



The next-generation space infrared astronomy mission

Space Infrared Telescope for Cosmology and Astrophysics

Windows on the Universe, ICISE,
RENCONTRES DU VIETNAM
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H. Matsuura, T. Nakagawa, Y. Kawakatsu (ISAS/JAXA)
for SPICA Team

Institute of Space & Astronautical Science, Japan Aerospace Exploration Agency



Wavelength

Submm
Far-IR
100 μ m

Development of Space IR Astronomy

Near-IR
10 μ m
Mid-IR



IRAS (surveyor)
1983 (USA/UK/Netherlands)
Diameter: 0.57m ϕ
Wavelength: 8~100 μ m



IRTS (surveyor)
1995 (Japan/US)
Diameter: 0.15m ϕ
Wavelength: 1~700 μ m

Initial Survey Era



ISO(Observatory)
1995-1998 (ESA)
Diameter: 0.60m ϕ
Wavelength: 2.4~240 μ m

AKARI (surveyor)
2006- 2011 (Japan)
Diameter: 0.685m ϕ
Wavelength: 2~180 μ m



Herschel (observatory)
2009- 2013 (ESA)
Diameter 3.5m ϕ
Wavelength: 55~700 μ m

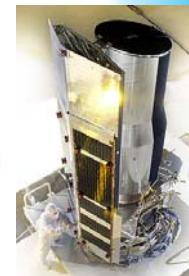
Resolution of
Evolution of
Matter in the
Universe



SPICA (Observatory)
2023- (Japan (Asia)& Europe)
Diameter: 3.2m ϕ
Wavelength: 5~210 μ m



New Generation All-Sky Survey



SPITZER (Observatory)
2003- (USA)
(2009 Liq. He boiling-off)
Diameter: 0.85m ϕ 、Wavelegh: 3~180 μ m



JWST (Observatory)
2018- (USA)
Diameter: 6.5m ϕ
Wavelength: 0.6~28 μ m

1980

1990

2000

2010

2020

Years

Outline of my Talk

- SPICA Scientific Goals
 - with a few slides on ‘Why IR Observatory in Space?’
- Mission Overview
- Programmatic Status

SPICA Science Goals

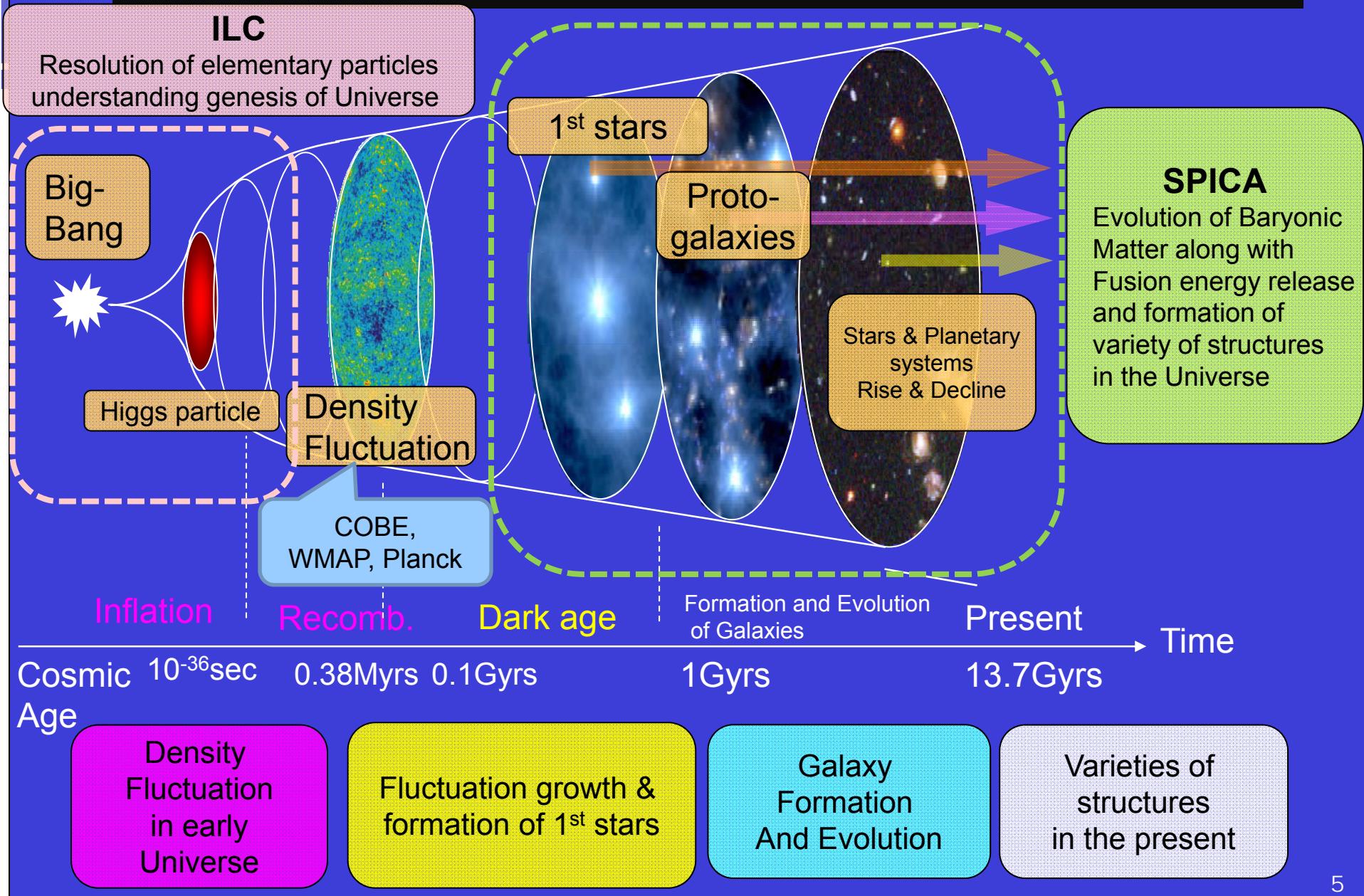
*Unveil the evolution of Baryonic
Matter in the Universe*

*Life cycle of interstellar &
Intergalactic matter*

Planetary Systems Formation

Birth and Evolution of Galaxies

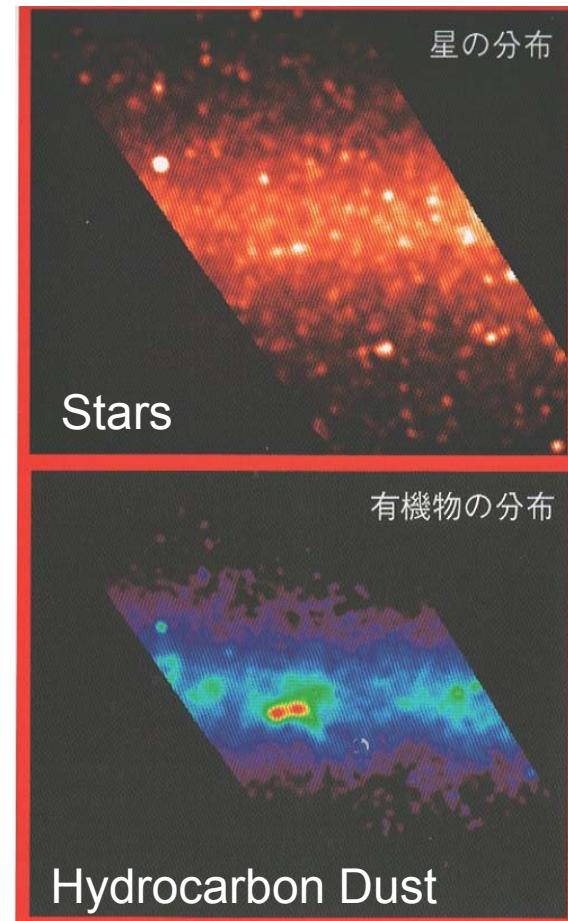
Cosmic History from Genesis of the Universe to Stars & Planetary systems Formation



