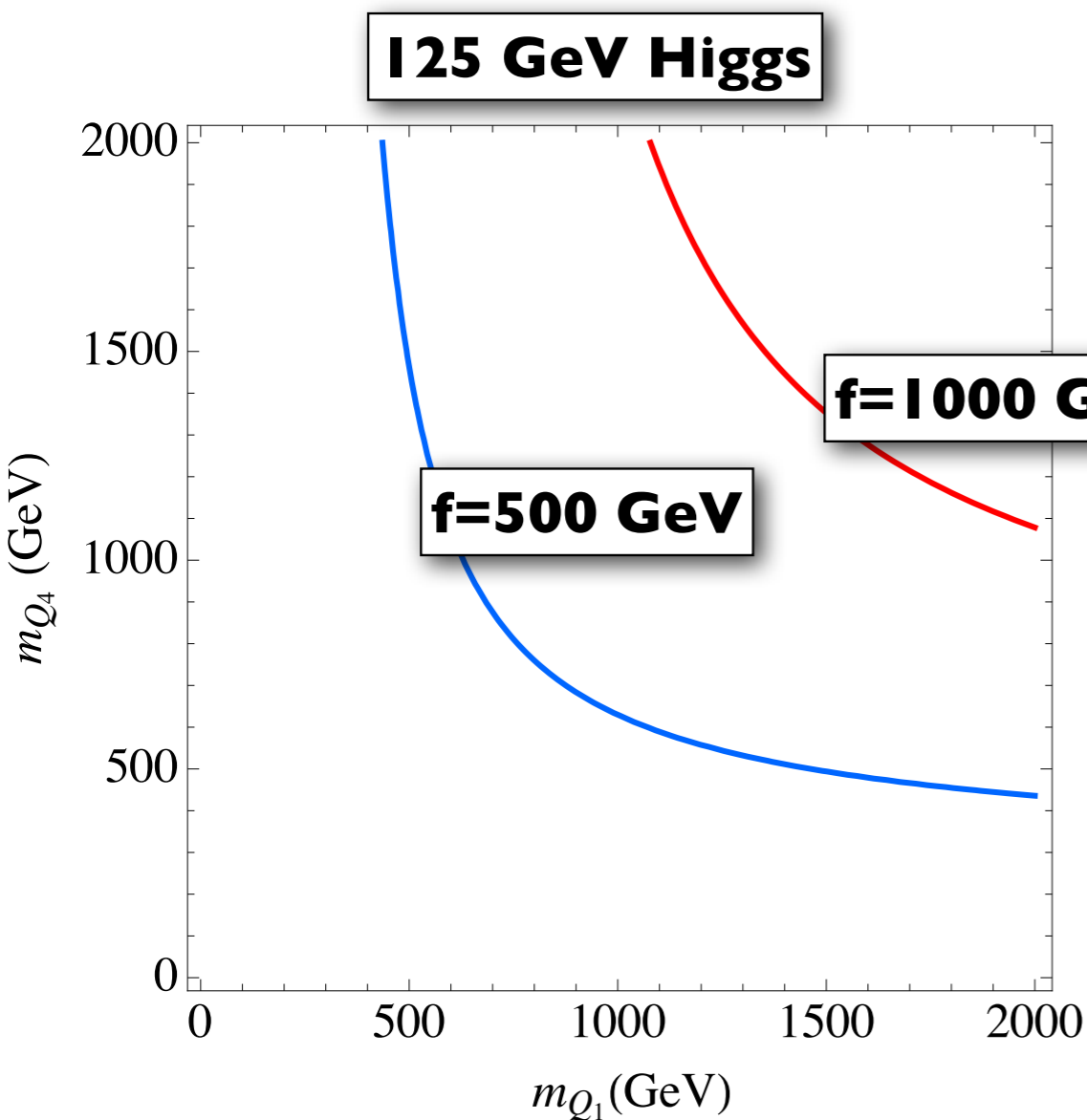
$$m_h^2 \simeq \frac{N_c}{\pi^2} \left[\frac{m_t^2}{f^2} \frac{m_{Q_4}^2 m_{Q_1}^2}{m_{Q_1}^2 - m_{Q_4}^2} \log \left(\frac{m_{Q_1}^2}{m_{Q_4}^2} \right) \right]$$

Pomarol, Riva 12

Light Higgs implies light fermionic top partners

$$m_h^2 \simeq \frac{N_c}{\pi^2} \left[\frac{m_t^2}{f^2} \frac{m_{Q_4}^2 m_{Q_1}^2}{m_{Q_1}^2 - m_{Q_4}^2} \log \left(\frac{m_{Q_1}^2}{m_{Q_4}^2} \right) \right]$$

Pomarol, Riva 12



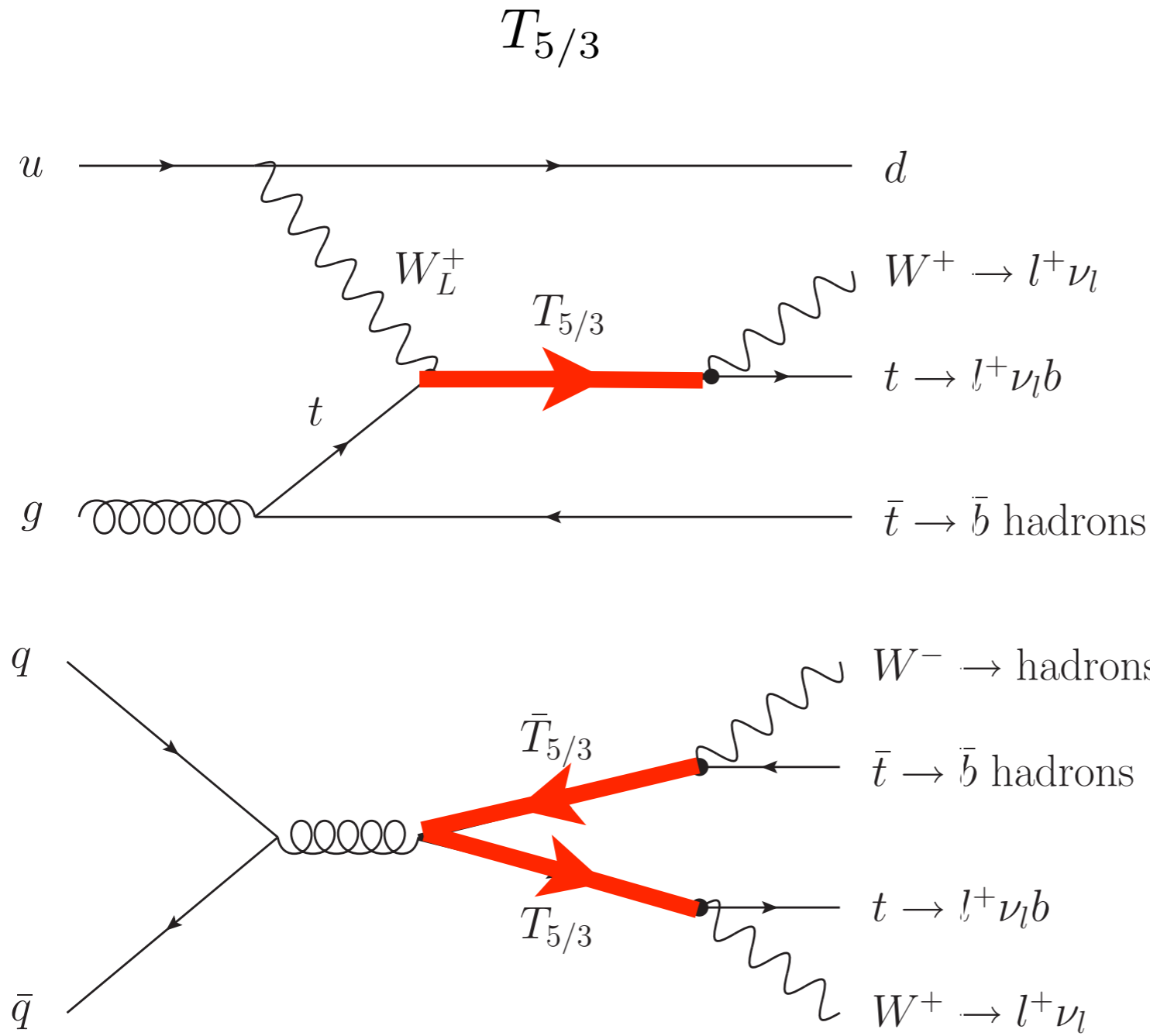
$$5 = 4 + 1$$

$Q_4 \quad Q_1$

with EM charges $5/3, 2/3, -1/3$

Contino et al; Pomarol, Riva;
Matsedonskyi, Panico, Wulzer; Redi, Tesi;
Marzocca, Serone, Shu;

e.g. Perelstein, Pierce, Peskin
 Contino, Servant; Mrazek, Wulzer;
 De Simone, Matsedonkyi, Rattazzi, Wulzer