SPICA reveals Evolution of Baryonic Matter in the Universe

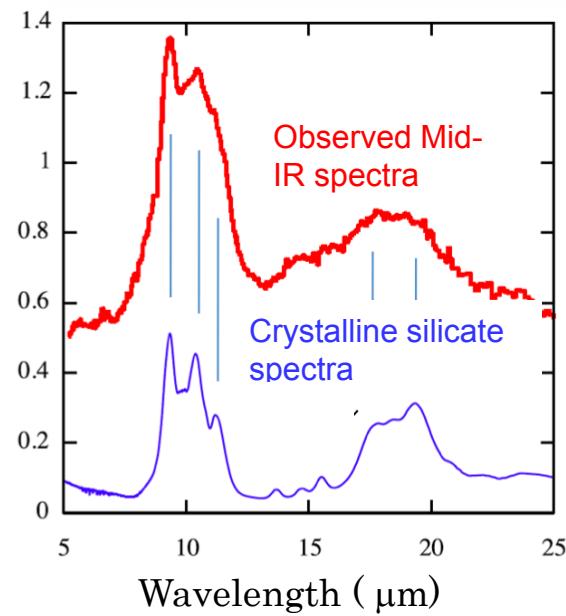
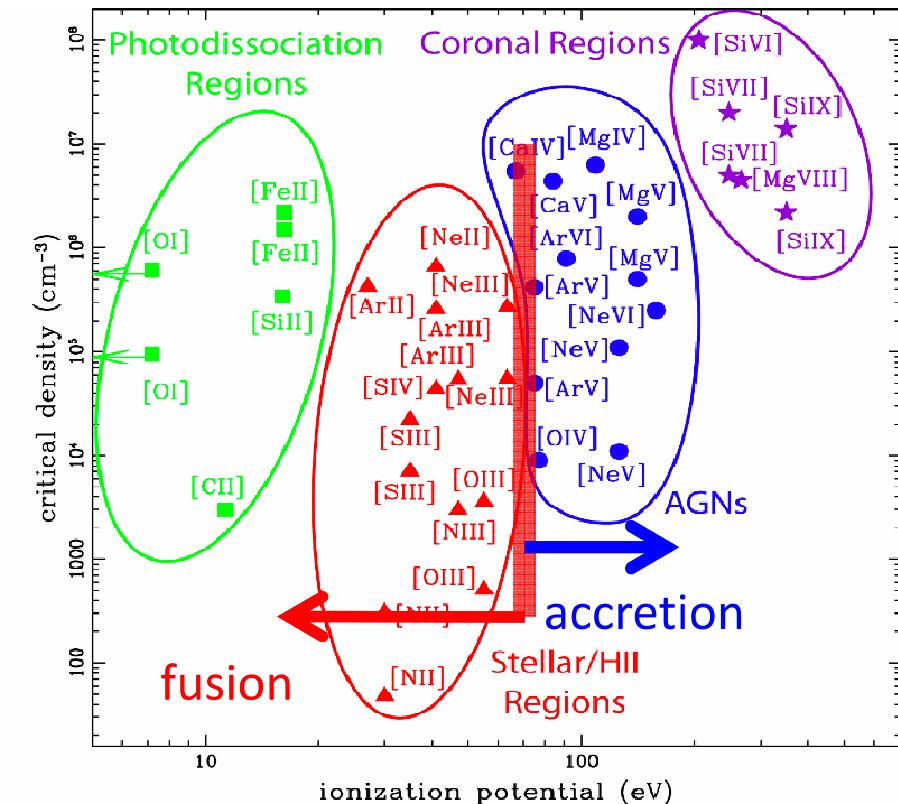
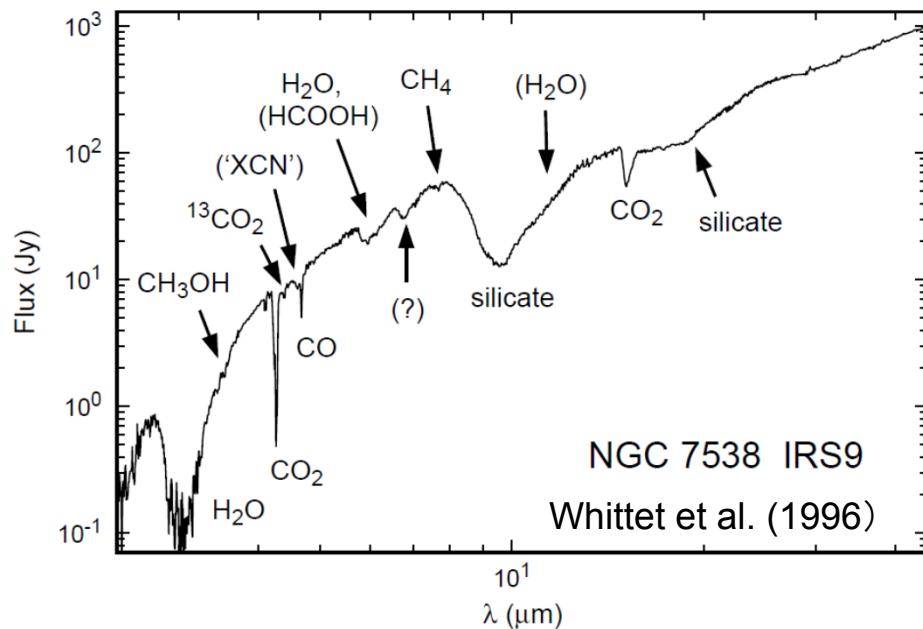
- Early Universe after Big-Bang
 - Only H & He
- Present Universe
 - Various metal elements exist in Gas and Dust (solid state particles)
 - Hydrocarbons are ubiquitously distributed in the Milky-way Galaxy
- Various metal elements were formed in stars, and matters were evolved through the life cycle of the stars

Results from
IRTS (1995)



Why Infrared spectra?