Single

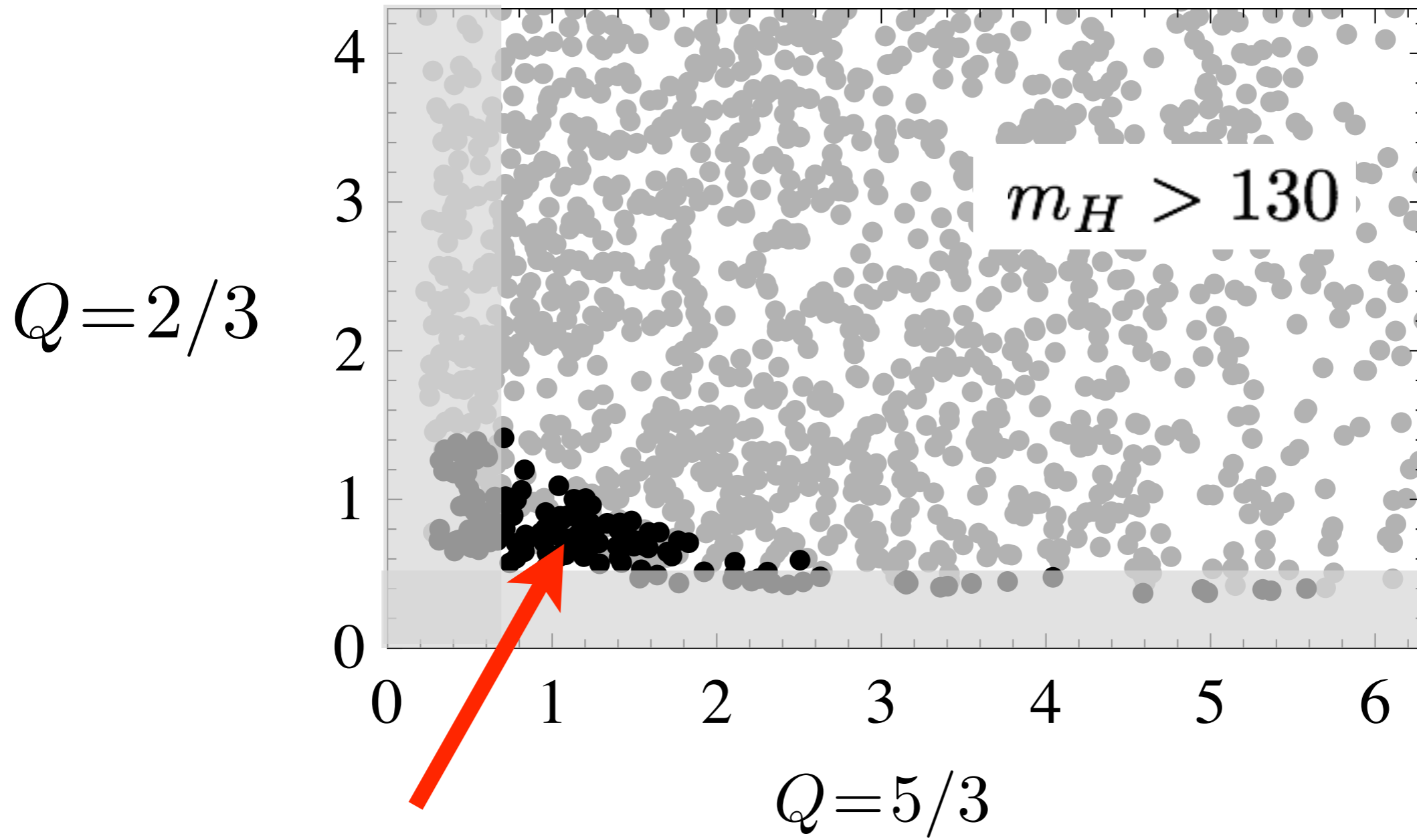
Spectrum:
 ——— B
 ——— T

$X_{2/3}$
 $X_{5/3}$

Double

from 1204.6333

$$\xi = 0.2$$

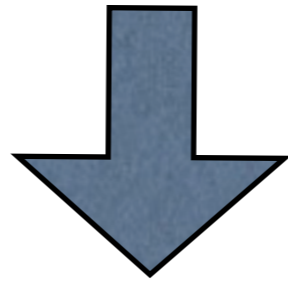


$m_H = 115 \dots 130$ GeV

Flavor used to be a show-stopper

CPV in Kaon mixing

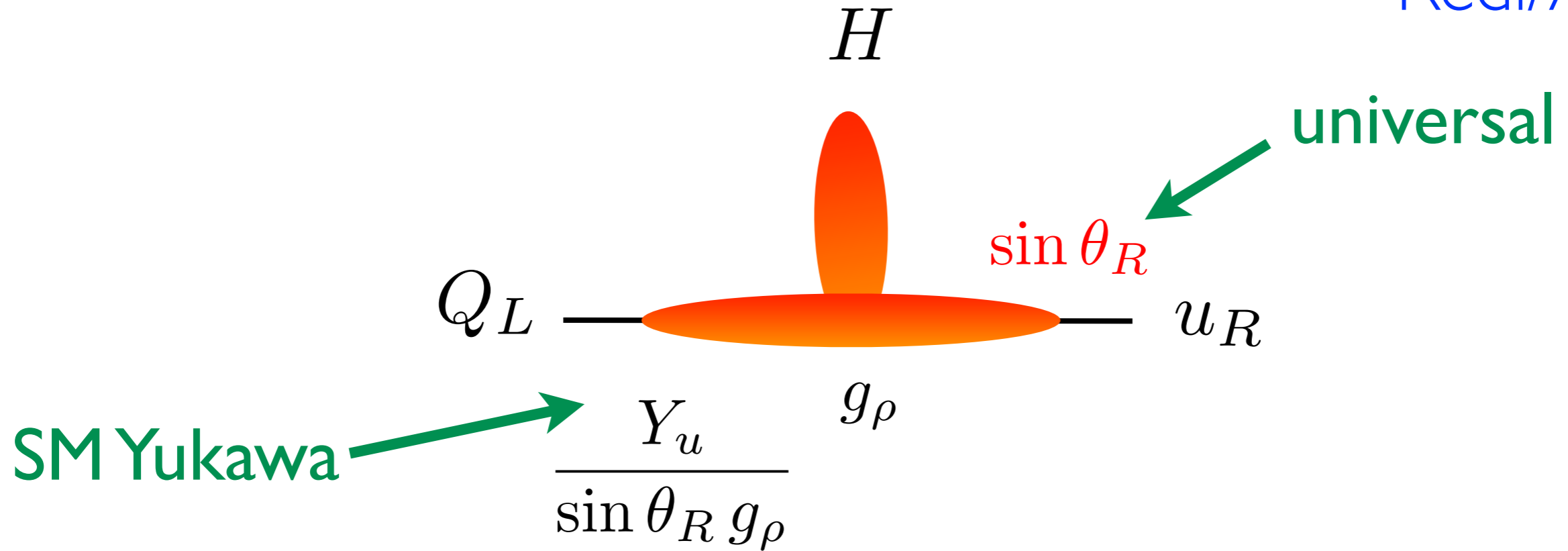
$$|\epsilon| = 2.3 \times 10^{-3} \implies \frac{M_{ETC}}{g_{ETC} \sqrt{\text{Im}(V_{sd}^2)}} \gtrsim 16,000 \text{ TeV}$$



$$m_{q,\ell,T}(M_{ETC}) \simeq \frac{g_{ETC}^2}{2M_{ETC}^2} \langle \bar{T}T \rangle_{ETC} \lesssim \frac{0.1 \text{ MeV}}{|V_{sd}|^2 N^{3/2}} \quad \text{vs. } m_{\text{top}}$$