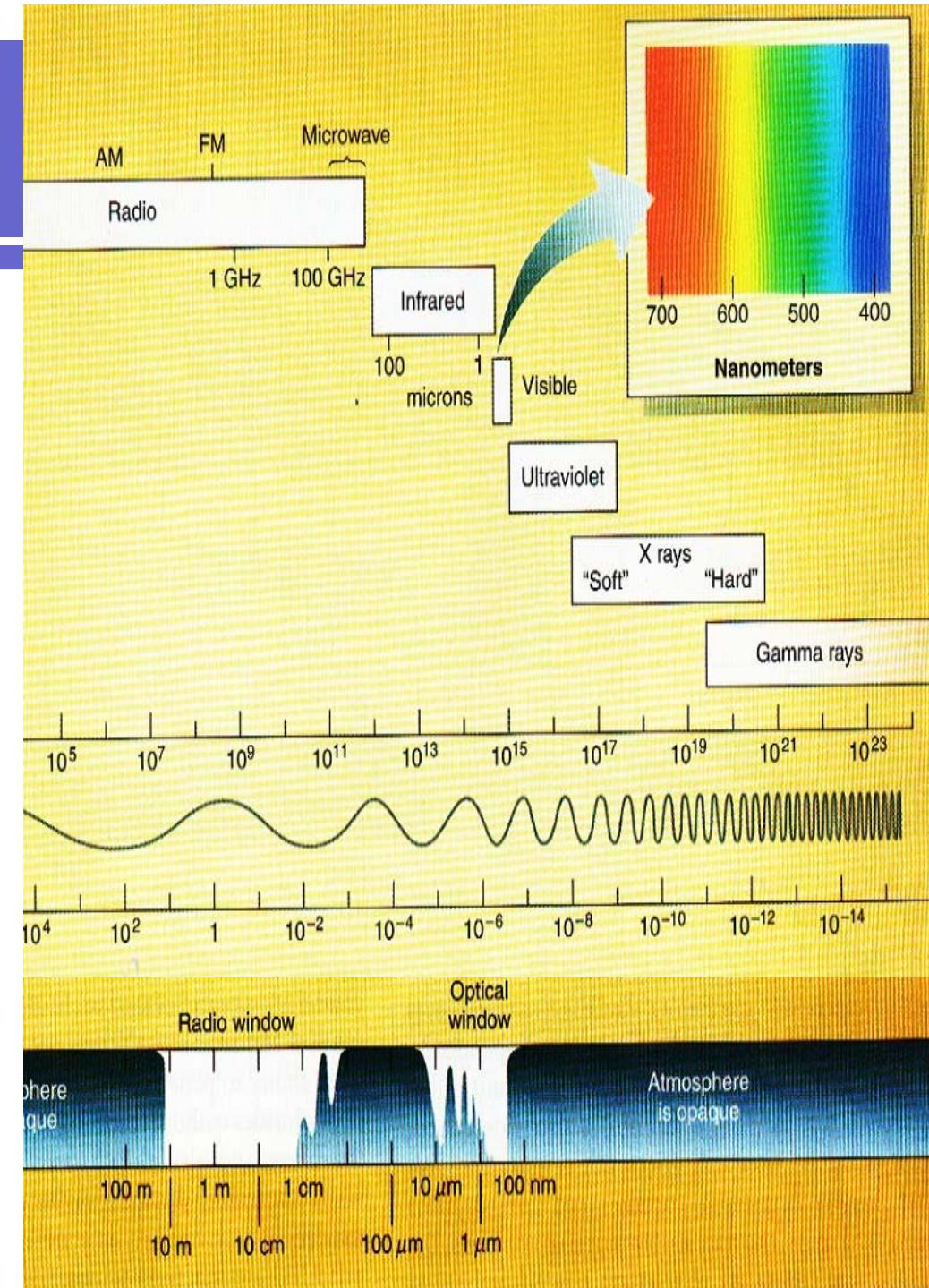
- Abundant spectral lines from atoms, ions and molecules for diagnostic tool.
- Broad-band spectral features provide the information of composition of solid state particles with ices



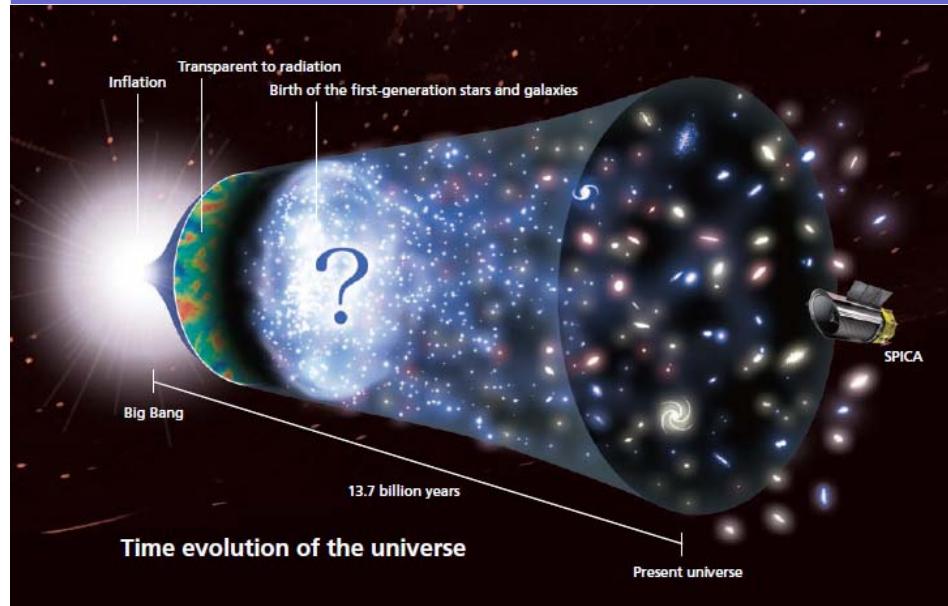
Infrared spectrum of the debris disk around a star HD154014 (Fujiwara et al. 2010)

Observation from
Space is essential:
very limited
atmospheric windows
on the Universe

Earth's atmosphere
blocks most of the
electromagnetic
waves

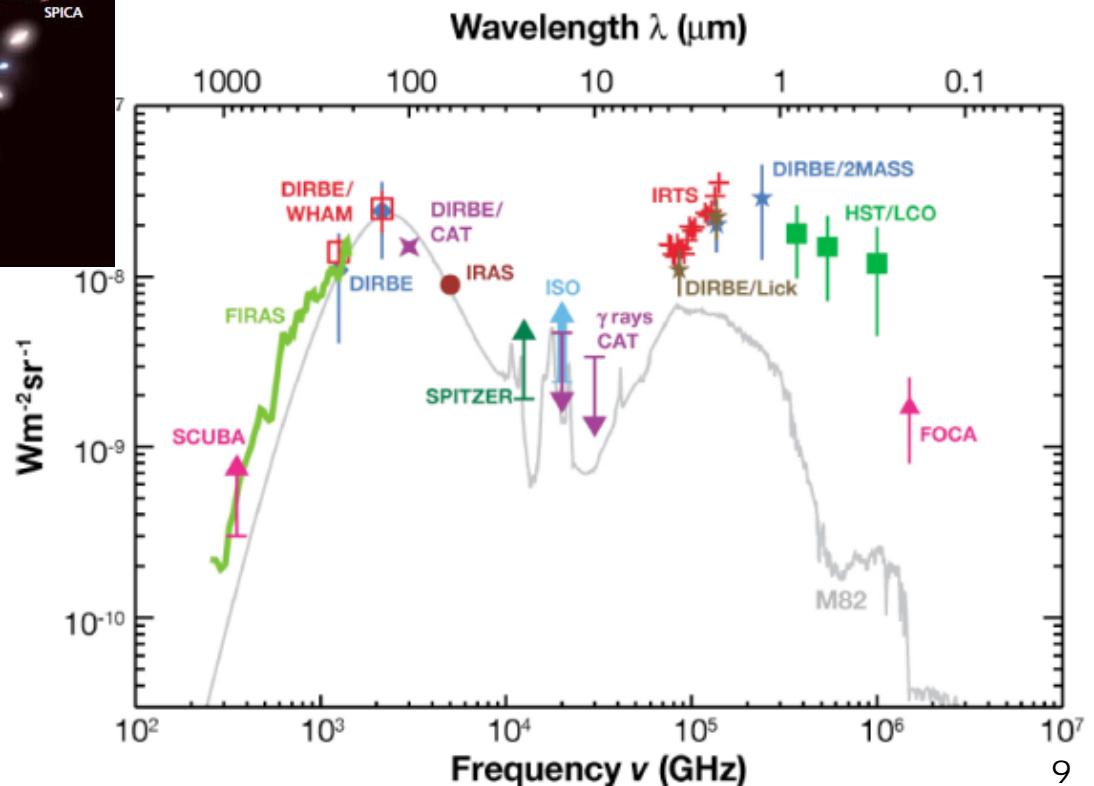


One of Primary Scientific Goals: the birth & evolution of galaxies



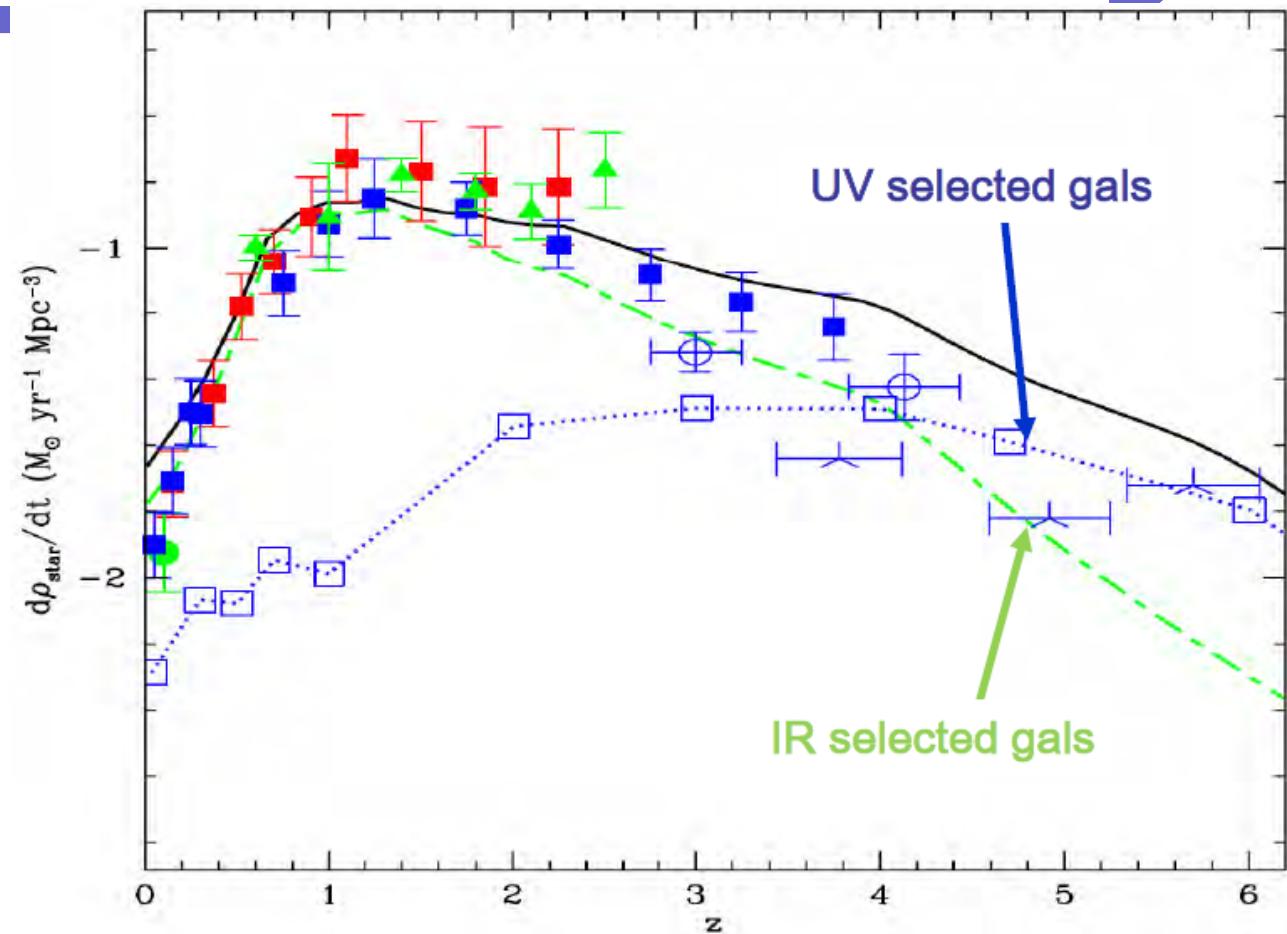
→ **Hidden Universe to be revealed by FIR Observations**

- Cosmic IR background tells us that about half of nuclear fusion energy is hidden by dust