A Minimal Flavor Violating Composite Higgs

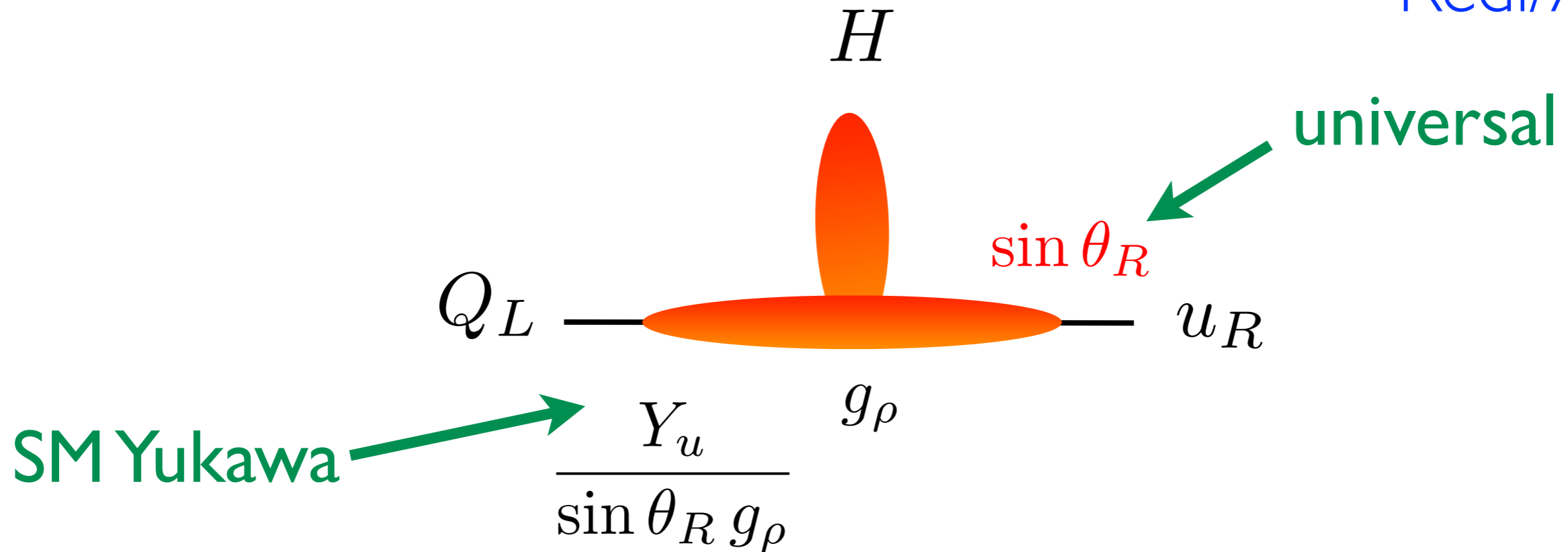
Redi/AW*



*for RS realization: Csaki,AW et al; Delaunay et al; da Rold; see also Barbieri et al

A Minimal Flavor Violating Composite Higgs

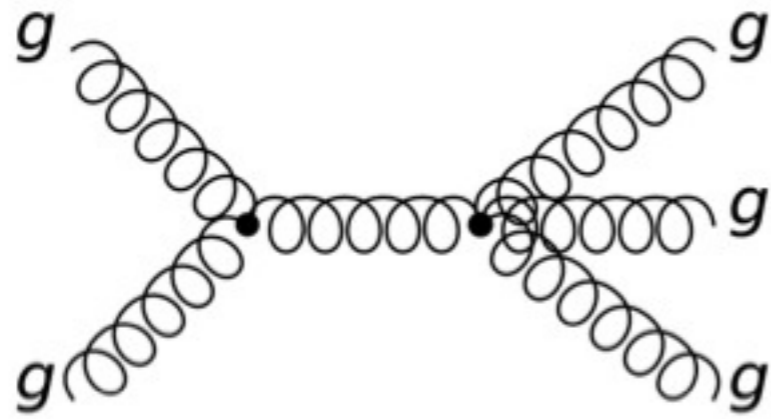
Redi/AW*



Composite u, d quarks, spectacular signals!

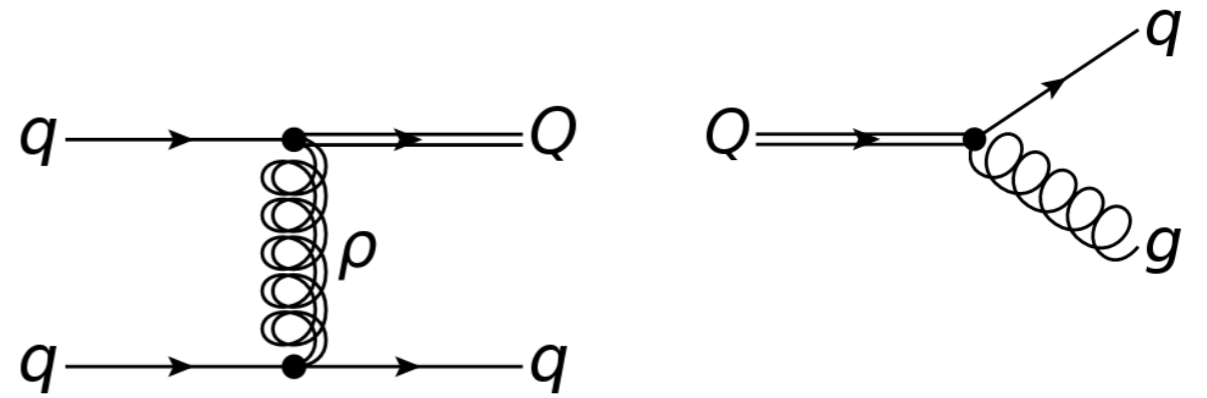
$$m_{top} : \quad \sin \theta_R \gtrsim \frac{1}{g_\rho} \sim \frac{1}{8}$$

*for RS realization: Csaki, AW et al; Delaunay et al; da Rold; see also Barbieri et al



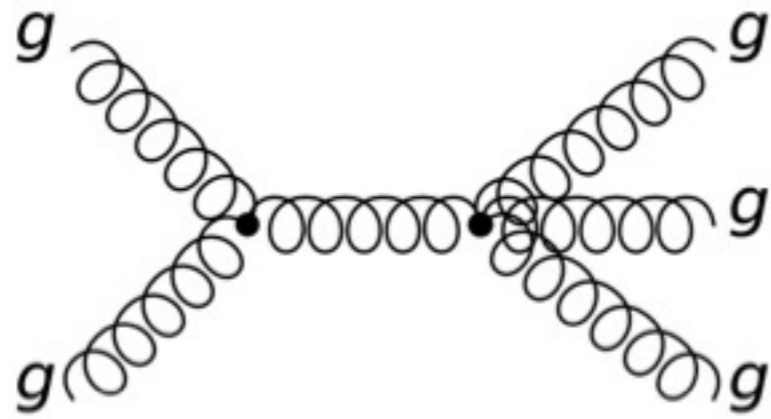
QCD

vs.



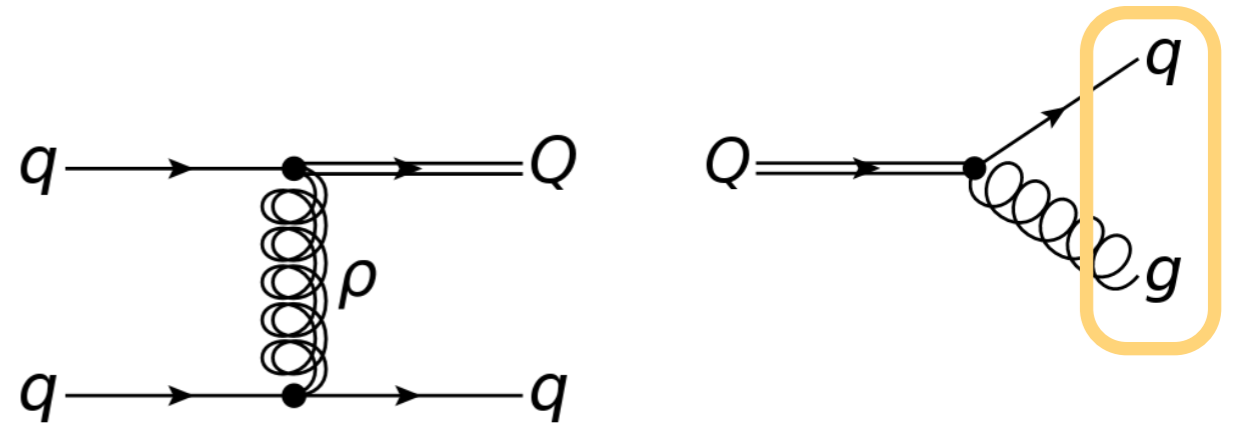
Composite Partners

bump in sub-leading jets



QCD

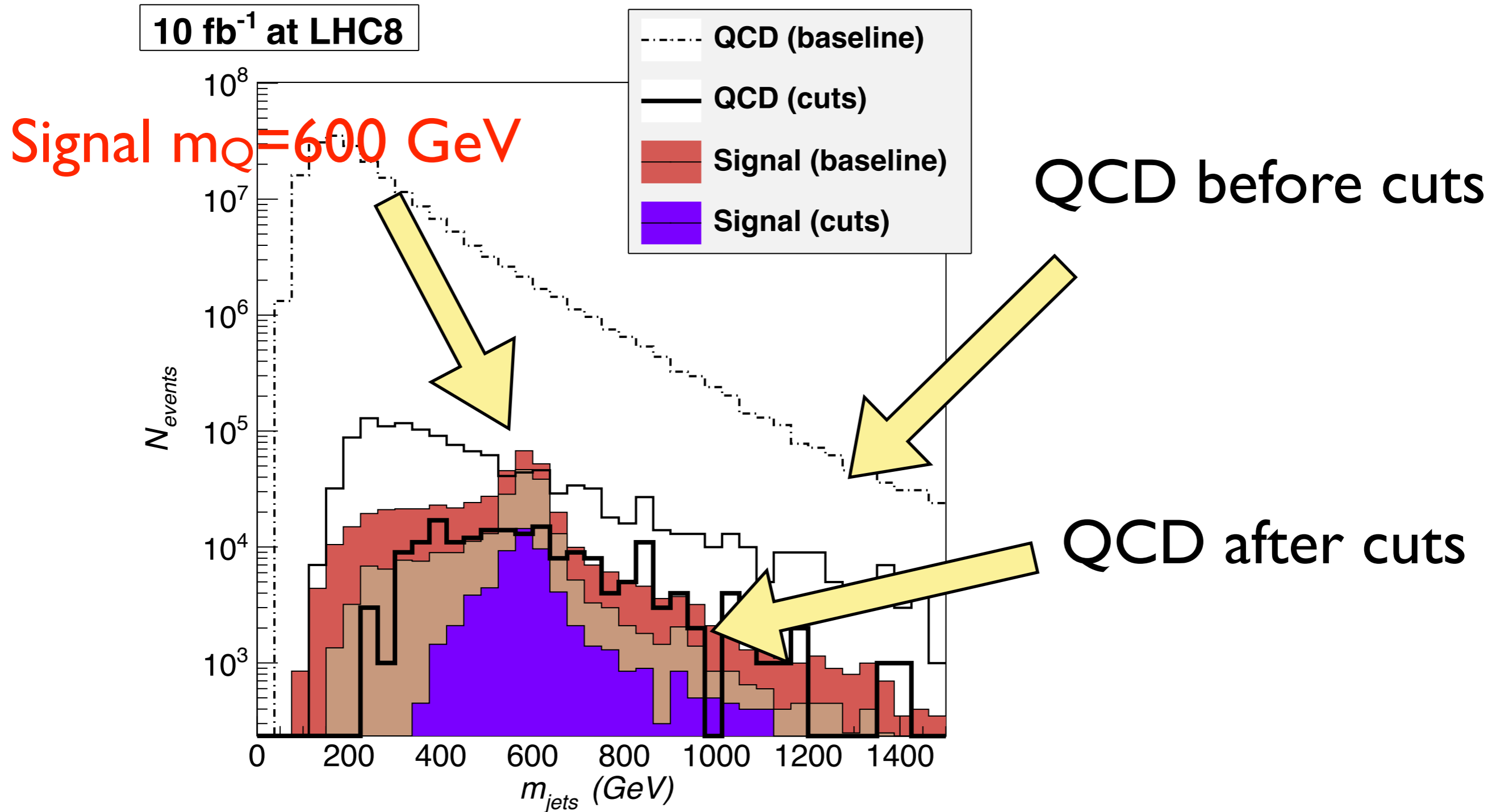
vs.



Composite Partners

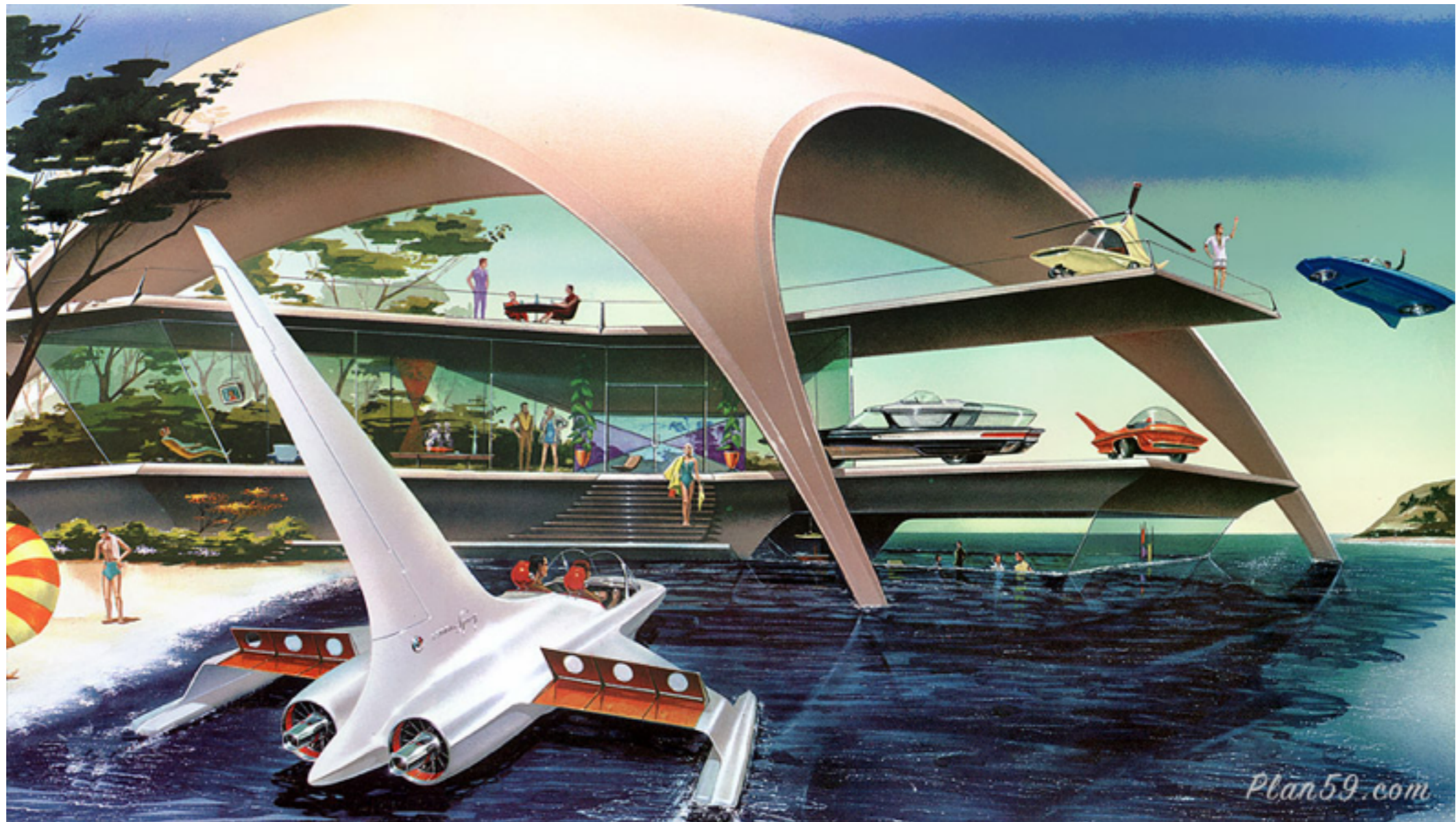
Discovery potential of a dedicated search