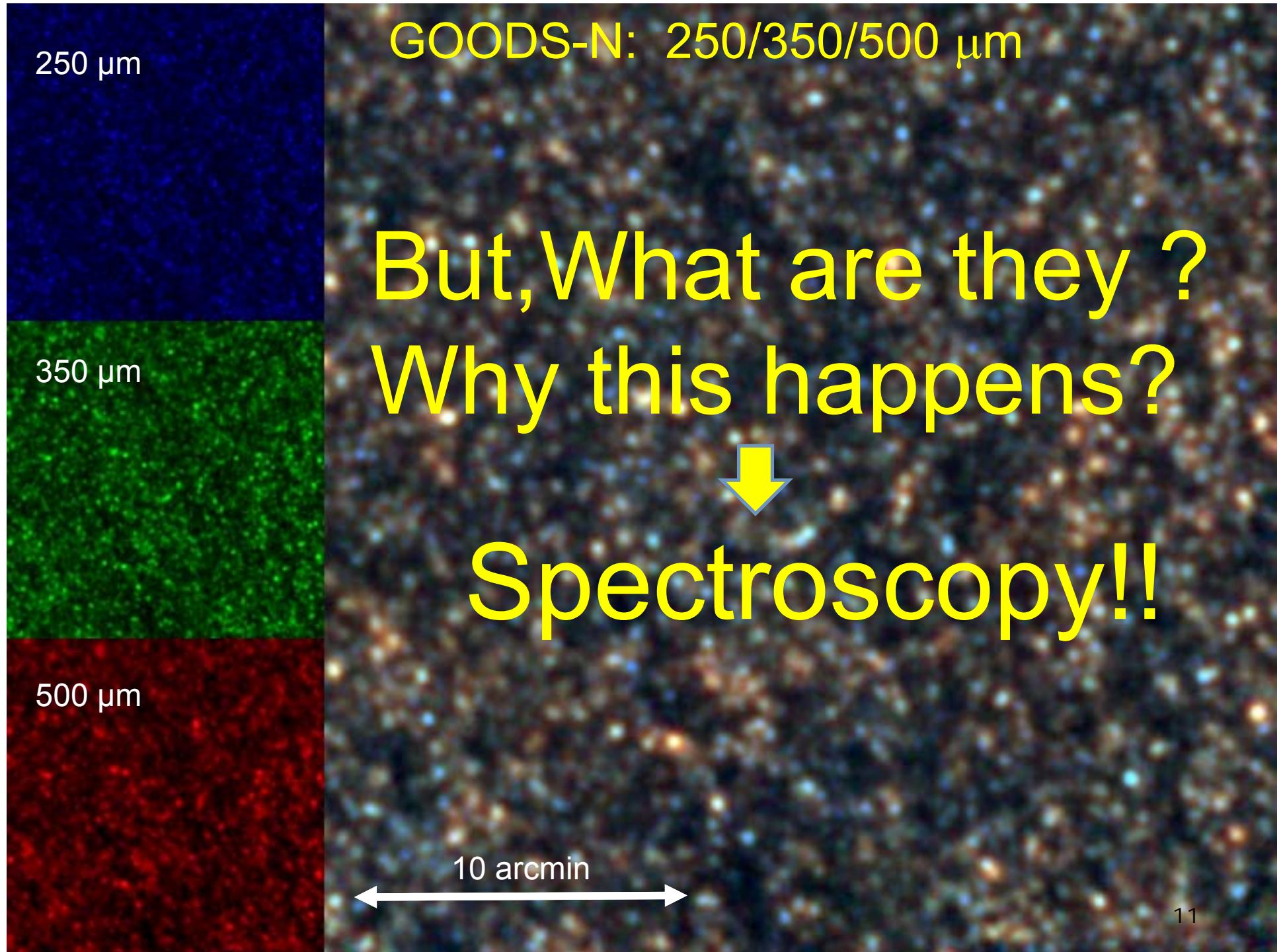


Cosmic Star-formation History revealed with Herschel

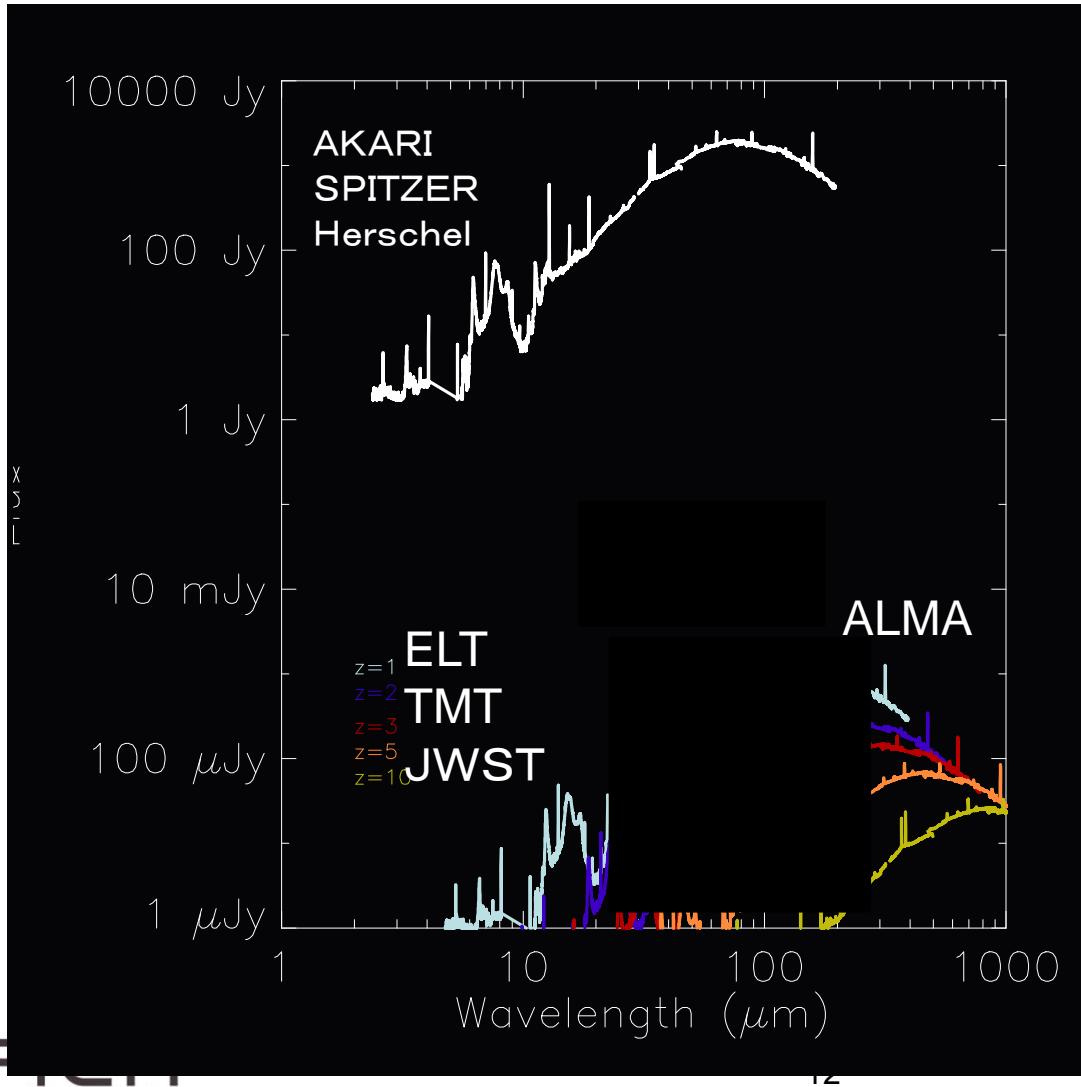
The bulk of SF activity at $z < 3-4$ appears to be produced by strongly dust-extinguished galaxies



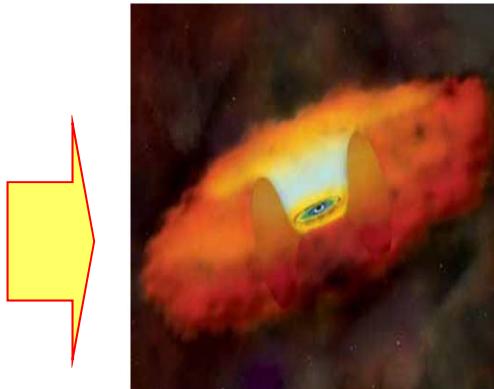
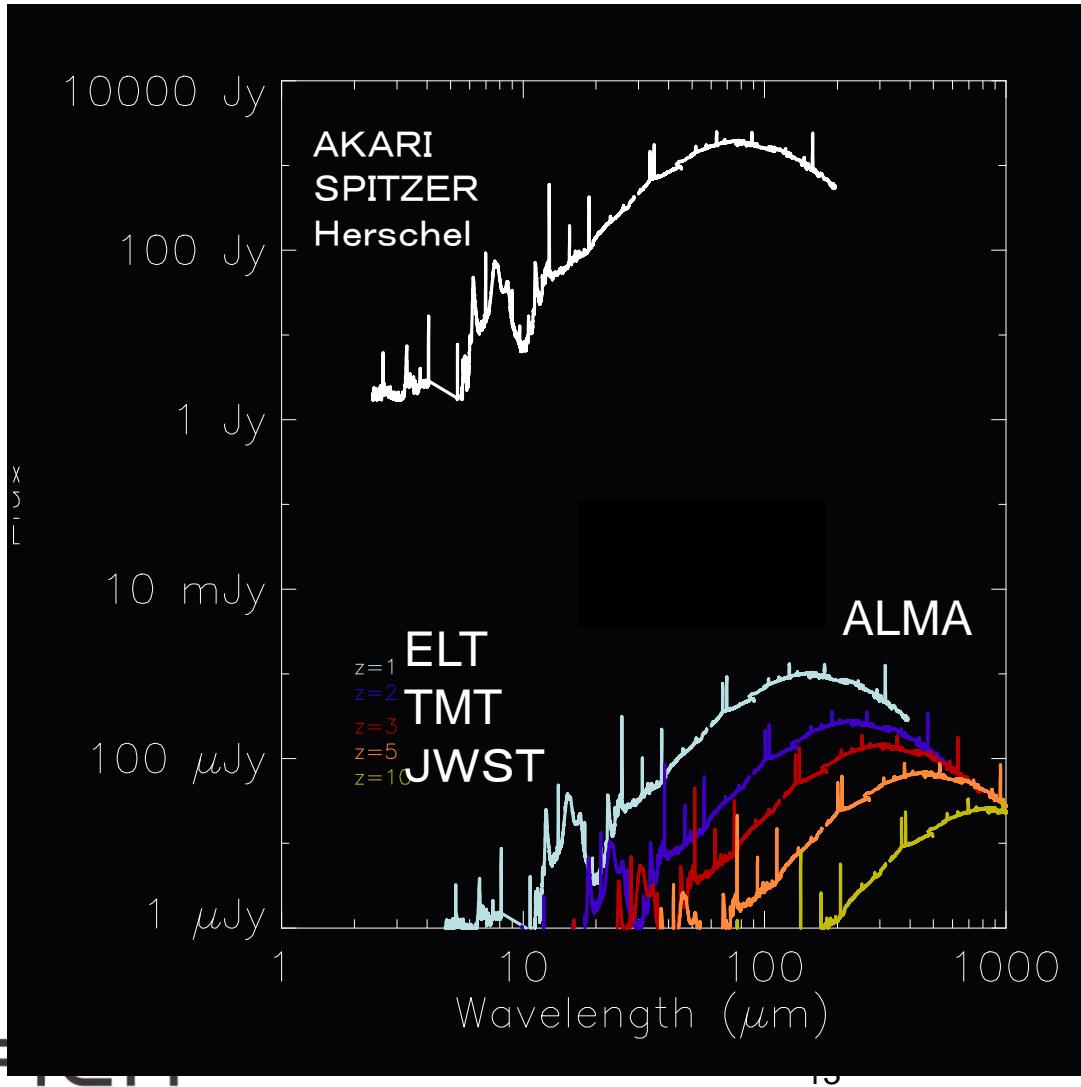
Herschel: COSMOS+WIDE
Vaccari et al. 2013; Marchetti et al. 2013



Characterizing Hidden Nature of Galaxies



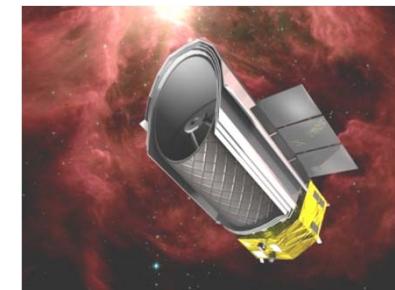
Characterizing Hidden Nature of Galaxies



Revealing true nature
of obscured galaxies,
Starburst and /or Super
Massive Black holes

**SPICA
observations
are essential**

Mission Overview



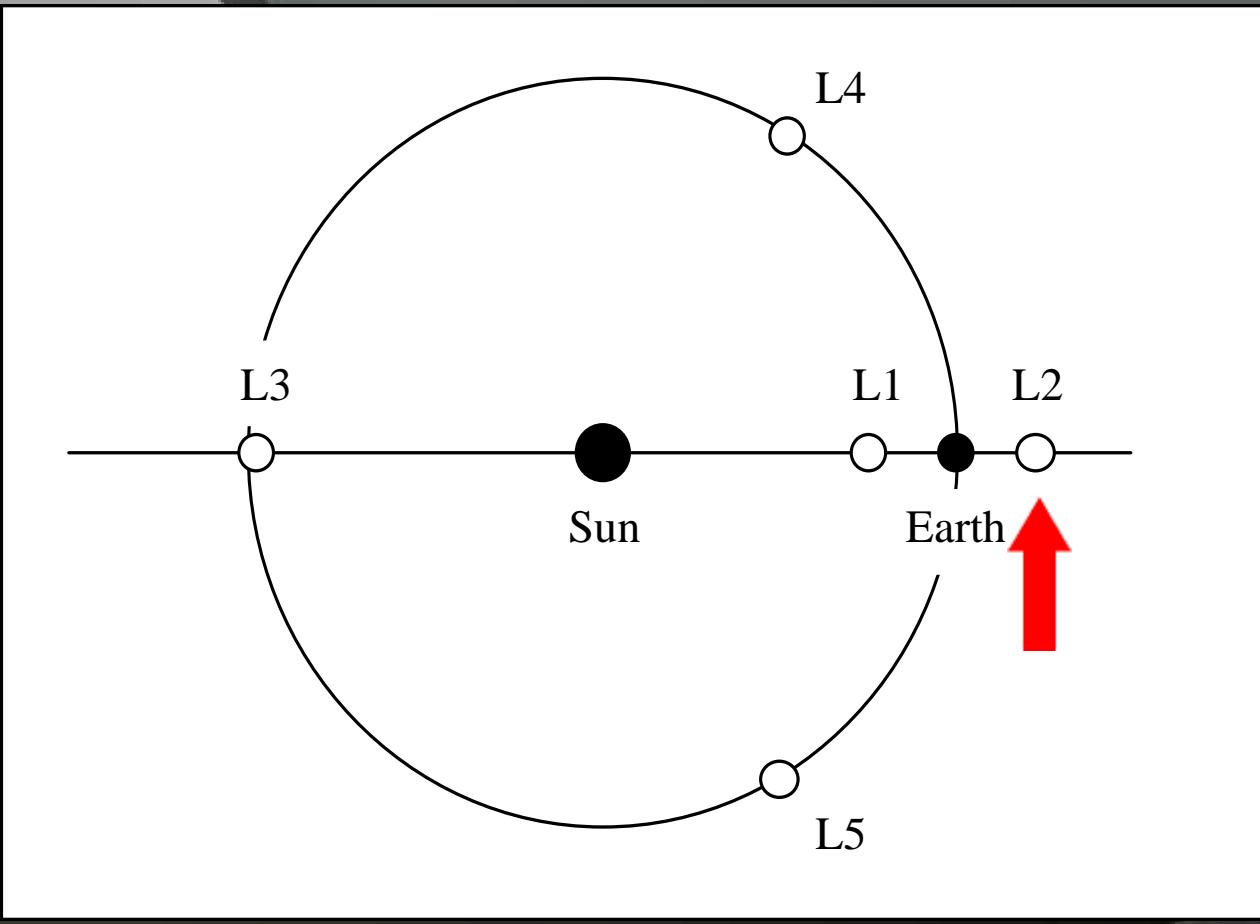
*Unveil the evolution of Matter
In the Universe*



SPICA Mission Overview

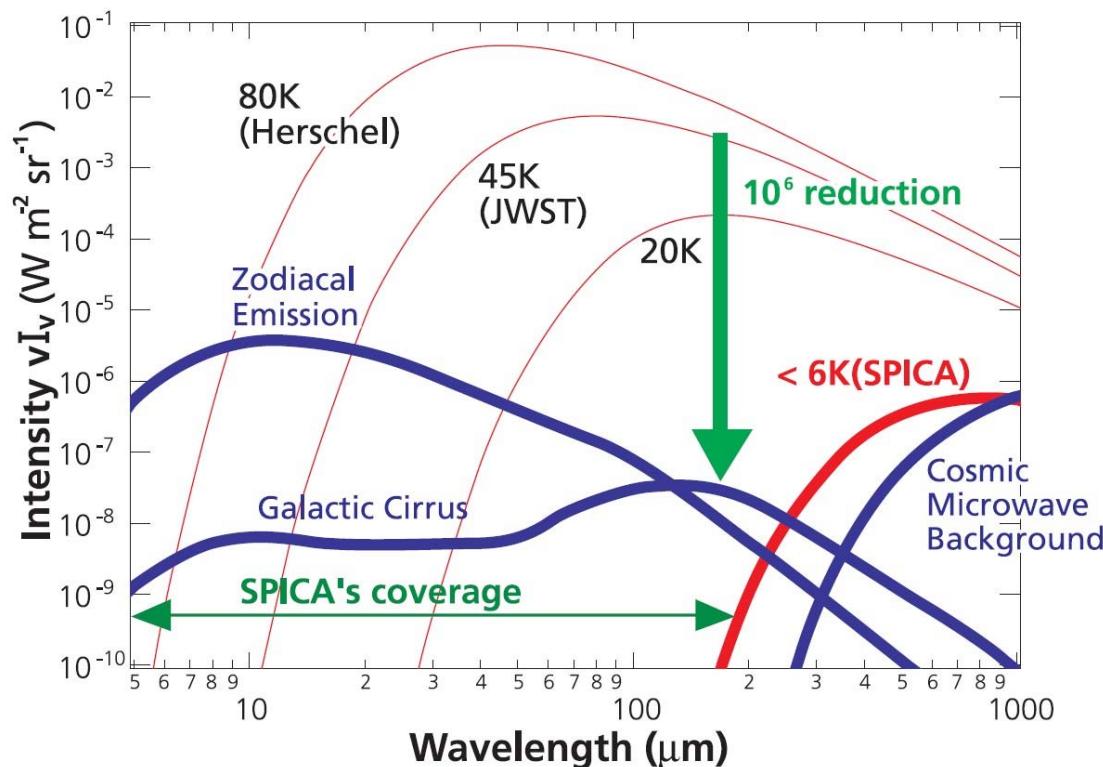
- Telescope 3.2m (EPD 3.0m), 6K
- Orbit: Sun-Earth L2 Halo
- Mission Life: 3 years (nominal), 5 years (goal)
- Weight: 3.7 t
- Launch: 2022
- Next-Generation Infrared Space Observatory with superior Sensitivity and good spatial resolution in mid- and Far-IR (5–210 μ m)
- International Mission
 - Japan, Europe, Korea, Taiwan (USA)

SPICA is Japan (Asian)-led space observatory operated in S-E L2 point,
1.5 Million km away