deVries, Redi, Sanz, AW, '13



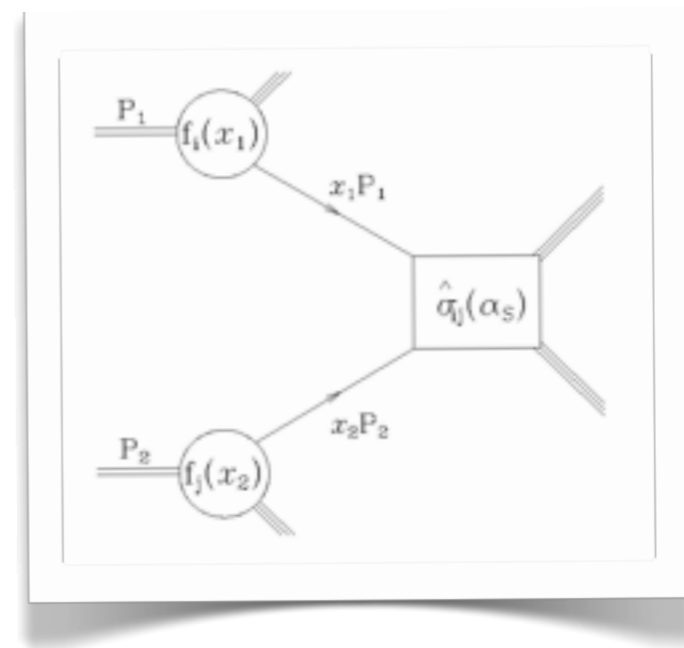
heavier signal easier, 3jet final state
easier, no optimization

The Future



What will we be sensitive to?

Simple exercise, **parton luminosities**:



$$\sigma(M^2) = \sum_{i,j} \int_{M^2/s}^1 d\tau \mathcal{L}_{ij}(\tau) \hat{\sigma}(s\tau)$$

Partonic cross-section scales as $\hat{\sigma} \propto \frac{1}{M^2}$

with $M = \sqrt{\hat{s}}, M_{Z'}, 2m_{\text{squark}}, \dots$

Mass reach?

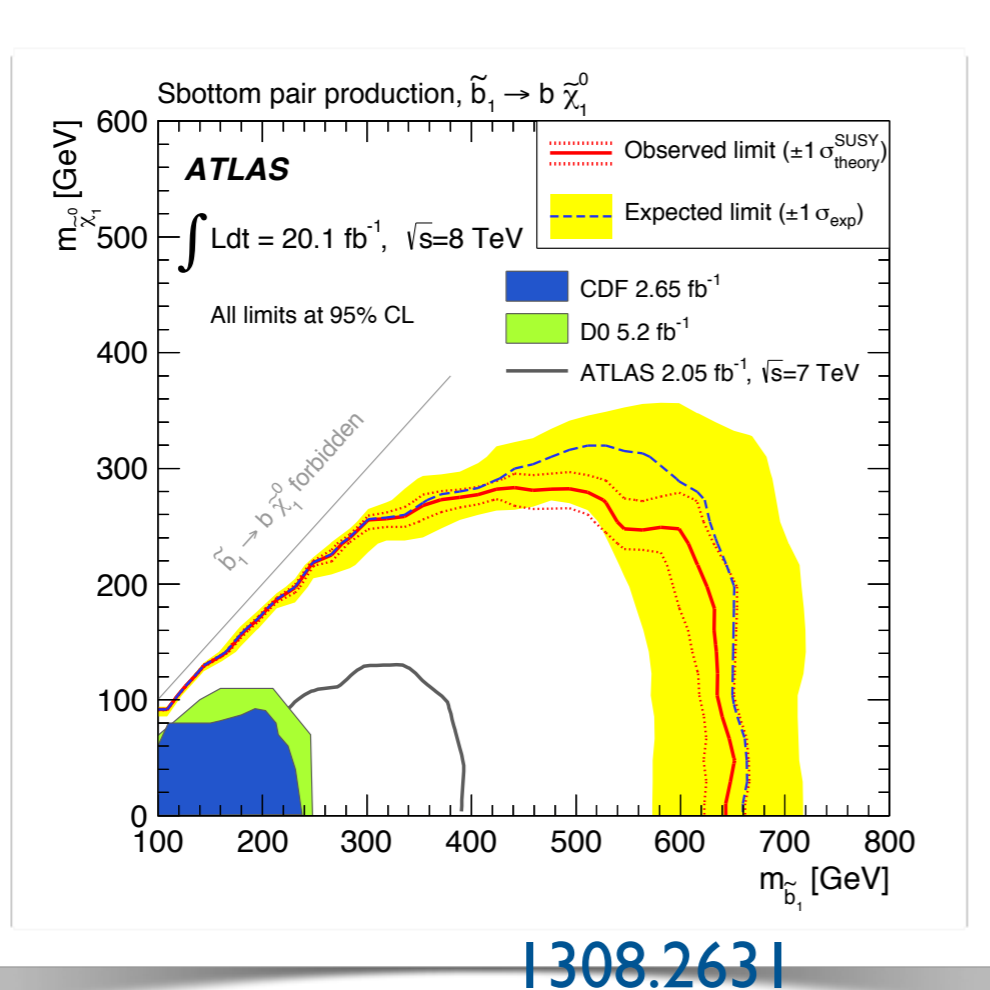
$$\frac{N_{\text{events}}(M_{\text{high}}^2, s_{\text{high}}, L_{\text{high}} [\text{fb}^{-1}])}{N_{\text{events}}(M_{\text{low}}^2, s_{\text{low}}, L_{\text{low}} [\text{fb}^{-1}])} = 1$$

Solve for M_{high}^2 .