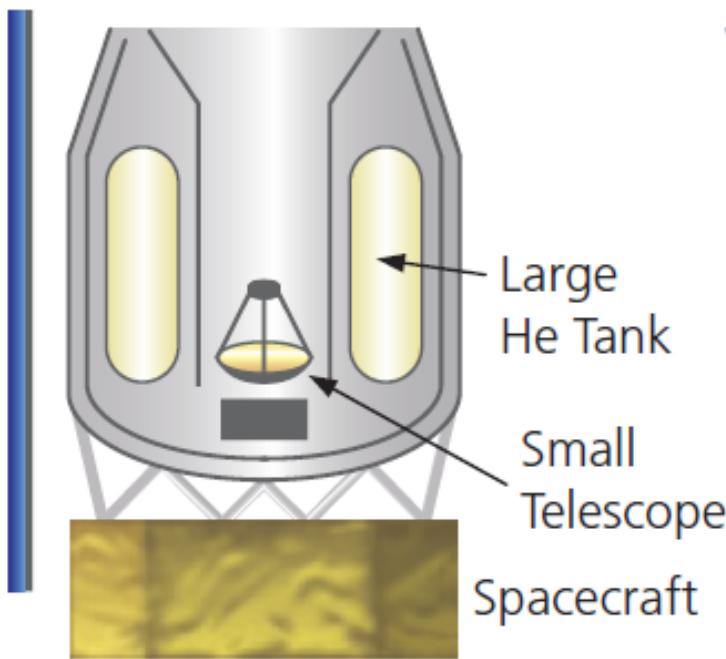
Requirements: Cooled Telescope

- $T < 10$ K is required to improve sensitivity
 - Background Radiation can be reduced by a factor of **one million** !

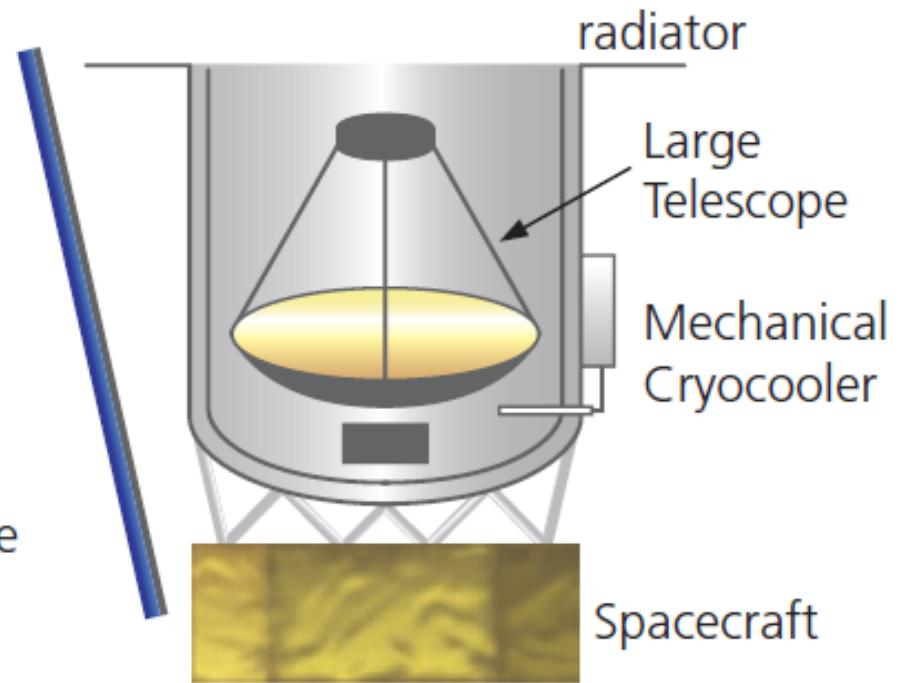


Cryogen-free mission

Today's Space Telescopes

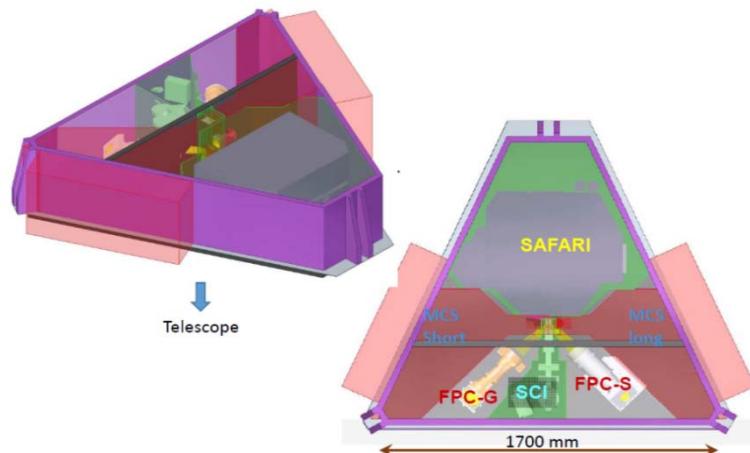
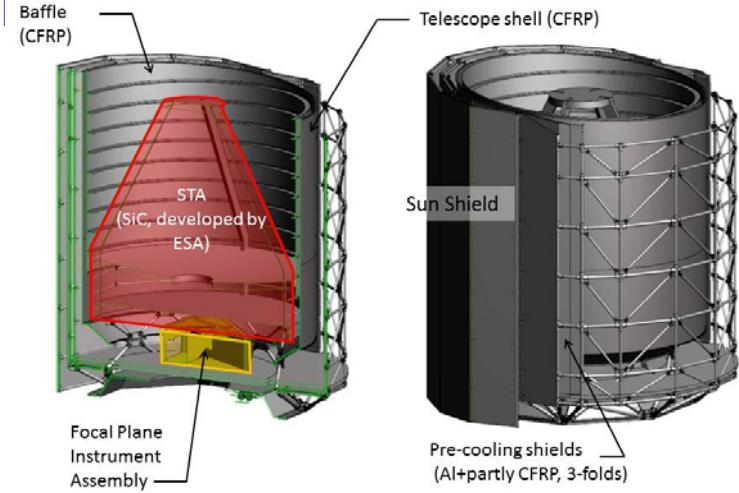