Check: ATLAS direct sbottom

[95%CL, $m_{LSP} = 0$ GeV]

7 TeV, 2.05 fb^{-1} >400 GeV
ATLAS



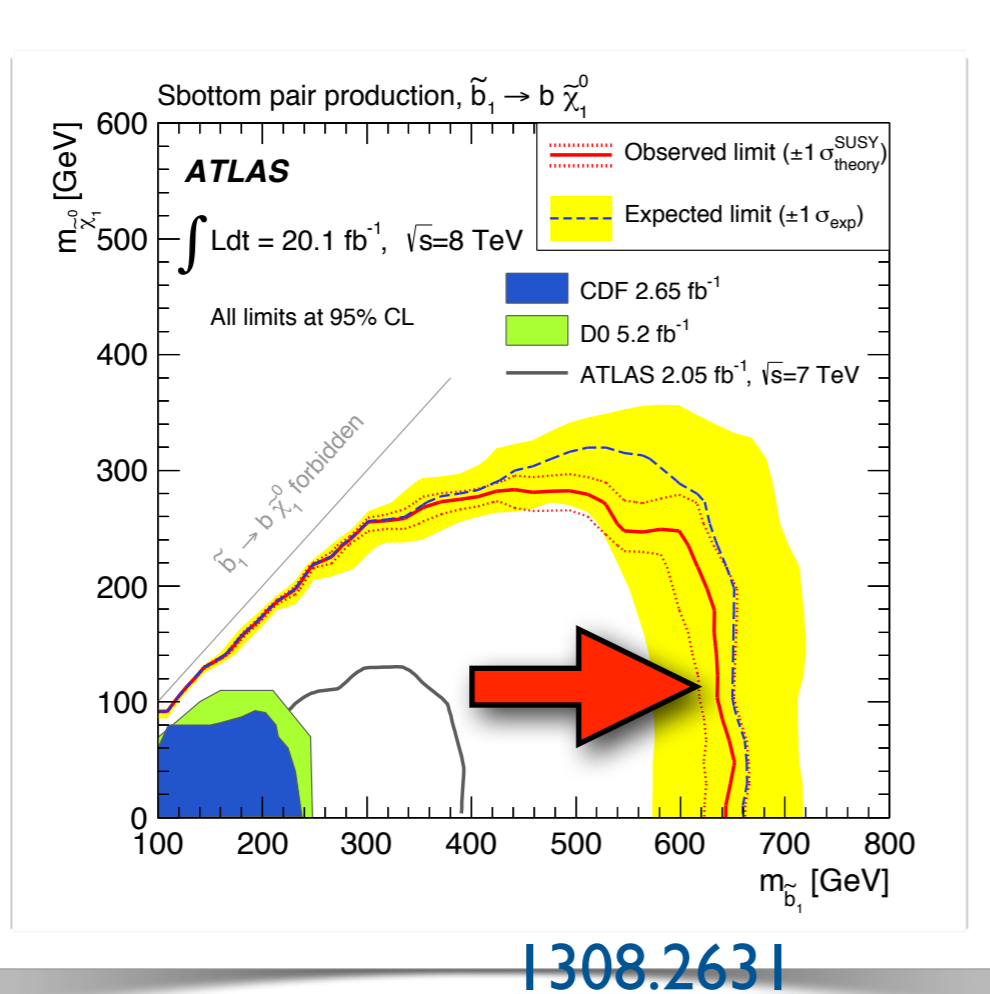
[1308.2631](#)

Check: ATLAS direct sbottom

[95%CL, $m_{LSP} = 0$ GeV]

7 TeV, 2.05 fb^{-1} >400 GeV
ATLAS

8 TeV, 20.1 fb^{-1} >640 GeV
ATLAS



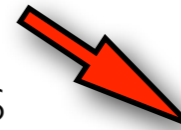
Check: ATLAS direct sbottom

[95%CL, $m_{LSP} = 0$ GeV]

7 TeV, 2.05 fb^{-1}

>400 GeV

ATLAS



8 TeV, 20.1 fb^{-1}

>640 GeV

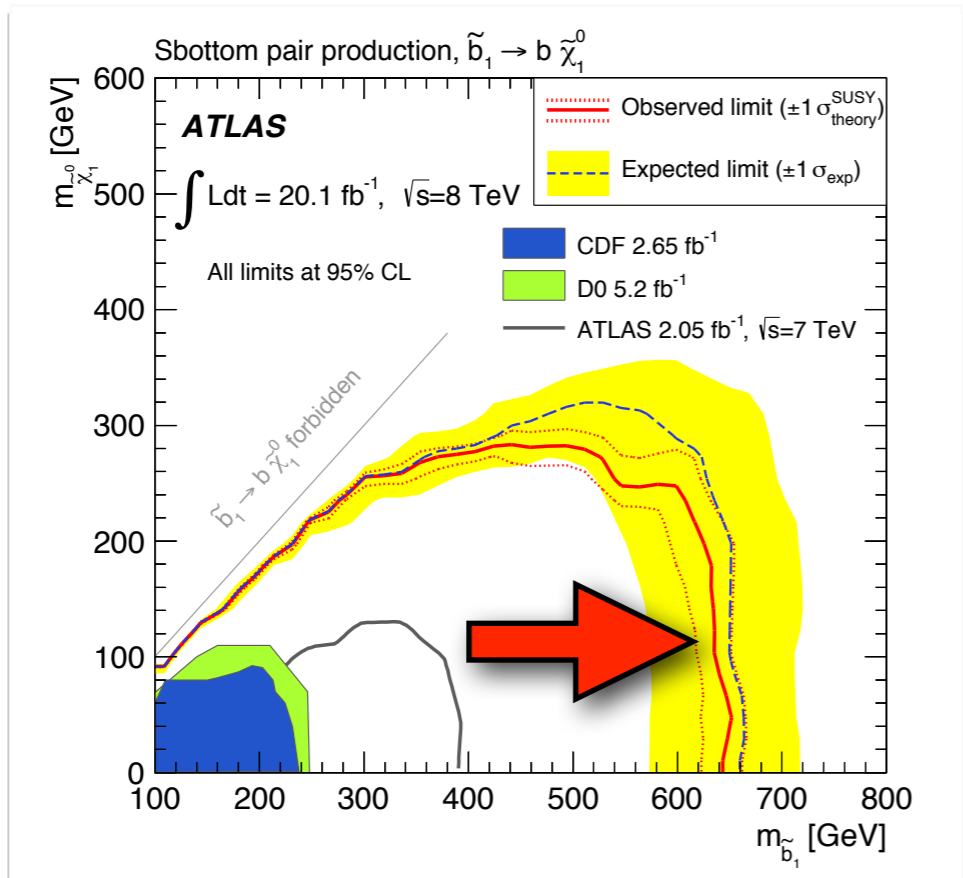
ATLAS

>620 GeV



parton lumi

Parton luminosity
estimate using gg-pdf's



[1308.2631](#)

8 TeV, 20 fb⁻¹ → 14 TeV, 300 fb⁻¹

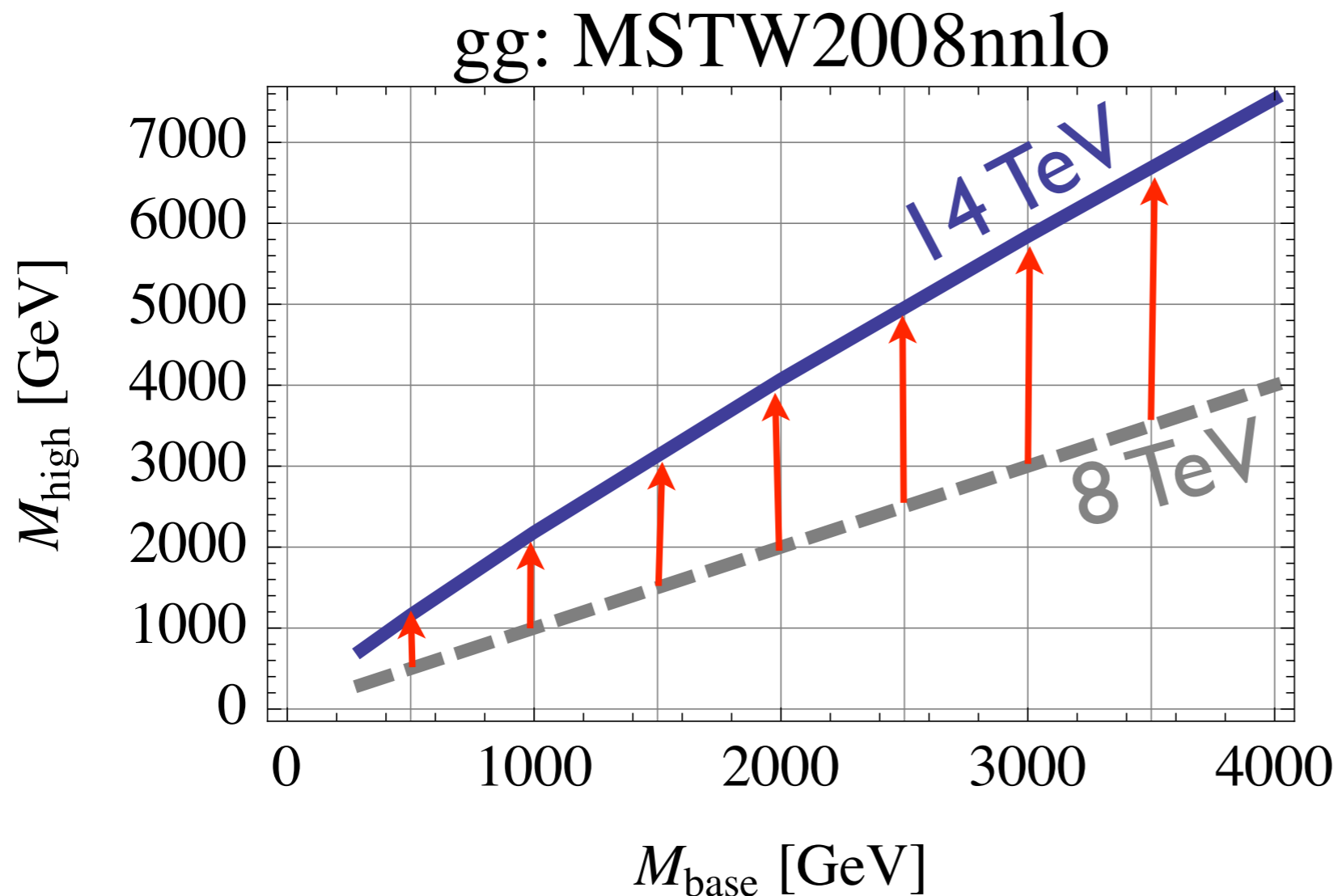
G. Salam, AW

Gluon initiated processes (e.g. direct stop/sbottom)

8 TeV, 20 fb⁻¹ → 14 TeV, 300 fb⁻¹

G. Salam, AW

Gluon initiated processes (e.g. direct stop/sbottom)

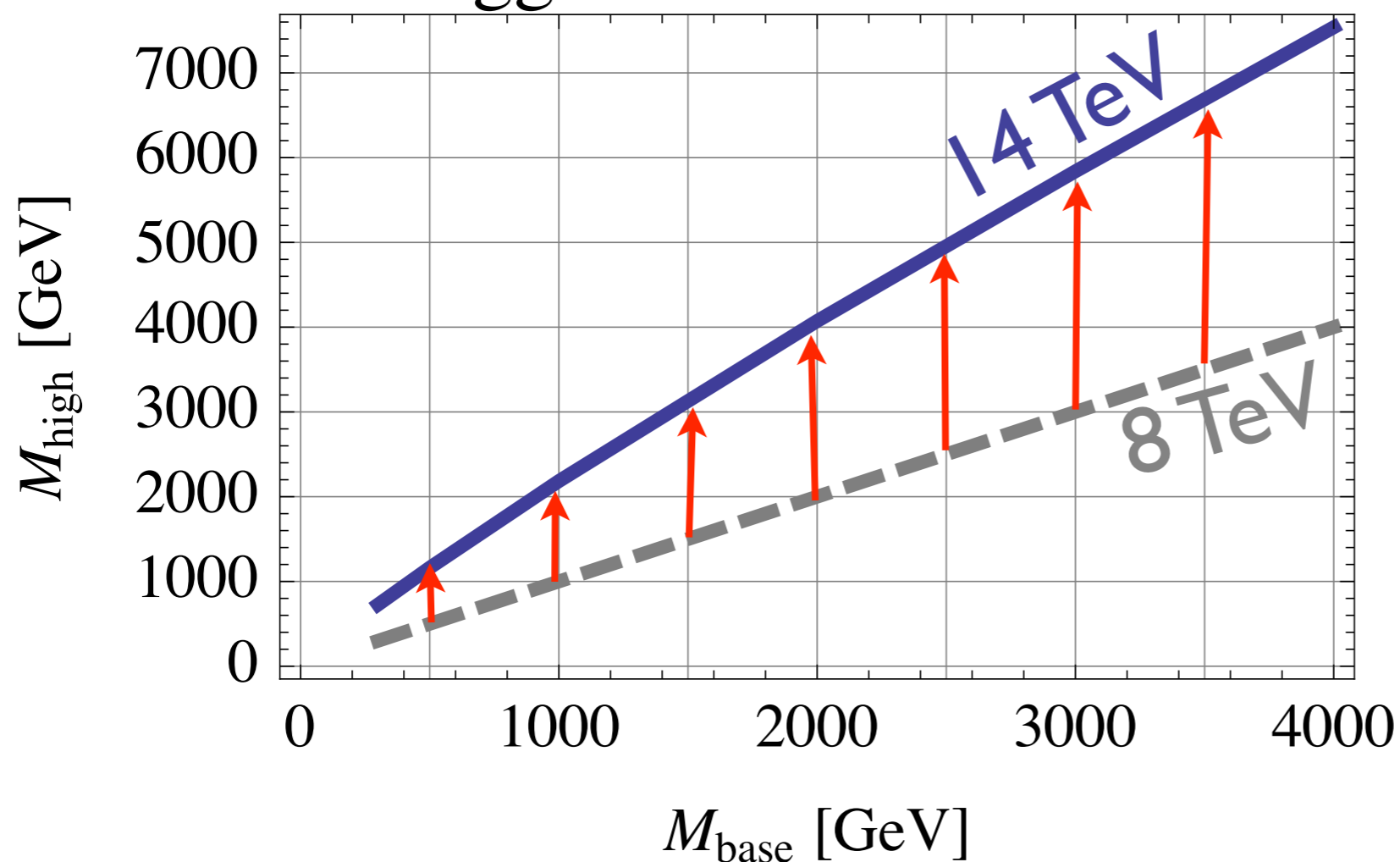


8 TeV, 20 fb⁻¹ → 14 TeV, 300 fb⁻¹

G. Salam, AW

Gluon initiated processes (e.g. direct stop/sbottom)