SPICA new design



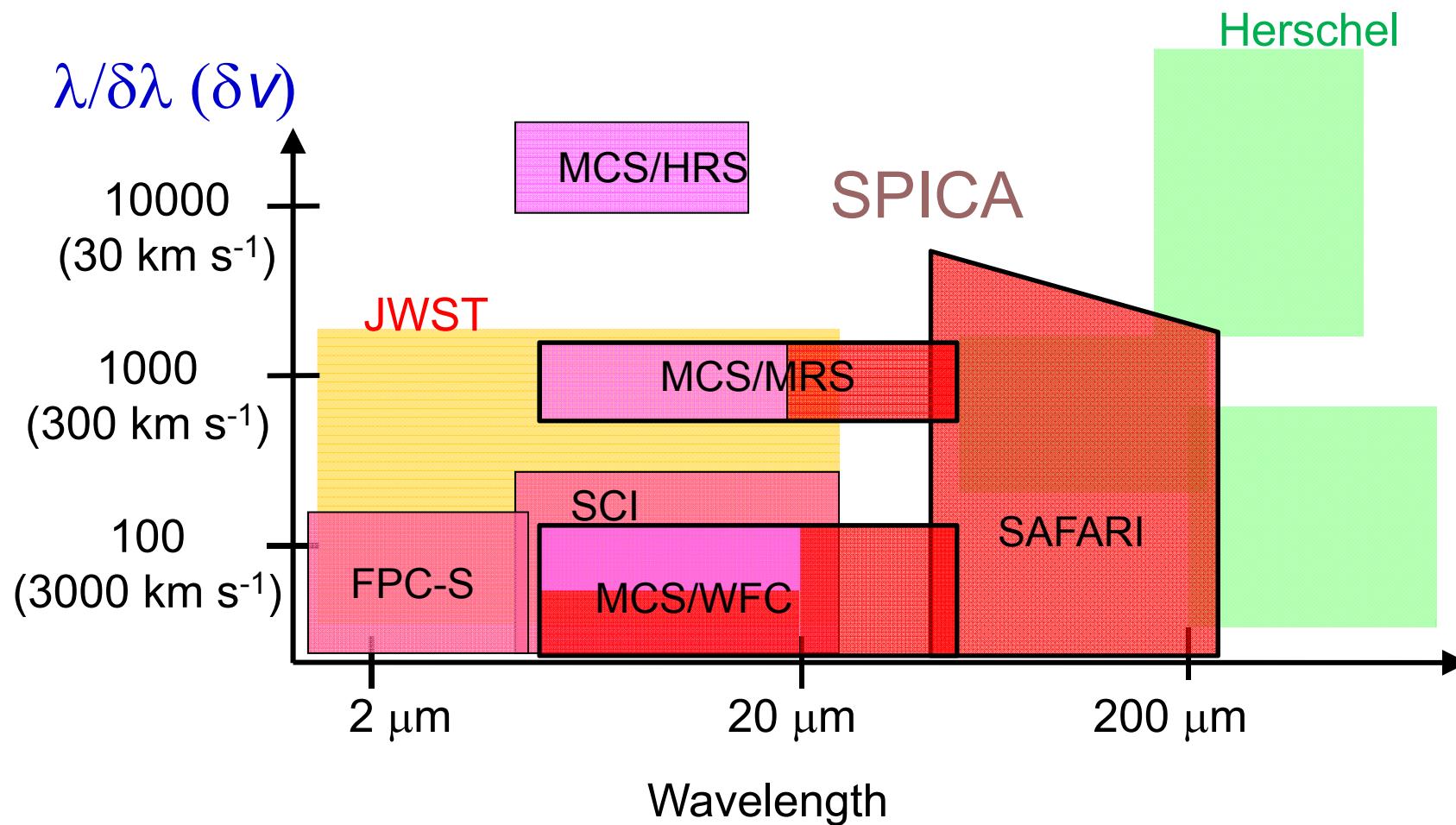
Lighter and Larger

SPICA Focal Plane Instruments

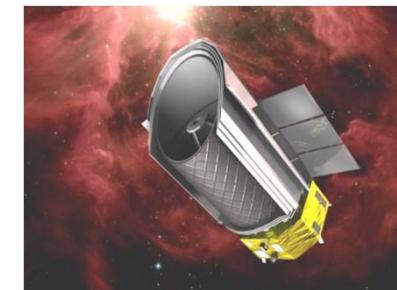


- **SAFARI**
 - Far-infrared imaging spectrometer
 - P.I. SRON (Netherlands) with SAFARI Consortium
- **MCS**
 - Mid-infrared camera & spectrometer
 - P.I. JAXA, Universities, and ASIAA (Taiwan)
- **SCI**
 - SPICA coronagraphic instrument
 - P.I. JAXA with Nagoya Univ.
- **FPC**
 - Near-infrared camera and spectrometer
 - P.I. KASI (Korea)

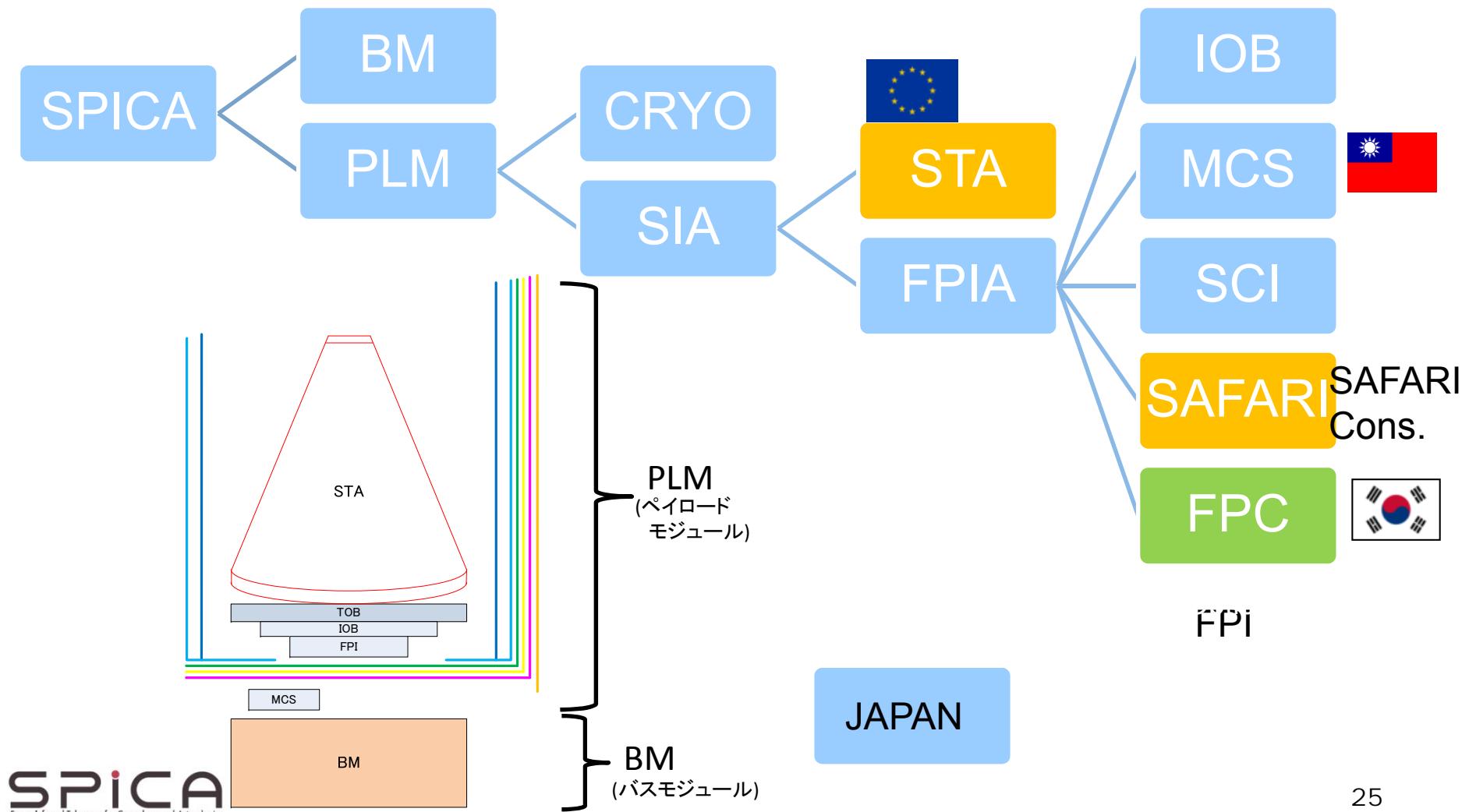
Focal Plane Instruments: baseline Wavelength coverage vs Resolving Power



Programmatic Aspects



International Collaboration



SPICA Project Status

- JAPAN (JAXA)
 - Official SPICA Preproject team (Phase A & B1) established in 2008
 - Planning SDR and Phase-up Review (final approval) in FY2013
- ESA
 - Assessment Study completed
 - Supporting I/F definition
- SAFARI Consortium (PI: SRON)
 - SRON got funded ! (90% of that required)
 - Dedicated team has been working actively
- Korea Status (PI KASI)
 - Official Study Team formed with KASI as PI
- Taiwan (PI: ASIAA)
 - Concrete collaboration started.
- US Status (PI: TBD)
 - Assessment Study by 3 teams funded by NASA in 2010
 - Strong recommendation in the US Decadal Survey in 2010

**Final Approval
Required !
Strong Support
by International
Community is
essential**

International SPICA Team

- 17 countries, regions and one International org.



Summary

- What's SPICA: next-generation space IR observatory, with cold (<6K) 3.2m diameter telescope
- *Much higher sensitivity in the thermal infrared than Herschel Space Observatory*
- SPICA can explore the evolution of Baryonic matter of various form in the history of the Universe
- The Japanese (Asian)-led international mission with ESA and European countries; international community support is essential !



SPICA
Space Infrared Telescope for Cosmology and Astrophysics

Space Odyssey