gg: MSTW2008nnlo



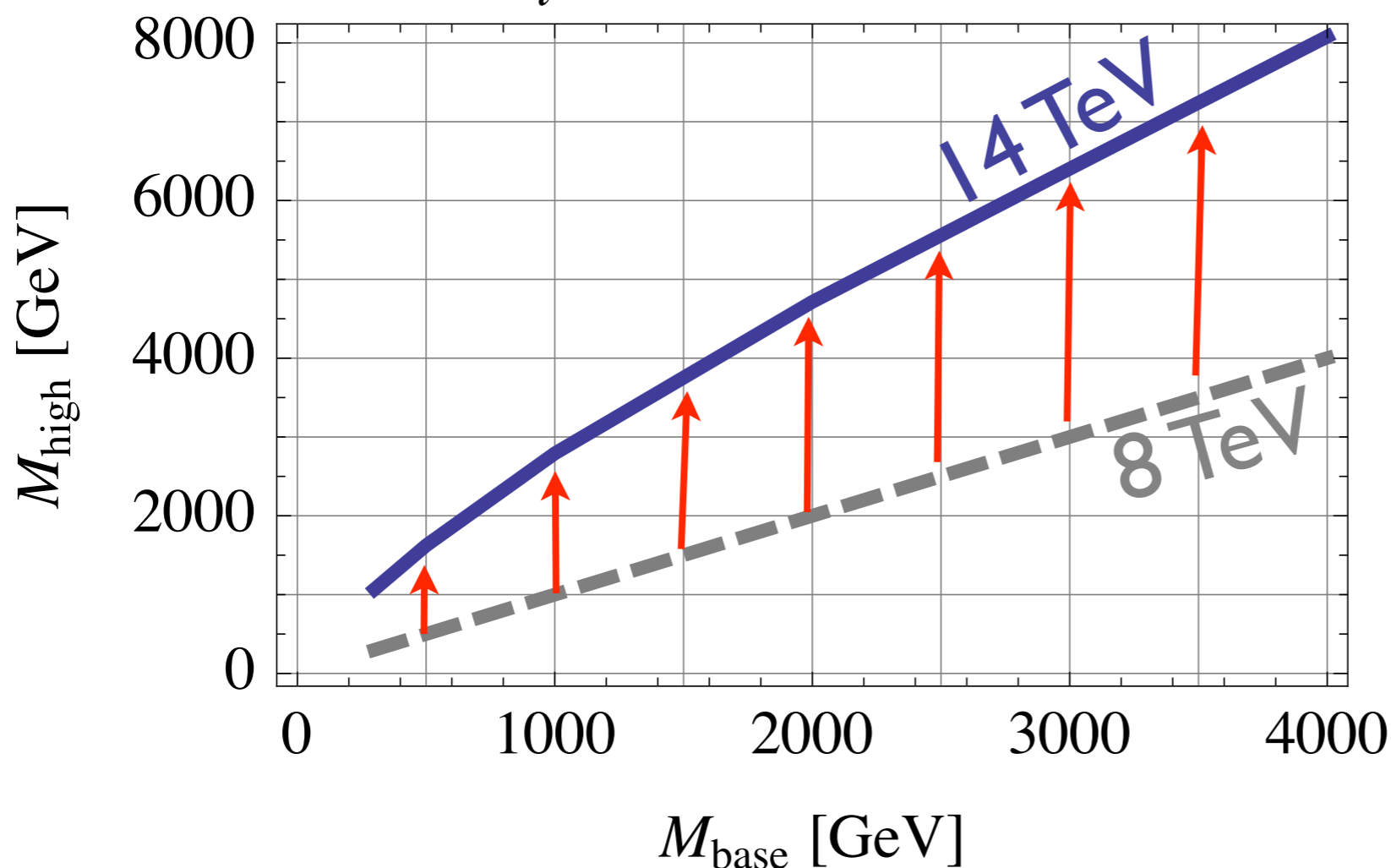
	20 fb ⁻¹	300 fb ⁻¹
	$M_{7\text{TeV}}$	$M_{14\text{TeV}}$
	500	1150
	1000	2200
	2000	4100

8 TeV, 20 fb⁻¹ → 14 TeV, 300 fb⁻¹

G. Salam, AW

$\sum_i \bar{q}_i q_i$ initiated processes (e.g. Z', ...)

$\Sigma \bar{q}_i q_i$ MSTW2008nnlo



	20 fb ⁻¹	300 fb ⁻¹
$M_{7\text{TeV}}$	$M_{14\text{TeV}}$	
500	1600	
1000	2800	
2000	4750	

High luminosity LHC

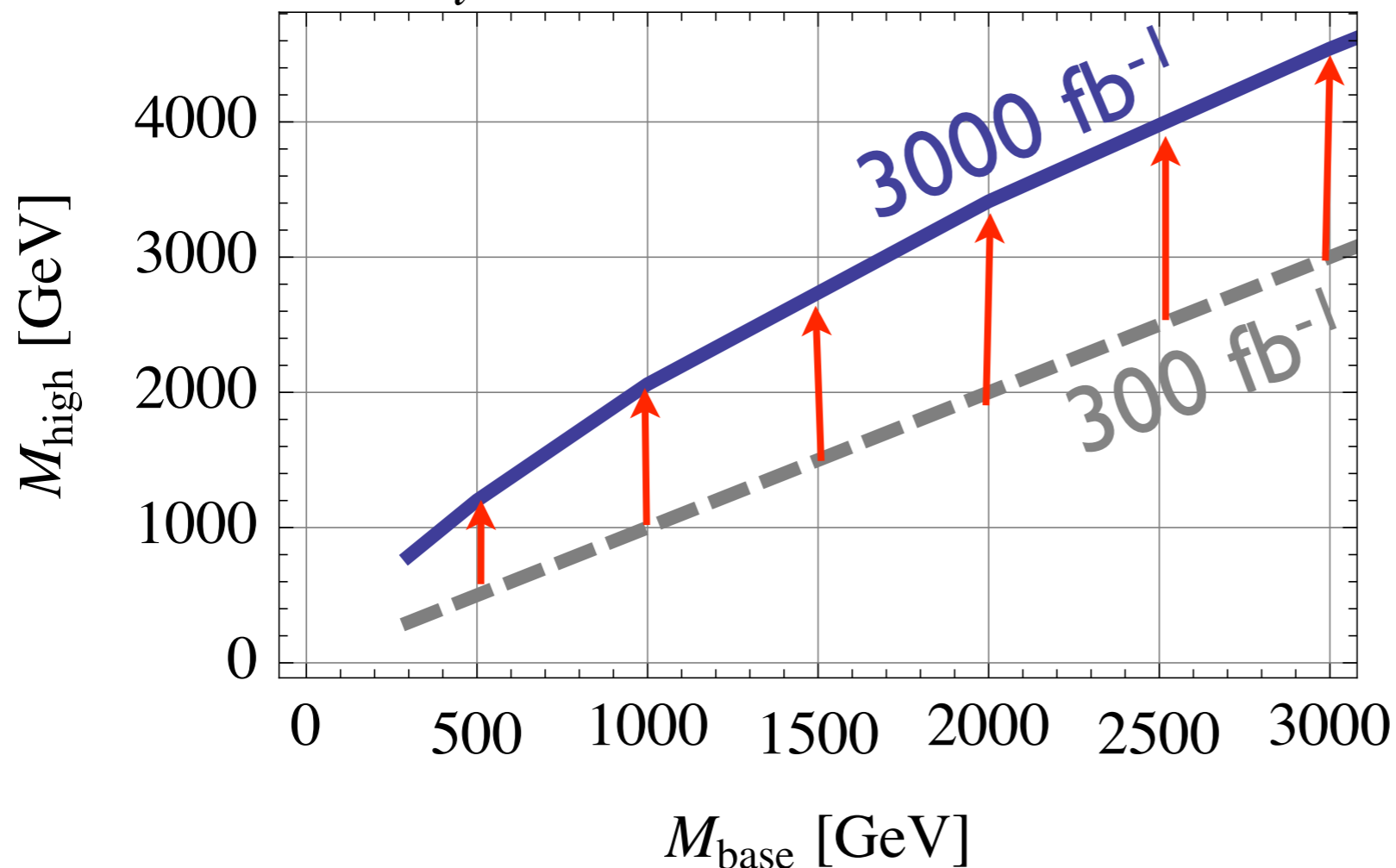
$300 \text{ fb}^{-1} \rightarrow 3000 \text{ fb}^{-1} @ 14 \text{ TeV}$

300 fb⁻¹ → 3000 fb⁻¹ @ 14 TeV

G. Salam, AW

$\sum_i \bar{q}_i q_i$ initiated processes (e.g. Z', ...)

$\Sigma \bar{q}_i q_i, 300 \text{ fb}^{-1} \rightarrow 3000 \text{ fb}^{-1}$



	14 TeV	14 TeV
	$M_{300 \text{ ifb}}$	$M_{3000 \text{ ifb}}$
	300	800
	500	1200
	1000	2050
	2000	3400

300 fb⁻¹ → 3000 fb⁻¹ @ 14 TeV

G. Salam, AW

$\sum_i \bar{q}_i q_i$ initiated processes (e.g. Z', ...)

$\Sigma \bar{q} q$ 300 fb⁻¹ , 3000 fb⁻¹

Impressive relative improvement
for lower x-sec processes
(e.g. EW production, rare decays,
difficult spectra, ...)

	14 TeV	14 TeV
	$M_{300\text{ ifb}}$	$M_{3000\text{ ifb}}$
	300	800
	500	1200
	1000	2050
	2000	3400

0 500 1000 1500 2000 2500 3000

M_{base} [GeV]

Outlook

The battle for a natural resolution of the hierarchy problem goes on

LHC14 will be decisive: $2\times\text{Energy} \rightarrow 4\times\text{Tuning}$

Flavor non-trivial signals to be explored,
charm tagging, bumps in sub-leading jets

‘Absence of evidence is not evidence of absence’,
still: some experimental guidance would be nice.

